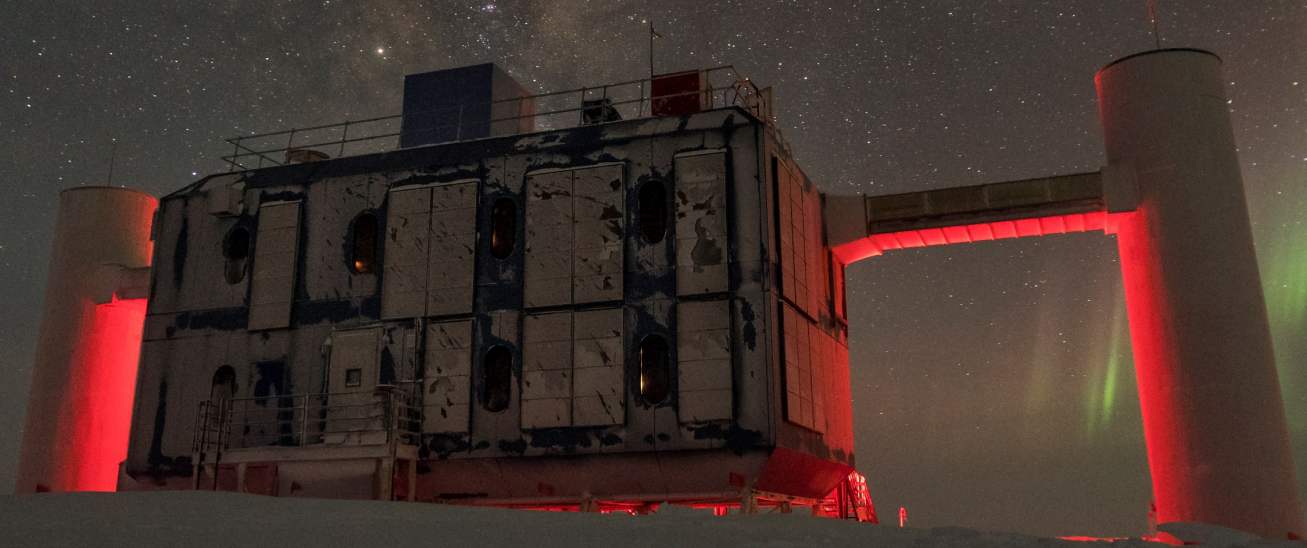




ICECUBE

MASTERCLASS

WORKSHOP ACTIVITY #2: LOOKING FOR NEUTRINO SOURCES

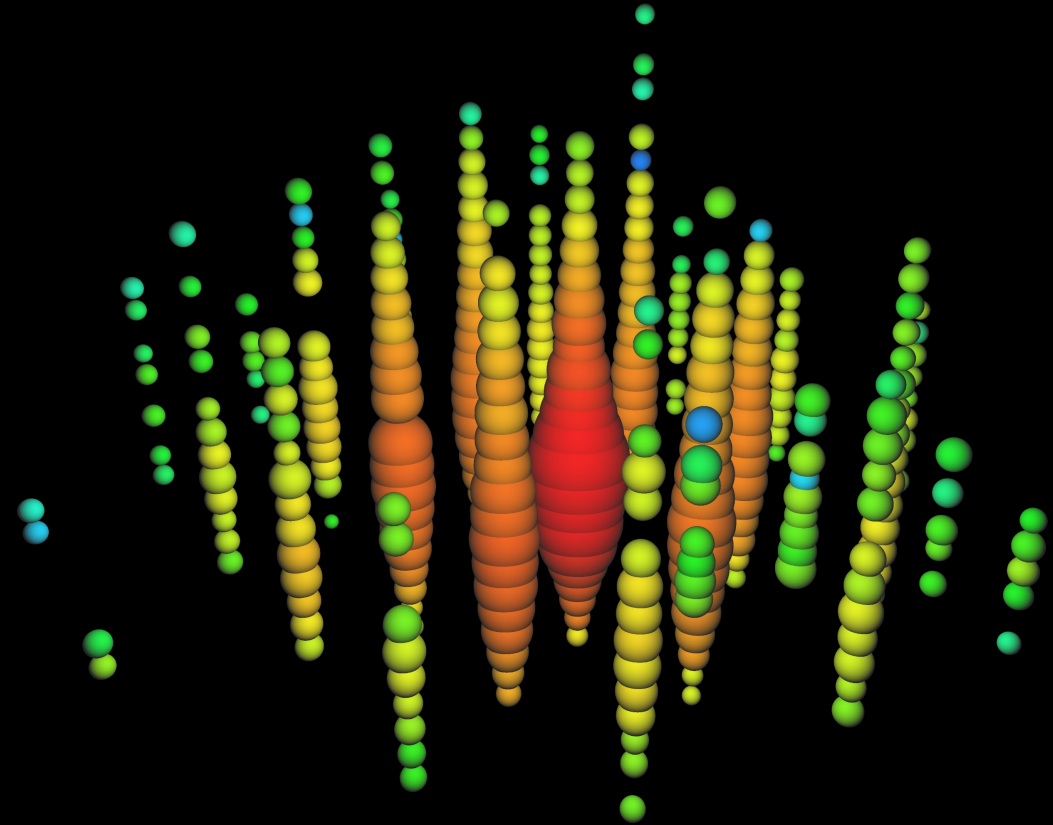


BEFORE LUNCH

We figured out how to find neutrinos in our data.

The questions we want to answer are:

- What can produce these neutrinos?
- Where do these neutrinos come from?



NEUTRINO PRODUCTION IN THE UNIVERSE

The electron-volt (eV) energy scale

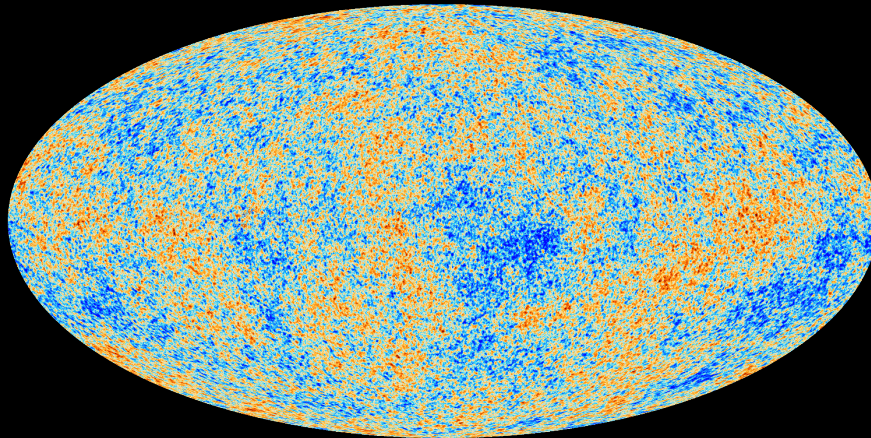
Low energy

High energy

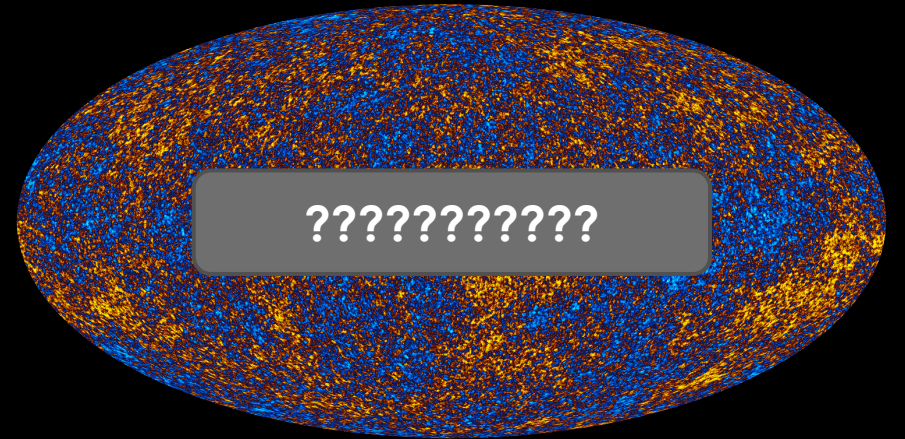


0.001
(meV)

The Cosmic
Neutrino
Background



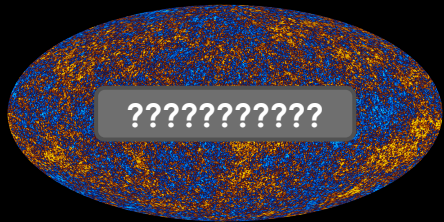
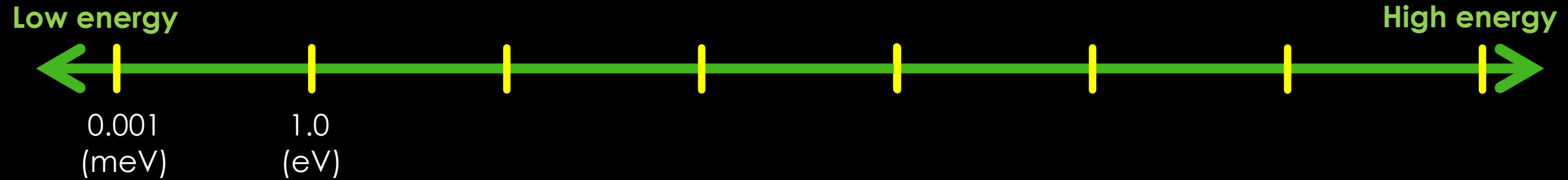
CMB (photons)



CνB (Neutrinos)

NEUTRINO PRODUCTION IN THE UNIVERSE

The electron-volt (eV) energy scale



The Cosmic Neutrino Background

Nothing to see!

NEUTRINO PRODUCTION IN THE UNIVERSE

The electron-volt (eV) energy scale

Low energy

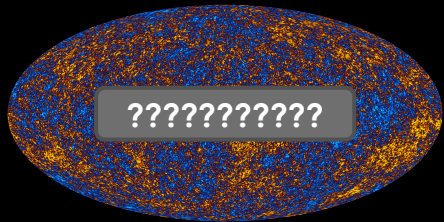
High energy



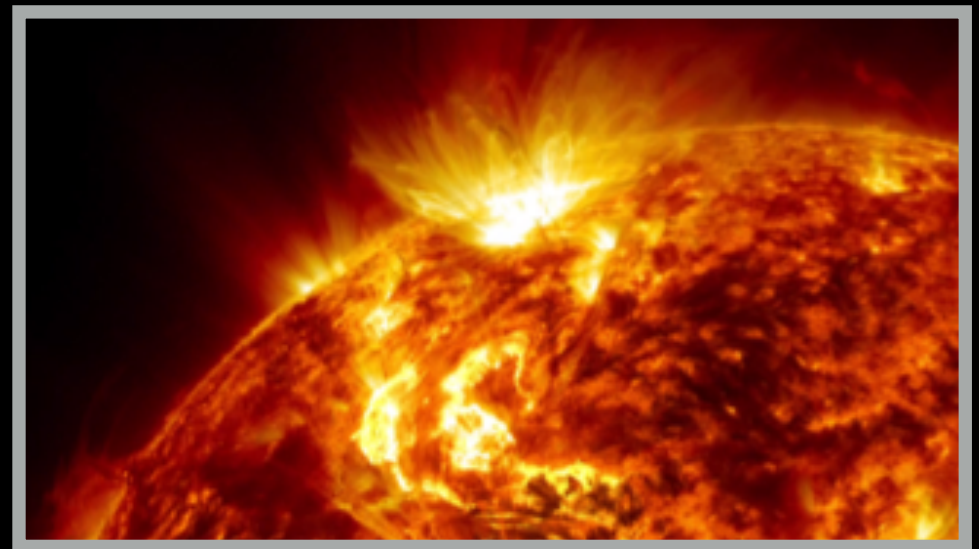
0.001
(meV)

1.0
(eV)

1000
(keV)



The Cosmic
Neutrino
Background



Solar Neutrinos

NEUTRINO PRODUCTION IN THE UNIVERSE

The electron-volt (eV) energy scale

Low energy

High energy

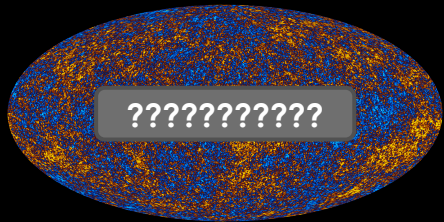


0.001
(meV)

1.0
(eV)

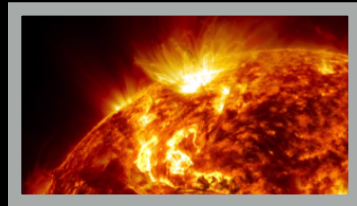
1000
(keV)

10^6
(MeV)

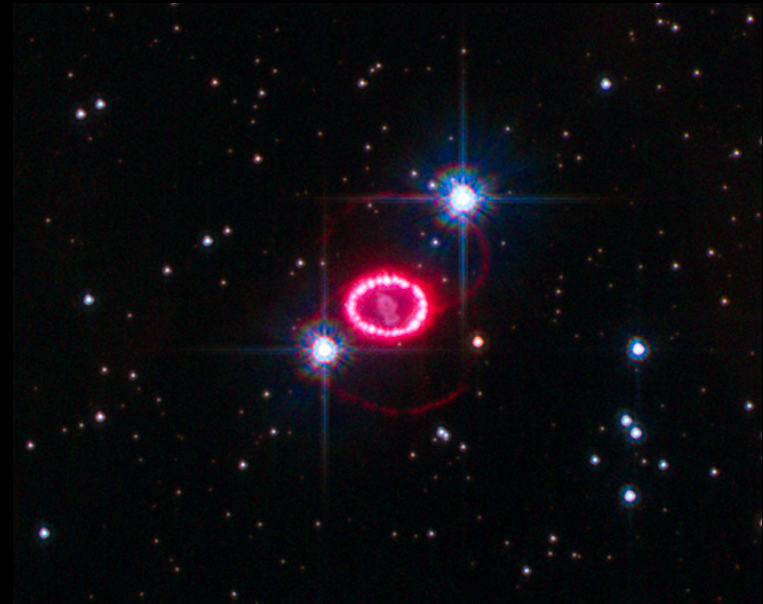


??????????

The Cosmic
Neutrino
Background



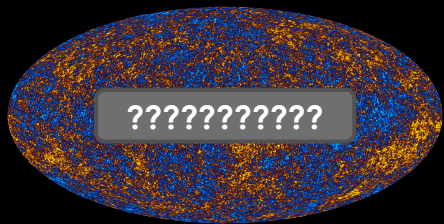
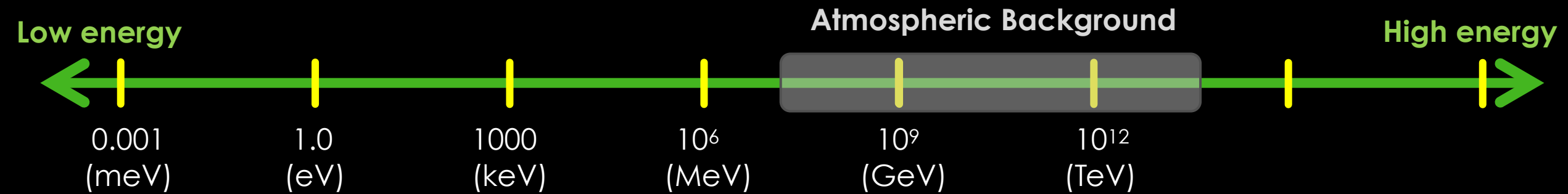
Solar Neutrinos



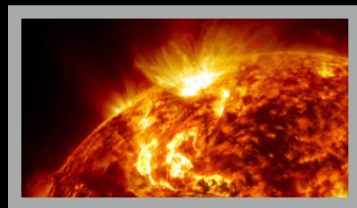
SUPERNOVA!

