

A Method to Distinguish **Quiescent** and **Dusty Star-forming** Galaxies **with t-SNE**

John R. Weaver

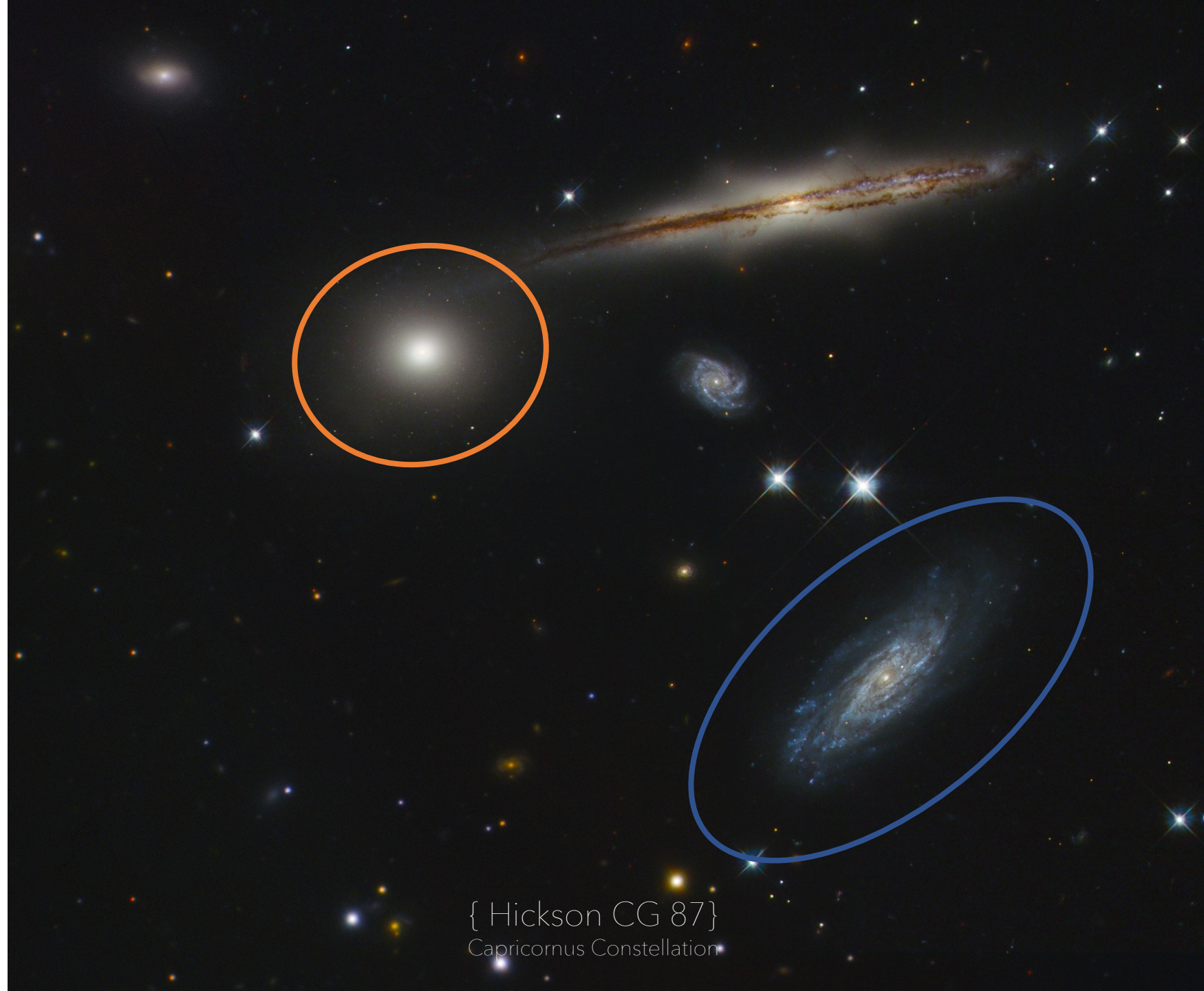
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Charles L. Steinhardt (DAWN)
Jack Maxfield (Caltech)
Iary Davidzon (DAWN)
Dan Masters (Caltech/JPL)
Madeline Schemel (Caltech)
Sune Toft (DAWN)



DAWN

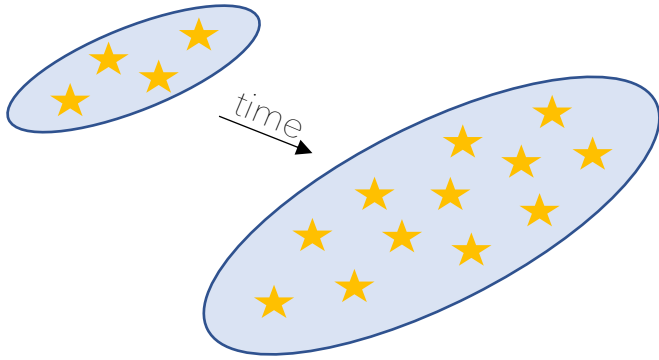
JPL Caltech
Jet Propulsion Laboratory



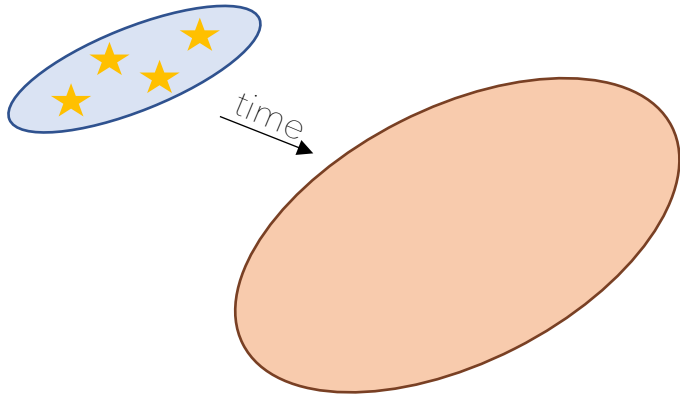
{ Hickson CG 87 }
Capricornus Constellation

