



INSTITUT
POLYTECHNIQUE
DE PARIS



ASYMPTOTIC SYMMETRIES, DOUBLE COPY & CONFORMAL DRESSINGS IN

◆ CELESTIAL DIAMONDS ◆

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\mathcal{S} -matrix *and holography*

→ basic observable of quantum gravity in asymptotically flat spacetimes.

Holographic flavor: use on-shell data to learn about bulk physics.

$$\text{boost} \langle out | \mathcal{S} | in \rangle_{\text{boost}} = \langle \mathcal{O}_{\Delta_1, J_1}^\pm(w_1, \bar{w}_1) \dots \mathcal{O}_{\Delta_n, J_n}^\pm(w_n, \bar{w}_n) \rangle_{\text{celestial CFT}}$$

Celestial Holography programme:

attempts to apply holographic principle to asymptotically flat spacetimes.

Observation: soft theorems in QFT = Ward ID of asymptotic symmetries.

Symmetry generators look like currents living on the celestial sphere.

\mathcal{S} -matrix as celestial amplitude

→ basic observable of quantum gravity in asymptotically flat spacetimes.

Holographic flavor: use on-shell data to learn about bulk physics.

$$\text{boost} \langle out | \mathcal{S} | in \rangle_{\text{boost}} = \langle \mathcal{O}_{\Delta_1, J_1}^\pm(w_1, \bar{w}_1) \dots \mathcal{O}_{\Delta_n, J_n}^\pm(w_n, \bar{w}_n) \rangle_{\text{celestial CFT}}$$

CFT-inspired approach to amplitudes:

Celestial CFT bootstrap machinery to compute and constrain scattering?

Collinear limits ⇒ OPE data for celestial CFT.

Conformally soft limits: constraints from symmetries & conformal dressings.

Anti-Wilsonian paradigm: celestial amplitudes probe all energy scales.

Plan of the talk

BASED ON

2012.15694 - SABRINA PASTERSKI

2105.03516 & 2105.09792 - SABRINA PASTERSKI & EMILIO TREVISANI

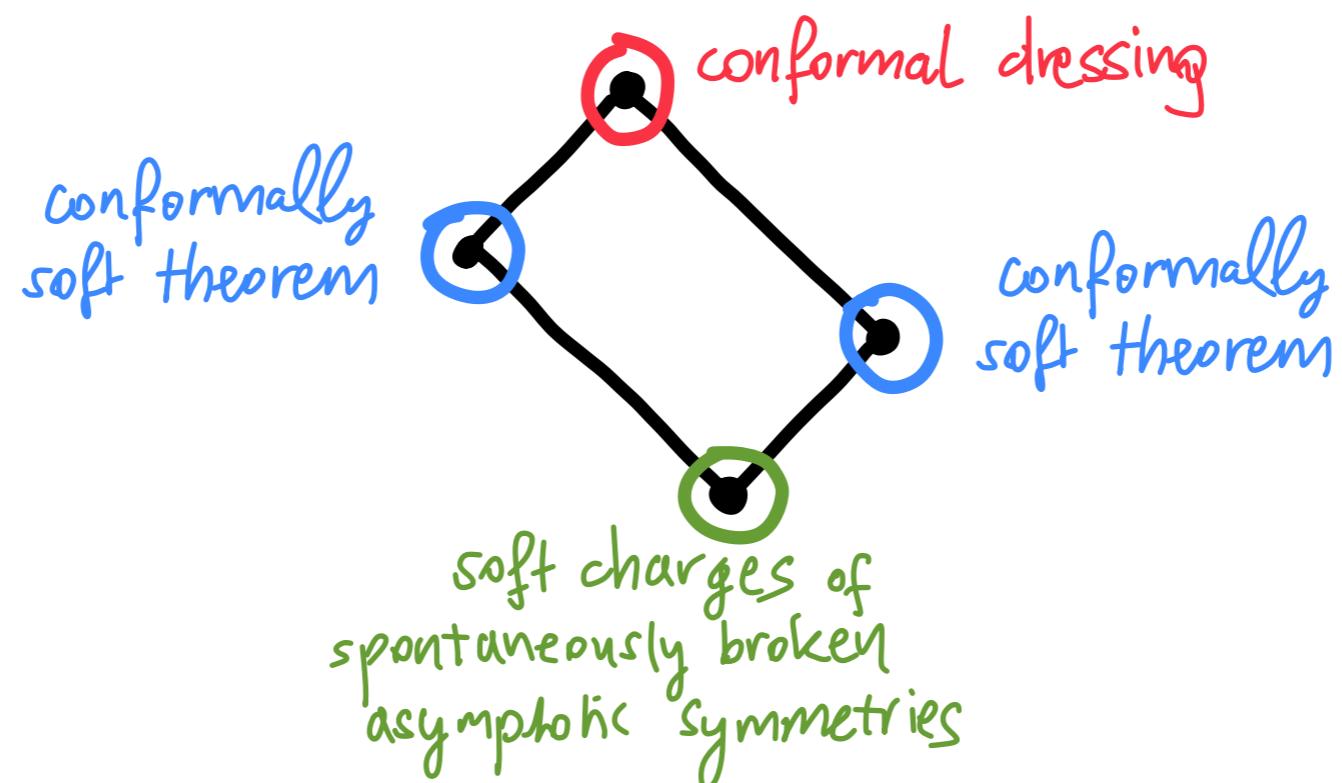
ALSO

2007.15027 - EDUARDO CASALI

2108.XXXXX - YORGO PANOS & SABRINA PASTERSKI

CELESTIAL DIAMONDS:

Natural structure in celestial CFT that unifies the discussion of soft sector:



Appearance of *celestial double copy* and *shockwaves* in the diamond.

Symmetries

$$\mathcal{L}\eta_{\mu\nu} = \partial_\mu\xi_\nu + \partial_\nu\xi_\mu = 0$$

Poincare generators in Bondi coordinates (u, r, z, \bar{z}) :

$$\begin{aligned} \xi = & (1 + \frac{u}{2r})Y^z\partial_z - \frac{u}{2r}D^{\bar{z}}D_zY^z\partial_{\bar{z}} - \frac{1}{2}(u+r)D_zY^z\partial_r + \frac{u}{2}D_zY^z\partial_u + c.c. \\ & + f\partial_u - \frac{1}{r}(D^z f\partial_z + D^{\bar{z}} f\partial_{\bar{z}}) + D^z D_z f\partial_r \end{aligned}$$

Translations:

$$f_0 = 1, f_1 = \frac{z + \bar{z}}{1 + z\bar{z}}, f_2 = \frac{i(\bar{z} - z)}{1 + z\bar{z}}, f_3 = \frac{1 - z\bar{z}}{1 + z\bar{z}}$$

$$Y_{12}^z = iz, Y_{13}^z = -\frac{1}{2}(1 + z^2), Y_{23}^z = -\frac{i}{2}(1 - z^2)$$

$$Y_{03}^z = z, Y_{02}^z = -\frac{i}{2}(1 + z^2), Y_{01}^z = -\frac{1}{2}(1 - z^2)$$

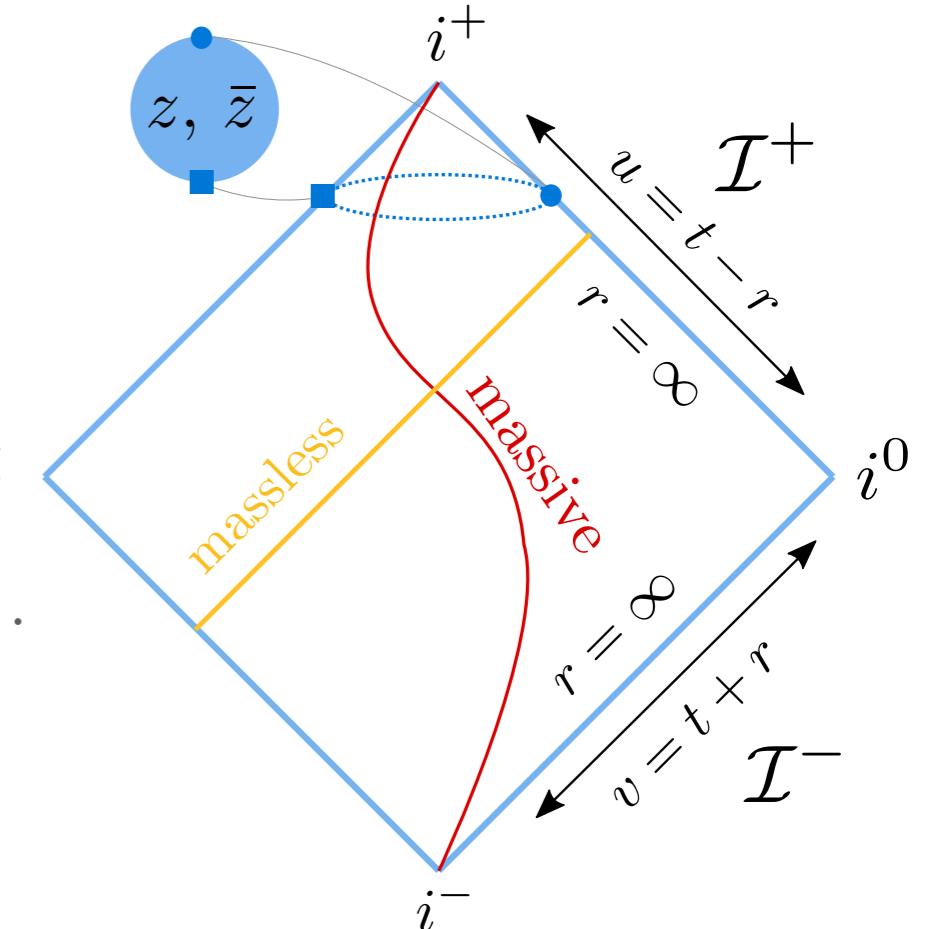
Asymptotic symmetry group of Einstein gravity:

