

# Progress toward inspiral-merger-ringdown waveforms at 1PA order

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# First-post-adiabatic (1PA) waveforms

- *Why?* Capra's longstanding goal: subradian waveform accuracy
- *What?* Solve Einstein equations through second order in mass ratio ( $\varepsilon$ ) using multiscale expansion

$$\Rightarrow h = \sum_{k^r, k^\theta, k^\phi} \left[ \varepsilon h_{k^A}^{(1)}(\mathcal{J}^A) + \varepsilon^2 h_{k^A}^{(2)}(\mathcal{J}^A) \right] e^{-ik^A \varphi_A}$$

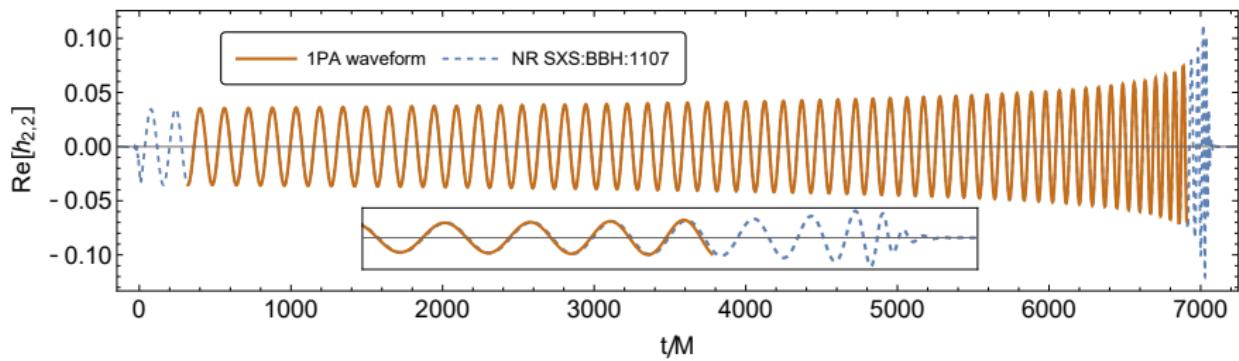
- orbital phases:  $\frac{d\varphi_A}{dt} = \Omega_A(\mathcal{J}^B)$
- slowly evolving parameters:

$$\frac{d\mathcal{J}^A}{dt} = \varepsilon \left[ F_{(0)}^A(\mathcal{J}^B) + \varepsilon F_{(1)}^A(\mathcal{J}^B) \right]$$

# 1PA waveforms for quasicircular, nonspinning binaries

[Wardell, AP, Warburton, Miller, Durkan, and Le Tiec 2021]

