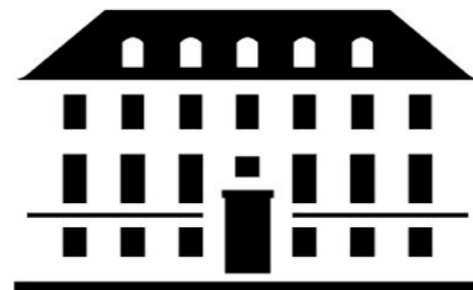


# Theoretical Astrophysics



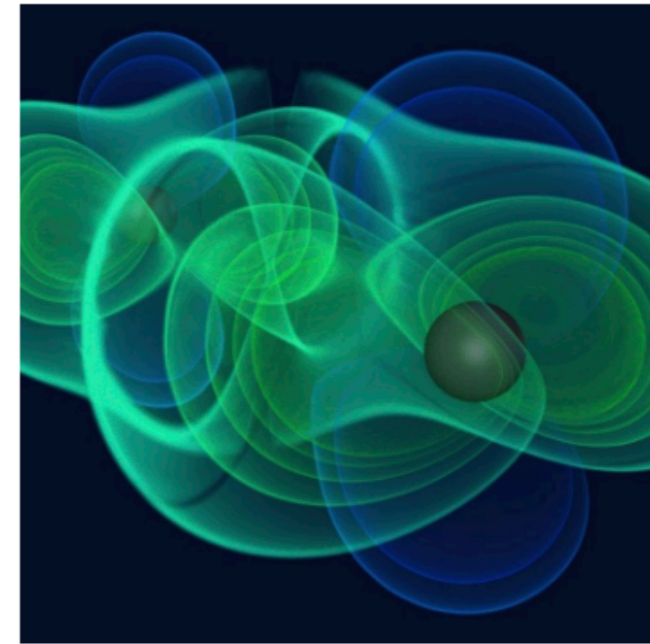
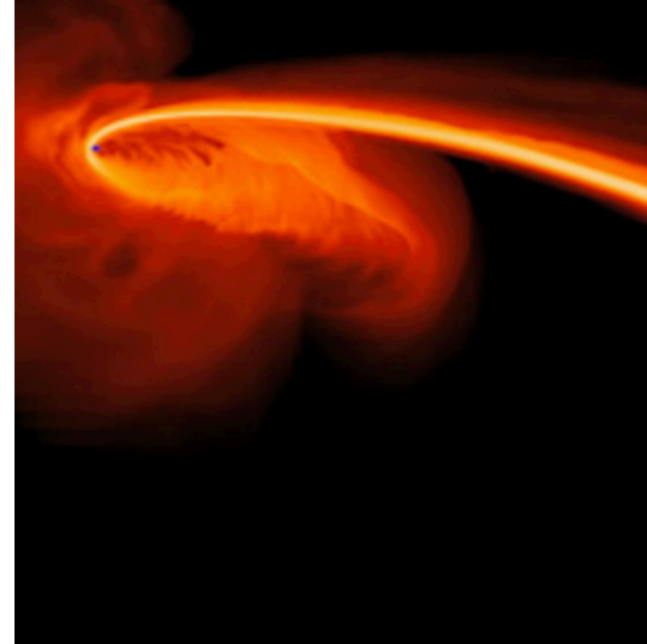
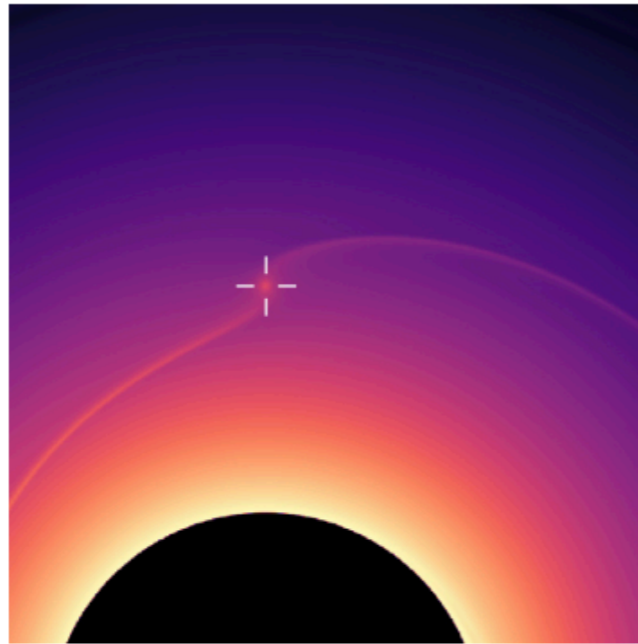
The Niels Bohr  
International Academy

Martin Pessah

[mpessah@nbi.dk](mailto:mpessah@nbi.dk)