$$\eta_{flat} \Rightarrow \eta_{flat} + \dots \quad \text{e.g. allow radiation}$$

[Bondi,van der Burg,Metzner,Sachs'62]
[Barnich,Troessaert'11] ...

BMS supertranslations and superrotations: $f(z, \bar{z})$ & $Y^z(z, \bar{z})$

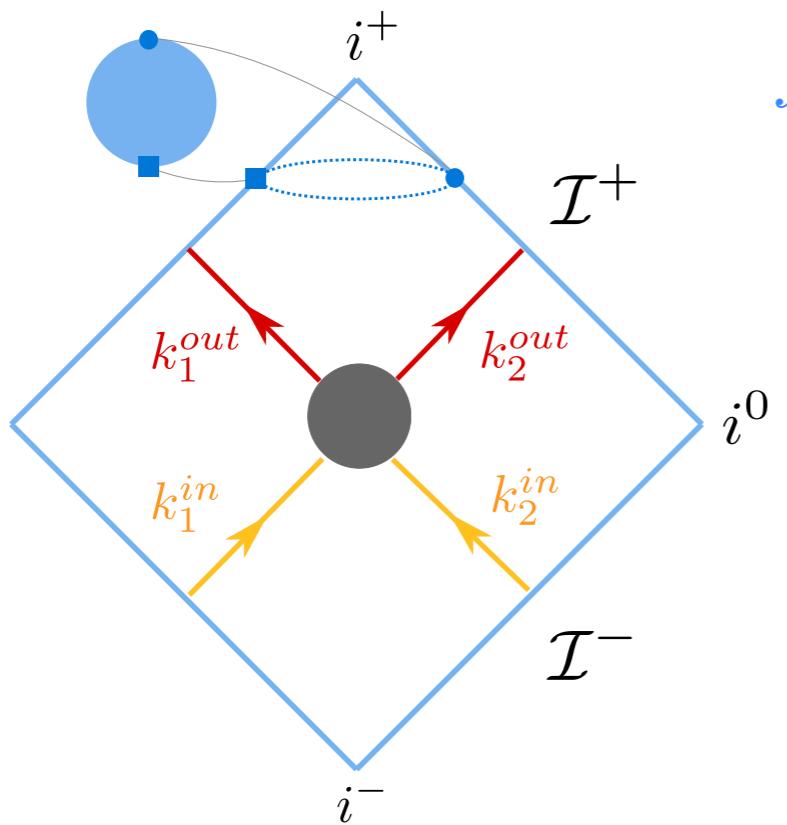
Asymptotic symmetries manifest in scattering amplitudes as soft theorems.



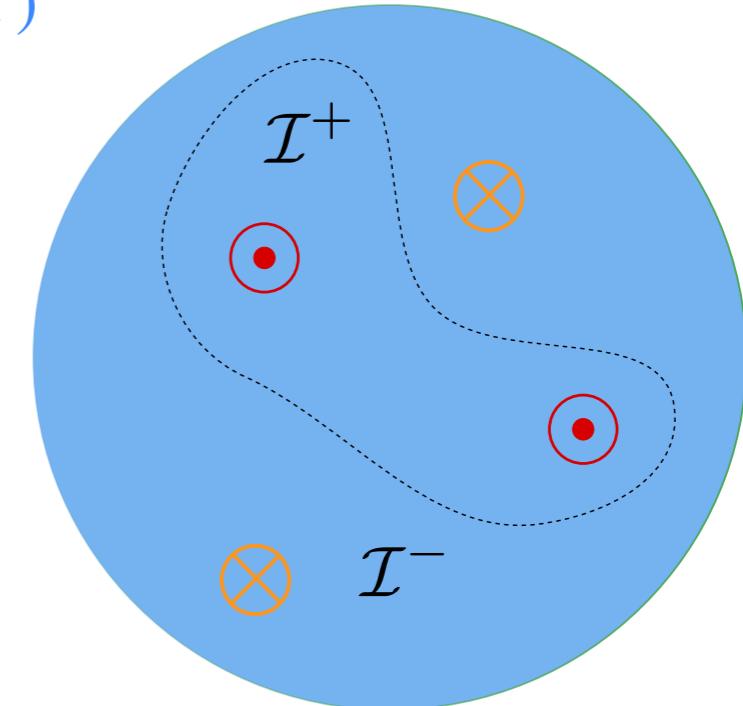
4D \mathcal{S} -matrix \Rightarrow 2D correlator

[de Boer,Solodhukin'03] [Cheung,de la Fuente,Sundrum'16][Pasterski,Shao,Strominger'16]

For massless scattering the map is a Mellin transform:



$$\mathcal{M}(.) = \int_0^{\infty} \frac{d\omega}{\omega} \omega^{\Delta}(.)$$



Lorentz symmetry

$$\langle \prod_{i=1}^n \mathcal{O}_{\Delta_i, J_i}(w_i, \bar{w}_i) \rangle$$

spin point on S^2

conformal symmetry

Conformal Primary Operators

Given a 4D operator $O^s(X^\mu)$ of spin- s in the Heisenberg picture, we can define a 2D operator $\mathcal{O}_{\Delta,J}^{s,\pm}(w, \bar{w})$ in the celestial CFT via a suitable inner product with *conformal primary wavefunctions* $\Phi_{\Delta,J}^s(X_\pm^\mu; w, \bar{w})$.

2D celestial CFT operator

$$\curvearrowleft \mathcal{O}_{\Delta,J}^{s,\pm}(w, \bar{w}) \equiv i(\underbrace{O^s(X^\mu)}, \underbrace{\Phi_{\Delta^*, -J}^s(X_\mp^\mu; w, \bar{w})})_\Sigma$$

[Donnay,Pasterski,AP'20]

[Pasterski,AP,Trevisani'21]

4D bulk operator
expanded into modes

4D wavefunction

inner product = integral over
Cauchy slice Σ in the 4D bulk

The \pm label indicates *in* vs *out* states selected by prescription for analytically continuing the wavefunctions as $X_\pm^\mu = X^\mu \pm i\varepsilon\{-1,0,0,0\}$.