The Current Paradigm

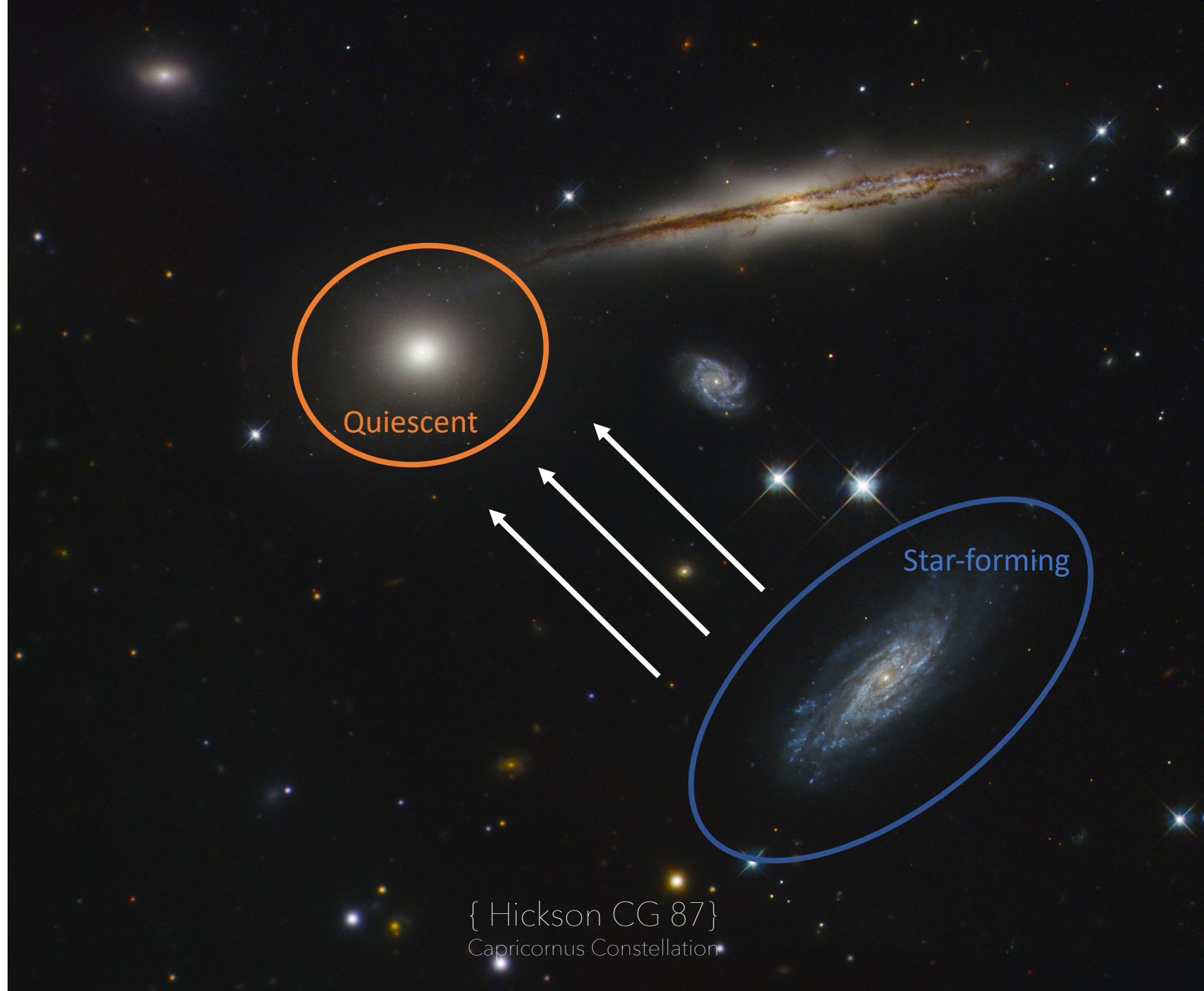
Galaxies grow by forming stars



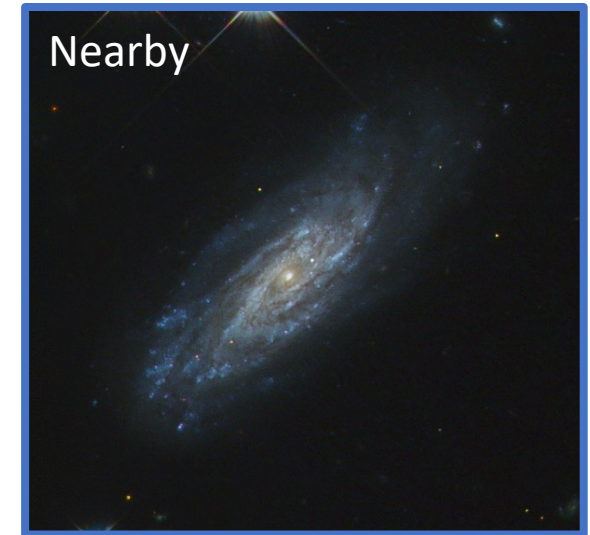
Sometimes they cease forming stars