<https://sites.google.com/view/martin-pessah>

# Astrophysical Dynamics



Astrophysical  
Magnetohydrodynamics

Protoplanetary Disks &  
Planet Formation

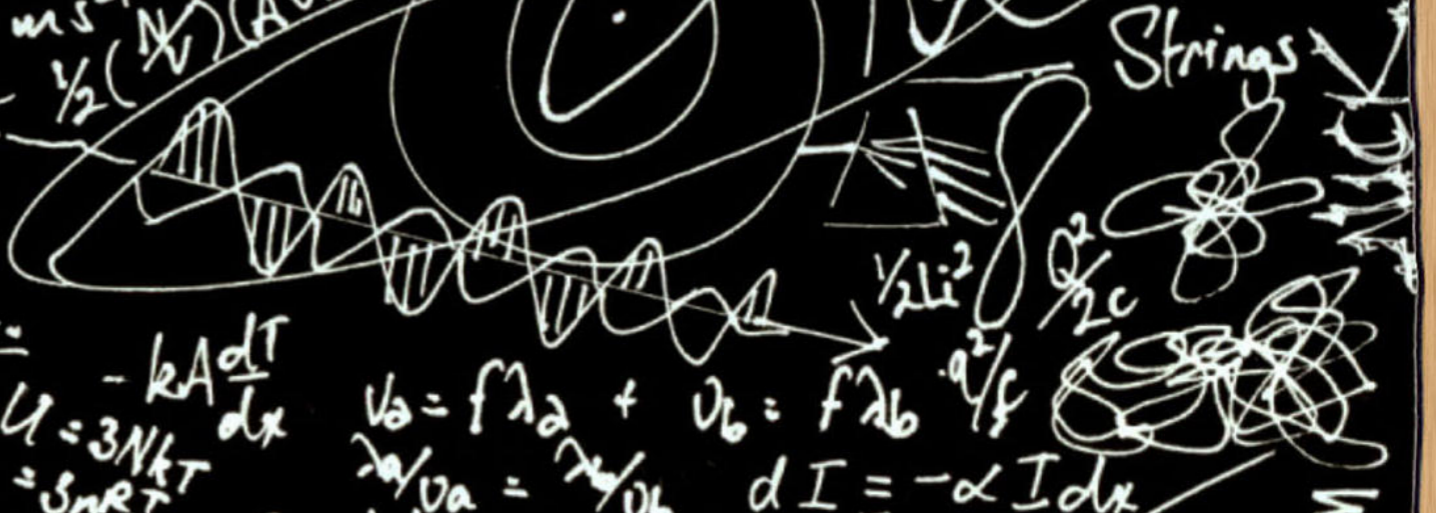
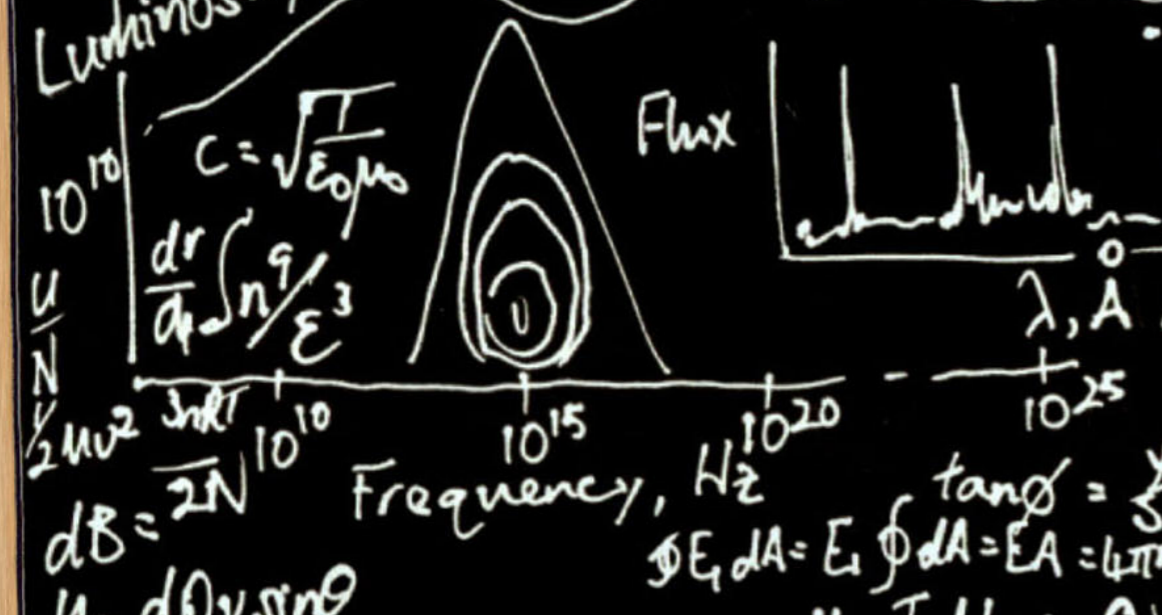
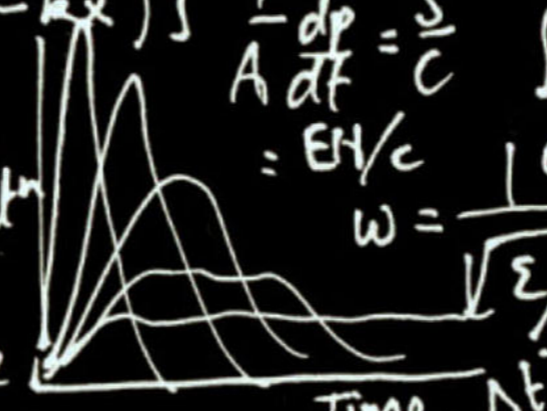
Tidal Disruptions Events

Gravitational Wave  
Astrophysics

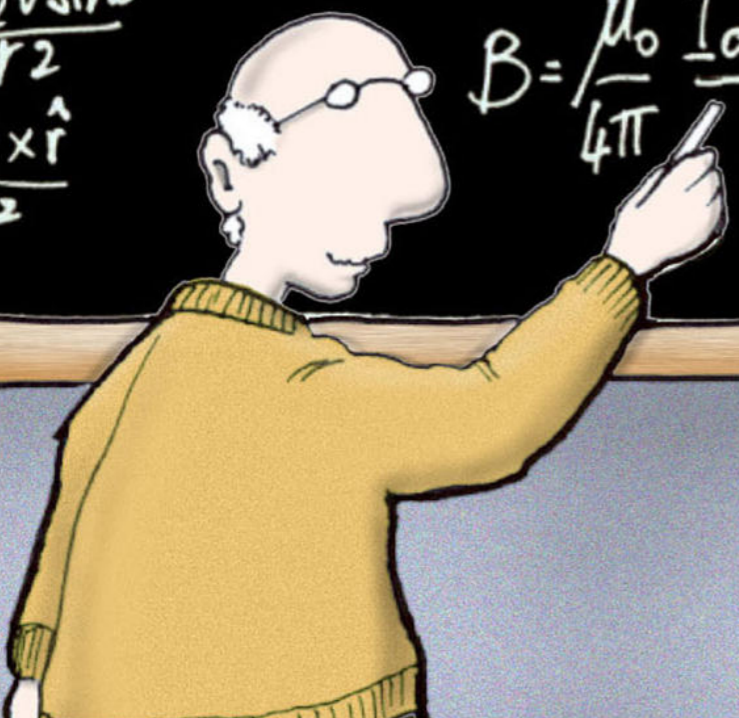
In a nutshell, we care to understand the dynamics of various astrophysical systems and how we can use observations to learn about the physical properties of planets, stars, and black holes.

How?

$E = E_{max} [-\sin(\omega t + kx) + \sin(\omega t - kx)]$   
 $H = H_{max} [\sin(\omega t + kx) + \sin(\omega t - kx)]$   
 $H = \frac{2\pi kL(T_2 - T_1)}{\ln(b/a)}$   
 $E = -2E_{max} \cos \omega t \sin kx$   
 $\frac{1}{A} \frac{dp}{dt} = \frac{S}{c}$   
 $\int \frac{dr}{r} = - \int \frac{2\pi kL}{H} dt$   
 $\omega = \frac{1}{\sqrt{\epsilon\mu}} = \frac{1}{\sqrt{\epsilon_0\mu_0}}$   
 $x = A \cos(\sqrt{\epsilon/\mu} t) = A \cos(\omega t)$   
 $\frac{c}{\omega} = n$   
 $\sin \phi_1 = \frac{n_0}{n_b} \sin \phi_2$   
 $\sin \phi_{crit} = \frac{n_b}{n_a}$



$\frac{dB}{dB} = \frac{2N}{2N}$   
 $\frac{1}{4\pi r^2} \int I d\Omega \sin \theta$   
 $\Phi_E dA = E_1 \oint dA = EA = 4\pi r^2 E$   
 $B = \frac{\mu_0 I dl \sin \theta}{4\pi r^2}$   
 $\frac{1}{2} \frac{d\phi_2}{dt} = M \frac{di_1}{dt}$   
 $E_2 = M \frac{di_1}{dt}$   
 $\frac{1}{\infty} + \frac{1}{s'} = \frac{2}{R}$   
 $s' = R/2$   
 $\frac{4\mu_0 L}{\pi} \int \frac{I(dy_j)}{-L(y^2 + z^2)^{3/2}}$   
 $M = \frac{1}{2} L T^2 = 2 \left( \frac{\mu_0 N^2 A}{L} \right)$   
 Surface density



