

RELATIVE SPUTTERING RATES OF FeS, MgS, AND Mg SILICATES: IMPLICATIONS FOR ISM GAS PHASE DEPLETIONS OF ROCK-FORMING ELEMENTS

Lindsay P. Keller¹ and Z. Rahman² ¹ARES, NASA/JSC, Houston, TX 77058 (Lindsay.P.Keller@nasa.gov), and ²Jacobs, NASA/JSC, TX 77058

- FeS and MgS sputter at a much higher rate than typical silicate minerals. ~14 S atoms sputtered from FeS for every Si atom from pyroxene or olivine.
- Preferential sputtering may account for the lack of significant depletion of S from the gas phase in the ISM.
- Irradiation effects limited to preferential sputtering of S and the formation a thin (2-3 nm) Fe^o layer on the surface. FeS doesn't amorphize.
- Rare nanophase FeS grains occur as inclusions in circumstellar amorphous silicate grains found in comet dust particles analyzed in the laboratory (Keller and Messenger 2011). These results show that a finite amount of S in the ISM occurs in solid grains.

