

The near-infrared universe
as shown by UltraVISTA

BUFFALO Meeting | July 12 2021

COSMOS2020

new insights into galaxy assembly and evolution at high- z

John R. Weaver

PhD Fellow @ Cosmic Dawn Center, Univ. of Copenhagen



@astroweave

with Olivier Kauffmann, Marko Shuntov, Iary Davidzon,
Olivier Ilbert, Gabe Brammer, Peter Capak,
Clotilde Laigle, Bo Milvang-Jensen, Paul B.C. Hsieh,
Andrea Moneti, Henry J. McCracken, Sune Toft
and the rest of the COSMOS team

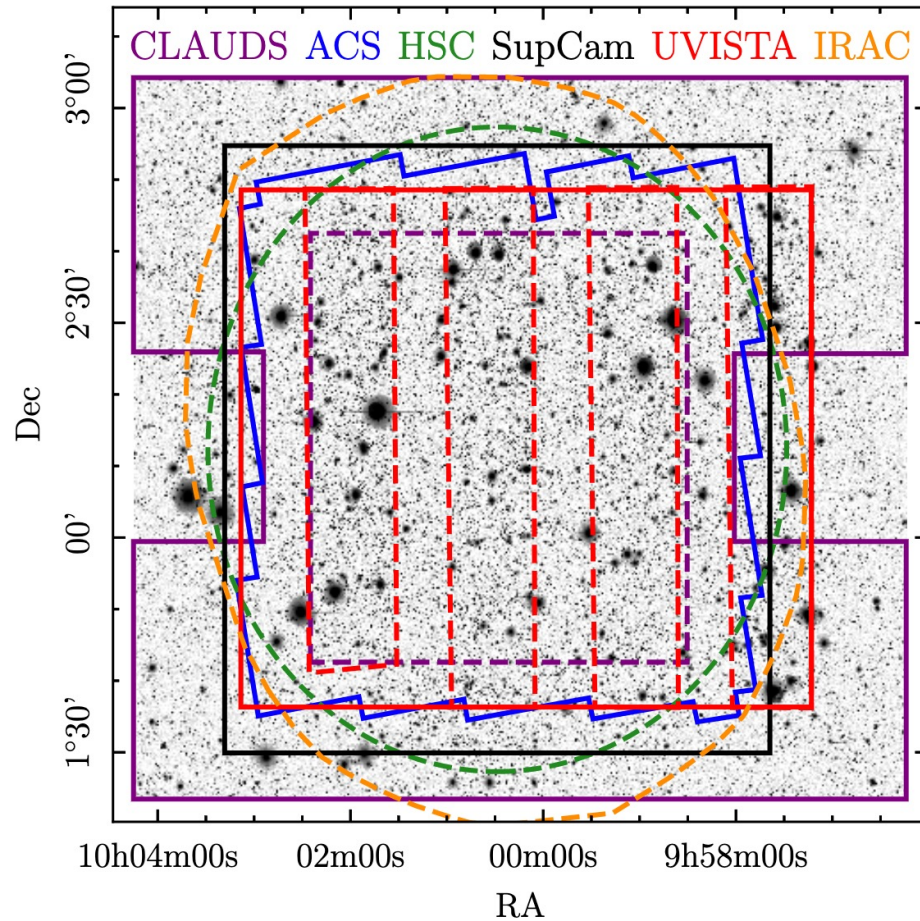


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COSMOS2020 *at a glance*

NIR-selected catalog of ~1M galaxies over 2 deg²

{Weaver et al., submitted}



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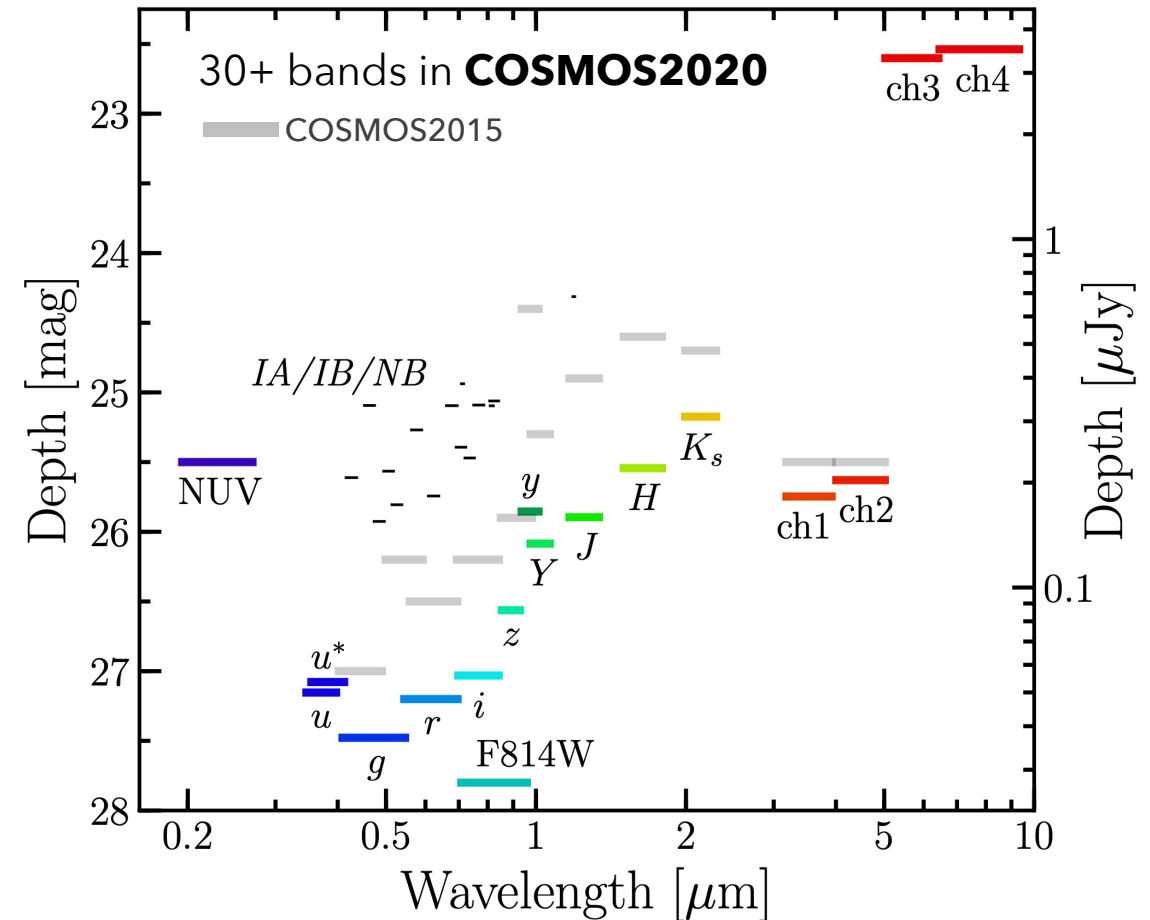
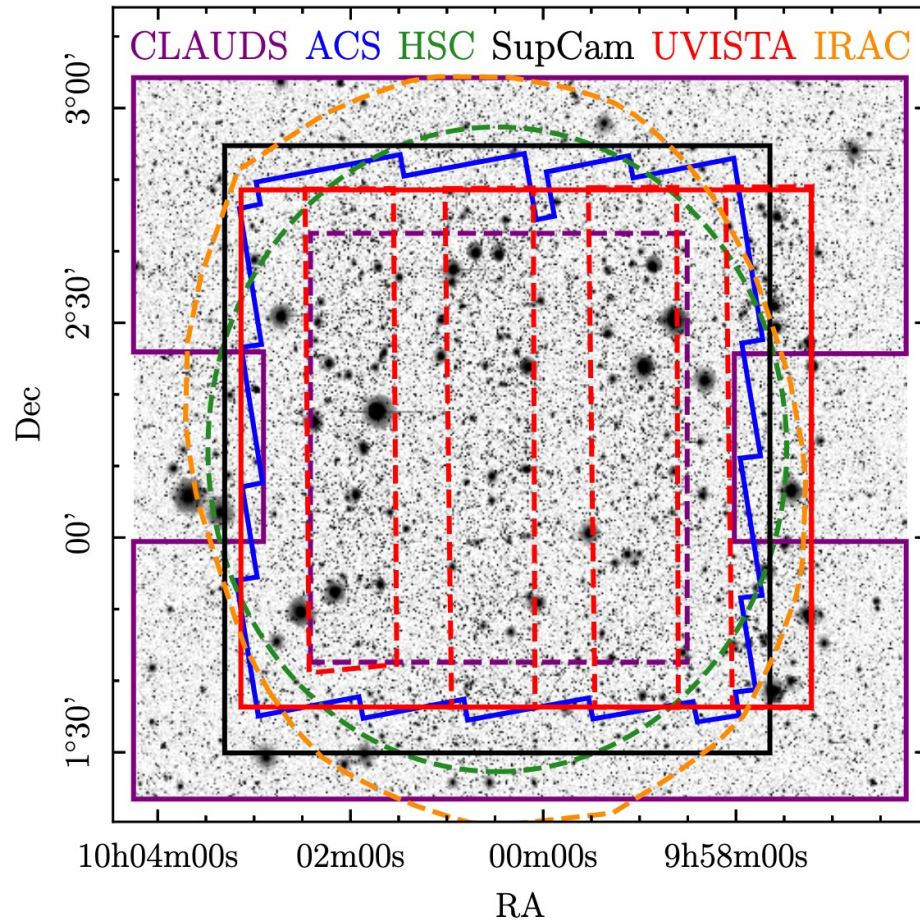
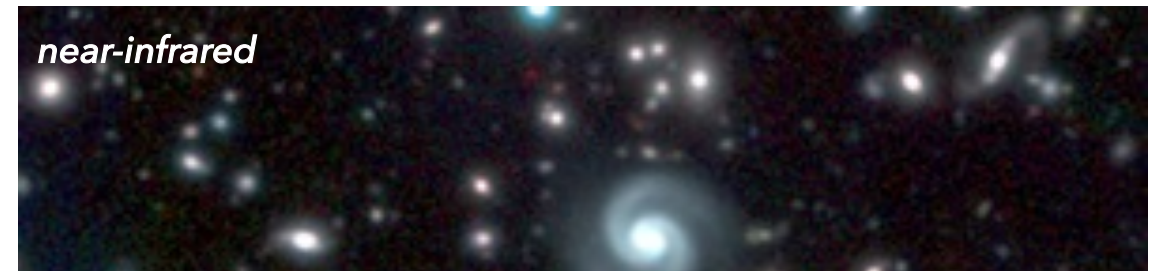
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COSMOS2020 *at a glance*