NEUTRINO PRODUCTION IN THE UNIVERSE

The electron-volt (eV) energy scale



The Cosmic Neutrino Background



Solar Neutrinos



Supernova

NEUTRINO PRODUCTION IN THE UNIVERSE

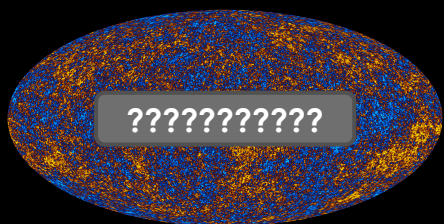
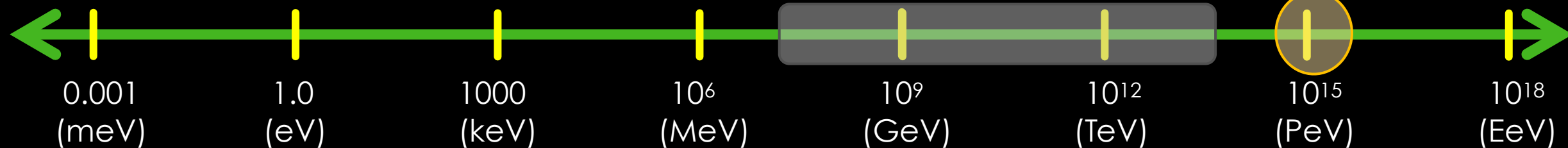
The electron-volt (eV) energy scale

Highest energy
seen by
IceCube

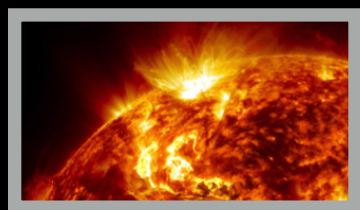
Atmospheric Background

High energy

Low energy



The Cosmic
Neutrino
Background



Solar Neutrinos



Supernova

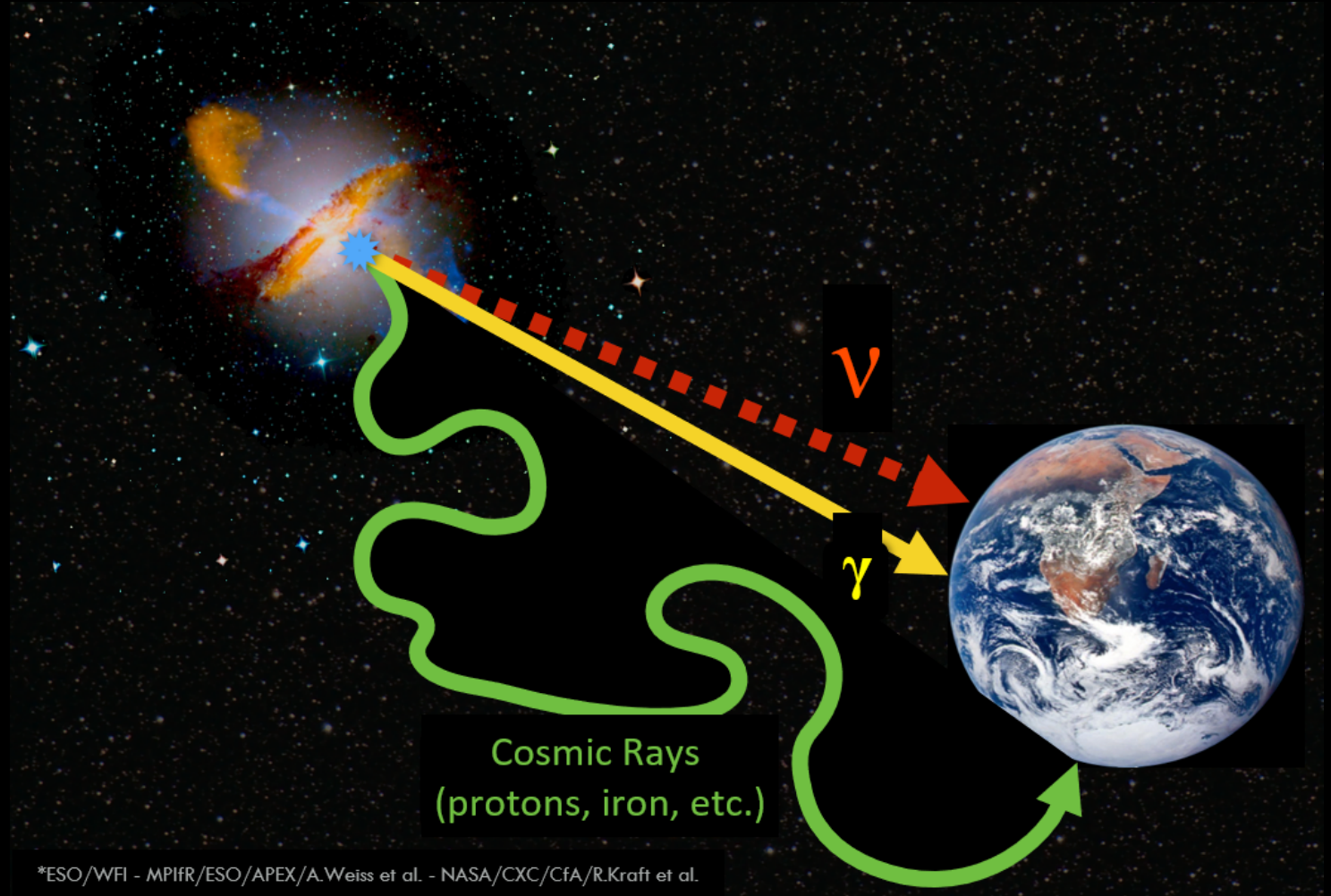


WE DON'T KNOW!

BACK TO NEUTRINO ASTRONOMY

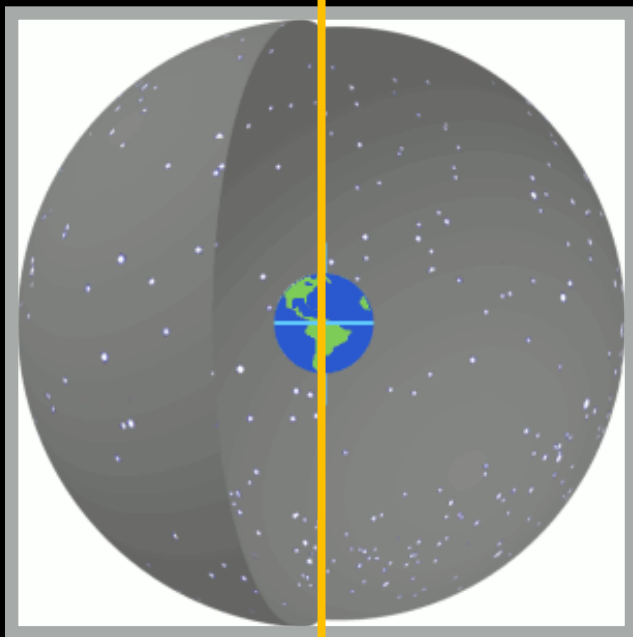
THE PRINCIPLES OF NEUTRINO ASTRONOMY

1. Very High-Energy neutrinos come from *cosmic accelerators*
2. **Neutrinos travel in a straight line from their source**

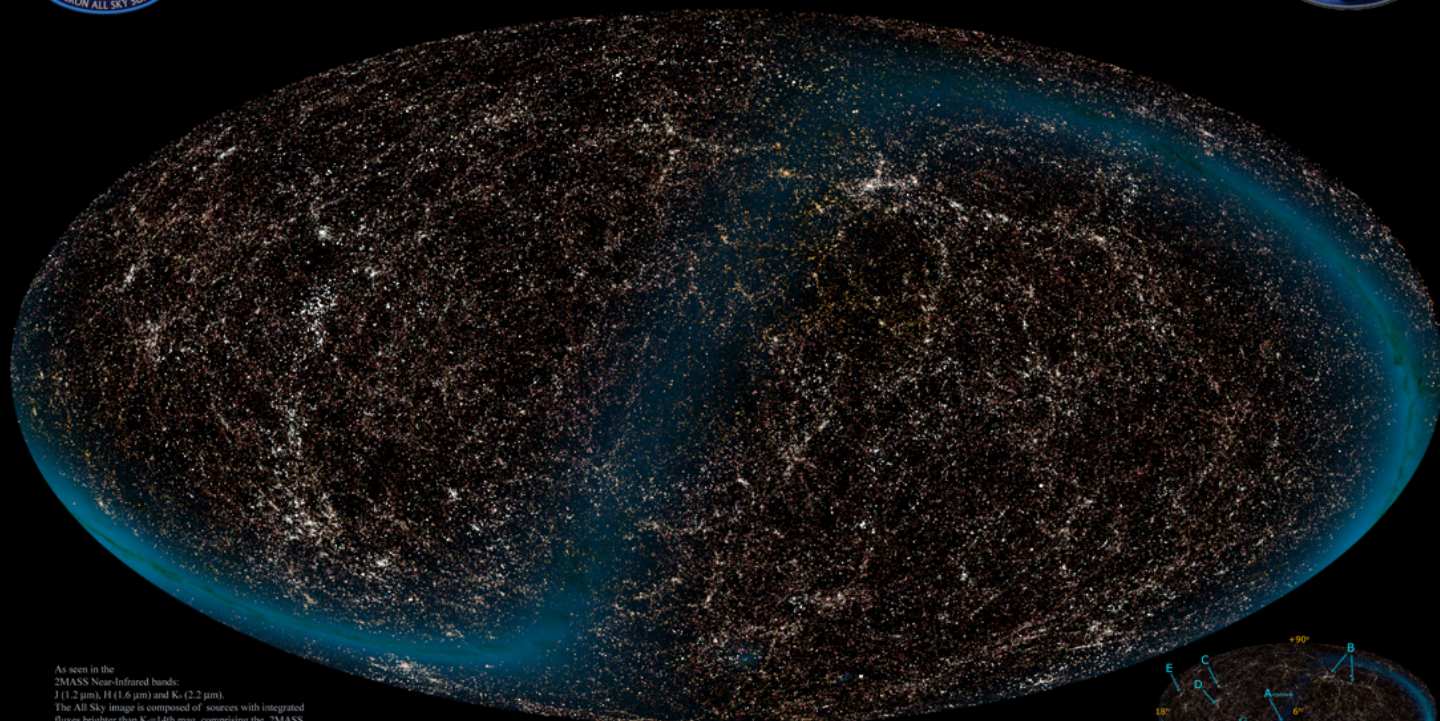


AN INTRODUCTION TO SKYMAPS

Axis of rotation
of the Earth

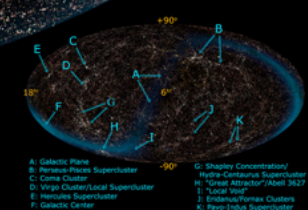


The Near-Infrared Local Universe



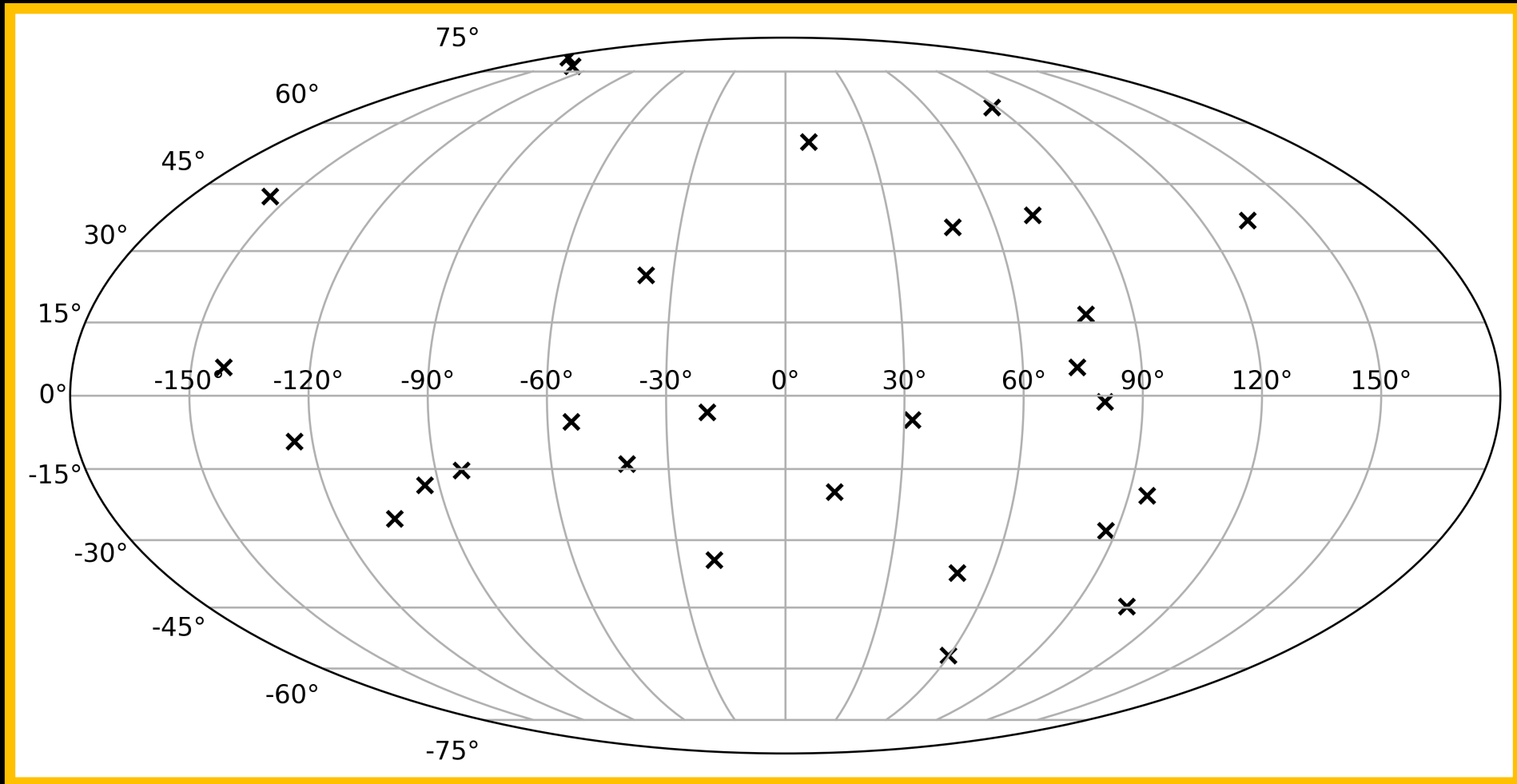
As seen in the 2MASS Near-Infrared bands: J (1.2 μm), H (1.6 μm) and K_s (2.2 μm).
The All Sky image is composed of sources with integrated fluxes brighter than K_s=14th mag, comprising the 2MASS Extended Source Catalog (XSC) -- more than 1.6 million galaxies, and the Point Source Catalog (PSC) -- nearly 0.5 billion Milky Way stars (here tinted in blue to show contrast with the background galaxies.)
The map is projected with an equal area aloff in the Geo-equatorial system (centered at 0 hr Right Ascension). The plane of the Milky Way runs diagonally across the image, with the Galactic anti-center facing you.

