

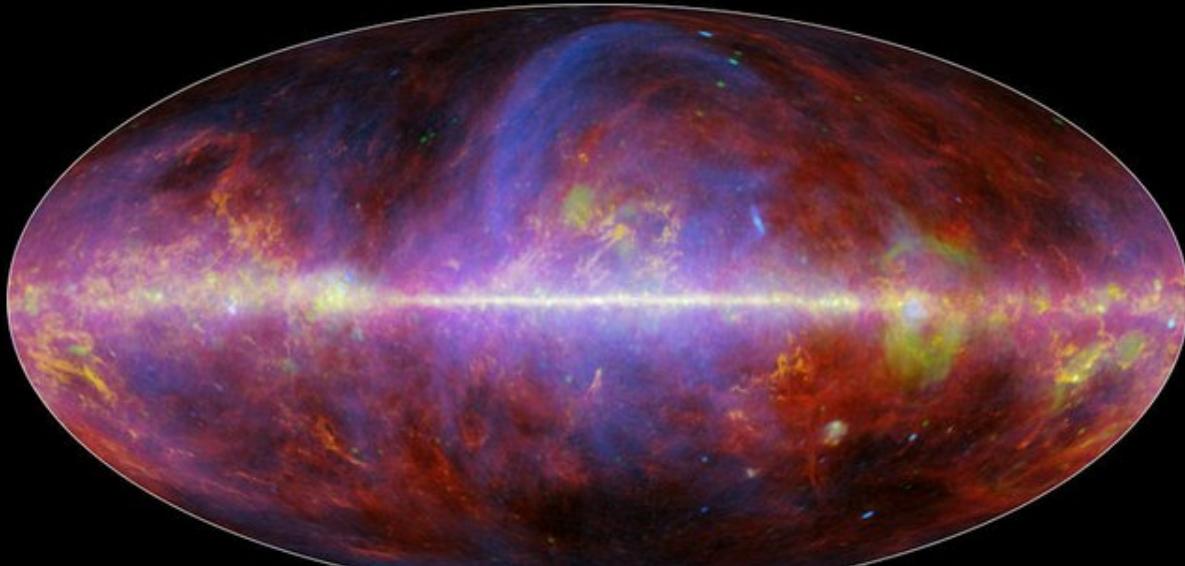
Exploring Galaxy Evolution With The High-Dimensional ISM

G. Zasowski
University of Utah

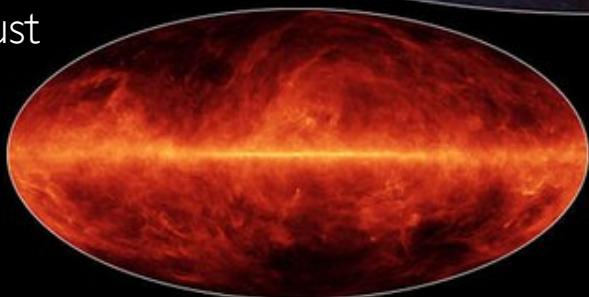


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UNIVERSITY
OF UTAH®

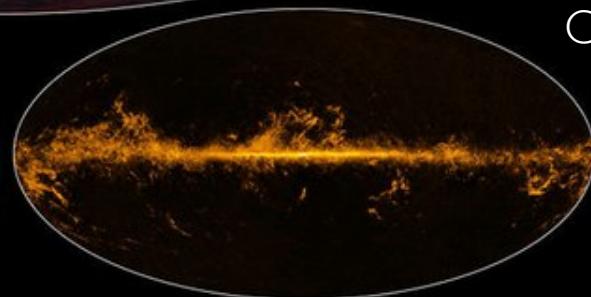
A Complex Interstellar Medium



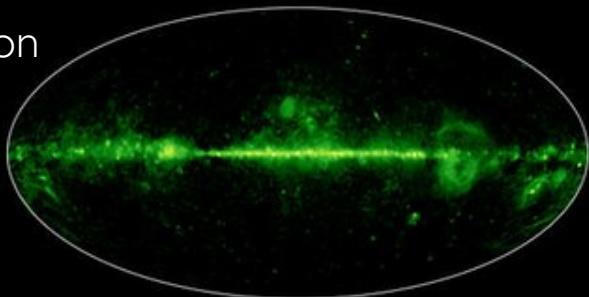
Cold dust



CO

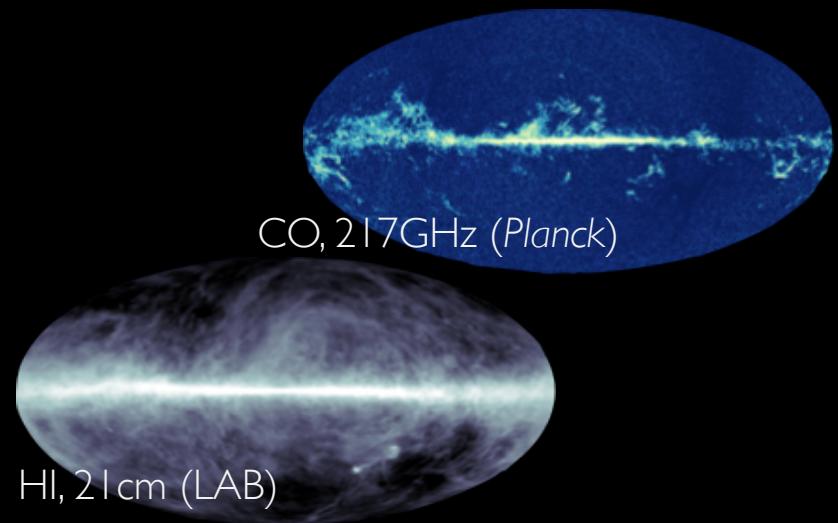


Free-free emission



Synchrotron emission

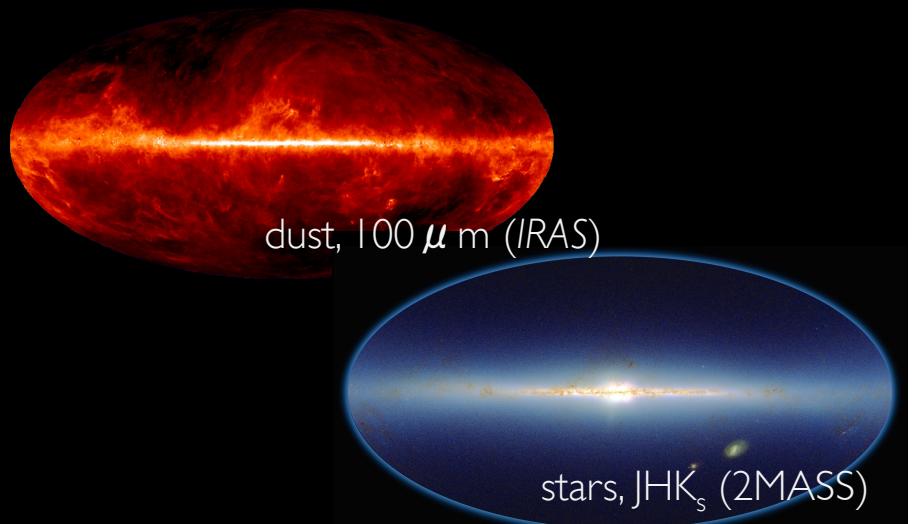




CO, 217GHz (Planck)

HI

21cm (LAB)

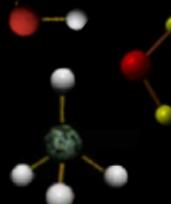


dust, 100 μ m (IRAS)

stars, JHK_s (2MASS)

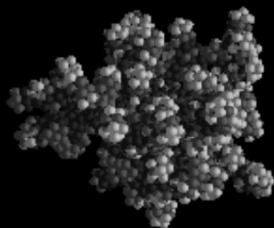
atoms

small molecules



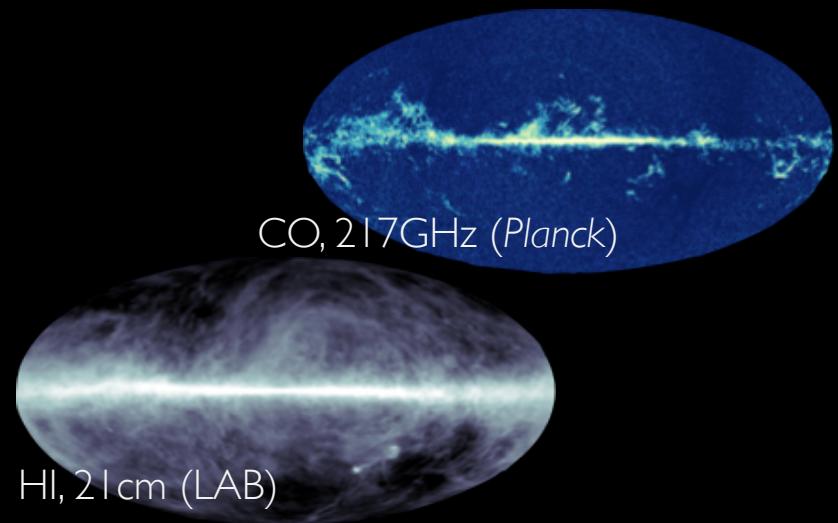
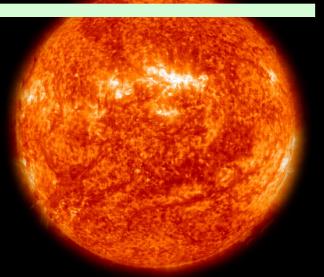
H

dust grains

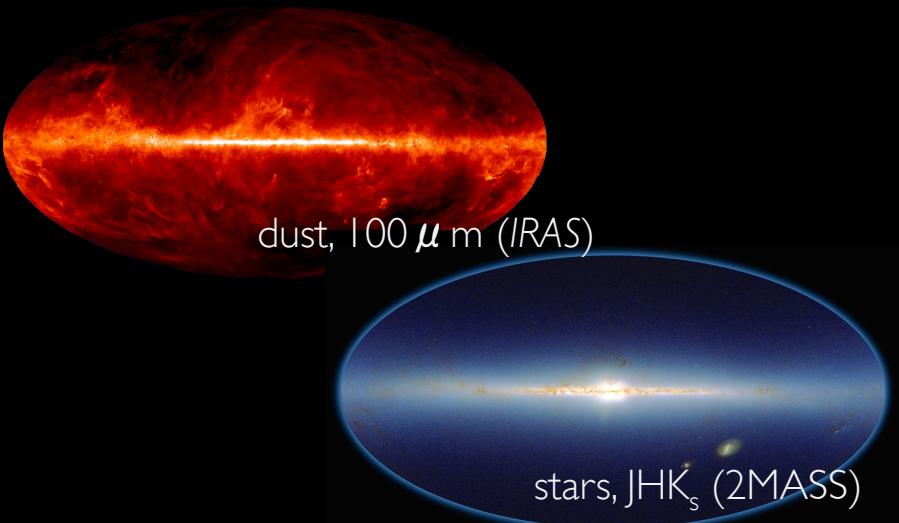


stars

complexity



CO, 217GHz (*Planck*)

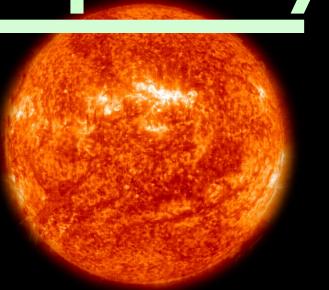
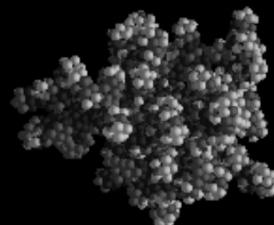
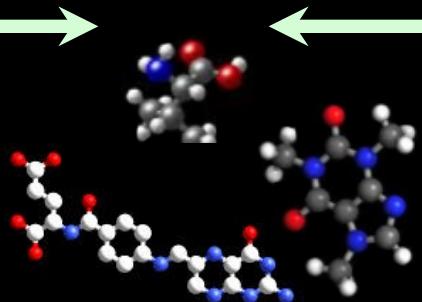
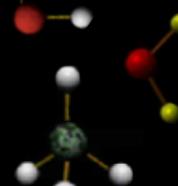


dust, 100 μ m (IRAS)

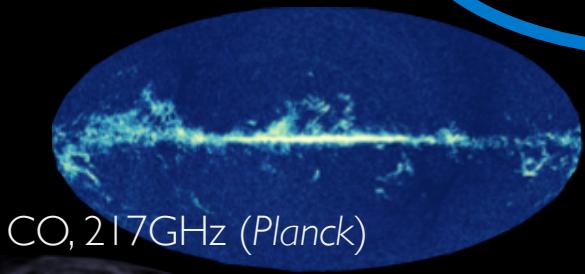
H I, 21 cm (LAB)

atoms small molecules large molecules dust grains stars
complexity

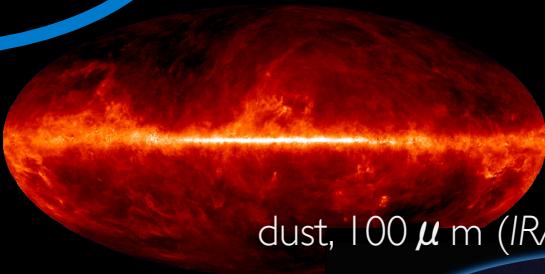
H



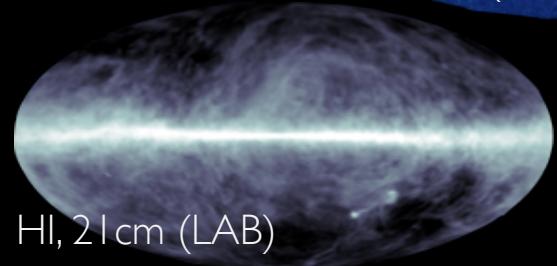
?



