

Effective Field Theories and Neutrinos

Tyler Corbett

Niels Bohr Institute

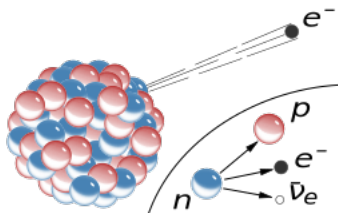
Largely inspired by Y. Cai, J. Herrero Garcia, M. Schmidt, A. Vicente, R. Volkas, arXiv:1706.08524

Outline

- 1 Basics of EFTs
- 2 Neutrino masses and EFTs
- 3 QFTs as a tower of EFTs
- 4 References

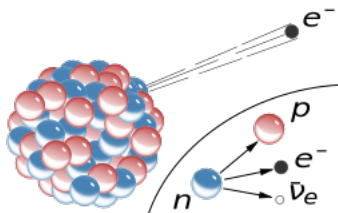
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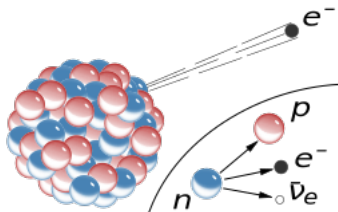
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- let's take the field content and symmetries of QED
- form operator that could explain neutron decay

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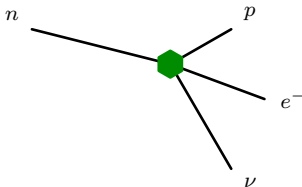


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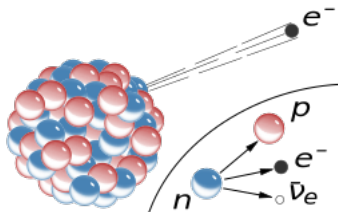
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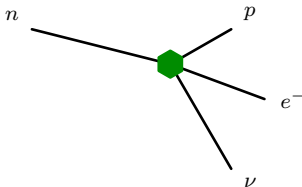
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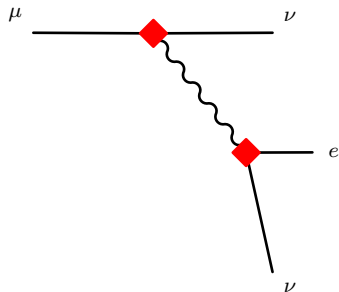
$$\mathcal{L} = G_F(\bar{n}p)(\bar{\nu}e)$$

$[G_F] = \frac{1}{M^2}$



Fermi-theory II: top down, muon decay

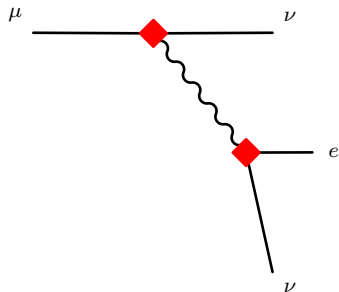
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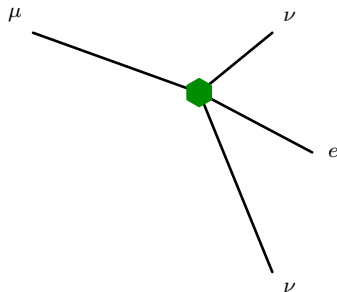
$$\mathcal{L} = \frac{g_W}{\sqrt{2}} \bar{\nu}_\mu \gamma^\mu P_L \mu W_\mu + \frac{g_W}{\sqrt{2}} \bar{e} \gamma^\mu P_L \nu_e W_\mu$$
$$\mathcal{M} \sim \frac{g_W^2}{2} (\bar{\nu}_\mu \gamma^\mu P_L \mu) \frac{g_{\mu\nu}}{k^2 - M_W^2} (\bar{e} \gamma^\nu P_L \nu_e)$$

