Exact solution for motion of spinning particles near static, spherically symmetric objects

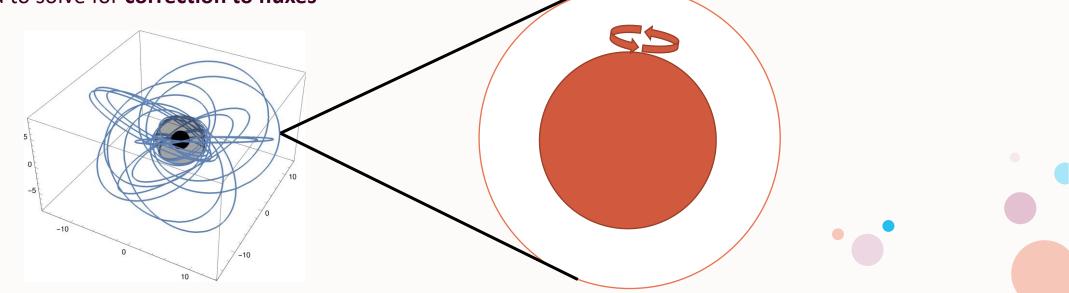
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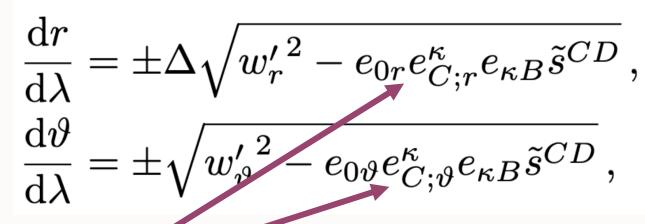
# . Why care about "spinning particle"?

- In compact object binaries, spin-orbital coupling appears as a conservative effect at linear order in mass ratio (assuming Kerr bound for black holes, mass-shedding limits for neutron stars)
- Also 1.5PN in a post-Newtonian expansion, 0.5PN earlier than non-trivial gravitational self-force (not obtainable from  $m \rightarrow \mu$  rescaling and Schwarzschild geodesics)
- That places it confidently into the **1PA iteration** of large-mass ratio waveforms.
  - Need to solve for motion perturbed by spin (linear order)
  - Need to solve for correction to fluxes



## Context

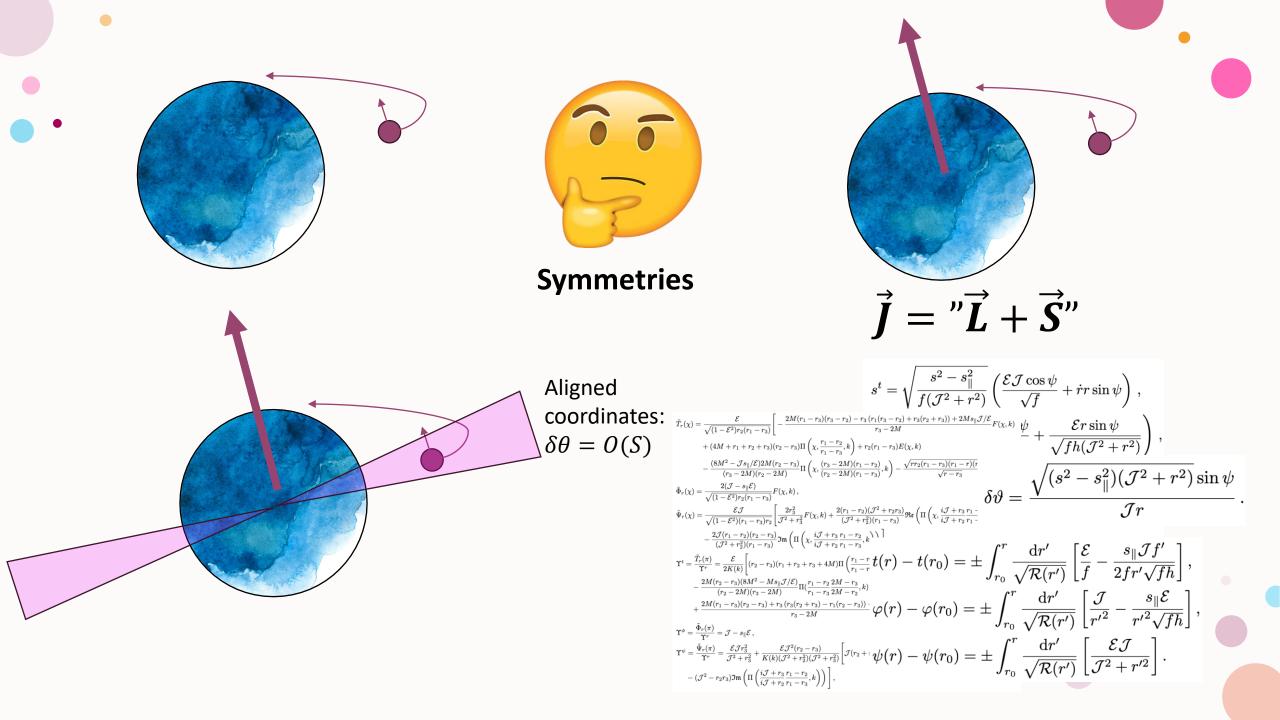
- 2018 (VW, Steinhoff, Lukes-Gerakopoulos <u>1808.06582</u>): Hamiltonians and canonical coordinates for spinning particles
- 2019 (VW, <u>1903.03651</u>): Hamilton-Jacobi equation, separability in Kerr, constants of motion  $\mathcal{E}_{so}, \mathcal{L}_{so}, \mathcal{K}_{so}, s_{\parallel}$ , solution depends on adapted Marck tetrad and its **connection**