The image was created by Thomas Jarrett & Robert Hurt (IPAC/Caltech).

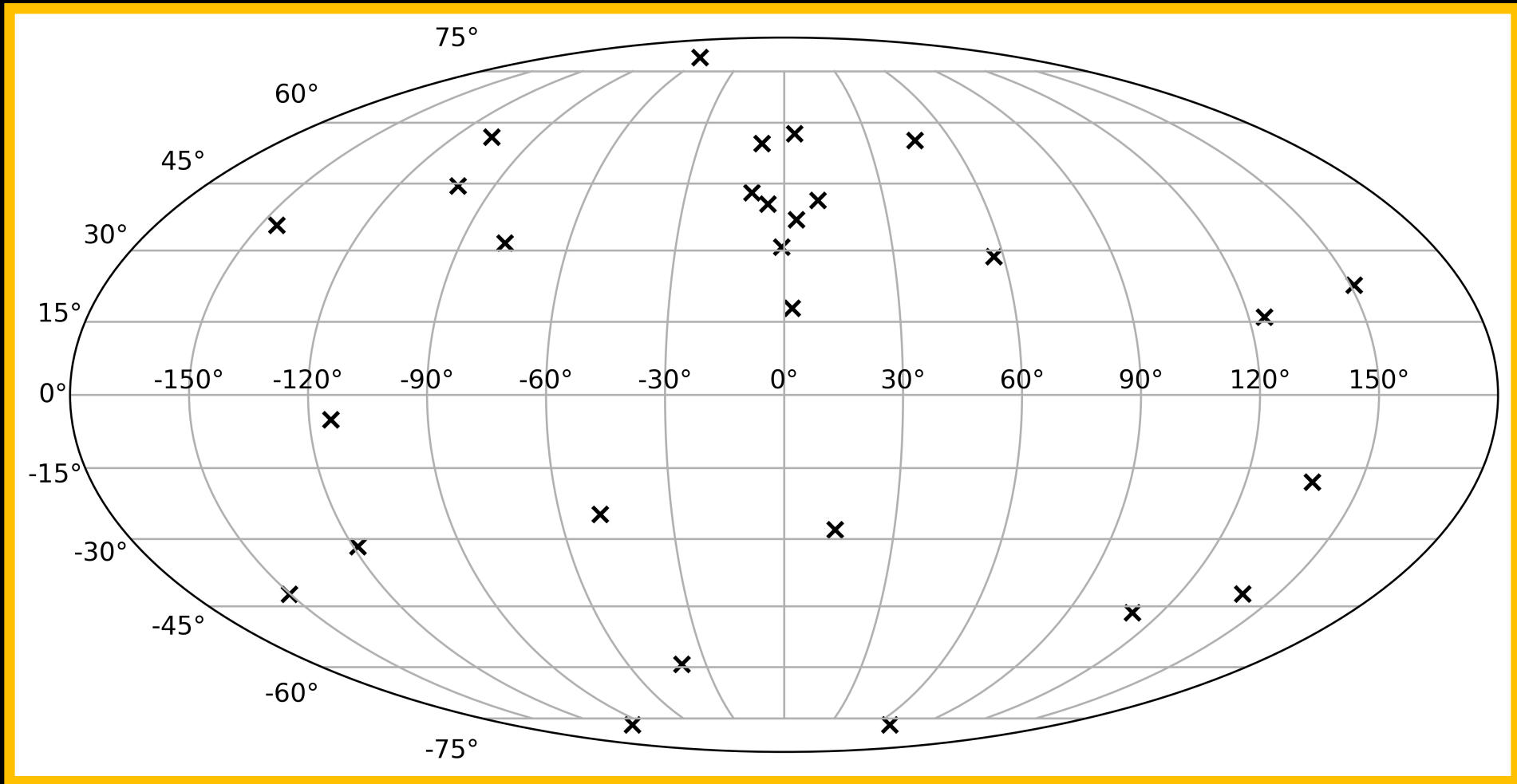


- A: Galactic Plane
- B: Perseus-Pisces Supercluster
- C: Coma Cluster
- D: Virgo Cluster/Local Supercluster
- E: Hercules Supercluster
- F: Galactic Center
- G: Shapley Concentration/ Hydra-Centaurus Supercluster
- H: "Great Attractor" (M83, M27)
- I: "Local Void"
- J: Indus/Taurus Clusters
- K: Pavo-Telus Supercluster

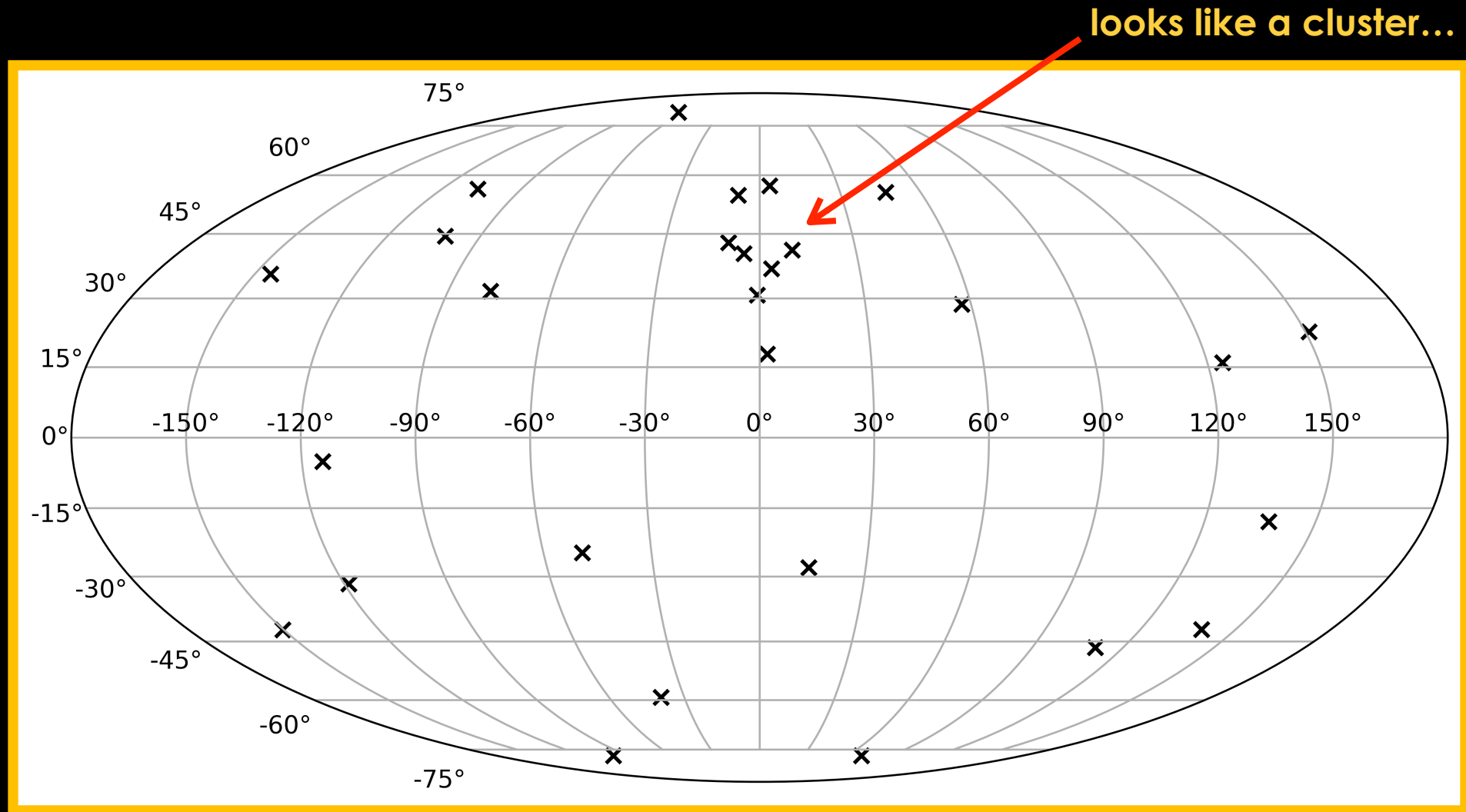
SKYMAP OF 28 EVENTS – RANDOMISED BACKGROUND EVENTS



SKYMAP OF 28 EVENTS – WITH A SIMULATED SOURCE

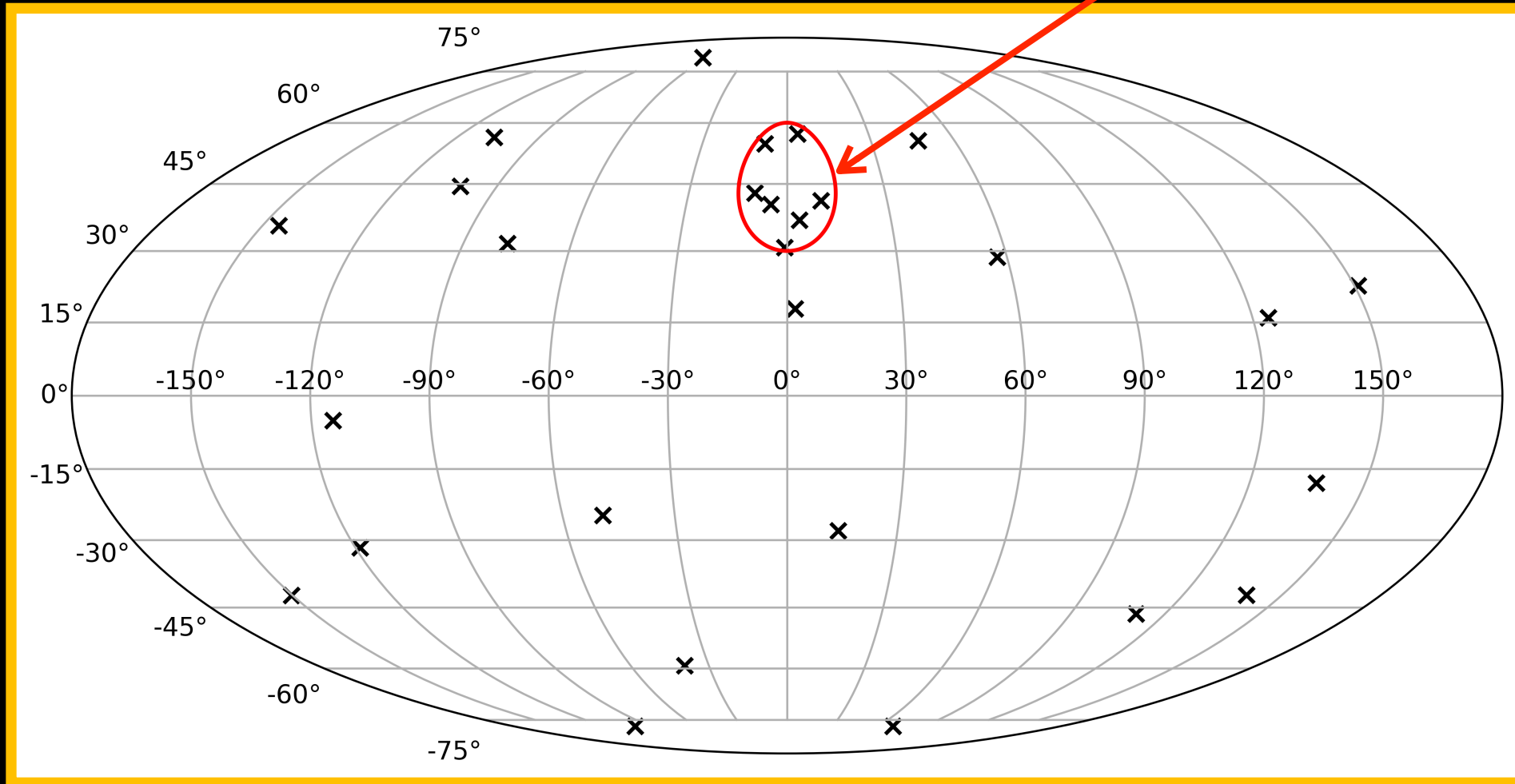


SKYMAP OF 28 EVENTS – WITH A SIMULATED SOURCE



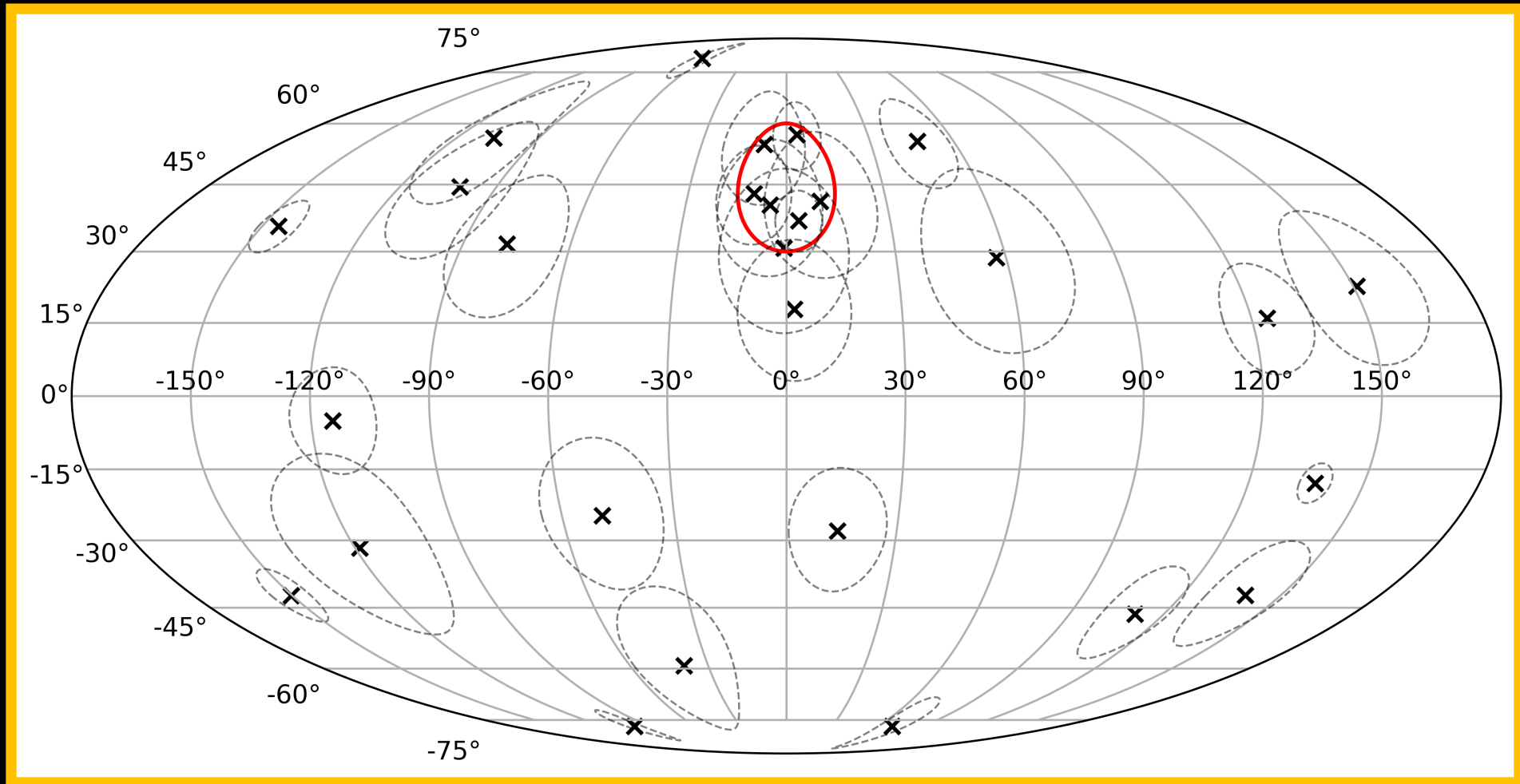
SKYMAP OF 28 EVENTS – WITH A SIMULATED SOURCE

And maybe it is?..