NIR-selected catalog of $\sim 1\text{M}$ galaxies over 2 deg^2

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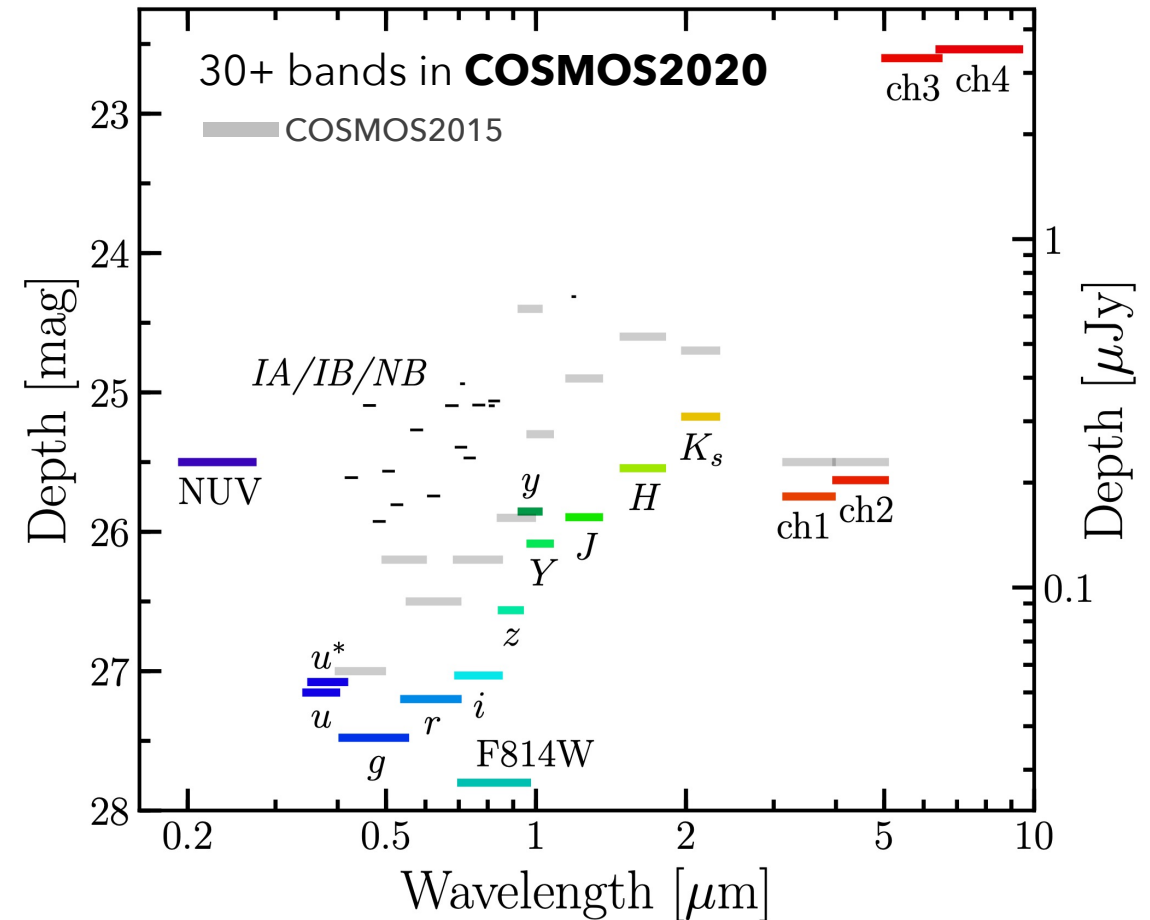
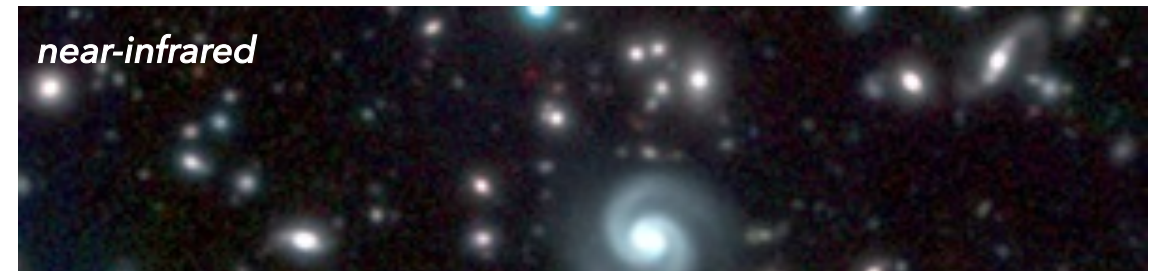
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➤ Two photometry codes:



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➤ Two photometry codes:



CLASSIC

{SExtractor & IRACLEAN}



THE FARMER

{Weaver et al., in prep}

➤ Uses SExtractor for detection + photometry

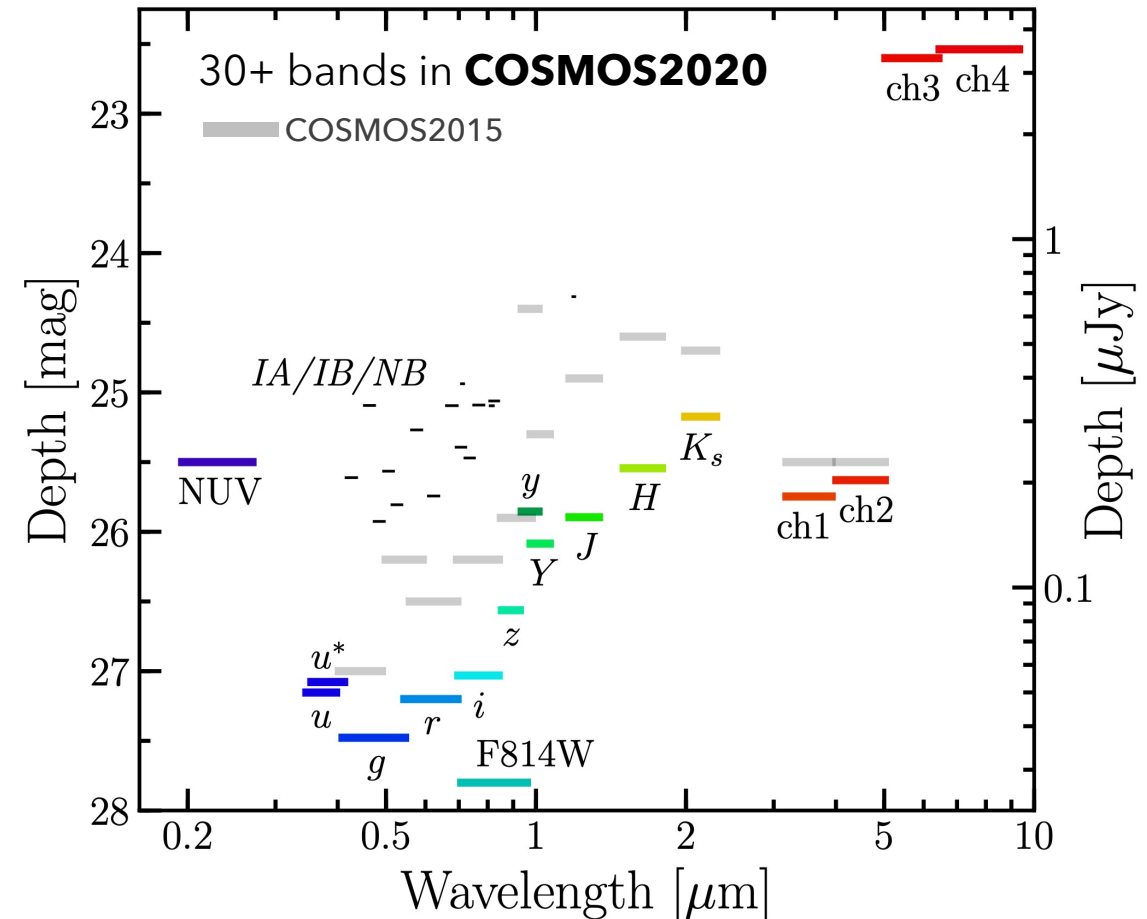
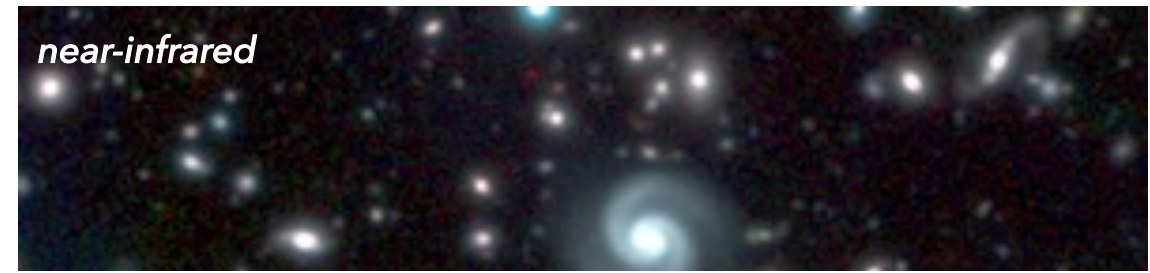
{Bertin & Arnouts, 1996}

➤ IRACLEAN for IRAC [3.6μm] [4.5μm]

{Hsieh et al. 2012}

➤ Same strategy as COSMOS2015

{Laigle et al. 2016}



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➤ Two photometry codes:



➤ Utilize parametric models to fit sources: The Tractor

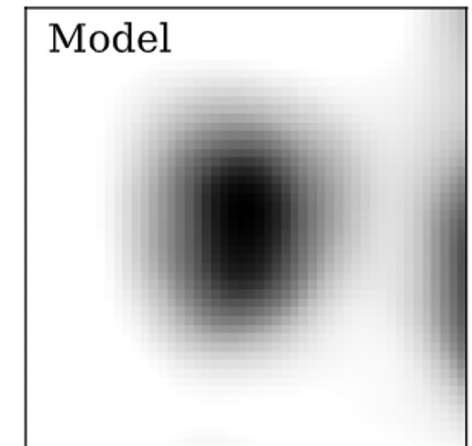
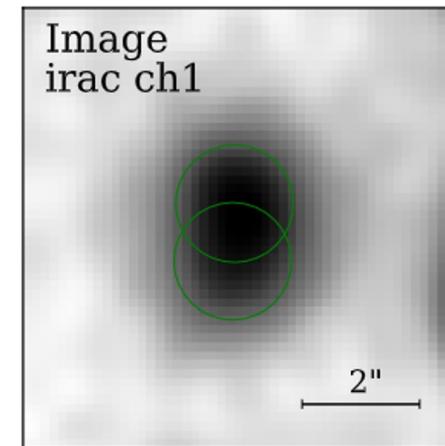
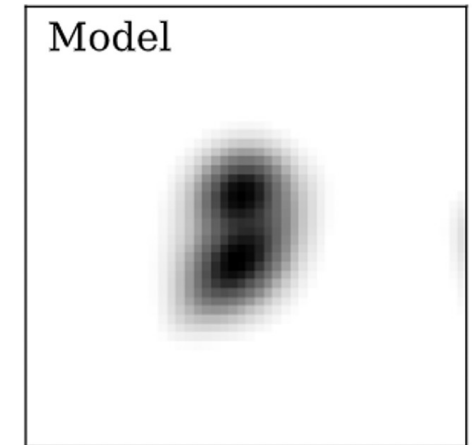
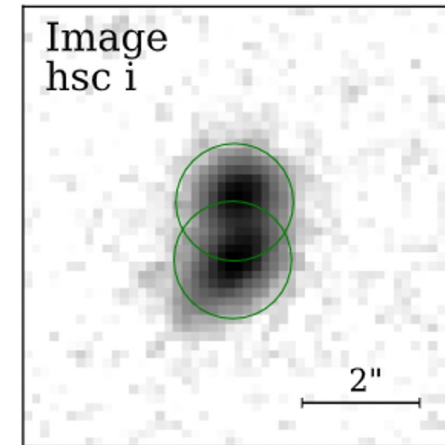
{Lang et al. 2016ab}