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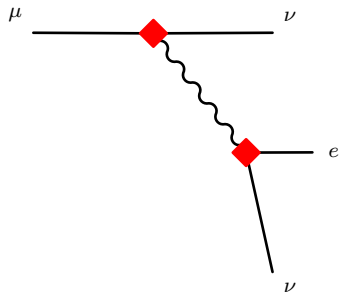
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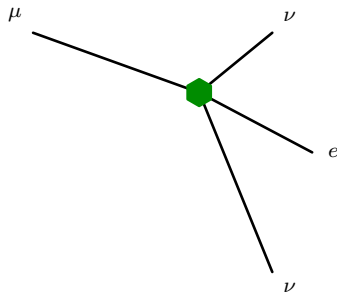
$$\mathcal{M} \sim -\frac{g_W^2}{2M_W^2} (\bar{\nu}_\mu \gamma^\mu P_L \mu) (\bar{e} \gamma^\mu P_L \nu_e) + \dots$$

Fermi-theory II: top down, muon decay

In the SM



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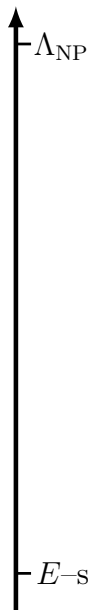
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$$\frac{1}{M_W^2} (\bar{\psi} \gamma^\mu P_L \psi)^2$$

$$\frac{1}{M_W^4} \partial^2 (\bar{\psi} \gamma^\mu P_L \psi)^2$$

$$\mathcal{M} \sim -\frac{g_W^2}{2M_W^2} (\bar{\nu}_\mu \gamma^\mu P_L \mu) (\bar{e} \gamma^\mu P_L \nu_e) - \frac{g_W^2 k^2}{2M_W^4} (\bar{\nu}_\mu \gamma^\mu P_L \mu) (\bar{e} \gamma^\mu P_L \nu_e) + \dots$$



The major underlying assumption of any EFT

$\Lambda_{NP} \gg E$ of the scale of experiments/measurements

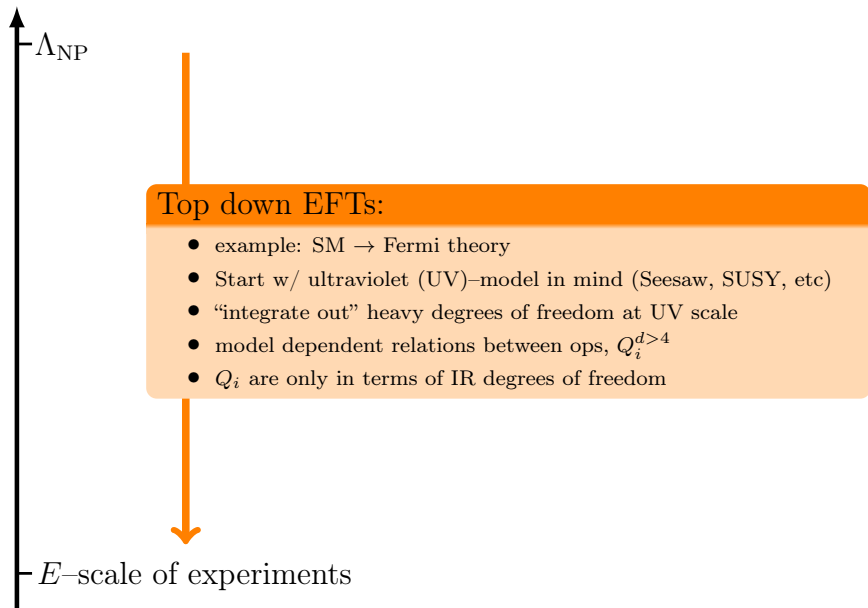


Λ_{NP}

E -scale of experiments

Bottom up EFTs:

- example: Fermi theory
- Start w/ infrared (IR)-model in mind (QED, SM)
- using symmetries of model put together ops, $Q^{d>4}$
- truncate EFT at some $\mathcal{O}(1/\Lambda)$
- constrain Q in experiment & infer properties of NP at Λ
- Q s are unrelated \rightarrow model independent
- Q s are only in terms of IR degrees of freedom