SKYMAP OF 28 EVENTS – WITH A SIMULATED SOURCE

Or were we just lucky to find 7 events inside this particular circle?

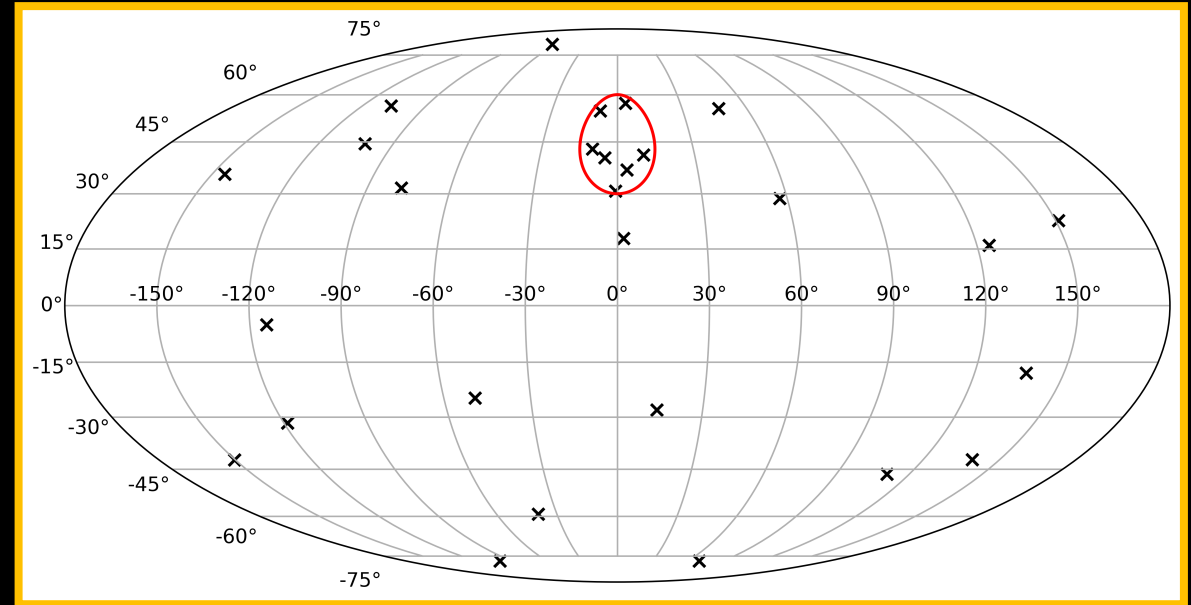


ACTIVITY!

- Is this a significant signal?
- We want to evaluate the likelihood that this clustering comes from a single source in the sky

In your group:

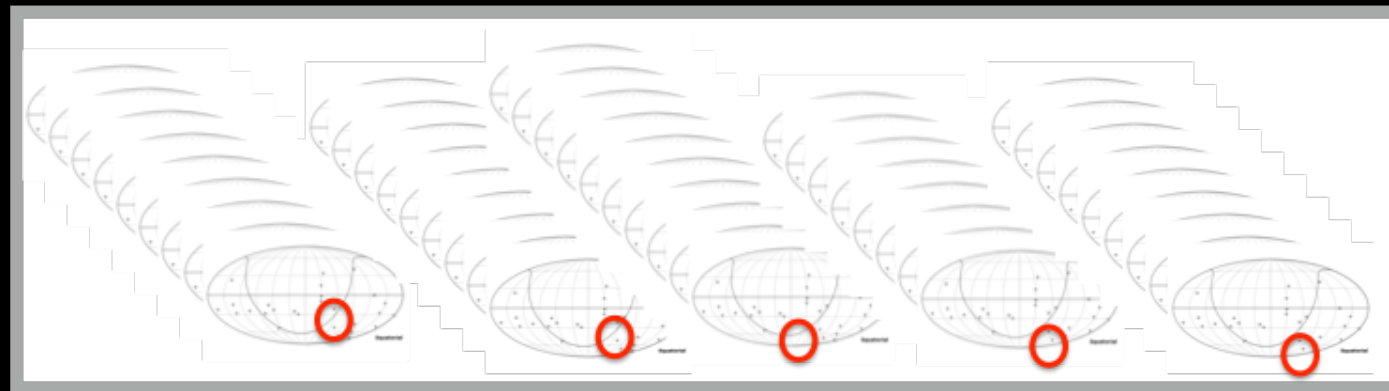
- Can you think of a method we could use to calculate a significance?



THE P-VALUE SIGNIFICANCE

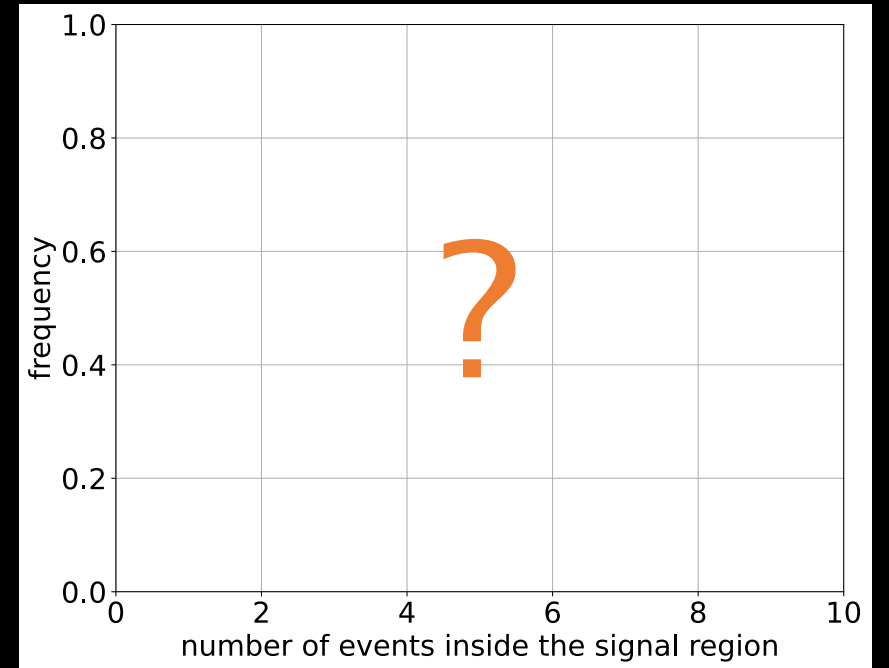
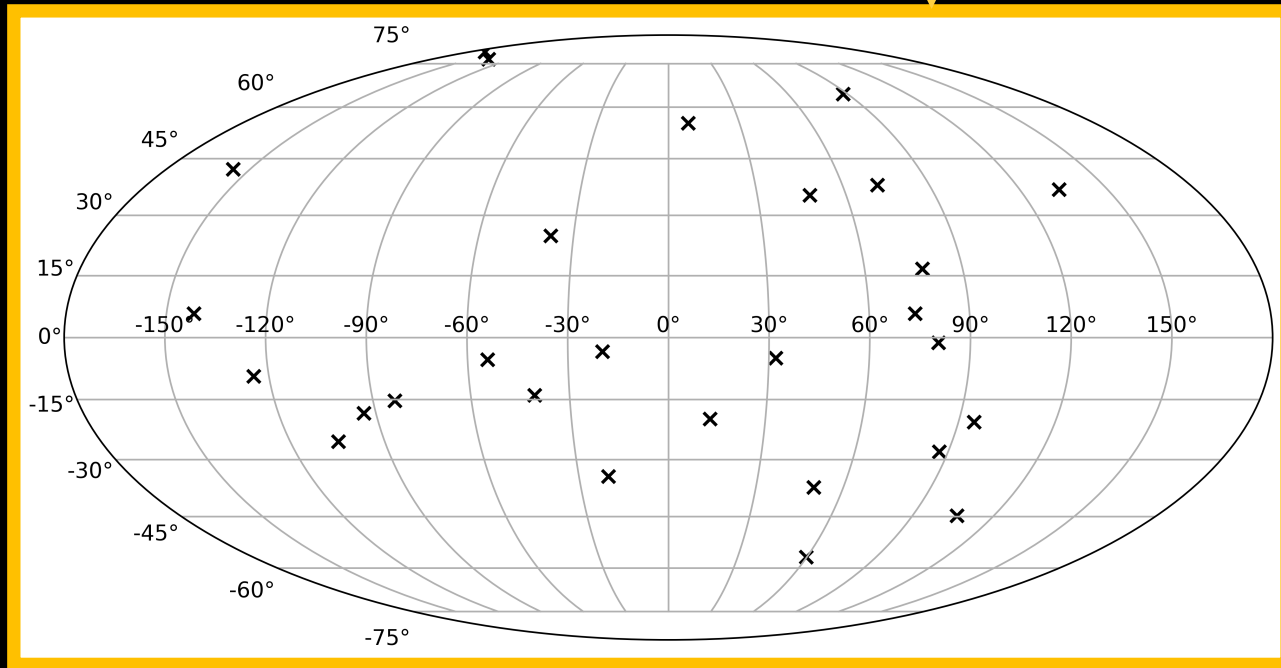
$$\text{P-value} = \frac{\text{\# of randomized « maps » with at least as many events as our data inside a circle}}{\text{total \# of randomized maps}}$$

One needs to compare the distribution of the signal to a large number of randomized maps



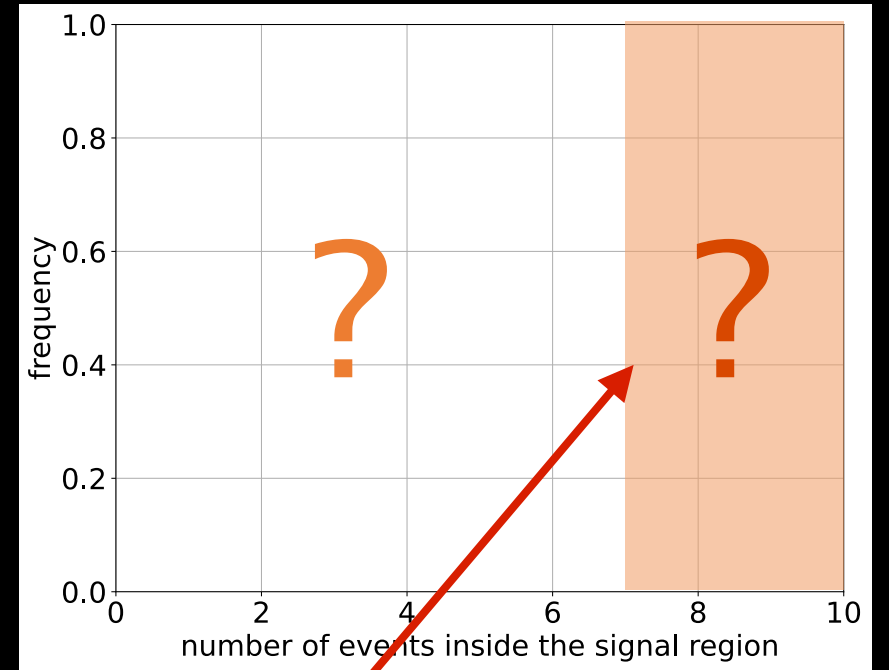
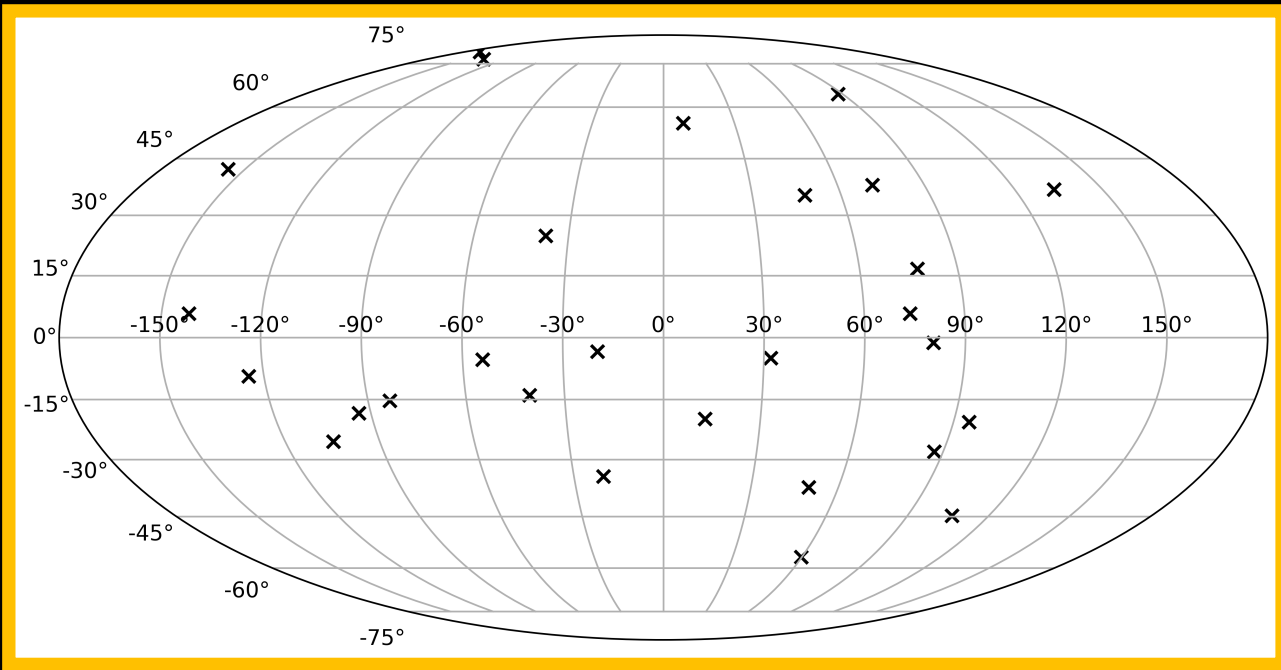
ACTIVITY!

We will make multiple random maps like this...



...and count how often we get 1, 2, 3, etc. events inside our circle.

ACTIVITY!



Eventually, we want to find how often we get **7 events or more** to conclude whether our map has a significant association with a neutrino source!

ACTIVITY!