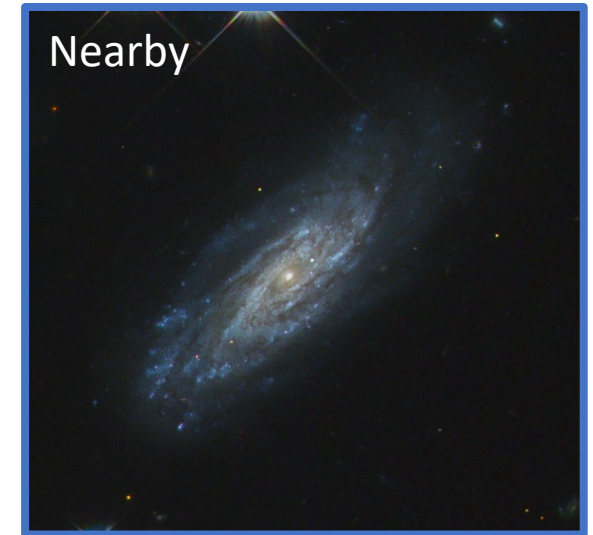
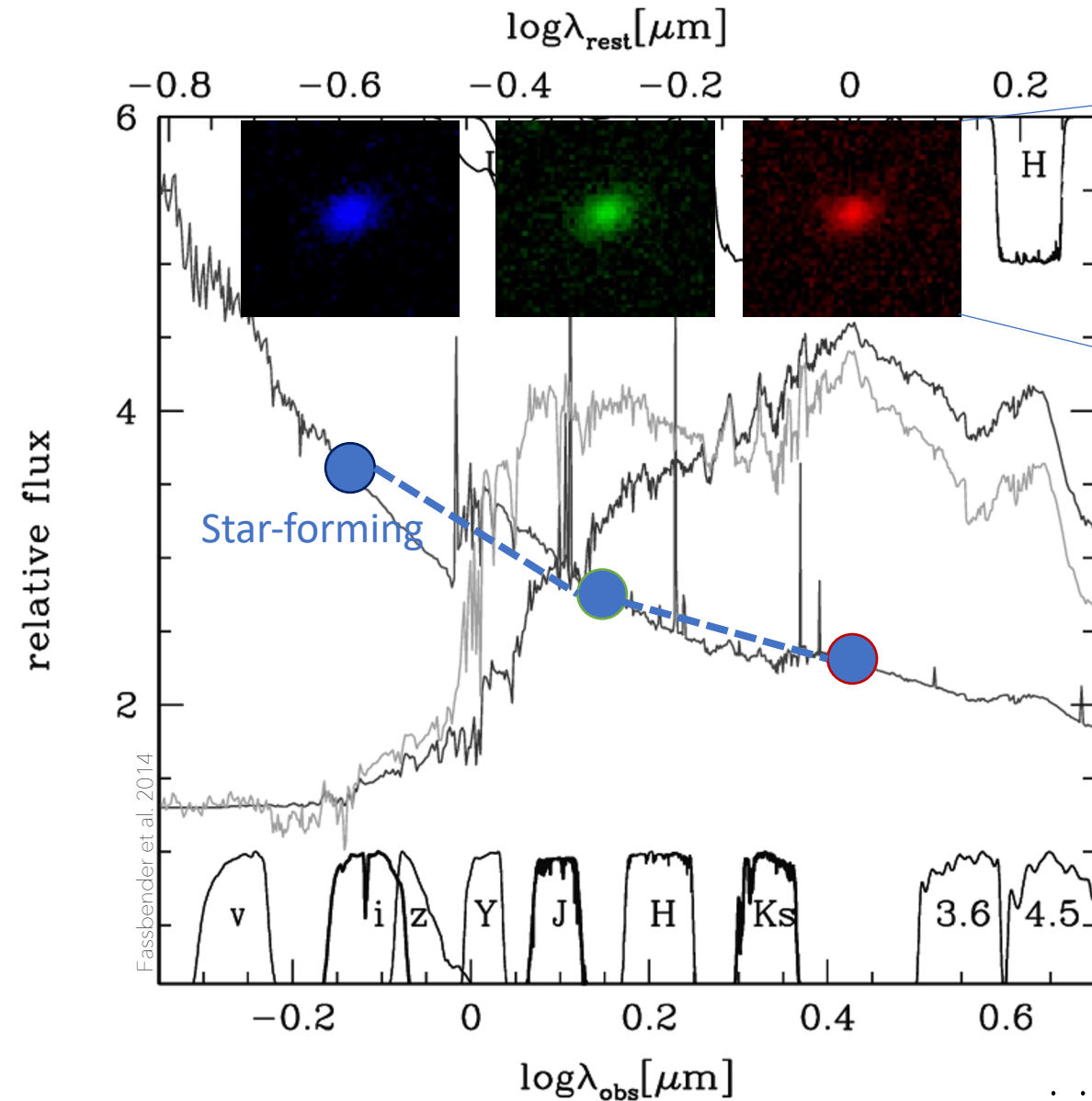
We want to find both kinds
to understand why!



We cannot resolve distant galaxies...