# Equations for a Magnetized Fluid

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

continuity

$$\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} = -\nabla \Phi - \frac{\nabla P}{\rho} + \frac{1}{\rho} \mathbf{J} \times \mathbf{B}$$

momentum

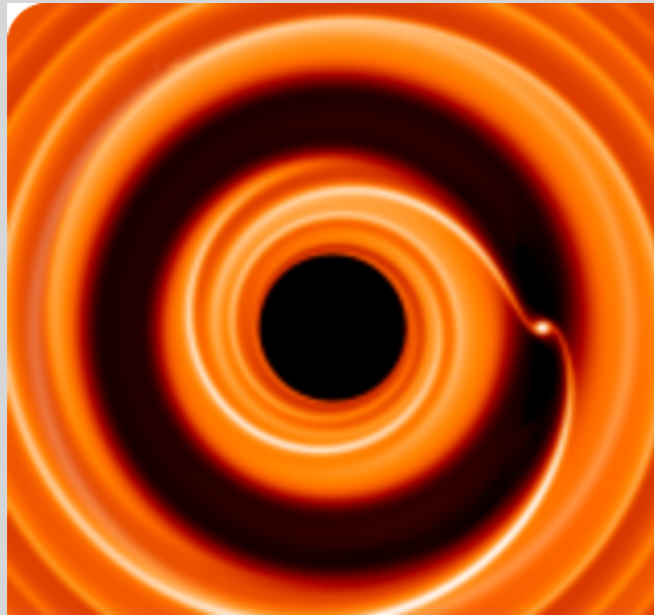
$$\frac{\partial \mathbf{B}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{B} = (\mathbf{B} \cdot \nabla) \mathbf{v} - \mathbf{B} (\nabla \cdot \mathbf{v})$$

induction

$$\frac{\partial e}{\partial t} + \nabla \cdot (e \mathbf{v}) = -P (\nabla \cdot \mathbf{v})$$

energy

# State-of-the-art Numerical Codes



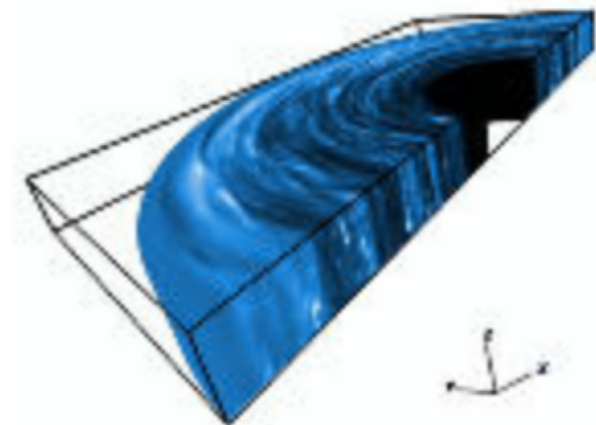
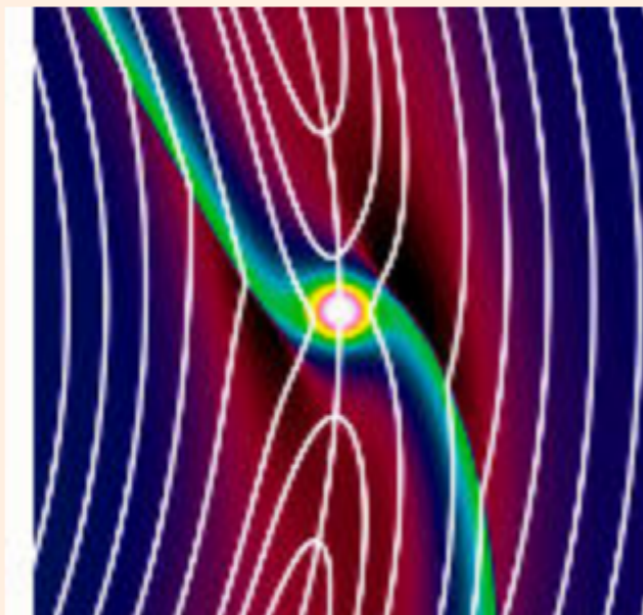
## FARGO3D

Benitez-Llambay & Masset (2016) FARGO3D is a very optimized code to study protoplanetary disks. <http://fargo.in2p3.fr/>



Pablo Benitez-Llambay

**A versatile HD/MHD code that runs on clusters of CPUs or GPUs, with special emphasis on protoplanetary disks.**



# Supercomputers



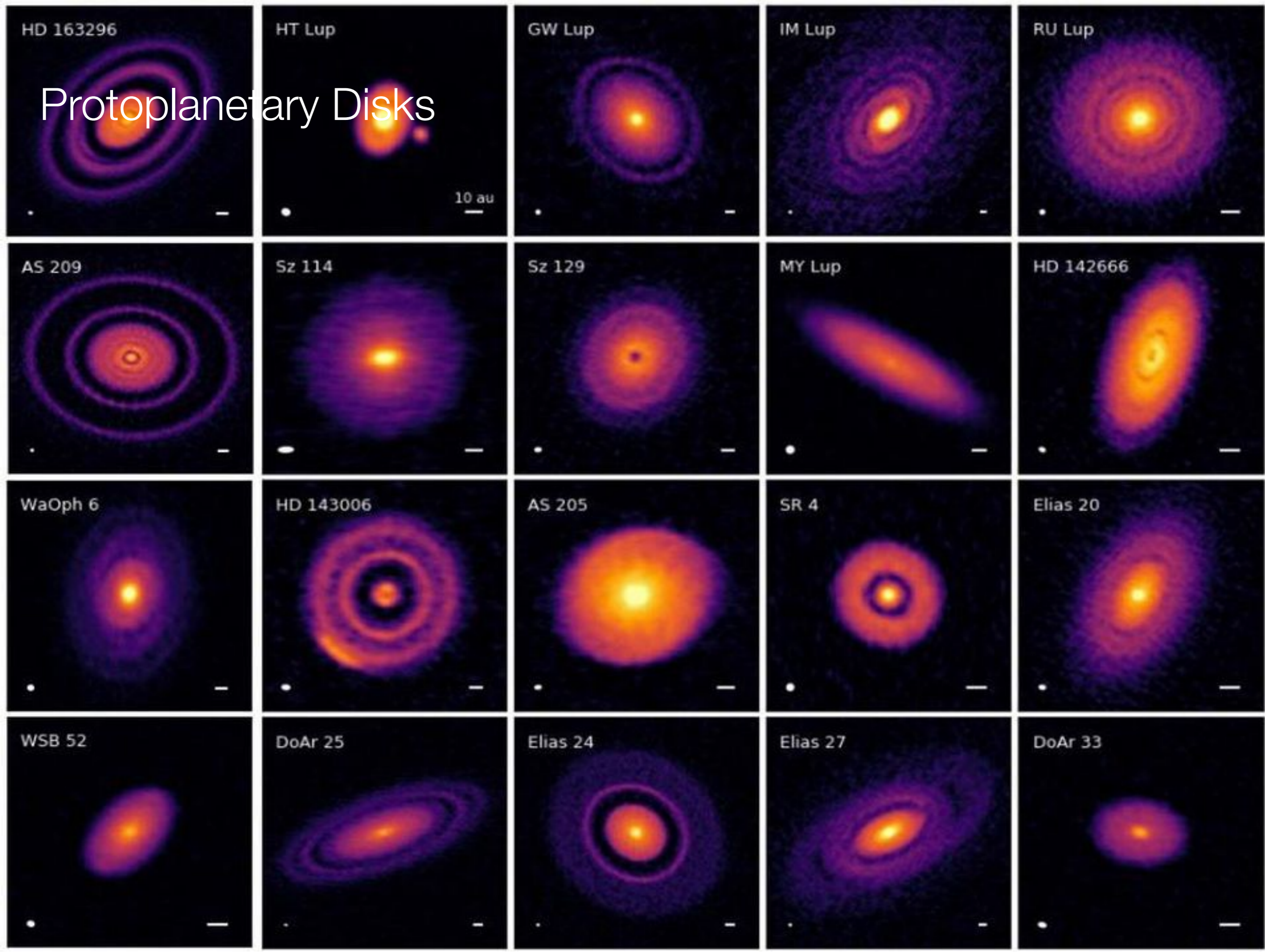
# Some Examples of Recent Projects on Protoplanetary Disks