SMEFT

In studying NP at $\Lambda_{\text{NP}} \gg v$, we (mostly) employ the [Standard Model EFT](#)

$$\mathcal{L}_{\text{SMEFT}} = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda} \mathcal{L}_5 + \frac{1}{\Lambda^2} \mathcal{L}_6 + \dots \qquad \mathcal{L}_d = \sum_i c_i Q_i$$

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D6 terms are more collider
pheno relevant

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The leading operator:

$$\mathcal{L}_5 = c_{\alpha\beta} (\bar{L}_\alpha^c \tilde{H}^\dagger) (\tilde{H}^\dagger L_\beta) \sim v^2 \bar{\nu}_\alpha^c \nu_\beta$$

$$\Rightarrow m_\nu \sim v^2 / \Lambda$$

The Weinberg operator

S. Weinberg, Phys.Rev.Lett. 43 (1979) 1566-1570

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below EWSB this gives:

$$H \rightarrow \left[\begin{array}{c} \phi^+ \\ \frac{v+h+\phi^0}{\sqrt{2}} \end{array} \right] \quad L_\alpha \rightarrow \left[\begin{array}{c} \nu_{L,\alpha} \\ e_{L,\alpha} \end{array} \right]$$

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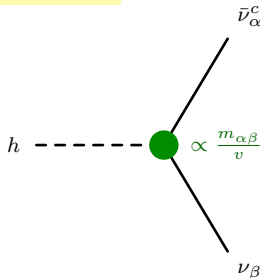
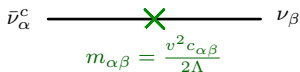
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Top down neutrino masses

Previously:

- Bottom up: Fermi Theory $\frac{1}{M_W^2}(\bar{\psi}\psi)(\bar{\psi}\psi) \Rightarrow W\text{-boson \& SM}$
- Top down: $W\text{-boson} \Rightarrow \text{Fermi theory } \frac{1}{M_W^2}(\bar{\psi}\psi)(\bar{\psi}\psi)$

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Now:

- Bottom up: Weinberg Operator $\frac{1}{\Lambda}(\bar{L}^c H)(HL)$
- Top down: ???

Type-I seesaw

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Let's try Type-I seesaw (heavy majorana ν_R):

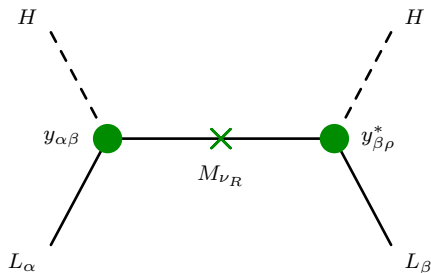
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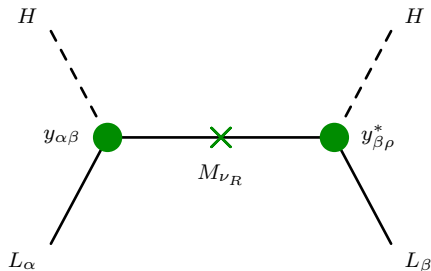
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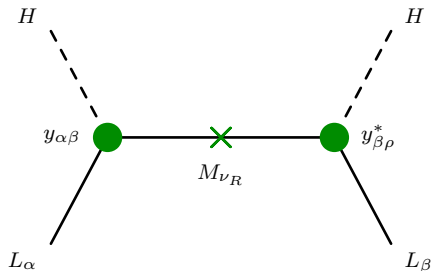
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This corresponds to:

$$\mathcal{L}_5 = \frac{c_{\alpha\rho}}{\Lambda} (\bar{L}_\alpha^c \tilde{H})(\tilde{H}^\dagger L_\beta) \quad c_{\alpha\rho} \sim y_{\alpha\beta} y_{\beta\rho}^*, \quad \Lambda = M_{\nu_R}$$

Type II Seesaw

What about the other seesaws?

