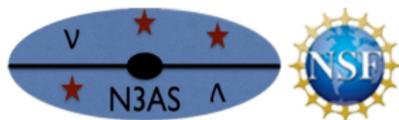


# Eigenvalues & eigenstates of the many-body collective neutrino oscillation problem

Amol V. Patwardhan

N3AS postdoctoral fellow  
UW-Madison, UC Berkeley

(@INT from Sep 2019)



August 28, 2019

## References

-  Amol V. Patwardhan, Michael J. Cervia, and A. B. Balantekin  
Phys. Rev. D 99, 123013 (2019), arXiv:1905.04386
-  Michael J. Cervia, Amol V. Patwardhan, and A. B. Balantekin  
IJMPE 28 (2019) 1950032, arXiv:1905.00082
-  Michael J. Cervia, Amol V. Patwardhan, A. B. Balantekin,  
S. N. Coppersmith, and Calvin W. Johnson  
arXiv:1908.03511
-  Ermal Rrapaj  
arXiv:1905.13335

## Bedtime reading

-  Savas Birol, Y. Pehlivan, A. B. Balantekin, and T. Kajino  
Phys. Rev. D 98, 083002 (2018), arXiv:1805.11767
-  Y. Pehlivan, A. B. Balantekin, Toshitaka Kajino, and Takashi Yoshida  
Phys. Rev. D 84, 065008 (2011), arXiv:1105.1182
-  Alexandre Faribault, Omar El Araby, Christoph Sträter, and Vladimir Gritsev  
Phys. Rev. B 83, 235124 (2011), arXiv:1103.0472
-  Pieter W. Claeys  
arxiv:1809.04447 (PhD thesis)

# Today's talk

- Collective neutrino oscillations as a many-body problem
- Eigenvalues and eigenstates of neutrino many-body Hamiltonian, for a two-flavor, single-angle system
- Talks by Michael Cervia and Ermal Rrapaj:  
flavor evolution of a neutrino many-body system, and comparison with the mean-field description

# Outline

- 1 Many-body treatment of neutrino oscillations
- 2 Solving the Bethe Ansatz equations
- 3 Eigenvalues and Eigenstates

# Neutrino oscillations: flavor/mass isospin operators

- Denote Fermionic operators for neutrino flavor/mass states as  $a_\alpha(\mathbf{p})$ ,  $a_j(\mathbf{p})$ , where  $\alpha = e, x$ , and  $j = 1, 2$

$$a_e(\mathbf{p}) = \cos \theta a_1(\mathbf{p}) + \sin \theta a_2(\mathbf{p})$$

$$a_x(\mathbf{p}) = -\sin \theta a_1(\mathbf{p}) + \cos \theta a_2(\mathbf{p})$$

- Introduce the mass-basis isospin operators

$$J_{\mathbf{p}}^+ = a_1^\dagger(\mathbf{p})a_2(\mathbf{p}) , \quad J_{\mathbf{p}}^- = a_2^\dagger(\mathbf{p})a_1(\mathbf{p}) ,$$

$$J_{\mathbf{p}}^z = \frac{1}{2} \left( a_1^\dagger(\mathbf{p})a_1(\mathbf{p}) - a_2^\dagger(\mathbf{p})a_2(\mathbf{p}) \right) ,$$

which obey the usual  $SU(2)$  commutation relations

$$[J_{\mathbf{p}}^+, J_{\mathbf{q}}^-] = 2\delta_{\mathbf{p}\mathbf{q}}J_{\mathbf{p}}^z , \quad [J_{\mathbf{p}}^z, J_{\mathbf{q}}^\pm] = \pm\delta_{\mathbf{p}\mathbf{q}}J_{\mathbf{p}}^\pm .$$