CO, 217GHz (Planck)



dust, 100 μ m (IRAS)

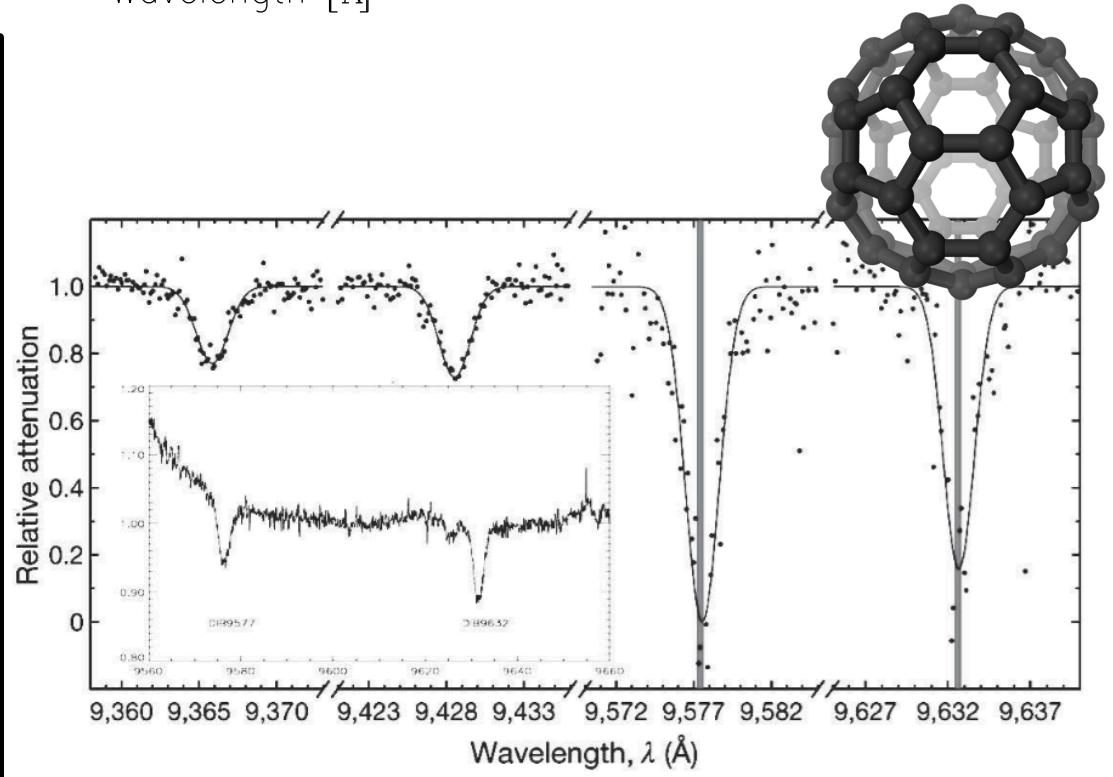
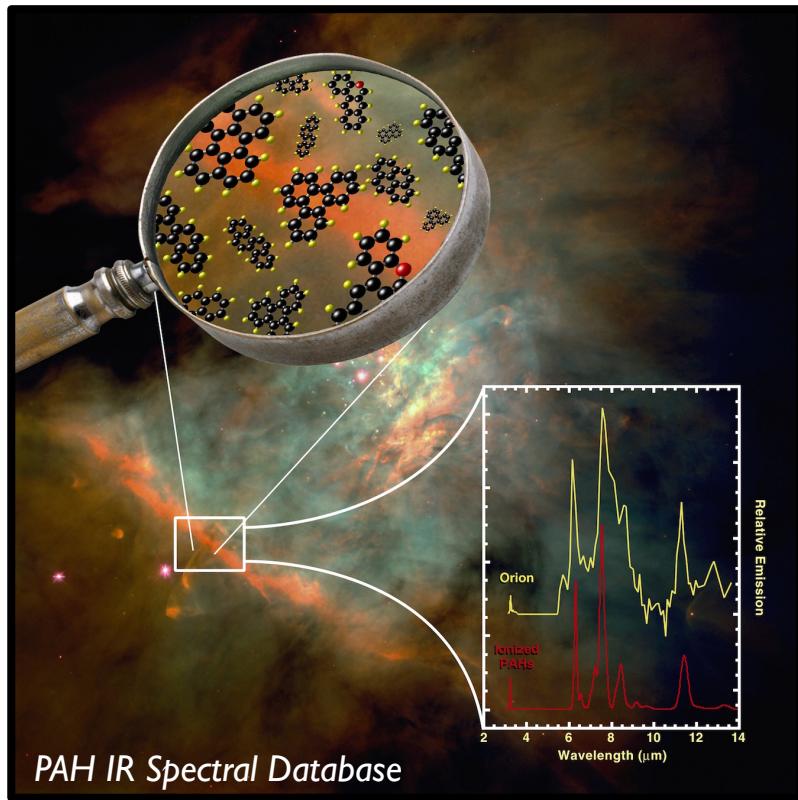
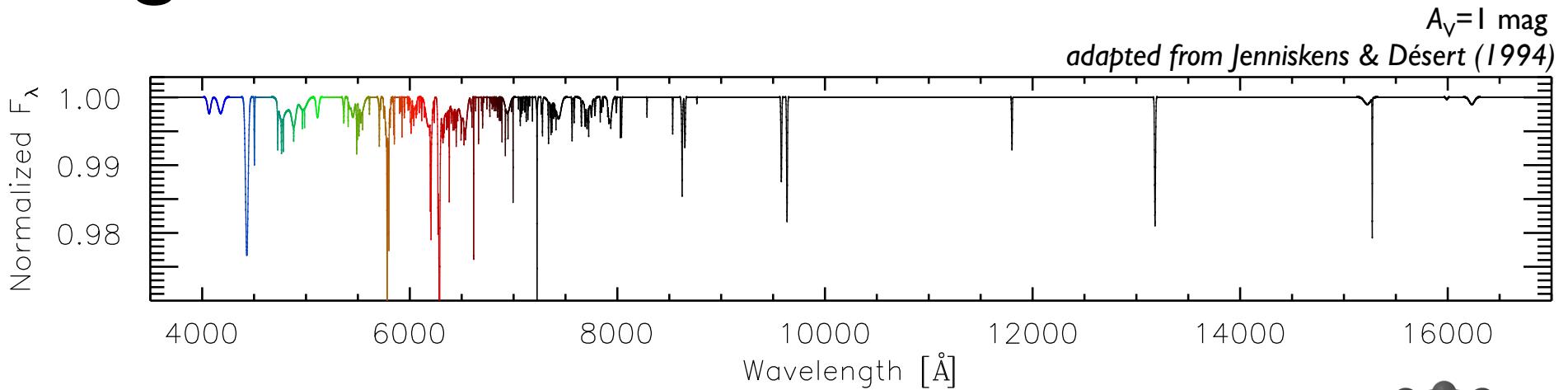


HI, 21cm (LAB)



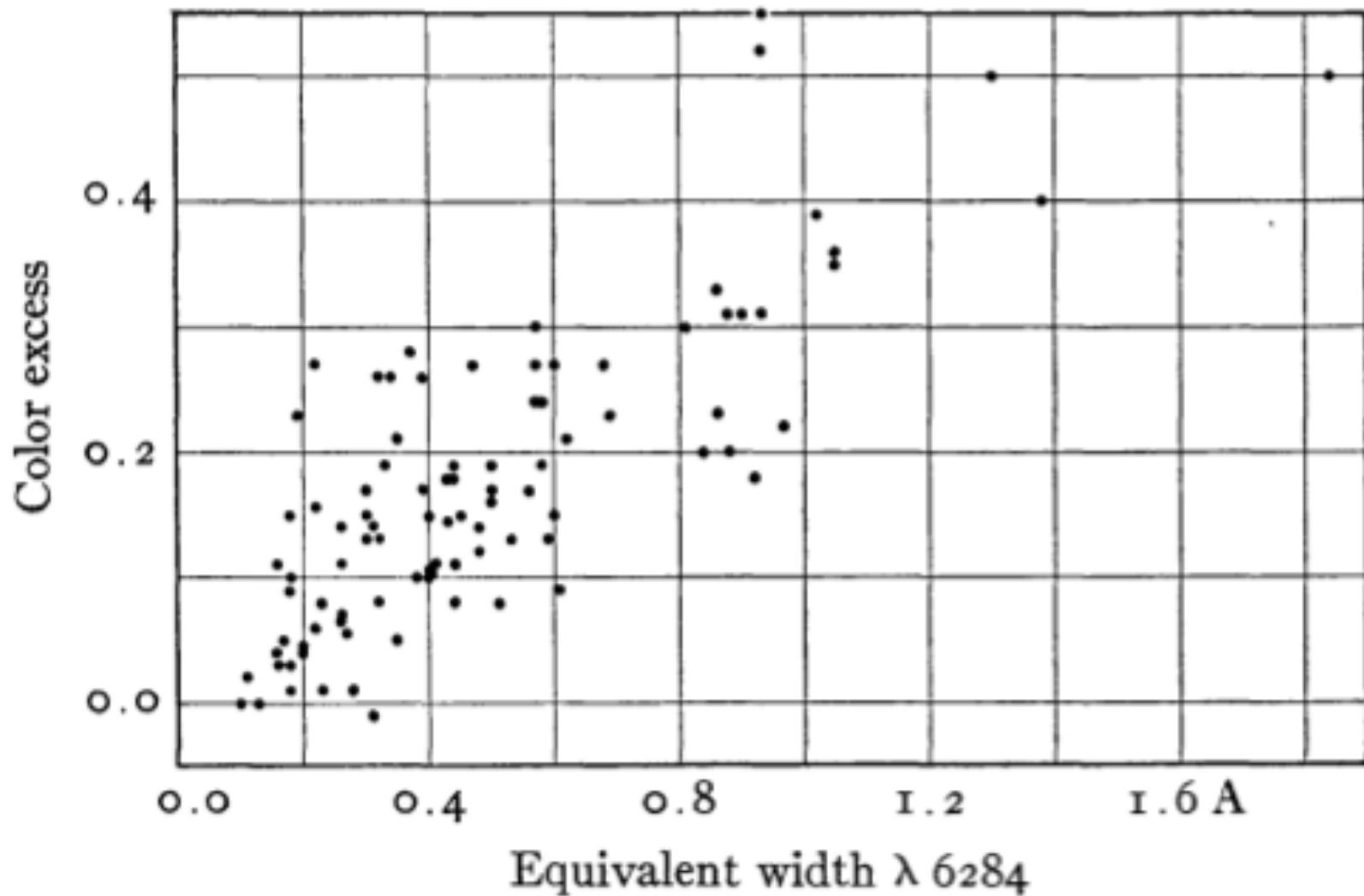
stars, JHK_s (2MASS)

Large Molecules \longleftrightarrow DIBs?

