

Constraining dust mineralogy from mid-IR spectra

Peter Scicluna

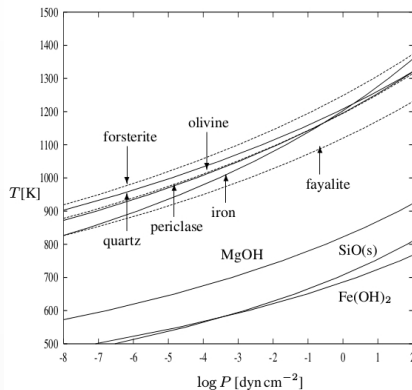
ASIAA, Taiwan

CPHDust 2018

Ciska Kemper, Sundar Srinivasan, Lapo Fanciullo, Alfonso Trejo,
Thavisha Dharmawardena, Jonty Marshall (ASIAA),
Sacha Hony (Heidelberg)

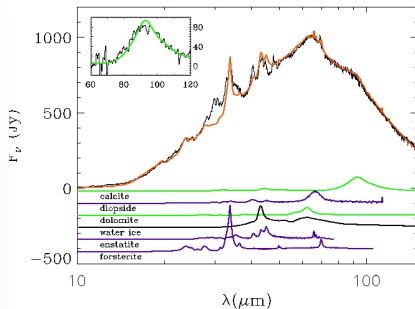
Why do we care?

- Traces physical conditions in dust-forming region
- History of dust processing
- Mineralogy determines optical properties, important for wind driving



Identifying minerals

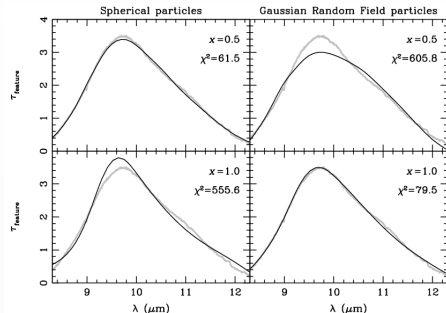
- Observe features and match with lab data.



Kemper et al, 2002b

Identifying minerals

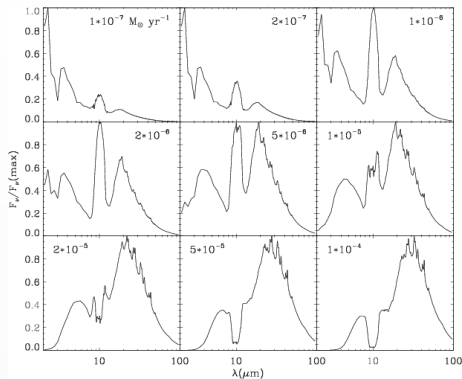
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- Difficult:
 - Grain size and shape



Min et al, 2005

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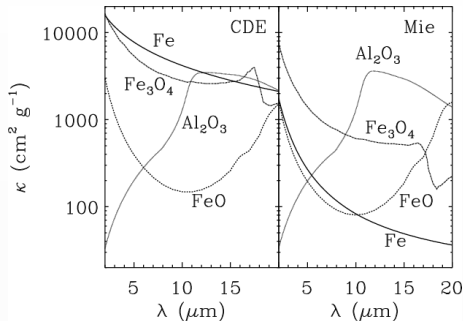
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- Difficult:
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 - temperature & radiative-transfer effects! \Rightarrow Must use models



Kemper et al, 2001

Identifying minerals

- Observe features and match with lab data.
- Difficult:
 - Grain size and shape
 - temperature & radiative-transfer effects! \Rightarrow Must use models
 - not everything has features!