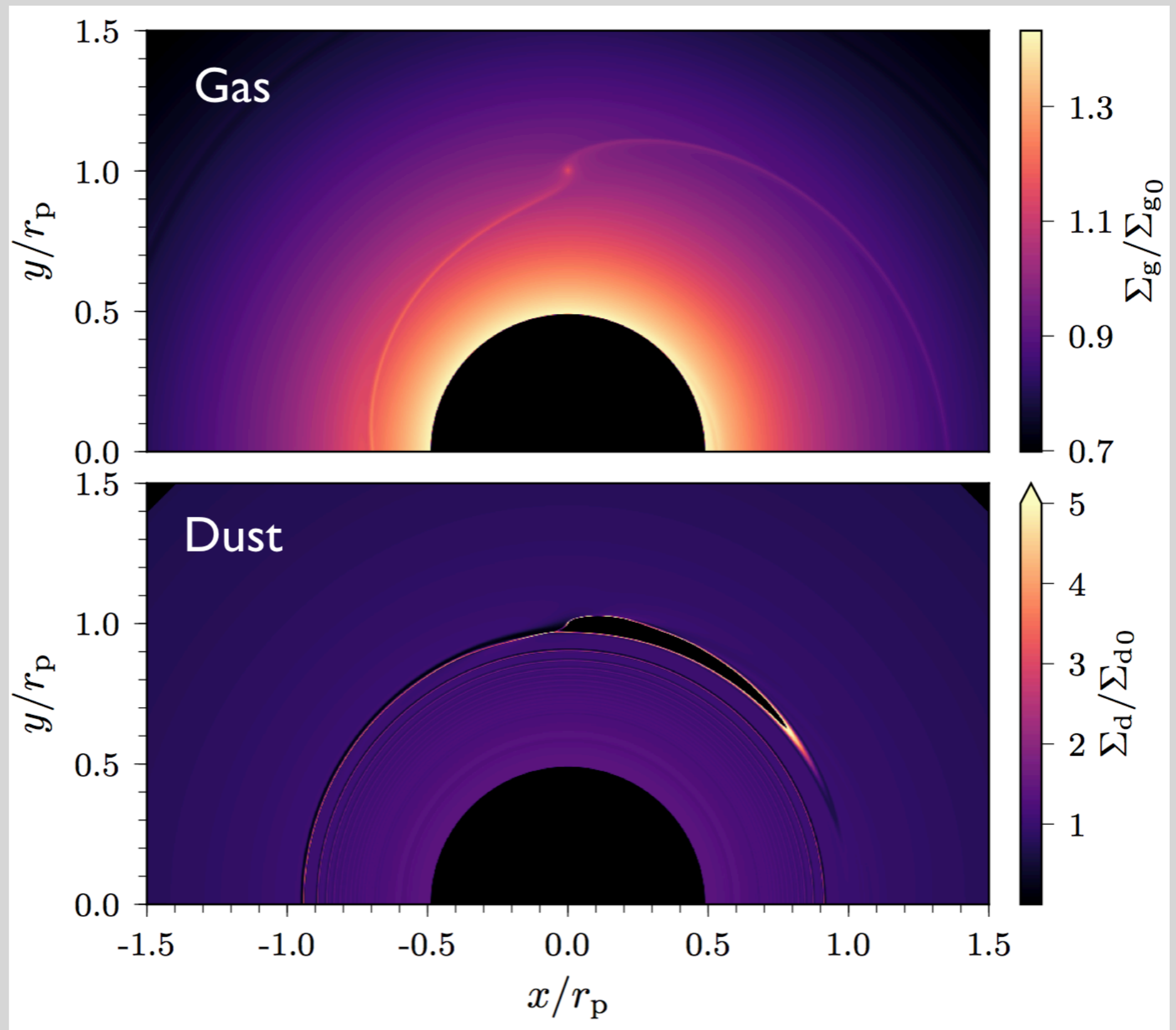
ALMA



# Protoplanetary Disks

