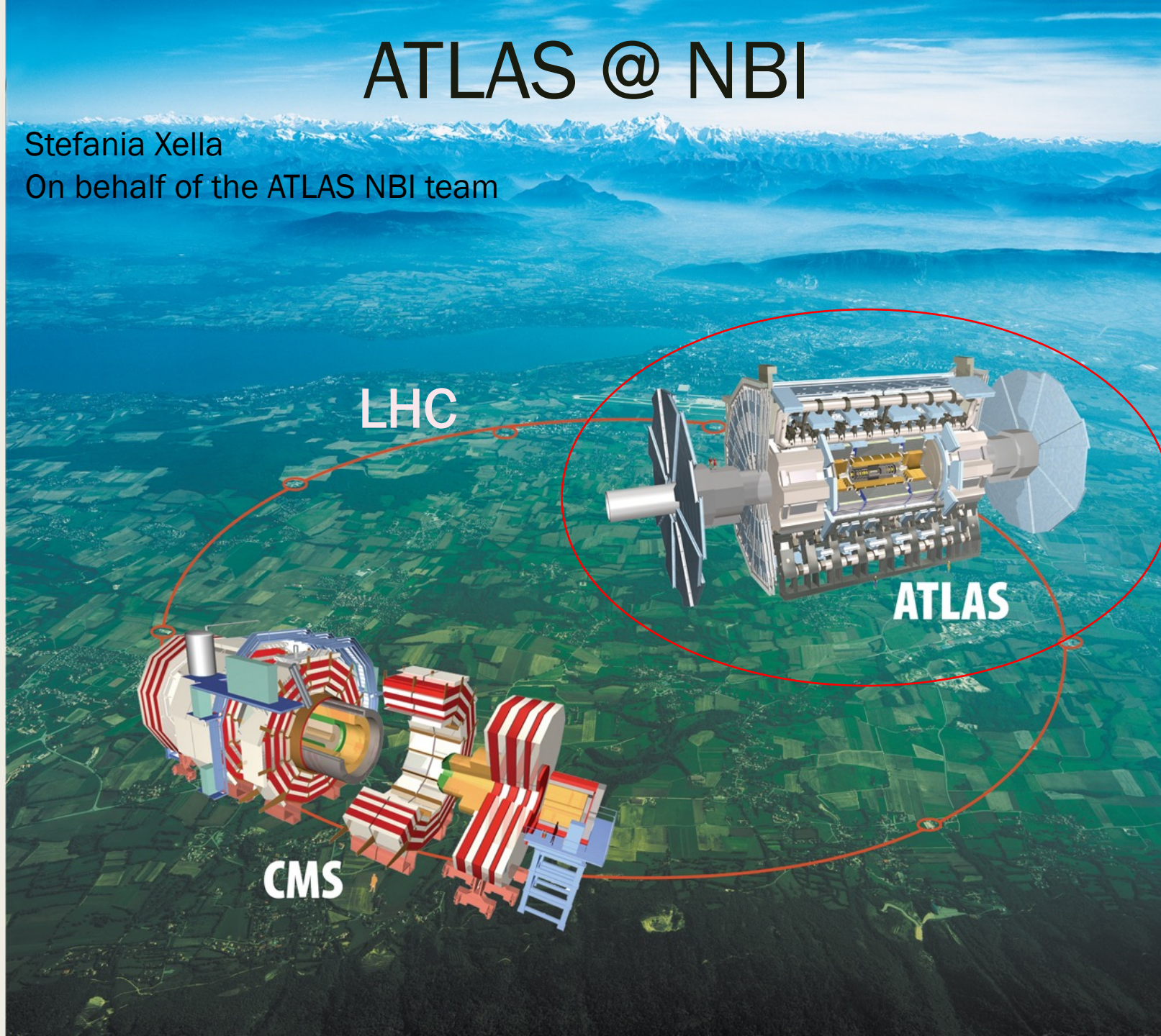


# ATLAS @ NBI

Stefania Xella

On behalf of the ATLAS NBI team



# LHC and ATLAS

Luminosity  $\mathcal{L}$  = how many particles we're able to squeeze through a given space in a given time

$$\mathcal{L} \cdot \sigma = \text{collision rate}$$

Two successful LHC Runs so far :

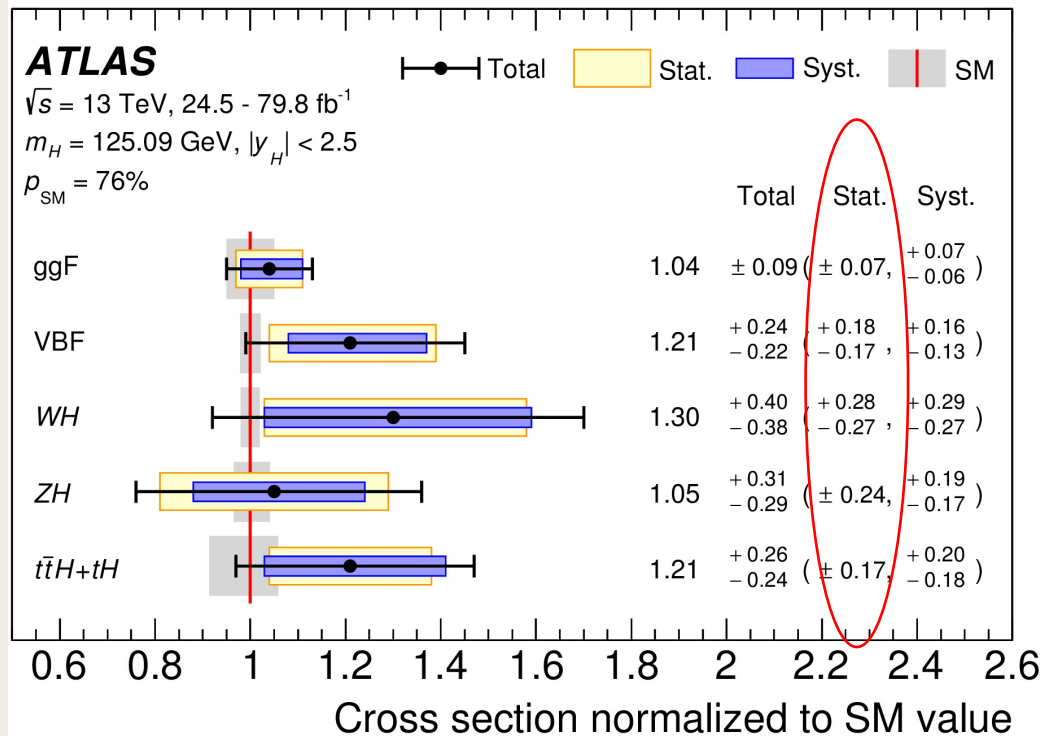
- 2008-2012, pp collision cms energy 7-8 TeV ,  $\mathcal{L}$  peak  $0.75 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
- 2015-2018, pp collision cms energy 13 TeV ,  $\mathcal{L}$  peak  $1.7 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$

ATLAS is a multi-purpose experiment at the LHC (the other is CMS). Some of the successes (imo):

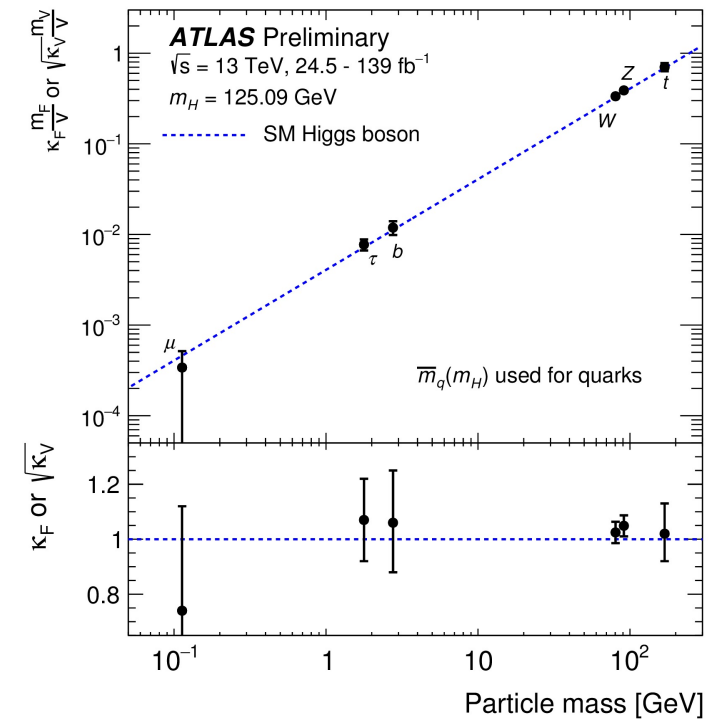
- Observed the Standard Model (SM) Higgs decaying in Z, W,  $\gamma$ ,  $\tau$ , b ,  $\mu$  particles and produced in gluon-gluon or Vector Boson fusion processes, or associated with a W or Z or top particle.
- Measured several rare SM processes, uncharted before: 1-3 bosons, 1-4 top quarks,...  
This is important for searches for rare new phenomena, beyond the SM.
- Excluded most obvious (=less likely?) BSM scenarios (in supersymmetry or other exotic theories).  
new particles minimum 1-3 TeV or super-weakly interacting with SM particles
- Showed complementarity with direct and indirect dark matter search experiments.  
collider experiments important role in the quest for dark matter.
- Sensitivity to lepton flavour/lepton number/lepton universality violation.  
In some cases doing better than LEP experiments

# Higgs properties [H]

## Higgs production modes



## Higgs decay modes



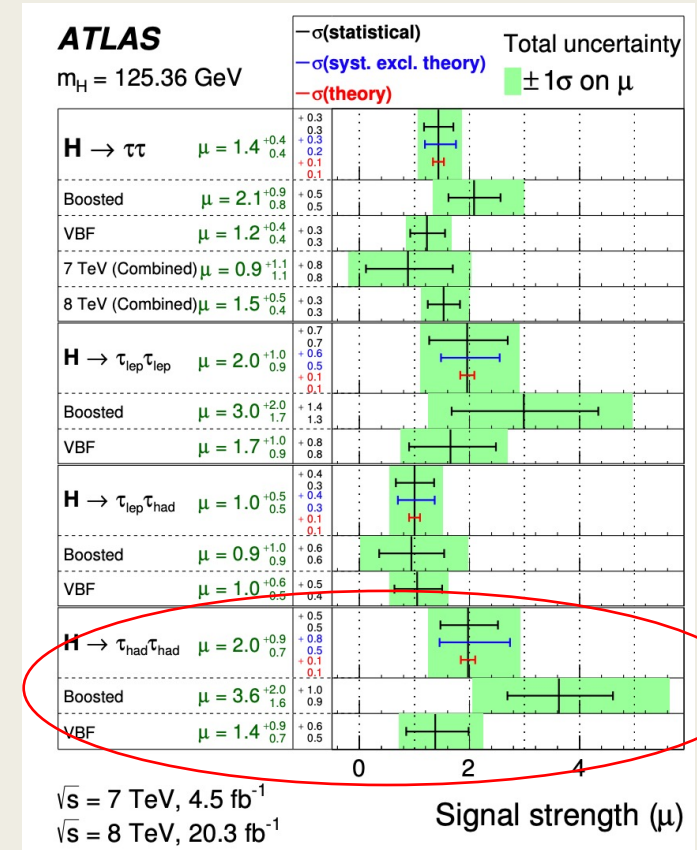
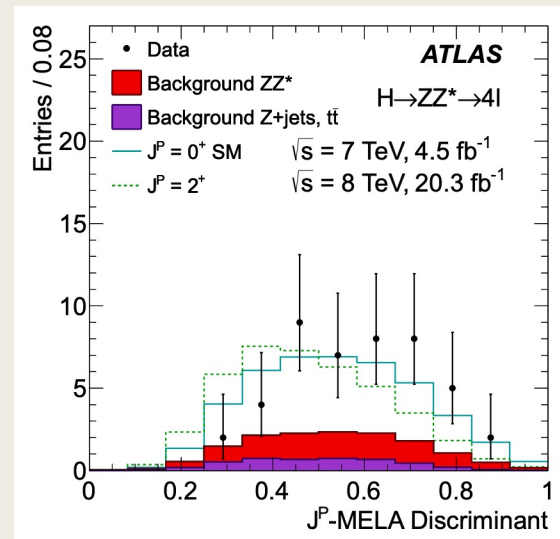
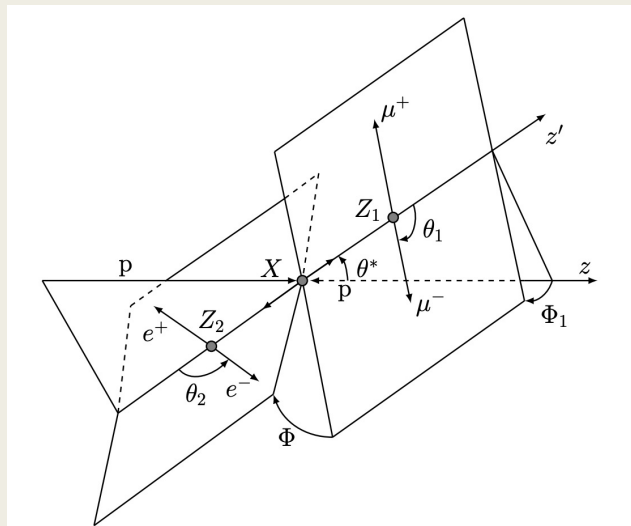
No indication of deviations from SM or additional Higgs particles yet  
 Statistical and systematic errors are comparable -> room for improvements