Conformal Primary Wavefunctions

$$\Phi_{\Delta,J}^{s=|J|}(X^\mu; w, \bar{w}) \simeq \int_0^\infty \frac{d\omega}{\omega} \omega^\Delta \epsilon_{\mu_1 \dots \mu_s} e^{\pm i\omega q \cdot X}$$

4D spin- s field under Lorentz transformations

2D conformal primary with conformal dimension Δ and spin J



Lorentz transformation
bulk point $X^\mu \mapsto \Lambda^\mu_\nu X^\nu$

boundary point

conformal transformation
 $w \mapsto \frac{aw + b}{cw + d}$ $\bar{w} \mapsto \frac{\bar{a}\bar{w} + \bar{b}}{\bar{c}\bar{w} + \bar{d}}$
 $ad - bc = 1 = \bar{a}\bar{d} - \bar{b}\bar{c}$

$$\Phi_{\Delta,J}^s\left(\Lambda^\mu_\nu X^\nu; \frac{aw + b}{cw + d}, \frac{\bar{a}\bar{w} + \bar{b}}{\bar{c}\bar{w} + \bar{d}}\right) = (cw + d)^{\Delta+J}(\bar{c}\bar{w} + \bar{d})^{\Delta-J} D_s(\Lambda) \Phi_{\Delta,J}^s(X^\mu; w, \bar{w})$$

3+1D spin- s representation of the Lorentz algebra

[Pasterski,Shao'17]

Radiative: $J = \pm s$ & solve the linearized eom for massless spin- s particles

Generalized: $|J| \leq s$ & allow sources and distributions

[Pasterski,AP'20]

Celestial amplitudes

[Pasterski,Shao,Strominger'16+'17]

\mathcal{S} -matrix elements constructed as

$$\widetilde{\mathcal{A}}(\Delta_i, z_i, \bar{z}_i) = \prod_{i=1}^n \int_0^\infty \frac{d\omega_i}{\omega_i} \omega_i^{\Delta_i} \mathcal{A}(\omega_i, z_i, \bar{z}_i)$$

transform by construction as correlators of (quasi)-primaries

$$\widetilde{\mathcal{A}}(\Delta_i, \frac{az_i + b}{cz_i + d}, \frac{\bar{a}\bar{z}_i + \bar{b}}{\bar{c}\bar{z}_i + \bar{d}}) = \prod_{j=1}^n (cz_j + d)^{\Delta_j + J_j} (\bar{c}\bar{z}_j + \bar{d})^{\Delta_j - J_j} \widetilde{\mathcal{A}}(\Delta_i, z_i, \bar{z}_i).$$

J_j = j-th helicity



since the external particles are in boost eigenstates.

Spectrum vs Symmetries

The transformation can be inverted easily when Δ is on the principal continuous series of the Lorentz group $\Delta = 1 + i\mathbb{R}$ which captures finite energy radiation:

$$\mathcal{A}(\omega_i, z_i, \bar{z}_i) = \prod_{i=1}^n \int_{1-i\infty}^{1+i\infty} \frac{d\Delta_i}{2\pi i} \omega_i^{-\Delta_i} \widetilde{\mathcal{A}}(\Delta_i, z_i, \bar{z}_i).$$

Analytic continuation to $\Delta \in \mathbb{C}$ of great interest to celestial CFT.

More on this later!

Translations shift the conformal dimension:

$$\mathcal{P}^\mu = q^\mu e^{\partial_\Delta} \Leftrightarrow \Delta \mapsto \Delta + 1$$

[Donnay,AP,Strominger'18]
[Stieberger,Taylor'18]

What happens to properties of amplitudes that exploit manifest translation symmetry?

Gravity = Gauge Theory²

[Bern,Carrasco,Johansson'10]

manifest translation symmetry

gluon amplitude

$$\mathcal{A}_n^{YM} = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{c_\gamma n_\gamma}{\prod_\gamma} \text{ propagator}$$

double
copy

$$c \mapsto n$$

$$\mathcal{A}_n^G = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{n_\gamma^2}{\prod_\gamma}$$

graviton amplitude

Celestial Gravity = Gauge Theory²

[Bern,Carrasco,Johansson'10]

manifest translation symmetry

$$\mathcal{M}(\cdot) = \int_0^\infty \frac{d\omega}{\omega} \omega^\Delta(\cdot)$$

[Casali,AP'20]

manifest conformal symmetry

gluon amplitude

$$\mathcal{A}_n^{YM} = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{c_\gamma n_\gamma}{\prod_\gamma} \text{color propagator}$$

double copy

$$c \mapsto n$$

$$\mathcal{A}_n^G = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{n_\gamma^2}{\prod_\gamma}$$

graviton amplitude

celestial gluon amplitude

$$\widetilde{\mathcal{A}}_n^{YM} = \mathcal{M}(\mathcal{A}_n^{YM})$$

$$\Phi_{\Delta,J}$$

?

$$\widetilde{\mathcal{A}}_n^G = \mathcal{M}(\mathcal{A}_n^G)$$

celestial graviton amplitude

Celestial Gravity = Gauge Theory²

[Bern,Carrasco,Johansson'10]

manifest translation symmetry

$$\mathcal{M}(.) = \int_0^\infty \frac{d\omega}{\omega} \omega^\Delta(.)$$

[Casali,AP'20]

manifest conformal symmetry

gluon amplitude

$$\mathcal{A}_n^{YM} = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{c_\gamma n_\gamma}{\prod_\gamma} \text{color propagator kinematic}$$

double copy

$$c \mapsto n$$

$$\mathcal{A}_n^G = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{n_\gamma^2}{\prod_\gamma}$$

graviton amplitude

kinematic functions	\mapsto	kinematic operators
$n(\omega)$		$\mathcal{N}(e^{\partial_\Delta})$

details: QCD meets gravity 2020

celestial gluon amplitude

$$\widetilde{\mathcal{A}}_n^{YM} = \sum_{\gamma \in \Gamma} c_\gamma \mathcal{N}_\gamma \widetilde{\mathcal{A}}_n^{\text{scalar}}$$

$$\Phi_{\Delta,J}$$

$$c \mapsto \mathcal{N}$$

celestial double copy

$$\widetilde{\mathcal{A}}_n^G = \sum_{\gamma \in \Gamma} \mathcal{N}_\gamma^2 \widetilde{\mathcal{A}}_n^{\text{scalar}}$$

celestial graviton amplitude

Celestial Gravity = Gauge Theory²

[Bern,Carrasco,Johansson'10]

manifest translation symmetry

$$\mathcal{M}(.) = \int_0^\infty \frac{d\omega}{\omega} \omega^\Delta(.)$$

[Casali,AP'20]

manifest conformal symmetry

gluon amplitude

$$\mathcal{A}_n^{YM} = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{c_\gamma n_\gamma}{\prod_\gamma} \text{color propagator kinematic}$$

double copy

$$c \mapsto n$$

$$\mathcal{A}_n^G = \delta^{(4)}\left(\sum_{i=1}^n k_i^\mu\right) \sum_{\gamma \in \Gamma} \frac{n_\gamma^2}{\prod_\gamma}$$

graviton amplitude

kinematic functions $n(\omega)$	\mapsto	kinematic operators $\mathcal{N}(e^{\partial_\Delta})$
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details: QCD meets gravity 2020

celestial gluon amplitude

$$\widetilde{\mathcal{A}}_n^{YM} = \sum_{\gamma \in \Gamma} c_\gamma \mathcal{N}_\gamma \widetilde{\mathcal{A}}_n^{\text{scalar}}$$

celestial double copy

$$\widetilde{\mathcal{A}}_n^G = \sum_{\gamma \in \Gamma} \mathcal{N}_\gamma^2 \widetilde{\mathcal{A}}_n^{\text{scalar}}$$

celestial graviton amplitude

Double copy is **fundamental property** of gauge and gravity amplitudes.

Operator-valued celestial double \Rightarrow *curved* spacetimes?

Building blocks for celestial CFT

[Pasterski,Shao'17]

[Pasterski,AP'20]

Embedding S^2 into $\mathbb{R}^{1,3}$ lightcone: $q^\mu = (1 + w\bar{w}, w + \bar{w}, -i(w - \bar{w}), 1 - w\bar{w})$

$$\sqrt{2}\epsilon_+^\mu = \partial_w q^\mu, \sqrt{2}\epsilon_-^\mu = \partial_{\bar{w}} q^\mu$$

Mellin transform of plane waves :

$$\mathcal{M}(e^{\pm ik \cdot X}) = \int_0^\infty \frac{d\omega}{\omega} \omega^\Delta e^{\pm i\omega q \cdot X_\pm} = \frac{(\mp i)^\Delta \Gamma(\Delta)}{(-q \cdot X_\pm)^\Delta}$$

&

null tetrad

$$\{l, \underbrace{n}_{J=0}, \underbrace{m}_{J=\pm 1}, \bar{m}\}$$

spin frame

$$\{o, \underbrace{\bar{o}}_{J=\pm \frac{1}{2}}, \underbrace{l}_{J=\pm \frac{1}{2}}, \bar{l}\}$$