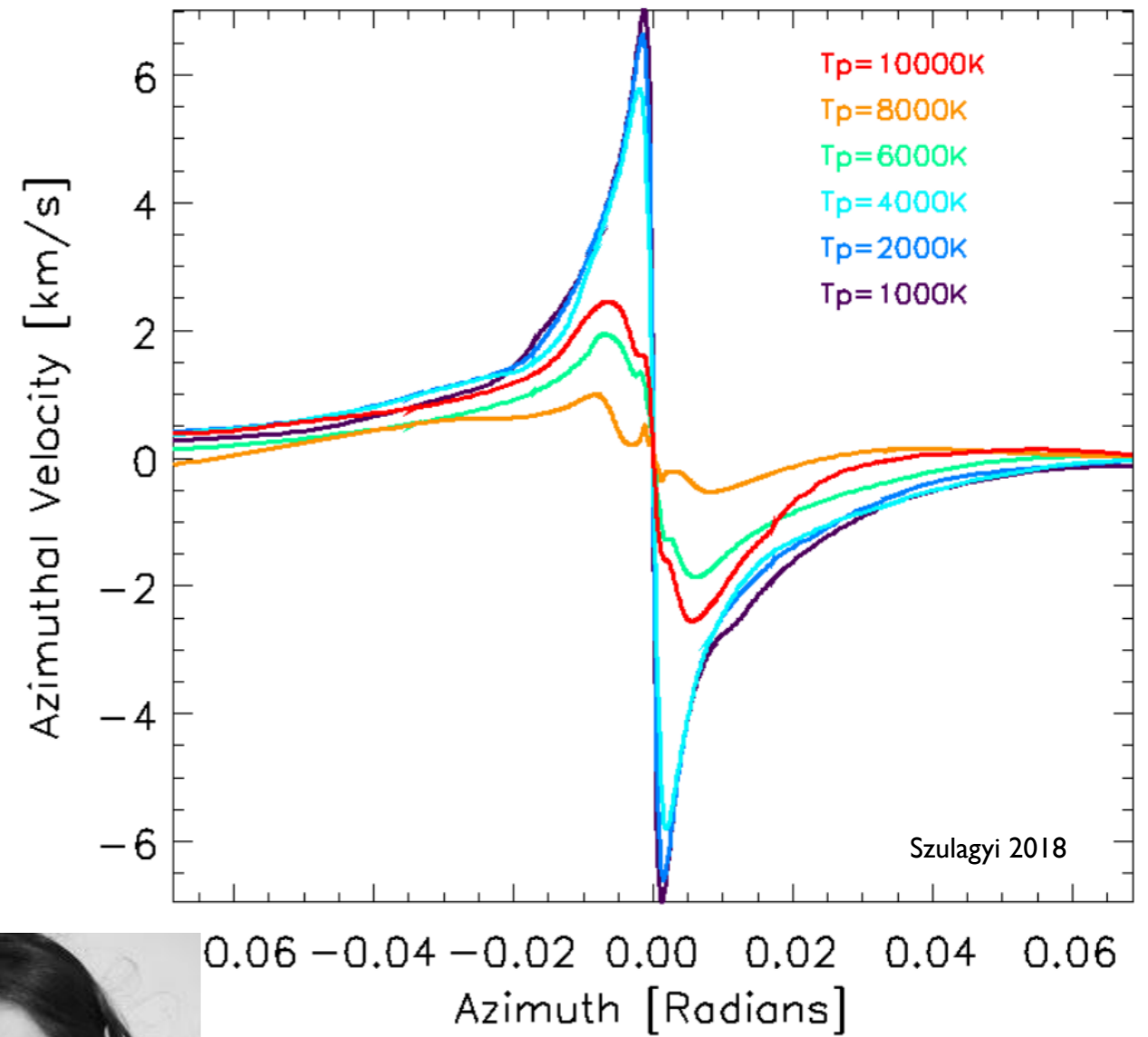
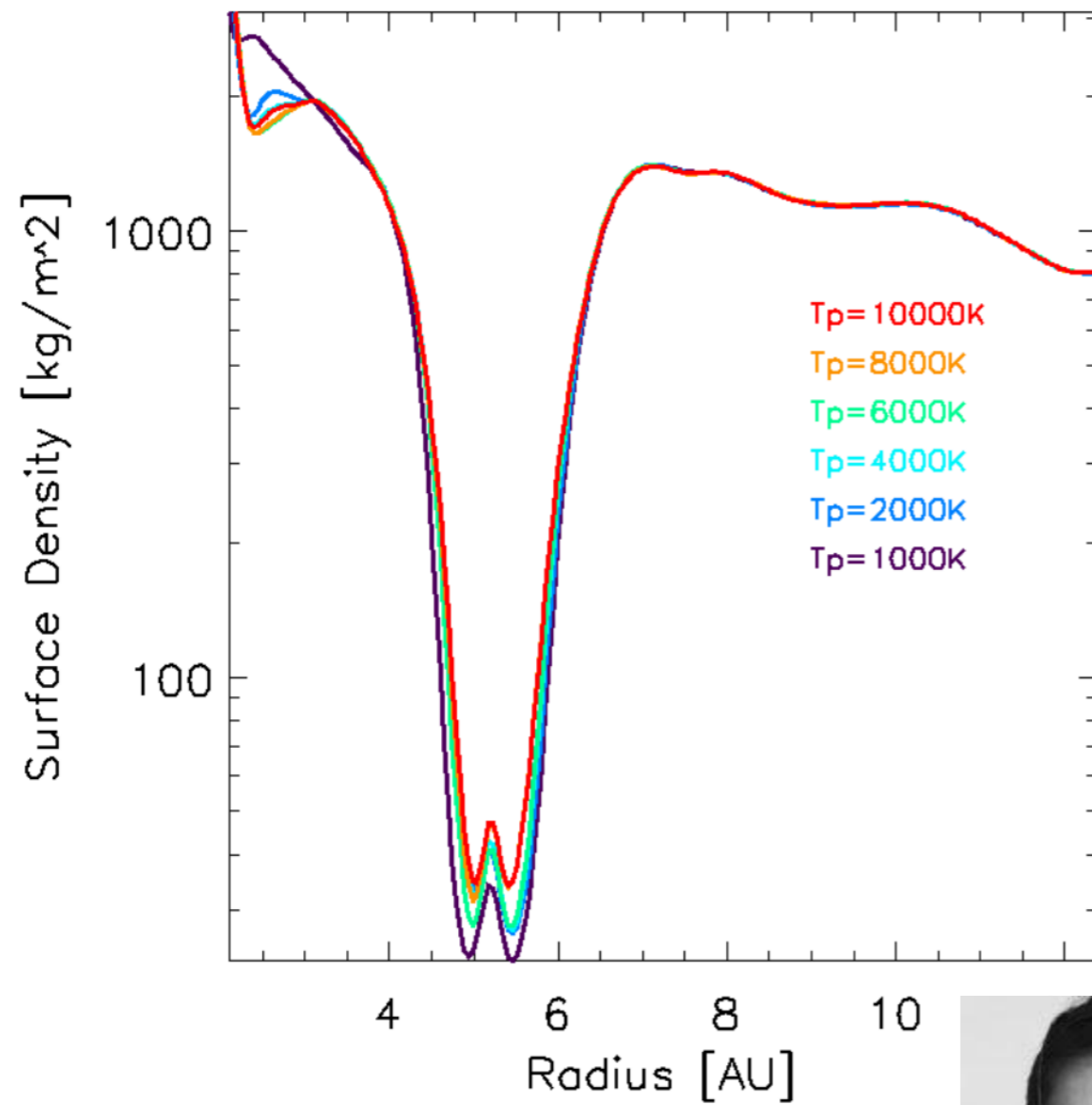