# Higgs decay modes [[htt](#), [hzz](#)]

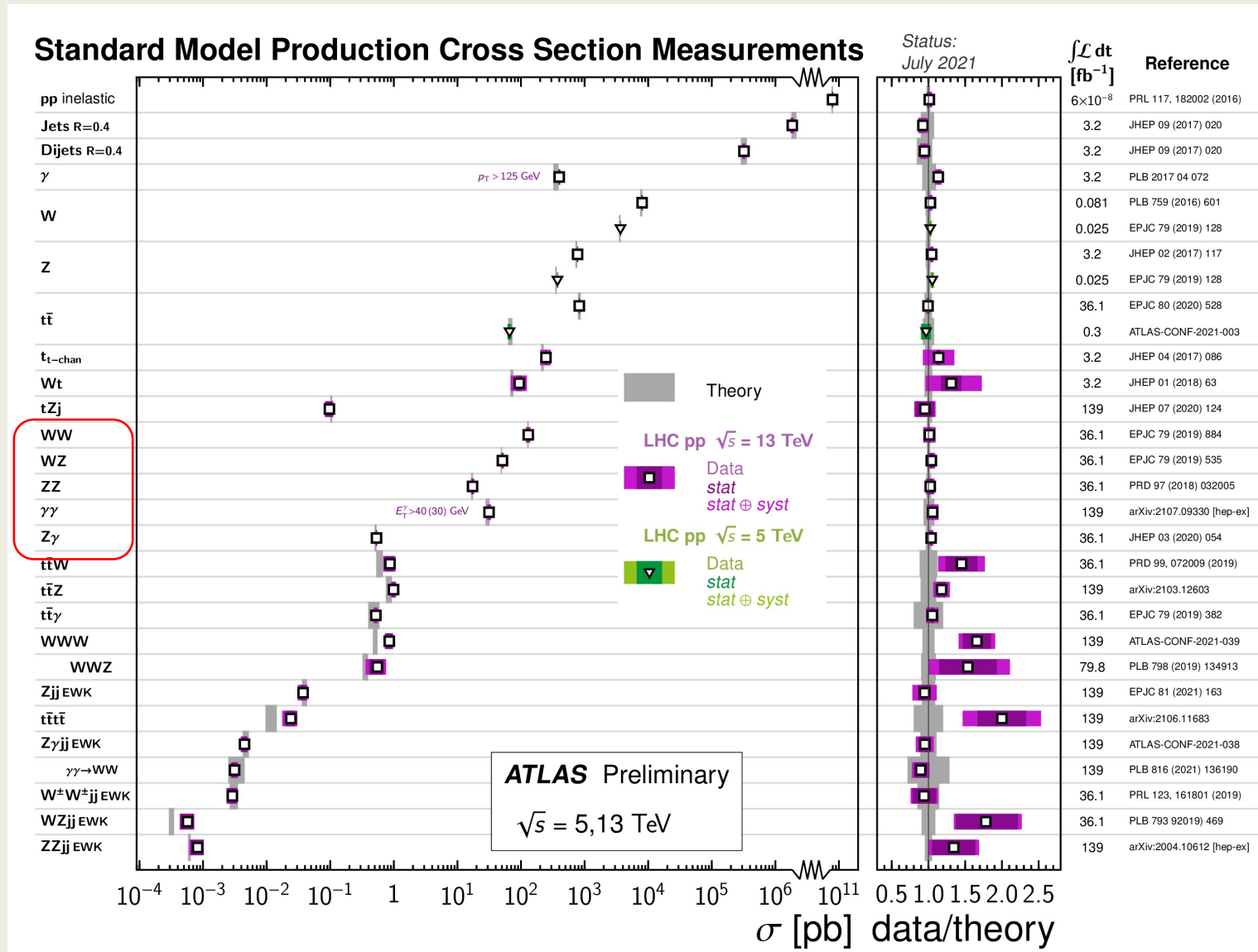
NBI

- Developed tau trigger algorithms and tau reconstruction and identification algorithms in Run 1, instrumental for the observation of the Higgs decay to tau leptons with Run 1 data, before CMS.
- Developed the parity and spin analysis of the Higgs boson, and excluded the spin-2 hypothesis, as well as other BSM spin-0 hypotheses, already with Run 1 data alone.



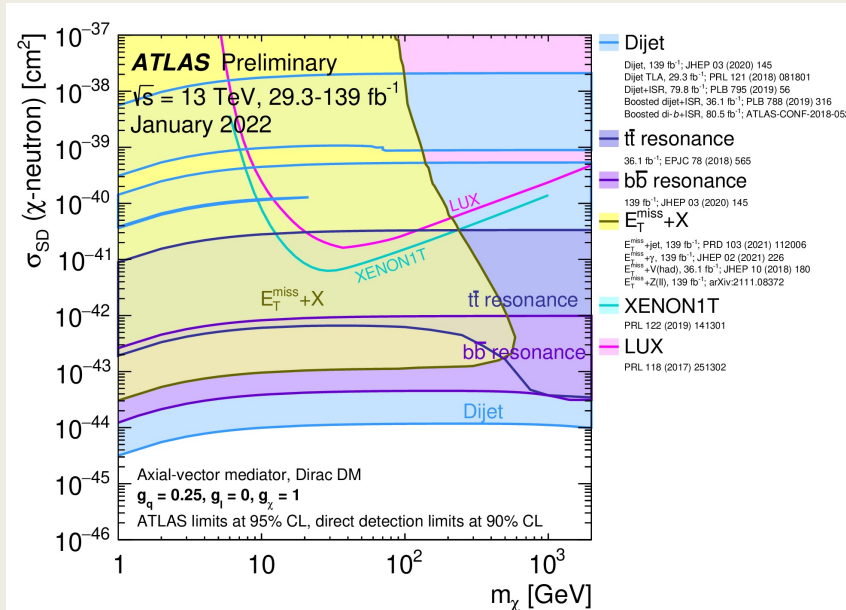
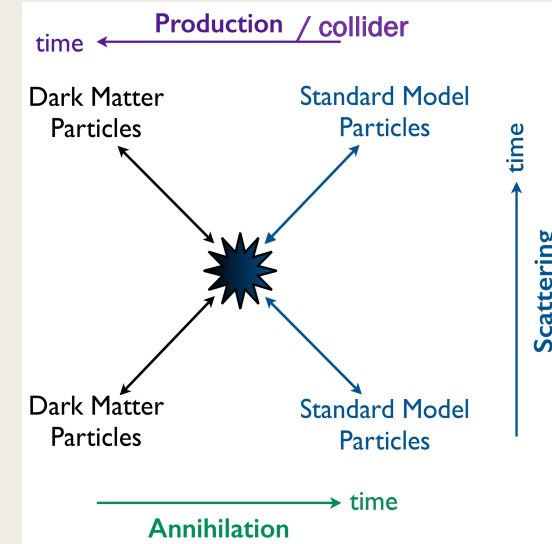
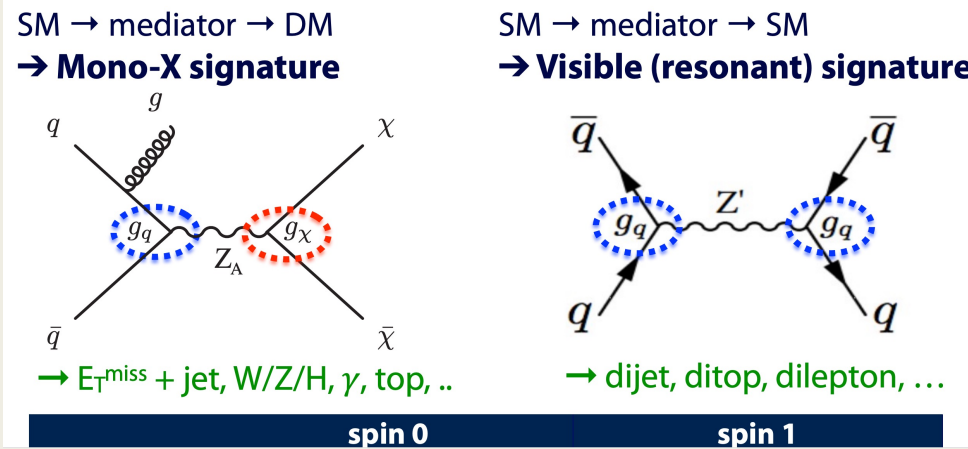


# Rare standard Model processes [[sm](#)]



Performed the Run 1 ZZ measurements and set the first limits on anomalous gauge couplings with LHC data

# Dark matter

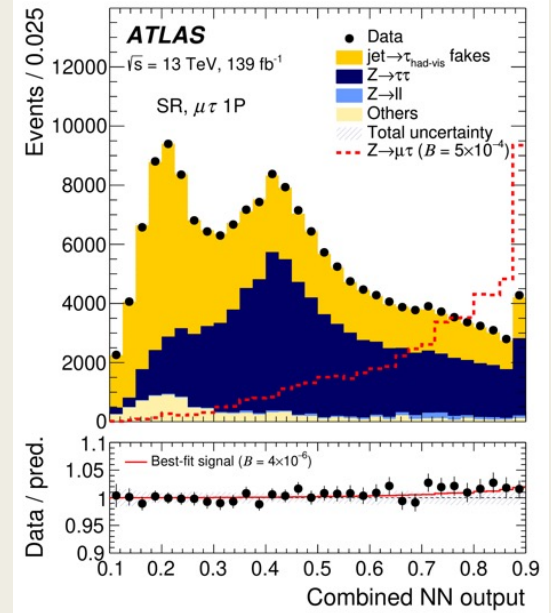
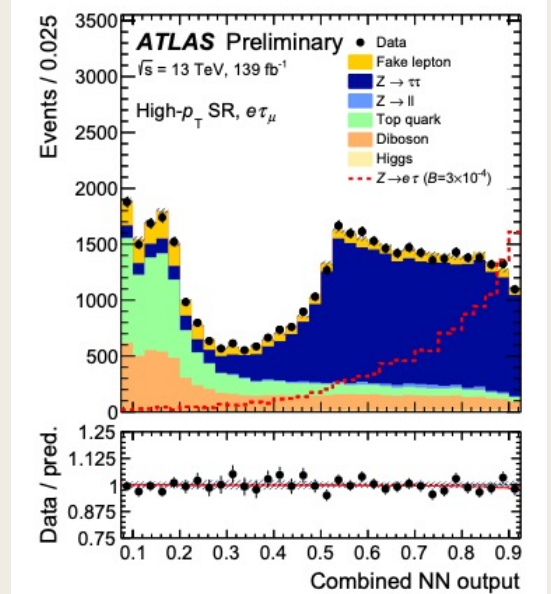
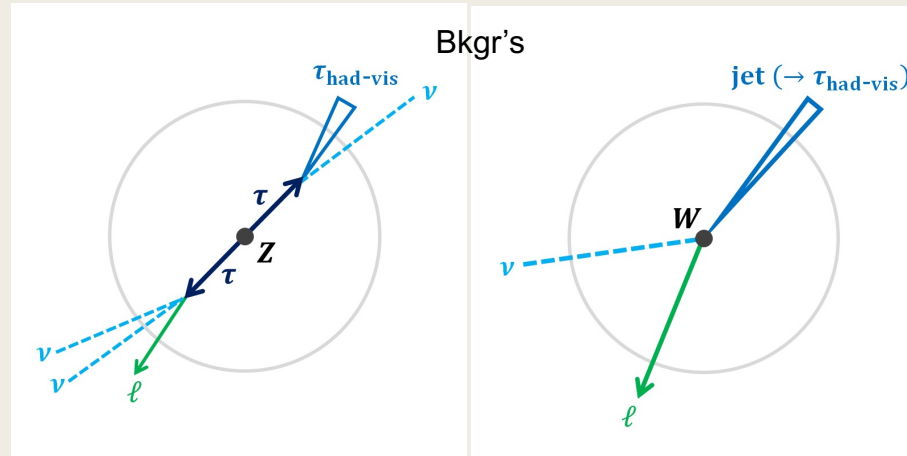


