

Based on:

OD, Prahar Mitra, Jorge Santos, 2207.07134 & 2210.xx

See also: Bhattacharyya, Minwalla, Papadodimas (2011) Markeviciute, Santos (2016, 2018)

NBI mini workshop: What is new in gravity? Helsingoer, August 2022

## $\rightarrow$ AdS<sub>5</sub> / CFT<sub>4</sub> duality

Type IIB supergravity theory on  $AdS_5 \times S^5$  with radius L and N units of  $F_{(5)}$  flux on  $S^5$ 

Large N and strong t'Hooft coupling  $\lambda = g_{YM}^2 N$  limit of  $\mathcal{N} = 4$  Super Yang-Mills (SYM) theory with gauge group SU(N) & YM coupling  $g_{YM}$ 



Thermal states of  $\mathcal{N}=4$  SYM

with temperature T, chemical potentials  $\mu_j$  and energies  $O(N^2)$ living on the Einstein static Universe  $\mathbb{R}_t \times S^3$ 

> Asymptotically global AdS<sub>5</sub>xS<sup>5</sup> BHs of IIB supergravity with Hawking temperature T and chemical potentials μ<sub>j</sub>

#### → Motivations

- We should find all the BHs and map them into thermal states in the dual SYM => identify the dominant phases (as saddle points) in the thermodynamic ensembles.
- Necessary to reproduce microscopically the Bekenstein-Hawking entropy of AdS BHs. [See Benini's review talk at Strings 2022]
- Contribute to understand a puzzle of SO(6) gauged supergravity: its most general SUSY BH known so far — Kunduri-Lucietti-Reall BH — has <u>only</u> 4 independent parameters.

However, asymptotically  $AdS_5 \times S^5$  BHs are characterized by 6 conserved charges with the BPS relation constraint E=  $Q_1 + Q_2 + Q_3 + J_1 + J_2$ 

=> the most general SUSY BH should be a 5-parameter solution. From dual CFT perspective, most general SUSY states also expected to be characterized by 5 parameters.

So, what is the missing gravitational parameter?

[Gutowski, Reall '04] [Kunduri, Lucietti, Reall '06]

→ AIM of this Talk: identify new thermal phases with a finite chemical potential that can dominate some thermodynamic ensembles.

#### → Known SYM phases

#### The massless bosonic fields of type IIB supergravity: metric tensor g<sub>ab</sub>, dilaton Φ, axion C, NS-NS antisymmetric 2-tensor B<sub>(2)</sub>, RR 2-form potential C<sub>(2)</sub>, and RR 4-form C<sub>(4)</sub> with a 5-form field strength F<sub>(5)</sub> =dC<sub>(4)</sub> satisfying a self-duality condition. Fermionic superpartners: complex Weyl gravitino & complex Weyl dilatino.

- Known solutions of IIB supergravity with only { g<sub>ab</sub>, F<sub>(5)</sub> }:
  - Global AdS<sub>5</sub>×S<sup>5</sup> Schwarzschild & its rotating black hole (BH) partners.
     Everywhere (not only at bdry) the direct product of two base spaces
     have horizon topology S<sup>3</sup>×S<sup>5</sup>.



- Asymptotically globally AdS5 x S<sup>5</sup> BHs that break the SO(6) symmet down to SO(5):
  - lumpy BHs with polar deformations along the  $S^5$  and
  - localized BHs on the  $5^5$  (with  $5^8$  horizon topology).