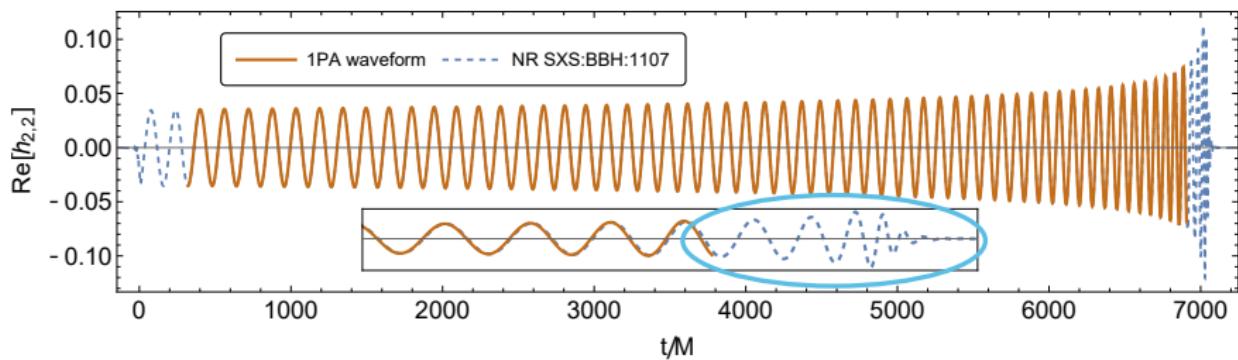
Mass ratio  $\varepsilon = 1/10$



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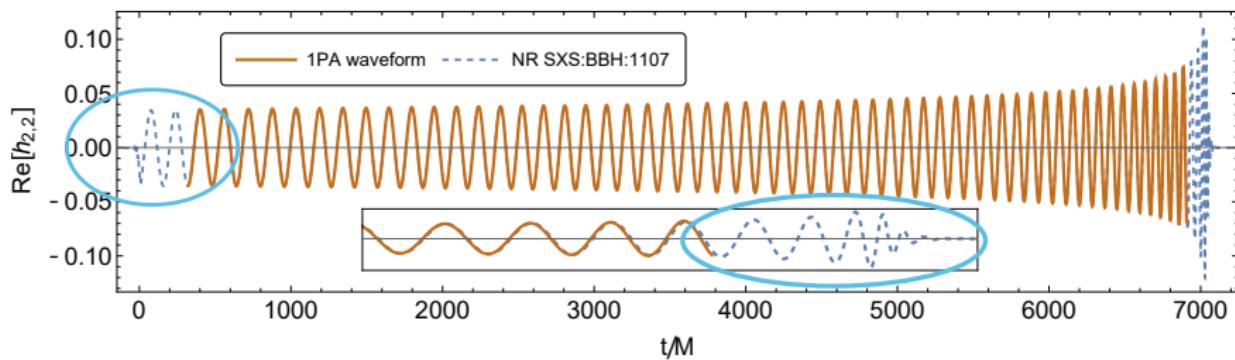
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# 1PA waveforms for quasicircular, nonspinning binaries

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# Multiscale expansion for quasicircular orbits in Schwarzschild [Miller & AP 2020]

- parameters:  $\mathcal{J}_A = (\Omega, M_{BH}, J_{BH})$ ,  $J_{BH} \sim \varepsilon$
- expansions:

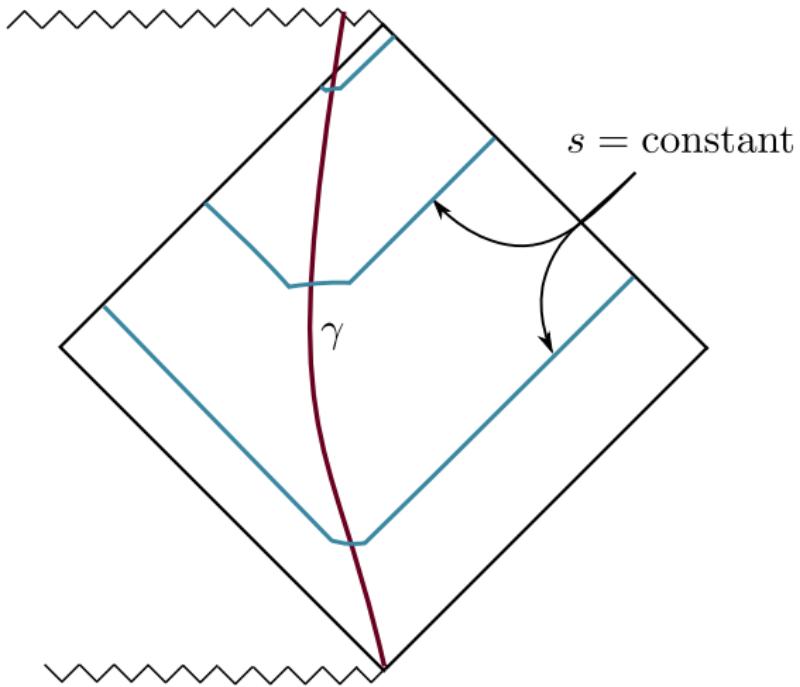
$$h_{\mu\nu}^{(n)} = \sum_{ilm} h_{ilm}^{(n)}(\mathcal{J}_A, r) e^{-im\phi_p} Y_{\mu\nu}^{ilm}$$

$$\frac{d\phi_p}{ds} = \Omega$$

$$\frac{d\Omega}{ds} = \nu \left[ F_\Omega^{(0)}(\Omega) + \nu F_\Omega^{(1)}(\mathcal{J}_A) \right]$$

- in Einstein equations,  $\frac{\partial}{\partial s} = \Omega \frac{\partial}{\partial \phi_p} + \frac{d\mathcal{J}_A}{ds} \frac{\partial}{\partial \mathcal{J}_A} = -im\Omega + O(\varepsilon)$
- solve field equations for amplitudes  $h_{ilm}^{(n)}$ . Precompute and store  
⇒ fast waveform generation.