Annoying to deal with (ask about connection terms and watch Gabriel's expression)

Sanity checks in equatorial plane, independent methods in special limits (Schwarzschild), contact with

Drummond & Hughes (<u>2201.13334</u>, <u>2201.13335</u>), Skoupý & Lukes-Gerakopoulos (<u>2102.04819</u>) Saijo+ (1998),...



Metric (why not general -> bosons stars and whatnot):

$$\mathrm{d}s^2 = -f(r)\mathrm{d}t^2 + h(r)\mathrm{d}r^2 + r^2\left(\mathrm{d}\theta^2 + \sin\theta^2\mathrm{d}\phi^2\right)$$

Killing constants of motion for spinning particle:

$$P^{\mu}\xi_{\mu}-\xi_{
ho;\sigma}S^{
ho\sigma}/2$$

Static, spherically symmetric (x,y,z axis rotations):

$$\mathcal{E} = -f\dot{t} + \frac{r^2\sin\theta f' s^{\phi}\dot{\theta} - s^{\theta}\dot{\phi}}{2\sqrt{fh}} \,,$$

$$\begin{aligned} \mathcal{J}_{x} &= -r^{2}(\sin\phi\theta + \cos\phi\cos\theta\sin\theta\phi) \\ &+ \sqrt{\frac{f}{h}} \Big[ \sin\theta\cos\phi h(s^{t}\dot{r} - s^{r}\dot{t}) + r\sin\phi\sin\theta(s^{\phi}\dot{t} - s^{t}\dot{\phi}) + r\cos\phi\cos\theta(s^{t}\dot{\theta} - s^{\theta}\dot{t}) \Big], \\ \mathcal{J}_{y} &= r^{2}(\cos\phi\dot{\theta} - \sin\phi\cos\theta\sin\theta\dot{\phi}) + \\ &\sqrt{\frac{f}{h}} \Big[ \sin\theta\sin\phi h(s^{t}\dot{r} - s^{r}\dot{t}) + r\cos\phi\sin\theta(s^{t}\dot{\phi} - s^{\phi}\dot{t}) + r\sin\phi\cos\theta(s^{t}\dot{\theta} - s^{\theta}\dot{t}) \Big], \\ \mathcal{J}_{z} &= r^{2}\sin^{2}\theta\dot{\phi} \\ &+ \sqrt{\frac{f}{h}} \Big[ \cos\theta h(s^{t}\dot{r} - s^{r}\dot{t}) + r\sin\theta(s^{\theta}\dot{t} - s^{t}\dot{\theta}) \Big]. \end{aligned}$$