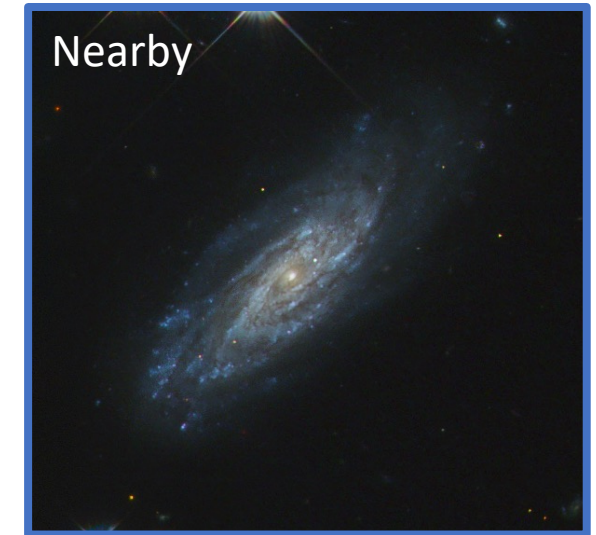
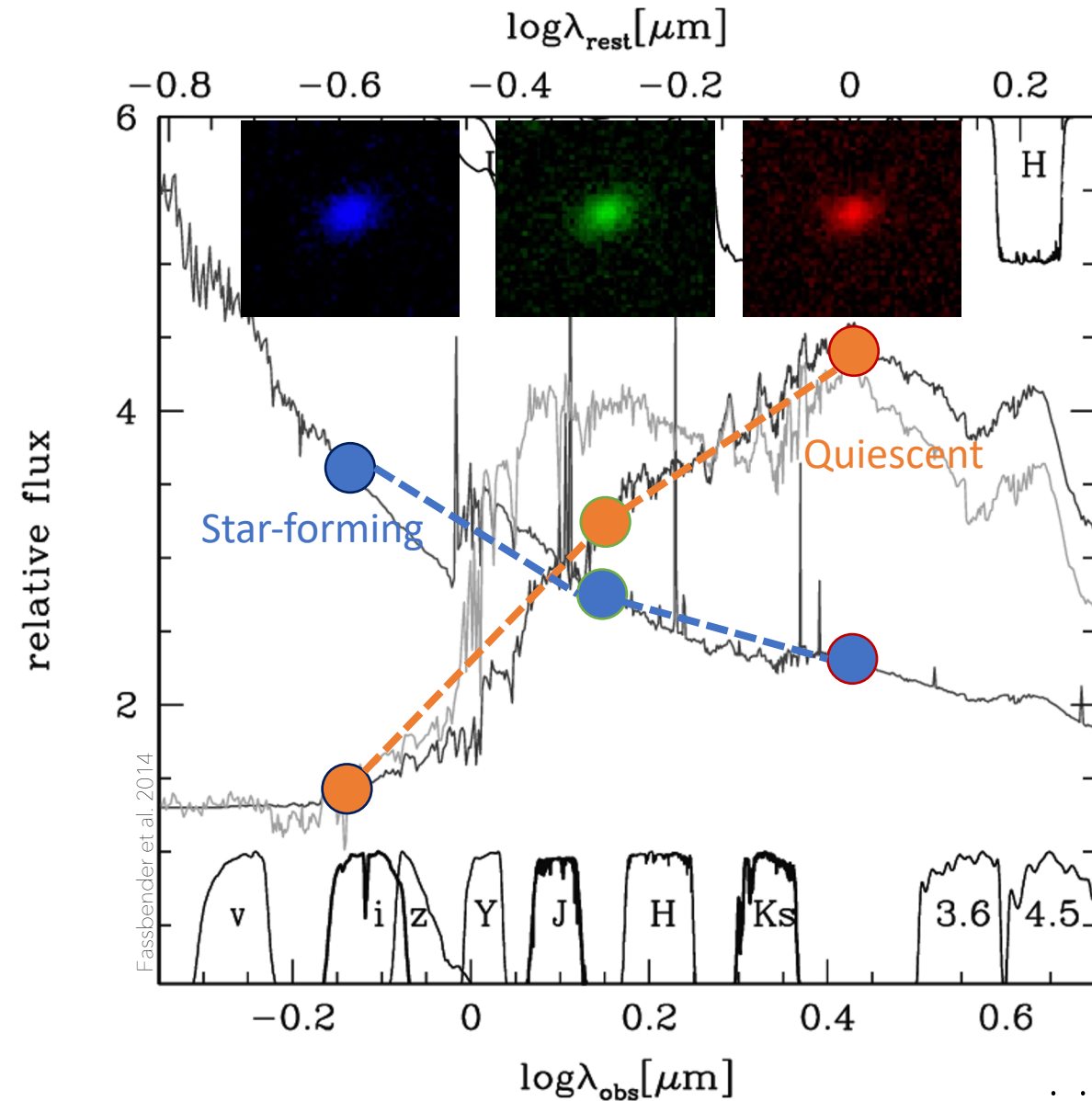


We cannot resolve distant galaxies...