# Choice of foliation



# Early inspiral

- PN expansion:  $x := (M_{\text{tot}}\Omega)^{2/3} \ll 1$
- evolution:

$$\frac{d\phi_p}{dt} = \Omega$$

$$\frac{d\Omega}{dt} = \nu x^{11/2} \left[ f_\Omega^{(0)}(\nu) + x f_\Omega^{(1)}(\nu) + x^{3/2} f_\Omega^{(1.5)}(\nu) + \dots + O(x^5) \right]$$

- $h = \nu x \sum_{n \geq 0} \sum_{\ell m} x^{n/2} H_{\ell m}^{(n)}(\nu) e^{-im\phi_p} {}_{-2}Y_{\ell m}$

# Transition to plunge

- evolution on timescale  $\sim 1/(\varepsilon^{1/5}\Omega)$  on frequency interval  $(\Omega - \Omega_{\text{isco}}) \sim \varepsilon^{2/5}$
- parameters:  $\mathcal{J}_A = \{\Delta\tilde{\Omega}, M_{BH}, J_{BH}\}$ ,  $\Delta\tilde{\Omega} := \frac{\Omega - \Omega_{\text{isco}}}{\varepsilon^{2/5}}$
- evolution:

$$\frac{d\phi_p}{dt} = \Omega_{\text{isco}} + \varepsilon^{2/5} \Delta\tilde{\Omega}$$

$$\frac{d\Delta\tilde{\Omega}}{dt} = \varepsilon^{1/5} \left[ F_{\Delta\tilde{\Omega}}^{(0)}(\Delta\tilde{\Omega}) + \varepsilon^{1/5} F_{\Delta\tilde{\Omega}}^{(1)}(\Delta\tilde{\Omega}) + O(\varepsilon^{2/5}) \right]$$

- $h_{\mu\nu} = \varepsilon \sum_{n \geq 0} \sum_{i\ell m} \varepsilon^{n/5} j_{i\ell m}^{(n)}(\mathcal{J}_A, r) e^{-im\phi_p} Y_{\mu\nu}^{i\ell m}$

# Plunge

- evolution on short time scale  $\sim 1/(\varepsilon^0 \Omega)$
- parameters:  $\mathcal{J}_A = \{\Omega, M_{BH}, J_{BH}\}$
- evolution:

$$\frac{d\phi_p}{dt} = \Omega$$

$$\frac{d\Omega}{dt} = F_\Omega^{(0)}(\Omega) + \varepsilon F_\Omega^{(1)}(\Omega) + O(\varepsilon^2)$$

- $h_{\mu\nu} = \sum_{i\ell m} \left[ \varepsilon h_{i\ell m}^{(1)}(\mathcal{J}^A, r) + \varepsilon^2 h_{i\ell m}^{(2)}(\mathcal{J}^A, r) + O(\varepsilon^3) \right] e^{-im\phi_p} Y_{\mu\nu}^{i\ell m}$

# Composite expansions

- Example: GSF & PN

$$\frac{d\Omega}{dt} = \nu F_\Omega^{(0)}(\Omega) + O(\nu^2)$$

$$\frac{d\Omega}{dt} = \nu x^{11/2} f_\Omega^{(0)}(\nu) + O(x^{13/2})$$

- Composite:

$$\frac{d\Omega}{dt} = \nu \left[ F_\Omega^{(0)}(\Omega) + x^{11/2} f_\Omega^{(0)}(\nu) - x^{11/2} f_\Omega^{(0)}(0) + O(\nu) \right]$$

# Continued-fraction resummation

$$M_{\text{Bondi}}(s) = E_{\text{bind}}(\mathcal{J}_A(s)) + m + M_{BH}(s)$$

$$\Rightarrow -\mathcal{F}_{\infty} = \frac{\partial E_{\text{bind}}}{\partial \Omega} \dot{\Omega} + \frac{\partial E_{\text{bind}}}{\partial M_{BH}} \dot{M}_{BH} + \frac{\partial E_{\text{bind}}}{\partial J_{BH}} \dot{J}_{BH} + \dot{M}_{BH}$$

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$$\Rightarrow \dot{\Omega} = \frac{-\mathcal{F}}{\partial E_{\text{bind}} / \partial \Omega}$$