Type-II seesaw (heavy triplet scalar Δ):

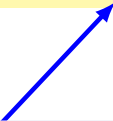
$$\mathcal{L} = \mathcal{L}_{SM} + \bar{\Delta}^a (D^2 - M^2) \Delta^a + g H^\dagger \sigma^a \tilde{H} \Delta^a + y_{\alpha\beta} \bar{L}_\alpha^c \sigma_2 \tau^a L_\beta \Delta^a$$

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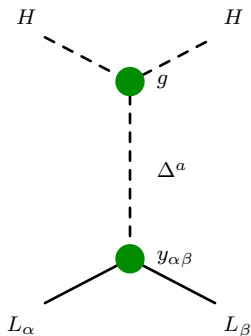


g has units of mass

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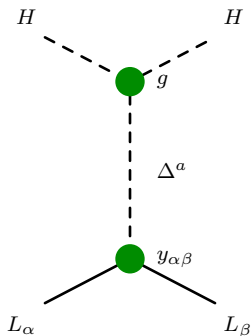


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Again we obtain the Weinberg operator:

$$\mathcal{L}_5 = \frac{c_{\alpha\beta}}{\Lambda} (\bar{L}_\alpha^c \tilde{H})(\tilde{H}^\dagger L_\beta) \quad c_{\alpha\beta} \sim y_{\alpha\beta}, \quad \Lambda = \frac{M_\Delta^2}{g}$$

Type III Seesaw

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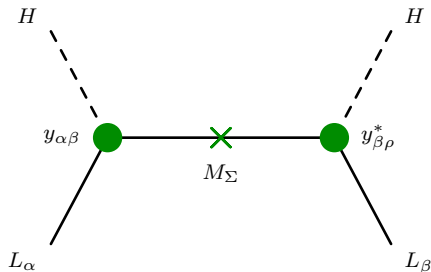
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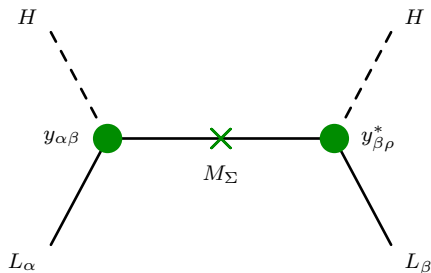


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Model Independence

3 different models of ν mass generation give the same (leading) operator.

Implicit assumptions:

- New particles were much heavier than SM particle content:

$$m_\nu \sim .05 \text{ eV} \sim \frac{v^2}{\Lambda_{\text{seesaw}}} \Rightarrow \Lambda \sim \frac{(246 \text{ GeV})^2}{.05 \cdot 10^{-9} \text{ GeV}} \sim 10^{15} \text{ GeV}$$

- We assumed SM gauge symmetries and fields (well justified, except maybe Higgs doublet)
- We assumed the scale of generation of ν masses was above EWSB, i.e. used SM doublet fields like L and H instead of components ν and h

THE SEESAW SCALE



What if I want to see the particles responsible for ν masses in my lifetime?

Kumerički et al. arXiv:1204.6599

Fermionic Quintuplet, $(1, 5, 0)$: Σ^A

Scalar Quadruplet, $(1, 4, -1)$: Φ^I

$$\mathcal{L} = \mathcal{L}_{SM} + i\bar{\Sigma}^A(\not{D} - M)\Sigma^A - (\Phi^I)^\dagger(D^2 - M^2)\Phi^I - (y_{\alpha\beta}\bar{L}_\alpha\Phi\Sigma_\beta + h.c.) + \lambda\Phi H^3 + \dots$$

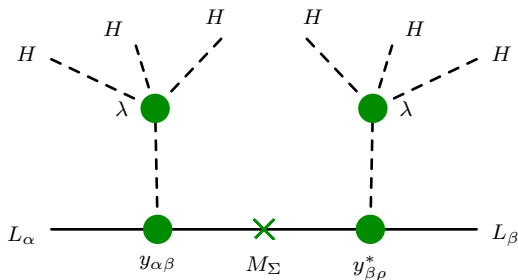
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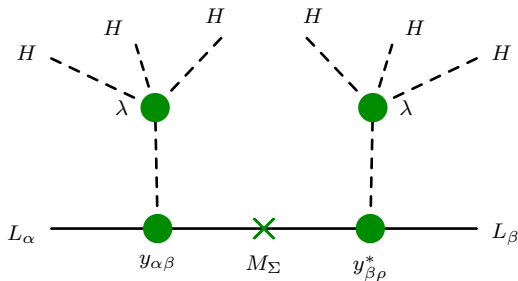


