

# Resolving Heavy Neutral Lepton oscillations at the intensity frontier

*(Based on Coherent Oscillations of Heavy Neutral Leptons, soon on arXiv)*

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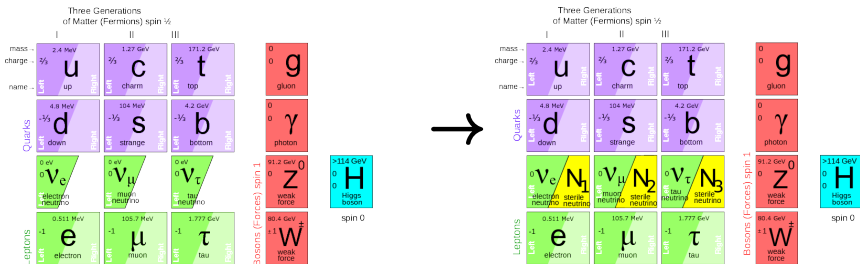
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# Why Heavy Neutral Leptons?

The SM is tremendously successful at explaining collider results, but...

- 1 Predicts massless neutrinos, but we know at least 2 are massive.
- 2 Not enough  $CP$ -violation to account for the observed BAU.
- 3 Does not provide a (particle type) Dark Matter candidate.
- 4 Its particle content hints at singlet (sterile), “right-handed”<sup>1</sup> counterparts to neutrinos  $\rightarrow N^\dagger$ .



1. Since HNLs are sterile under  $SU(2)_L$ , it does not make much sense to associate them with a specific chirality, but this denomination remains common.

# Massive neutrinos

- The SM allows two types<sup>2</sup> of masses terms for neutrinos<sup>3 4</sup>:
  - 1 Dirac:  $m_D(\nu N + N^\dagger \nu^\dagger)$ .
  - 2 Majorana (for the singlets):  $\frac{M_R}{2}(NN + N^\dagger N^\dagger)$ .
- For three generations and multiple HNLs<sup>5</sup>:
  - 1  $(m_D)_{\alpha I} \nu_\alpha N_I + \text{h.c.} \stackrel{\text{def}}{=} \nu^T m_D N + \text{h.c.}$
  - 2  $(M_R)_{IJ} N_I N_J + \text{h.c.} \stackrel{\text{def}}{=} N^T M_R N + \text{h.c.}$
- In general, both terms can be present:
  - $\rightarrow$  Dirac-Majorana mass term (here in matrix form):

$$\mathcal{L}_{D+M} = \begin{pmatrix} \nu^T & N^T \end{pmatrix} \begin{pmatrix} 0 & m_D^T \\ m_D & M_R \end{pmatrix} \begin{pmatrix} \nu \\ N \end{pmatrix} + \text{h.c.}$$

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2. Including only the existing representations of the SM gauge groups.
  3. Everything here is in two-component spinor formalism (see Dreiner, Haber, and Martin 2010 for a review), using the  $(\frac{1}{2}, 0)$  representation.
  4. For convenience, we define the left-handed HNLs  $N_a = \varepsilon_{ab}(N^\dagger \dot{a})^\dagger$ .
  5. The number of HNLs is not constrained by anything, but we often postulate one HNL per generation.

# The Type-I Seesaw mechanism

(Minkowski 1977; Gell-Mann, Ramond, and Slansky 1979; Mohapatra and Senjanović 1980; Yanagida 1980)

- Diagonalize mass matrix with an orthogonal field rotation <sup>6</sup>.

$$\mathcal{L}_{D+M} = \frac{m_i}{2} (n_i n_i + n_i^\dagger n_i^\dagger)$$

- Singular values are the physical *Majorana* masses.
- Mixing between flavor states  $\nu_\alpha$  and mass eigenstates  $n_i$ :

$$\nu_\alpha = \mathcal{U}_{\alpha i} n_i = U_{\alpha i}^{\text{PMNS}} \nu_i + \Theta_{\alpha I} N_I$$

→ *Sterile* neutrinos can still interact through (small) mixing!

- Approximate block diagonalization leads to:

$$\begin{aligned}\hat{M}_{\text{light}} &\cong -m_D^T (M_R)^{-1} m_D \\ \hat{M}_{\text{heavy}} &\cong M_R\end{aligned}$$

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6. Takagi factorization.

# Seesaw models

Different choices of parameters lead to different phenomenologies.