# Neutrino oscillations: many-body Hamiltonian

- Vacuum oscillations:

$$\begin{aligned} H_{\text{vac}} &= \sum_{\mathbf{p}} \left( \frac{m_1^2}{2p} a_1^\dagger(\mathbf{p}) a_1(\mathbf{p}) + \frac{m_2^2}{2p} a_2^\dagger(\mathbf{p}) a_2(\mathbf{p}) \right) \\ &= \sum_{\omega} \omega \vec{B} \cdot \vec{J}_{\omega} , \end{aligned}$$

where  $\omega = \frac{\delta m^2}{2p}$ ,  $\vec{J}_{\omega} = \sum_{p = \frac{\delta m^2}{2\omega}} \vec{J}_{\mathbf{p}}$ , and

$$\vec{B} = (0, 0, -1)_{\text{mass}} = (\sin 2\theta, 0, -\cos 2\theta)_{\text{flavor}} .$$

- Neutrino-neutrino interactions

$$H_{\nu\nu} = \frac{\sqrt{2}G_F}{V} \sum_{\mathbf{p}, \mathbf{q}} (1 - \cos \vartheta_{\mathbf{p}\mathbf{q}}) \vec{J}_{\mathbf{p}} \cdot \vec{J}_{\mathbf{q}} .$$

# Neutrino Hamiltonian: single-angle approximation

- Suitable averaging over the angle  $\vartheta_{\mathbf{pq}}$  to simplify the problem

$$H_{\nu\nu} \approx \frac{\sqrt{2}G_F}{V} \langle (1 - \cos \vartheta_{\mathbf{pq}}) \rangle \vec{J} \cdot \vec{J} \quad (\text{includes } J_p^2 \text{ terms})$$
$$\equiv \mu(r) \vec{J} \cdot \vec{J}, \quad \text{where } \vec{J} = \sum_{\omega} \vec{J}_{\omega}$$

- Combined neutrino Hamiltonian (vacuum + self-interactions)

$$H_{\nu} = \sum_{\omega} \omega \vec{B} \cdot \vec{J}_{\omega} + \mu(r) \vec{J} \cdot \vec{J}.$$

# Mean-field (random phase) approximation

- This method yields the effective one-particle neutrino Hamiltonian

$$H \sim H^{\text{RPA}} = \sum_{\omega} \omega \vec{B} \cdot \vec{J}_{\omega} + \mu \vec{P} \cdot \vec{J},$$

where  $\vec{P}_{\omega} = 2\langle \vec{J}_{\omega} \rangle$  is the “Polarization vector”, and  $\vec{P} = \sum_{\omega} \vec{P}_{\omega}$

- The self-consistency requirement of the mean-field approach then implies that  $\vec{P}_{\omega}$  must satisfy

$$\frac{d}{dt} \vec{P}_{\omega} = (\omega \vec{B} + \mu \vec{P}) \times \vec{P}_{\omega}$$

# Many-body eigenstates: limiting $\mu$ cases

- Many-body neutrino Hamiltonian (two-flavor, single-angle):

$$H_\nu = \sum_{p=1}^M \omega_p \vec{B} \cdot \vec{J}_p + \mu(r) \vec{J} \cdot \vec{J},$$

where  $p$  is an index for the  $\omega$ s in the system,  $M$  in number

- As  $\mu \rightarrow 0$ , eigenstates are simply tensor products of single-particle vacuum mass eigenstates, e.g.,

$$|\nu_1 \nu_1 \nu_1 \dots\rangle, \quad |\nu_2 \nu_1 \nu_1 \dots\rangle, \quad |\nu_1 \nu_2 \nu_1 \dots\rangle, \quad \dots$$

- As  $\mu \rightarrow \infty$ , the total isospin states  $|j, m\rangle_{\text{mass/ flavor}}$  are eigenstates, with eigenvalues  $\mu j(j+1)$

# Many-body eigenstates: Richardson-Gaudin procedure

- For  $0 < \mu < \infty$ , the extremal states  $|\nu_1, \dots, \nu_1\rangle$  and  $|\nu_2, \dots, \nu_2\rangle$  are eigenstates, with energies

$$E_{(\pm N/2)} = \mp \sum_{p=1}^M \omega_p \frac{N_p}{2} + \mu \frac{N}{2} \left( \frac{N}{2} + 1 \right),$$

$N_p$  the number of neutrinos at frequency  $\omega_p$ , and  $N = \sum_p N_p$ .