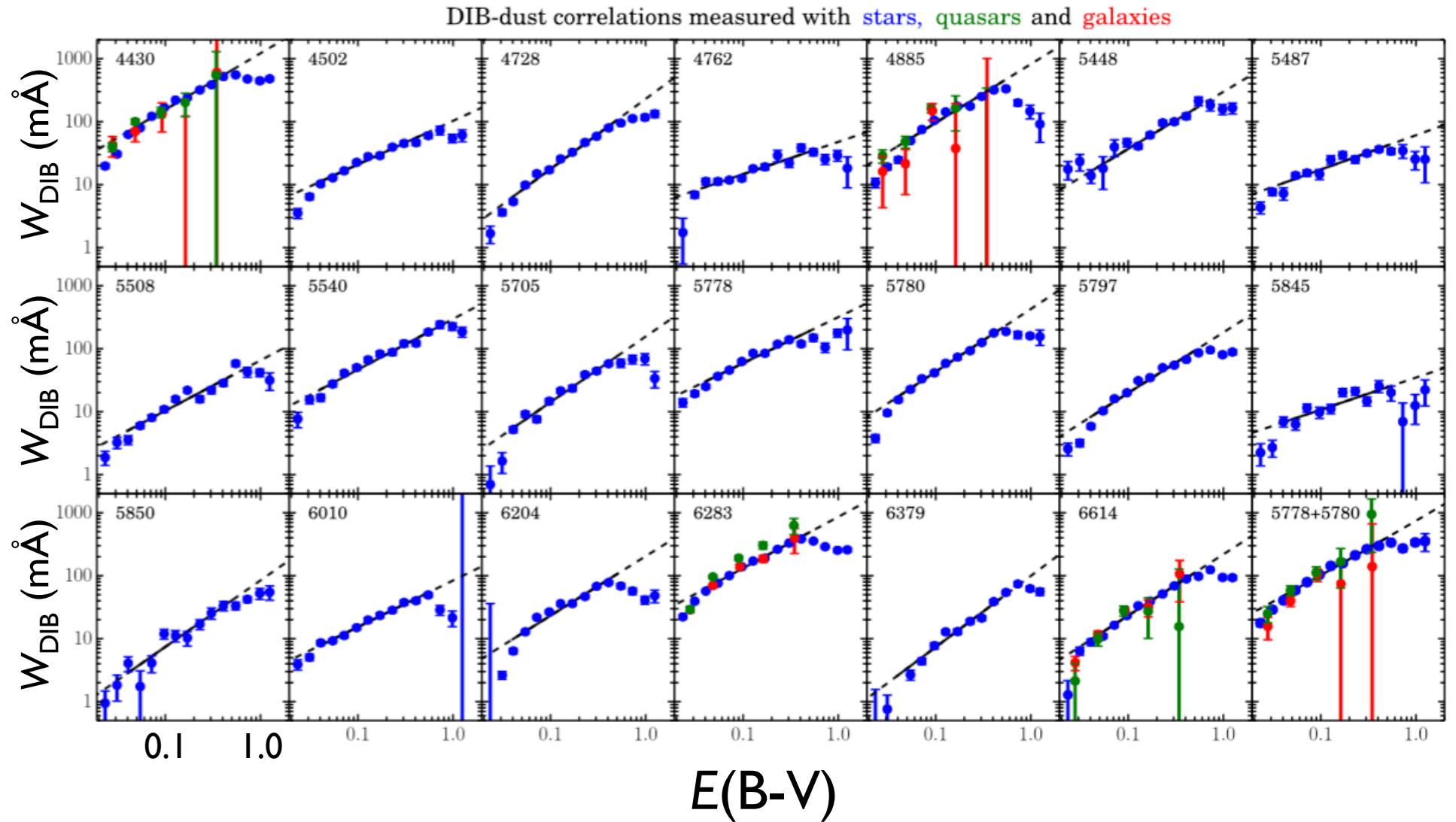


Campbell et al. (2015)

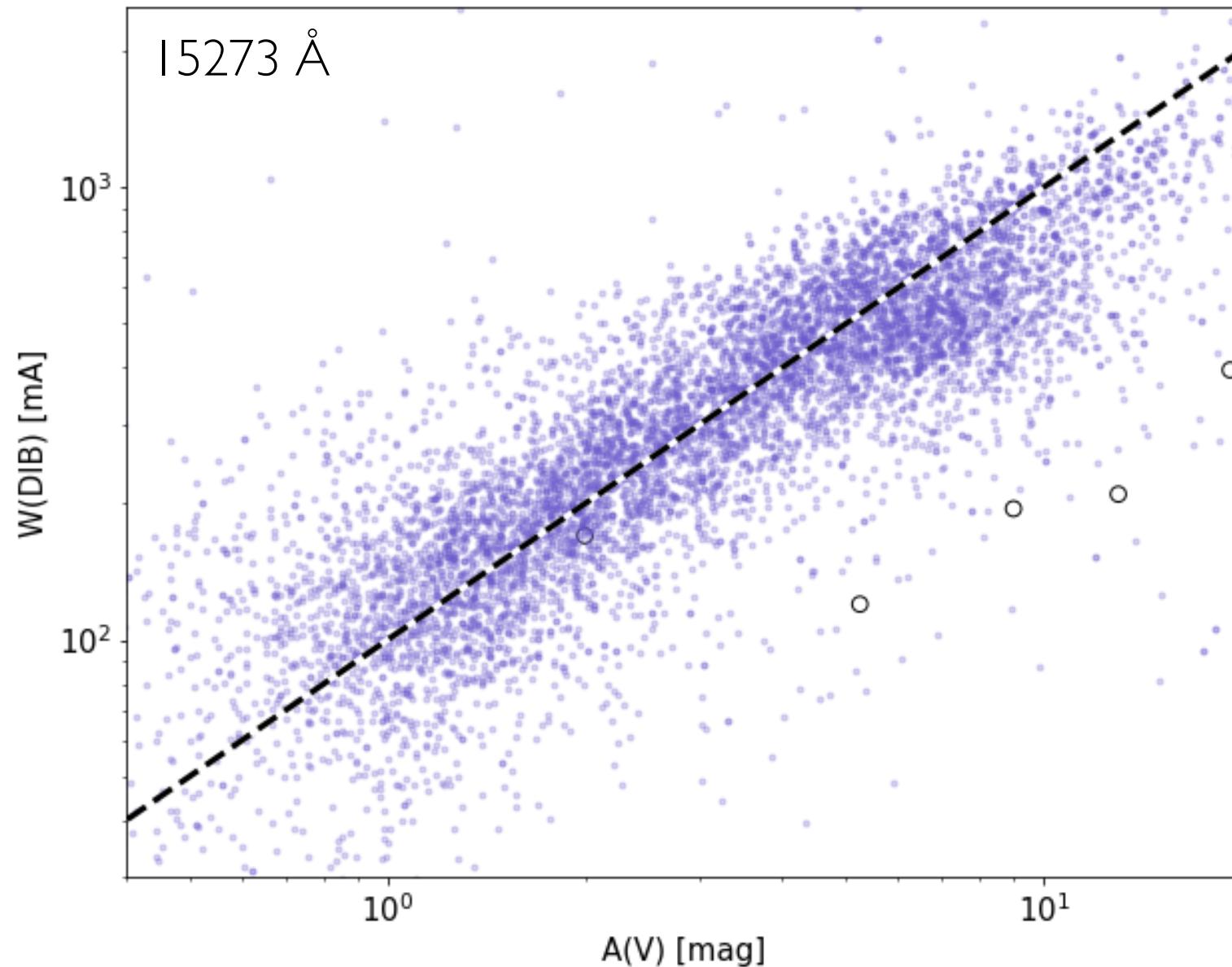
DIBs + Dust



DIBs + Dust

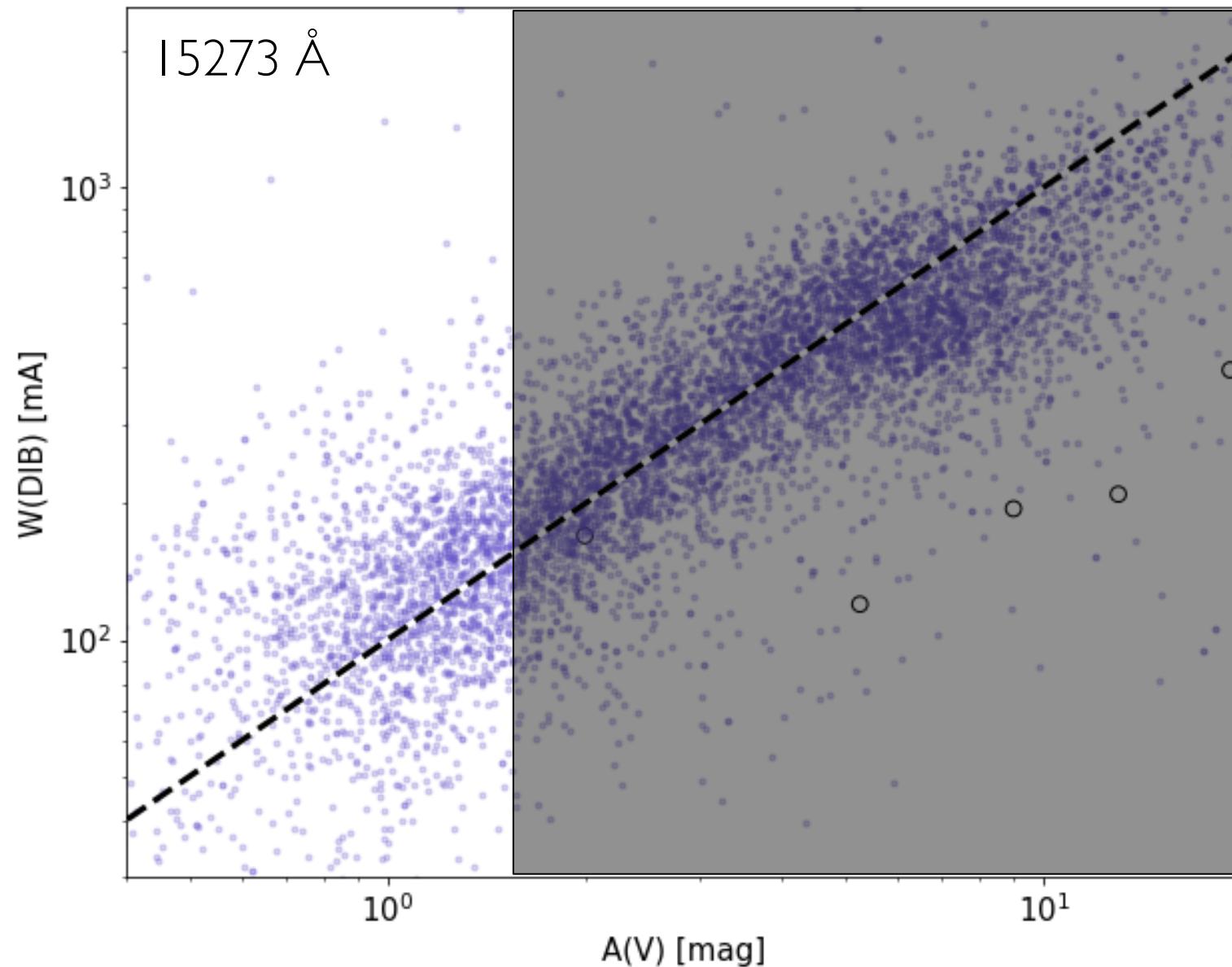


DIBs + Dust



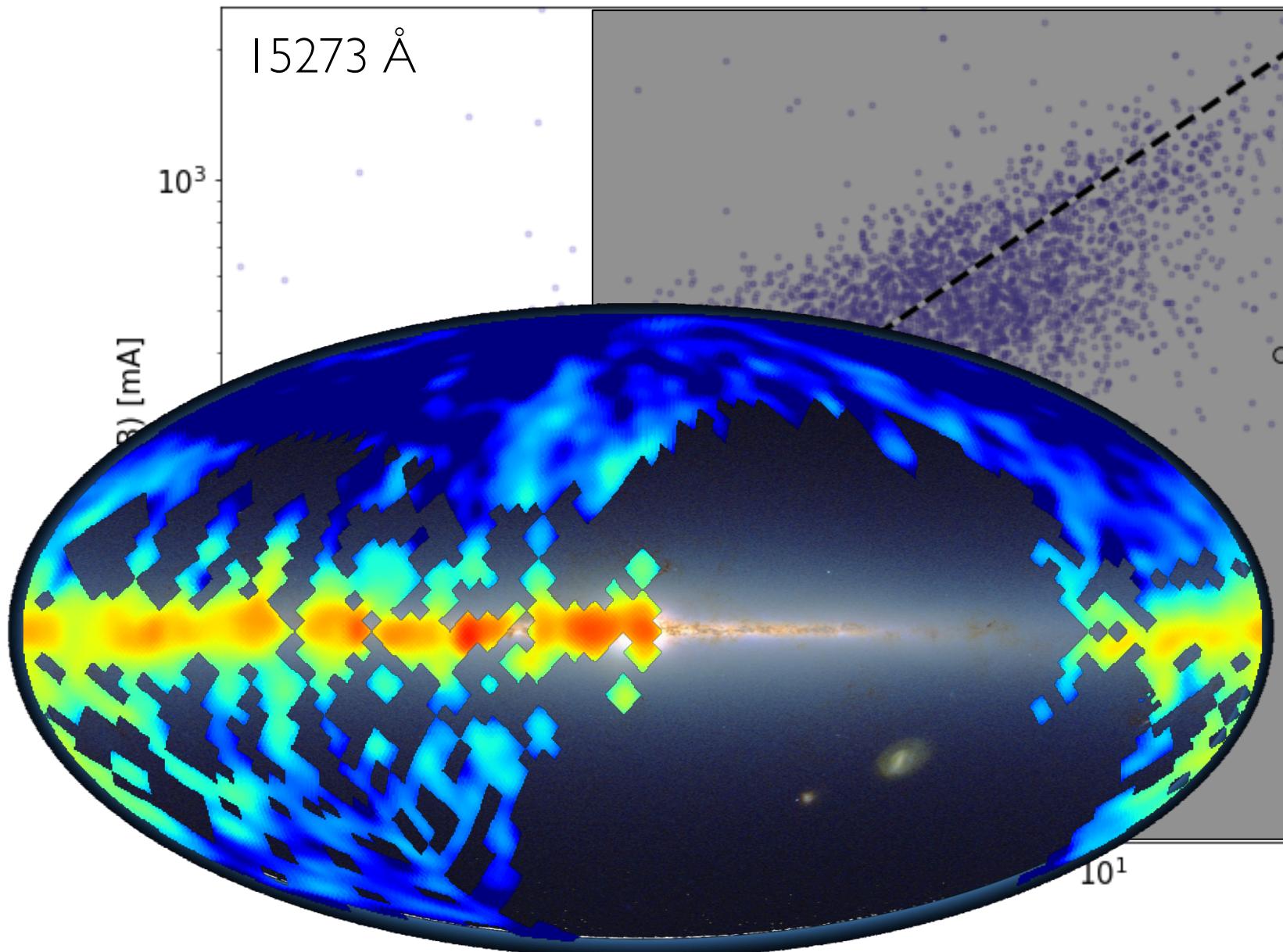
GZ et al. (2015)