Spin $s = 0$:

$$\varphi^\Delta = \frac{1}{(-q \cdot X)^\Delta}$$

Spin $s = \frac{1}{2}, 1, \frac{3}{2}, 2$:

$$\psi_{\Delta, J=+\frac{1}{2}} = o\varphi^\Delta$$

$$A_{\Delta, J=+1} = m\varphi^\Delta$$

$$\chi_{\Delta, J=+\frac{3}{2}} = om\varphi^\Delta$$

$$h_{\Delta, J=+2} = mm\varphi^\Delta$$

$$o_a = \sqrt{q \cdot X} |q]_a$$

$$o_a \mapsto (cw + d)^{\frac{1}{2}} (\bar{c}\bar{w} + \bar{d})^{-\frac{1}{2}} (Mo)_a$$

$$m^\mu = \epsilon_+^\mu + (\epsilon_+ \cdot X) l^\mu$$

spacetime dependent
polarization vector

$$m^\mu \mapsto \frac{cw + d}{\bar{c}\bar{w} + \bar{d}} \Lambda_\nu^\mu m^\nu$$

Spin- s via double copy and supersymmetry

Supersymmetry: relates primaries by $\frac{1}{2}$ integer spin steps

$$\mathcal{Q}_a = \partial_\theta |q]_a e^{\partial_\Delta/2} \quad \bar{\mathcal{Q}}_{\dot{a}} = \theta \langle q|_{\dot{a}} e^{\partial_\Delta/2} \quad \{\mathcal{Q}_a, \bar{\mathcal{Q}}_{\dot{a}}\} = -\sigma_{a\dot{a}}^\mu \mathcal{P}_\mu \quad \mathcal{P}_\mu = q_\mu e^{\partial_\Delta}$$

[Fotopoulos,Stieberger,Taylor,Zhu'20]

Kerr-Schild double copy:

$$g_{\mu\nu} = \eta_{\mu\nu} + \underbrace{m_\mu m_\nu \varphi^\Delta}_{h_{\Delta,J=\pm 2;\mu\nu}}$$

[Pasterski,AP'20]

Kerr-Schild vector:

$$\begin{aligned} g^{\mu\nu} m_\mu m_\nu &= 0 && \text{null} \\ \eta^{\mu\nu} m_\mu m_\nu &= 0 \\ m^\mu \nabla_\mu m_\nu &\propto m_\nu && \text{geodesic} \\ m^\mu \partial_\mu m_\nu &\propto m_\nu \end{aligned}$$

[Monteiro,O'Connell,White'14]

$h_{\Delta,J=\pm 2;\mu\nu}$ is *exact* sol to Einstein: Petrov type N

$$R_{\mu\nu}^{\Delta,J} = 0$$

$$\Psi_4 = -\frac{1}{2}\Delta(\Delta-1)\varphi^\Delta$$

$\Delta = 0,1$: asymptotic symmetries!

Finite conformal primary backgrounds

[Pasterski,AP'20]

- Aichelburg-Sexl shockwave or ultraboosted Schwarzschild [Aichelburg,Sexl'71]

$$g_{\mu\nu} = \eta_{\mu\nu} - 4G_N \alpha q_\mu q_\nu \log(X^2) \delta(q \cdot X)$$

$$h_{\Delta=-1, J=0; \mu\nu}^{gen}$$

$$E = \alpha q^0$$

generalized conformal primary metric

- Dray-'t Hooft planar shell [Dray,t'Hooft'86]

$$g_{\mu\nu} = \eta_{\mu\nu} - 4G_N \alpha q_\mu q_\nu \frac{(X^2)}{D-2} \delta(q \cdot X)$$

$$\rho = \alpha q^0$$

- Beam-like gravitational wave [Ferrari,Pendenza,Veneziano'88]

$$g_{\mu\nu} = \eta_{\mu\nu} - 4G_N \alpha q_\mu q_\nu f(X^2) \delta(q \cdot X)$$

$$E = \alpha q^0 \sqrt{X^2} f'(X^2)$$

- Kerr gyraton or ultraboosted Kerr see [Cristofoli'20] [Arkani-Hamed,Huang,O'Connell'20]...

$$g_{\mu\nu} = \eta_{\mu\nu} - 4G_N \alpha q_\mu q_\nu \log(|X^2 - a^2|) \delta(q \cdot X)$$

$$a^\mu = s^\mu/m$$

→ *Amplitudes methods to explore non-perturbative bulk physics in celestial CFT.*

Soft limits as celestial currents

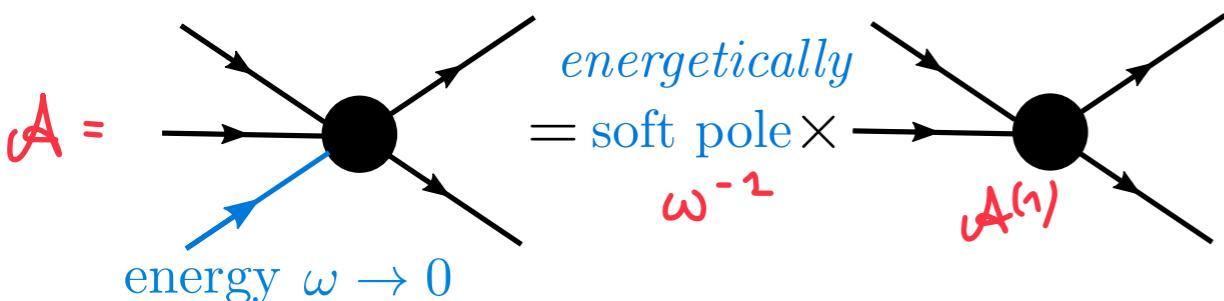
Ward ID of asymptotic symmetries = soft theorems in QFT.

[Strominger'13] ...

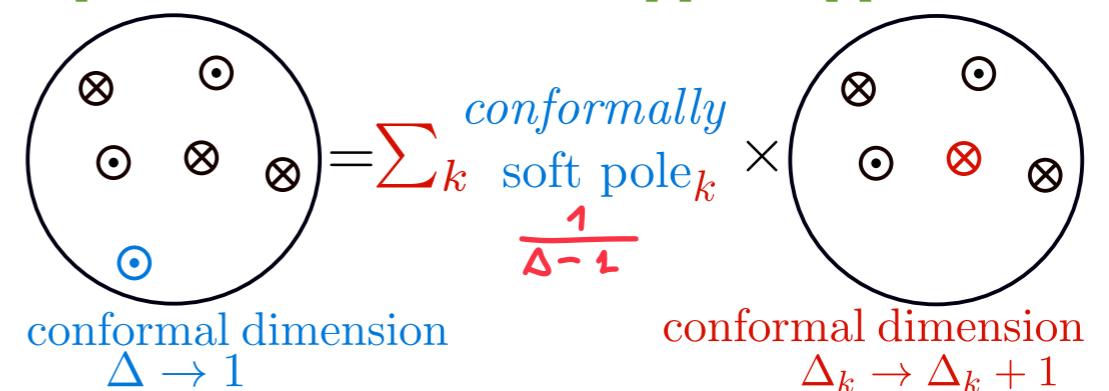
$$\lim_{\omega \rightarrow 0} \mathcal{A} = \omega^{-1} \mathcal{A}^{(1)} + \omega^0 \mathcal{A}^{(0)} + \dots$$

Soft limits at different orders in $\omega \rightarrow 0 \Leftrightarrow$ "conformally soft" Δ poles.

e.g. leading soft graviton: [Weinberg'65]



[Adamo,Mason,Sharma'19] [AP'19] [Guevara'19]



$$\mathcal{P}_w \mathcal{O}_\omega(z, \bar{z}) \sim \frac{\omega}{w-z} \mathcal{O}_\omega(z, \bar{z})$$

supertranslation current

$$\mathcal{P}_w \mathcal{O}_\Delta(z, \bar{z}) \sim \frac{1}{w-z} \mathcal{O}_{\Delta+1}(z, \bar{z})$$

[Donnay,AP,Strominger'18]

Conformally soft primaries

$$\mathcal{O}_{\Delta,J} \equiv i(\hat{O}, \Phi_{\Delta^*, -J}^G)_{\Sigma}$$

symmetry generator pure gauge

[Donnay,AP,Strominger'18]
 [Donnay,Pasterski,AP'20]
 [Pasterski,AP'20]
 [Pano,Pasterski,AP-to appear]

	$A_{\Delta,J=\pm 1}$	$\chi_{\Delta,J=+\frac{3}{2}}$	$h_{\Delta,J=\pm 2}$	
Δ symmetry	1 large U(1)	$\frac{1}{2}$ large SUSY	1 supertranslation	0 shadow superrotation

Goldstones of spontaneously broken asymptotic symmetries for $1 \leq s \leq 2$

$\text{Diff}(S^2)$

Kerr-Schild double copy: $\text{BMS} = (\text{large U}(1))^2$

[Huang,Kol,O'Connell'19]

	$\psi_{\Delta,J=+\frac{1}{2}}$	$A_{\Delta,J=\pm 1}$	$\chi_{\Delta,J=+\frac{3}{2}}$	$h_{\Delta,J=\pm 2}$
Δ	$\frac{1}{2}$	0	$-\frac{1}{2}$	-1

not pure gauge

[Fotopoulos,Taylor,
Stieberger,Zhu'20]
[Adamo,Mason,Sharma'19]
[Guevara,19]