- Other eigenstates can be constructed from the extremal states by applying Gaudin operators

$$S^{\pm}(\zeta_{\alpha}) = \sum_{p=1}^M \frac{J_p^{\pm}}{\omega_p - \zeta_{\alpha}},$$

where  $\zeta_{\alpha}$  are parameters to be determined

# Many-body eigenstates: Bethe Ansatz equations

- Bethe Ansatz:  $|\zeta_\alpha\rangle \equiv S^-(\zeta_\alpha)|\frac{N}{2}, \frac{N}{2}\rangle$  is an eigenstate of  $H_\nu$  if  $\zeta_\alpha$  satisfies the equation

$$-\frac{1}{2\mu} - \sum_{p=1}^M \frac{j_p}{\omega_p - \zeta_\alpha} = 0,$$

where  $j_p = N_p/2$

- More generally,  $|\zeta_1, \dots, \zeta_\kappa\rangle \equiv S^-(\zeta_1) \dots S^-(\zeta_\kappa)|\frac{N}{2}, \frac{N}{2}\rangle$  is an eigenstate of the Hamiltonian with the energy

$$E(\zeta_1, \dots, \zeta_\kappa) = E_{N/2} + \sum_{\alpha=1}^{\kappa} \zeta_\alpha - \kappa\mu(N - \kappa + 1),$$

if  $\zeta_1, \dots, \zeta_\kappa$  obey the system of equations

$$-\frac{1}{2\mu} - \sum_{p=1}^M \frac{j_p}{\omega_p - \zeta_\alpha} = \sum_{\substack{\beta=1 \\ (\beta \neq \alpha)}}^{\kappa} \frac{1}{\zeta_\alpha - \zeta_\beta}.$$

# Outline

- 1 Many-body treatment of neutrino oscillations
- 2 Solving the Bethe Ansatz equations**
- 3 Eigenvalues and Eigenstates

# Finding roots of Bethe Ansatz equations

- For a given  $\kappa$ , the Bethe Ansatz (BA) equations are a system of simultaneous equations in  $\kappa$  complex variables  $\{\zeta_1, \dots, \zeta_\kappa\}$ . Singularities present as  $\zeta_\alpha$ s approach one another
- To avoid singularities, can be converted to polynomial equations where the order of each polynomial is  $M + \kappa - 2$  ( $M$  is the number of frequencies in the system)
- To obtain a complete set of eigenstates, one must solve BA equations  $\forall \kappa \in \{1, \dots, N\}$ : a different set of equations must be solved for each  $\kappa$

# An alternative approach: the 'Lambda' method

- Defining  $\Lambda(\lambda) = \sum_{\alpha=1}^{\kappa} \frac{1}{\lambda - \zeta_{\alpha}}$ , one can convert BA equations into an ODE, which thereafter can be reduced to another algebraic system by taking the limit  $\lambda \rightarrow \omega_q$

- One obtains the following set of equations for each  $\omega_q$ :

$$\Lambda^2(\omega_q) + (1 - 2j_q)\Lambda'(\omega_q) + \frac{1}{\mu}\Lambda(\omega_q) = 2 \sum_{\substack{p=1 \\ p \neq q}}^M j_p \frac{\Lambda(\omega_q) - \Lambda(\omega_p)}{\omega_q - \omega_p}.$$

- In particular, if  $j_q = 1/2$ , the  $\Lambda'(\omega_q)$  term disappears and the equation becomes purely algebraic in all the  $\Lambda(\omega_p)$ . If  $j_q > 1/2$ , taking derivatives of the ODE w.r.t  $\lambda$  before taking the  $\lambda \rightarrow \omega_q$  limit gives additional equations which can be used to eliminate all  $\Lambda'(\omega_p), \Lambda''(\omega_p), \dots$ , etc.