DIBs + Dust



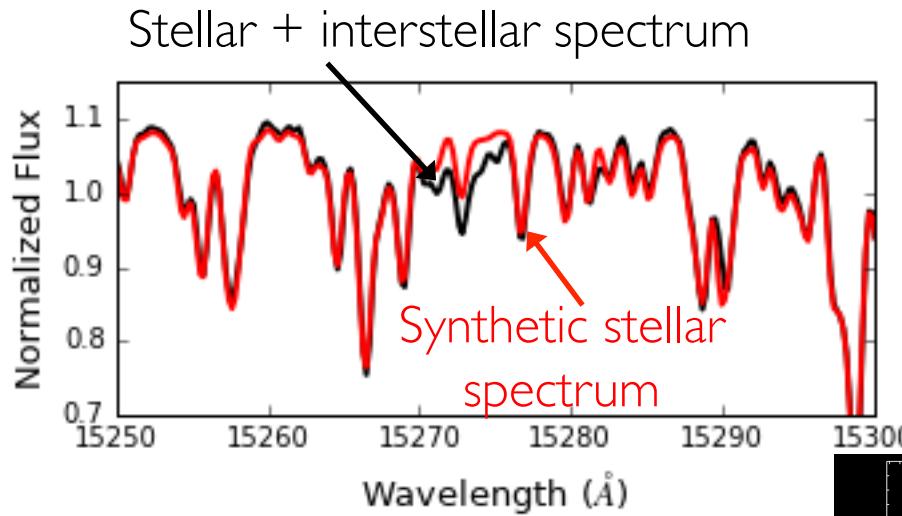
GZ et al. (2015)

DIBs + Dust



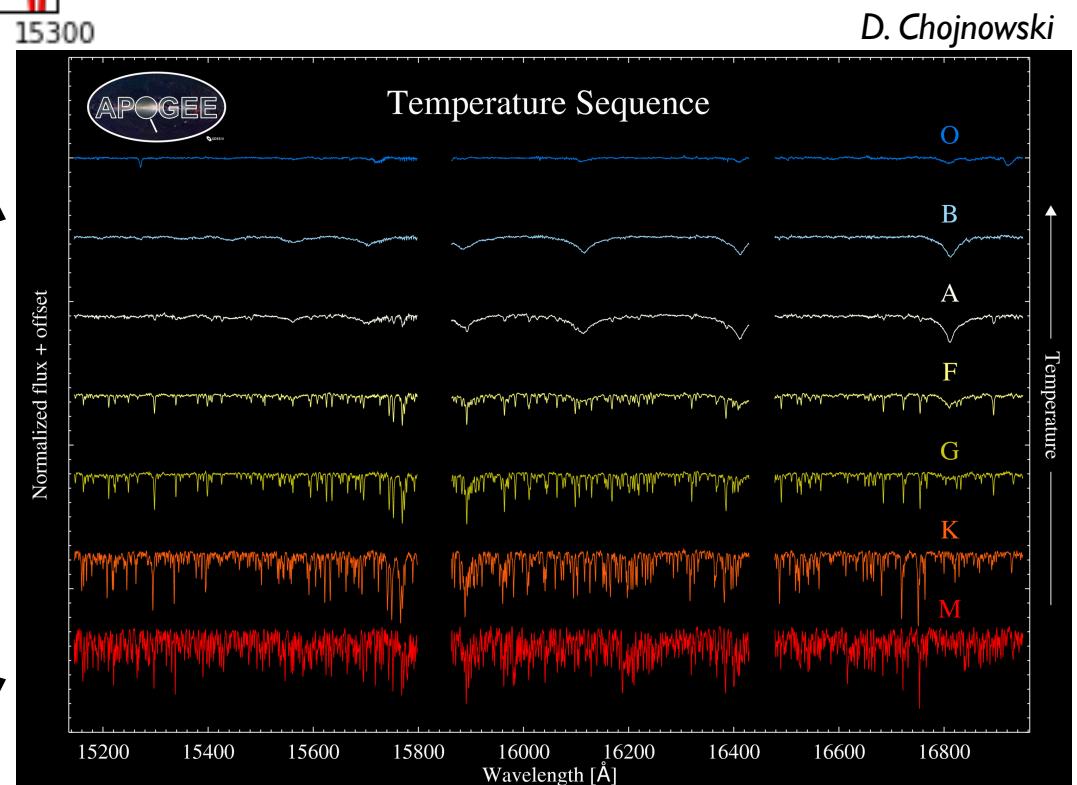
GZ et al. (2015)

