Navier-Stokes to Maxwell via Einstein

Cindy Keeler

Arizona State University

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arXiv:2005.04242 with T. Manton and N. Monga; forthcoming with N. Monga

Overview

Outline

- From Einstein to Maxwell: the classical double copy via Weyl
- From Navier-Stokes to Einstein: fluid-gravity duality via a cutoff
- Algebraic Speciality in Fluids
 - Type D Fluids: constant vorticity
 - Type N Fluids: potential flows
- Towards a general fluid?

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From Einstein to Maxwell: The Classical Double Copy

Yang-Mills amplitudes \mathcal{A}^{YM} (properly gauged) 'square' to gravity amplitudes $\mathcal{M}^{\text{grav}}$:

$$\mathcal{A}^{\sf YM} \sim \sum_k rac{n_k c_k}{props} \quad \longrightarrow \quad \mathcal{M}^{\sf grav} \sim \sum_k rac{n_k n_k}{props}$$

Also scalar theory with amplitudes $\mathcal{A}^s \sim \sum_k c_k \tilde{c}_k/props$ For review see ch 10 of Bern, Carrasco, Chiodaroli, Johansson, Roiban 1909.01358

Kerr-Schild Map (Monteiro, O'Connell, White 1410.0239)

Pick metric in Kerr-Schild coordinates (with $k^2 = 0$):

$$g_{\mu\nu} = \eta_{\mu\nu} + \phi k_{\mu} k_{\nu} \qquad \longrightarrow G_{\mu\nu} = 0$$

$$A_{\mu} = \phi k_{\mu} \qquad \longrightarrow \nabla_{\nu} F^{\mu\nu} = 0$$

$$\phi \qquad \longrightarrow \nabla^{2} \phi = 0$$

Note our color factors will always be trivial, so we are restricting to the U(1) sector.

From Einstein to Maxwell: The Weyl Classical Double Copy

For Type D/N spacetimes with principal null vectors aligning in pairs/all four align

■ Rewrite Weyl tensor in spinor notation:

$$C_{ABCD} = \frac{1}{4} W_{\mu\nu\lambda\gamma} \sigma_{AB}^{\mu\nu} \sigma_{CD}^{\dot{\lambda}\dot{\gamma}}$$

- Decompose in principle spinors $C_{ABCD} = \alpha_{(A}\beta_B\gamma_C\delta_{D)}$
- $\blacksquare C_{ABCD}^{D} \sim \alpha_{(A}\alpha_{B}\beta_{C}\beta_{D)}, C_{ABCD}^{N} \sim \alpha_{(A}\alpha_{B}\alpha_{C}\alpha_{D)}$

For these special spacetimes, can 'square root' the Weyl tensor:

$$C_{ABCD} = \frac{1}{S} f_{(AB} f_{CD)},$$
 with e.g. $f_{(AB)} = \alpha_{(A} \beta_{B)}$

and $\nabla_0^2 S = 0$. Spinor $f_{AB} \to F_{\mu\nu}$ which satisfies $\nabla_0^{\mu} F_{\mu\nu} = 0$.

Luna, Monteiro, Nicholson, O'Connell 1810.08183;

Godazgar², Monteiro, Veiga, Pope 2010.02925

Why Fluid-Gravity Duality?

Questions from the Classical Double Copy

- Why is there a spacetime (not momentum space) double copy?

 ∃ linearized derivation via twistors and the Penrose transform

 White 2012.02479; Chacon, Nagy, White 2103.16441
- Can we extend the classical copy to Petrov type II or type I solutions? 2012.02479: type III example, 2103.16441: (linearized) whenever ∃ a Penrose transform Fluid-gravity gives physically interesting type II solutions to test
- Can we go beyond the linearized level?
 Cutoff fluid-gravity duality is non-linear but still perturbative.
 Good forum to ask.
- Can we build a Weyl double copy in higher dimensions? Yes, for Schwarzschild Monteiro, Nicholson, O'Connell 1809.03906 Fluid-grav duals generalize to higher d; 3+1d fluid=5d grav.