Please follow this link to run the exercise in Python*:

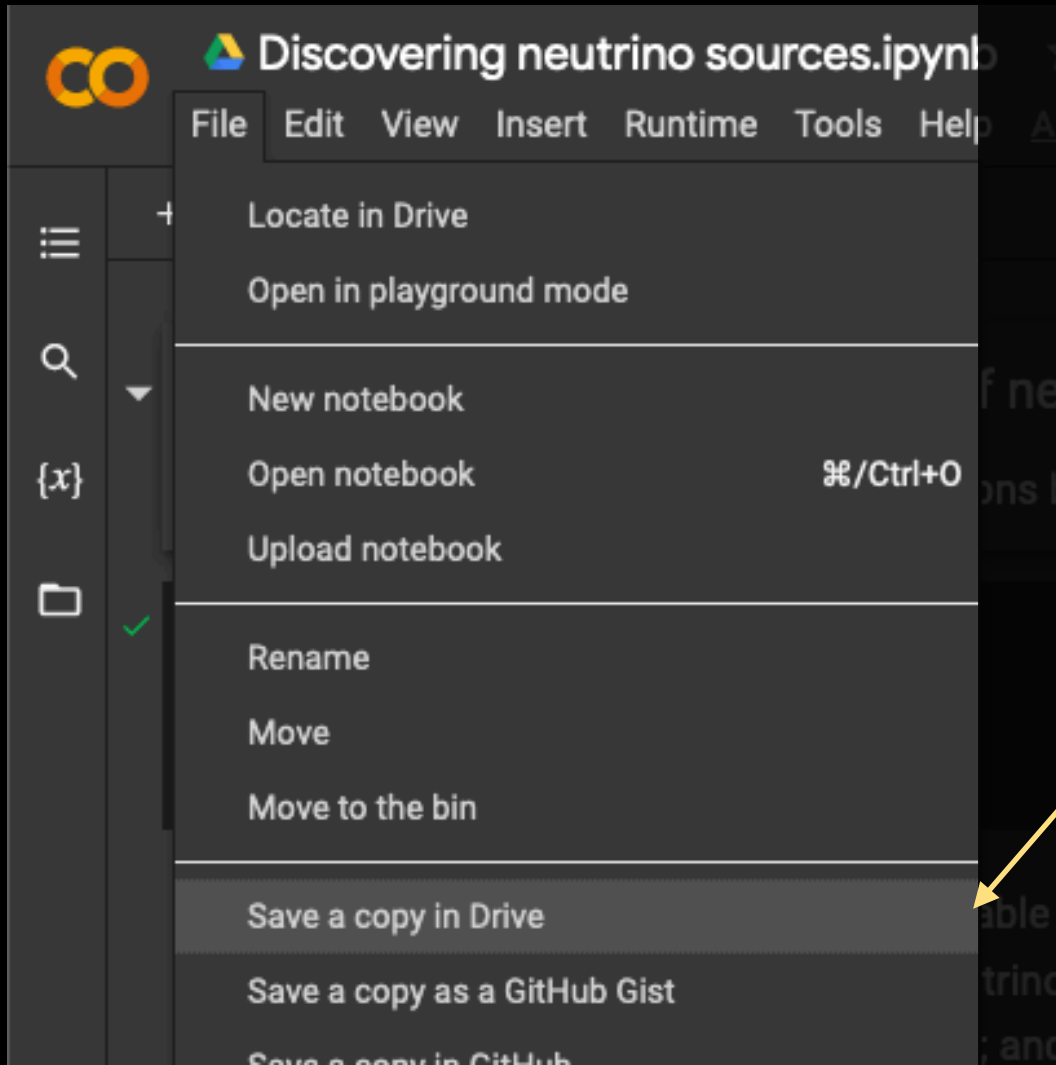
<http://bit.ly/IceCubeSources>

*PART 1: What does it mean to be random?

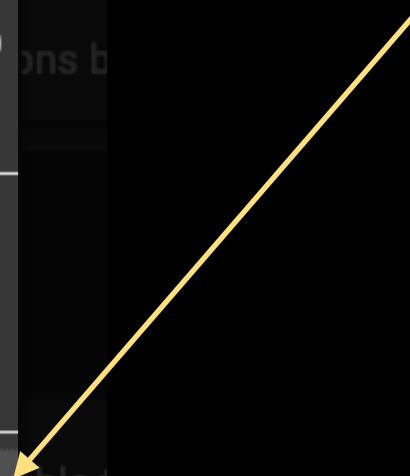
Question: What significance do you find?

IMPORTANT REMINDER

Please create your own copy of the notebook to make any edits:



click here!



WHAT CONSTITUTES A DISCOVERY?

How many events must come from the same region of the sky "discover" a source?

WHAT CONSTITUTES A DISCOVERY?

How many events must come from the same region of the sky "discover" a source?

Answer: The number of events in the signal region must have **less than 0.000000006%** (or 6×10^{-7}) chance of coming from randomized background

That is, only **one out of 1.7 million** randomized maps has more (or the same number of) events in the red circle than your data

THE POISSONIAN STATISTICS

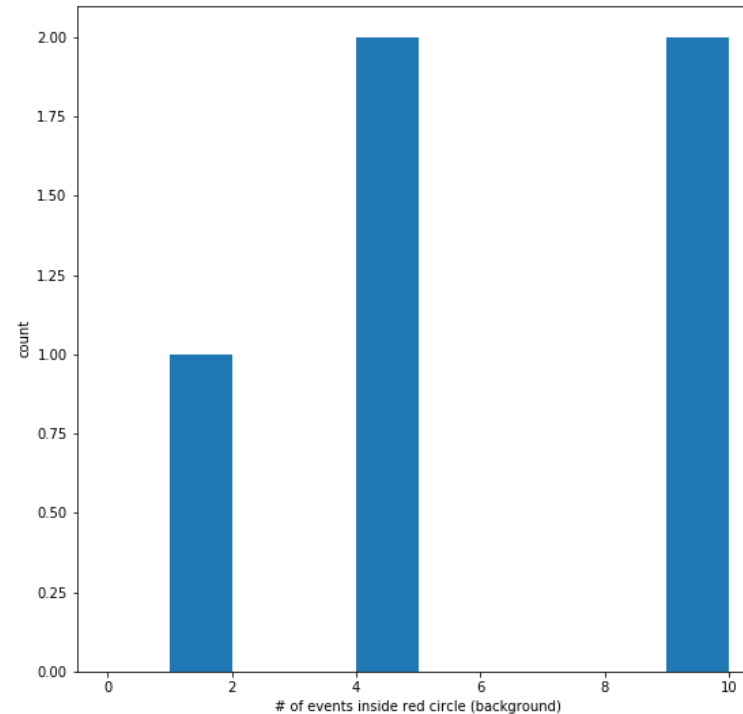
Say that we expect **on average 5** background events to occur within our patch of sky.

How many times are we going to see X events?

Here: 5 scrambles

```
[0]: fig,ax = plt.subplots(figsize=(9,9))  
  
bins = np.linspace(0,10,11)  
  
ax.hist(np.random.poisson(lam=5,size=5),bins=bins)  
ax.set_xlabel('# of events inside red circle (background)')  
ax.set_ylabel('count')
```

```
[0]: Text(0, 0.5, 'count')
```



THE POISSONIAN STATISTICS

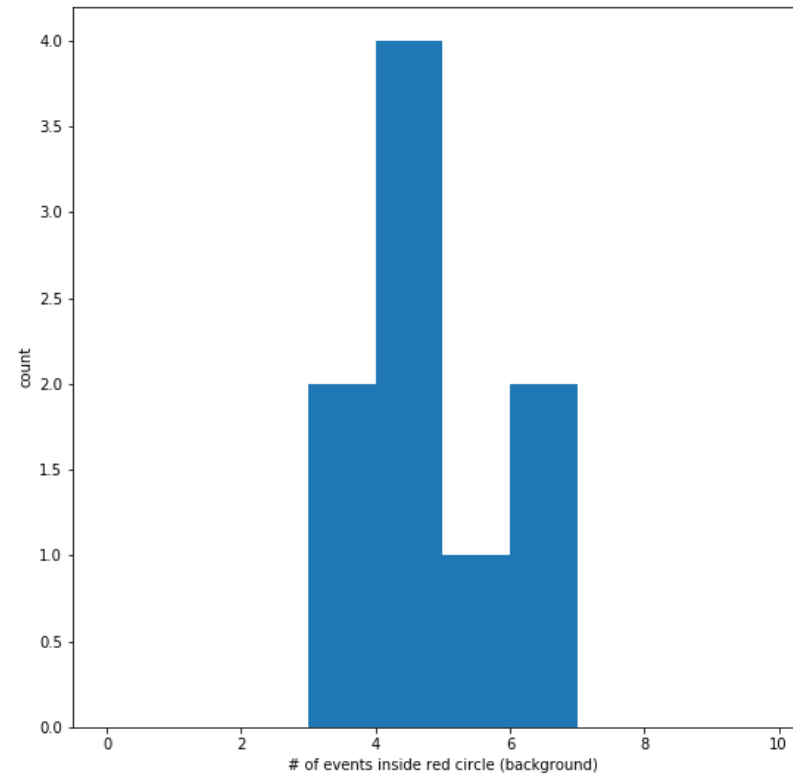
Say that we expect **on average 5** background events to occur within our patch of sky.

How many times are we going to see X events?

Here: 10 scrambles

```
[9]: fig,ax = plt.subplots(figsize=(9,9))  
  
bins = np.linspace(0,10,11)  
  
ax.hist(np.random.poisson(lam=5,size=10),bins=bins)  
ax.set_xlabel('# of events inside red circle (background)')  
ax.set_ylabel('count')
```

```
[9]: Text(0, 0.5, 'count')
```



THE POISSONIAN STATISTICS

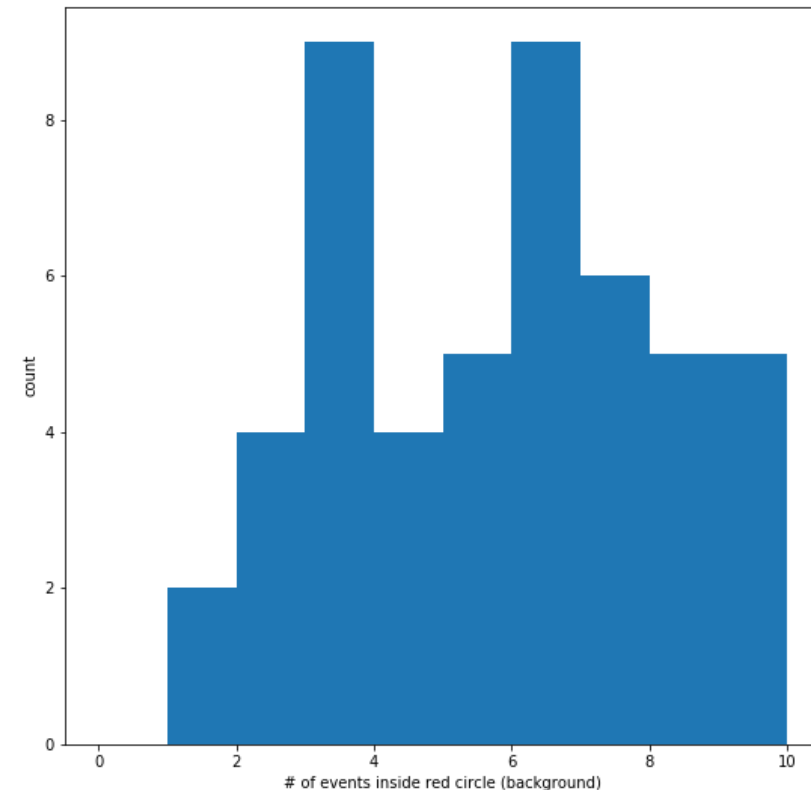
Say that we expect **on average 5** background events to occur within our patch of sky.

How many times are we going to see X events?

Here: 50 scrambles

```
[10]: fig,ax = plt.subplots(figsize=(9,9))  
  
bins = np.linspace(0,10,11)  
  
ax.hist(np.random.poisson(lam=5,size=50),bins=bins)  
ax.set_xlabel('# of events inside red circle (background)')  
ax.set_ylabel('count')
```

```
[10]: Text(0, 0.5, 'count')
```



THE POISSONIAN STATISTICS

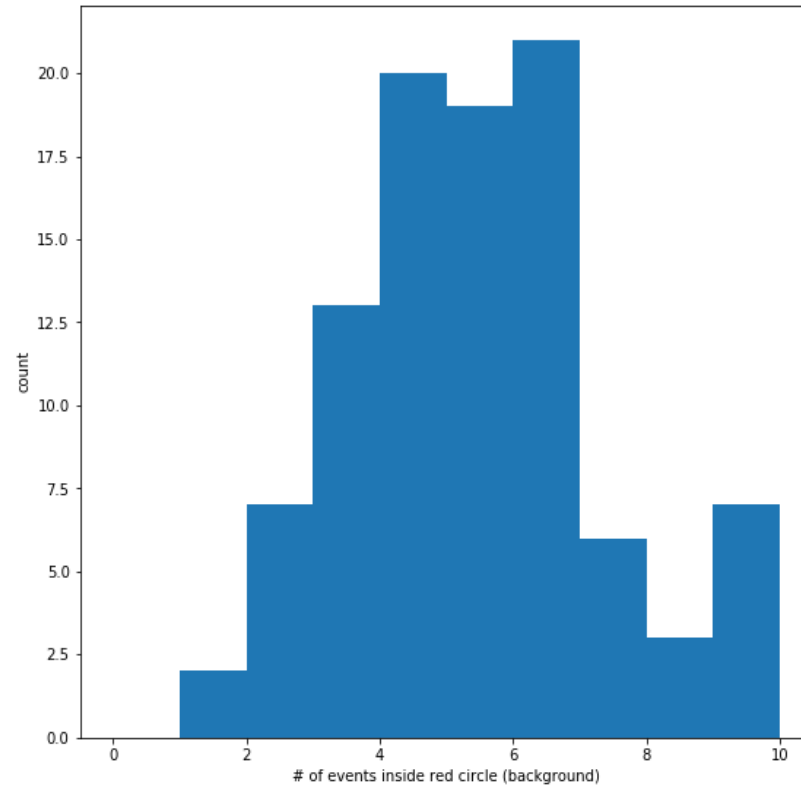
Say that we expect **on average 5** background events to occur within our patch of sky.

How many times are we going to see X events?

Here: 100 scrambles

```
[13]: fig,ax = plt.subplots(figsize=(9,9))  
  
bins = np.linspace(0,10,11)  
  
ax.hist(np.random.poisson(lam=5,size=100),bins=bins)  
ax.set_xlabel('# of events inside red circle (background)')  
ax.set_ylabel('count')
```

```
[13]: Text(0, 0.5, 'count')
```



THE POISSONIAN STATISTICS

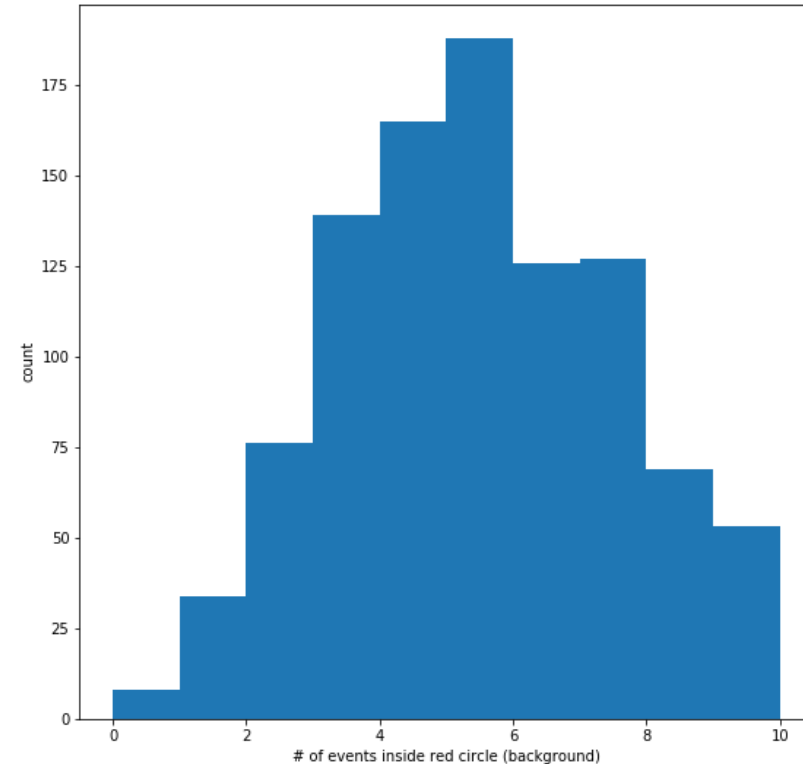
Say that we expect **on average 5** background events to occur within our patch of sky.