Extracting the 1.53 μ m DIB

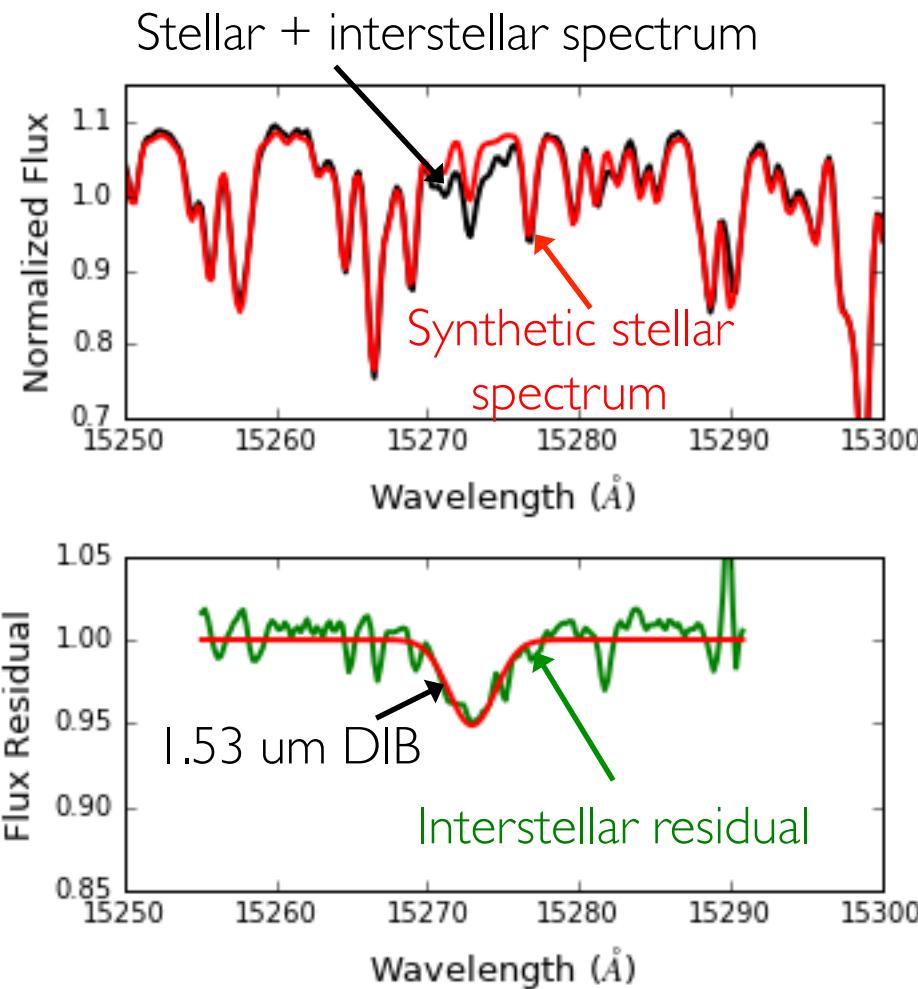


Relatively easy to use these

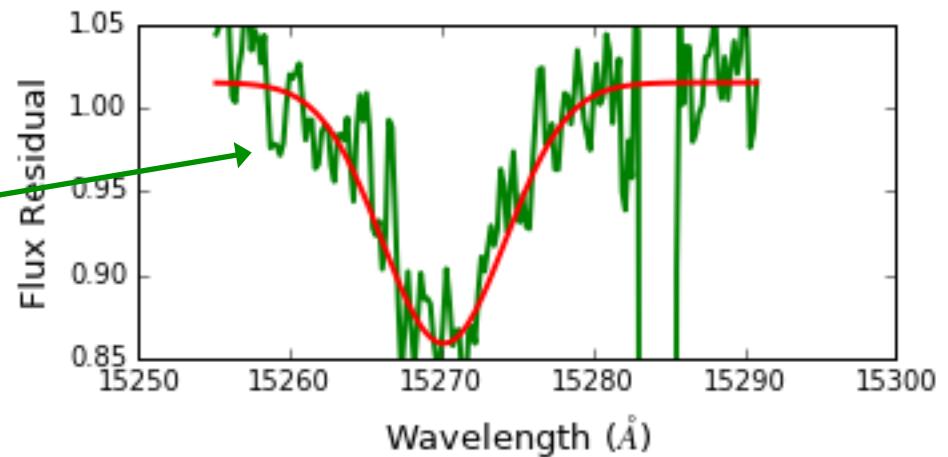
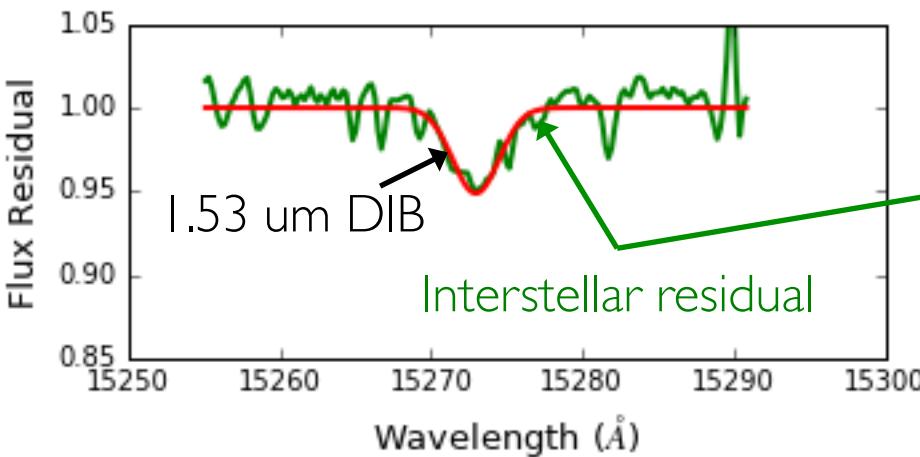
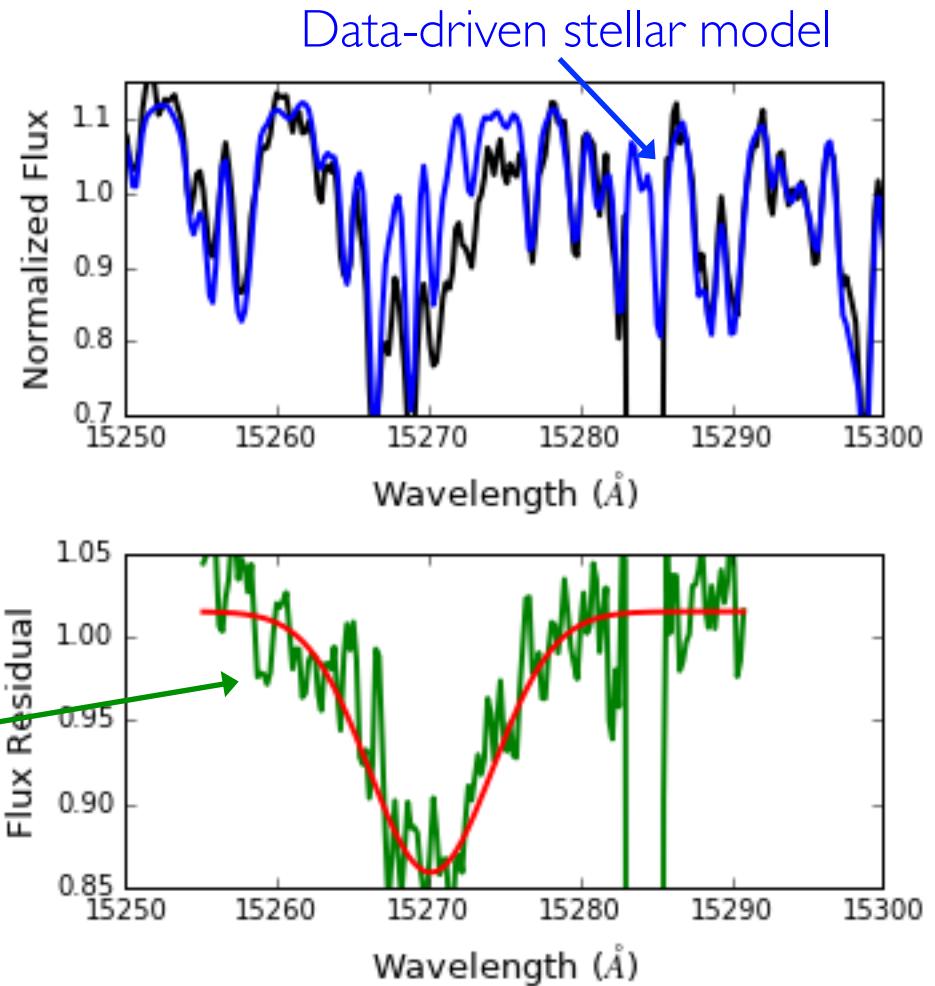
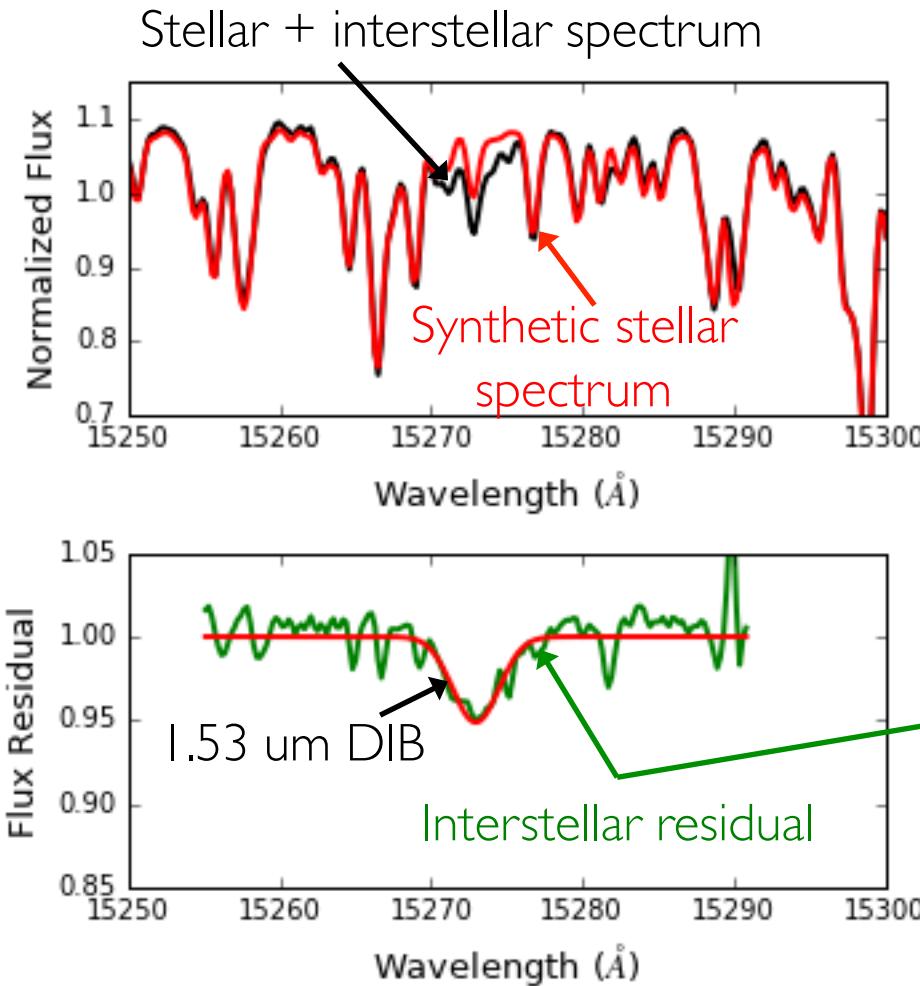
To probe large distances in the disk, need to use these