[ OD, Santos, Way 1501.06574 & 1605.0491

#### → Towards finding more SYM phases

- Explore even further the phase space of thermal states (to identify all the relevant saddle points for the thermodynamic partition functions of the theory).
- Useful: dimensional reduction of IIB along S<sup>5</sup> yields 5d N=8 gauged supergravity.
   It's believed (not proven) to be a <u>consistent</u> reduction of IIB on AdS<sub>5</sub>xS<sup>5</sup>.
   Gunaydin, Romans, Warner (1986)
- 10d fields  $g_{ab}$ ,  $\Phi$ , C,  $B_{(2)}$ ,  $C_{(2)}$ ,  $C_{(4)}$  are equivalently encoded in the 5d field content of gauged N =8 SUGRA: graviton  $g_{ab}$ , fifteen SO(6) gauge fields  $A^{ij}$ , twelve 2-form gauge potentials in the  $6 + \overline{6}$  representations of SO(6), 42 scalars in the  $1 + 1 + 20' + 10 + \overline{10}$  representations of SO(6) & the fermionic superpartners.
- But IIB with only  $g_{ab}$ ,  $F_{(5)}$  (relevant for AdS/CFT: source D3's) can be <u>consistently</u> dim reduced along the S<sup>5</sup> to yield 5d SO(6) gauged supergravity.

Cvetic-Lü-Pope-Sadrzadeh-Tran [hep-th/0003103]

This is itself is a <u>consistent</u> truncation of gauged N =8 SUGRA where we set the  $1+1+10+\overline{10}$  scalars and the  $6+\overline{6}$  2-form potentials to zero. The bosons that survive descend from  $\{g_{ab}, F_{(5)}\}$  of IIB:

the graviton  $g_{ab}$ , the 15 gauge fields  $A^{ij}$  & the 20' scalars.

 $\rightarrow$  U(1)<sup>3</sup> gauged supergravity:

[OD, Mitra, Santos '22]

<u>Consistent</u> truncation of SO(6) gauged SUGRA down to the Cartan subgroup of SO(6) with associated gauge fields  $\{A_{(1)}^{k}\}$  (K=1,2,3): U(1)<sup>3</sup> gauged supergravity.

The field content:

- Graviton  $g_{ab}$  + 2 neutral scalars { $\phi_1$ ,  $\phi_2$ } + 3 U(1)'s gauge fields { $A_{(1)}^{K}$ },
- + 3 complex scalar fields  $\{\Phi_k\}$  minimally coupled to  $\{A_{(1)}^k\}$  with charge qL =2.

All 5 scalars have mass m<sup>2</sup>L<sup>2</sup> = -2 => saturate AdS<sub>5</sub> Breitenlöhner-Freedman (BF) bound.

$$S = \frac{1}{16\pi G_5} \int dx^5 \sqrt{-g} \left\{ R - V(\varphi_{\tilde{K}}, \lambda_K) - \frac{1}{2} \sum_{\tilde{K}=1}^2 (\nabla \varphi_{\tilde{K}})^2 - \frac{1}{4} \sum_{K=1}^3 \frac{\left(F_{(2)}^K\right)^2}{X_K^2(\varphi_{\tilde{K}})} - \frac{1}{8} \sum_{K=1}^3 \left[ (D_a \Phi_K) (D^a \Phi_K)^\dagger - \frac{(\nabla \lambda_K)^2}{4(4 + \lambda_K)} \right] \right\} - \frac{1}{16\pi G_5} \int F_{(2)}^1 \wedge F_{(2)}^2 \wedge A_{(1)}^3$$
$$D_a \Phi_K = \partial_a \Phi_K - i \frac{2}{L} A_a^K \Phi_K, \qquad F_{ab}^K \equiv \partial_a A_b^K - \partial_b A_a^K, \qquad \lambda_K \equiv \Phi_K \Phi_K^\dagger$$
$$V(\varphi_{\tilde{K}}, \Phi_K) = \frac{1}{2L^2} \left[ \sum_{K=1}^3 X_K^2(\varphi_{\tilde{K}}) \lambda_K - \sum_{I \neq J \neq K=1}^3 \frac{\sqrt{4 + \lambda_J} \sqrt{4 + \lambda_K}}{X_I(\varphi_{\tilde{K}})} \right]$$

• EoM of U(1)<sup>3</sup> gauged SUGRA are complicated.