Simplified model for DM :  
One new force and one new particle, DM and SM particles interact

Complementarity of colliders and direct searches (eg LUX, Xenon1T) for dark matter

# LFV Z decays to $\ell + \tau$ [[lfvz1](#), [lfvz2](#)]

e.g. tau hadronic decays



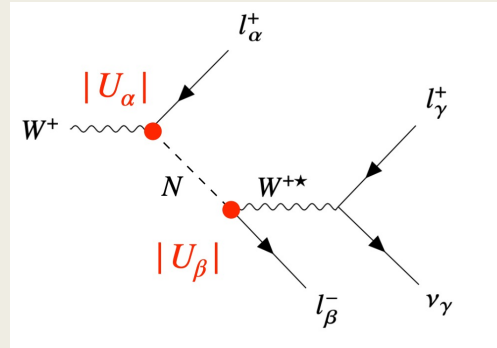
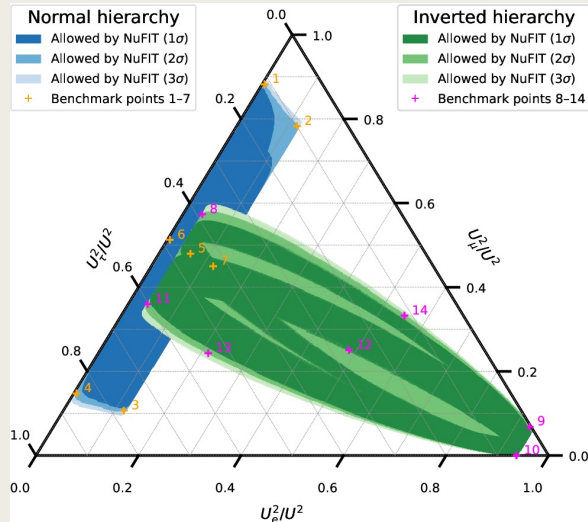
Final state, polarization assumption	Observed (expected) upper limit on $\mathcal{B}(Z \rightarrow \ell\tau)$ [ $\times 10^{-6}$ ]	
	$e\tau$	$\mu\tau$
$\ell\tau_{\text{had}}$ Run 1 + Run 2, unpolarised $\tau$ [10]	8.1 (8.1)	9.5 (6.1)
$\ell\tau_{\text{had}}$ Run 2, left-handed $\tau$ [10]	8.2 (8.6)	9.5 (6.7)
$\ell\tau_{\text{had}}$ Run 2, right-handed $\tau$ [10]	7.8 (7.6)	10 (5.8)
$\ell\tau_{\ell'}$ Run 2, unpolarised $\tau$	7.0 (8.9)	7.2 (10)
$\ell\tau_{\ell'}$ Run 2, left-handed $\tau$	5.9 (7.5)	5.7 (8.5)
$\ell\tau_{\ell'}$ Run 2, right-handed $\tau$	8.4 (11)	9.2 (13)
<b>Combined <math>\ell\tau</math> Run 1 + Run 2, unpolarised <math>\tau</math></b>	<b>5.0 (6.0)</b>	<b>6.5 (5.3)</b>
Combined $\ell\tau$ Run 2, left-handed $\tau$	4.5 (5.7)	5.6 (5.3)
Combined $\ell\tau$ Run 2, right-handed $\tau$	5.4 (6.2)	7.7 (5.3)
LEP OPAL, unpolarised $\tau$ [10]	9.8	17
LEP DELPHI, unpolarised $\tau$ [11]	22	12

Measurement still statistically limited

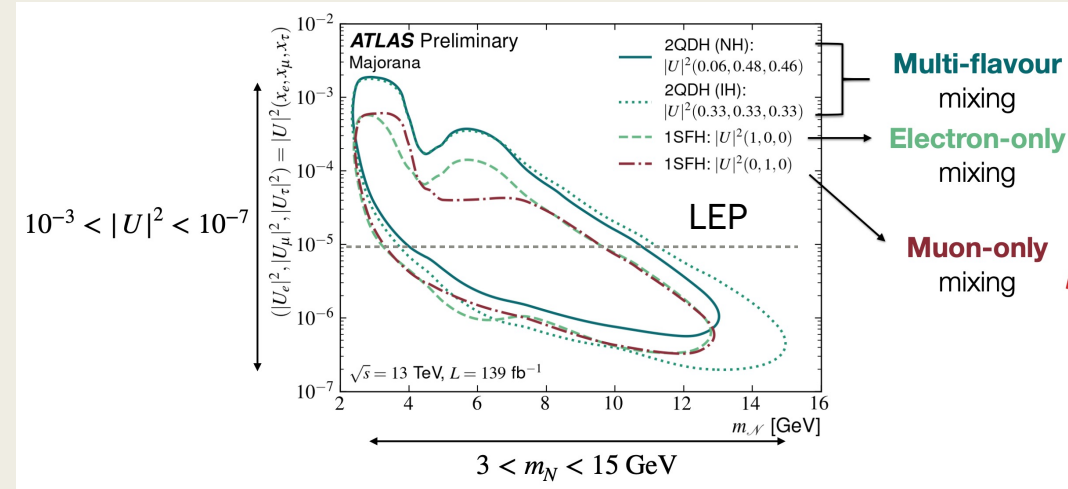
ATLAS sets the world best limit



# Heavy Neutral Leptons [[hnl1](#), [hnl2](#)]



$$\tau_N \propto \frac{1}{m_N^5 |U_\alpha|^2}$$



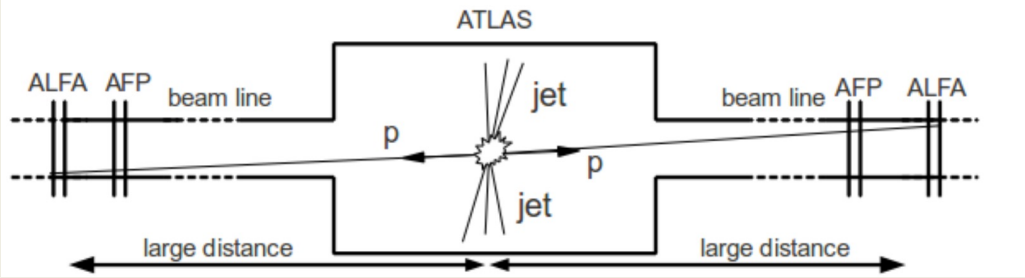
“Realistic” multi-flavour mixing models consistent with neutrino oscillations data

- Introduce 3 right-handed states, and use seesaw mechanism to explain low mass for weak-interaction neutrinos (and dark matter and baryon asymmetry).
- ATLAS has demonstrated sensitivity to e-N and  $\mu$ -N mixing at the same level as LEP experiments. Has also published a realistic interpretation of data for full Run 2 data analysis
- Investigate now the Run 2 data sensitivity for tau-N mixing and understand how to improve before Run 3 starts (new triggers)

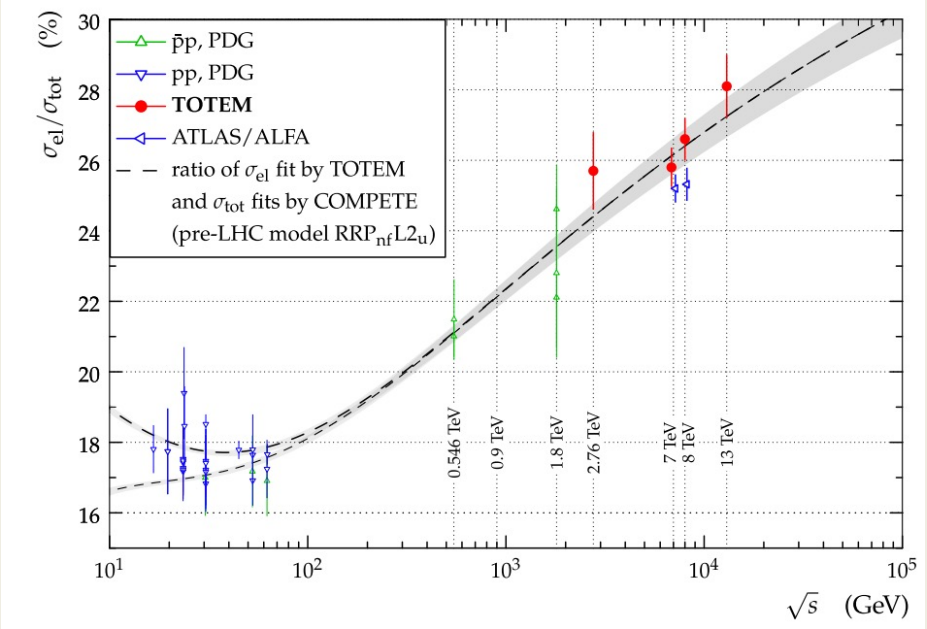
# Total elastic cross section [alfa, tot]

$$\sigma_{\text{tot}}^2 = \frac{16\pi(\hbar c)^2}{1 + \rho^2} \left. \frac{d\sigma_{\text{el}}}{dt} \right|_{t=0}, \quad \rho = \frac{\text{Re}[F_{\text{el}}(t)]}{\text{Im}[F_{\text{el}}(t)]},$$

t=4-momentum transfer between protons



- measurement of the differential cross section down to small angles gives not only the elastic cross section by integration, but also the total cross section from which the inelastic cross section can be deduced.
- Dedicated runs with low pileup and high  $\beta^*$  (low t) , with ALFA detector , allow to measure unfolded differential elastic Xsection, and then fit it with  $\sigma_{\text{tot}}$  as free parameters ( $\rho$  from models).
- Done measurement at 8 TeV. Working on measurement at 13 TeV.
- A final data taking with ALFA is planned in run 3 with very high- $\beta^*$  optics for a measurement of elastic scattering and associated parameters  $\sigma_{\text{tot}}$ ,  $\rho$  and slope. Worth only if cms energy > 13.5 TeV