Kemper et al, 2002

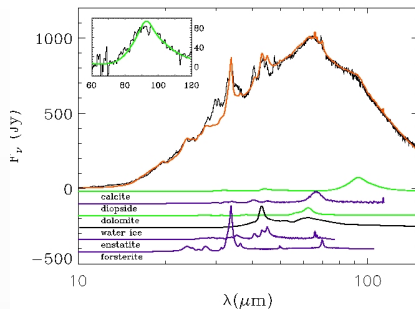
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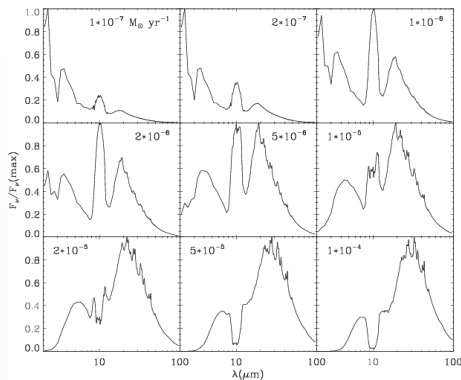
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 - need statistics to see the big picture

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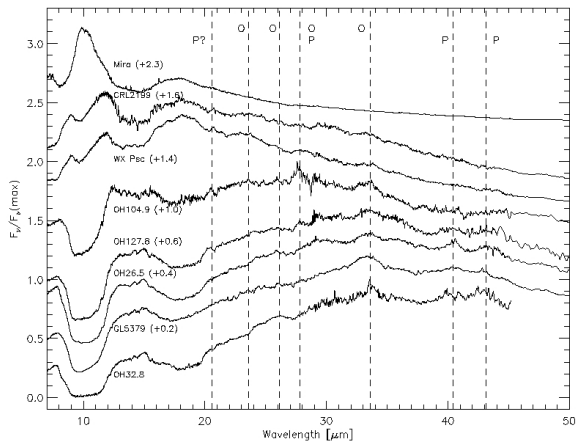
- Features small compared to overall emission
- Individual sources or small samples
 - need statistics to see the big picture
- Many confounding factors
 - RT effects,
 - grain shape
 - variability
 - ...



Kemper et al, 2001

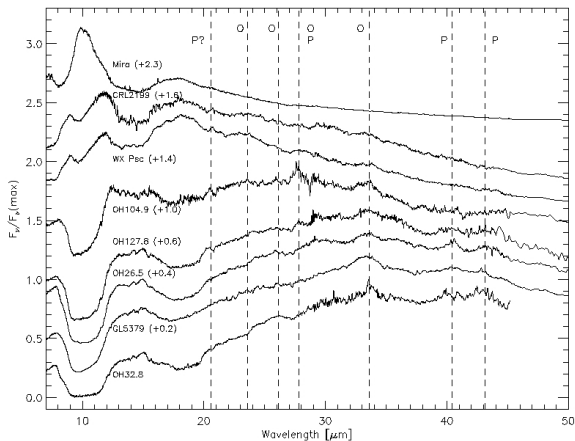
Previous attempts

e.g. Sylvester et al., 1999

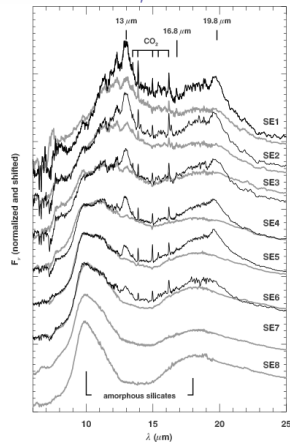


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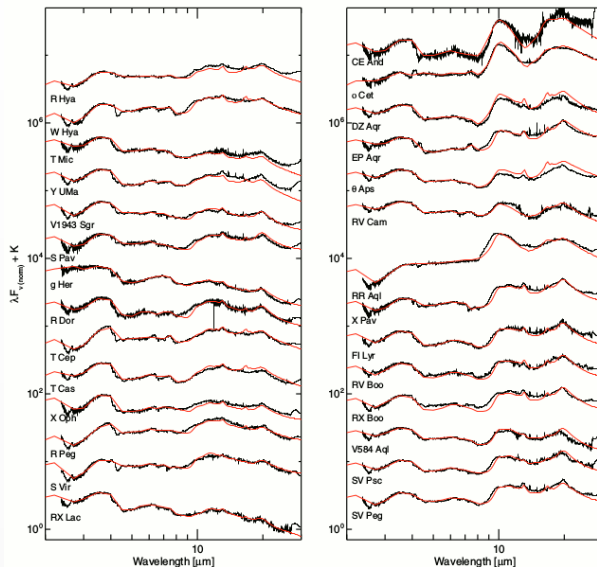