- Three special cases of U(1)<sup>3</sup> gauged supergravity where the EoM simplify:
  - 1. Truncation with three equal charges =>  $A^1 = A^2 = A^3 \equiv A, \Phi_1 = \Phi_2 = \Phi_3 \equiv \Phi$

 $\varphi_1 = \varphi_2 = 0$ 

Liu,Lü,Pope,Vazquez-Poritz '07 Bhattacharyya, Minwalla, Papadodimas (2011) Markeviciute, Santos (2016, 2018)

2. Truncation with a single charge  $\Rightarrow A_{(1)}^1 = A_{(1)}^2 \equiv 0, A_{(1)}^3 \equiv A_{(1)} \text{ and } \Phi_1 = \Phi_2 \equiv 0, \Phi_3 \equiv \Phi$   $\varphi_1 \equiv \varphi, \varphi_2 = 0$ This talk [OD, Mitra, Santos '22]

3. Truncation with two equal charges =>  $A^1 = A^2 \equiv A, A^3 \equiv 0$  and  $\Phi_1 = \Phi_2 \equiv \Phi, \Phi_3 \equiv 0$   $\varphi_1 \equiv \varphi, \varphi_2 = 0$ This talk [OD, Mitra, Santos '22]  $\rightarrow$  'Bald' BHs (Cvetic-Lü-Pope 2004):  $\Phi_{\rm K}$  =0 (K=1,2,3)  $\Rightarrow$   $\langle \Phi_{\rm K} \rangle = 0$ 

- Parameterized by an energy E and three electric charges  $\{Q_1, Q_2, Q_3\} \leftarrow A_{(1)}^{K}$ Can be viewed as the "Reissner-Nordström-AdS<sub>5</sub>" (RNAdS) BHs of the theory. But, also have non-trivial neutral scalar fields  $\{\varphi_1, \varphi_2\}$  supporting them.
  - <u>Single</u> charge truncation



 <u>Two equal</u> charge truncation (equal 3-charge truncation is similar)



## → Truncation of U(1)<sup>3</sup> gauged supergravity with a single charge

$$A_{(1)}^1 = A_{(1)}^2 \equiv 0, A_{(1)}^3 \equiv A_{(1)} \text{ and } \Phi_1 = \Phi_2 \equiv 0, \Phi_3 \equiv \Phi$$
  
 $\varphi_1 \equiv \varphi, \, \varphi_2 = 0$ 

→ Truncation of U(1)<sup>3</sup> gauged supergravity with a single charge

 Scalar condensation instability of '<u>Bald</u>' CLP BHs (**Φ** = 0):

Perturb CLP BH with charged

scalar field  $\Phi_3 = \Phi$ :  $\Phi(t,r) = \widehat{\Phi}_{\omega}(r) e^{-i\omega t}$ 

- & solve the associated "Klein-Gordon" eqn with appropriate BCs.
- => Unstable when Im ( $\omega$ L) >0
- We can also solve directly for the
   <u>onset</u> of the instability, Q<sub>o</sub>(E), where ω = 0.
   CLP unstable above onset.
- There is also a supersymmetric Soliton with Q=E and ⟨Φ⟩ ≠ 0 (analytical soln)
   [Liu, Lü, Pope, Vazquez-Poritz '07]
   [OD, Mitra, Santos '22]



• CLP BHs are linearly unstable to condensation of charged scalar  $\Phi$ .

⇒ <u>beyond leading order</u> in perturbation theory, there should exist hairy BHs with charged hair bifurcating (in a 2nd order phase transition) from the onset of the CLP instability, in a phase diagram of static solutions.