➤ Flux and position are now *model parameters*

- ✓ Sensitivity to ultra-faint sources
- ✓ Precise de-blending
- ✓ Free fitting + residual statistics, shapes, sizes

➤ Developed a scalable + reproducible framework

THE FARMER {Weaver et al., in prep}



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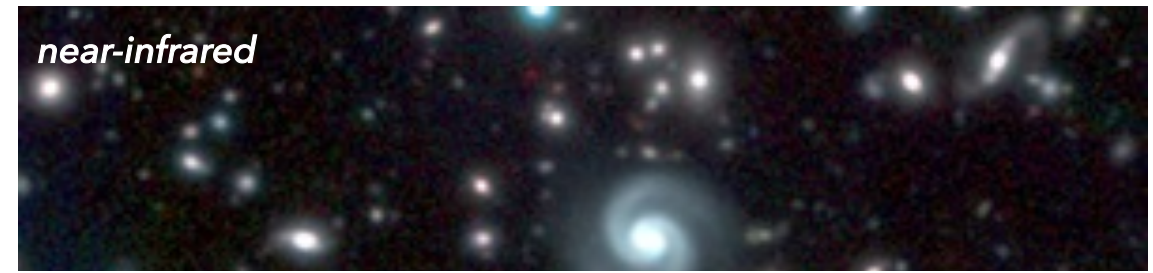
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➤ Two photometry codes:



➤ Each paired with both two photo-z codes:

Le Phare
{Ilbert et al. 2006}

EAzY
{Brammer et al. 2008}



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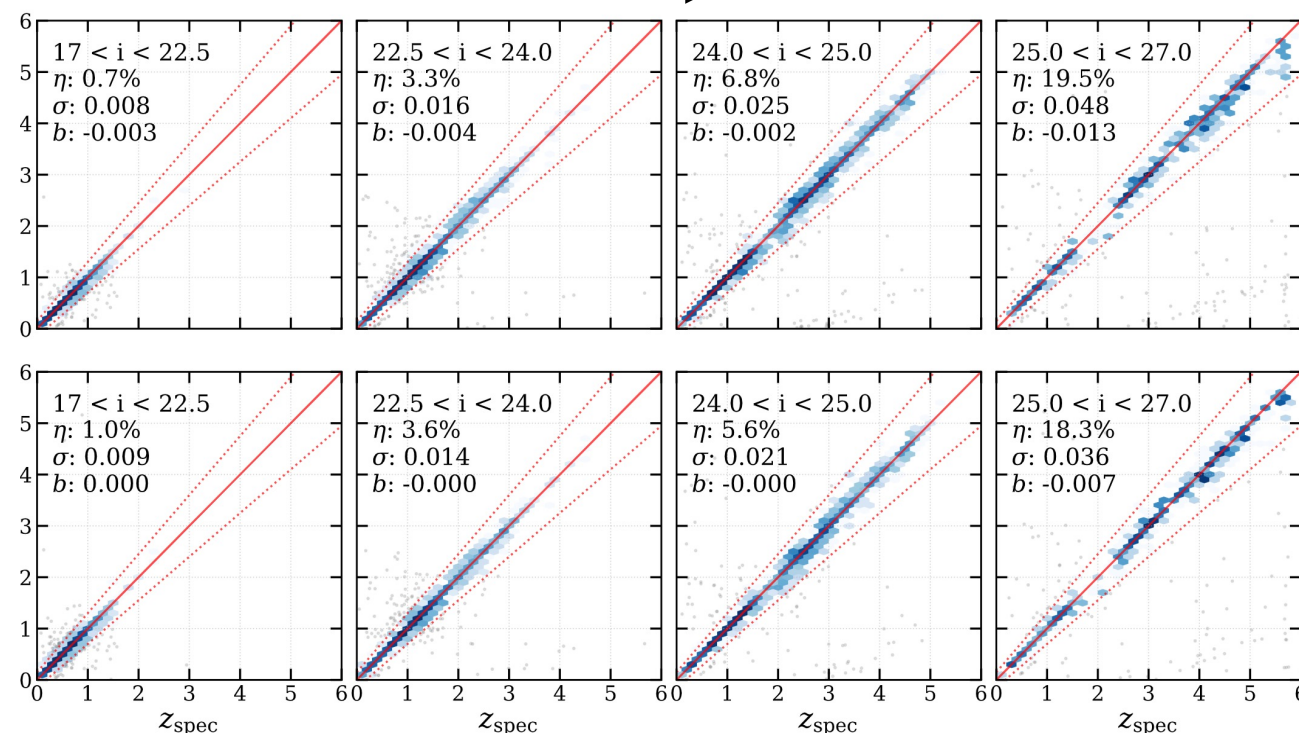
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η – outlier fraction
 σ – standard deviation
 b – median bias

Classic {apertures}



The Farmer {models}



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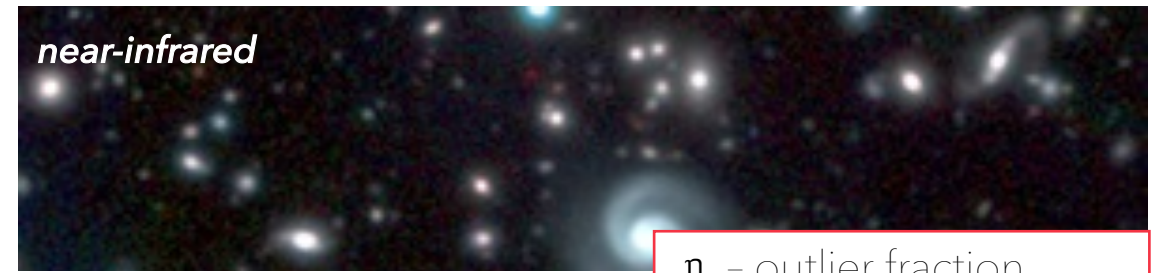
EAzY

{Brammer et al. 2008}

➤ Unprecedented photometric accuracy

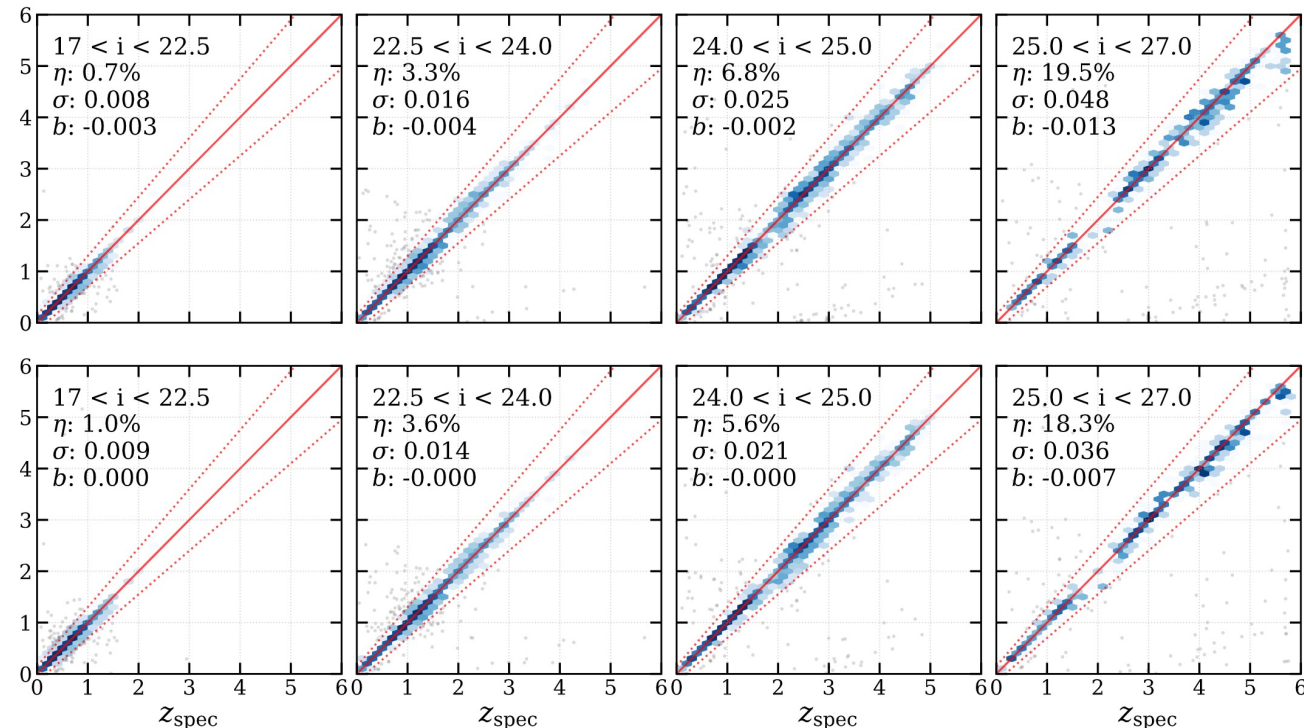
$\sigma < 1\%$ at $i < 22.5$ AB; $< 5\%$ at ~ 26 AB

➤ Low bias, low failure rate



η – outlier fraction
 σ – standard deviation
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Classic {apertures}



The Farmer {models}



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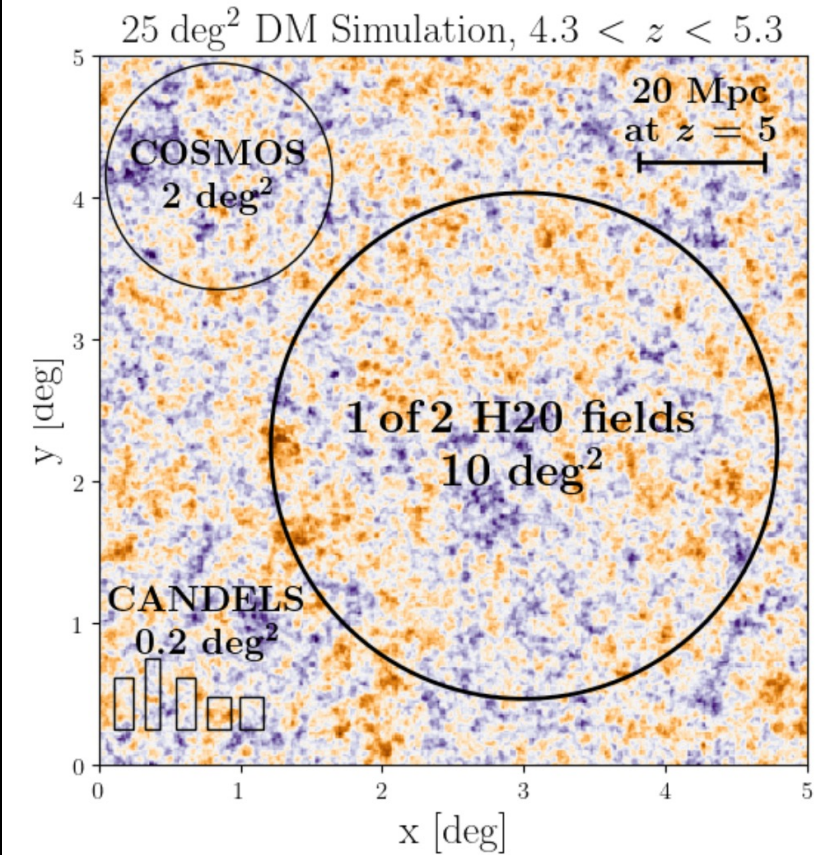
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Environments dense enough to support the most massive galaxies
are found in only the largest & deepest surveys → **COSMOS2020**