How many times are we going to see X events?

Here: 1000 scrambles

```
[14]: fig,ax = plt.subplots(figsize=(9,9))  
  
bins = np.linspace(0,10,11)  
  
ax.hist(np.random.poisson(lam=5,size=1000),bins=bins)  
ax.set_xlabel('# of events inside red circle (background)')  
ax.set_ylabel('count')
```

```
[14]: Text(0, 0.5, 'count')
```



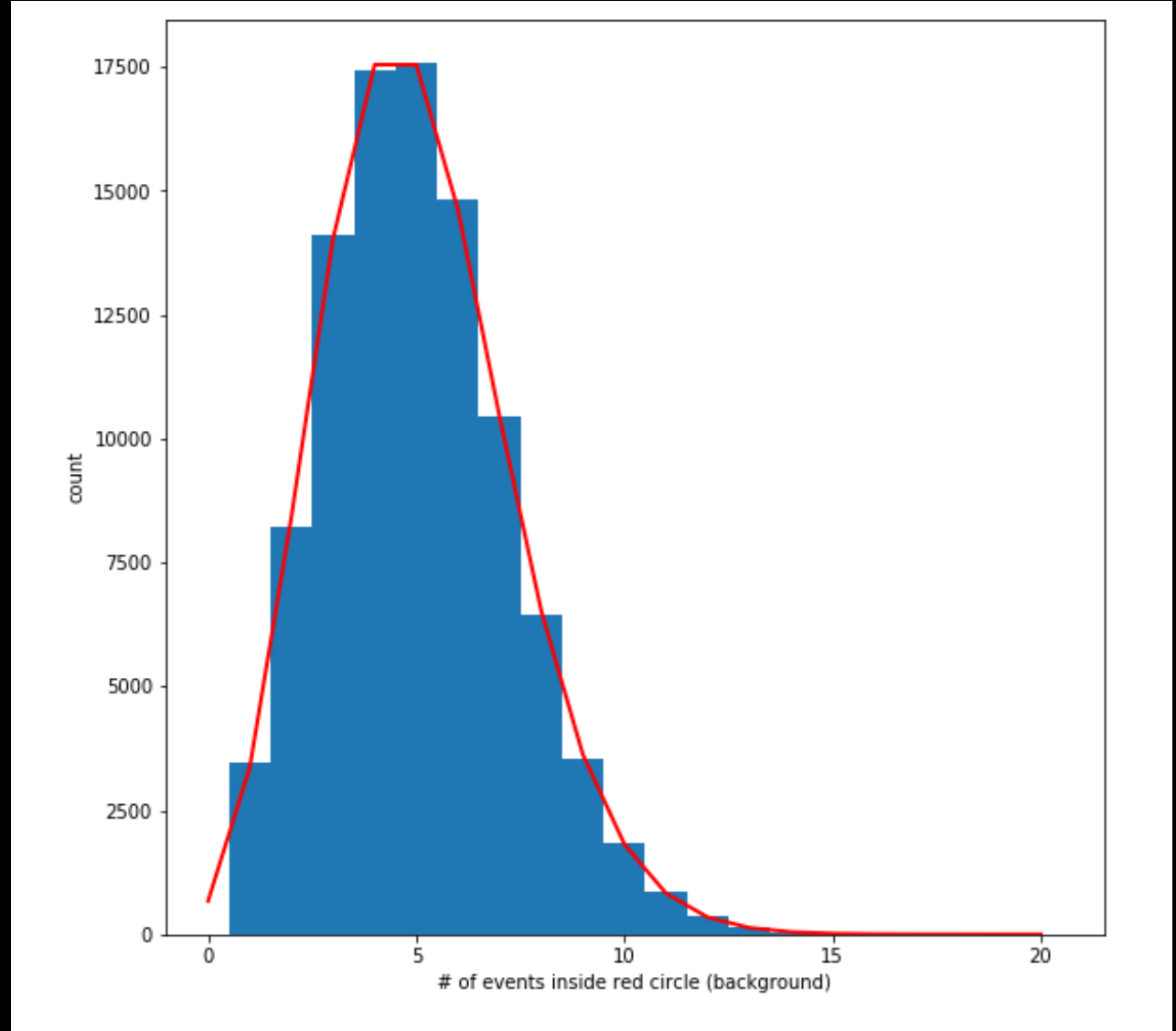
THE POISSONIAN STATISTICS

After a looooot of trials, a clear shape appears... and that shape follows the equation:

$$P(n) = \frac{\lambda^n}{n!} e^{-\lambda}$$

Probability of getting n background events in your red circle

Expected number of events



QUANTIFYING SIGNIFICANCE

- You can save yourself a lot of computation time by guessing how your background behaves. This behaviour is characterized by a **probability distribution function (PDF)**
- For a purely random background, the number of events in a region is determined by a **Poisson distribution**:

$$P(n) = \frac{\lambda^n}{n!} e^{-\lambda}$$

Expected number of events

Probability of getting n
background events in your
red circle

ACTIVITY!

- What is the expected number of background events for a signal region of $\sim 0.0685\pi$ rad² (the area of your circle)?

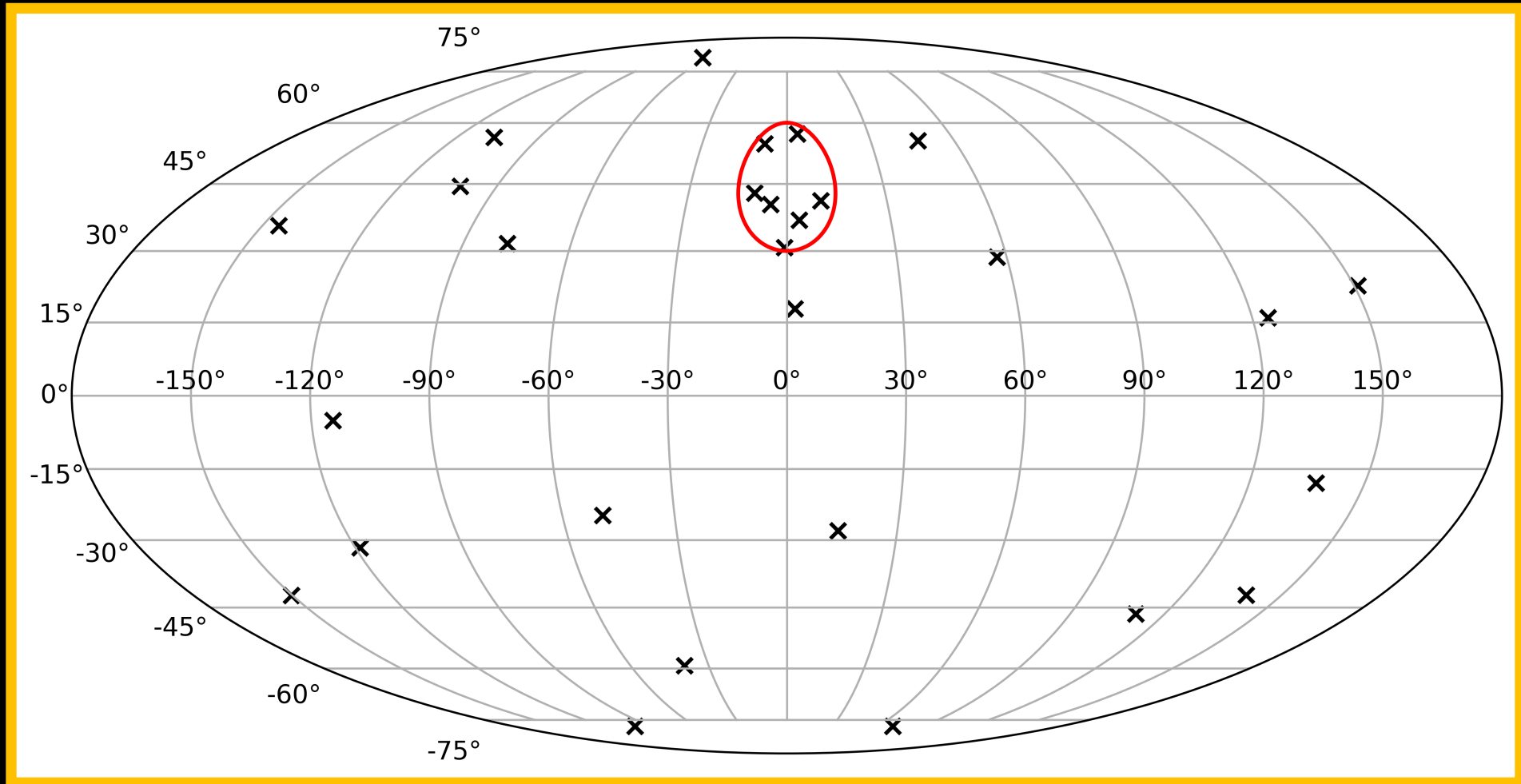
ACTIVITY!

- What is the expected number of background events for a signal region of $\sim 0.0685\pi$ rad² (the area of your circle)?

$$\text{Answer} = 28 \text{ events} \cdot 0.0685\pi / 4\pi \approx 0.48$$

(It's OK to expect non-integer values!)

WOULD THIS BE A SIGNIFICANT EVENT?



WOULD THIS BE A SIGNIFICANT EVENT?



Going back to our Python exercise*:

<http://bit.ly/IceCubeSources>

*PART 2: Poisson statistics: Quantifying how lucky we need to be

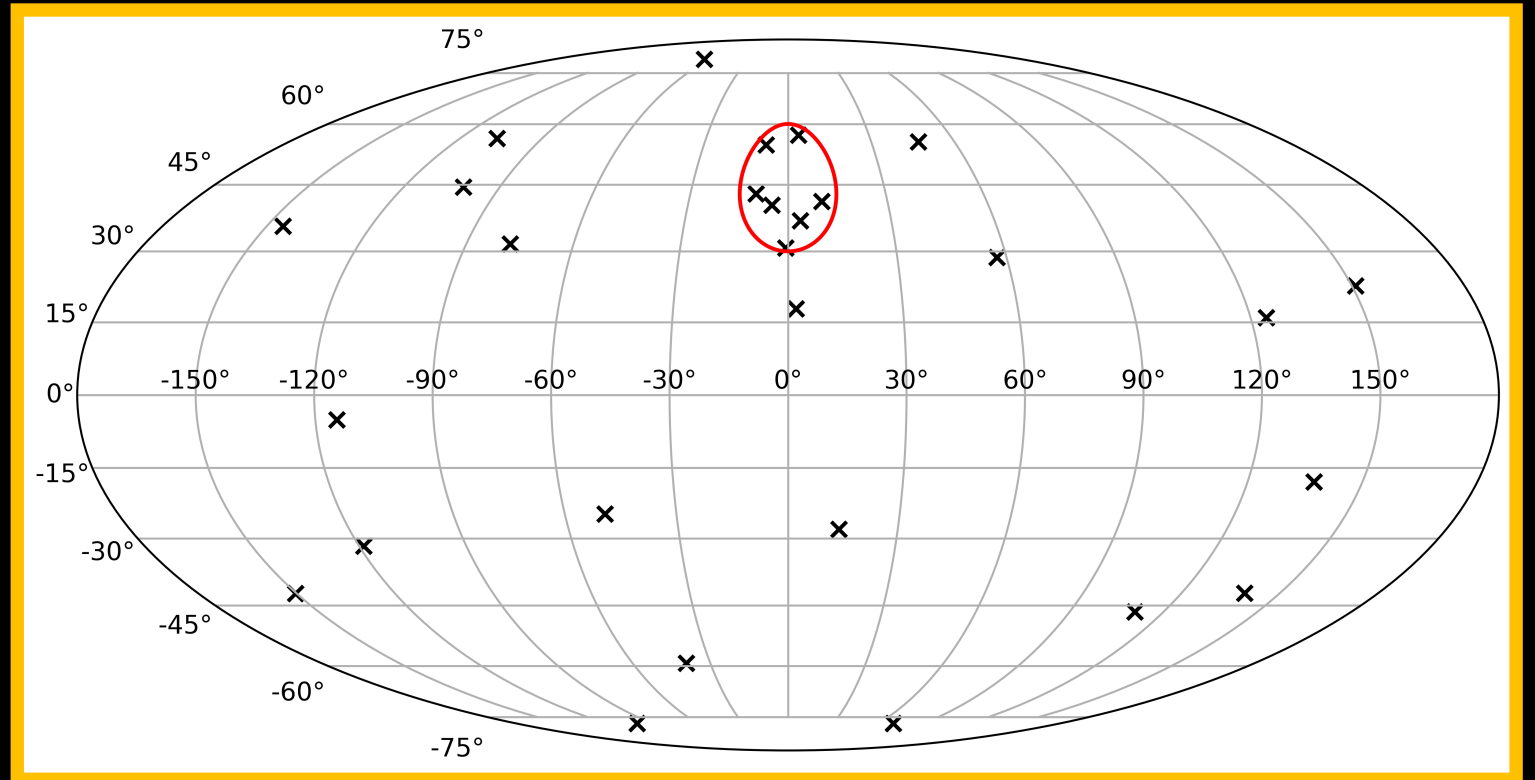
WOULD THIS BE A SIGNIFICANT EVENT?

Not quite!