- Solve EoM (5 coupled nonlinear ODEs) either **numerically** (at full nonlinear level) or within **perturbation theory** (matched asymp expansion) with boundary conditions:
  - Regularity at horizon (hairy BH) or at origin (supersymmetric soliton)
  - Asymp global AdS5 BCs for metric: bdry is Einstein Static Universe

$$\mathrm{d}s_{\partial\mathcal{M}}^2 = -\mathrm{d}t^2 + L^2\,\mathrm{d}\Omega_3^2$$

Leave the homogeneous chemical potential (source of U(1) gauge field) free :

$$\lim_{z \to 0} A^K = \mu_K \, \mathrm{d}t \qquad (z \text{ is FG coordinate with bdry at } z=0)$$

NO sources for the charged & neutral scalar fields (all saturate the BF bound).

$$\begin{array}{ll}
\varphi_{\tilde{K}} \xrightarrow{}_{z \to 0} V_{\tilde{K}} z^{2} \log z + \langle \mathcal{O}_{\varphi_{\tilde{K}}} \rangle z^{2} + \dots \\
\Phi_{K} \xrightarrow{}_{z \to 0} V_{K} z^{2} \log z + \langle \mathcal{O}_{\Phi_{K}} \rangle z^{2} + \dots \\
\end{array} \qquad \Rightarrow \qquad \begin{array}{ll}
\varphi_{\tilde{K}} \xrightarrow{}_{z \to 0} \langle \mathcal{O}_{\varphi_{\tilde{K}}} \rangle z^{2} + \dots \\
\Phi_{K} \xrightarrow{}_{z \to 0} \langle \mathcal{O}_{\Phi_{K}} \rangle z^{2} + \dots \\
\end{array}$$

• Thermodynamic quantities computed using Holographic Renormalization

→ Microcononical phase diagram (of truncation with a <u>single</u> charge)

Fix energy E and charge Q: dominant phase is the one with highest entropy S



- Hairy BHs always have higher S than CLP BH with same E,Q
- Hairy BH temperature approaches T L=1/ $\pi$  (& µ=1) in the singular (S=0) BPS lim !! This is also min T that CLP can reach (at A) in the singular BPS limit.

→ Microcononical phase diagram (of truncation with a <u>single</u> charge)

Limit E—>Q of Hairy BHs ? Singular limit (S=0) coincides with soliton at E=Q.

Charged and neutral scalar field VEVs for the line of Hairy BH solutions closest to the supersymmetric bound (blue disks) Solid red lines are analytical results for for the SUSY soliton.



 $\rightarrow$  Truncation of U(1)<sup>3</sup> gauged supergravity with <u>two equal</u> charges

$$A^1 = A^2 \equiv A, A^3 \equiv 0$$
 and  $\Phi_1 = \Phi_2 \equiv \Phi, \Phi_3 \equiv 0$ 

$$\varphi_1 \equiv \varphi, \, \varphi_2 = 0$$

Physics of equal three charge truncation is similar (General arbitrary charge case should be similar)

→ Truncation of U(1)<sup>3</sup> gauged supergravity with two equal charges

- Scalar condensation instability of 'Bald' CLP BHs ( $\Phi = 0$ ): <u>Perturb</u> CLP BH with charged scalar field  $\Phi_{1,2} = \Phi$ :  $\Phi(t,r) = \widehat{\Phi}_{\omega}(r) e^{-i\omega t}$ & solve the associated "Klein-Gordon" eqn in the CLP BH with appropriate BCs.
  - => Unstable when Im ( $\omega$ L) >0

- We can also solve directly for the <u>onset</u> of instability, Q<sub>o</sub>(E), where ω = 0.
   CLP unstable above onset.
- There is also a supersymmetric Soliton with Q=E/2 & ⟨Φ⟩ ≠ 0 (numerical soln)
  [Liu, Lü, Pope, Vazquez-Poritz '07]
  [OD, Mitra, Santos '22]



Microcononical phase diagram (of truncation with <u>two equal</u> charges)
Fix energy E and charge Q: dominant phase is the one with highest entropy S