# Continued-fraction resummation

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$$\Rightarrow \dot{\Omega} = \frac{-\mathcal{F}}{\partial E_{\text{bind}} / \partial \Omega}$$

$$= \frac{\nu^2 \mathcal{F}^{(1)}(\Omega) + \nu^3 \mathcal{F}^{(2)}(\Omega) + \dots}{\nu \partial E_{\text{bind}}^{(0)} / \partial \Omega + \nu^2 \partial E_{\text{bind}}^{(1)} / \partial \Omega + \dots}$$

# Continued-fraction resummation

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$$\Rightarrow \dot{\Omega} = \frac{-\mathcal{F}}{\partial E_{\text{bind}} / \partial \Omega}$$

$$= \frac{\nu^2 \mathcal{F}^{(1)}(\Omega) + \nu^3 [\mathcal{F}_R^{(2)}(\Omega) + \dot{\Omega}_0 \mathcal{F}_S^{(2)}(\Omega)] + \dots}{\nu \partial E_{\text{bind}}^{(0)} / \partial \Omega + \nu^2 \partial E_{\text{bind}}^{(1)} / \partial \Omega + \dots}$$

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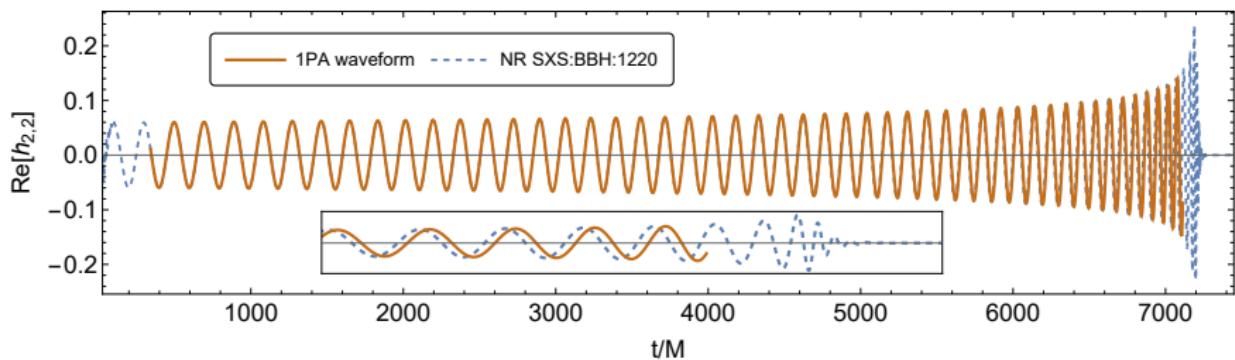
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$$\Rightarrow \dot{\Omega} = \frac{\nu^2 \mathcal{F}^{(1)}(\Omega) + \nu^3 \left[ \mathcal{F}_R^{(2)}(\Omega) + \frac{\mathcal{F}^{(1)}(\Omega)}{\partial E_{\text{bind}}^{(0)}/\partial \Omega + \nu \partial E_{\text{bind}}^{(1)}/\partial \Omega + \dots} \mathcal{F}_S^{(2)}(\Omega) \right] + \dots}{\nu \partial E_{\text{bind}}^{(0)}/\partial \Omega + \nu^2 \partial E_{\text{bind}}^{(1)}/\partial \Omega + \dots}$$