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History of Fluid/Gravity Duality

Membrane Paradigm

- Began with prescient thesis of Damour in 1978
- Fluctuations of a black hole horizon act like a viscous fluid
- Fluid viscosity is computed to be $\eta = 1/16\pi G$
- \blacksquare Dividing by the entropy density s=1/4G gives $\eta/s=1/4\pi$
- Always considers fluctuations at the black hole horizon $r=r_h$ itself; produces Damour-Navier Stokes equation

AdS/CFT Method

- Policastro, Son, Starinets hep-th/0205052 considered the hydrodynamics of $\mathcal{N}=4\,SU(N)$ SYM via AdS/CFT
- Again find $\eta/s = 1/4\pi$
- lacksquare Performed at AdS spatial infinity $r=\infty$
- Requires string theory, SUSY gauge theory, and AdS/CFT

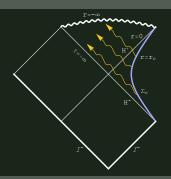
A 'Wilsonian' Approach

Fluid-gravity duality in the cutoff approach relates solutions of the incompressible Navier-Stokes equation

$$\partial^i v_i = 0, \quad \partial_\tau v_i - \bar{\eta} \partial^2 v_i + \partial_i P + v^j \partial_j v_i = 0$$

to solutions of the Einstein equation:

$$G_{\mu\nu} = 0$$



Fixing cutoff surface $r = r_c$, then perturbing:

- lacktriangle induced metric at $r=r_c$ is Ricci flat
- lacktriangle waves are infalling at $r=r_h$
- lacktriangle extrinsic curvature at $r=r_c$ becomes fluid stress tensor . . .
- in a hydrodynamic limit

Satisfying the Einstein Constraints

The Nonlinear Metric in the Hydrodynamic Limit

$$ds^{2} = -rd\tau^{2} + 2d\tau dr + dx_{i}dx^{i}$$

$$-2\left(1 - \frac{r}{r_{c}}\right)v_{i}dx^{i}d\tau - 2\frac{v_{i}}{r_{c}}dx^{i}dr$$

$$+\left(1 - \frac{r}{r_{c}}\right)\left[\left(v^{2} + 2P\right)d\tau^{2} + \frac{v_{i}v_{j}}{r_{c}}dx^{i}dx^{j}\right] + \left(\frac{v^{2}}{r_{c}} + \frac{2P}{r_{c}}\right)d\tau dr$$

$$-\frac{\left(r^{2} - r_{c}^{2}\right)}{r_{c}}\partial^{2}v_{i}dx^{i}d\tau + \dots \mathcal{O}(\epsilon^{3})$$

with
$$v_i \sim \mathcal{O}(\epsilon)$$
, $P \sim \mathcal{O}(\epsilon^2)$, $\partial_i \sim \mathcal{O}(\epsilon)$, $\partial_\tau \sim \mathcal{O}(\epsilon^2)$.

- Induced metric at $r = r_c$ cutoff is flat
- lacksquare constraint eqns at $\mathcal{O}(\epsilon^2)$ are $\partial^i v_i = 0$
- constraint eqns at $\mathcal{O}(\epsilon^3)$ are $\partial_{\tau}v_i r_c\partial^2 v_i + \partial_i P + v^j\partial_j v_i = 0$, Navier-Stokes with viscosity $\bar{\eta} = r_c$
- $\blacksquare G_{ra}, G_{ab}, G_{rr} = \mathcal{O}(\epsilon^4)$

Cutoff Approach

Highlights

- Does not require AdS, but is connectible to the AdS approach (Brattan, Camps, Loganayagam, Rangamani 1106.2577)
 Toy to test how double copy relates to AdS/CFT?
- Extendible to higher orders (Compere, McFadden, Skenderis, Taylor, 1103.3022; Pinzani-Fokeeva, Taylor 1401.5975)
- Hydrodynamic limit can be recast as near horizon limit
- Spacetime is algebraically special! Generic 4d fluid-duals are Petrov type II through $\mathcal{O}(\epsilon^{14})$ (Bredberg, Keeler, Lysov, Strominger 1101.2451
- More restricted fluids are more special!