## The $j_q = 1/2$ case: algebraic $\Lambda$ equations

- Denote  $\Lambda_q = \Lambda(\omega_q)$  and define  $\tilde{\Lambda}_q = \mu\Lambda_q$ . The case where  $j_q = 1/2 \forall q$  (i.e., only one neutrino at each  $\omega$ ) reduces to

$$\tilde{\Lambda}_q^2 + \tilde{\Lambda}_q = \mu \sum_{\substack{p=1 \\ p \neq q}}^N \frac{\tilde{\Lambda}_q - \tilde{\Lambda}_p}{\omega_q - \omega_p}.$$

- System of size  $N$  represented by  $N$  coupled polynomial equations, of quadratic order regardless of  $N$  or  $\kappa$ . In fact, the same set of equations admits solutions for all  $\kappa = 0, \dots, N$ . And there are no singularities.
- Solutions corresponding to a particular  $\kappa$  satisfy

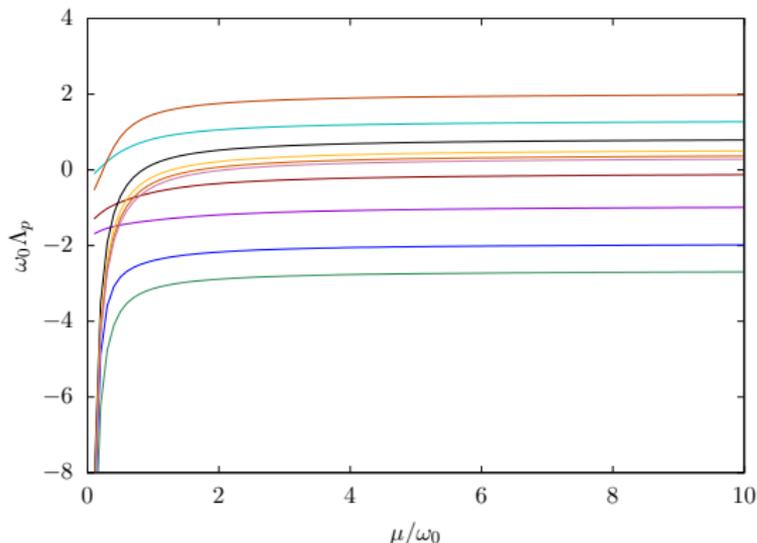
$$\sum_p \tilde{\Lambda}_p = -\kappa,$$

# Solving $\Lambda$ equations by homotopy continuation

$$\tilde{\Lambda}_q^2 + \tilde{\Lambda}_q = \mu \sum_{\substack{p=1 \\ p \neq q}}^N \frac{\tilde{\Lambda}_q - \tilde{\Lambda}_p}{\omega_q - \omega_p}.$$

- In the  $\mu = 0$  limit, each  $\tilde{\Lambda}_q$  can take the value 0 or  $-1$ , giving a total of  $2^N$  solutions. In each solution, the number of  $\tilde{\Lambda}_q$ s that are  $-1$  is the  $\kappa$  of that particular solution.
- Each of the  $2^N$  solutions for  $\mu > 0$  can be constructed incrementally, starting from the corresponding  $\mu = 0$  solution:
  - Use each  $\tilde{\Lambda}_q$  solution obtained at  $\mu = \mu_n$  to construct a starting guess for the corresponding solution at  $\mu = \mu_n + \delta\mu$
  - Improve the starting guess using iterative numerical methods such as Newton-Raphson

# Ten neutrino system: $\{\Lambda(\omega_q)\}$ solutions



[Patwardhan, Cervia, Balantekin, Phys. Rev. D 99 123013 (2019)]

**Figure:** A solution  $\{\Lambda(\omega_1), \dots, \Lambda(\omega_{10})\}$  as a function of  $\mu$ , for a system with ten neutrinos, distributed uniformly across frequencies  $\omega_q = q\omega_0$ . Shown here is one sample solution out of the 1024 for this system.