Soft theorems without conformal Goldstones for $1 \leq s \leq 2$

Conformally soft shadow primaries

$$\widetilde{\Phi}_{\Delta,J} \equiv \widetilde{\Phi_{2-\Delta,-J}}$$

	$\tilde{A}_{\Delta,J=\pm 1}$	$\tilde{\chi}_{\Delta,J=-\frac{3}{2}}$	$\tilde{h}_{\Delta,J=\pm 2}$
Δ	1	$\frac{3}{2}$	1
symmetry	large U(1)	large SUSY	supertranslation

Goldstones of spontaneously broken asymptotic symmetries for $1 \leq s \leq 2$

[Donnay,AP,Strominger'18]

[Donnay,Pasterski,AP'20]

[Pasterski,AP'20]

[Pano,Pasterski,AP-to appear]

SUSY current

[Fotopoulos,Taylor,
Stieberger,Zhu'20]

stress tensor for celestial CFT

[Kapac,Mitra,Raclariu,Strominger'16]

	$\tilde{\psi}_{\Delta,J=-\frac{1}{2}}$	$\tilde{A}_{\Delta,J=\pm 1}$	$\tilde{\chi}_{\Delta,J=-\frac{3}{2}}$	$\tilde{h}_{\Delta,J=\pm 2}$
Δ	$\frac{3}{2}$	2	$\frac{5}{2}$	3

Soft theorems without conformal Goldstones for $1 \leq s \leq 2$



not pure gauge

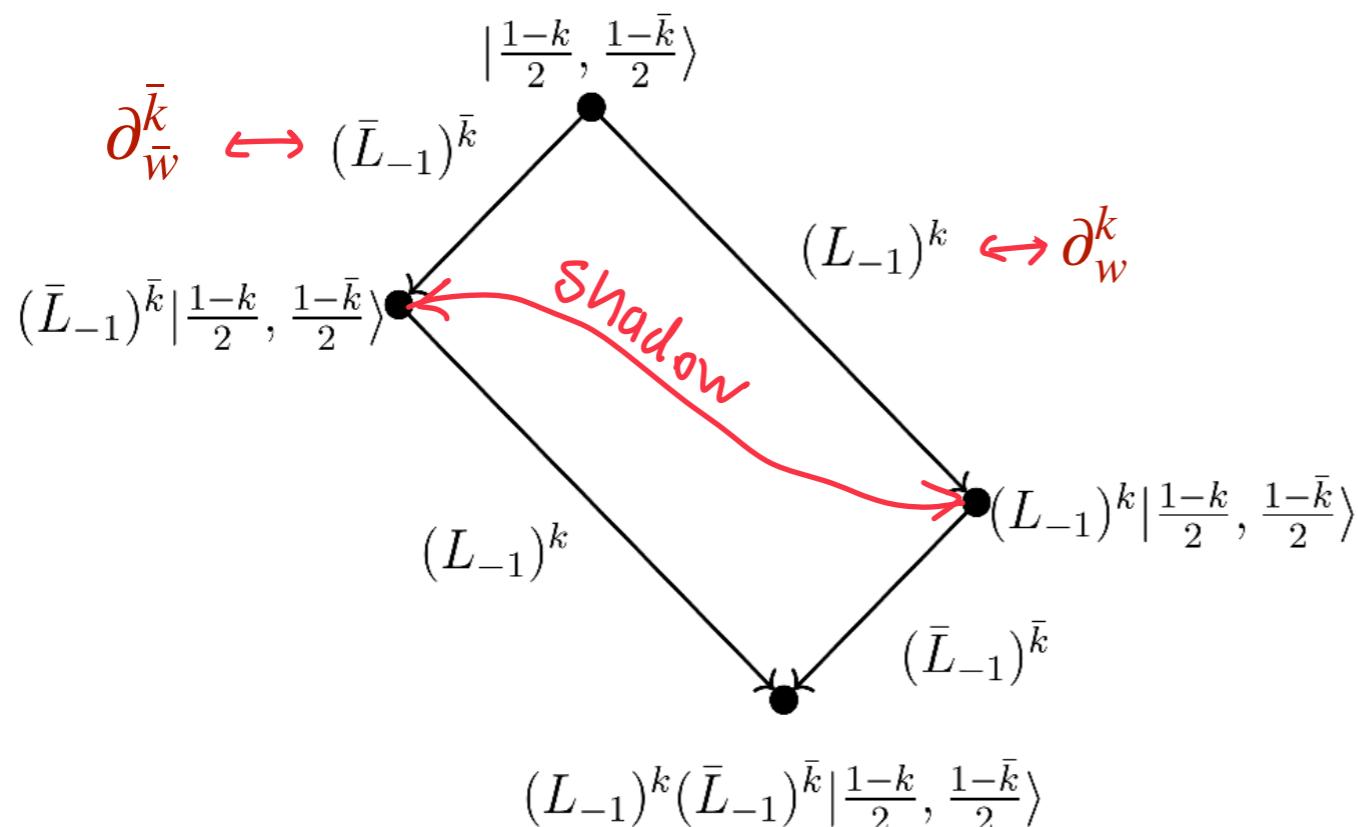
A unified treatment of soft modes

Asymptotic symmetry generators, their associated soft charges as well as conformal dressings for celestial amplitudes are all captured by the structure of the conformal multiplets of global $SL(2, \mathbb{C})$.

Primary $|h, \bar{h}\rangle$ annihilated by action of L_1, \bar{L}_1 .

Primary descendants from action of L_{-1}, \bar{L}_{-1}

$$\begin{aligned} & (L_{-1})^k |h, \bar{h}\rangle \text{ when } h = \frac{1-k}{2} \text{ with } k \in \mathbb{Z}_> \\ & (\bar{L}_{-1})^{\bar{k}} |h, \bar{h}\rangle \text{ when } \bar{h} = \frac{1-\bar{k}}{2} \text{ with } k \in \mathbb{Z}_> \end{aligned}$$



"celestial diamonds"

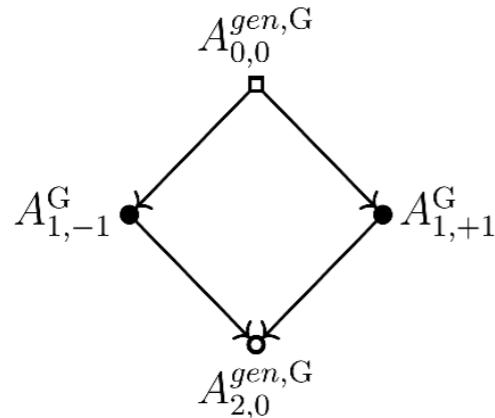
[Pasterski,AP,Trevisani'21]

Celestial diamonds

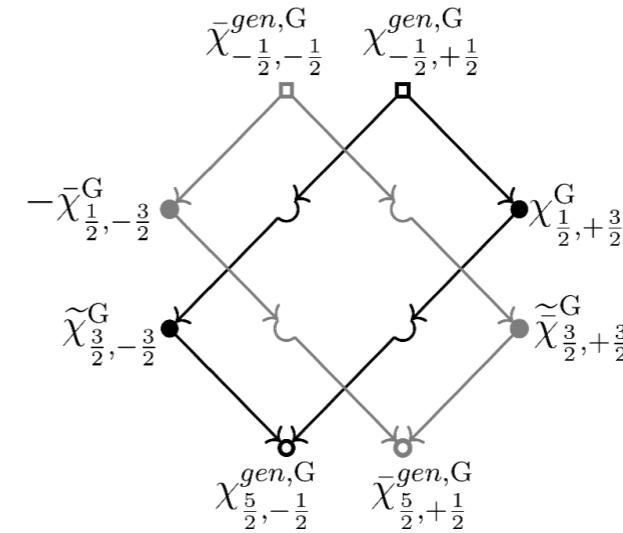
$$\Phi^{gen,s} \mapsto \Phi^{s=|J|} \text{ or } \Phi^{s=|J|} \mapsto \Phi^{gen,s}$$

[Pasterski,AP,Trevisani'21]