Simulation of The Epoch of Reionization
{M. Alvarez et al.}

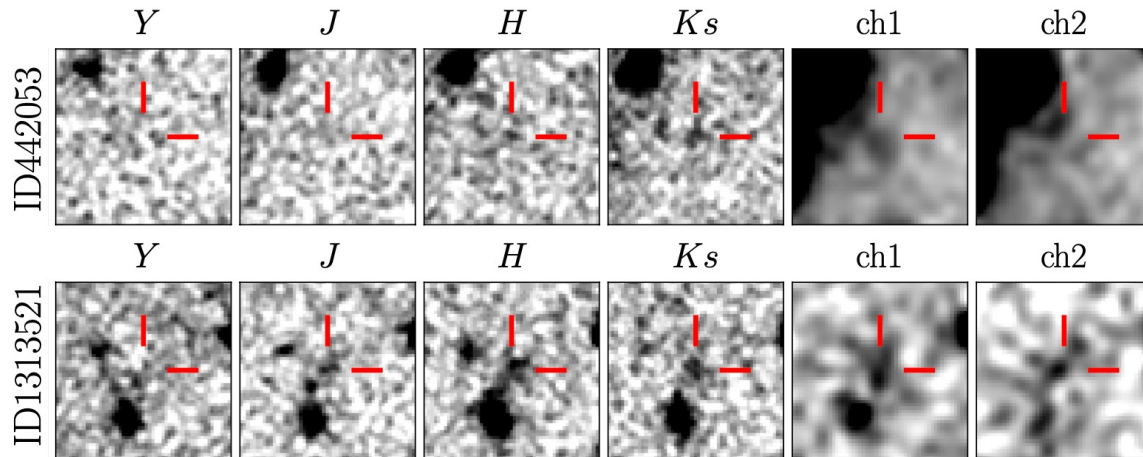


The UV Luminosity Function

New constraints on the first ultra-luminous galaxies

{Kauffmann, Ilbert, Weaver et al., in prep.}

- 31 galaxy candidates @ $z > 7.5$ over 0.8 deg^2
- New sources only found with **THE FARMER**
Not selected in apertures due to crowding

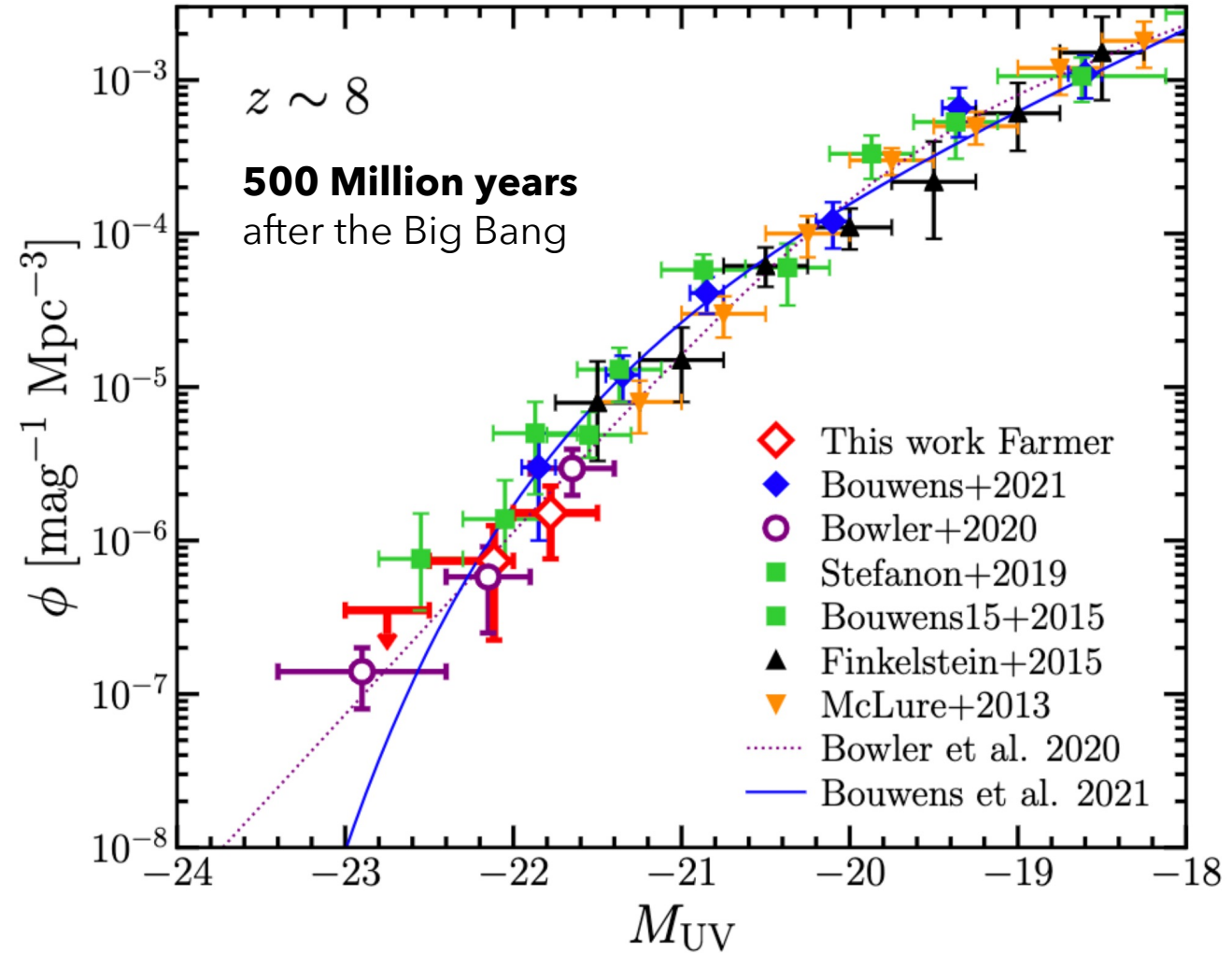
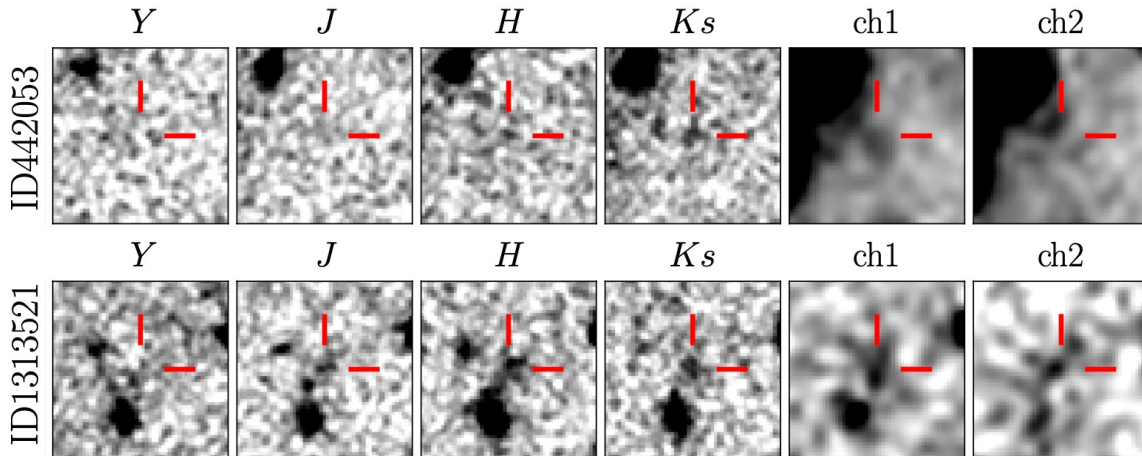


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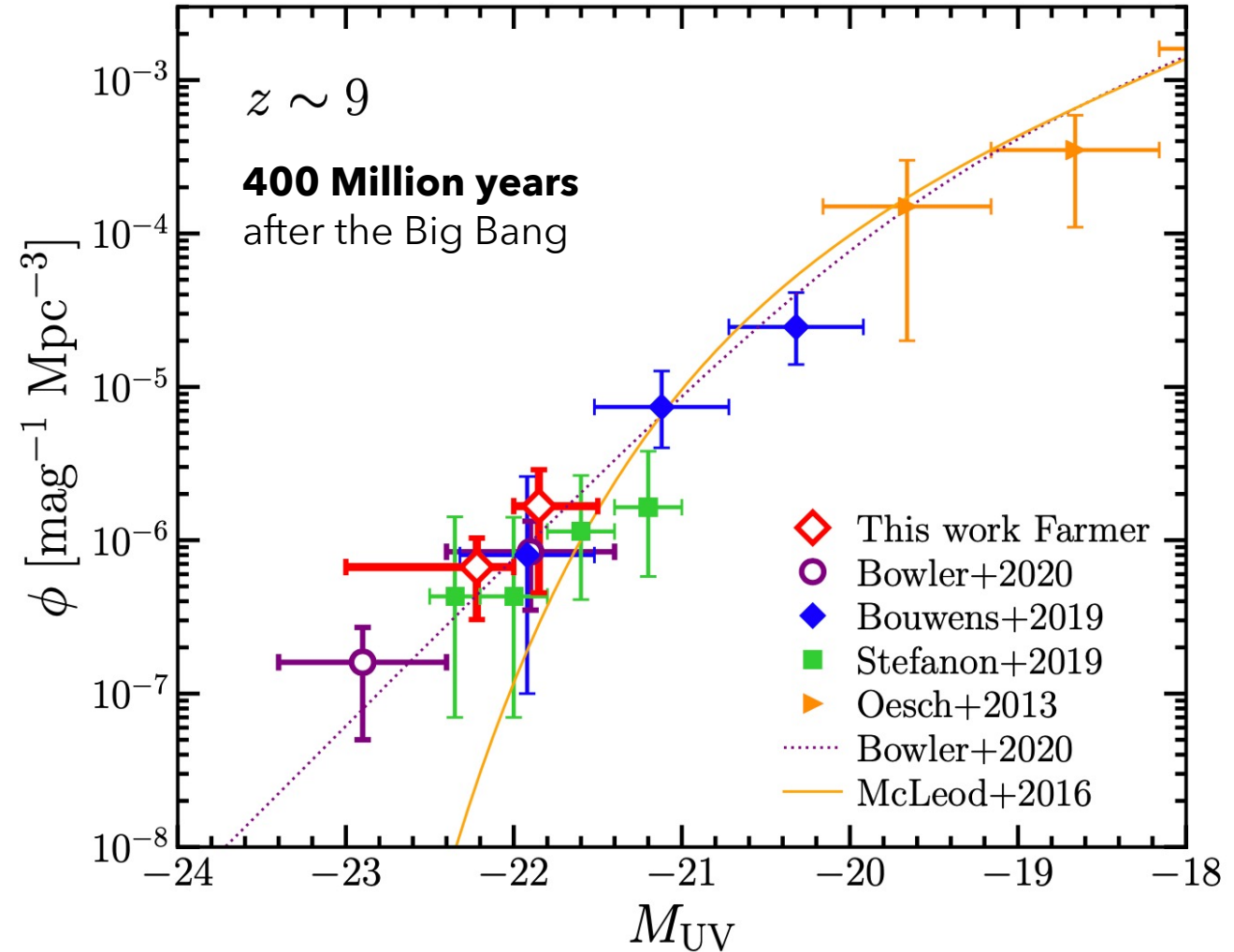