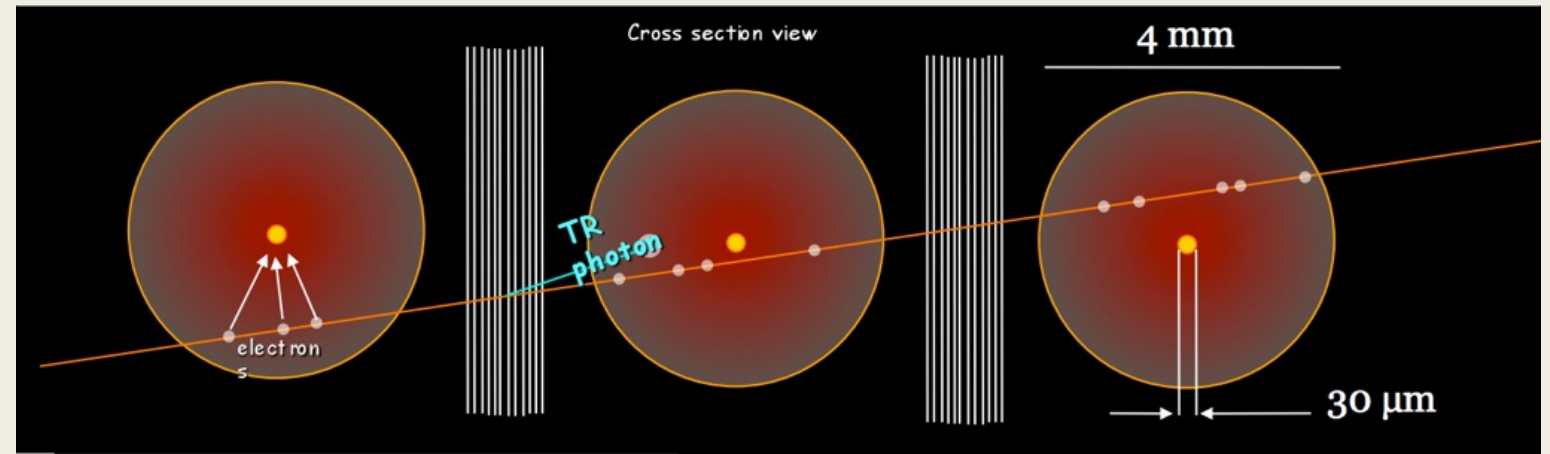
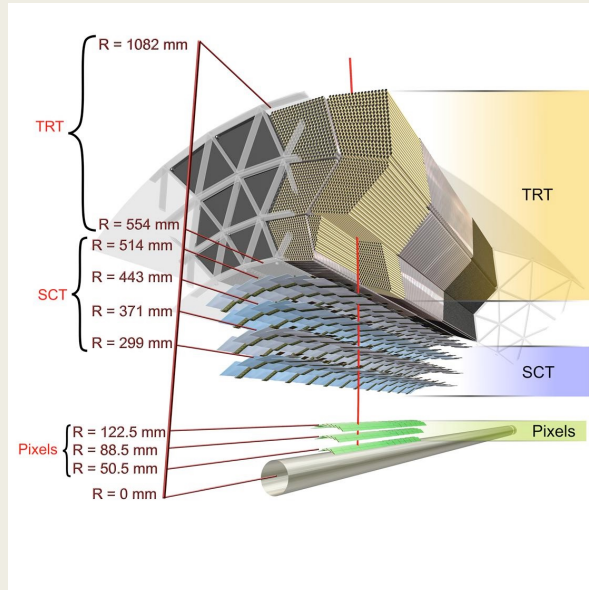


# Run 3 (2022-2025)

- A new run starts in May, pp collision cms energy 13.5 TeV and  $\mathcal{L}$  peak  $2 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
- Run 3 will not add a lot of data, but will be used strategically to focus on new type of searches : eg
  - *long lived particles, super-weakly interacting particles (continue from Run2)*
  - *dijet resonances, etc..*
  - *B physics*
  - *Rare Higgs decays*
- Run 3 will also test completely new software and hardware, which will stay with ATLAS until the end of HL-LHC
  - *New L1 trigger hardware and Central Trigger Processor*
  - *New multi threaded software (eg. tracking) and extensive use + establishment of machine learning for data analysis*



# Transition Radiation Tracker (TRT)



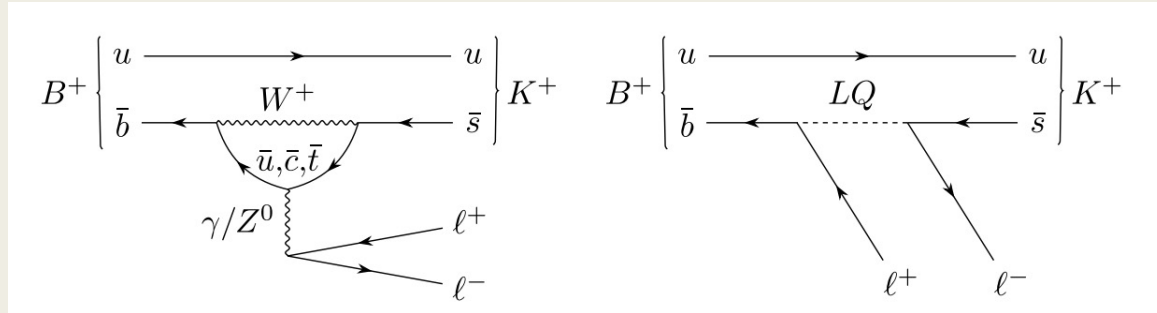
NBI contributes to TRT operations and maintenance from start Run 1 until end of Run 3

TRT provides half the momentum resolution and significant  $e/\pi$  rejection, capabilities which are crucial for several analyses

Challenges in Run 3 : deal with leaks – partly radiation induced – of both the cooling gas (a potent greenhouse gas) and the active Xenon gas (important for the electron identification, but at present very expensive).

The effects of closing modules and switching other modules to Argon will be mitigated by improvements in software.

# Bphysics



$$R_K = \frac{\mathcal{B}(B^+ \rightarrow K^+ \mu^+ \mu^-)}{\mathcal{B}(B^+ \rightarrow J/\psi (\rightarrow \mu^+ \mu^-) K^+)} \bigg/ \frac{\mathcal{B}(B^+ \rightarrow K^+ e^+ e^-)}{\mathcal{B}(B^+ \rightarrow J/\psi (\rightarrow e^+ e^-) K^+)}.$$

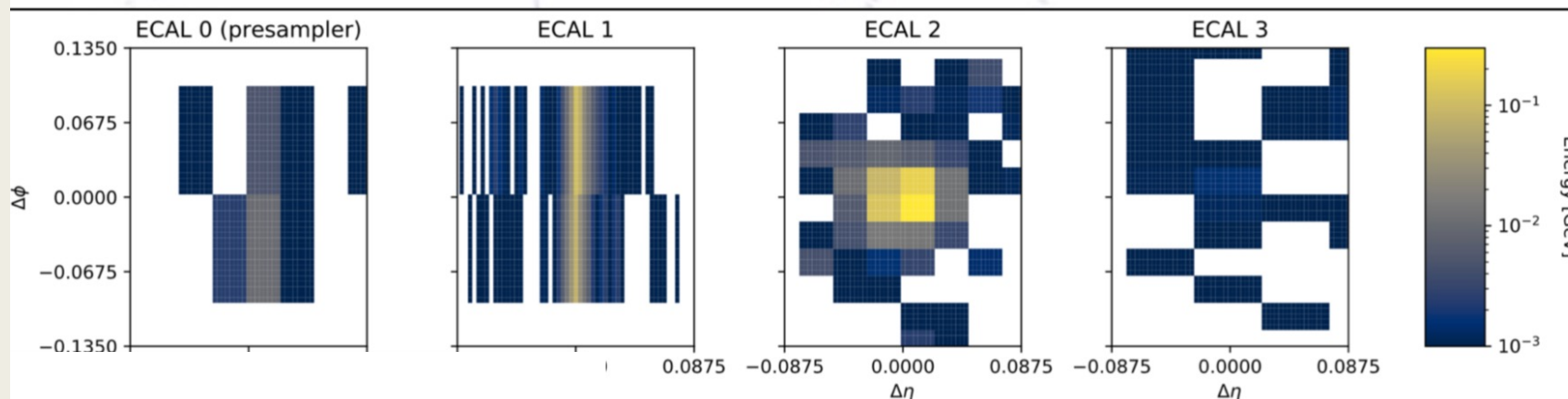
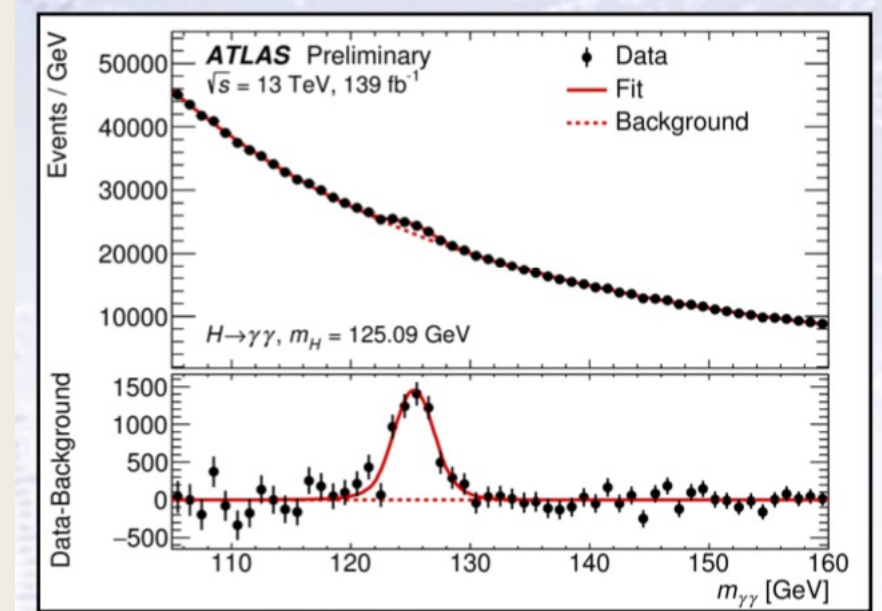
K could be other hadrons, eg  $K^{*0}$   
 B+ could be other B hadron, eg  $B^0$ ,  $B_s^0$

- This is a rare decay, and the low electron momentum is a trigger challenge for ATLAS.
- Dedicated triggers were included in the last part of Run 2 data taking.
- New triggers are studied for Run 3, and improved data acquisition and rate allocation should allow to make ATLAS competitive.
- In Run 2 we collected enough events with  $\ell = e, \mu$  to do a first measurement of  $R_{K^*}$
- The analysis strategy is to select events (with ML) independently of the invariant masses, and fit the reconstructed B mass to estimate signal size and separate non-resonant (signal) from resonant ( $J/\psi$  control) contributions.

# Photon /electron energy reconstruction

NBI

- The photon energy resolution  $\sigma(E_\gamma)$  is the **dominant performance parameter** for the  $H \rightarrow \gamma\gamma$
- Our team wishes to improve this resolution in ATLAS.
- Consider the cell energies in the LAr calorimeter as pixels in four images (one for each layer), and train a Convolutional Neural Network (CNN) to estimate the energy of electron/photon candidates.



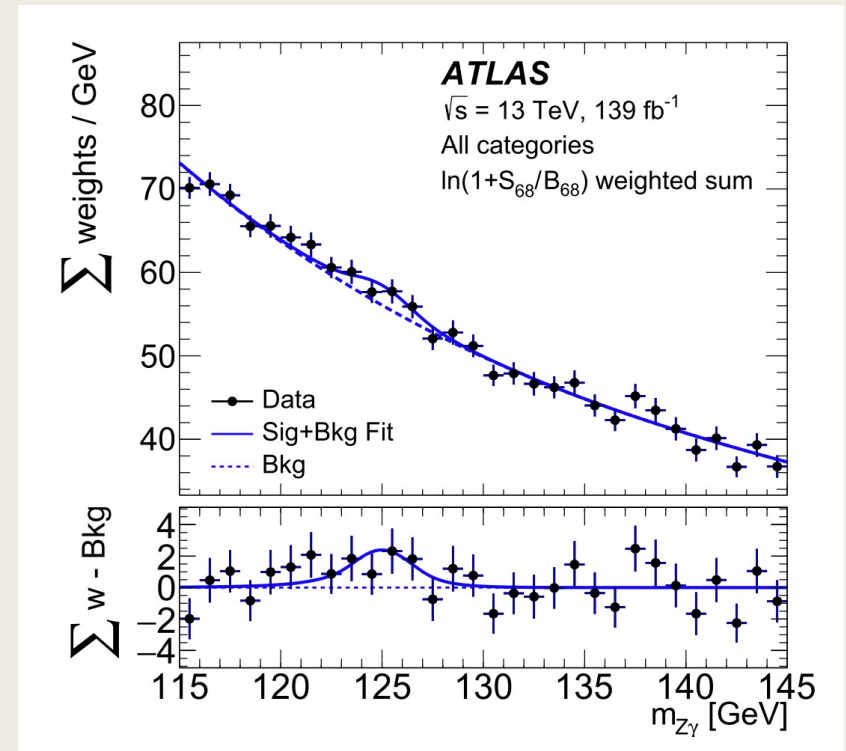
Measured the improvement in data, using electrons (we have billions of  $Z \rightarrow ee$ )