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Petrov type II: C^{II}_{ABCD} \sim \alpha_{(A}\alpha_B\beta_C\gamma D)

Type D: C^{D}_{ABCD} \sim \alpha_{(A}\alpha_B\beta_C\beta_D) Type N: C^{N}_{ABCD} \sim \alpha_{(A}\alpha_B\alpha_C\alpha_D)
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Algebraic Speciality in Fluids

Can prove algebraic speciality by writing

$$C_{ABCD} = \Psi_0 \iota_A \iota_B \iota_C \iota_D - 4\Psi_1 o_{(A} \iota_B \iota_C \iota_D) + 6\Psi_2 o_{(A} o_B \iota_C \iota_D) - 4\Psi_3 o_{(A} o_B o_C \iota_D) + \Psi_4 o_A o_B o_C o_D$$

If only Ψ_2 is nonzero, then the spacetime is type D. If only Ψ_4 is nonzero, then the spacetime is type N. For general fluid-dual spacetimes, $\Psi_0, \Psi_1, \Psi_3 = 0 + \mathcal{O}(\epsilon^3)$,

$$\Psi_{2} = -i\epsilon^{2} (\partial_{x}v_{y} - \partial_{y}v_{x}) / 4r_{c} + \mathcal{O}(\epsilon^{3})$$

$$\Psi_{4} = -\epsilon^{2} (\partial_{x}v_{x} - \partial_{y}v_{y} + i(\partial_{x}v_{y} + \partial_{y}v_{x})) / 2r + \mathcal{O}(\epsilon^{3}).$$

Algebraically special fluid-dual spacetimes (τ -independent)

- Type D fluids have constant vorticity
- Type N fluids are potential flows

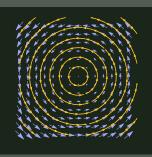
Type D fluids: Constant Vorticity

Only nonzero Ψ_I is

$$\Psi_2 = -i\epsilon^2 \omega / 2r_c + \mathcal{O}(\epsilon^3)$$

with natural background

$$ds_{(0)}^2 = -rd\tau^2 + 2drd\tau + dx^2 + dy^2$$



Single and Zeroth Copies

$$S = i\omega r_c e^{2i\theta}, \quad f_{AB} = e^{i\theta}\omega \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \rightarrow \begin{cases} F^{\tau r} = -\omega\cos\theta \\ F^{xy} = -\omega\sin\theta \end{cases}$$

- \blacksquare all other $F^{\mu\nu}$ components are zero
- lacksquare S is constant so trivially solves $\nabla^2_{(0)}S=0$

$$\nabla_{\nu}^{(0)} F^{\mu\nu} = 0, \quad \nabla_{[\mu}^{(0)} F_{\rho\sigma]} = 0$$

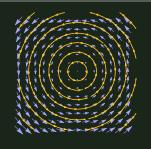
Type D fluid single copy: A giant solenoid

Choosing $\theta = 3\pi/2$ we have

$$v_x = -\omega y, \quad v_y = \omega x$$

$$F^{\tau r} = 0, \quad F^{xy} = \omega$$

$$E_{\mu} = 0, \quad B_{\mu} = \omega \delta_{\mu}^r$$



Type D Fluid Double Copy Summary

- Fluid is solution inside of slowly rotating cylinder with no-slip conditions at the wall
- Magnetic field $\vec{B} = \omega \hat{r}$ is uniform field inside a big solenoid with current proportional to ω
- lacktriangle zeroth copy field S is constant and thus plays a passive role
- Fluid only in hydro regime for $x,y \sim \epsilon^{-1}$; can fix by going to near-horizon expansion instead

Type N fluids: Potential flow: The Double Copy Story

The potential ϕ resides in the zeroth copy scalar S .