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We find an excess of $M_{\text{UV}} < -21.5$ sources

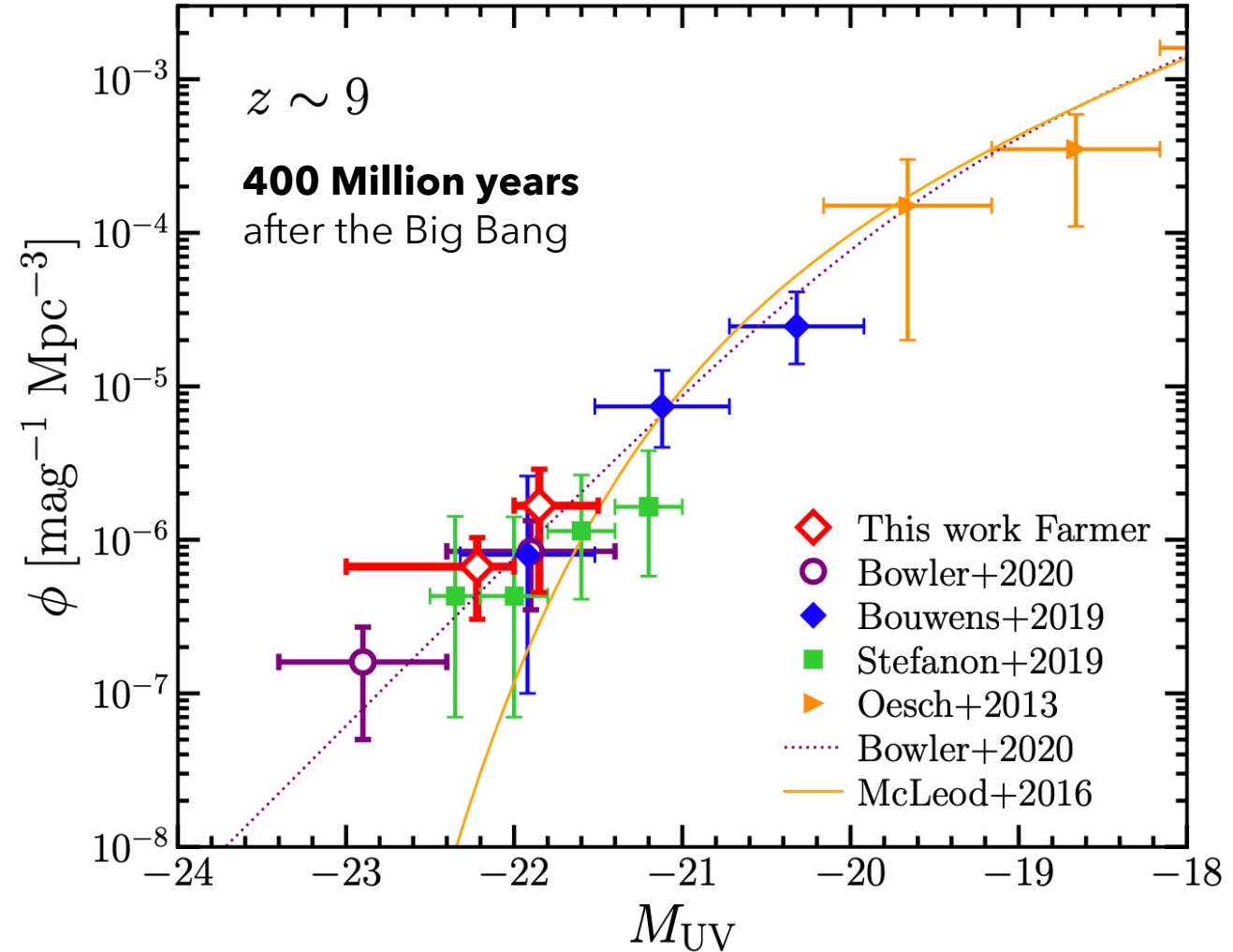
{Consistent with previous work; e.g., Stefanon et al. 2019, Bowler et al. 2020}

Could **quenching** have not begun yet?

{halo quenching; Peng et al. 2010}

Or does **dust evolution** complicate this picture?

{e.g., Bowler et al. 2015, Finkelstein et al. 2015}



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The UV Luminosity Function

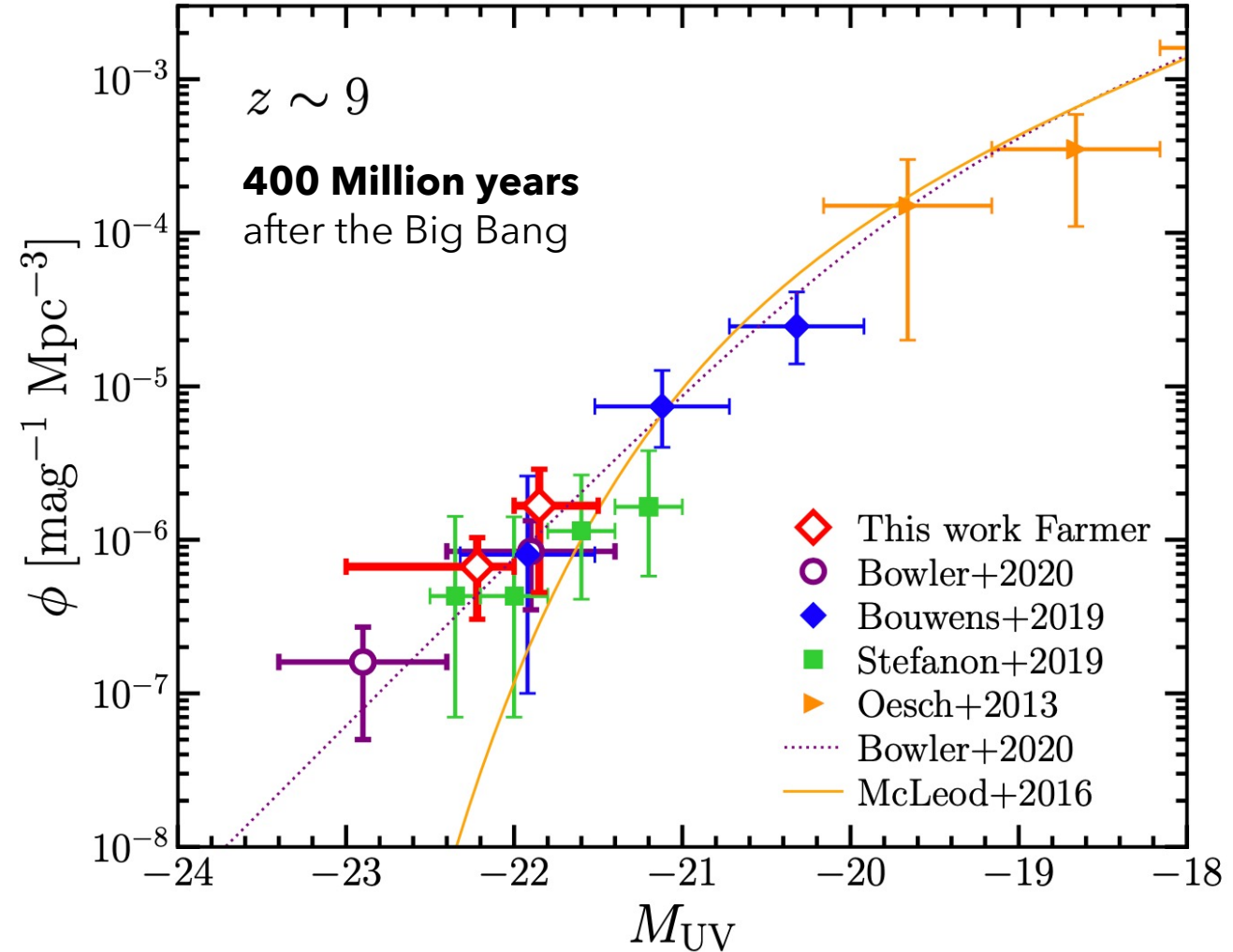
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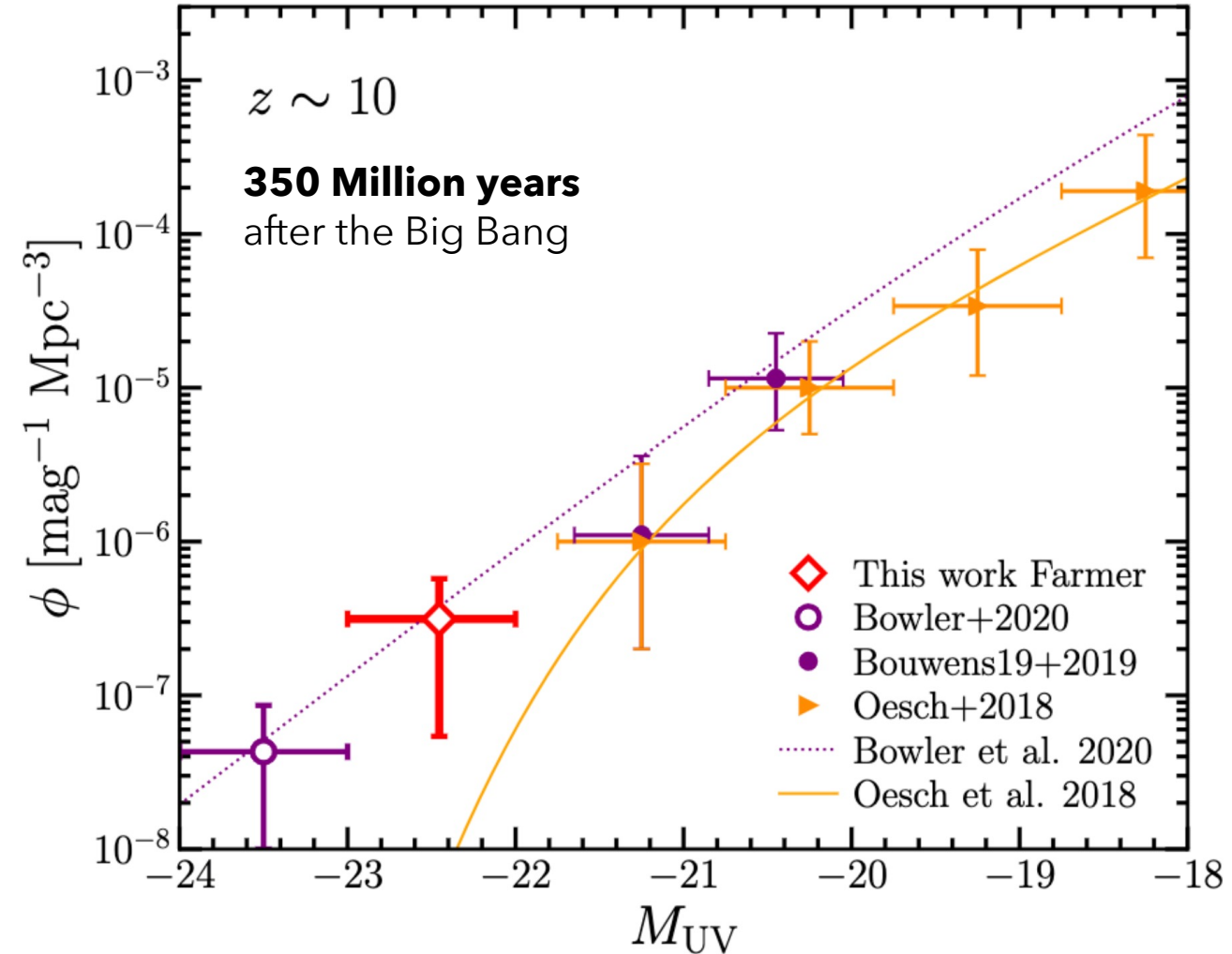
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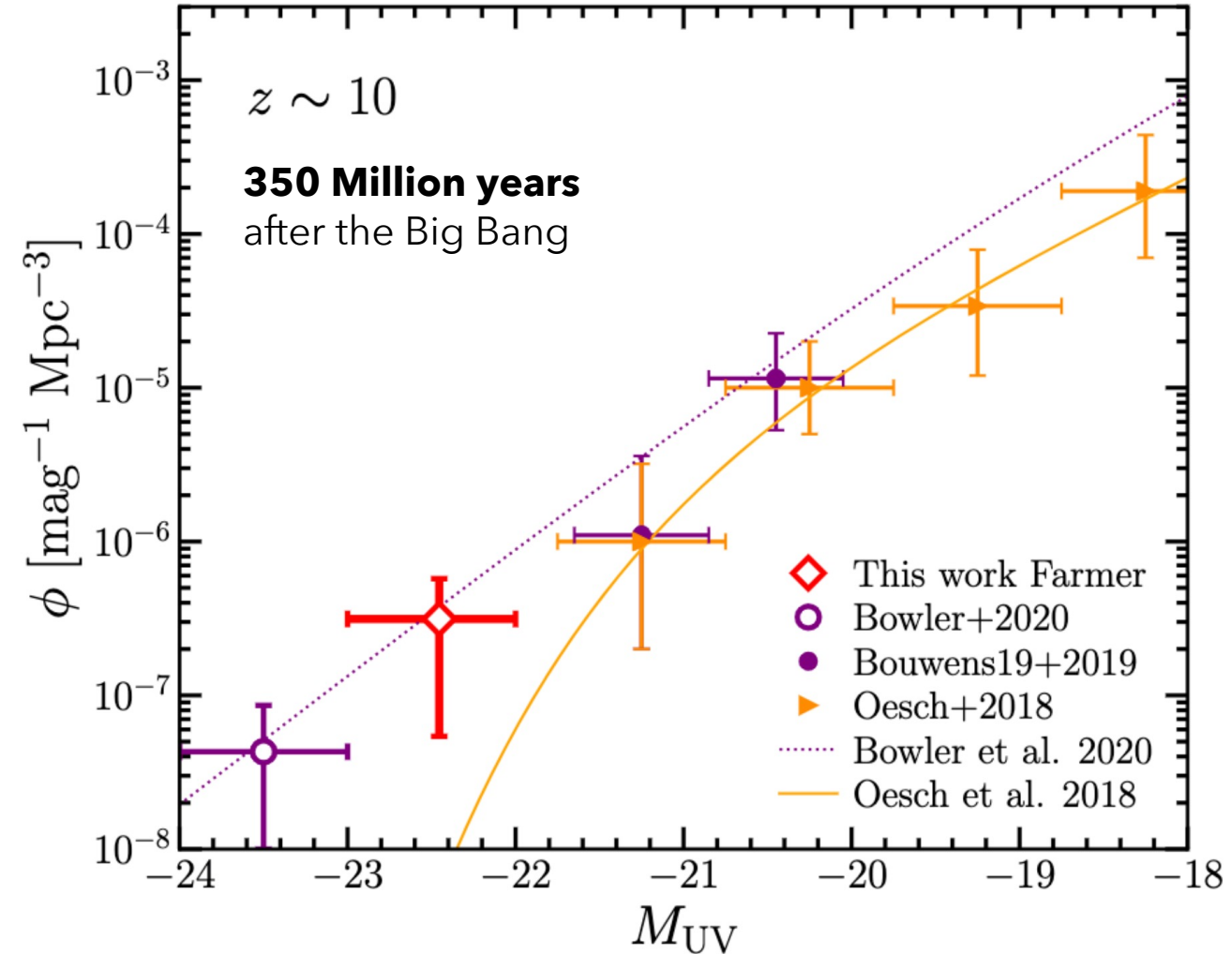
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{Consistent with previous work; e.g., Stefanon et al. 2019, Bowler et al. 2020}