Expected improvement on photons  $E$  resolution of 20%

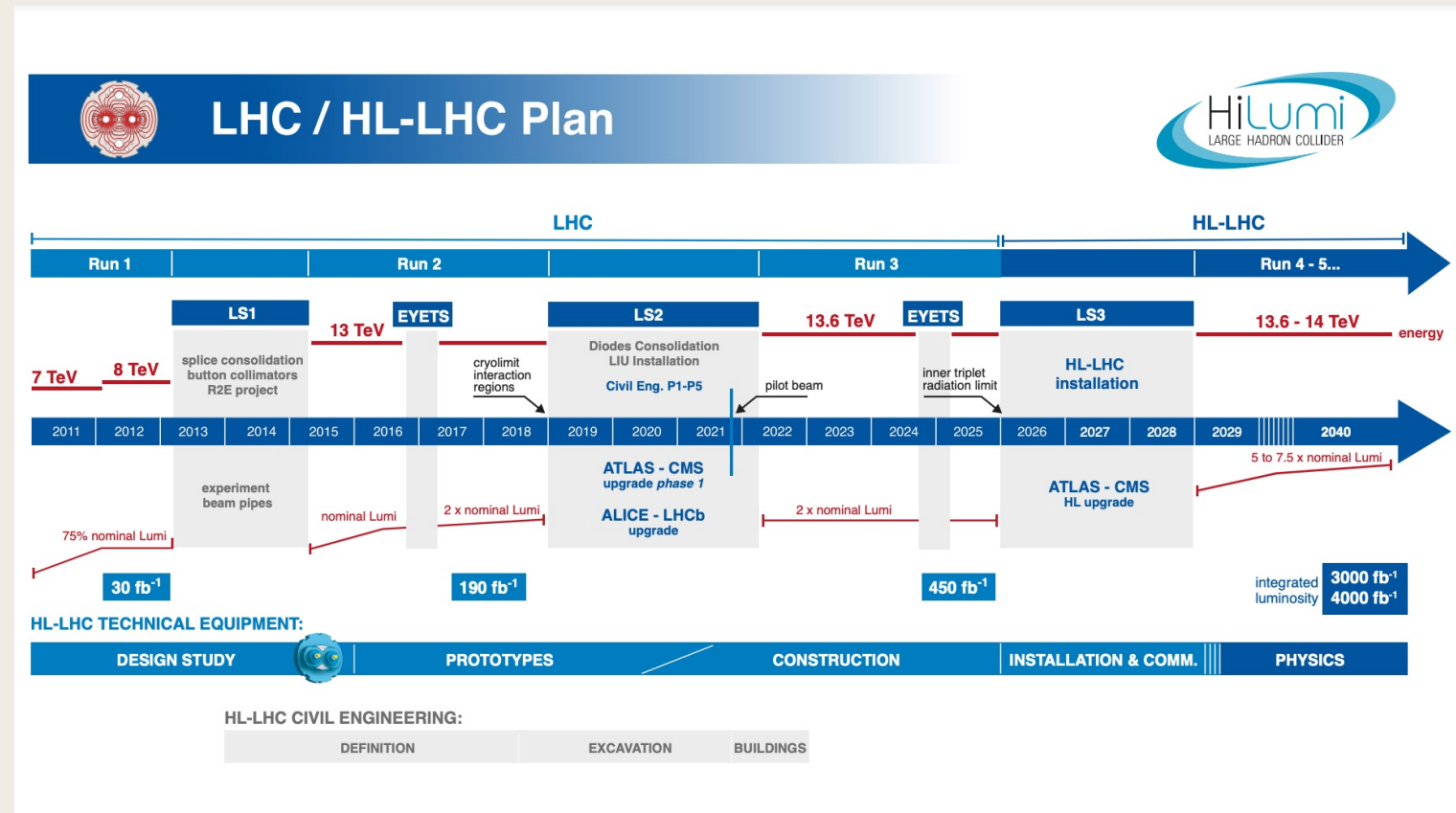


# Rare Higgs decays : $Z \gamma$

- Current result, no optimized photon energy resolution yet (see right)
- The observed 95% CL upper limit on the  $\sigma(pp \rightarrow H) \cdot B(H \rightarrow Z\gamma)$  is 3.6 times the SM prediction for a Higgs boson mass of 125.09 GeV [zg].
- Main background : non resonant Z production in association with a photon
- With expected improvement in photon energy resolution just shown, one can be sure to have a clear observation of this H decay within Run 3 end



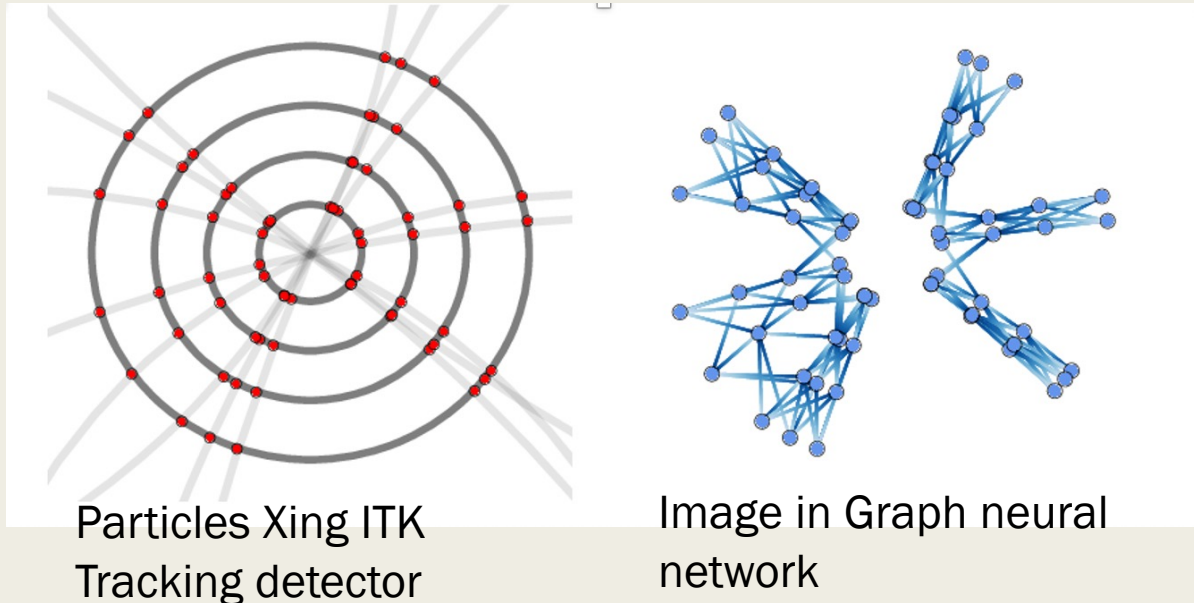
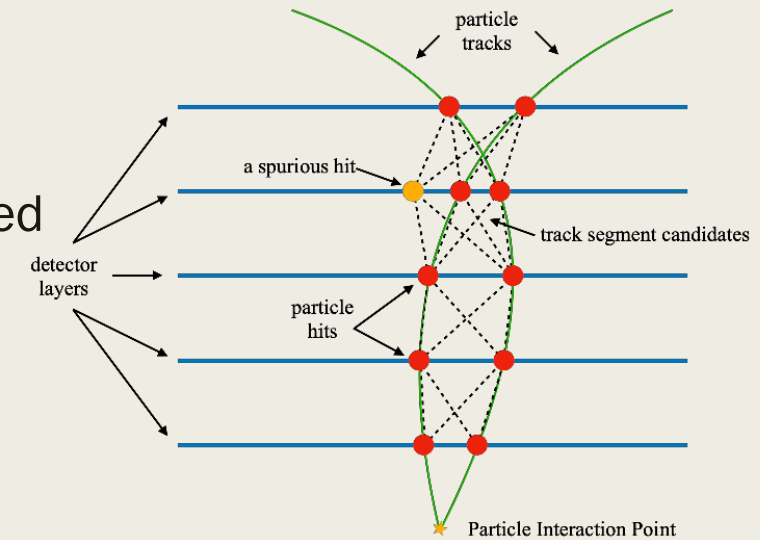
# Beyond Run 3: High Luminosity LHC



Big game changer for improving the SM and beyond SM physics measurements shown earlier. The high luminosity LHC which will bring 10 times more data, and additionally the ATLAS tracker will have a significant increase in angular acceptance.

# New tracking trigger

- At HL-LHC the collision data rate will increase by a factor 10
- the ATLAS real-time data selection (trigger) is challenged : speed and efficiency are the critical parameters.
- charged particle track reconstruction is the bottleneck



New pattern recognition algorithms development

Focus on choice of

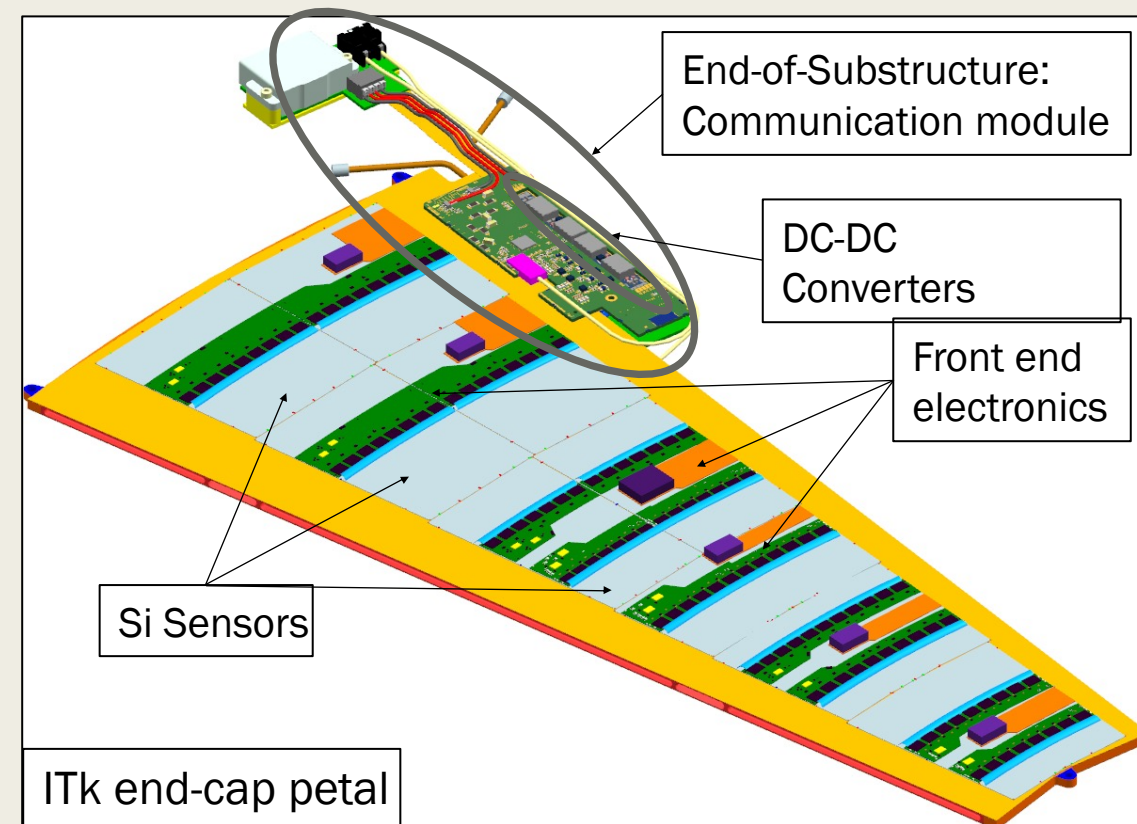
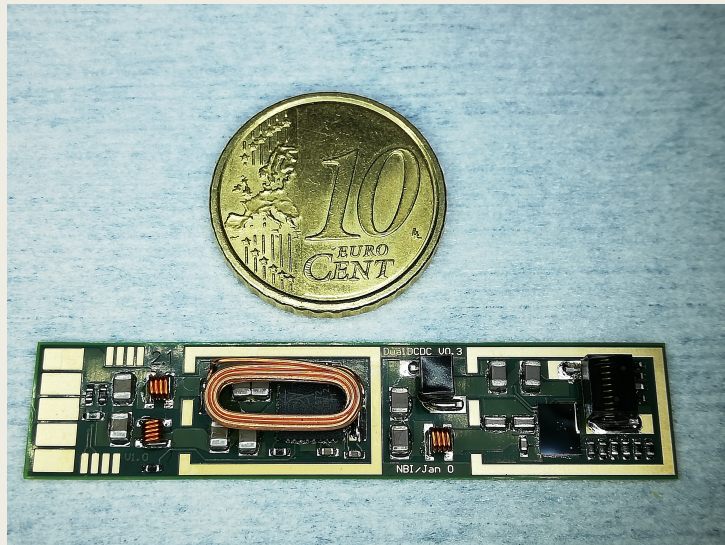
- Hardware implementation (optimal combination of CPU/GPU and possibly FPGAs)
- Algorithm type (neural networks is the option studied now, CNN vs GNN)

# ITK DC-DC converter

NBI

DC-DC Converter is part of the End-of-Substructure communication module , and provides power for this.