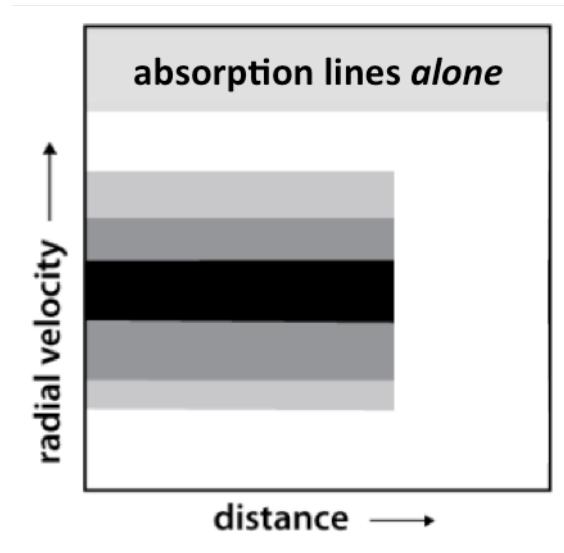
Extracting the 1.53 μ m DIB



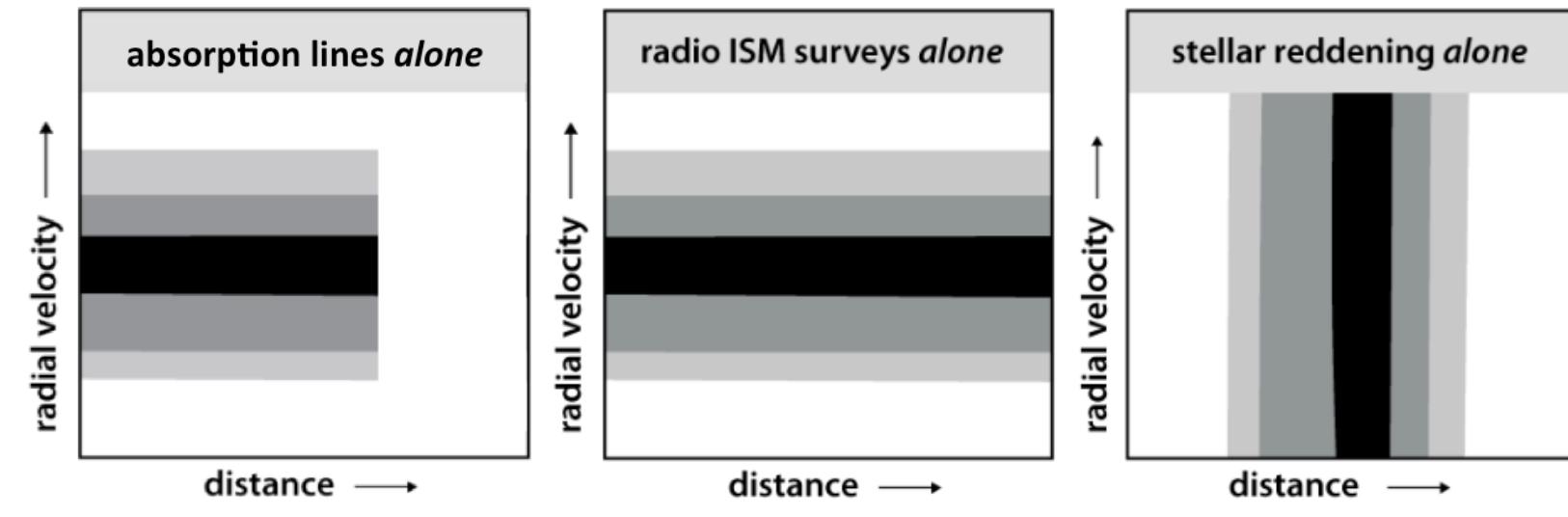
Extracting the 1.53 μ m DIB



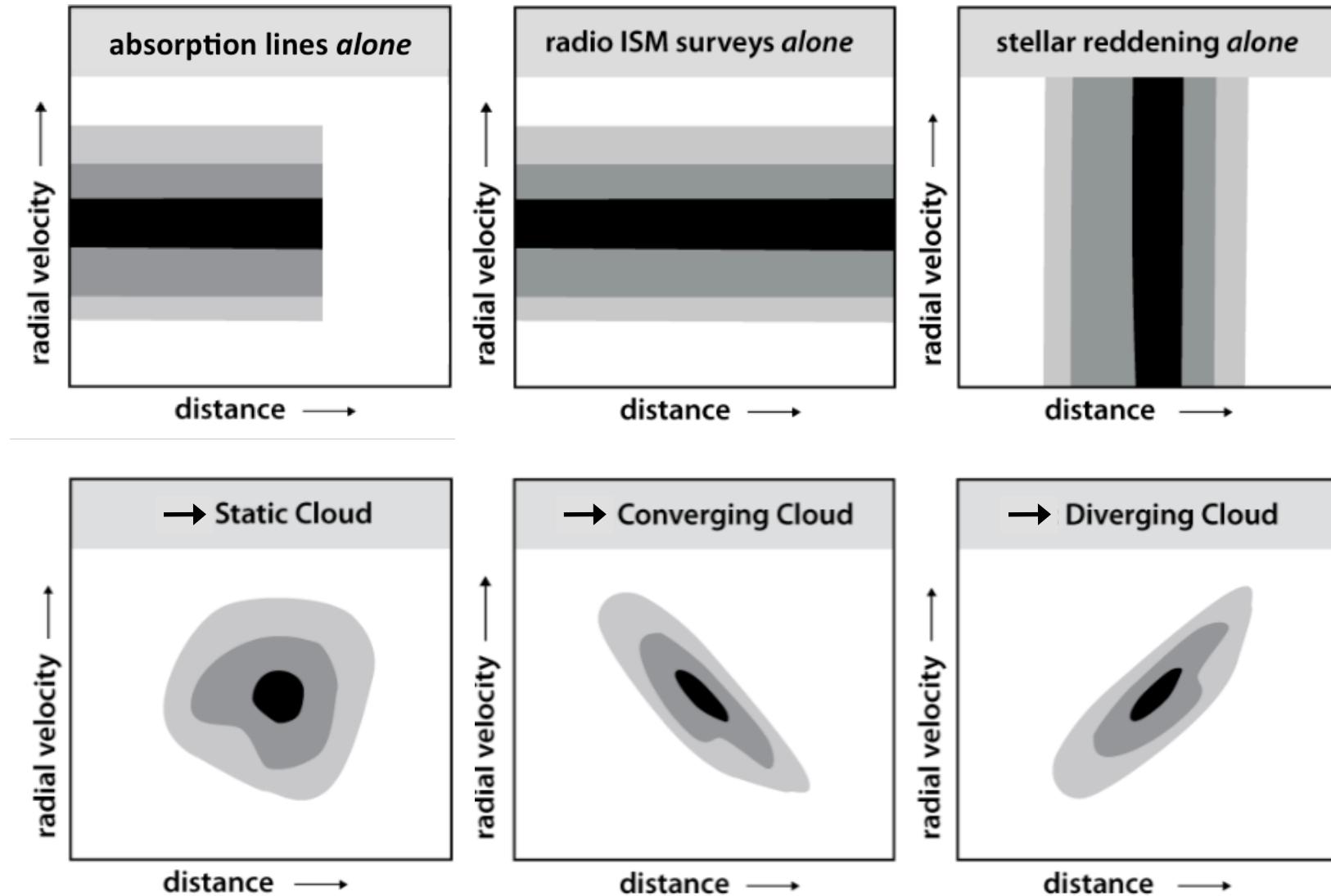
Galaxy Evolution via the 4-D ISM



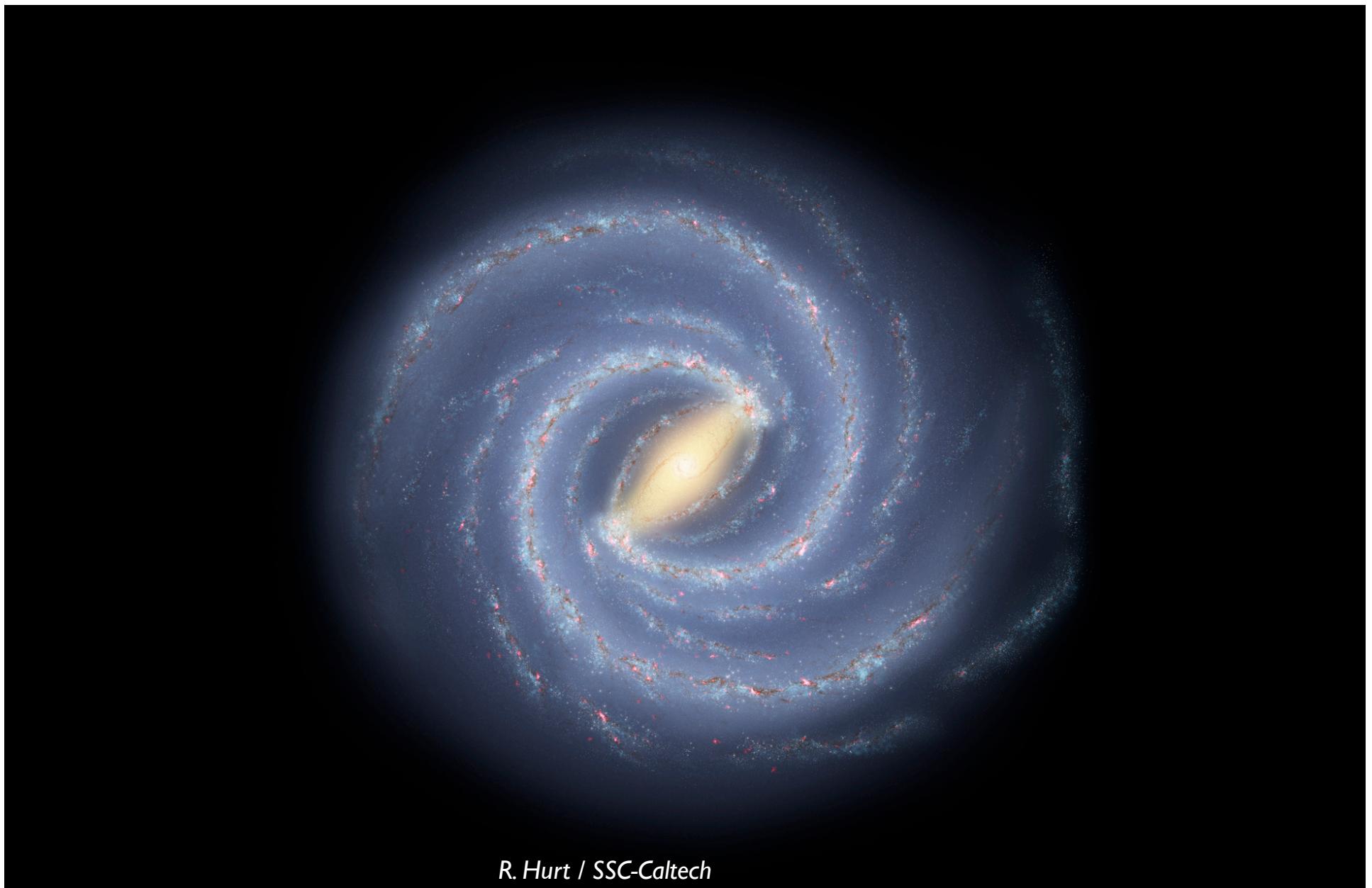
Galaxy Evolution via the 4-D ISM



Galaxy Evolution via the 4-D ISM



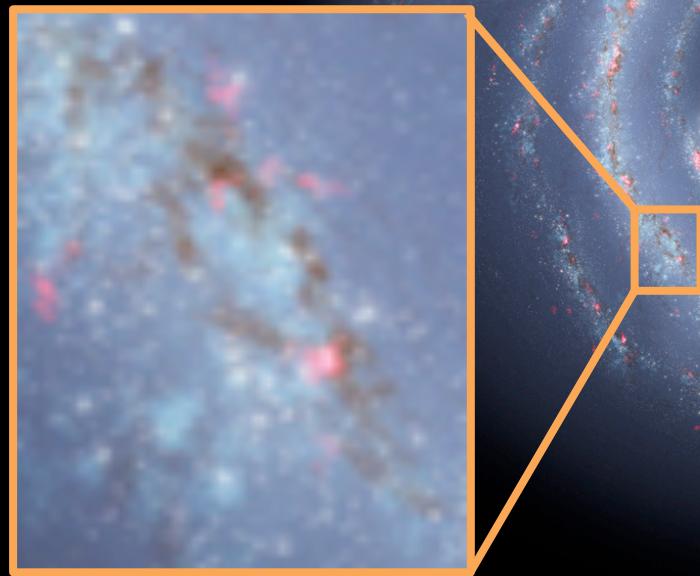
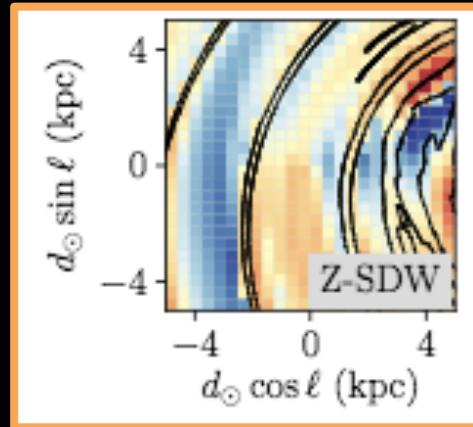
Galaxy Evolution via the 4-D ISM



R. Hurt / SSC-Caltech

Galaxy Evolution via the 4-D ISM

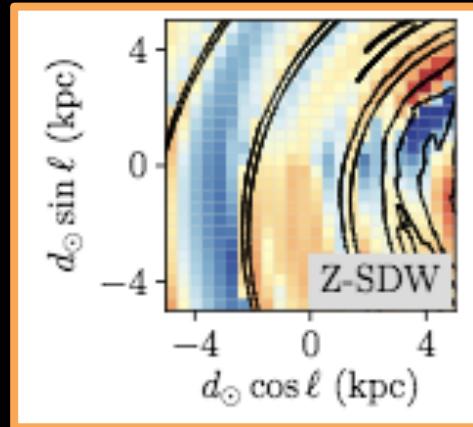
What drives spiral arms?



R. Hurt / SSC-Caltech

Galaxy Evolution via the 4-D ISM

What drives spiral arms?

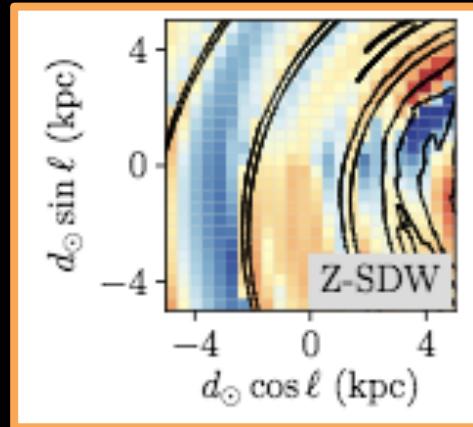


How do GMCs evolve?



Galaxy Evolution via the 4-D ISM

What drives spiral arms?



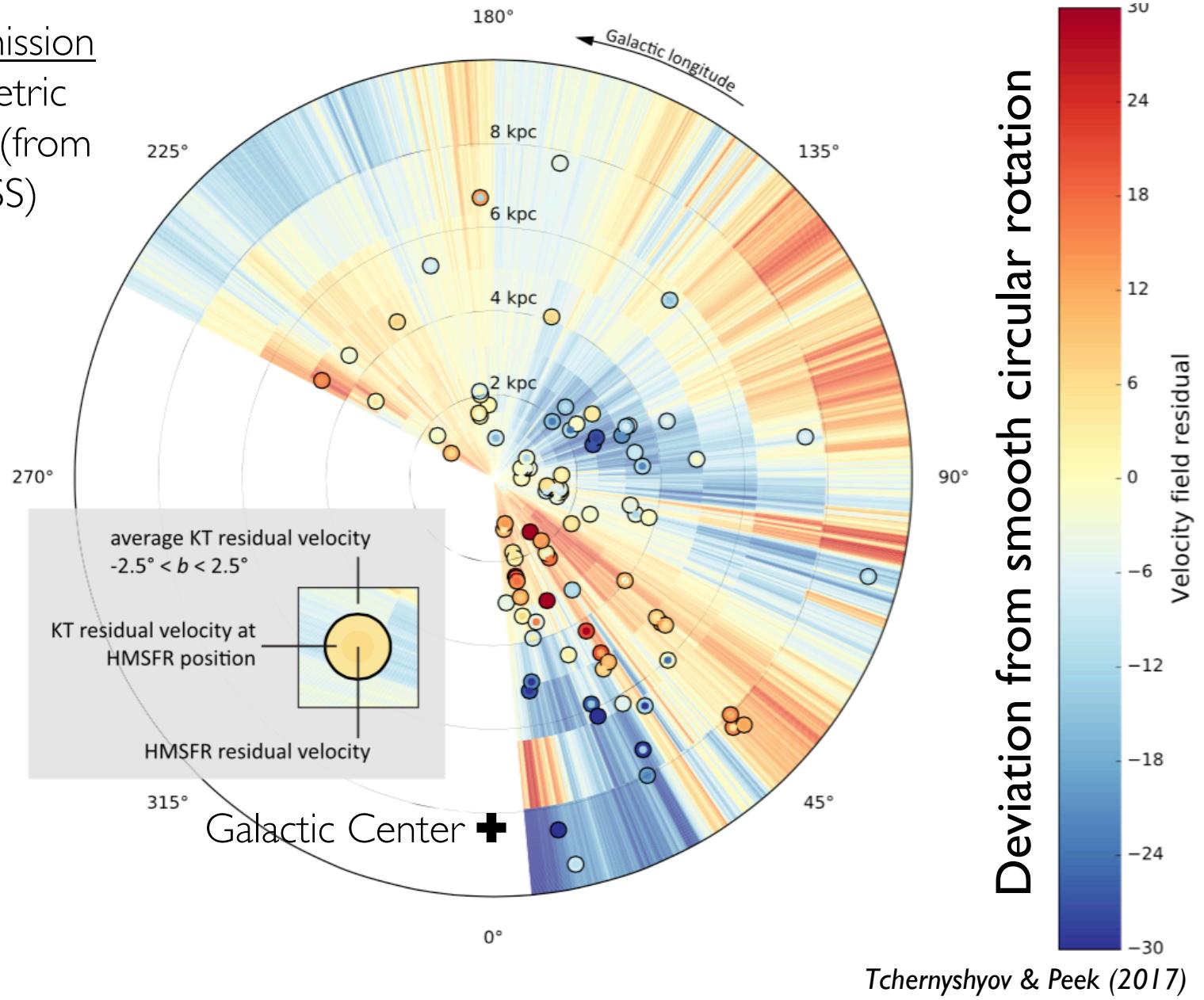
How do GMCs evolve?



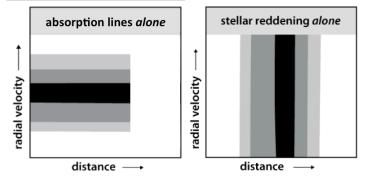
How does the ISM move around the Galaxy?