# • Aligned frame $\vec{z'} \parallel \vec{J}$ : decoupling!!

$$\delta \vartheta = rac{\sqrt{fh}(s^r \dot{t} - s^t \dot{r})}{r^2 \dot{\varphi}} + \mathcal{O}(s^2) \,.$$

$$\begin{split} \dot{r}^2 &= \frac{1}{h} \left( -1 + \frac{\mathcal{E}^2}{f} - \frac{\mathcal{J}^2}{r^2} \right) - \frac{s_{\parallel} \mathcal{E} \mathcal{J} (2f - rf')}{(fh)^{3/2} r^2} ,\\ \dot{t} &= \frac{\mathcal{E}}{f} - \frac{s_{\parallel} \mathcal{J} f'}{2fr \sqrt{fh}} ,\\ \dot{\varphi} &= \frac{\mathcal{J}}{r^2} - \frac{s_{\parallel} \mathcal{E}}{r^2 \sqrt{fh}} . \end{split}$$

$$s_{\parallel} \equiv rac{l^{\mu}s_{\mu}}{\sqrt{l^{\nu}l_{\nu}}} \,, \; rac{\mathrm{d}s_{\parallel}}{\mathrm{d} au} = 0 + \mathcal{O}(s^2) \,.$$

### Spin sector? Parallel transport

Generalization of the tetrad of Marck (1983) for Kerr geodesics to any static, spherically symmetric space-time:

$$\begin{split} e^{\mu}_{(0)} &= \left( \mathcal{E}/f, \dot{r}, \mathcal{J}/r^{2}, 0 \right) ,\\ e^{\mu}_{(3)} &= \left( 0, 0, 0, \frac{1}{r} \right) , \end{split} \quad \tilde{e}^{\mu}_{(1)} &= \left( \frac{\dot{r}r\sqrt{h}}{\sqrt{f(\mathcal{J}^{2} + r^{2})}}, \frac{\mathcal{E}r}{\sqrt{fh(\mathcal{J}^{2} + r^{2})}}, 0, 0 \right) ,\\ \tilde{e}^{\mu}_{(2)} &= \left( \frac{\mathcal{E}\mathcal{J}}{f\sqrt{\mathcal{J}^{2} + r^{2}}}, \frac{\mathcal{J}\dot{r}}{r}, \frac{\sqrt{\mathcal{J}^{2} + r^{2}}}{r^{2}}, 0 \right) , \end{split}$$

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$$\dot{\psi} = rac{\mathcal{E}\mathcal{J}}{\mathcal{J}^2 + r^2}$$

$$s^{t} = \sqrt{\frac{s^{2} - s_{\parallel}^{2}}{f(\mathcal{J}^{2} + r^{2})}} \left(\frac{\mathcal{E}\mathcal{J}\cos\psi}{\sqrt{f}} + \dot{r}r\sin\psi\right), \quad s^{\varphi} = \frac{\sqrt{(s^{2} - s_{\parallel}^{2})(\mathcal{J}^{2} + r^{2})}\cos\psi}{r^{2}} \\ s^{r} = \sqrt{s^{2} - s_{\parallel}^{2}} \left(\frac{\mathcal{J}\dot{r}\cos\psi}{r} + \frac{\mathcal{E}r\sin\psi}{\sqrt{fh(\mathcal{J}^{2} + r^{2})}}\right) \\ s^{\vartheta} = \frac{s_{\parallel}}{r}.$$