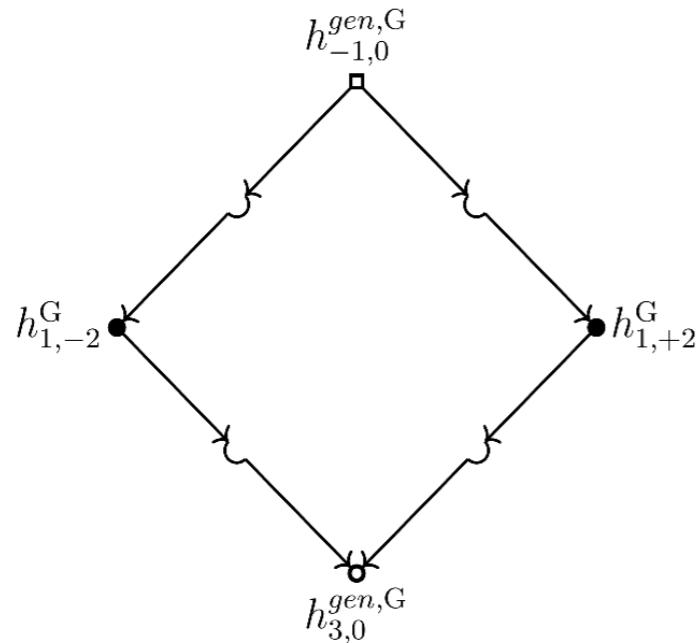
Leading photon:



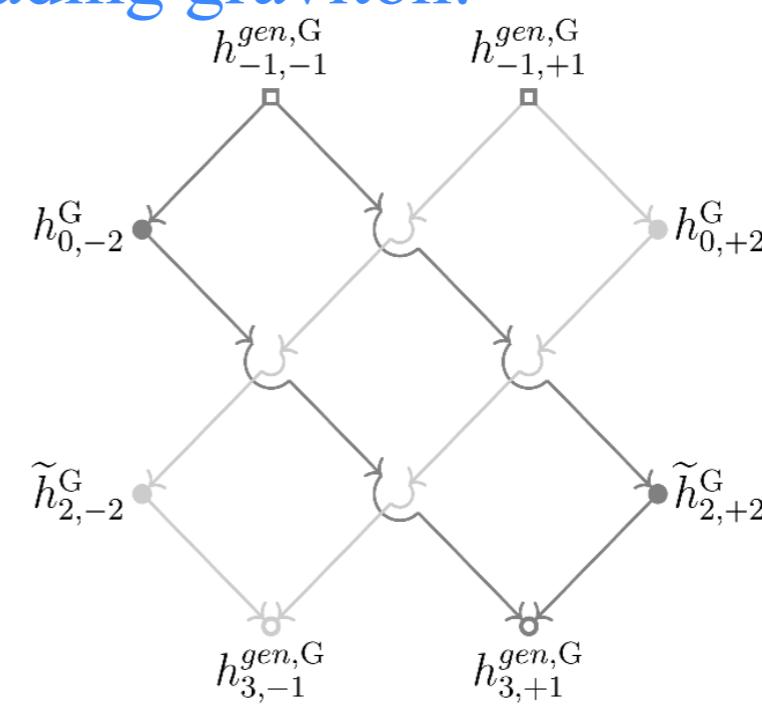
Leading gravitino:



Leading graviton:



Subleading graviton:



⇒ Capture most leading (conformally) soft theorems.

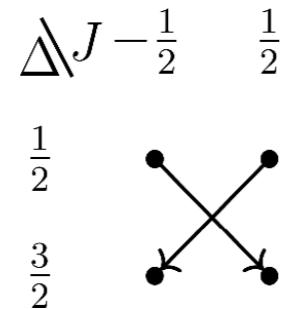
Shockwaves are at the top corner of the associated memory diamonds.

Degenerate diamonds

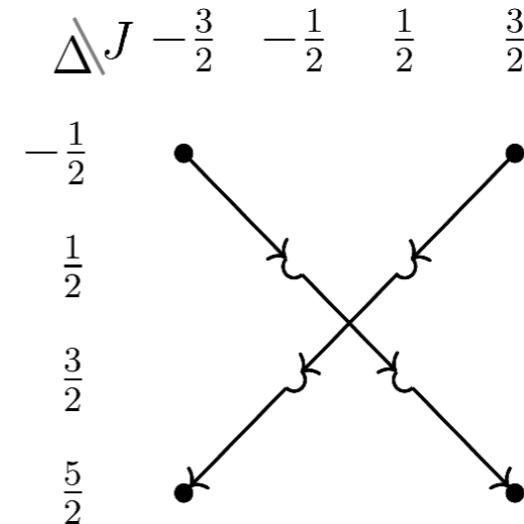
$$\Phi^{s=|J|} \mapsto \widetilde{\Phi}^{s=|J|}$$

[Pasterski,AP,Trevisani'21]

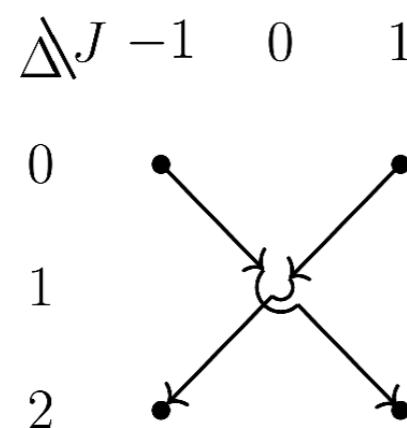
Leading photino:



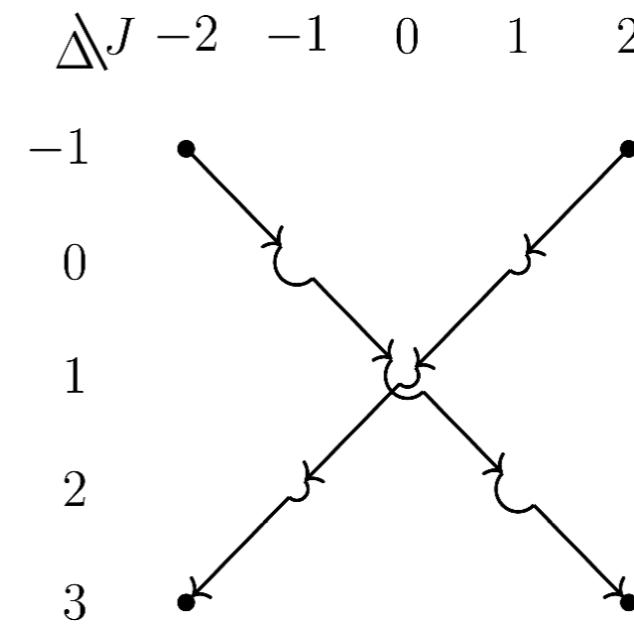
Subleading gravitino:



Subleading photon:



Subsubleading graviton:



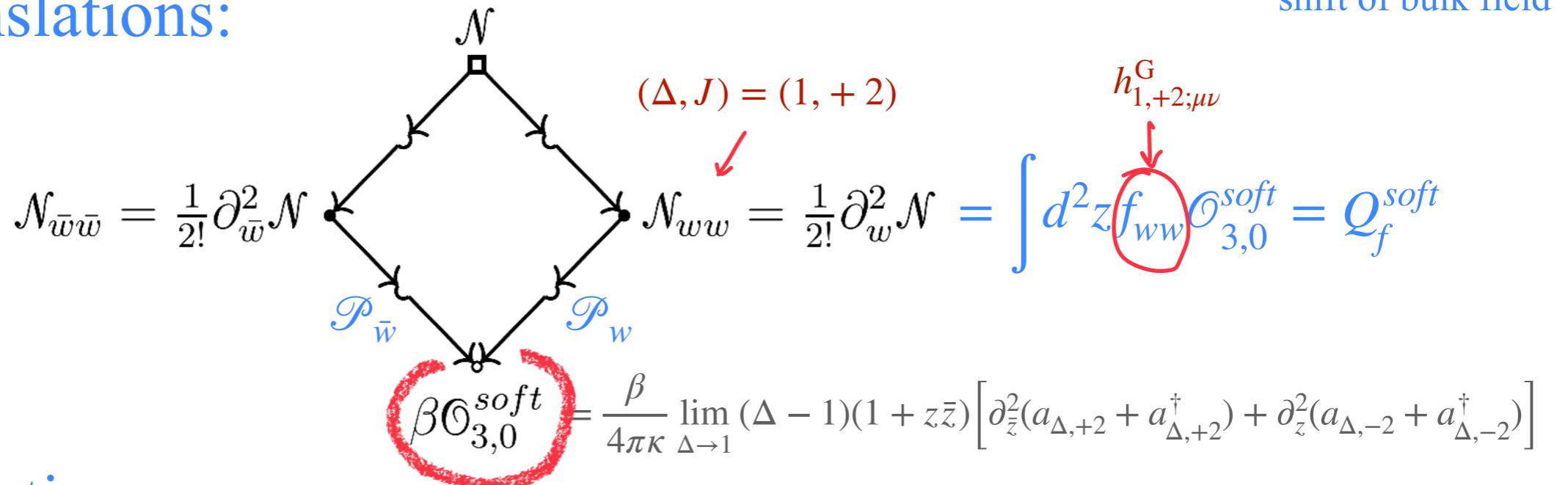
⇒ Capture most subleading (conformally) soft theorems.