Case Study: Kinetic Tomography

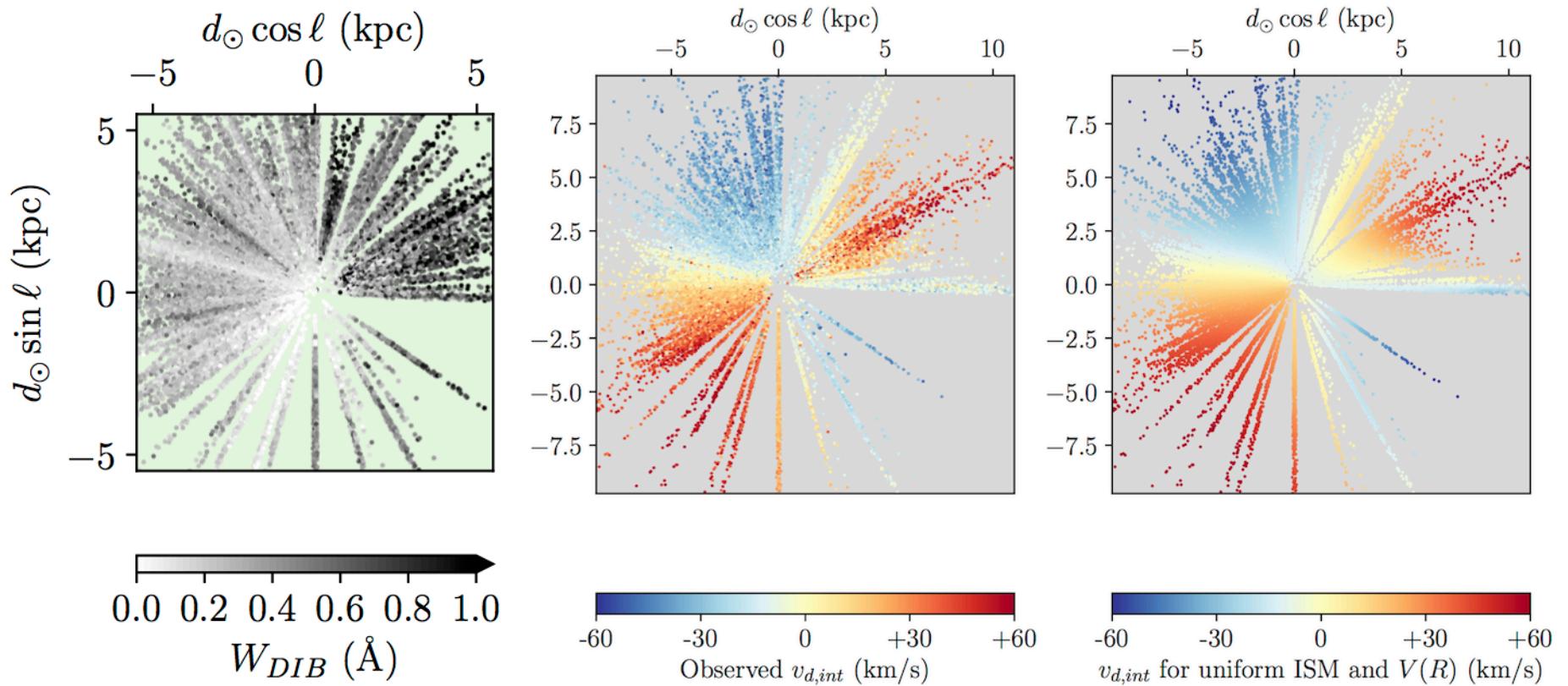
Using HI/CO emission
and 3D photometric
extinction maps (from
Panstarrs+2MASS)



Case Study: Kinetic Tomography



- Advantages to using absorption lines

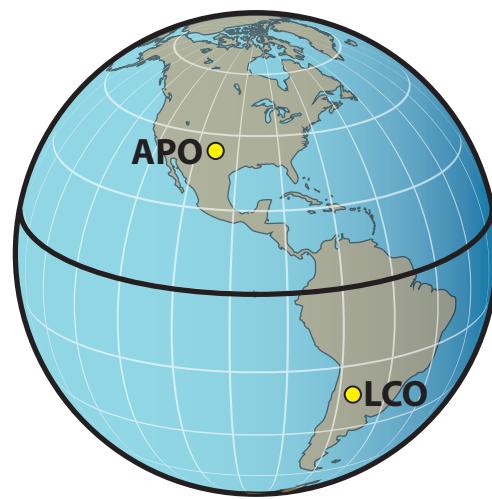


Quite new -- much to explore!

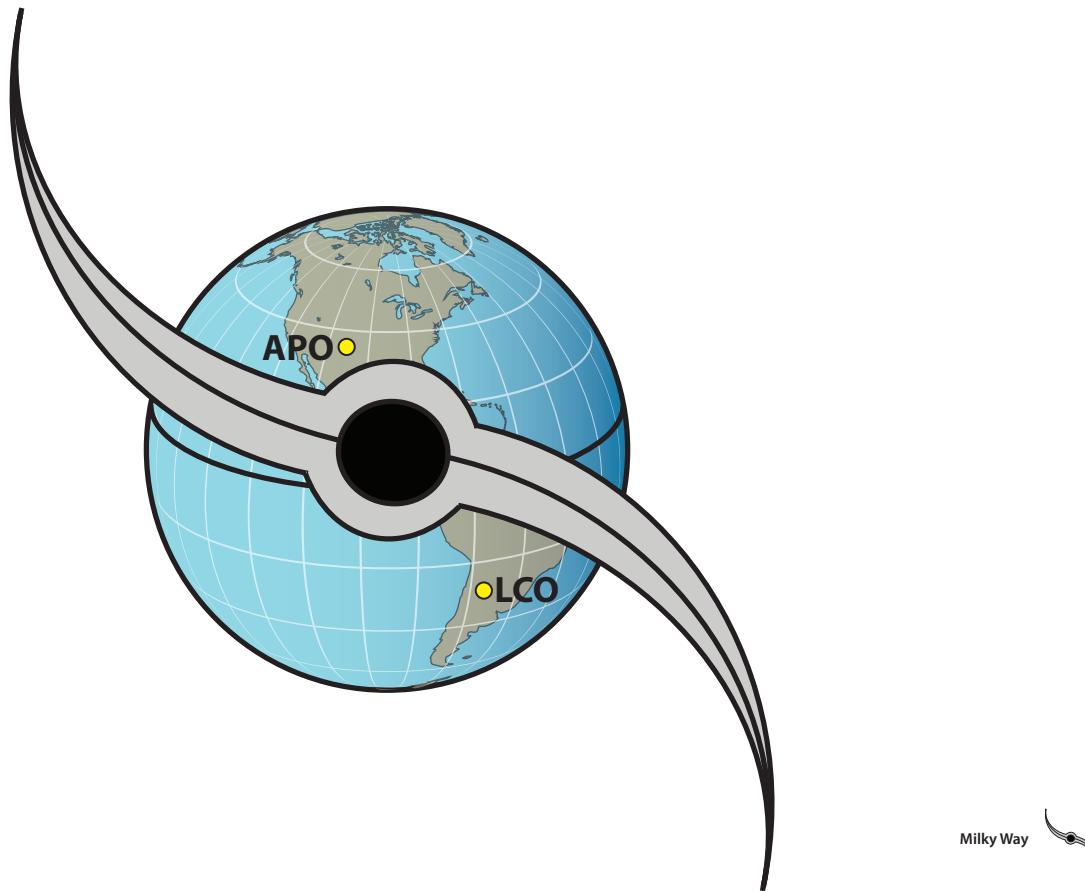
The Future is...Dark! (in a good way)

- Improved 3-D dust maps with *Gaia*
- 1.53 μ m DIBs along \sim 5M sightlines with SDSS-V