#### Taster: radial motion in Schwarzschild

$$\lambda(r) - \lambda(r_2) = \int_{r_2}^r \frac{\mathrm{d}r'}{\sqrt{R(r)}} = \frac{2F(\chi, k)}{\sqrt{(1 - \mathcal{E}^2)(r_1 - r_3)r_2}},$$

$$R(r) \equiv (1 - \mathcal{E}^2)(r_1 - r)(r - r_2)(r - r_3)r,$$

$$r_{1,2} = \frac{p}{1 \mp e}$$

$$r_3 = \frac{2Mp}{p - 4M} - \frac{2s_{\parallel}\sqrt{Mp\left[p^2 - 4Mp + 4M^2(1 - e^2)\right]}}{(p - 4M)^2}$$

$$\begin{split} \mathcal{E}^2 &= \frac{(p-2M)^2 - 4M^2 e^2}{p \left[ p - M(3+e^2) \right]} \\ &- s_{\parallel} \frac{(e^2-1)^2 M \sqrt{Mp \left[ p^2 - 4Mp - 4M^2(e^2-1) \right]}}{p^2 \left[ p - M(3+e^2) \right]^2} \end{split}$$

ISCO position in closed form:

 $k^2 = rac{(r_1 - r_2)r_3}{(r_1 - r_3)r_2},$ 

$$p_{\rm c} = (6+2e)M - 2s_{\parallel}\sqrt{\frac{2(1+e)}{3+e}}$$

 $\sin \chi \equiv \sqrt{\frac{(r_1 - r_3)(r - r_2)}{(r_1 - r_2)(r - r_3)}},$ 

#### Inverting and t-motion (cf. van de Meent, <u>1906.05090</u>)

$$egin{aligned} r(\lambda) &= rac{r_3(r_1-r_2) ext{sn}^2 \left(rac{K(k)}{\pi} q^r, k
ight) + r_2(r_1-r_3)}{(r_1-r_2) ext{sn}^2 \left(rac{K(k)}{\pi} q^r, k
ight) - (r_1-r_3)} \,, \ q^r &\equiv \Upsilon^r \lambda + q_0^r \,, \ \Upsilon^r &\equiv rac{\pi \sqrt{(1-\mathcal{E}^2)(r_1-r_3)r_2}}{2K(k)} \,, \end{aligned}$$

$$t(\lambda) = q^t + \tilde{T}_r \left( \operatorname{am} \left( \frac{q^r}{\pi} K(k), k \right) \right) - \frac{\tilde{T}_r \left( \pi \right)}{2\pi} q^r, \ q^t \equiv \Upsilon^t \lambda + q_0^t, \ \Upsilon^t = \frac{\tilde{T}_r \left( \pi \right)}{\Upsilon^r}$$

$$\begin{split} \tilde{T}_{r}(\chi) &= \frac{\mathcal{E}}{\sqrt{(1-\mathcal{E}^{2})r_{2}(r_{1}-r_{3})}} \left[ -\frac{2M(r_{1}-r_{3})(r_{3}-r_{2})-r_{3}\left(r_{1}(r_{3}-r_{2})+r_{3}(r_{2}+r_{3})\right)+2Ms_{\parallel}\mathcal{J}/\mathcal{E}}{r_{3}-2M}F(\chi,k) \right. \\ &+ \left(4M+r_{1}+r_{2}+r_{3}\right)(r_{2}-r_{3})\Pi\left(\chi,\frac{r_{1}-r_{2}}{r_{1}-r_{3}},k\right)+r_{2}(r_{1}-r_{3})E(\chi,k) \\ &- \frac{\left(8M^{2}-\mathcal{J}s_{\parallel}/\mathcal{E}\right)2M(r_{2}-r_{3})}{(r_{3}-2M)(r_{2}-2M)}\Pi\left(\chi,\frac{(r_{3}-2M)(r_{1}-r_{2})}{(r_{2}-2M)(r_{1}-r_{3})},k\right)-\frac{\sqrt{rr_{2}(r_{1}-r_{3})(r_{1}-r)(r-r_{2})}}{\sqrt{r-r_{3}}}\right], \end{split}$$

#### Conclusions

- Spinning particle motion in spherically symmetric, static space-times fully solved by quadrature
- ISCO, energy, angular momentum shifts, turning points, all depend only on  $s_{\parallel}$
- $s_{\parallel}$ =0 means geodesic with precessing plane!
- For Schwarzschild full closed-form solution in terms of angle variables  $q^r, q^t, q^{\phi}, q^{\psi}$
- Lookout on arxiv for Witzany & Piovano 2307.XXXXX