- Converting in two steps from input voltage of 11 V first to 2.5 V used for optical chips and further to 1.2 V for electronics
- PCBs (panels of 10 each) by GS Swiss PCB AG, CH
- Automatized component mounting at Danchell A/S, DK
- Functionality testing and burn-in in NBI electronics lab



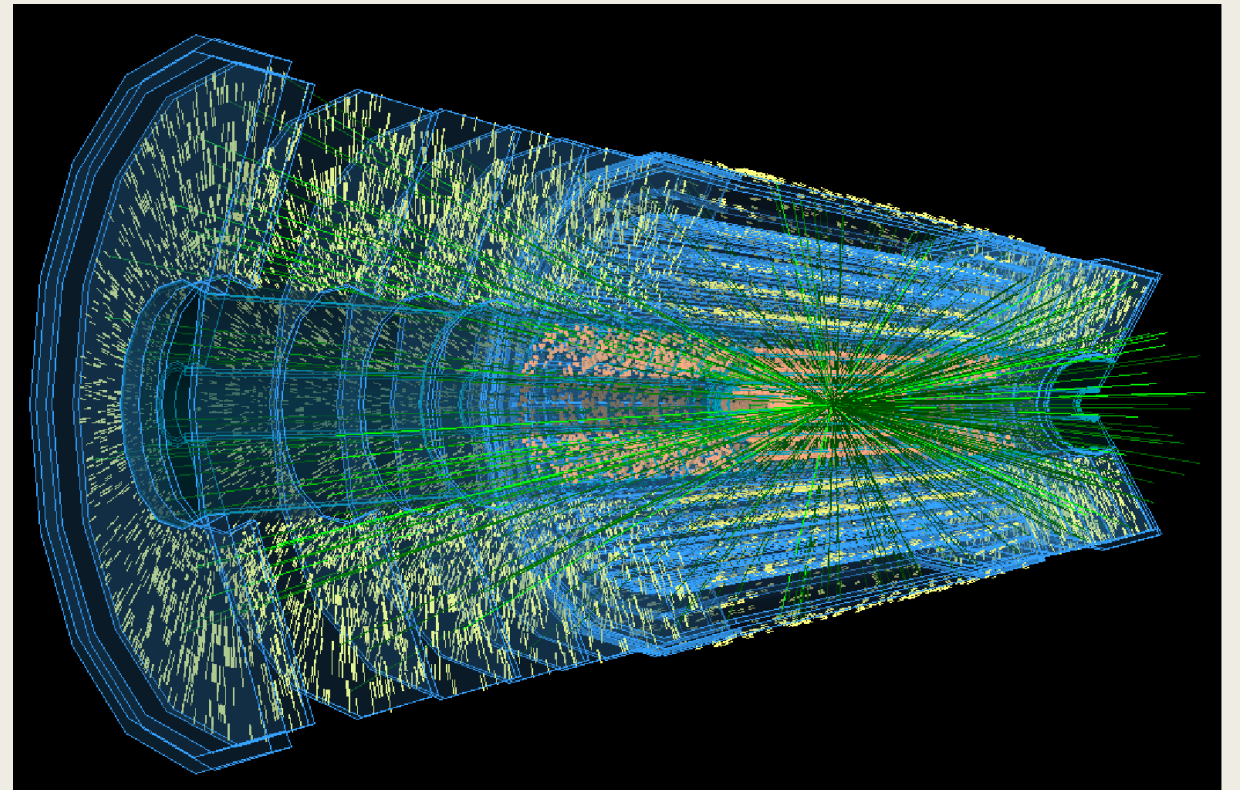


# ITK endcap strip modules

- **600** modules will be produced [ $\sim 10\%$ ] in Scandinavia (in two different geometries)
- Production will run from **2022** to **2025**
- NBI responsible for all module

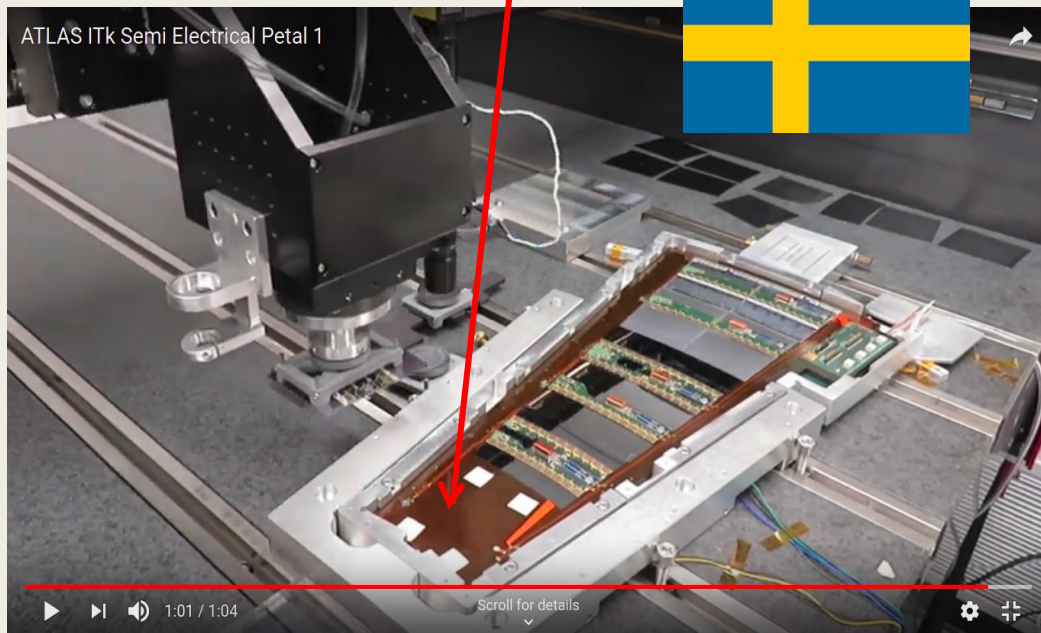
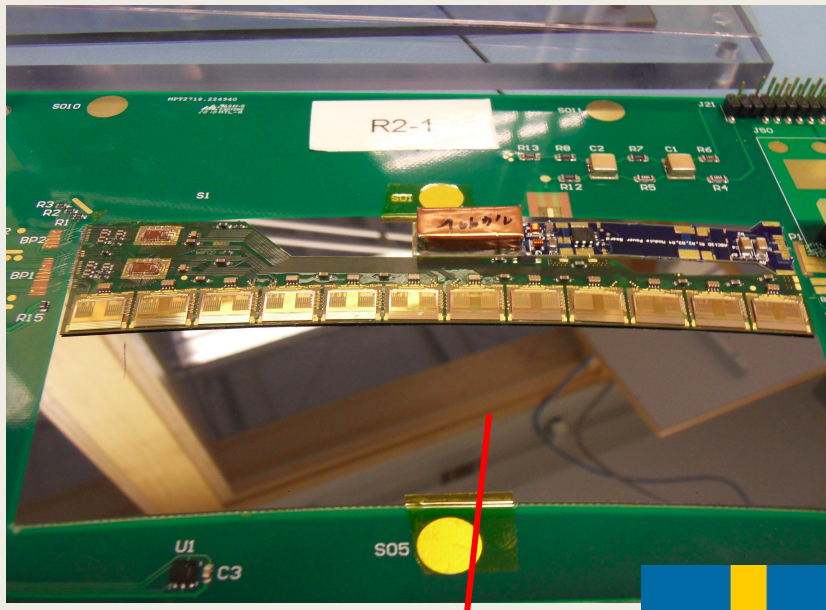
## Quality Control & Assurance

- *visual inspection*
- *electrical testing*
- *metrology*
- *thermal cycling*





# ITK endcap strip modules



New NBI clean-room lab for module production

# Who are we ?

Name	Position	Interests
Peter Hansen	Prof. Emeritus	TRT calibration
Mogens Dam	Ass. Prof.	ITK HL-LHC : DC-DC converters , design & production in Denmark
Stefania Xella	Ass.Prof.	Trigger HL-LHC : optimal algorithms for tracking, design of final Trigger system Tau leptons reconstruction
Jørgen Beck Hansen	Ass.Prof.	Lepton flavour/number, Lepton universality violation Total cross section measurement
Troels Petersen	Ass.Prof.	Photon/electron reconstruction  Lepton universality violation Higgs rare decays
Oleg Ruchayskyi	Ass. Prof.	Lepton flavour/number violation (STA)
Craig Wiglesworth	Assist. Prof.	ITK HL-LHC: module production & quality control
Alessandra Camplani	Assist. Prof.	Trigger HL-LHC: optimal algorithms for tracking, optimal hardware for tracking
Jimmy Hansen	technician	ITK HL-LHC : DC-DC converters, technical help in testing
	Master students	Lepton flavour violation, photon energy reconstruction, lepton universality violation

# Summary

- LHC and ATLAS are an excellent machine and experiment for investigating fundamental new particle and forces at the energy frontier.
- The performance is exceeding what expected for this stage of the physics program thanks to the excellent accelerator team, the inventiveness of the experimental physicists (new software or hardware, new analysis methods) and the tight collaboration of the theoretical and experimental physicists
  - *We are zooming in on the Higgs particle*
  - *We are testing theory extensions of the SM, looking for the reasons for neutrino mass, dark matter, fine tuning of the SM*
- The ATLAS NBI team contributes strongly to these achievements



Back up

# Which responsibilities did/do we have?

- Stefania : ATLAS NBI team leader, representant in TDAQ IB
  - *Exotics subgroup convenor for searches for new physics with leptons + X 2020-2022*
  - *Trigger coordinator 2017-2019*
  - *Tau Trigger & Tau identification coordinator during Run 1 (<2015)*
- Mogens : ATLAS NBI deputy team leader, representant in ITK IB
  - *Responsible for DC-DC converters for the whole endcap strips detector, HL-LHC, ITK*
- Alessandra : ATLAS representant in TDAQ IB
  - *Firmware coordinator, HL-LHC, Trigger (EF)*
- Peter :
  - *TRT project leader 2019-2021*
- Craig :
  - *Trigger Calorimeter reconstruction coordinator 2014-2016*
  - *Overall coordinator for Quality control & assurance procedures, and responsible for scandinavian endcap strip modules. HL-LHC, ITK*
- Jørgen:
  - *Total cross section  $pp$  , 8 & 13 TeV , analysis coordinator*