SDSS-V



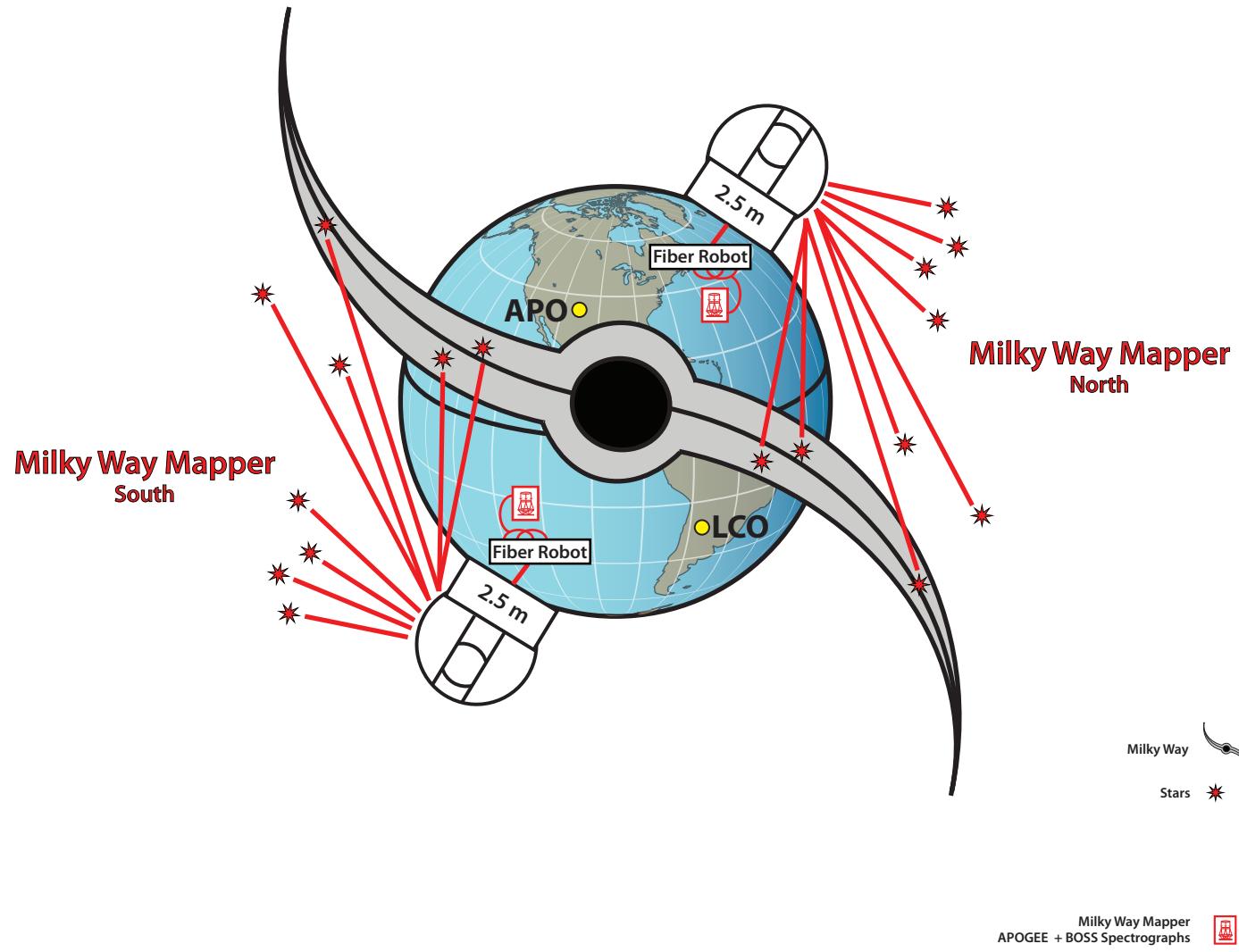
SDSS-V



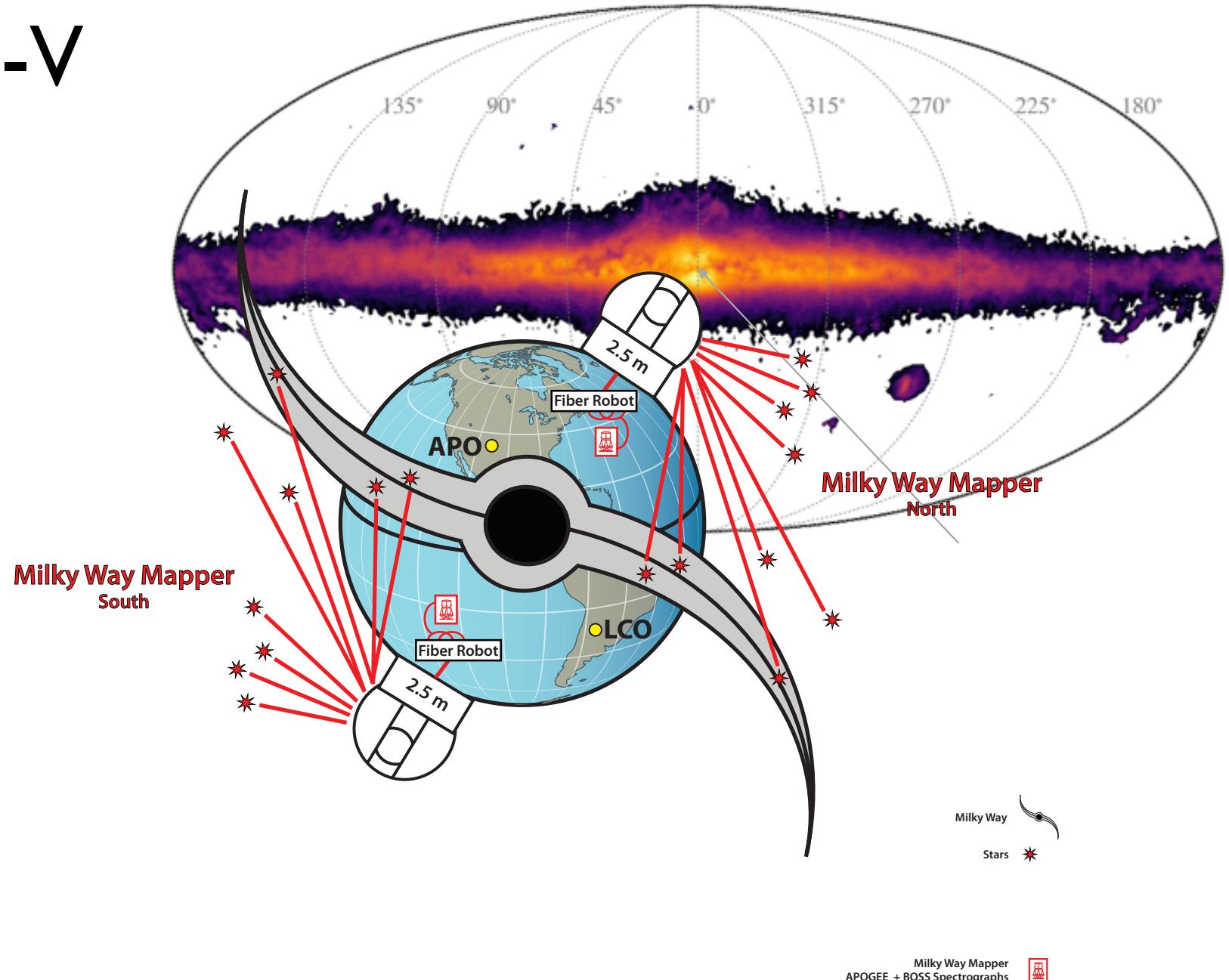
<http://www.sdss.org/future/>

M. Seibert

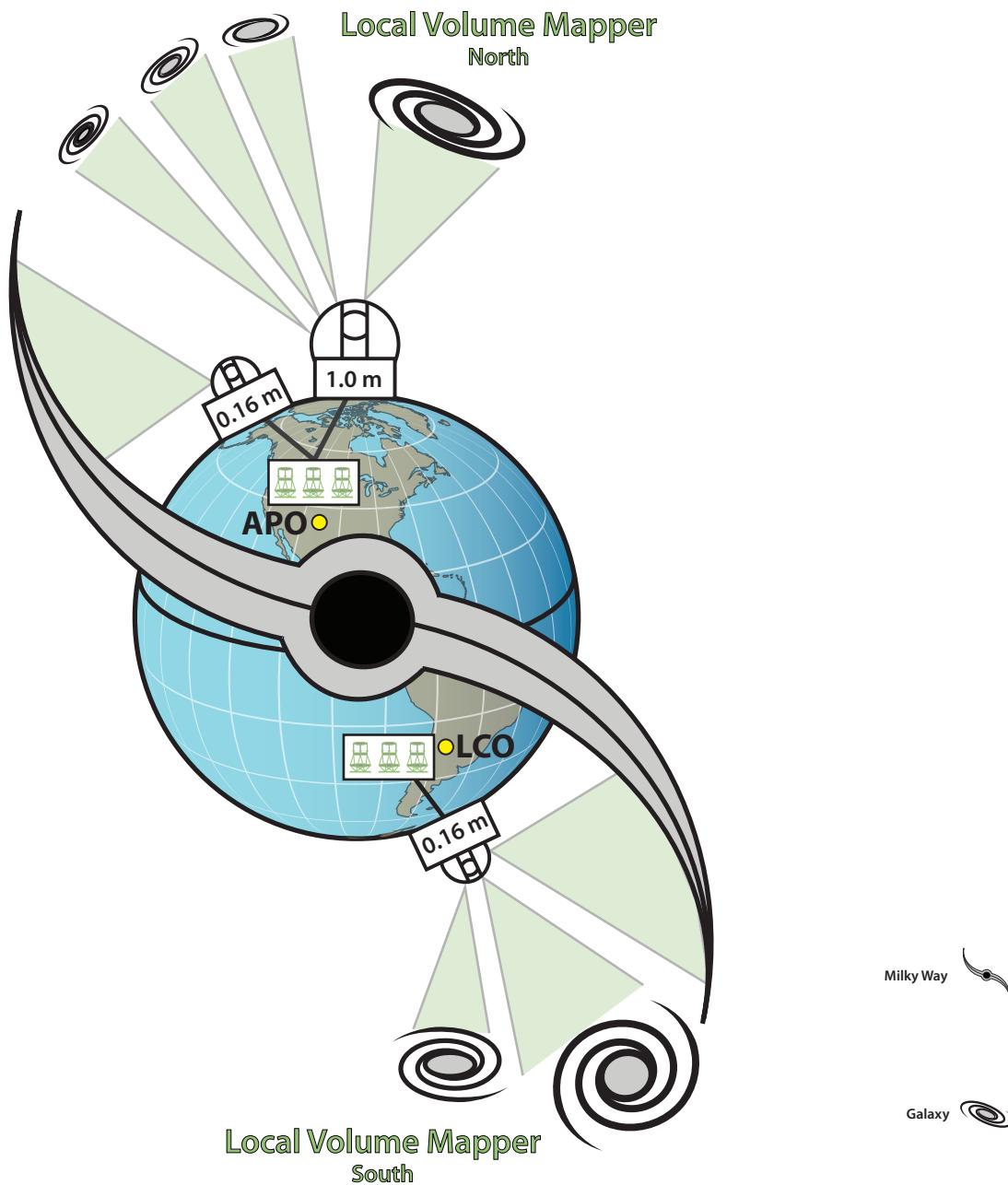
SDSS-V



SDSS-V



SDSS-V



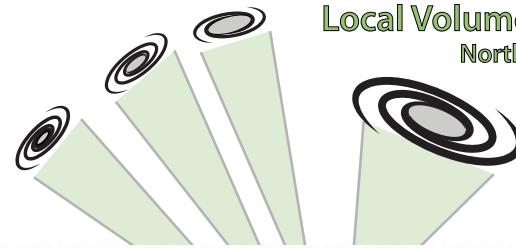
<http://www.sdss.org/future/>

Local Volume Mapper
Integral Field Spectrograph



M. Seibert

SDSS-V



Local Volume Mapper
North

100 pc/pixel

50 pc/pixel

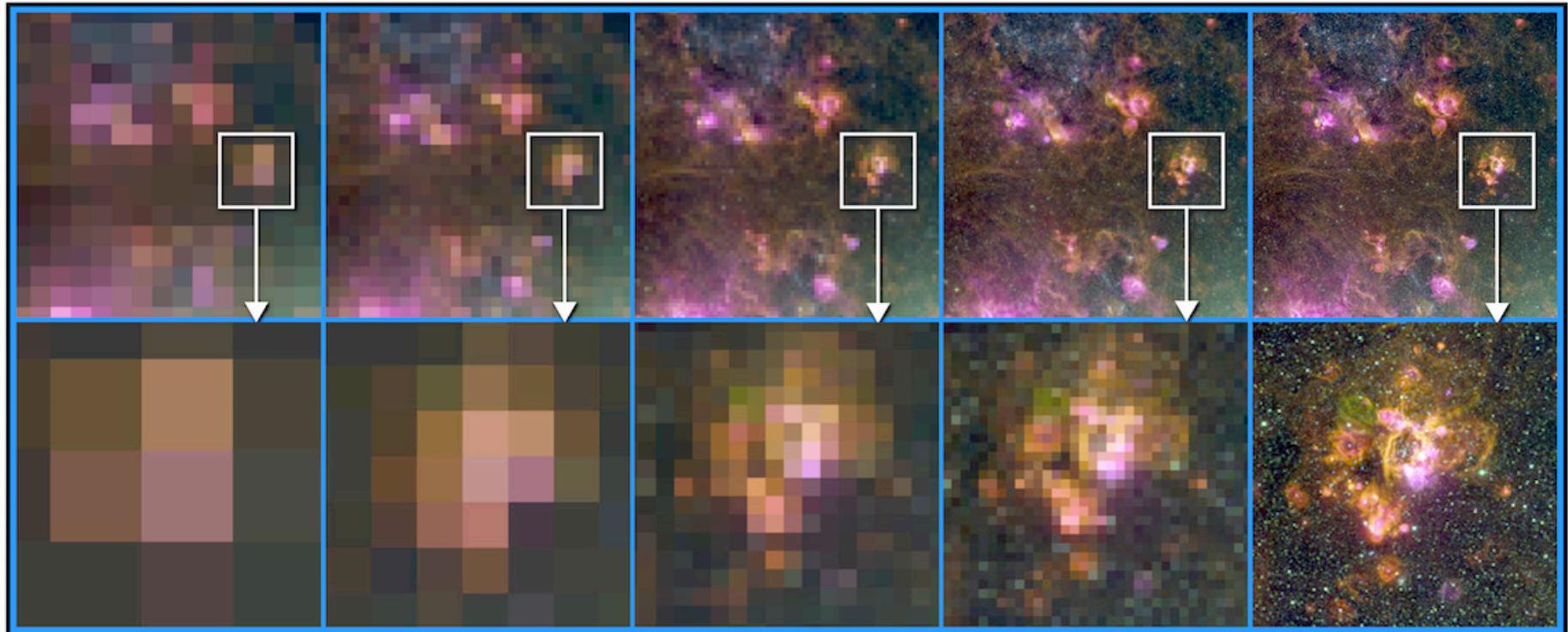
25 pc/pixel

10 pc/pixel

1.6 pc/pixel

1.4² sqr. kpc

320² sqr. pc



Local Volume Mapper
South

Galaxy

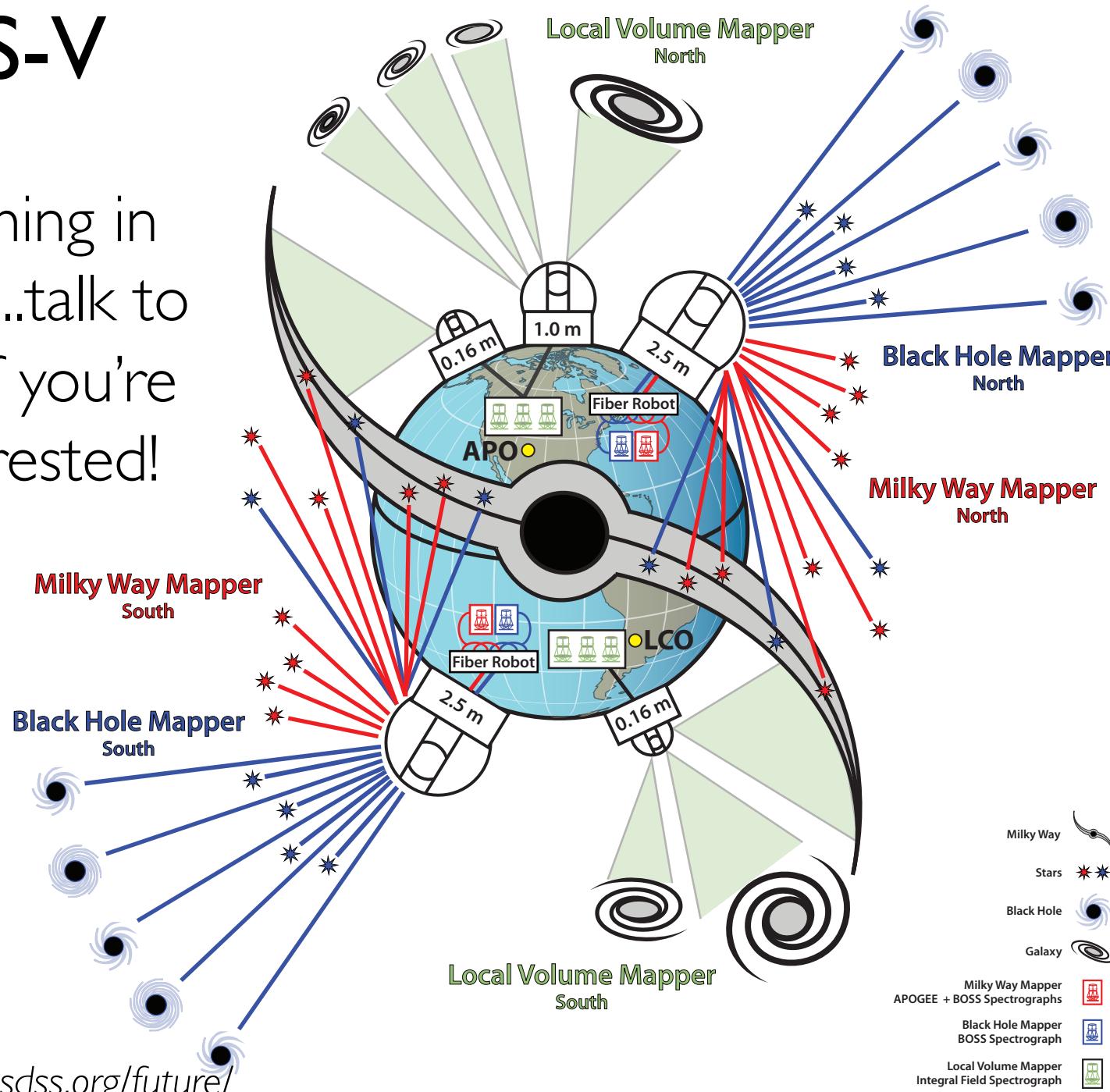
Local Volume Mapper
Integral Field Spectrograph



M. Seibert

SDSS-V

Coming in
2020...talk to
me if you're
interested!



<http://www.sdss.org/future/>

M. Seibert

Thank You



R. Dienel