Probability of getting 7
events or more by pure
luck:

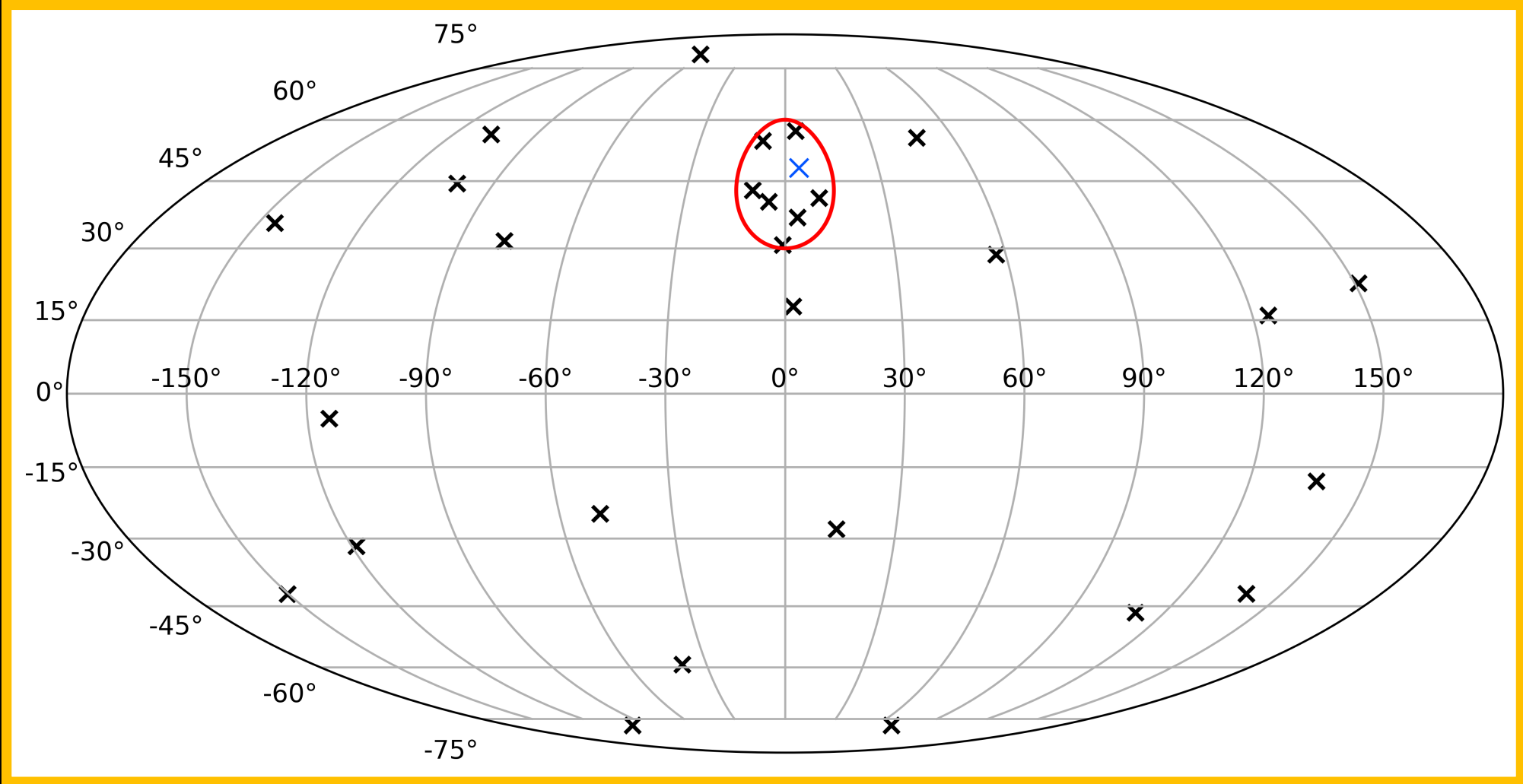
$$P \sim 7.64 \times 10^{-7}$$

Small, but still greater
than 6×10^{-7}



...We still cannot rule out a background at our desired confidence :(

WHAT IF ANOTHER EVENT ENDS UP IN THE CIRCLE?



YOU KNOW WHAT TO DO!



<http://bit.ly/IceCubeSources>

WHAT IF ANOTHER EVENT ENDS UP IN THE CIRCLE?

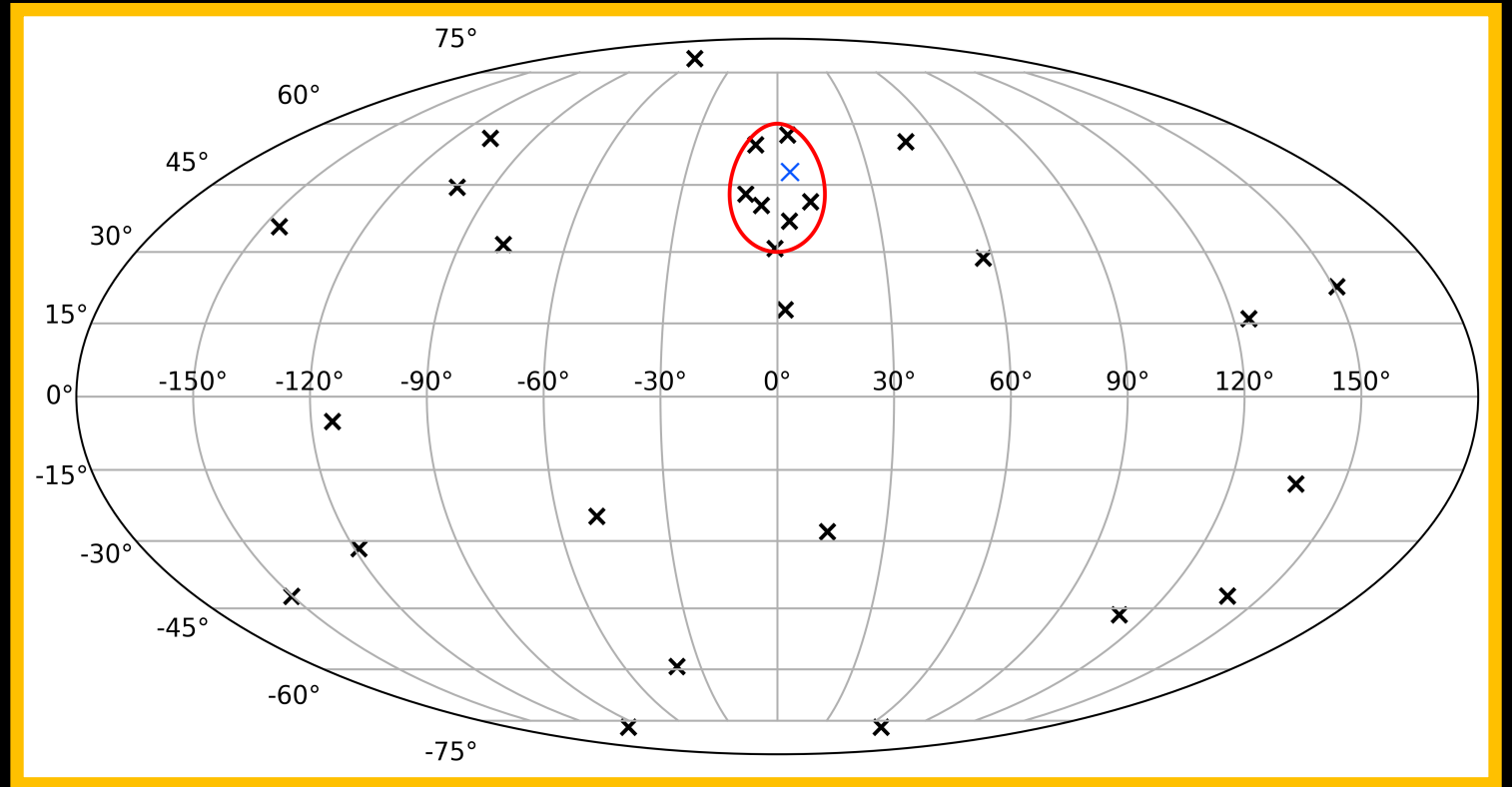
First, we need to update our background expectation:

We now expect $29 \times 0.0685 / 4 \approx 0.496$ events per circle

Plug in numbers...

$P \sim 5.9 \times 10^{-8}$

Now we have a discovery!

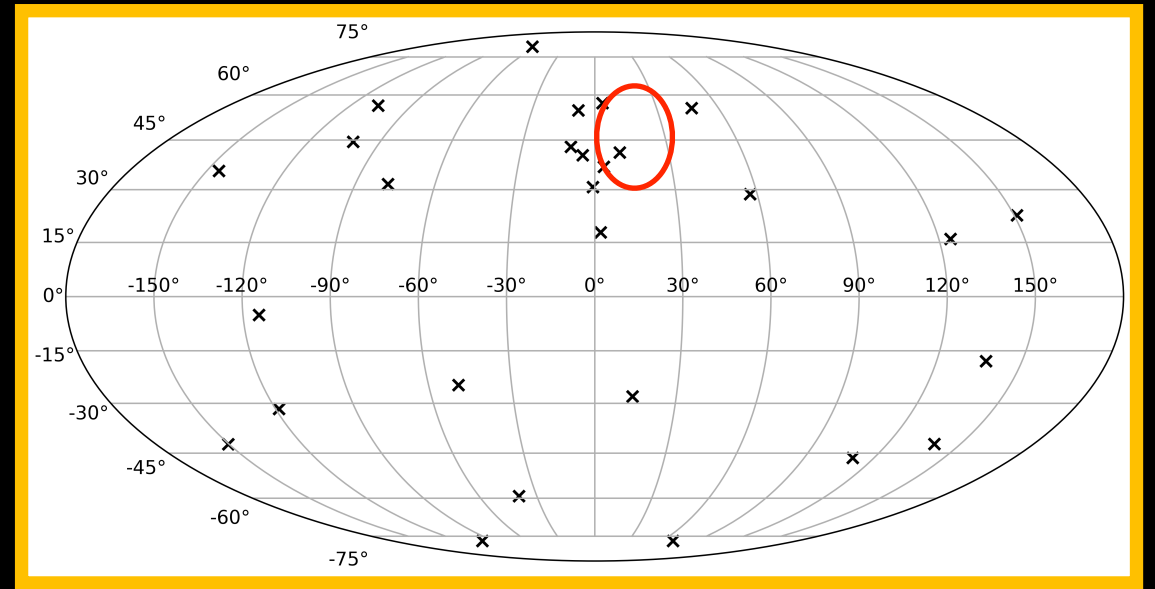
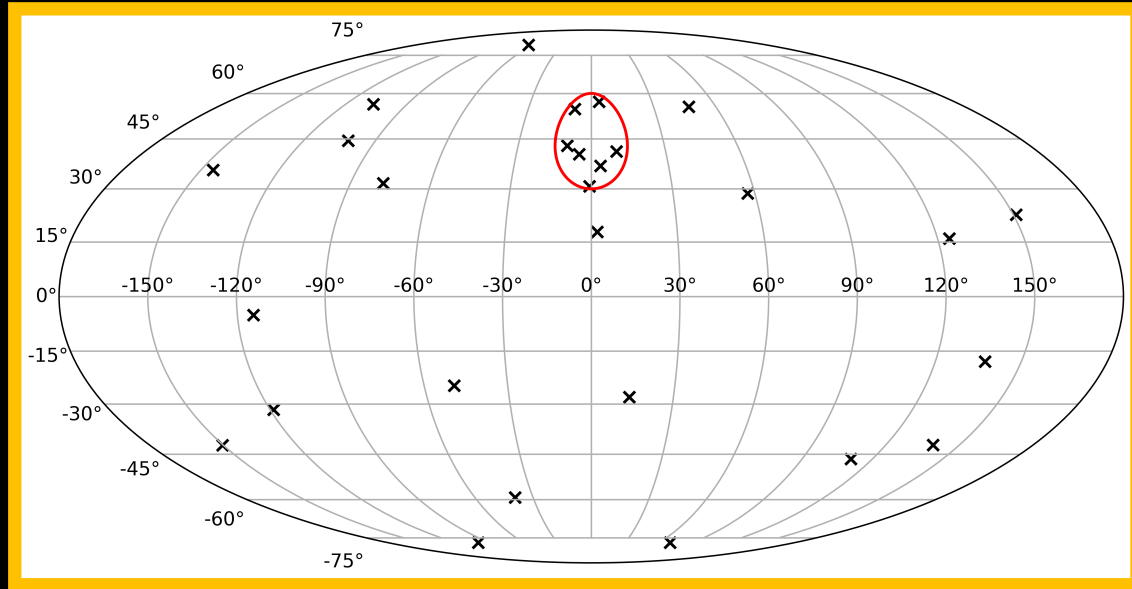


Every single neutrino matters.

MAKING THINGS A LITTLE MORE REALISTIC...

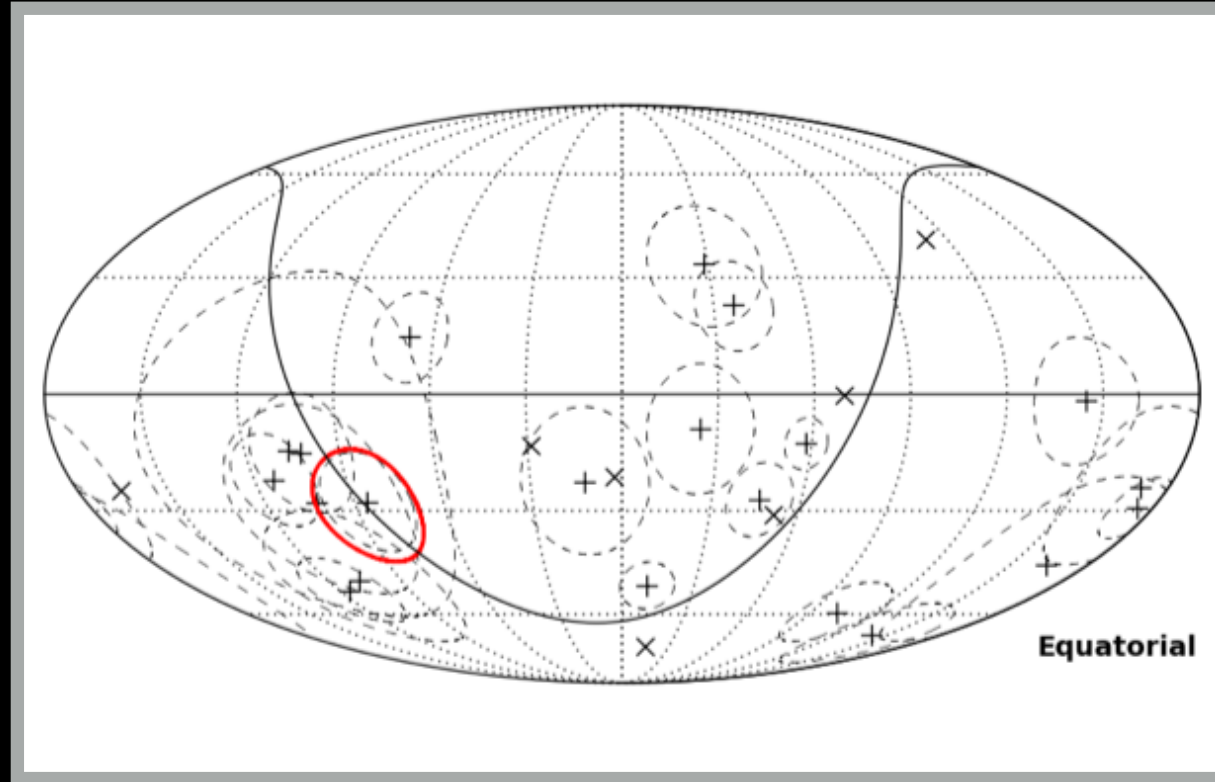
- What are the limitations of this approach?
- Do you think we could use more information to get our probability?
- Can you think of cases where binning may cause a problem?

MAKING THINGS A LITTLE MORE REALISTIC...



Would I still be able to detect a source using the bin on the right?

MAKING THINGS A LITTLE MORE REALISTIC...