We have, using z = x + iy,

$$v_x = \partial_x \phi, \quad v_y = \partial_y \phi \text{ with } \phi = f(z) + \bar{f}(\bar{z})$$

The zeroth and single copy fields become

$$S = -rac{e^{2i heta}}{2\partial_{ar{z}}^2ar{f}(ar{z})}, \quad f_{AB} = rac{e^{i heta}}{\sqrt{r}} \left(egin{array}{cc} 1 & 1 \ 1 & 1 \end{array}
ight)$$

Type N Fluid Double Copy Summary

- $\blacksquare \ \nabla^2_{(0)} S = 0$ nontrivially; because $\phi = f(z) + \bar{f}(\bar{z})$
- lacksquare 'Background' single copy field is $ar{E}=-\hat{x},\ ar{B}=\hat{y}$
- Poynting vector of single copy is $\vec{S} = -\hat{r}$.
- Gauge field is single copy necessary to build up any fluid with a potential component.

Algebraically Special Fluid Double Copy Summary

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Type N Fluid Double Copy Summary

- lacksquare $\nabla^2_{(0)}S=0$ nontrivially; because $\phi=f(z)+ar{f}(ar{z})$
- lacksquare 'background' single copy field is still $ec{E}=-\hat{x},\ ec{B}=\hat{y}$
- Poynting vector of single copy is $\vec{S} = -\hat{r}$.
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Can we generalize to all Fluid Dual spacetimes?

Generic fluid-dual spacetime is type II, so its Weyl spinor is

$$C_{ABCD} = 6\Psi_2 o_{(A} \iota_B o_C \iota_{D)} + \Psi_4 o_A o_B o_C o_D = \frac{1}{S} f_{(AB}^{(1)} f_{CD)}^{(2)}$$
 for

$$f_{AB}^{(1)} = \beta \left(i\sqrt{6\Psi_2}o_{(A}\iota_{B)} + \sqrt{\Psi_4}o_{A}o_{B} \right)$$

$$f_{AB}^{(2)} = \frac{S}{\beta} \left(-i\sqrt{6\Psi_2}o_{(A}\iota_{B)} + \sqrt{\Psi_4}o_{A}o_{B} \right)$$

but we cannot pick a single β and S for which both gauge fields solve Maxwell and the scalar solves Klein-Gordon for all fluids.

Possible Solutions

- **Extension:** $C = \frac{1}{s_1} f^{(1,1)} f^{(1,2)} + \frac{1}{s_2} f^{(2,1)} f^{(2,2)}$
- Twistor method? cf. White 2012.02479; Chacon, Nagy, White 2103.16441

Details on General Type II Fluid-Dual spacetimes

The Weyl Spinor

In two dimensions for incompressible fluids: $v_i = \epsilon_{ij}\partial_j \chi$, where χ is the stream function and z = x + iy:

$$C_{ABCD} = 6\Psi_2 o_{(A} \iota_B o_C \iota_{D)} + \Psi_4 o_A o_B o_C o_D$$

$$\Psi_2 = i\partial_z \partial_{\bar{z}} \chi, \quad \Psi_4 = 2i\partial_{\bar{z}} \partial_{\bar{z}} \chi.$$

(In near horizon, λ expansion)

Factorizing
$$C_{ABCD}=lpha_{(A}lpha_{B}eta_{C}\gamma_{D)}$$
:
$$lpha_{A}=rac{P}{i\sqrt{6\Psi_{2}}}o_{A}$$

$$eta_{A}=i\sqrt{6\Psi_{2}}\iota_{A}+\sqrt{\Psi_{4}}o_{A}$$

$$\gamma_{A}=-iC\sqrt{6\Psi_{2}}\iota_{A}+C\sqrt{\Psi_{4}}o_{A}$$

Setting $S=-CP^2/6\Psi_2$, we can factorize either as $f_1=\alpha_{(A}\beta_{B)},\ f_2=\alpha_{(A}\gamma_{B)}$ or as $f_1=\alpha_{(A}\alpha_{B)},\ f_2=\beta_{(A}\gamma_{B)}$.