If SFRE at $z \sim 10$ is similar to $z \sim 9$,
we **should not find** such bright galaxies here.



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The UV Luminosity Function

New constraints on the first ultra-luminous galaxies

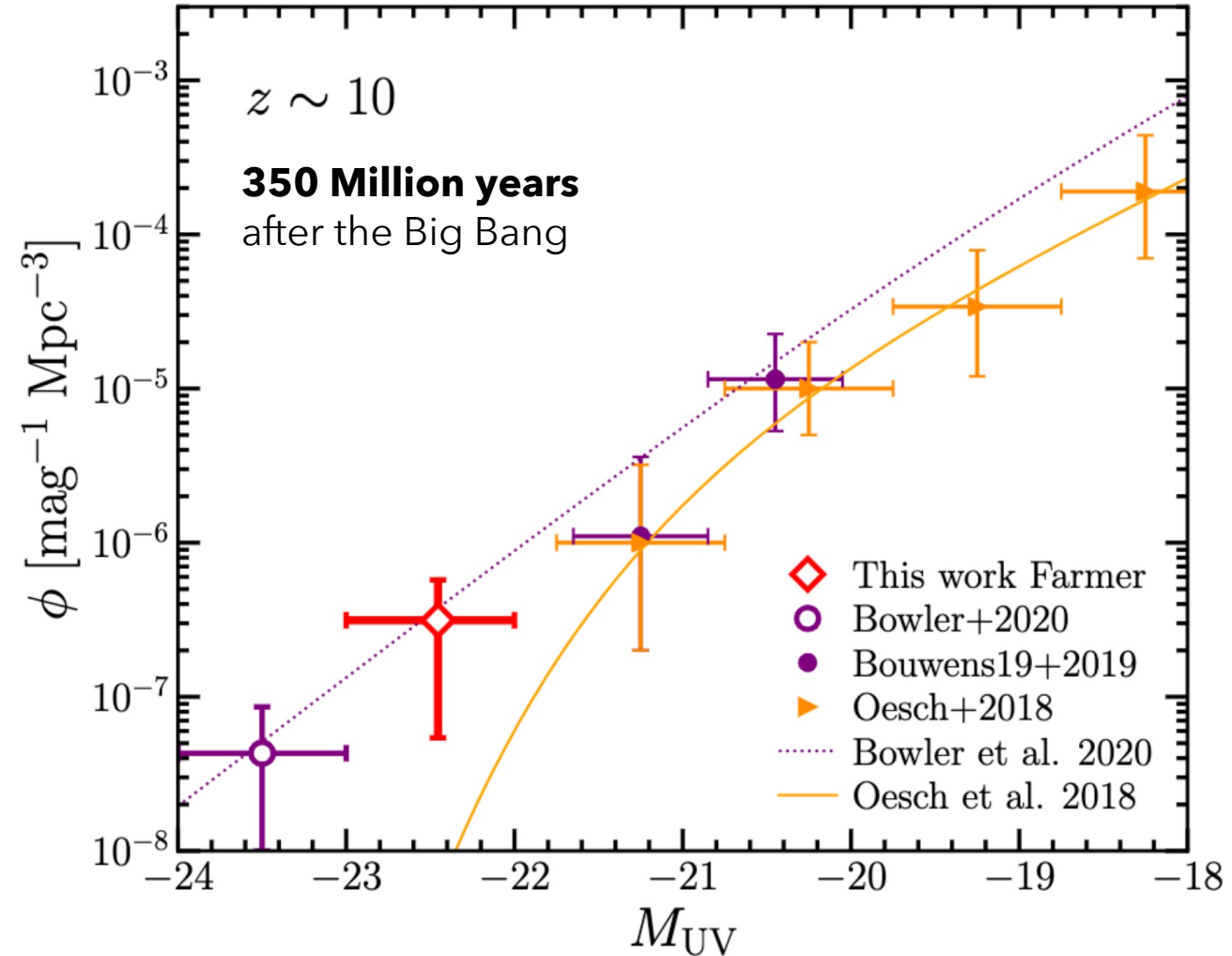
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While **lensing** estimates are too weak at $z \sim 8-9$
it may contribute to the lack of evolution at $z \sim 10$
in addition to **contamination**

