# Simulating AGN jet feedback in galaxy clusters



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- Jet propagation
- Feedback effects at t < t<sub>cool</sub>
- Self-regulation
- Multi-phase gas

Jet



Jet's head & contact discontinuity



Bromberg et al. (2011)





RW et al. (in prep.)





#### Feedback mechanism





50 Myr





rw jet 2

rw heavyjet

#### Self-regulated heating & cooling





### Simulations of galaxy clusters



Ehlert et al. (arXiv:2204.01765)

 $R_{200} = 1031.9 \text{ kpc/h}$   $M_{200} = 1.2e+14 \text{ M}_{\odot}/\text{h}$   $M_{stars} = 3.4e+11 \text{ M}_{\odot}/\text{h}$ 



 $10^1$   $10^2$   $10^3$   $10^4$ Entropy (keV cm<sup>2</sup>)

Box: 800 kpc



 $10^{-3}$ Column density (g/cm<sup>2</sup>)



Jlassi et al. (in prep.)

#### Multi-phase medium









### 2-fluid hydrodynamics



#### Multi-phase outflows



velocity

Mass exchange between phases Source term



phase 1





#### Jet propagation in the interstellar medium



• Flux in hot phase

- Evaporation rate
- Flux in cold phase

Mukherjee et al. (2018)

#### Multi-phase intra-cluster medium



Nelson et al. (2020)

## Cosmological simulations



#### Summary

Jets in isolation: immediate feedback effect via central cocoon

• Light vs. heavy jet: momentum flux matters for feedback effects

Self-regulated jet simulations: non-uniqueness of AGN feedback

• Correct star formation rate does not imply correct effects on gas structure

Multi-phase medium: 2-fluid modeling for unresolved structure

• Connect cosmological simulations to small-scale studies

#### Modeling of resolved multi-fluid flows



### Modeling of resolved multi-fluid flows



RW & Hernquist (arXiv:2204.05316)