## 1 “Traditional Type-I Seesaw”:

- $m_D \sim v$  (natural Yukawas),  $M_R \sim 10^{15}$  GeV.
- $m_\nu \sim \frac{m_D^2}{M_R}$ : the larger  $M_R$ , the smaller  $m_\nu \rightarrow$  seesaw.
- Leads to observed  $\nu$  masses:  $\hat{M}_\nu \sim 10^{-2}$  eV.

## 2 $\nu$ MSM (Asaka and Shaposhnikov 2005):

- Some variants can explain neutrino masses, BAU and Dark Matter.
- Complete the SM with 2 nearly-degenerate HNLs  $N_{1,2}$ .<sup>7</sup>
- Include all renormalizable terms allowed by the SM symmetries.
- $m_D = -\frac{v}{\sqrt{2}} Y_{\alpha I}^\nu$  and  $M_R \sim v$ .
- Neutrinos masses are small because Yukawas are small...
- ... or new symmetry: approximate lepton number conservation.

## 3 And more (radiative seesaw, ...)

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7. And optionally add a DM candidate.

## Why HNL oscillations?

- In the  $\nu$ MSM, BAU is produced through the ARS mechanism (Akhmedov, Rubakov, and Smirnov 1998):
  - 1  $CP$ -violating oscillations between  $N_{1,2}$  produce a lepton *flavor* asymmetry.
  - 2 Difference between couplings leads to a lepton *number* asymmetry.
  - 3 This lepton number asymmetry is partially processed into a *baryon* number asymmetry by sphalerons.
- If the physical splitting  $\delta M$  is small enough, HNL oscillations can produce the observed Dark Matter abundance (Shaposhnikov 2008).

*Can we resolve oscillations and measure  $\delta M$ ?*

We need an accurate model of HNL oscillations to answer this question.

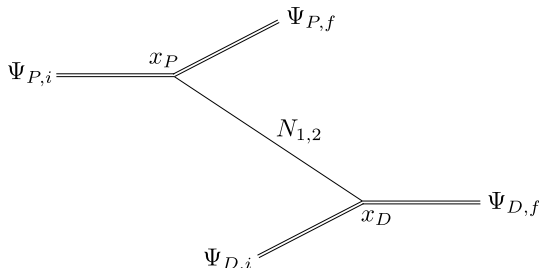
→ *External wave packet model*

# External wave packet model I

Sachs 1963; Giunti, Kim, and Lee 1993; Beuthe 2003; Akhmedov and Kopp 2010; Akhmedov and Smirnov 2011; Akhmedov, Hernandez, and Smirnov 2012

## 1 External:

- Compute the amplitude for the whole process, including the HNL production and its decay.



- Keep the phase factor  $e^{-iq_I(x_D-x_P)}$  with each internal line.
- Sum the partial amplitudes:

$$\mathcal{P}(\Psi_i \rightarrow \Psi_f) = |\mathcal{A}(\Psi_i \rightarrow \Psi_f)|^2 = |\mathcal{A}_1|^2 + |\mathcal{A}_2|^2 + 2\text{Re}(\mathcal{A}_1^* \mathcal{A}_2)$$

# External wave packet model II

## 2 Wave packet:

- Both the initial and final states have associated wave packets. For example, for a HNL produced in a semileptonic decay  $H \rightarrow H' l_\alpha N$ , then decaying to as  $N \rightarrow H'' l_\beta$ :

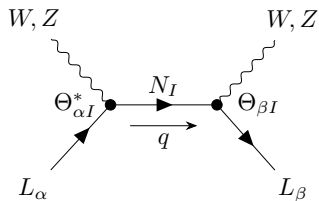
$$|\psi_I\rangle = \int d\Omega_p \psi_I(\mathbf{p}) |H(\mathbf{p})\rangle$$

$$|\psi_F\rangle = \int d\Omega_{k_1} d\Omega_{k_2} d\Omega_{k_3} d\Omega_{k_4} \psi_F(\mathbf{k}_1, \mathbf{k}_2, \mathbf{k}_3, \mathbf{k}_4) \cdot |H'(\mathbf{k}_1), l_\alpha(\mathbf{k}_2), H''(\mathbf{k}_3), l_\beta(\mathbf{k}_4)\rangle$$

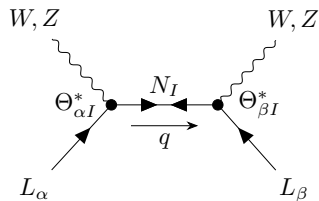
- Interaction vertices are *localized*: no integration over  $x_{P,D}$ .
- Necessary to have overlap between initial and final states while conserving momentum at *each* vertex (see Cohen, Glashow, and Ligeti 2009).
- If it is possible to tell the HNLs apart with a sufficiently precise measurement of the external momenta or of the propagation time (due to dispersion), then decoherence occurs.