# Dust & Planetary Migration



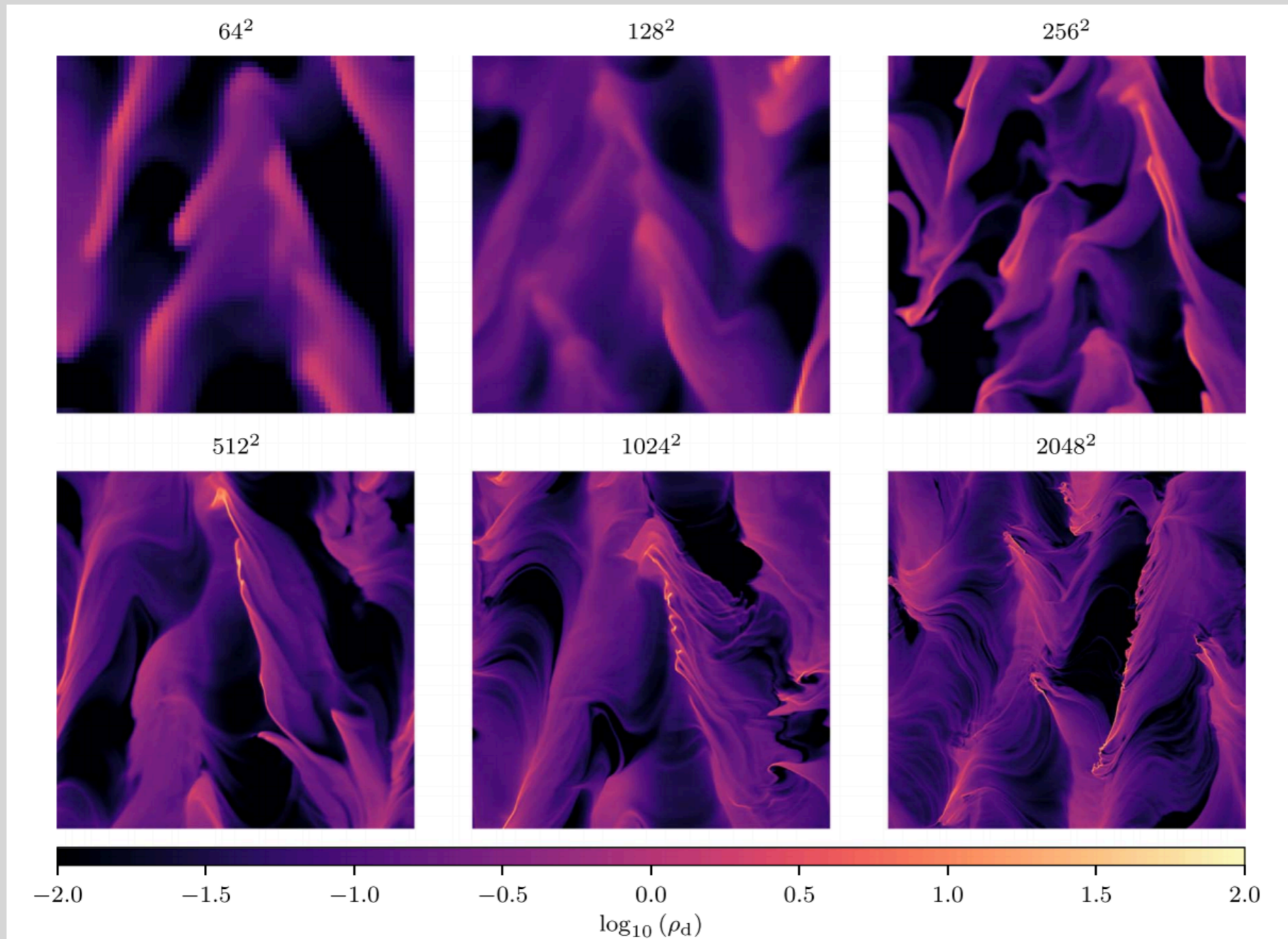
Pablo Benitez-Llambay