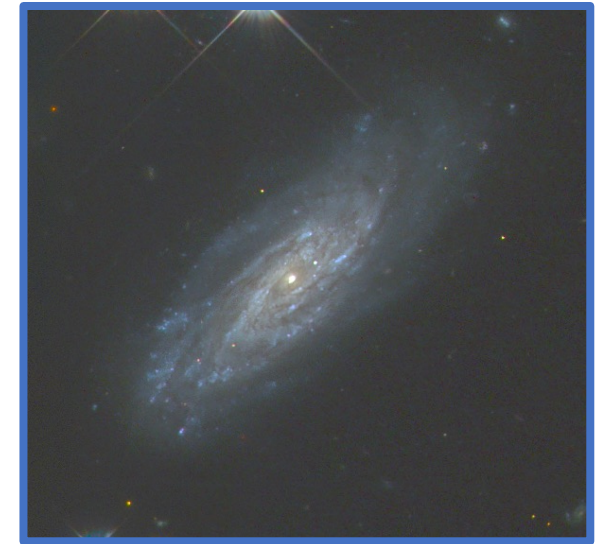
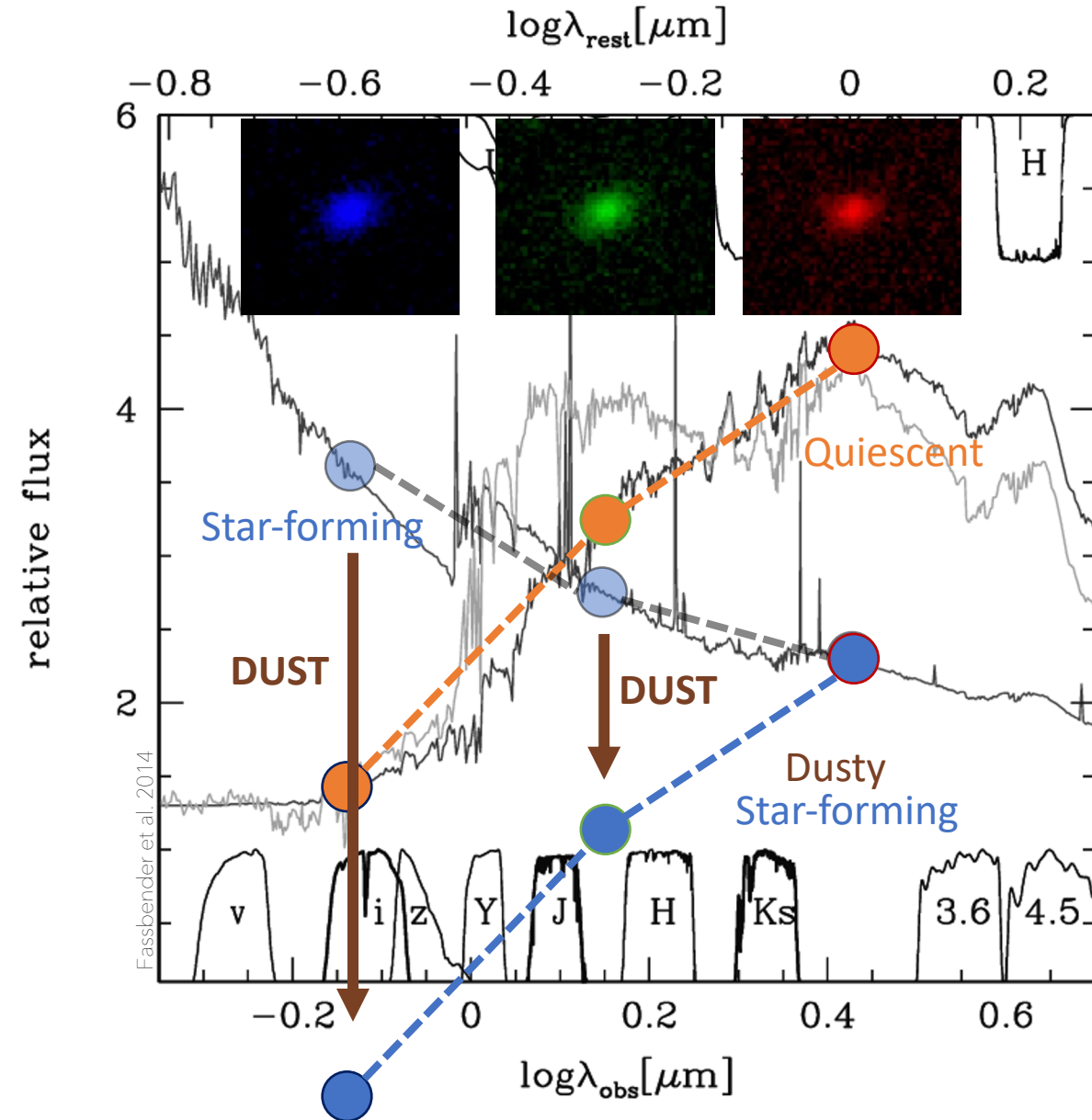
... but we can capture their light!

We cannot resolve distant galaxies...