Physics of equal three charge 2.0 truncation is similar Bhattacharyya, Minwalla, Papadodimas (2011) SUSY Markeviciute Santos (2016, 2018) Soliton, 1.5 Hairy Q=E/2Singular (S=0) Extremal CLP BH  $QL/N^2$ 1.0 Onset of **CLP** instability 0.5 CLP BH CLP BHs exist for 0<Q<Qext Hairy BHs exist for Onset<Q<E/2 (Singular - S=0 - in BPS lim)0.0 2 1 3 4  $EL/N^2$ 

• Hairy BHs always have <u>higher</u> entropy S than CLP BH with same E,Q

• Hairy BH temperature approaches T L=1/( $2\pi$ ) &  $\mu$ =1 in the singular (s=0) BPS lim !! This is also min T that CLP can reach in the singular extremal limit. → Microcononical phase diagram (of truncation with <u>two equal</u> charges)

Limit  $E \rightarrow 2Q$  of Hairy BHs ? Singular limit (S=0) coincides with soliton at E=2Q that we constructed numerically.

Charged scalar VEVs for the line of Hairy BH at fixed Q L/N<sup>2</sup>=0.5 (blue disks).

Red triangle is the SUSY soliton result at E=2Q=1.

• Soliton  $\exists$  for <u>any</u>  $E = 2Q \ge 0!$ 

Unlike when  $Q_1 = Q_2 = Q_3$ , where soliton  $\exists$  for  $E < E_{max}$ 

The Q<sub>1</sub>=Q<sub>2</sub>=Q<sub>3</sub> behaviour
 is believed to be generic.



Liu, Lü, Pope, Vazquez-Poritz '07 Bhattacharyya, Minwalla, Papadodimas (2011) Markeviciute Santos (2016, 2018)

[ Liu, Lü, Pope, Vazquez-Poritz '07] [ OD, Mitra, Santos '22 ]

# -> What happens when we <u>add</u> rotation?

Example: U(1)<sup>3</sup> gauged supergravity with two equal charges

→ Microcononical phase diagram (of truncation with two equal Q's): adding rotation

- When  $J_1 = J_2 = J$ , we can keep co-homogeneity <u>one</u> (10 coupled nonlinear ODEs)
- BPS relation is now E = 2Q + 2J.
- Work at finite temperature and approach  $T \rightarrow 0$ : want to find novel SUSY BHs!



→ Microcononical phase diagram (of truncation with two equal Q's): adding rotation

• Work at finite temperature and approach  $T \rightarrow 0$ : want to find novel SUSY BHs!



(as in the static case) & have higher entropy S than CLP for given {E,Q,J} !

 Hairy BHs have a non-singular lim (where T → 0, µ →1) in the BPS lim: novel SUSY BHs (this time with hair)! => can be missing grav parameter!

### → Conclusions & Future work

- Our hairy BHS with O(N<sup>2</sup>) entropy are dominant in the microcanonical ensemble
   => should be important for the microstate counting of the entropy of SUSY BHs:
   [See Benini's review talk at Strings 2022]
  - Static hairy BHs do <u>not</u> have a (smooth) BPS limit with  $O(N^2)$  entropy.
  - However, **rotating** hairy BHs **have** smooth BPS limit with  $O(N^2)$  entropy.
  - Microstate counting should compute the entropy of bald <u>and</u> hairy BHs!
- → Future work:
  - Complete Rotation study.
  - In the absence of charged scalar hair ( $\Phi$ K =0), all SUSY BHs are **1/16-BPS**. However, recent computations of a twisted SYM index in the so-called Macdonald limit ( $Q_3 + J_2 = 0$ ) suggest an O(N<sup>2</sup>) entropy for 1/8-BPS states

[Choi, Kim, Kim, Nahmgoong, 1810.12067]

It is unclear what solutions this index corresponds to in gauged supergravity. Can we have 1/8-BPS hairy BHs ?

- Classification scheme for SUSY BHs ?
- Generalizes to AdS4xS7?