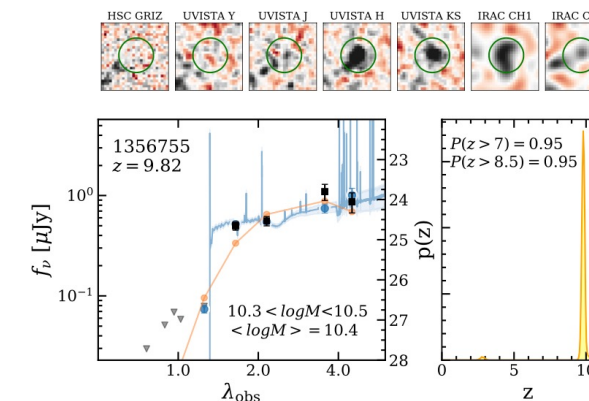
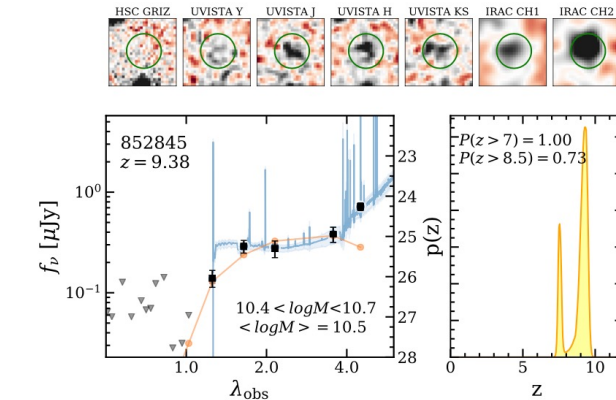
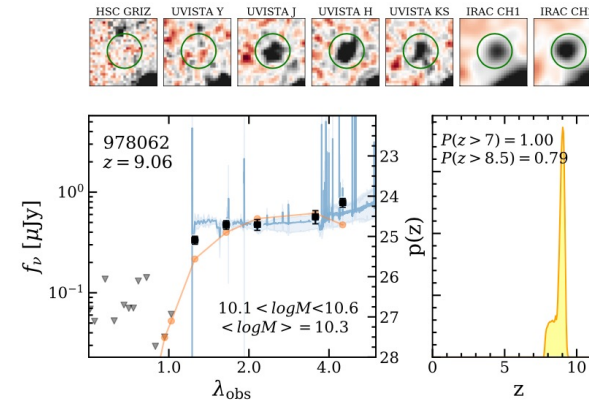
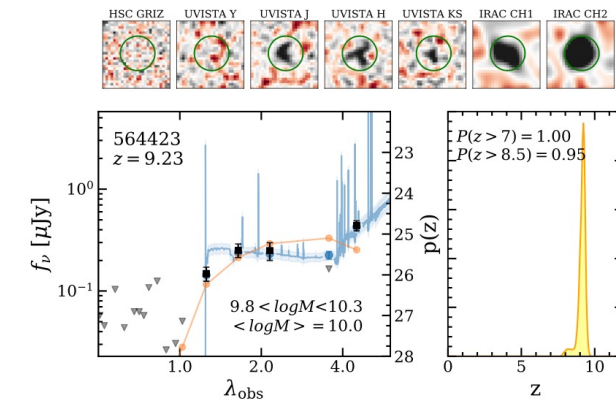
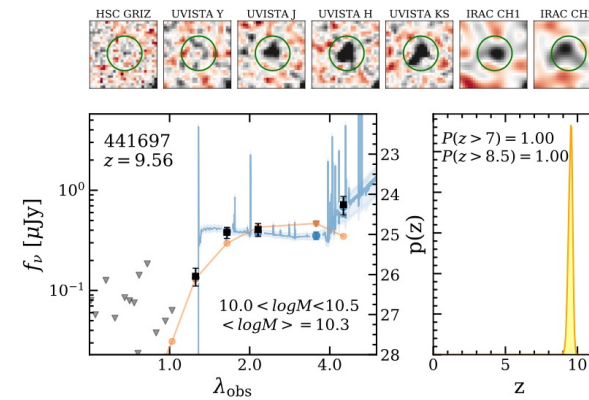


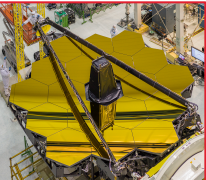
"Beasts" Cycle 1 JWST Program

{PI: Weaver, PID 2659}

Detailed spectroscopic study of $z \sim 9$ galaxies

- **Five** robust, *ultra*-luminous galaxies ($M_{UV} \sim -22.5$)
- Will not be seen in smaller surveys (e.g. JADES)
- Characterize their mass assembly and star-formation



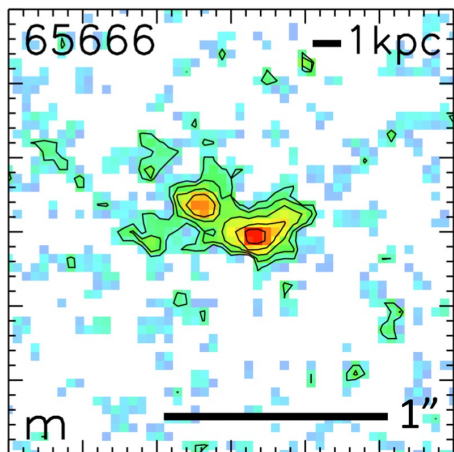


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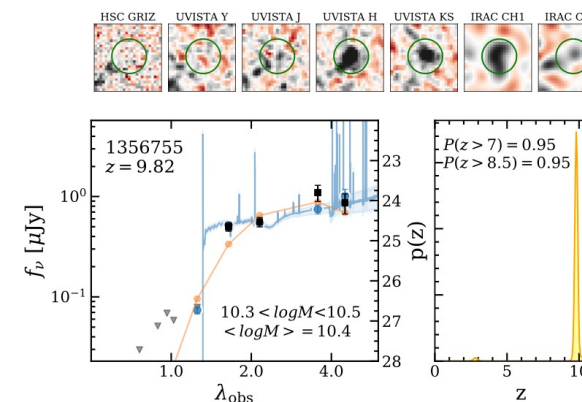
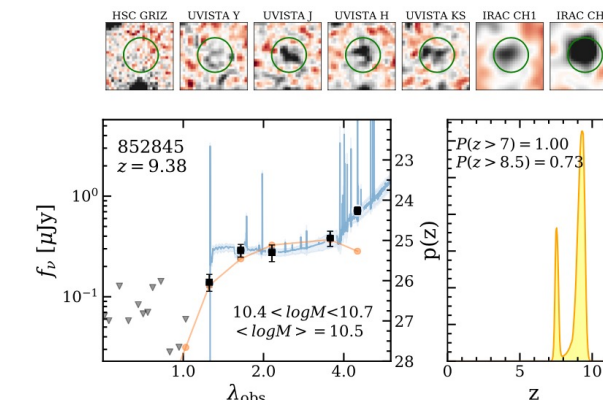
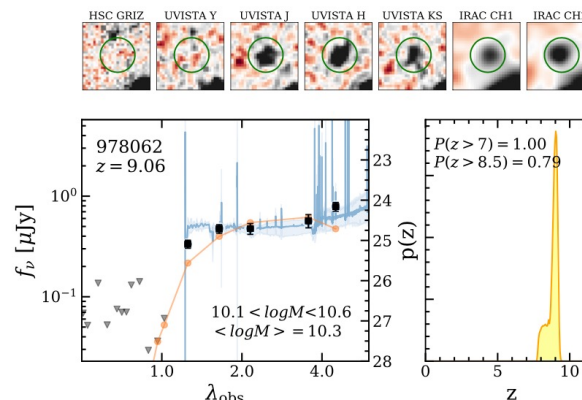
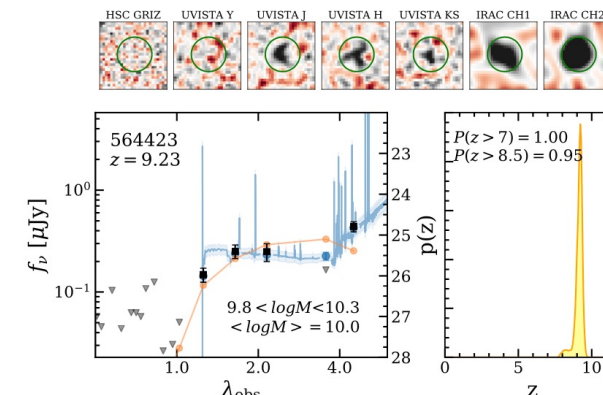
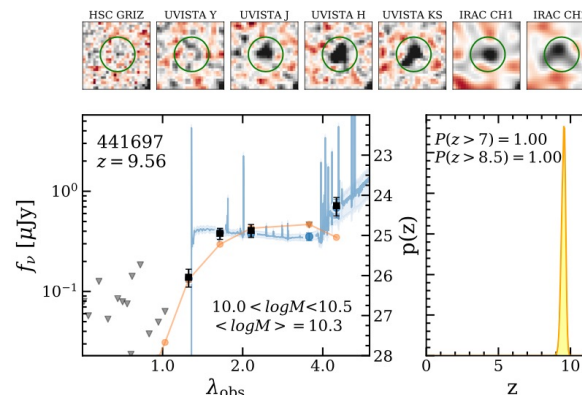
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- Will not be seen in smaller surveys (e.g. JADES)
- Characterize their mass assembly and star-formation
- Deep ($Y>26$) spatially resolved spectra; CLEAR/PRISM



$z \sim 7$; Bowler et al. 2018



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A 50 deg² Survey to map the high-redshift universe

<https://dawn.ipac.caltech.edu/>

- Includes 20 deg² of the Euclid Deep Field North and South

Deep near-infrared selection

- Optical imaging underway from the Hawaii 2-0 Program

{McPartland, ... Weaver et al., in prep.} {Zalesky, ... Weaver et al., in prep.}

- Supported by largest Spitzer mission ever

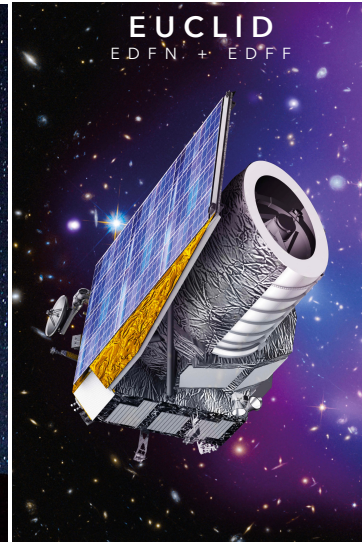
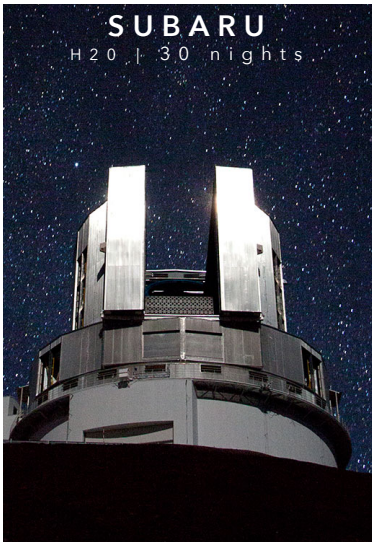
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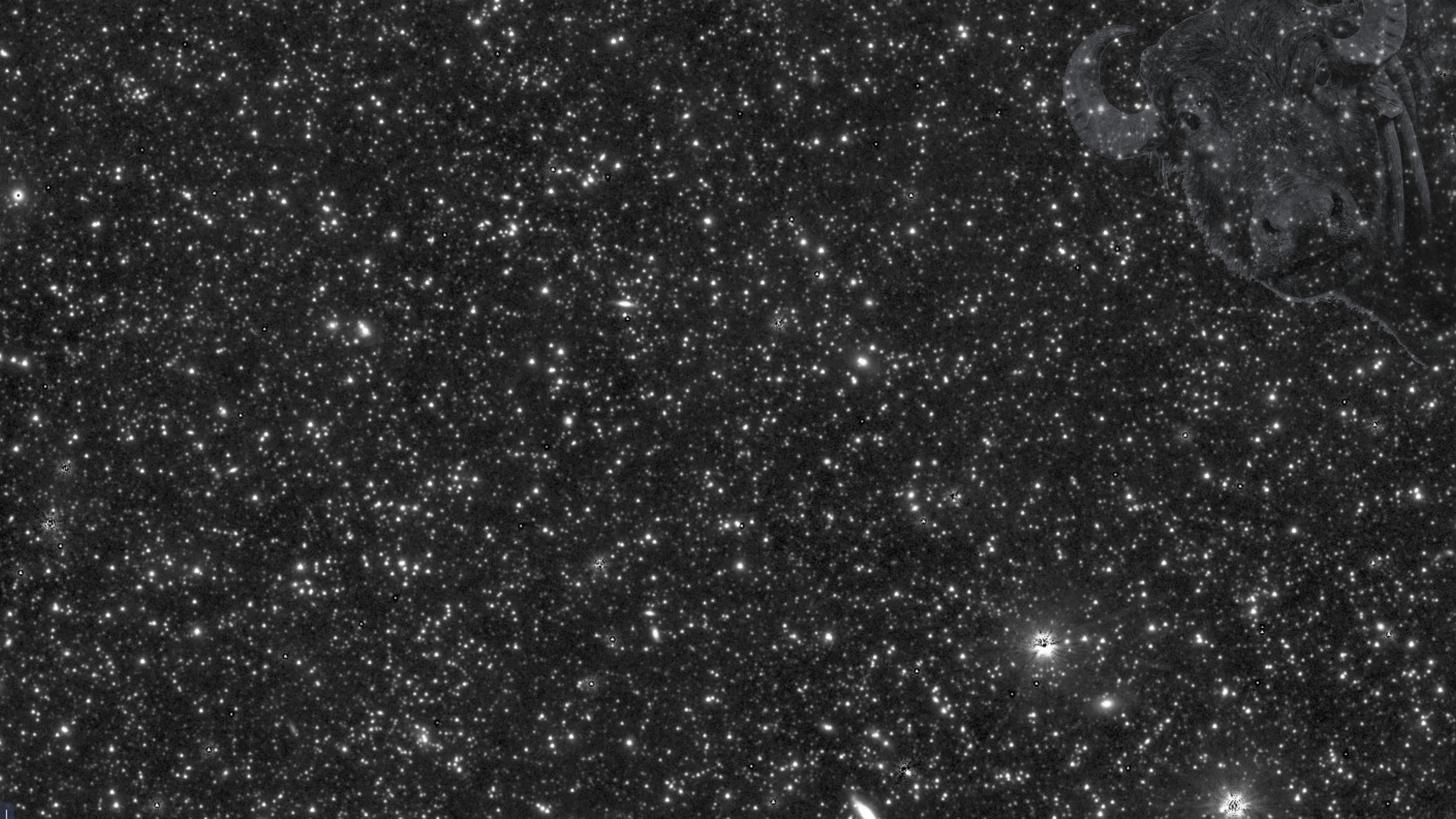
A view into the deep universe