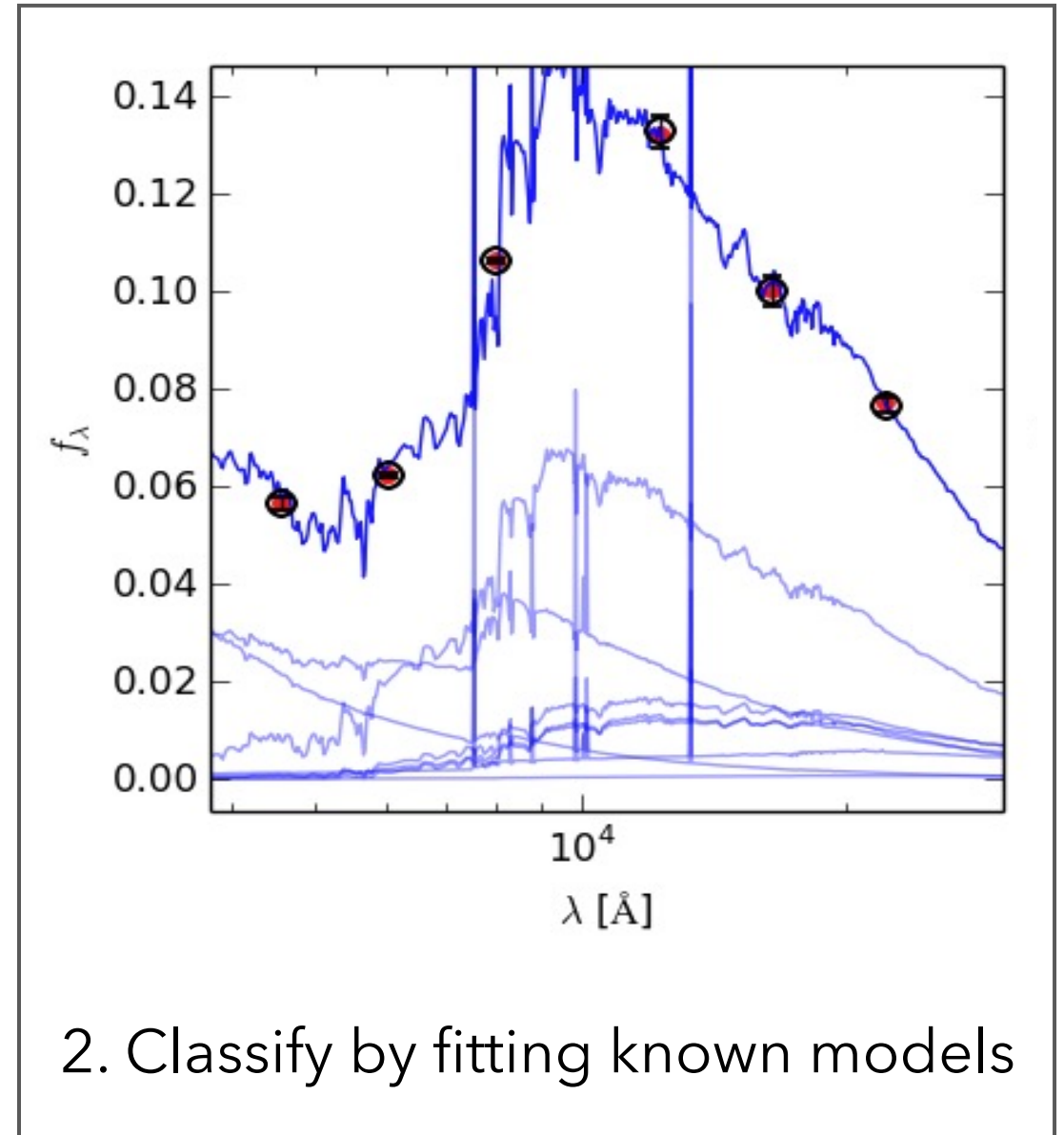
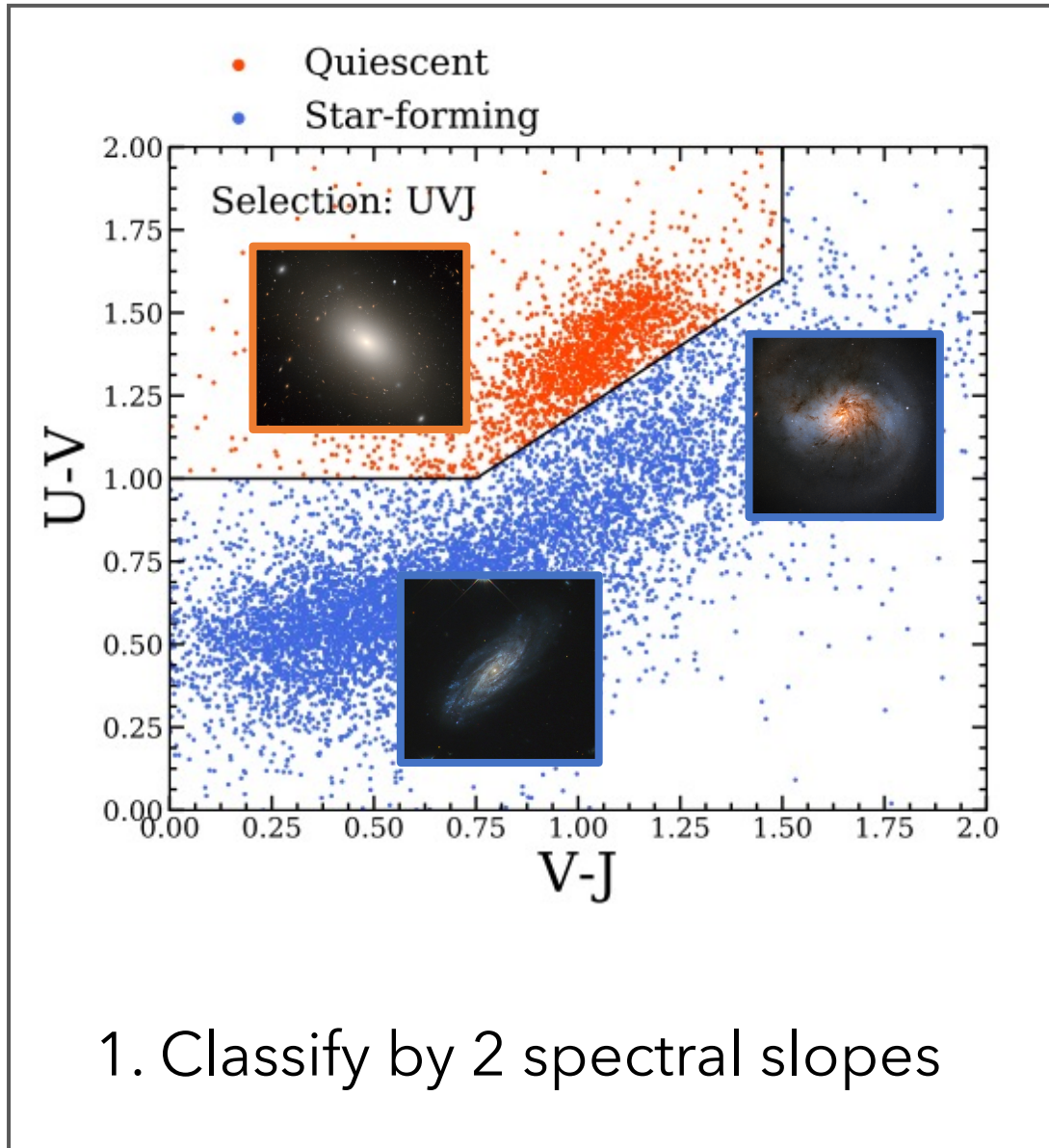
... but we can capture their light!

The Problem: But if you add dust...