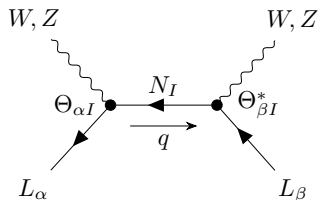
# Propagators for Majorana particles



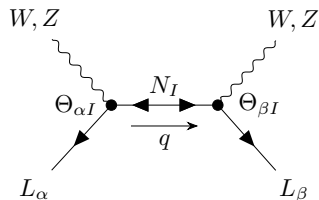
(a)  $L \rightarrow L$



(b)  $L \rightarrow R$



(c)  $R \rightarrow R$



(d)  $R \rightarrow L$

## Coherent cross-section

- For the coherent cross-section, the extra phases  $e^{-i\mathbf{q}_I \cdot (x_D - x_P)}$  coming from on-shell propagators are relevant.
- If we integrate wave packets<sup>8</sup> over 3-momentum  $\mathbf{q}$ , then for each eigenstate the energy is  $E_I = \sqrt{M_I^2 + \mathbf{q}^2}$ .
- In the ultra-relativistic limit:  $q_I \cdot (x_D - x_P) \cong \frac{M_I^2}{2|\mathbf{q}|} L$ .
- Squaring the amplitude and including self-energy corrections  $M_I^2 \rightarrow M_I^2 - iM_I\Gamma_I$ , we can express the (differential) coherent cross-section in terms of the incoherent one:

$$(d\sigma_{\text{coh}})_{\alpha\beta}^{\pm\pm} = \frac{\left| \sum_I \Theta_{\alpha I}^{\mp} \Theta_{\beta I}^{\pm} e^{-i\frac{M_I^2}{2E} L - \frac{\Gamma_I}{2} \tau} \right|^2}{\sum_J |\Theta_{\alpha J}|^2 |\Theta_{\beta J}|^2 e^{-\Gamma_J \tau}} (d\sigma_{\text{inc}})_{\alpha\beta}^{\pm\pm}$$

where we have defined  $\Theta^- \stackrel{\text{def}}{=} \Theta$  and  $\Theta^+ \stackrel{\text{def}}{=} \Theta^*$ .

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8. There is neither fixed  $E$  nor fixed  $\mathbf{q}$ , but rather overlap between wave packets.

## Quasi-Dirac limit

- Current facilities can only probe HNLs with large (“unnatural”) mixing with flavor fields.
- In the large-couplings limit, 2 HNLs form a quasi-Dirac pair<sup>9</sup>, i.e.  $\Theta_{\alpha 2} = \pm i\Theta_{\alpha 1}$ .
- A new symmetry, leading to *approximate* lepton-number conservation, can be postulated to produce nearly-degenerate HNLs with large couplings while avoiding fine-tuning (Shaposhnikov 2007).
- The maximal mixing between the two HNLs leads to LNC-LNV oscillations (Anamiati, Hirsch, and Nardi 2016; Antusch, Cazzato, and Fischer 2017).

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9. This is easily seen from the Casas-Ibarra parameterization, and more generally this is a consequence of the argument from Kersten and Smirnov 2007.

# LNC-LNV asymmetry

## 1 Long-lived HNLs at colliders:

- Displaced vertex.
- Primary and secondary leptons may have opposite (LNC) or same sign (LNV), and in general they can have different flavors.
- LNC-LNV asymmetry oscillates as a function of  $\tau = \sqrt{(x_D - x_P)^2}$ .

## 2 Short-lived HNLs at colliders:

- Integrated LNC-LNV asymmetry  $R_{ll}$  can tell us whether  $\delta M \lesssim \Gamma$ .