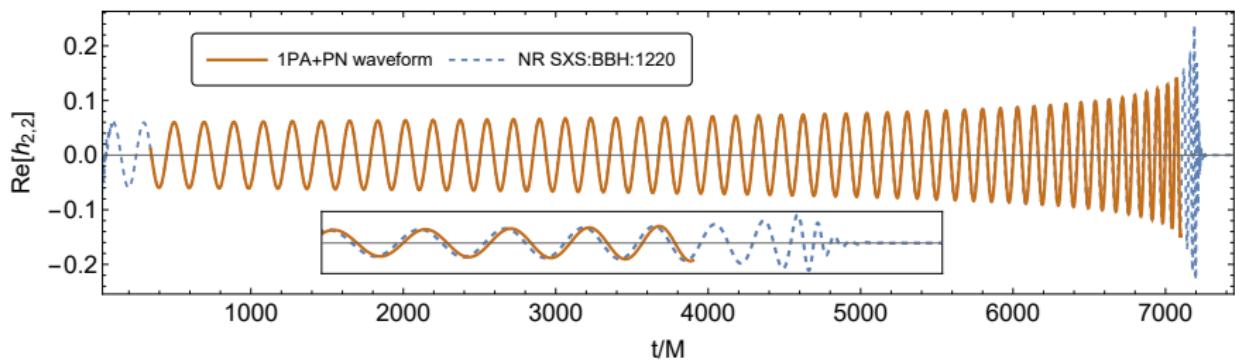
# 1PA waveforms + PN information

Mass ratio  $\varepsilon = 1/4$



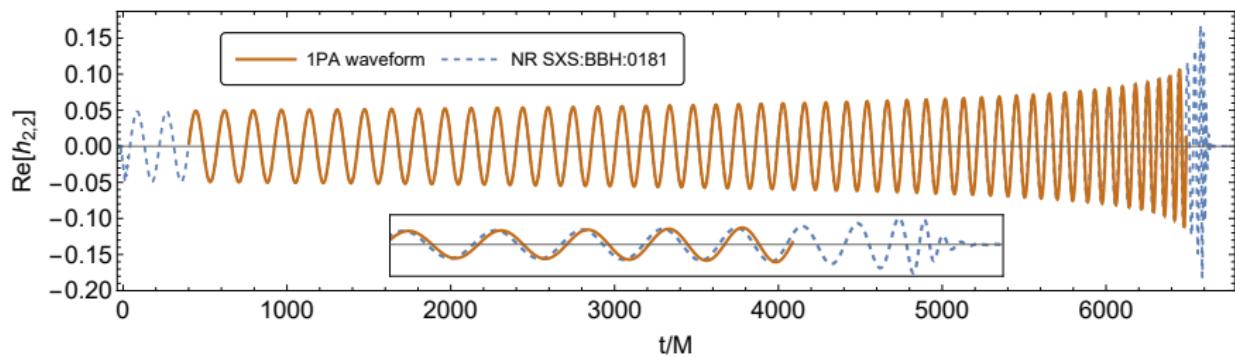
# 1PA waveforms + PN information

Mass ratio  $\varepsilon = 1/4$



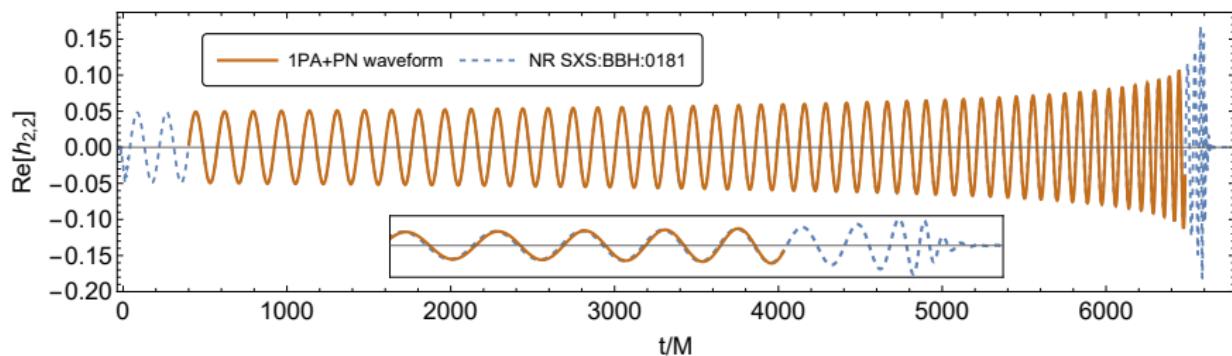
# 1PA waveforms + PN information

Mass ratio  $\varepsilon = 1/6$



# 1PA waveforms + PN information

Mass ratio  $\varepsilon = 1/6$



# Conclusion

- Progress toward completing 1PA waveforms
- What's next?
  - merger & ringdown implementation (talk by Lorenzo Kuchler)
  - Kerr (morning talks, discussion session)
  - eccentricity (talk by Ben Leather)

## Second-order discussion session

- public release of 2SF tools on BHPToolkit
- upcoming workshops:
  - online series on 1PA waveform generation
  - Princeton, Oct 10-13 (<https://pcts.princeton.edu/events/2023/nonlinear-aspects-general-relativity>)
- paths to Kerr

# Paths to Kerr

- Difficulties
  - standard CCK reconstruction produces pathological singularities  
⇒ problematic source at second order
  - numerical burden in any case is calculation of quadratic source term
    - Have we made our lives harder by seeking maximum separability?
- Possibilities
  - Easier in an  $m$ -mode decomposition? (Osburn, Panosso Macedo)
  - GHZ puncture scheme (Bourg)
  - Use Lorenz gauge, either reconstructed (Dolan) or directly (Osburn)