Celestial diamonds unify pure gauge & non-pure gauge soft modes.

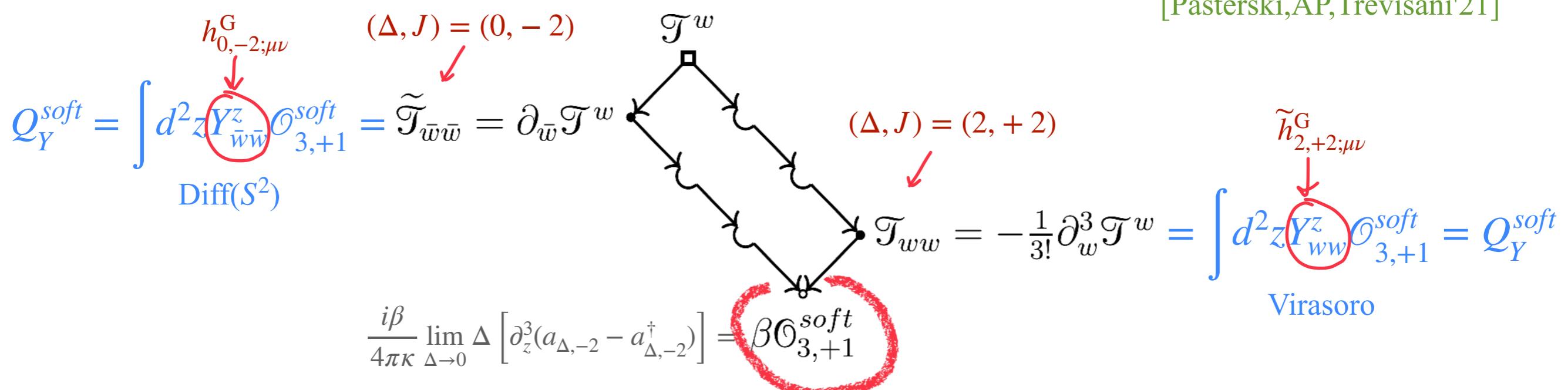
Soft charges

$$[\mathcal{O}_{\Delta,\pm 2}^G, \hat{h}_{\mu\nu}] = i(\nabla_\mu \xi_\nu + \nabla_\nu \xi_\mu) = i h_{\Delta,\pm 2;\mu\nu}^G$$

Supertanslations:



Superrotations:



The symmetry parameters $\{f_{ww}, Y_{ww}^z, Y_{\bar{w}\bar{w}}^z\}$ are Green's fcts $\{\partial_{\bar{z}}^{-2}, \partial_{\bar{z}}^{-1}, \partial_z^{-3}\}$.

Conformal dressings

$$W_{G,j}[\phi_G] | p_j \rangle$$

Faddeev-Kulish dressings render IR divergent amplitudes finite.

j -th hard momentum

Conformally soft dressing (subleading order in soft expansion, leading in coupling κ): [Akhouri,Choi'19]

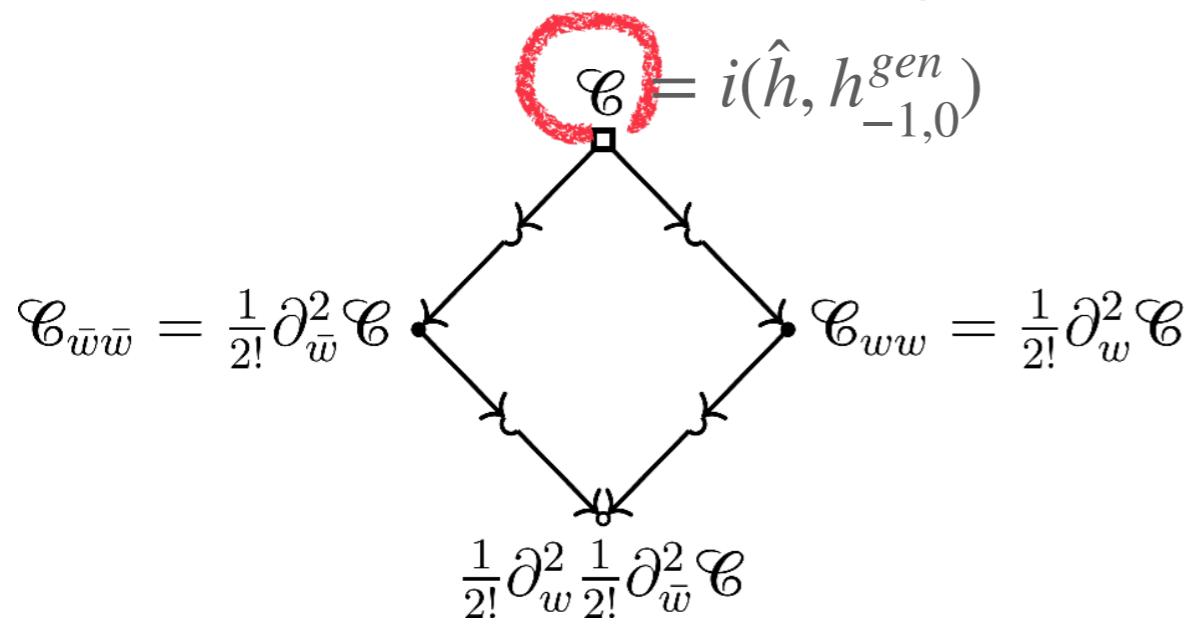
$$W_{G,j} = \exp \left\{ \frac{\kappa}{2} \int \frac{d^3 k \phi_G(\vec{k})}{(2\pi)^3 2k^0} \frac{p_j^\mu}{p_j \cdot k} \left[(p_j^\nu - ik_\rho J_j^{\rho\nu}) a_{\mu\nu}^\dagger - (p_j^\nu + ik_\rho J_j^{\rho\nu}) a_{\mu\nu} \right] \right\}$$

$$\phi_G(\vec{k}) = 1 \quad \overset{\textcolor{red}{\uparrow}}{\textcolor{red}{\mathcal{C}}} = \exp \left\{ -i\kappa \left[\omega_j \mathcal{C} + \bar{h}_j \partial_{\bar{w}_j} \mathcal{Y}^{\bar{w}_j} + \mathcal{Y}^{\bar{w}_j} \partial_{\bar{w}_j} + h_j \partial_{w_j} \mathcal{Y}^{w_j} + \mathcal{Y}^{w_j} \partial_{w_j} \right] \right\}$$

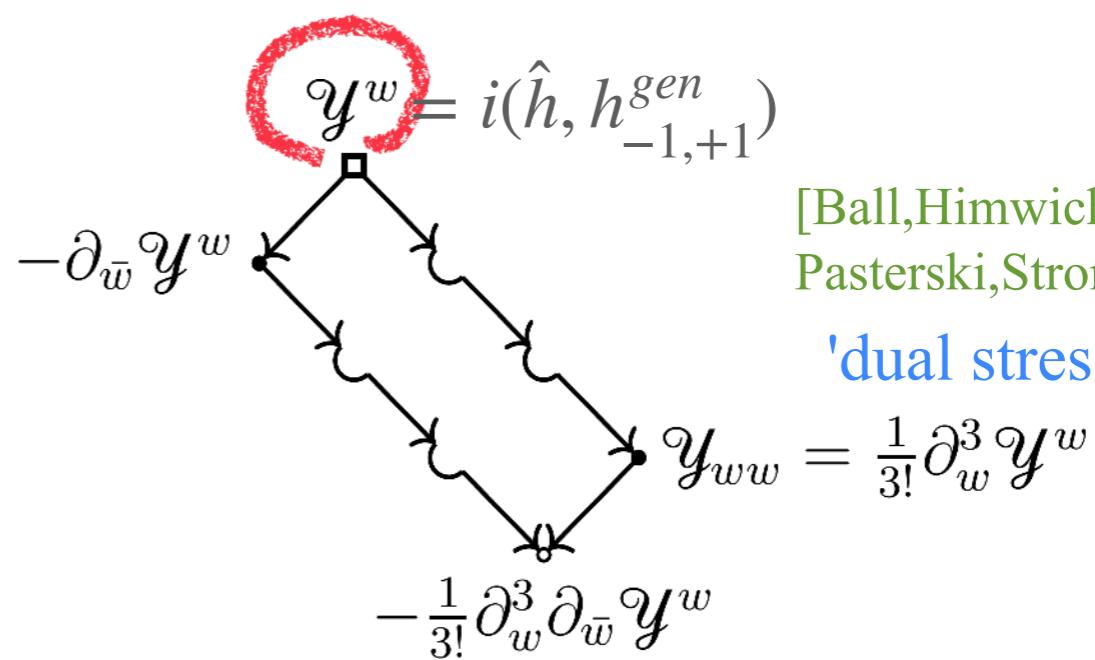
$(\Delta, J) = (-1, 0)$ $(\Delta, J) = (-1, \pm 1)$

conformal dressing

[Arkani-Hamed,Pate,
Raclariu,Strominger'20]



[Pasterski,AP,Trevisani'21]



[Ball,Himwich,Narayanan,
Pasterski,Strominger'19]

'dual stress tensor'

Summary

Celestial diamonds reveal power of symmetry to organize conformally soft physics. Stack via SUSY into celestial pyramids.