# Outline

- 1 Many-body treatment of neutrino oscillations
- 2 Solving the Bethe Ansatz equations
- 3 Eigenvalues and Eigenstates**

# Energy eigenvalues from $\Lambda_s$

- Recall that energy eigenvalues given by:

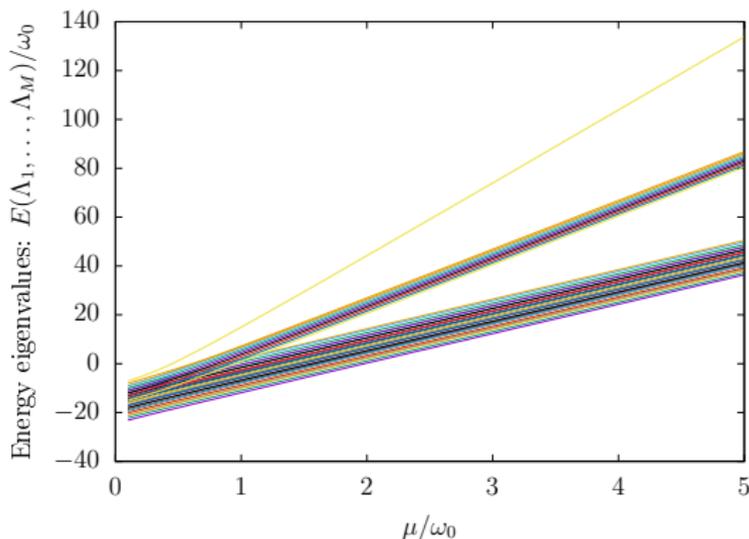
$$E(\zeta_1, \dots, \zeta_\kappa) = E_{N/2} + \sum_{\alpha=1}^{\kappa} \zeta_\alpha - \kappa\mu(N - \kappa + 1)$$

- In terms of  $\tilde{\Lambda}_q$ , these may be rewritten as

$$E(\tilde{\Lambda}_1, \dots, \tilde{\Lambda}_N) = - \sum_p \frac{\omega_p}{2} + \mu \frac{N}{2} \left( \frac{N}{2} + 1 \right) - \sum_p \omega_p \tilde{\Lambda}_p$$

- Instructive to group energy eigenvalues according to  $\kappa$ : solutions are eigenstates of  $J_z$ , with eigenvalue  $m = N/2 - \kappa$ . Within each  $\kappa$ , energy eigenvalues split into branches at large  $\mu$ . Each branch associates to a particular  $|j, m\rangle$  as  $\mu \rightarrow \infty$ , with  $j = |m|, \dots, N/2$ .

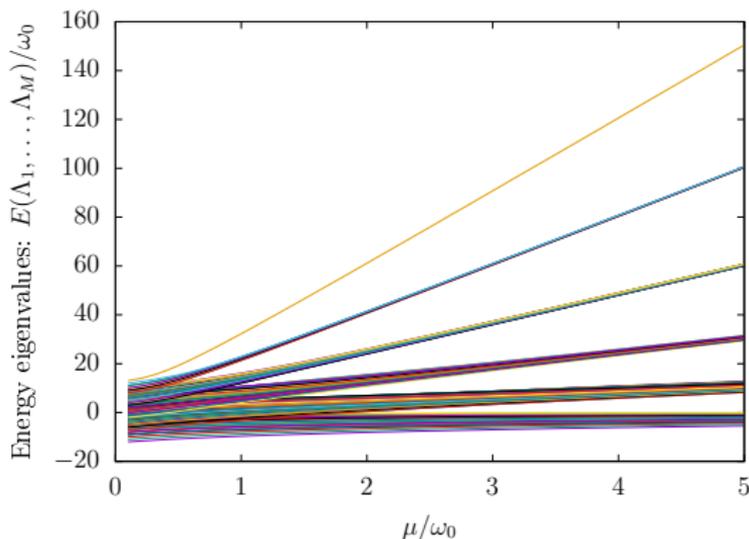
# Ten neutrino system: energy eigenvalues



[Patwardhan, Cervia, Balantekin, Phys. Rev. D 99 123013 (2019)]

**Figure:** Energy eigenvalues corresponding to all  $\kappa = 2$  ( $m = 3$ ) solutions of the BA equations, for an  $N = 10$  neutrino system, as functions of  $\mu$ .