Sloan et al., 2003



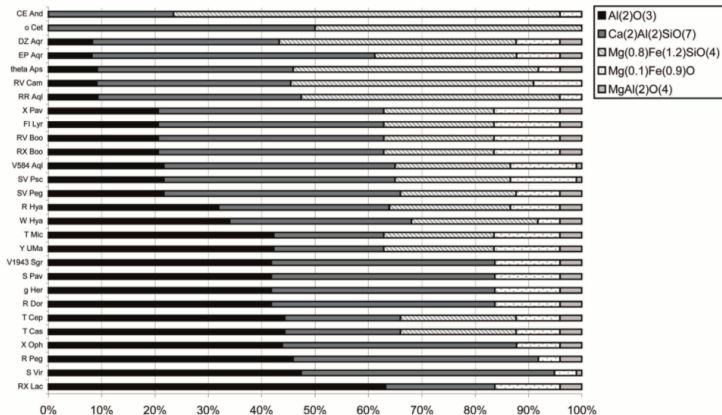
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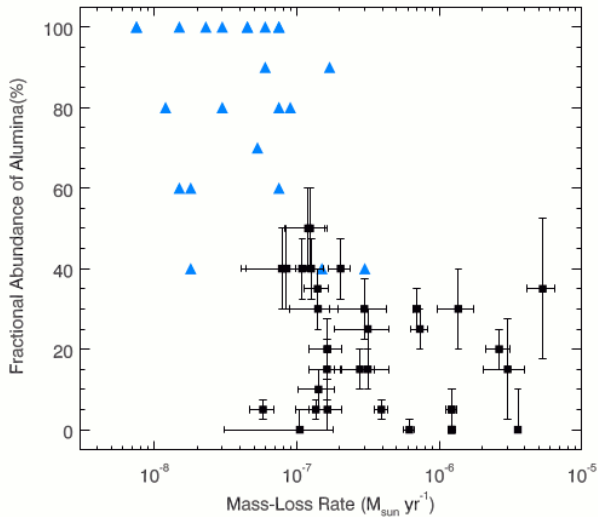
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Previous attempts - being quantitative

Jones et al., 2014



Statistical problems

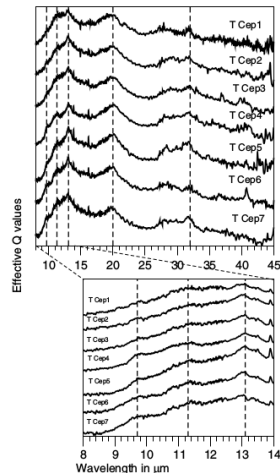
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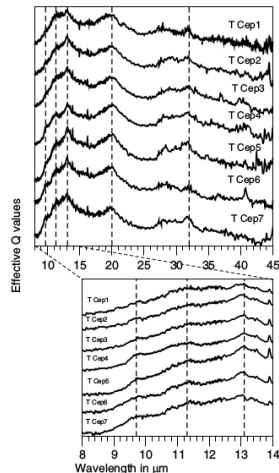
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 - Often “ χ by eye”, doesn't give any idea of uncertainties
 - Or grids, sampling often too sparse

AMPERE: A new framework

Toolkit to combine data, models and optimisers

Pure python, GPL licenced

Bayesian inference - include the problems in the model!