# Structure of Disks and Gaps



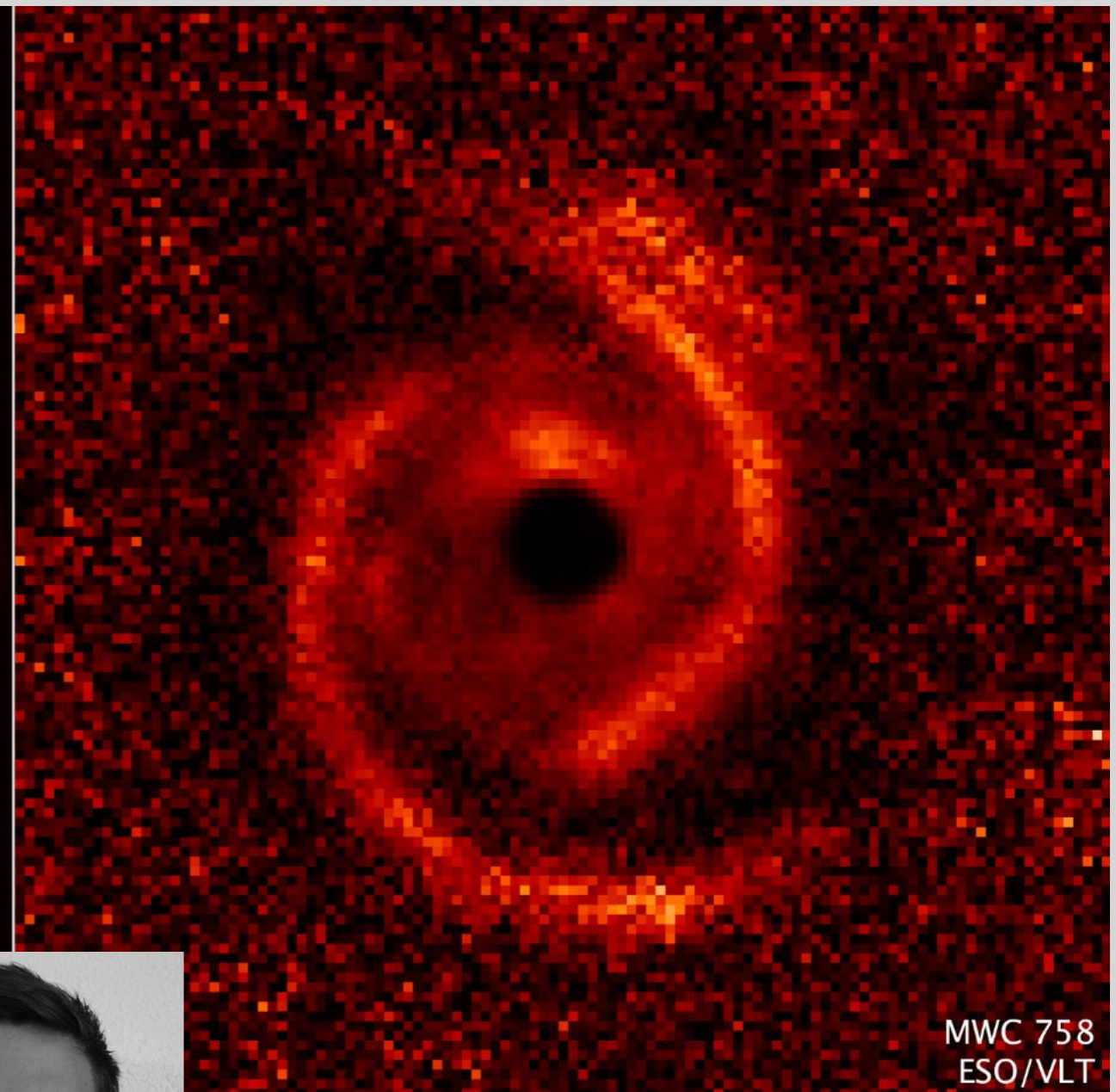
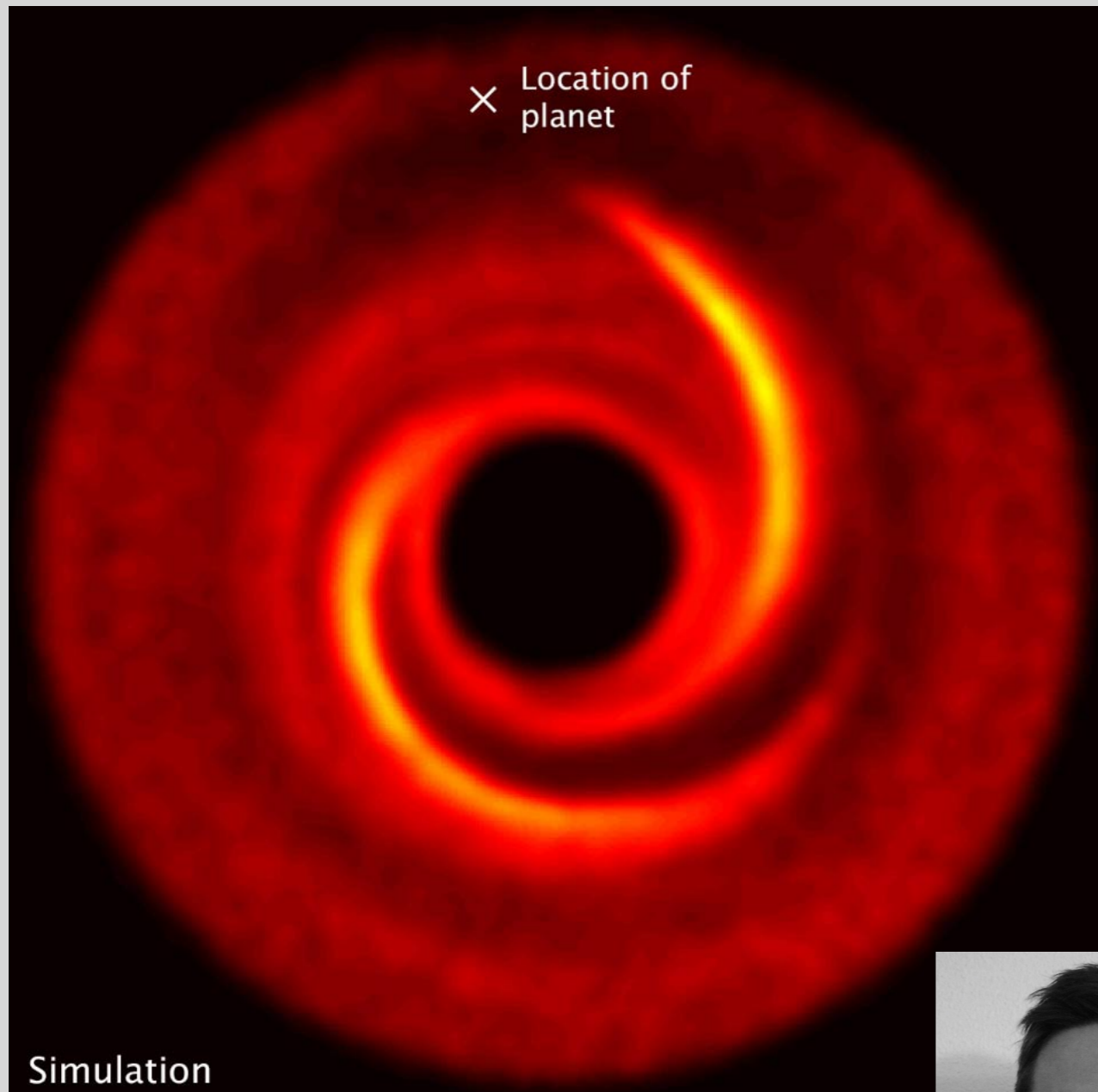
Ximena Ramos