Details on General Type II Fluid-Dual spacetimes, Pt 2

The Weyl Spinor

$$C_{ABCD} = 6\Psi_2 o_{(A} \iota_B o_C \iota_D) + \Psi_4 o_A o_B o_C o_D$$
$$\Psi_2 = i\partial_z \partial_{\bar{z}} \chi, \quad \Psi_4 = 2i\partial_{\bar{z}} \partial_{\bar{z}} \chi.$$

Cases with $\Psi_2 \neq 0$ and $\Psi_4 \neq 0$:

- $\chi_{\text{Couette}} = A(z-\bar{z})^3 + B(z-\bar{z})^2 + C(z-\bar{z})$
- \blacksquare $\chi_{\text{Oseen-Lamb}} = \operatorname{Ei}\left[-z\bar{z}/4\eta t\right]/4\pi$
- Recover type D constant vorticity and type N potential flow
- For $\alpha\beta$, $\alpha\gamma$ factorization, Maxwell's give $\partial_{\bar{z}} \left[\Psi_4/\Psi_2 \right] = 0$. But (e.g. for Couette flow) can't always get $\nabla^2 S \neq 0$ (instead have $\nabla^2 (1/S) = 0$)
- For $\alpha\alpha$, $\beta\gamma$ factorization: Maxwell's: $\partial_{\bar{z}} \left[\Psi_4/\Psi_2 \right] = 0$, again Couette flow gives $\nabla^2(1/S) = 0$).
- Oseen-Lamb vortex doesn't work under either factorization

Future Directions

Future Questions

- Solving for Type II fluids (e.g. Couette or Oseen-Lamb):
 - Consider extension to sum of terms:

$$C = \frac{1}{s_1} f^{(1,1)} f^{(1,2)} + \frac{1}{s_2} f^{(2,1)} f^{(2,2)}$$

- Use Penrose transform/twistor story?
- Perturbative but nonlinear in Navier-Stokes
- higher orders in ϵ or λ ?
- relate to other fluid-gravity dualities
 - large D and near horizon physics
 - AdS/CFT: study fluid modes?
- Larger dimensions: 5d gravity= 3+1 d fluid forthcoming: (S. Chawla+C.K.) on general separable spacetimes as a double copy

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- Type D fluids have constant vorticity
- Type N fluids are potential flows

$$o_i - o_i \varphi, \quad o_{i} = o_i o_{\tau} \varphi \quad o_i \varphi o_i o_j \varphi.$$

Type N fluids: Planar Extensional Flow

The simplest Type N fluid has $\phi = \frac{\alpha}{2}(y^2 - x^2)$, so

$$v_x = \partial_x \phi = -\alpha x, v_y = \partial_y \phi = \alpha y$$

The zeroth and single copy fields become

$$S = \frac{e^{2i\theta}}{\alpha}, \quad f_{AB} = \frac{e^{i\theta}}{\sqrt{r}} \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix}$$

Again choosing $\theta = 3\pi/2$ the nonzero components of F become

$$F^{rx} = 1, F^{\tau x} = \frac{2}{r} \longrightarrow \vec{E} = -\hat{x}, \vec{B} = \hat{y}.$$

On the background $ds_{(0)}^2 = -rd\tau^2 + 2drd\tau + dx^2 + dy^2$ again both Klein-Gordon and Maxwell's are solved. Poynting vector is

$$\vec{S} = -\hat{r}$$
.

Gauge field is single copy necessary to build up any fluid with a potential component.

What if we consider a different potential ϕ ?

Type N fluids: Potential flow: The Double Copy Story

We already studied extensional flow: $\phi = \frac{\alpha}{2}(y^2 - x^2)$, so

$$v_x = \partial_x \phi = -\alpha x, v_y = \partial_y \phi = \alpha y$$

$$\vec{E} = -\hat{x}, \vec{B} = \hat{y}$$

But there are many other potential flow fluids!

	Potential ϕ	v_x	$ v_y $
Ext. flow	$-\frac{\alpha}{2}(x^2-y^2)$	$-\alpha x$	αy
Source/Sink	$\ln(x^2 + y^2)$	$2x/(x^2+y^2)$	$2y/(x^2+y^2)$
Dipole	$x/(x^2+y^2)$	$(y^2 - x^2)/(x^2 + y^2)^2$	$-2xy/(x^2+y^2)^2$
Line Vortex	$\arctan(y/x)$	$-y/(x^2+y^2)$	$x/(x^2+y^2)$

If $F_{\mu\nu}$ is just a 'support' single copy, then what distinguishes these fluids from each other? S!