# supersymmetry

## ATLAS SUSY Searches\* - 95% CL Lower Limits

March 2022

ATLAS Preliminary

$\sqrt{s} = 13$  TeV

Model	Signature	$\int \mathcal{L} dt$ [fb $^{-1}$ ]	Mass limit	Reference		
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0 $e, \mu$ mono-jet	$E_T^{miss}$ 139 $E_T^{miss}$ 139	$\tilde{q}$ [1x, 8x Degen.] 1.0 $\tilde{q}$ [8x Degen.] 0.9	$m(\tilde{\chi}_1^0) < 400$ GeV $m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV	2010.14293 2102.10874
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 $e, \mu$ 2-6 jets	$E_T^{miss}$ 139	$\tilde{g}$ 2.3 Forbidden 1.15-1.95	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{g}) = 1000$ GeV	2010.14293 2010.14293
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}W\tilde{\chi}_1^0$	1 $e, \mu$ 2-6 jets	$E_T^{miss}$ 139	$\tilde{g}$ 2.2	$m(\tilde{\chi}_1^0) < 600$ GeV	2101.01629
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(l)\tilde{\chi}_1^0$	$ee, \mu\mu$ 2 jets	$E_T^{miss}$ 139	$\tilde{g}$ 2.2	$m(\tilde{\chi}_1^0) < 700$ GeV	CERN-EP-2022-014
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	0 $e, \mu$ 7-11 jets	$E_T^{miss}$ 139	$\tilde{g}$ 1.97	$m(\tilde{\chi}_1^0) < 600$ GeV	2008.06032
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	SS $e, \mu$ 6 jets	$E_T^{miss}$ 139	$\tilde{g}$ 1.15	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV	1909.08457
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 $e, \mu$ 3 $b$	$E_T^{miss}$ 79.8	$\tilde{g}$ 2.25	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2018-041
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	SS $e, \mu$ 6 jets	$E_T^{miss}$ 139	$\tilde{g}$ 1.25	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV	1909.08457
3 <sup>rd</sup> gen. squarks direct production	$\tilde{b}_1\tilde{b}_1$	0 $e, \mu$ 2 $b$	$E_T^{miss}$ 139	$\tilde{b}_1$ 1.255 $\tilde{b}_1$ 0.68	$m(\tilde{\chi}_1^0) < 400$ GeV $10 \text{ GeV} < \Delta m(\tilde{b}_1, \tilde{\chi}_1^0) < 20$ GeV	2101.12527 2101.12527
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_2^0 \rightarrow bh\tilde{\chi}_1^0$	0 $e, \mu$ 6 $b$	$E_T^{miss}$ 139	$\tilde{b}_1$ 0.23-1.35	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV	1908.03122
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_2^0 \rightarrow bh\tilde{\chi}_1^0$	2 $\tau$ 2 $b$	$E_T^{miss}$ 139	$\tilde{b}_1$ 0.13-0.85	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV	2103.08189
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	0-1 $e, \mu$ $\geq 1$ jet	$E_T^{miss}$ 139	$\tilde{t}_1$ 1.25	$m(\tilde{\chi}_1^0) = 1$ GeV	2004.14060, 2012.03799
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$	1 $e, \mu$ 3 jets/1 $b$	$E_T^{miss}$ 139	$\tilde{t}_1$ 0.65	$m(\tilde{\chi}_1^0) = 500$ GeV	2012.03799
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{\tau}_1 b\nu, \tilde{\tau}_1 \rightarrow \tau\tilde{G}$	1-2 $\tau$ 2 jets/1 $b$	$E_T^{miss}$ 139	$\tilde{t}_1$ 1.4	$m(\tilde{\tau}_1) = 800$ GeV	2108.07665
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$	0 $e, \mu$ 2 $c$	$E_T^{miss}$ 36.1	$\tilde{t}_1$ 0.85	$m(\tilde{\chi}_1^0) = 0$ GeV	1805.01649
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$	0 $e, \mu$ mono-jet	$E_T^{miss}$ 139	$\tilde{t}_1$ 0.55	$m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 5$ GeV	2102.10874
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z/h\tilde{\chi}_1^0$	1-2 $e, \mu$ 1-4 $b$	$E_T^{miss}$ 139	$\tilde{t}_1$ 0.067-1.18	$m(\tilde{\chi}_2^0) = 500$ GeV	2006.05880
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow t_1 + Z$	3 $e, \mu$ 1 $b$	$E_T^{miss}$ 139	$\tilde{t}_2$ 0.86	$m(\tilde{t}_1) = 360$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 40$ GeV	2006.05880
EW direct	$\tilde{\chi}_1^+ \tilde{\chi}_2^0$ via WZ	Multiple $\ell$ /jets $ee, \mu\mu$	$E_T^{miss}$ 139 $E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_2^0$ 0.96 $\tilde{\chi}_1^+ / \tilde{\chi}_2^0$ 0.205	$m(\tilde{\chi}_1^0) = 0$ , wino-bino $m(\tilde{\chi}_1^+) - m(\tilde{\chi}_1^0) = 5$ GeV, wino-bino	2106.01676, 2108.07586 1911.12606
	$\tilde{\chi}_1^+ \tilde{\chi}_2^0$ via WW	2 $e, \mu$	$E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_2^0$ 0.42	$m(\tilde{\chi}_1^0) = 0$ , wino-bino	1908.08215
	$\tilde{\chi}_1^+ \tilde{\chi}_2^0$ via Wh	Multiple $\ell$ /jets	$E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_2^0$ 1.06	$m(\tilde{\chi}_1^0) = 70$ GeV, wino-bino	2004.10894, 2108.07586
	$\tilde{\chi}_1^+ \tilde{\chi}_2^0$ via $\tilde{\ell}_1/\tilde{\nu}$	2 $e, \mu$	$E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_2^0$ 1.0	$m(\tilde{\ell}_1, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^+) + m(\tilde{\chi}_1^0))$	1908.08215
	$\tilde{\tau}_1, \tilde{\tau}_1 \rightarrow \tau\tilde{\chi}_1^0$	2 $\tau$	$E_T^{miss}$ 139	$\tilde{\tau}_1$ 0.16-0.31 0.12-0.39	$m(\tilde{\chi}_1^0) = 0$	1911.06660
	$\tilde{\ell}_{1,R}\tilde{\ell}_{1,R}, \tilde{\ell} \rightarrow \tilde{\chi}_1^0$	2 $e, \mu$ $ee, \mu\mu$	$E_T^{miss}$ 139 $E_T^{miss}$ 139	$\tilde{\ell}$ 0.256 $\tilde{\ell}$ 0.7	$m(\tilde{\chi}_1^0) = 0$ $m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV	1908.08215 1911.12606
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 $e, \mu$ 4 $e, \mu$ 0 jets	$E_T^{miss}$ 36.1 $E_T^{miss}$ 139 $E_T^{miss}$ 139	$\tilde{H}$ 0.13-0.23 $\tilde{H}$ 0.55 $\tilde{H}$ 0.29-0.88 $\tilde{H}$ 0.45-0.93	$BR(\tilde{H} \rightarrow h\tilde{G}) = 1$ $BR(\tilde{H} \rightarrow Z\tilde{G}) = 1$ $BR(\tilde{H} \rightarrow Z\tilde{G}) = 1$	1806.04030 2103.11684 2108.07586
	Direct $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^+$	Disapp. trk 1 jet	$E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_1^-$ 0.21 $\tilde{\chi}_1^+ / \tilde{\chi}_1^-$ 0.66	Pure Wino Pure higgsino	2201.02472 2201.02472
	Stable $\tilde{g}$ R-hadron	pixel dE/dx	$E_T^{miss}$ 139	$\tilde{g}$ 2.05	$m(\tilde{\chi}_1^0) = 100$ GeV	CERN-EP-2022-029
	Metastable $\tilde{g}$ R-hadron, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	pixel dE/dx	$E_T^{miss}$ 139	$\tilde{g}$ [ $\tau(\tilde{g}) = 10$ ns ] 2.2	$m(\tilde{\chi}_1^0) = 100$ GeV	CERN-EP-2022-029
$\tilde{\ell}\tilde{\ell}, \tilde{\ell} \rightarrow \ell\tilde{G}$	Displ. lep	$E_T^{miss}$ 139	$\tilde{\ell}$ 0.7	$\tau(\tilde{\ell}) = 0.1$ ns $\tau(\tilde{\ell}) = 0.1$ ns	2011.07812 2011.07812	
$\tilde{\ell}\tilde{\ell}, \tilde{\ell} \rightarrow \ell\tilde{G}$	pixel dE/dx	$E_T^{miss}$ 139	$\tilde{\ell}$ 0.36	$\tau(\tilde{\ell}) = 10$ ns	CERN-EP-2022-029	
RPV	$\tilde{\chi}_1^+ \tilde{\chi}_1^- / \tilde{\chi}_1^0, \tilde{\chi}_1^+ \rightarrow Z\ell - \ell\ell$	3 $e, \mu$	$E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_1^-$ [BR(Z $\tau$ )=1, BR(Z $e$ )=1] 0.625 $\tilde{\chi}_1^+ / \tilde{\chi}_1^-$ [ $A_{33} \neq 0, A_{124} \neq 0$ ] 1.05	Pure Wino $m(\tilde{\chi}_1^0) = 200$ GeV	2011.10543 2103.11684
	$\tilde{\chi}_1^+ \tilde{\chi}_1^- / \tilde{\chi}_1^0 \rightarrow WWZZ\ell\ell\nu\nu$	4 $e, \mu$ 0 jets	$E_T^{miss}$ 139	$\tilde{\chi}_1^+ / \tilde{\chi}_1^-$ 0.95 $\tilde{\chi}_1^+ / \tilde{\chi}_1^-$ [ $m(\tilde{\chi}_1^0) = 200$ GeV, 1100 GeV ] 1.3 1.9	Large $A'_{12}$	1804.03568
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq$	4-5 large jets	36.1	$\tilde{g}$ 0.55 $\tilde{g}$ 1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	ATLAS-CONF-2018-003
	$\tilde{u}, \tilde{t} \rightarrow b\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs$	Multiple $\geq 4b$	36.1 139	$\tilde{u}$ 0.95 Forbidden	$m(\tilde{\chi}_1^0) = 500$ GeV	2010.01015
	$\tilde{u}\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	2 jets + 2 $b$	36.7	$\tilde{u}$ [ $gq, b\nu$ ] 0.42 0.61		1710.07171
	$\tilde{u}\tilde{t}_1, \tilde{t}_1 \rightarrow q\ell$	2 $e, \mu$ 2 $b$	36.1	$\tilde{u}$ 0.4-1.45	$BR(\tilde{u} \rightarrow b\ell/h\nu) > 20\%$	1710.05544
$\tilde{u}\tilde{t}_1, \tilde{t}_1 \rightarrow q\ell$	1 $\mu$ DV	136	$\tilde{u}$ [ $1e-10 < A'_{33} < 1e-8, 3e-10 < A'_{33} < 3e-9$ ] 1.0 1.6	$BR(\tilde{u} \rightarrow q\nu) = 100\%$ , $\cos\theta_0 = 1$	2003.11956	
$\tilde{\chi}_1^0 / \tilde{\chi}_2^0 / \tilde{\chi}_1^+, \tilde{\chi}_1^0 \rightarrow tbs, \tilde{\chi}_1^+ \rightarrow bbs$	1-2 $e, \mu$ $\geq 6$ jets	$E_T^{miss}$ 139	$\tilde{\chi}_1^0$ 0.2-0.32	Pure higgsino	2106.09609	

\*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10<sup>-1</sup>

1

Mass scale [TeV]

# exotics

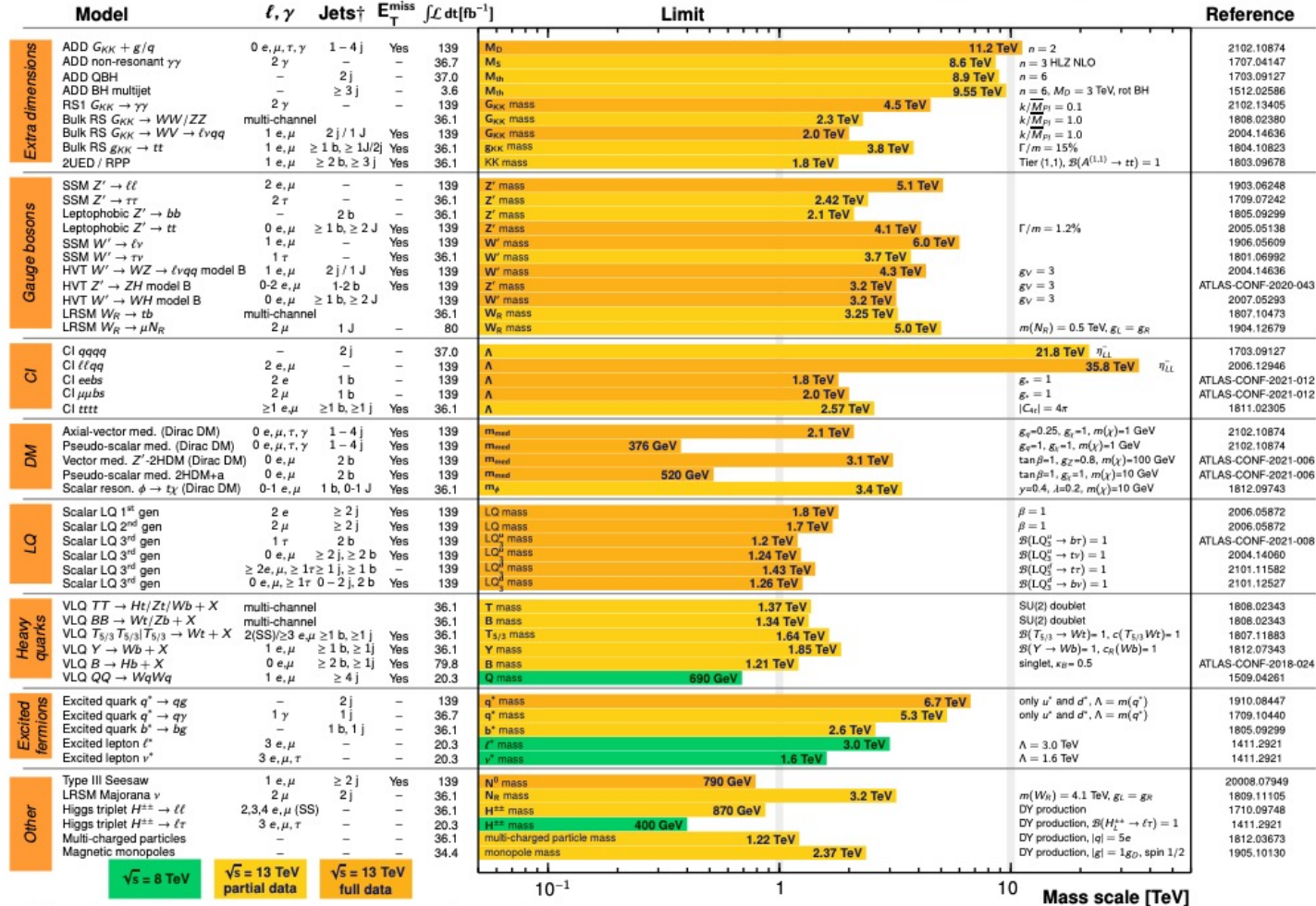
## ATLAS Exotics Searches\* - 95% CL Upper Exclusion Limits

Status: March 2021

ATLAS Preliminary

$\int \mathcal{L} dt = (3.6 - 139) \text{ fb}^{-1}$

$\sqrt{s} = 8, 13 \text{ TeV}$



\*Only a selection of the available mass limits on new states or phenomena is shown.