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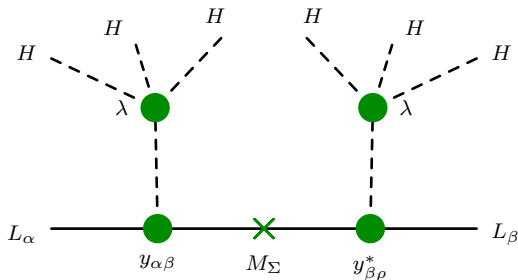
$$\mathcal{M} \sim y_{\alpha\beta} y_{\beta\rho}^* \lambda^2 \frac{\bar{\psi}(\not{p} + M)\psi}{(p^2 - M^2)^3} \rightarrow \frac{\lambda^2 y_{\alpha\beta} y_{\beta\rho}^*}{M^5} \bar{\psi}\psi + \mathcal{O}\left(\frac{p^2}{M^7}\right)$$

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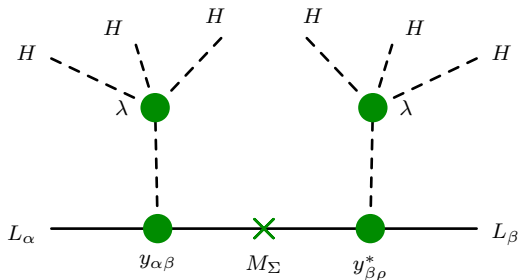
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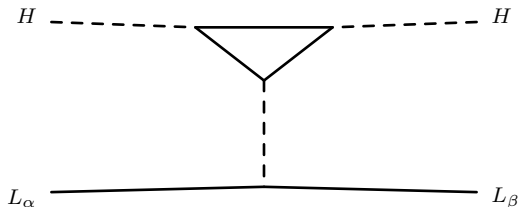
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$$m_\nu \sim .05 \text{ eV} \sim \frac{v^6}{\Lambda^5} \quad \Rightarrow \quad \Lambda \sim \sqrt[5]{\frac{(246 \text{ GeV})^6}{.05 \cdot 10^{-9} \text{ GeV}}} \sim 85 \text{ TeV}$$

What if I want to see the particles responsible for ν masses in my lifetime? II

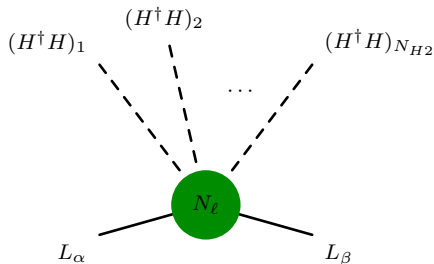
Could induce the Weinberg operator via loops:



$$m_\nu \sim .05 \text{ eV} \sim \frac{1}{16\pi^2} \frac{v^2}{\Lambda} \quad \Rightarrow \quad \Lambda \sim \frac{(246 \text{ GeV})^2}{16\pi^2 (.05 \cdot 10^{-9} \text{ GeV})} \sim 10^{12} \text{ GeV}$$

What if I want to see the particles responsible for ν masses in my lifetime? II

Or a combination of mass suppressions and loops:



$$m_\nu \sim .05 \text{ eV} \sim \left(\frac{1}{16\pi^2} \right)^{N_\ell} \left(\frac{v}{\Lambda} \right)^{2N_{H2}-1} v$$

N_ℓ	N_{H2}	Λ
0	1	10^{15} GeV
0	2	10^6 GeV
0	3	85 TeV
0	4	16 TeV
0	5	6 TeV
1	1	10^{12} GeV
1	2	700 TeV
1	3	31 TeV
1	4	8 TeV
2	1	10^{10} GeV
2	2	140 TeV
2	3	11 TeV
3	1	10^8 GeV
3	2	26 TeV
4	1	10^6 GeV
4	2	5 TeV