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- Build a *flexible* likelihood function
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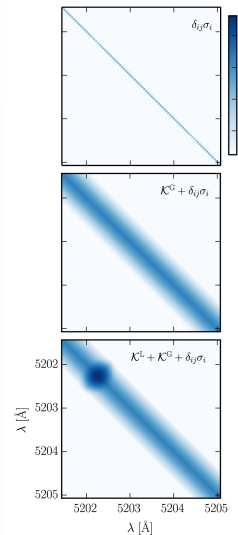
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- Also for scientific issues - variability, imperfect model (not all species, no atomic lines etc), foreground extinction
- Can use interpolation if model expensive (e.g. STARFISH, Czekala et al., 2015)

Modelling data

- Model covariance matrix to encapsulate noise
 - Simple but effective
 - Plenty of literature
- Simple additive and multiplicative terms for calibration

“Keep it simple stupid” – Kelly Johnson



Czekala et al., 2015

Current/future capabilities

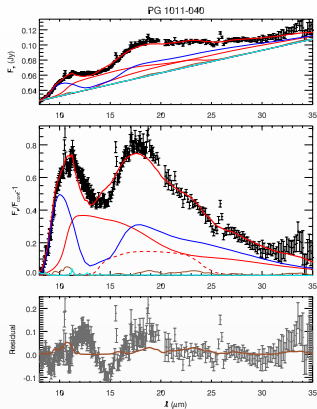
- Synthetic photometry, spectra
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- Optimisation with emcee
- simple analytical models
- ideas and contributions are welcome!

Current/future capabilities

- Synthetic photometry, spectra, *visibilities*, *images*
 - Global covariance structure, *non-stationary covariance*
- Optimisation with emcee, *parallel genetic algorithm*, *+more*
- simple analytical models, *interfaces to common RT codes*
- ideas and contributions are welcome!

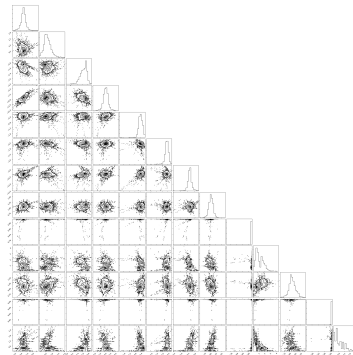
Current work

Replicating Srinivasan et al. (2017) - mineralogy of AGN winds.
Sundar's results:



“Best fit” is consistent, but uncertainty is $\gtrsim 20\%$
Definitely not Gaussian

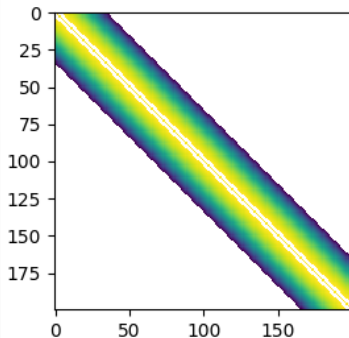
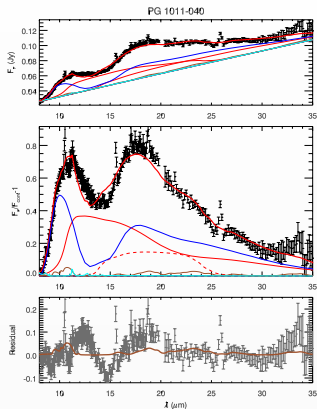
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AMPERE results:



“Best fit” is consistent, but uncertainty is $\gtrsim 20\%$

Simple covariance model, correlated uncertainties are small

Summary

- Mineralogy
 - Critical to understanding dust formation and processing
 - Many components known, but abundances are not
- Major obstacles are methodological
 - Fitting heterogeneous data
 - variability, correlated uncertainties, etc
- AMPERE is designed to tackle these issues
 - forward modelling
 - model covariance \Rightarrow optimum weighting
 - understand distributions of solutions
- Code will be public
- Input and contributions are welcome!