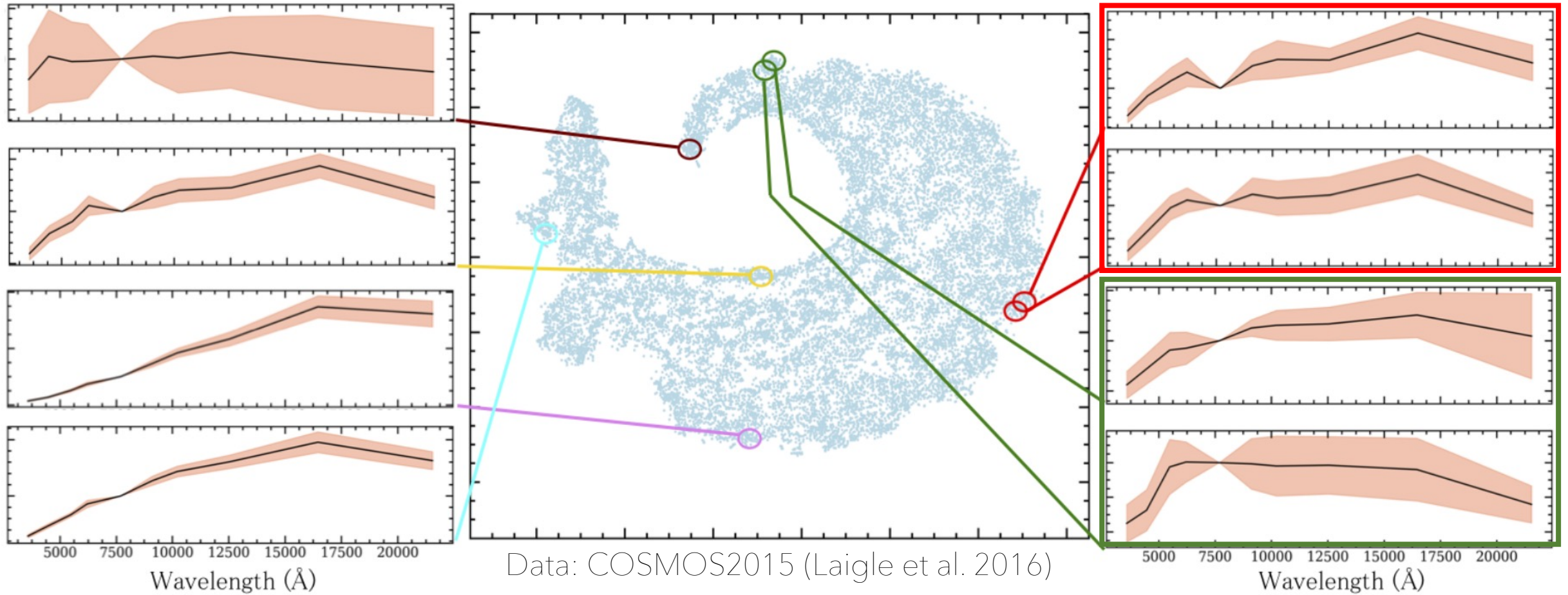
... their light looks the same!

Traditional approaches:



Application of t-SNE

We gave t-SNE 1,000+ photometric measurements of spectra

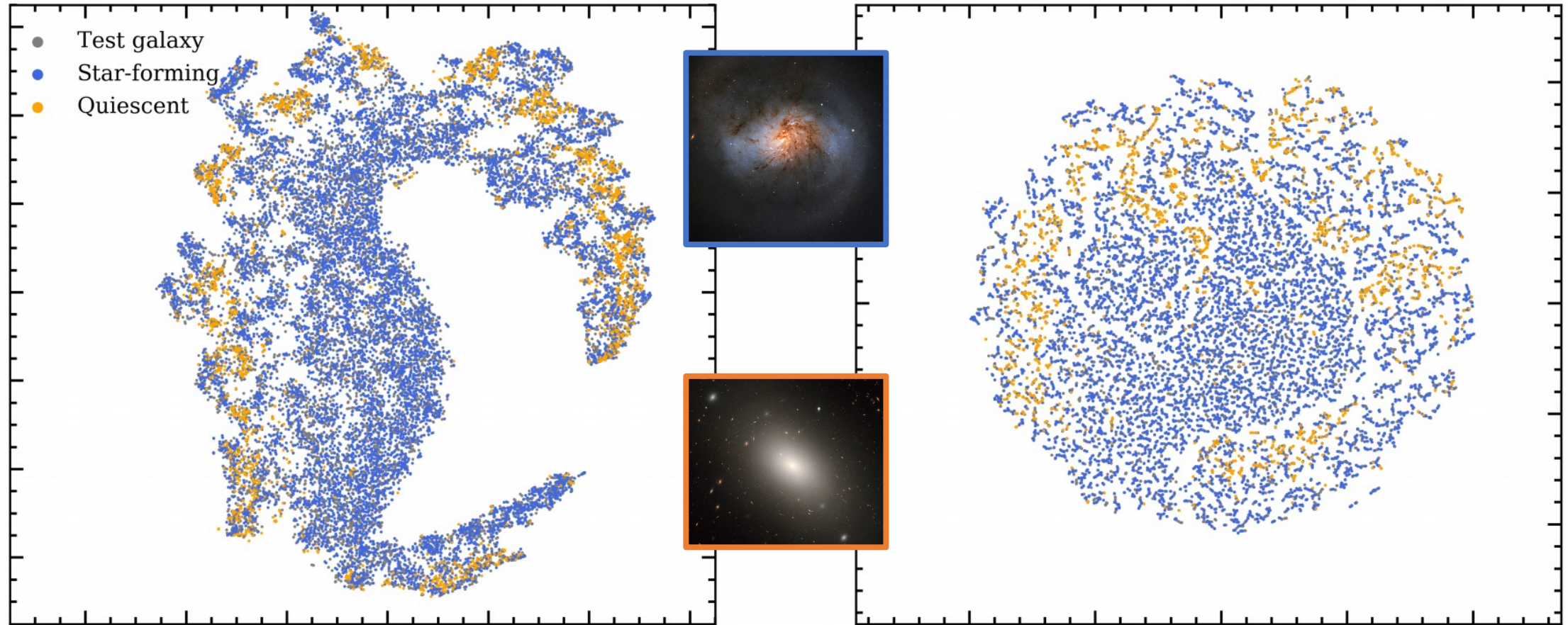


t-SNE successfully groups galaxies by their spectral shapes!