Discussed:

- EFTs are **model independent** (mostly)
3 tree level models with different field content/couplings \rightarrow 1 operator
- **Simpler low E dynamics**
can study the SM + one or some operators instead of many UV models
- “opening up” the operators
can motivate UV models by an operator you want to generate in the IR
If we want TeV scale neutrino mass generation, e.g. $N_\ell = 1$, $N_{H2} = 4$

Important, but not discussed:

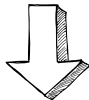
- EFTs allow us to deal with **non-perturbative physics**
e.g. if strong dynamics were to generate ν masses
e.g. the theory of π, κ, p, n (chiral perturbation theory) is an EFT
- can adapt by including RH neutrinos in ops for Dirac neutrinos
- modern view of QFTs is a tower of EFTs

Beyond the Standard Model

$$\mathcal{L}_{\text{BSM}} = \mathcal{L}(\Phi_{\text{NP}}, H, Q, L, u_R, d_R, e_R, W^I, B, G^A)$$

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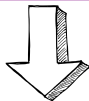


Standard Model + SMEFT

$$\mathcal{L}_{\text{SM}} + \sum_n \frac{1}{\Lambda^n} \mathcal{L}^{(n+4)} = \mathcal{L}(H, Q, L, u_R, d_R, e_R, W^I, B, G^A)$$

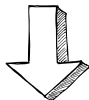
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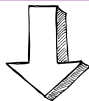


LEFT

$$\mathcal{L}_{\text{QED}} + \mathcal{L}_{\text{QCD}} + \sum_n \frac{1}{v^n} \mathcal{L}^{(n+4)} = \mathcal{L}(2 \times u, 3 \times d, 3 \times e, \gamma, G^A)$$

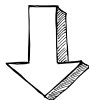
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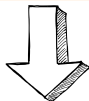
Standard Model + SMEFT

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QED+

$$\mathcal{L}_{\text{QED}} + \sum_n \frac{1}{\Lambda_{\text{QCD}}^n} \mathcal{L}^{(n+4)} = \mathcal{L}(\gamma, e^\pm, \pi^{\pm,0}, p, n)$$

A tower of EFTs

Top down \Rightarrow correlations between operators in the IR theory,
Above EWSB:

$$\mathcal{L} = i\bar{L}\not{D}L = i\bar{L}\gamma^\mu(\partial_\mu - \frac{i}{2}g_1 B_\mu + ig_2\tau^A W_\mu^A)L$$

Below EWSB:

$$\mathcal{L} \rightarrow -\frac{1}{2}Z_\mu [(g_1 s_W + g_2 c_W)\bar{\nu}_L\gamma^\mu\nu_L + (g_1 s_W - g_2 c_W)\bar{e}_L\gamma^\mu e_L]$$

An example interaction, Bustamante et al. arXiv:2001.04994

$$\mathcal{L} \sim g_{\alpha\beta} \phi \bar{\nu}_\alpha \nu_\beta$$

UV completing this theory requires:

- rewrite interaction in terms of $L_\alpha = \{\nu_{L,\alpha}, \ell_\alpha^-\}$
→ implies charged interactions with ϕ → more stringent constraints
- or assume the SM is ‘wrong’

So searches like these:

- look for new physics
- but also test the SM (via implied relations between interactions)

Some references

- ① On Effective Field Theories, Schwartz QFT book, Ch22
- ② Formalism of Integrating out particles, arXiv:1412.1837
- ③ On the Weinberg operator and UV completions, arXiv:1706.08524 (used in this talk)
- ④ On the SMEFT, arXiv:1706.08945
- ⑤ An alternative to the SMEFT, the HEFT, arXiv:1311.1823
- ⑥ Integrating out the W, Z, t, H from SMEFT to LEFT, arXiv:1908.05295

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