# Ten neutrino system: energy eigenvalues



[Patwardhan, Cervia, Balantekin, Phys. Rev. D 99 123013 (2019)]

**Figure:** Energy eigenvalues corresponding to all  $\kappa = 5$  ( $m = 0$ ) solutions of the BA equations, for an  $N = 10$  neutrino system, as functions of  $\mu$ .

# Energy eigenstates from $\Lambda_s$

- The process of calculating the eigenstates from the  $\Lambda_s$  involves first calculating the power sums  $P_f = \sum_{\alpha=1}^{\kappa} (S_{\alpha}^{-})^f$ , where  $S_{\alpha}^{-} \equiv S^{-}(\zeta_{\alpha})$  are the Gaudin operators defined earlier. In terms of  $\Lambda_s$ , the power sums are given by

$$\begin{aligned}
 P_f &= \sum_{\alpha=1}^{\kappa} (S_{\alpha}^{-})^f = \sum_{i=1}^{\kappa} \sum_{p_1=1}^M \cdots \sum_{p_f=1}^M \frac{J_{p_1}^{-} \cdots J_{p_f}^{-}}{(\omega_{p_1} - \zeta_{\alpha}) \cdots (\omega_{p_f} - \zeta_{\alpha})} \\
 &= \sum_{p_1=1}^M \cdots \sum_{p_f=1}^M J_{p_1}^{-} \cdots J_{p_f}^{-} \sum_{m=1}^f \Lambda_{p_m} \prod_{\substack{l=1 \\ l \neq m}}^f \frac{1}{\omega_{p_l} - \omega_{p_m}}.
 \end{aligned}$$

- For an eigenstate with a particular  $\kappa$ , the power sums  $P_f$  for  $f = 1, \dots, \kappa$  are required

# Energy eigenstates from $\Lambda_s$

- Recognize that

$$|\zeta_1, \dots, \zeta_\kappa\rangle = e_\kappa(S_1^-, \dots, S_\kappa^-) \left| \frac{N}{2}, \frac{N}{2} \right\rangle$$

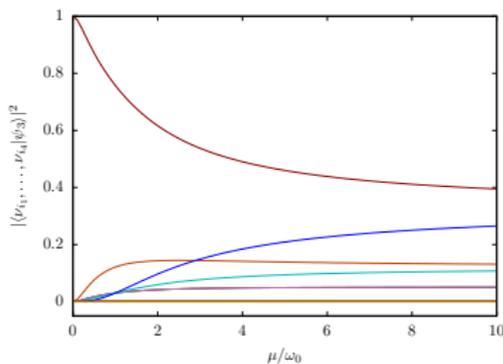
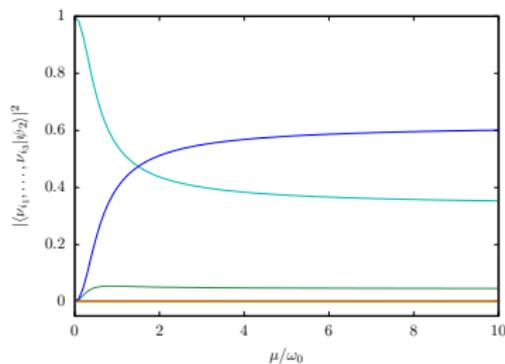
where  $e_\kappa$  is the  $\kappa$ -th elementary symmetric polynomial

- The elementary symmetric polynomials may be calculated recursively from the power sums using Newton's identities:

$$e_k = \frac{1}{k} \sum_{i=1}^k (-1)^{i-1} e_{k-i} P_i,$$

for  $k = 1, \dots, \kappa$ . Calculating  $e_\kappa$  for a particular set of  $\Lambda_s$  immediately yields the corresponding eigenstate