Can t-SNE separate
quiescent galaxies from dusty star-forming ones?

Application of t-SNE

Galaxy sample at 8 Billion years ago ($z \sim 1$)

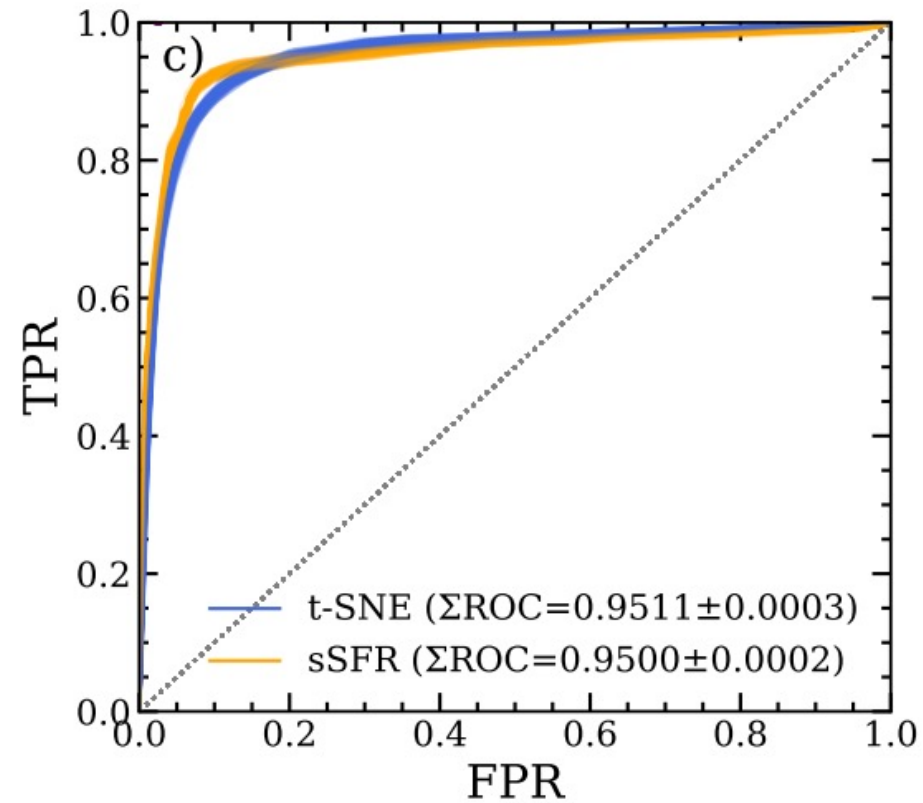


50% are labelled using scarce external data (i.e. 24 μ m)

The other 50% 'tests' are then classified by neighbors

Application of t-SNE

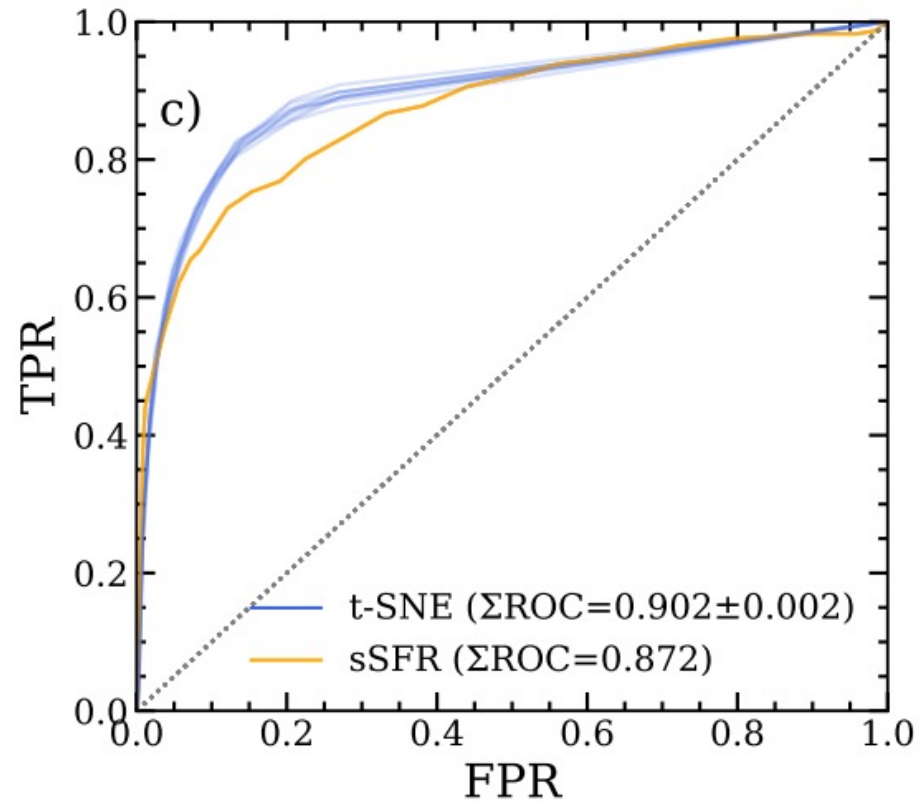
Galaxy sample at 8 Billion years ago ($z \sim 1$)
50% labelled and 50% test



t-SNE achieves better Σ ROC in 63% of trials

Application of t-SNE

Same galaxy labelled sample at 8 Billion years ago ($z \sim 1$)
New galaxy test sample at 10.5 Billion years ago ($z \sim 2$)



t-SNE achieves better ΣROC in 100% of trials

Can t-SNE separate
quiescent galaxies from dusty star-forming ones?



Yes



and better than traditional methods
for more distant, harder to classify galaxies!

{ see Steinhardt, Weaver et al. 2019 }



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