# Key Processes in Planet Formation



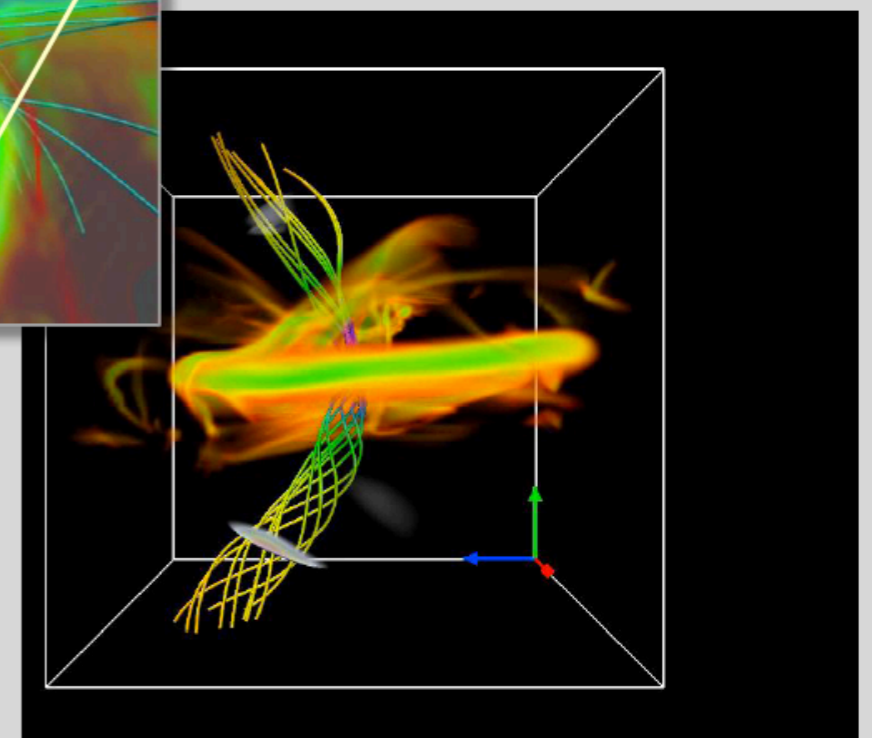
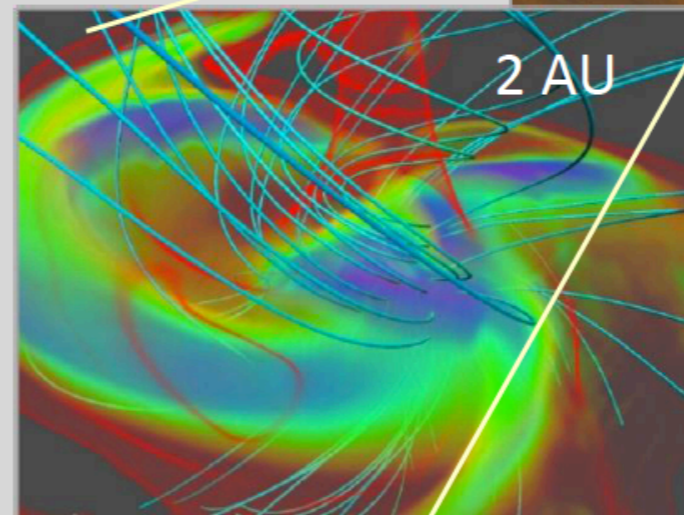
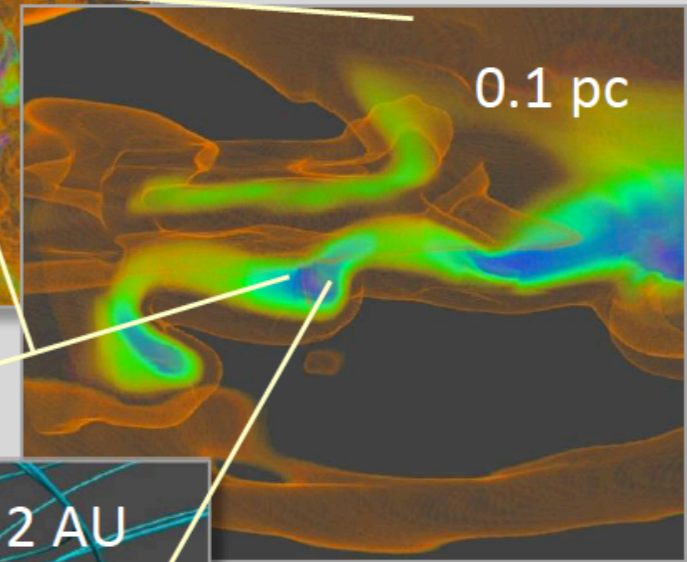
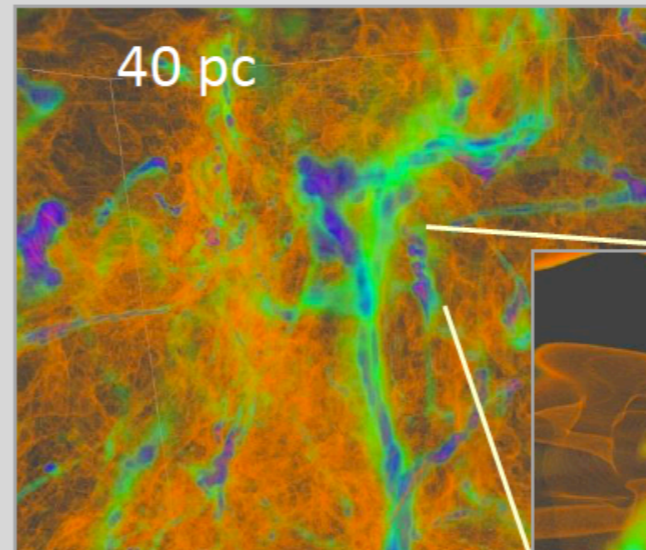
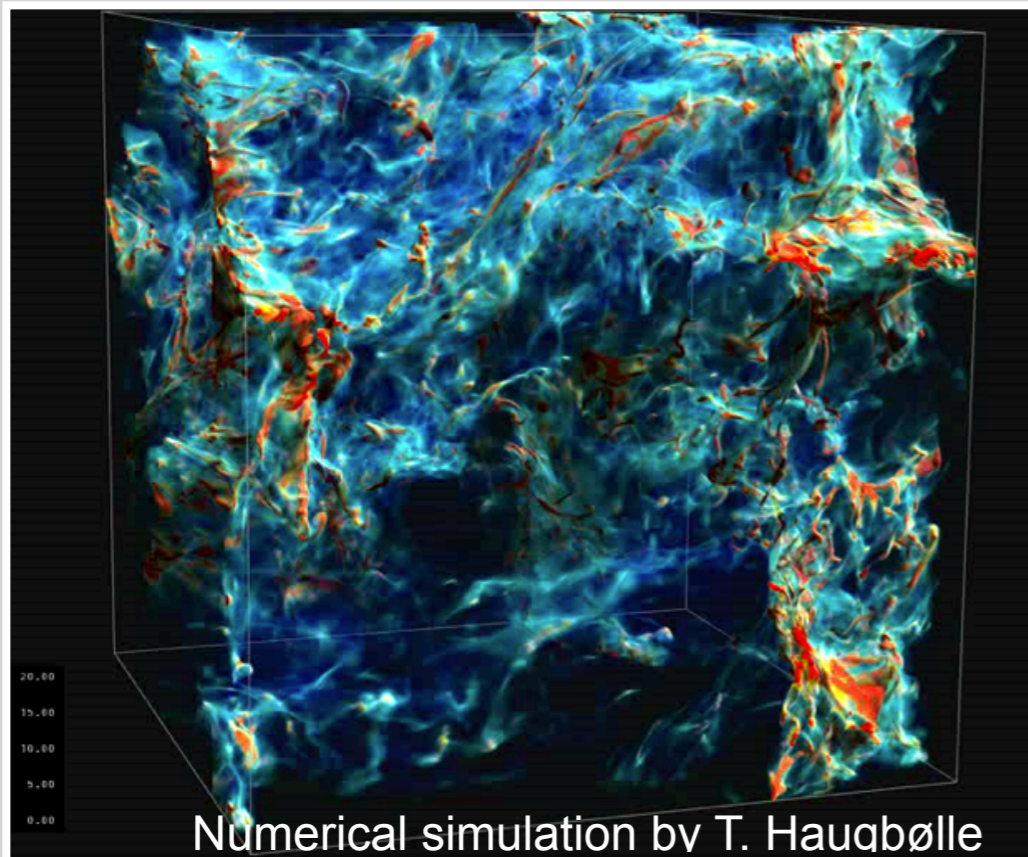
Leonardo Krapp

# Self-Gravitating Disks



Christopher Andersen

# Interstellar Medium & Star Formation



Troels Haugbølle

Troels Haugboelle <[haugboel@nbi.dk](mailto:haugboel@nbi.dk)>

Aake Nordlund <[aake@nbi.dk](mailto:aake@nbi.dk)>

# What You Can Learn

- Fundamental Physics
  - Fluid Dynamics + N-body Dynamics
  - Magnetohydrodynamics
  - Radiative Processes
- Numerical Skills
  - From writing your own scripts to running state-of-the-art numerical codes on thousands of processors
- Learn to think
  - How to use these building blocks for future projects



# Choose Something...

- Interesting  
So you will be motivated and get others excited
- Useful  
So you can build on this for future projects
- Fun  
So you will enjoy working on it!

GOOD LUCK!

# Relevant Webpages

- <https://sites.google.com/view/nbia-astrophysics>
- <https://sites.google.com/view/nbiagwastro>
- <http://www.astrophysics.nbi.ku.dk>

You are always welcome to email us or stop by!