# Energy eigenstates: examples



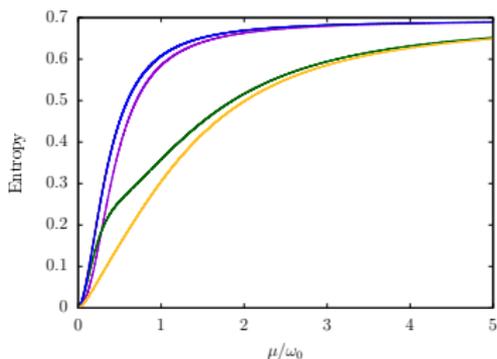
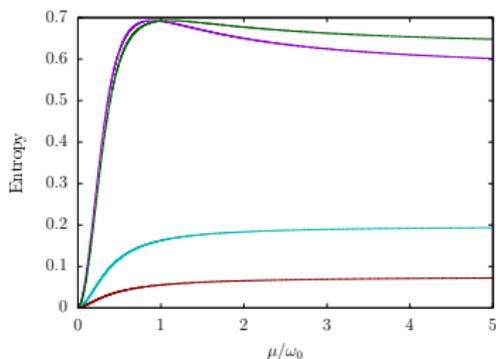
[Patwardhan, Cervia, Balantekin, Phys. Rev. D 99 123013 (2019)]

**Figure:** Overlaps,  $|\langle \nu_{i_1}, \dots, \nu_{i_N} | \psi_n \rangle|^2$ , of particular energy eigenstates  $|\psi_n\rangle$  with the mass basis states (with  $i_1, \dots, i_N = 1, 2$ ), as functions of  $\mu$ . Left: a particular eigenstate with  $\kappa = 1$ , of an  $N = 3$  system. Right: a particular eigenstate with  $\kappa = 2$ , of an  $N = 4$  system. Observe that, for  $\mu > 0$ , a state with a certain  $\kappa$  has  ${}^N C_\kappa$  nontrivial components.

# Conclusions

- Calculations of collective neutrino flavor evolution typically rely on a 'mean-field', i.e., effective one-particle description
- Important to test the efficacy and/or limitations of the mean-field by performing many-body calculations
- Evolution in the many-body case can be studied by calculating the eigenvalues and eigenvectors of the Hamiltonian by solving the Bethe Ansatz equations
- For certain simple systems, qualitative differences in flavor evolution observed between many-body and mean-field treatments, resulting from entangled states which are absent in the mean-field limit—**talks by Michael and Eermal to follow**

# Entanglement: a preview



[Cervia et al., arxiv:1908.03511]

**Figure:** Entropy of entanglement between the neutrino at frequency  $\omega_4$  and the rest of the ensemble, for all eigenstates of an  $N = 4$  neutrino system, corresponding to  $\kappa = 1$  (left) and 2 (right).

# The N3AS collaboration

- Network in Neutrinos, Nuclear Astrophysics, and Symmetries
- Multi-institutional network (3 centers + 8 sites) dedicated to recruiting and training postdocs, fostering collaborative efforts, and advancing research in the following areas:
  - Neutrino physics and astrophysics
  - Dense matter
  - Dark matter
- Funded by National Science Foundation (NSF) and Heising-Simons Foundation
- <https://n3as.wordpress.com/>

## Simple case: two neutrinos, $j_1 = j_2 = 1/2$

Defining  $\Lambda_p \equiv \Lambda(\omega_p)$  and  $\eta = \omega_2 - \omega_1$ , the equations are:

$$\Lambda_1^2 + \frac{1}{\mu}\Lambda_1 = \frac{\Lambda_1 - \Lambda_2}{-\eta}$$
$$\Lambda_2^2 + \frac{1}{\mu}\Lambda_2 = \frac{\Lambda_2 - \Lambda_1}{\eta}.$$

The following solutions can be obtained analytically

$$\Lambda_1 = \Lambda_2 = 0, \quad \text{or} \quad \Lambda_1 = \Lambda_2 = -\frac{1}{\mu}, \quad \text{or}$$

$$\Lambda_1 = -\left(\frac{1}{2\mu} + \frac{1}{\eta}\right) \pm \frac{1}{2} \sqrt{\frac{1}{\mu^2} + \frac{4}{\eta^2}}$$

$$\Lambda_2 = -\left(\frac{1}{2\mu} - \frac{1}{\eta}\right) \mp \frac{1}{2} \sqrt{\frac{1}{\mu^2} + \frac{4}{\eta^2}}.$$

The three sets of solutions correspond to  $\kappa = 0, 2$ , and  $1$ , respectively