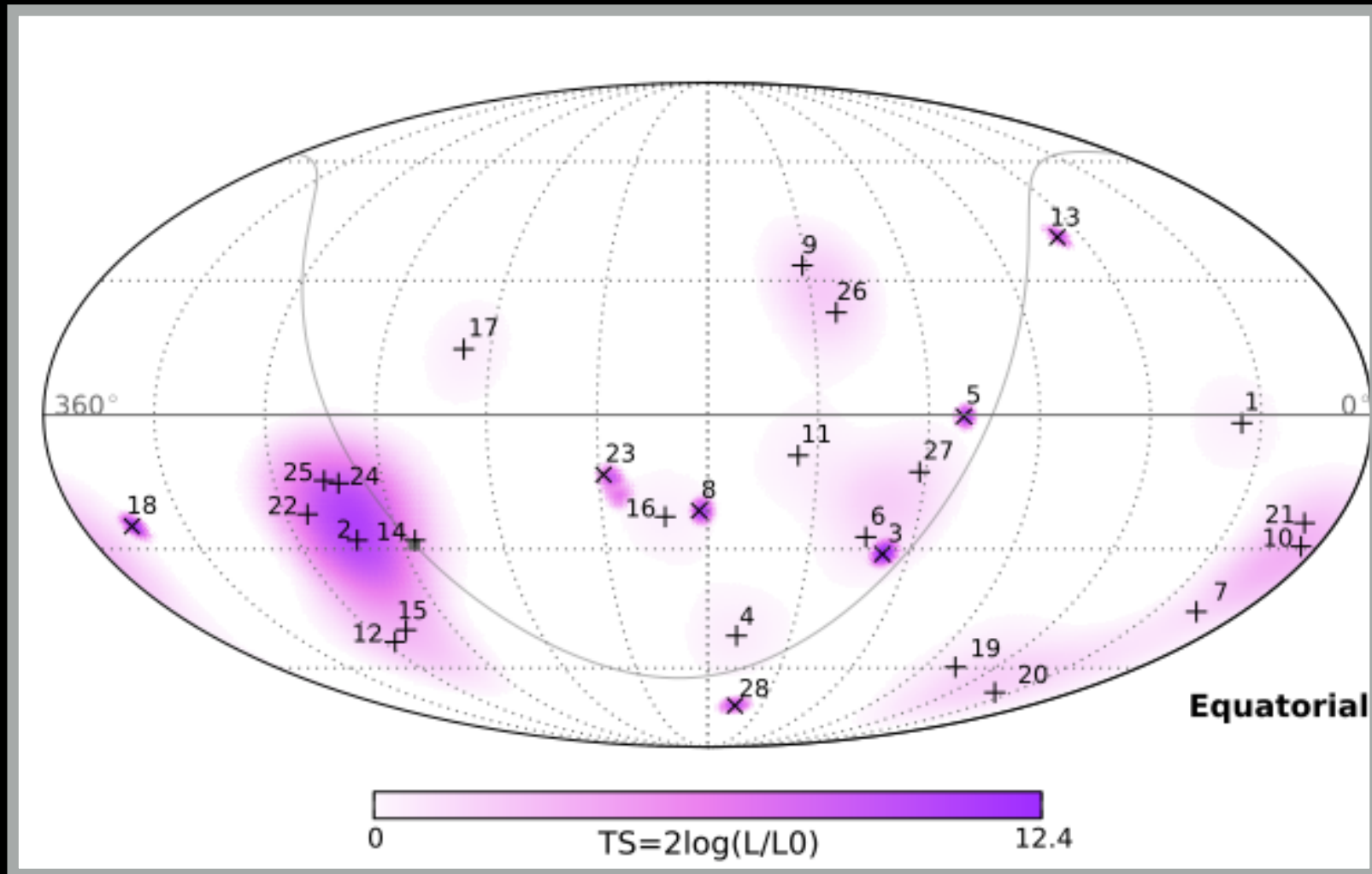


Should all events have the same contribution?

THE REAL DEAL: THE LIKELIHOOD METHOD

p-value of
the cluster:
0.08

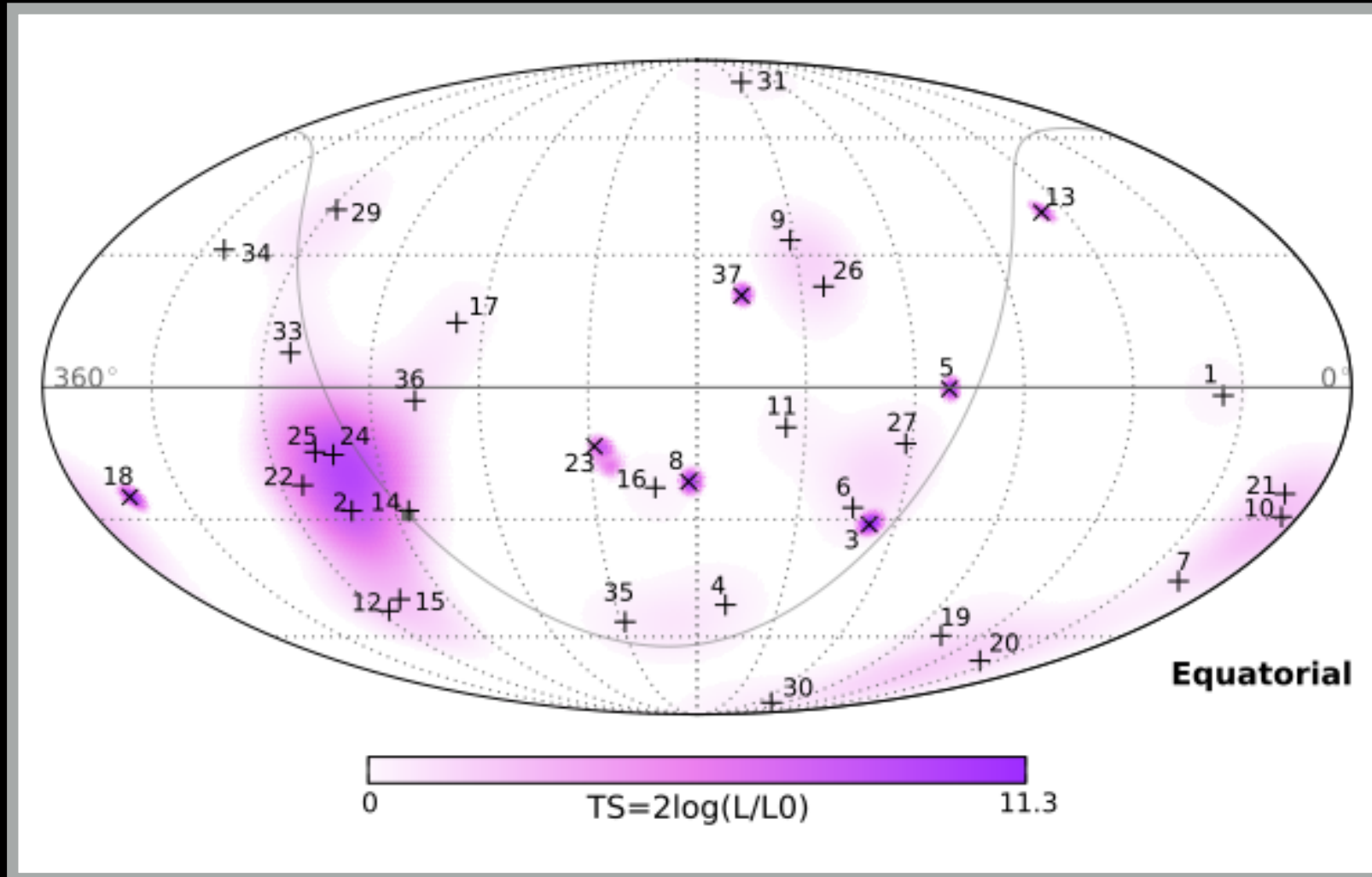
[So, only
expect the
same cluster
or weirder 8%
of the time!]



FULL ANALYSIS WITH 3 YEARS OF DATA

p-value of
the cluster:
0.84

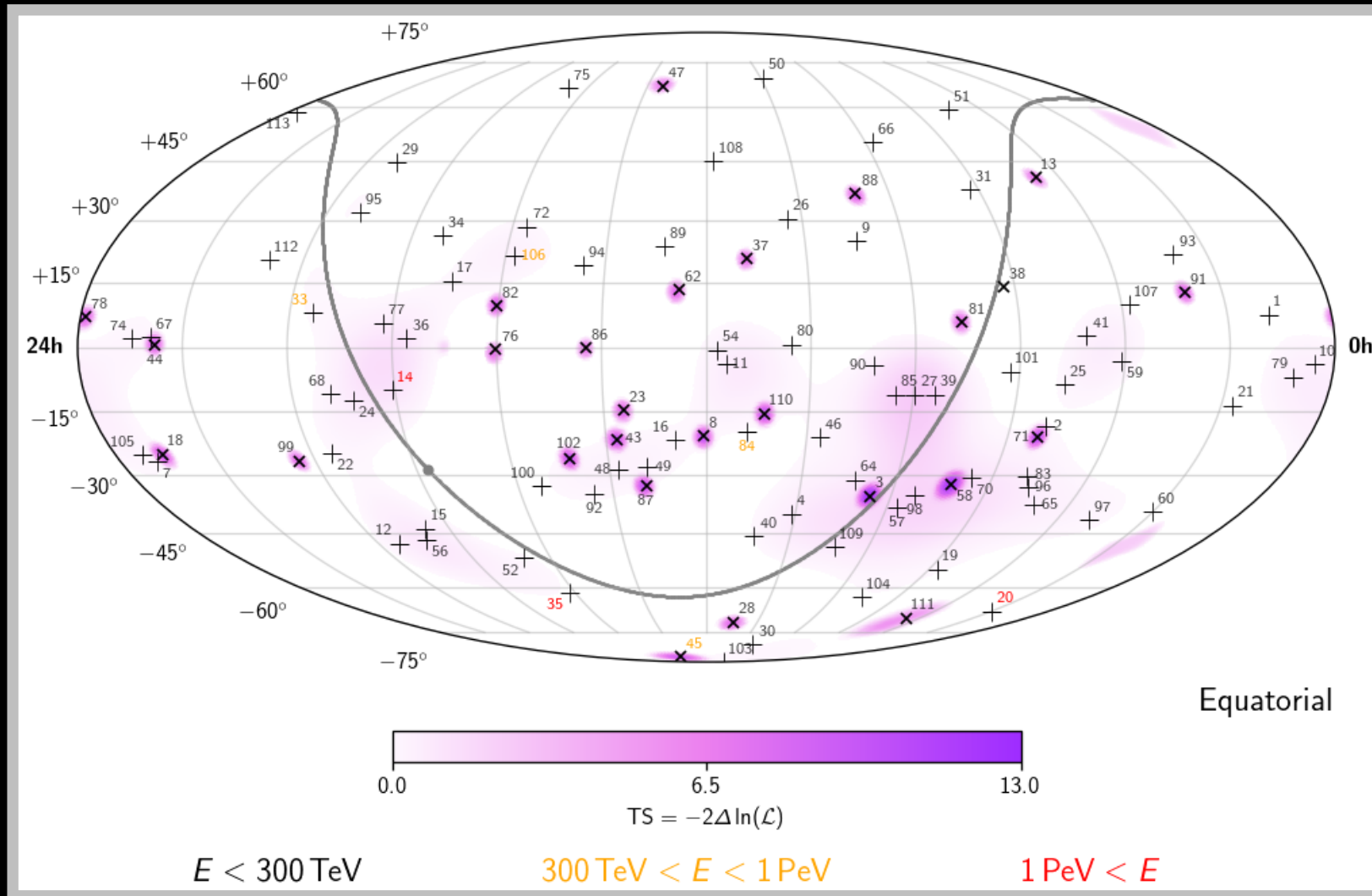
[after 3 years,
we see that
it's possible to
get the same
cluster of
weirders 84% of
the time...]

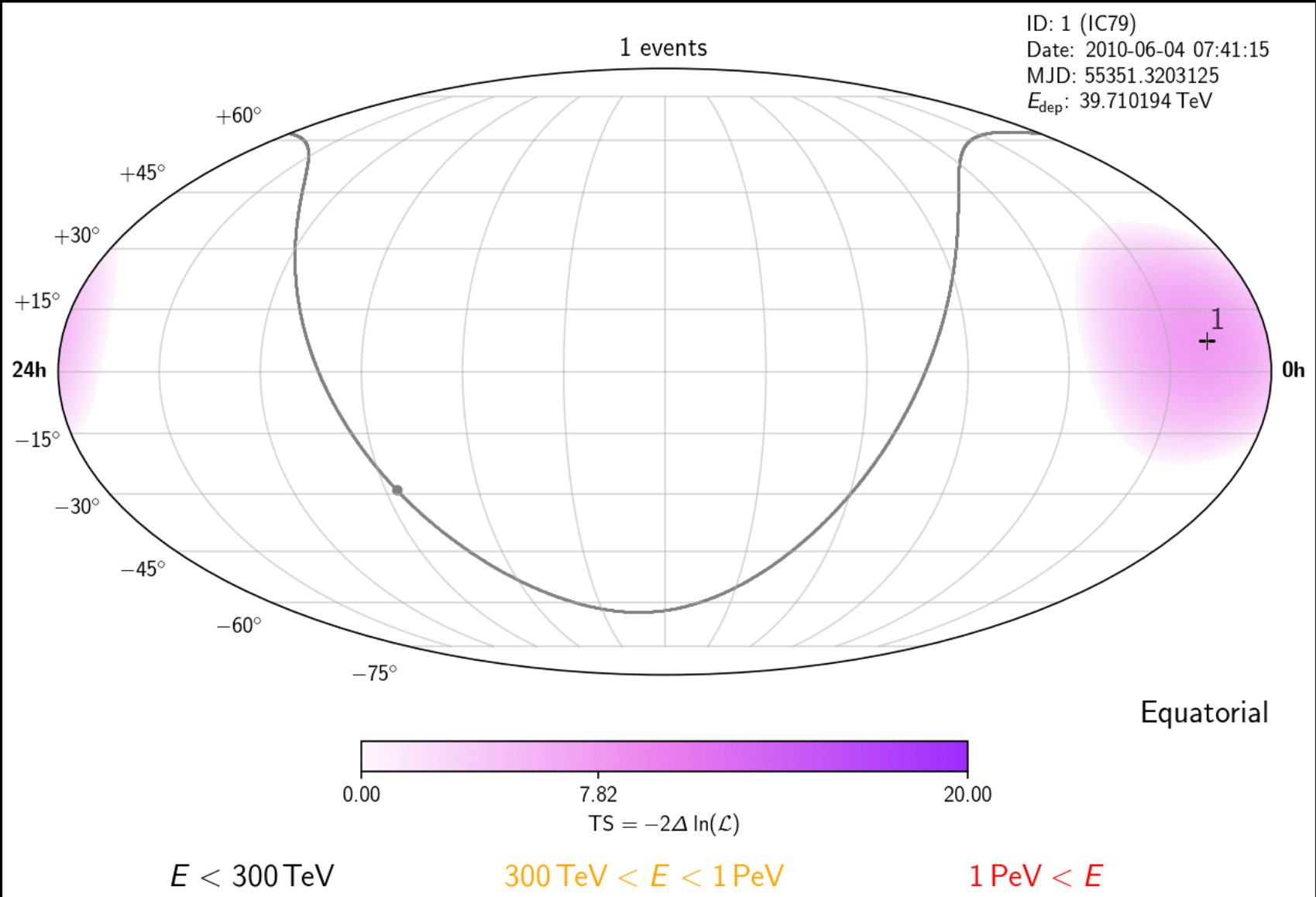


FULL ANALYSIS WITH 7 YEARS OF DATA

p-value of
the cluster:
0.81

[after 7 years,
we see that
it's possible to
get the same
cluster of
weirder 81% of
the time...]





LESSONS LEARNED?

- > We can't just find a cluster of neutrinos and claim that it is a significant source! It can also be "isotropic background", and we need to know how likely.
- > We can either generate random maps many times or use analytical background expectation (such as Poissonian probability) to quantify significance.
- > Often, if we know the analytical form of the background expectation, it can save us a lot of computational time!
- > Finally, it is better to know where we're looking instead of searching across the entire sky. Multimessenger observations can help guide our searches!

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- > Often, if we know the analytical form of the background expectation, it can save us a lot of computational time!
- > Finally, it is better to know where we're looking instead of searching across the entire sky. **Multimessenger observations can help guide our neutrino searches!**