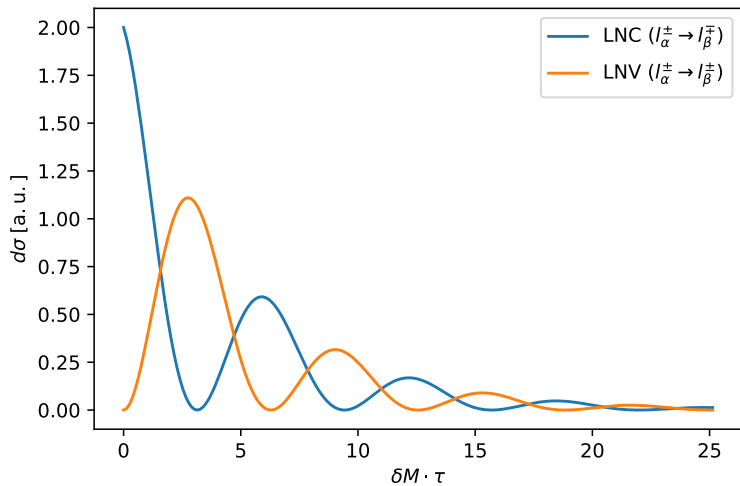
## 3 Beam-dump experiments:

- Primary lepton usually inside the target  $\rightarrow$  not visible.
- HNLs produced in the decay of heavy mesons:  $D, D_s, B, B_c \dots$
- Beam-dump experiments typically produce equal amounts of anti-mesons.  
 $\Rightarrow$  We cannot reliably use the LNC-LNV asymmetry here!<sup>10</sup>

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10. We can in principle see HNL disappearance close to the seesaw bound (lowest possible couplings), but no experiment is sensitive to this region yet.

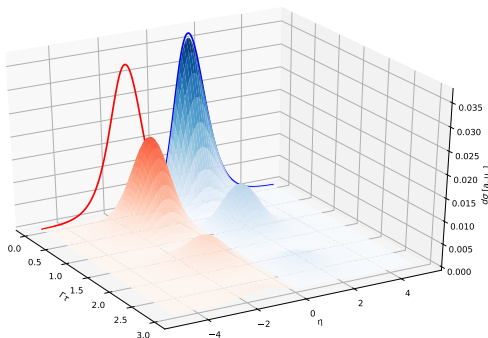
# LNC-LNV asymmetry



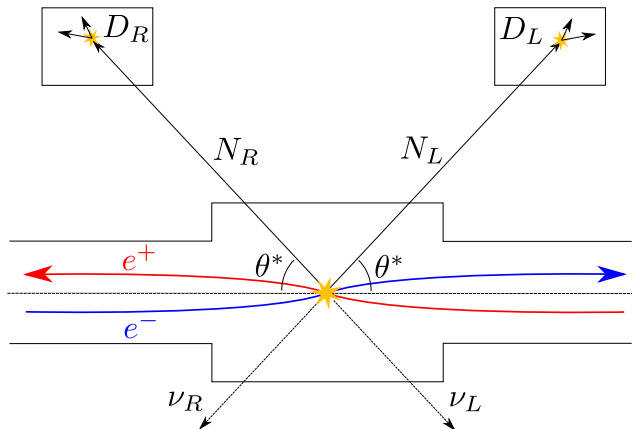
# Angular distribution

Here, we generalize results from Hernández, Jones-Pérez, and Suárez-Navarro 2018.

- Prompt HNLs produced at an  $e^+e^-$  collider, along with light  $\nu$ .
- Bin HNL candidate events by secondary lepton charge and flavor, as well as HNL proper lifetime  $\tau$  and pseudorapidity  $\eta$ .
- Conservation of total spin will lead to non-trivial correlations between these parameters. For a secondary  $l_{\beta}^-$ :



## Detecting small- $\delta M$ oscillations at $e^+e^-$ colliders



- Opposite, non-trivial charge asymmetries in both detectors would indicate that HNLs have oscillated.
- A similar effect might be used to break the accidental symmetry at beam-dump experiments.

# Conclusion

- HNLs are a primary target for future intensity frontier facilities.
- Observing their oscillations and measuring  $\delta M$  would put strong constraints on BAU / leptogenesis.
- There are several ways to detect their oscillations, using charge asymmetries and kinematics.
- HNL oscillations are best studied using scattering theory.
- When studying displaced vertices, the phases of on-shell propagators should be kept in the calculations, otherwise only the integrated effect of oscillations is visible.
- Oscillations must be allowed for in displaced vertex searches.



Questions?