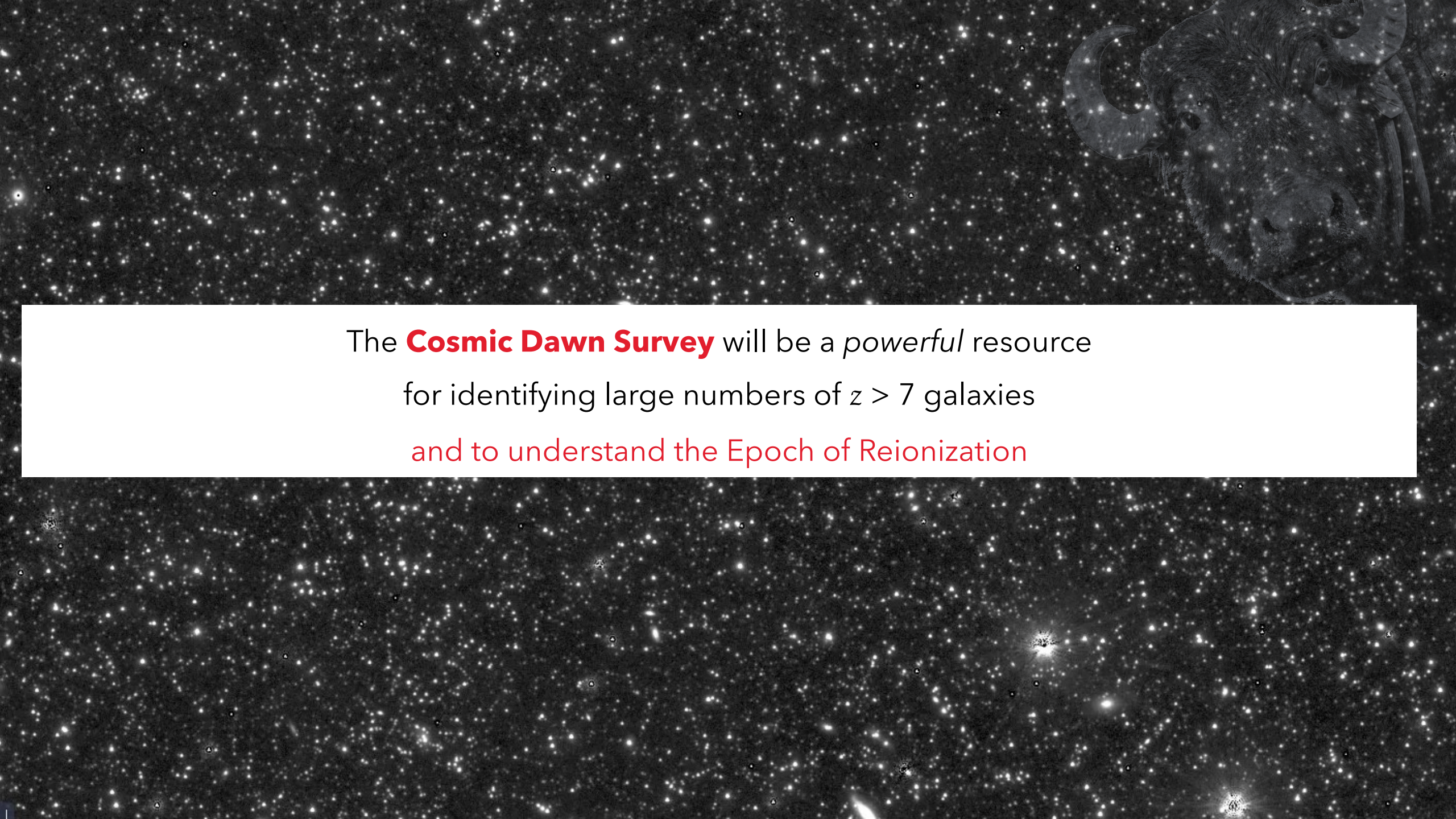
Spitzer/IRAC 3.6um @ the North Ecliptic Pole



Euclid FOV





The background of the slide is a deep space image filled with numerous stars of varying brightness. In the upper right corner, there is a detailed, dark-toned illustration of a bison's head, facing forward, which appears to be superimposed on the starry background. The bison has large, curved horns and a thick, shaggy coat.

The **Cosmic Dawn Survey** will be a *powerful* resource
for identifying large numbers of $z > 7$ galaxies
and to understand the Epoch of Reionization

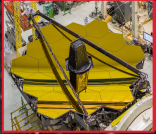


Interested in working together?
I'm on the market. **Get in touch!**

john.weaver@nbi.ku.dk
or find me on slack

**COSMIC
DAWN
SURVEY**

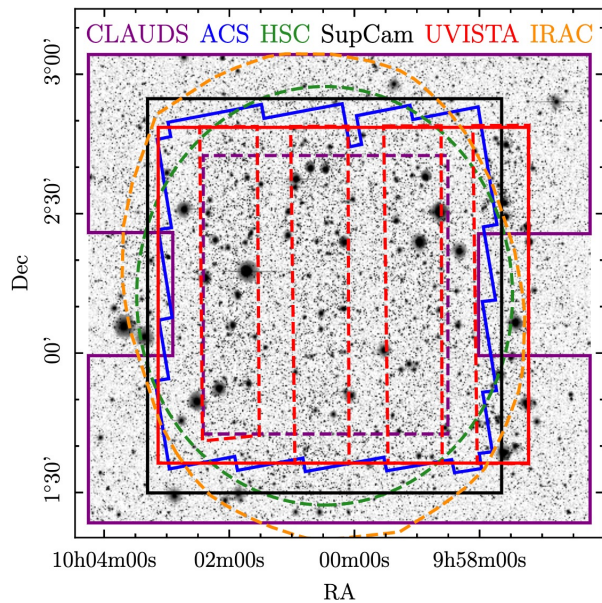
International collaborations
Euclid | Subaru | Keck | Spitzer
{Moneti, ... Weaver et al., subm.}
{McPartland, ... Weaver et al., in prep.}
{Zalesky, Weaver et al., in prep.}



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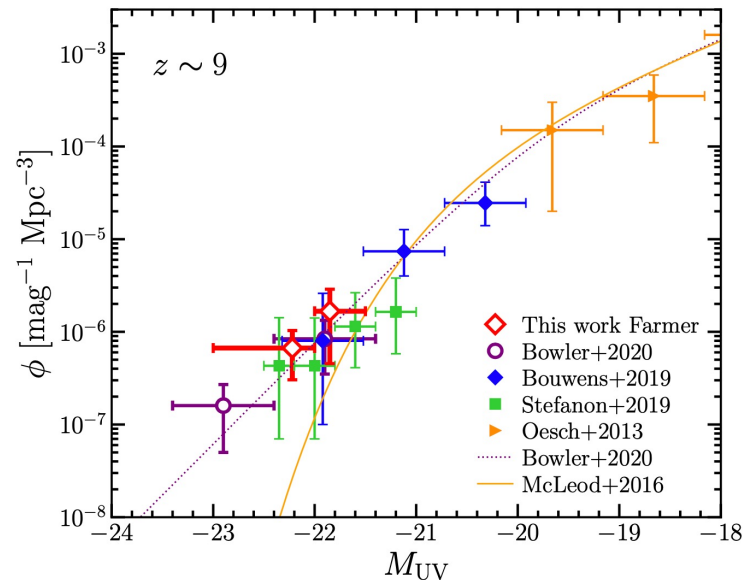
COSMOS2020

{Weaver et al., subm.}



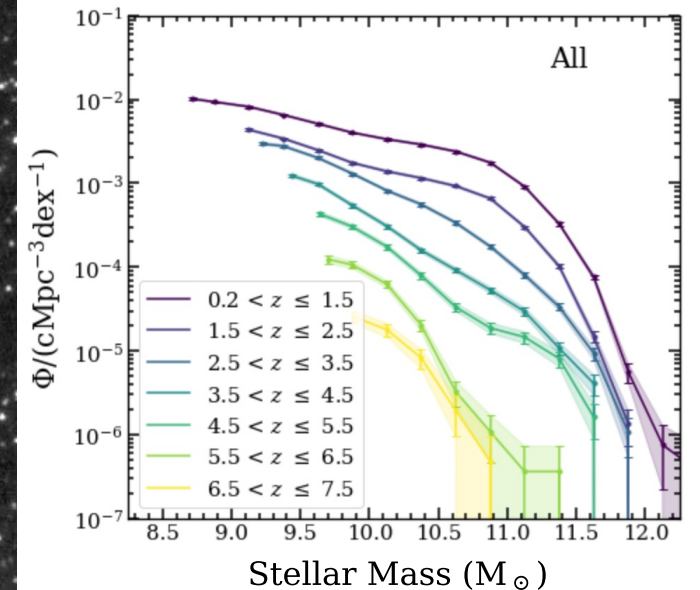
UV Luminosity Function

{Kauffmann, ... Weaver et al., in prep.}



Stellar Mass Function

{Weaver et al., in prep.}



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PhD Fellow @ Cosmic Dawn Center, Univ. of Copenhagen
john.weaver@nbi.ku.dk | astroweaver.github.io | @astroweave



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