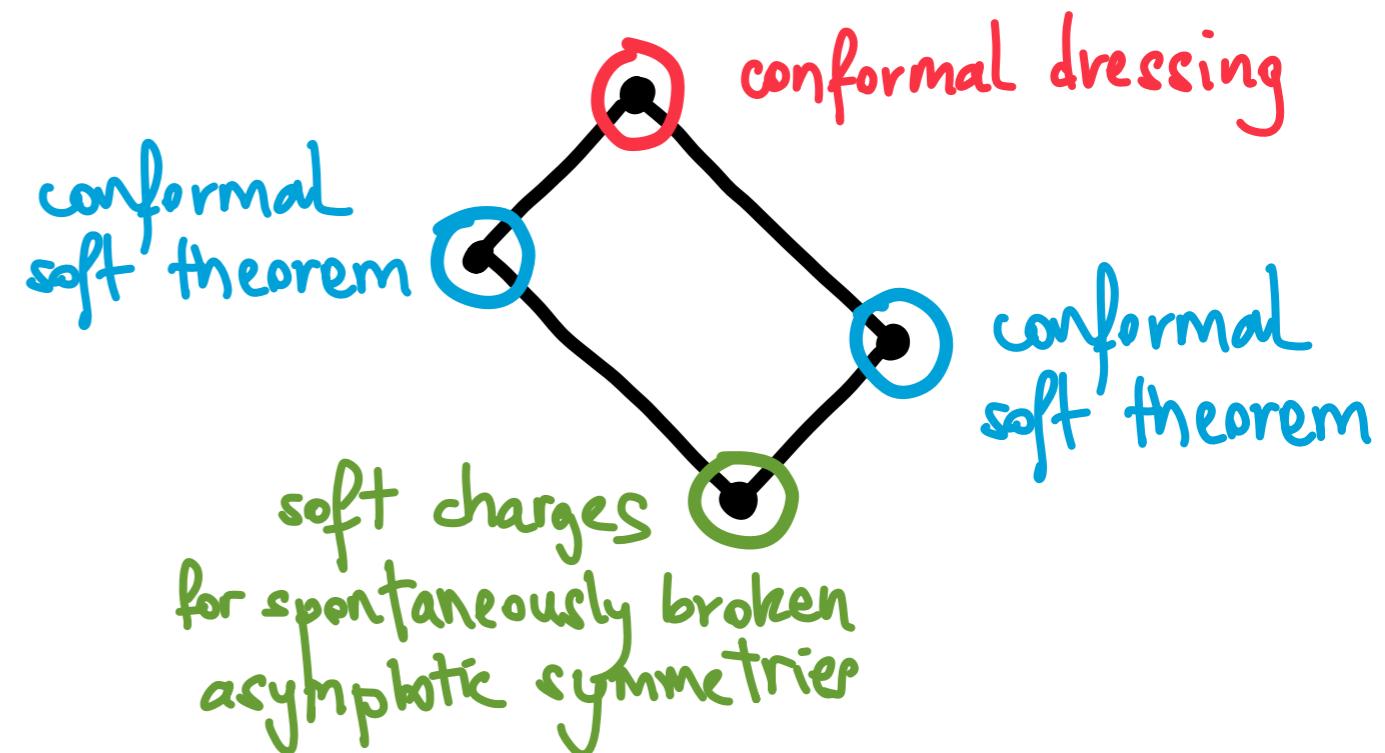
2D operators at the top of the diamond:

- non-trivial levels and central extensions
- effective 2D descriptions of soft sector

Conformal primary wavefunctions satisfy classical double copy in the variety of forms: Kerr-Schild, Weyl and operator-valued.

Celestial amplitudes satisfy operator-valued double copy.

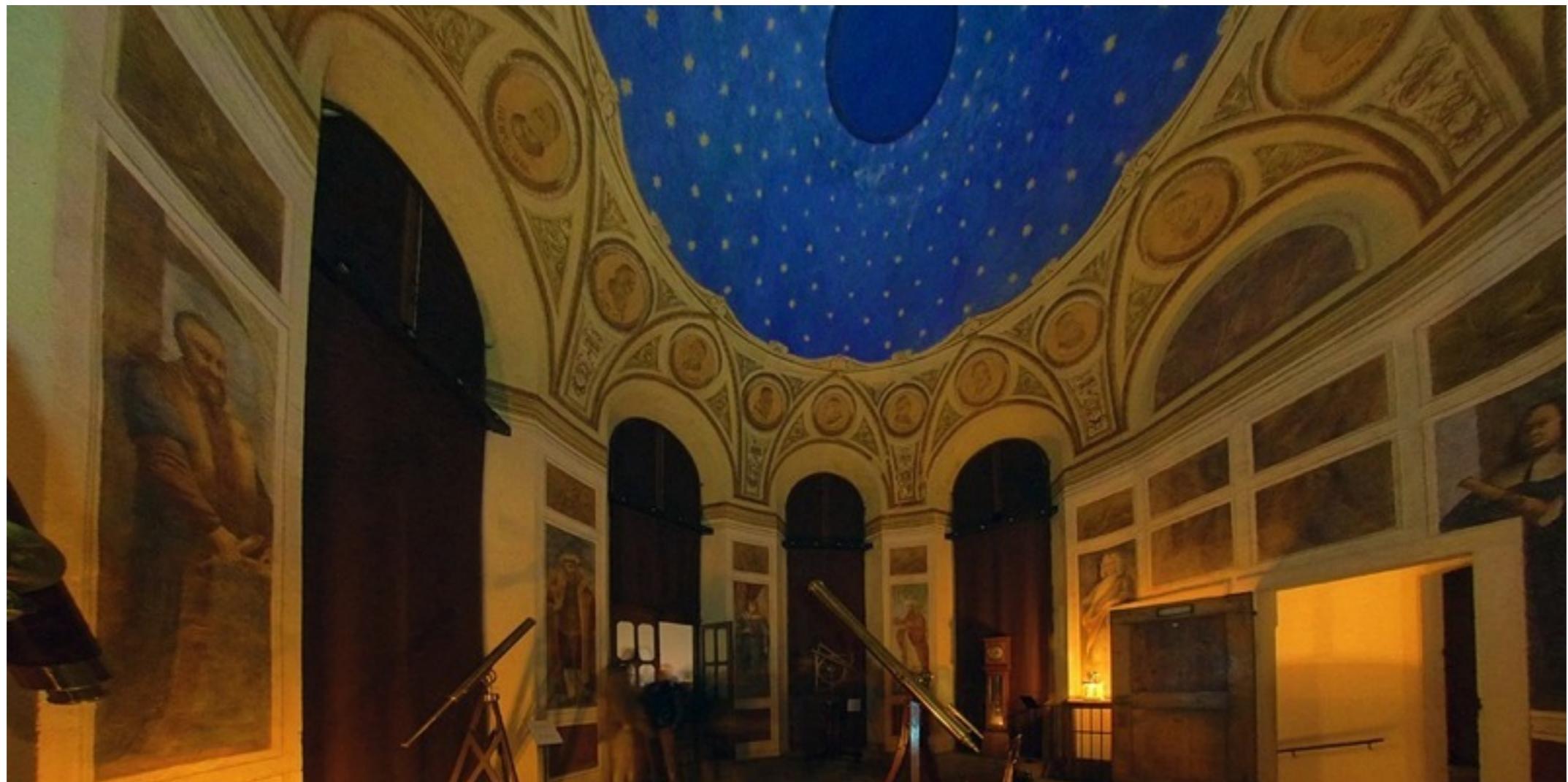
Generalized conformal primaries capture interesting bulk metrics as exact sol to Einstein's equations: ultraboosted black holes, shockwaves,...



→ *Combining tools from amplitudes, classical GR & CFT:*

insight into bulk physics from Celestial Holography programme.

Exploration of celestial territory has only begun!



Thank you!