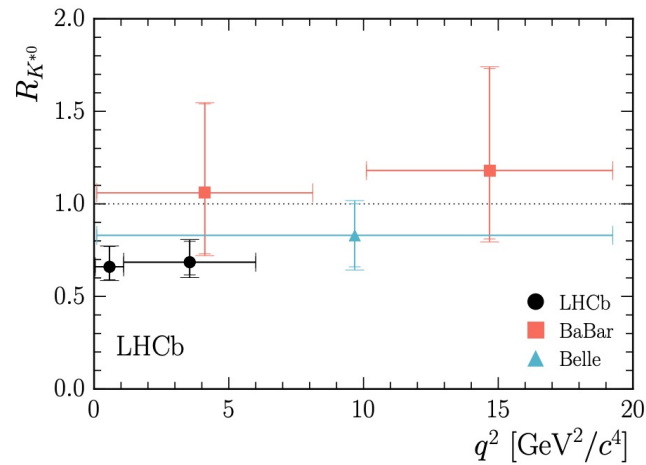
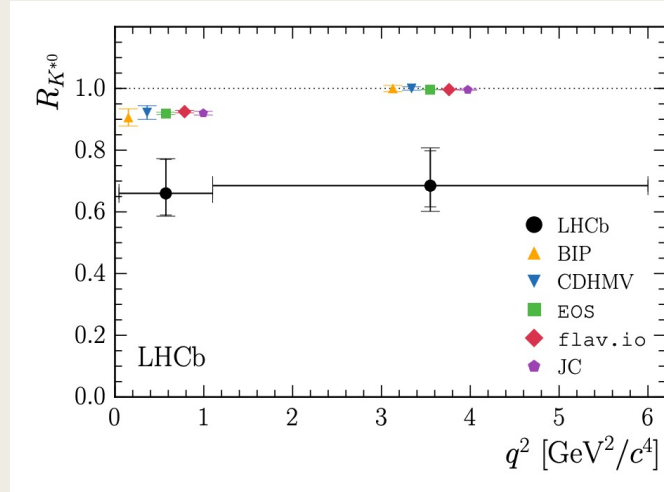
†Small-radius (large-radius) jets are denoted by the letter j (J).

$\sqrt{s} = 8 \text{ TeV}$   $\sqrt{s} = 13 \text{ TeV}$  partial data  $\sqrt{s} = 13 \text{ TeV}$  full data

10<sup>-1</sup> 1 10 Mass scale [TeV]

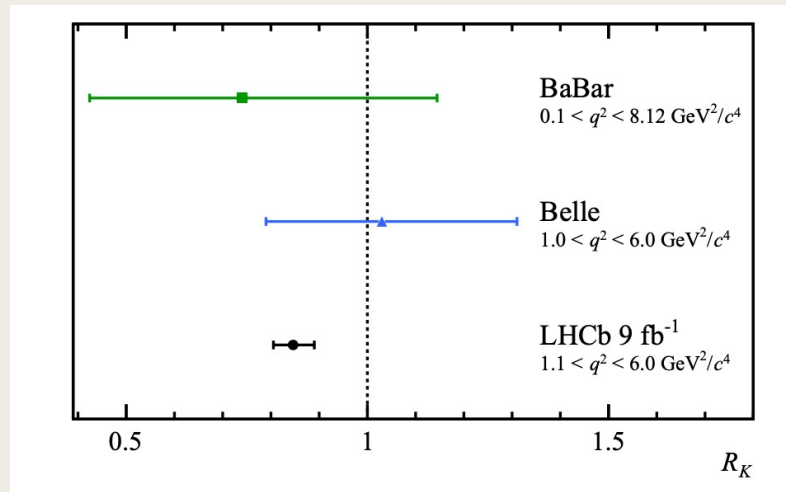


# B anomalies



LHCb Run 1 , 3 fb-1 [[lhcb1](#)]

LHCb Run 2, 9 fb-1 [[lhcb2](#)]

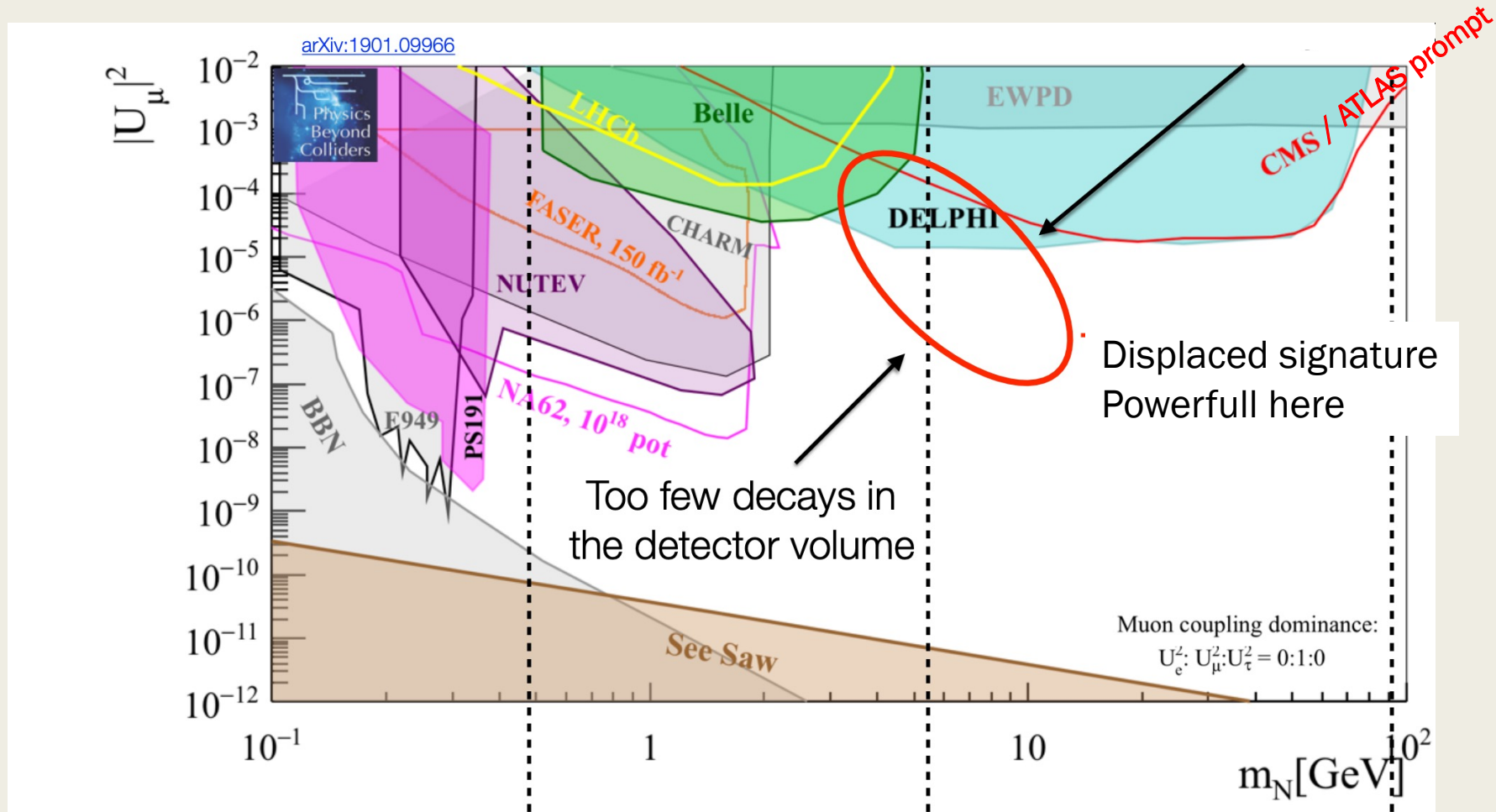


LHCb Run 2, 9 fb-1 [[lhcb3](#)]

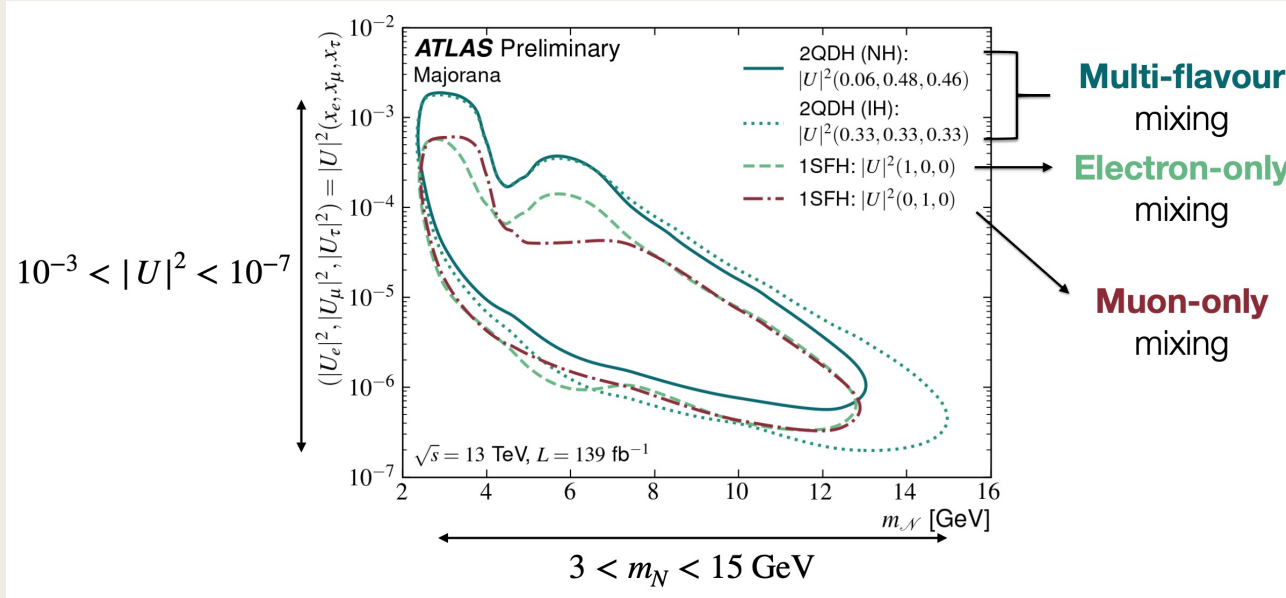
$$R_{K_S^0} = 0.66^{+0.20}_{-0.14} (\text{stat.})^{+0.02}_{-0.04} (\text{syst.}),$$

$$R_{K^{*+}} = 0.70^{+0.18}_{-0.13} (\text{stat.})^{+0.03}_{-0.04} (\text{syst.}).$$

# Heavy neutral leptons



# Heavy neutral leptons



Full Run 2 result - displaced signature search