# Multiscale expansion

[Miller & AP; AP & Wardell; Flanagan, Hinderer, Moxon, AP]

- all time dependence in  $h_{\mu\nu}$  follows from puncture's motion and black hole's evolution
- recall  $(z^\mu, u^\mu) \rightarrow (\varphi_A, J_A)$ . Define full set of system parameters  $\mathcal{J}_A \sim (J_A, M, a)$

$$\frac{d\varphi_A}{ds} = \Omega_A(\mathcal{J}_B)$$

$$\frac{d\mathcal{J}_A}{ds} = \varepsilon F_A^{(0)}(\mathcal{J}_B) + \varepsilon^2 F_A^{(1)}(\mathcal{J}_B) + O(\varepsilon^3)$$

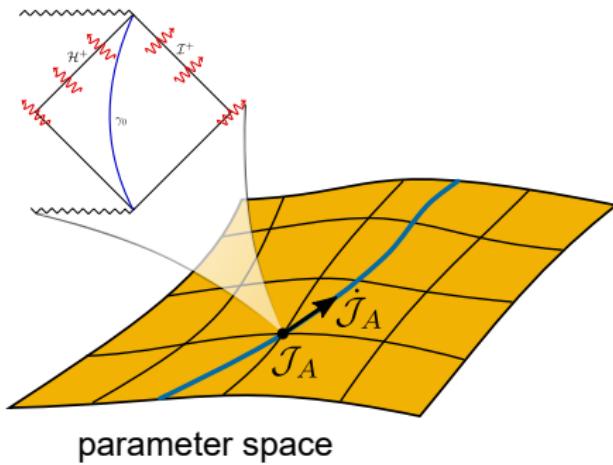
- treat  $h_{\mu\nu}$  as function on extended manifold: in spacetime coords  $(s, x^i)$ ,

$$h_{\mu\nu} = \varepsilon h_{\mu\nu}^{(1)}(\varphi_A, \mathcal{J}_A, x^i) + \varepsilon^2 h_{\mu\nu}^{(2)}(\varphi_A, \mathcal{J}_A, x^i) + O(\varepsilon^3)$$

- in Einstein equations,  $\frac{\partial}{\partial s} = \Omega_A \frac{\partial}{\partial \varphi_A} + \frac{d\mathcal{J}_A}{ds} \frac{\partial}{\partial \mathcal{J}_A}$

# Multiscale expansion

[Miller & AP; AP & Wardell; Flanagan, Hinderer, Moxon, AP]



- Fourier series:
$$h_{\mu\nu}^n = \sum_{k^A} h_{\mu\nu}^{(n,\omega_k)}(\mathcal{J}_A, x^i) e^{-ik^A \varphi_A}$$
$$\omega_k := k^A \Omega_A$$
- solve field equations for amplitudes  $h_{\mu\nu}^{(n,\omega_k)}$  on grid of  $\mathcal{J}_A$  values
- millisecond waveform generation when combined with FastEMRIWaveforms tools [Katz, Chua, Speri, Warburton, Hughes]

# Field equations

$$G_{\mu\nu}^{(1)}[h^{\mathcal{R}(1)}] = -G_{\mu\nu}^{(1)}[h^{\mathcal{P}(1)}]$$

$$G_{\mu\nu}^{(1)}[h^{\mathcal{R}(2)}] = -G_{\mu\nu}^{(2)}[h^{(1)}, h^{(1)}] - G_{\mu\nu}^{(1)}[h^{\mathcal{P}(2)}]$$

# Field equations

$$G_{\omega_k \ell m}^{(1,0)}[h_{\omega_k \ell m}^{\mathcal{R}(1)}] = -G_{\omega_k \ell m}^{(1,0)}[h_{\omega_k \ell m}^{\mathcal{P}(1)}]$$

$$G_{\mu\nu}^{(1)}[h^{\mathcal{R}(2)}] = -G_{\mu\nu}^{(2)}[h^{(1)}, h^{(1)}] - G_{\mu\nu}^{(1)}[h^{\mathcal{P}(2)}]$$

# Field equations

$$G_{\omega_k \ell m}^{(1,0)}[h_{\omega_k \ell m}^{\mathcal{R}(1)}] = -G_{\omega_k \ell m}^{(1,0)}[h_{\omega_k \ell m}^{\mathcal{P}(1)}]$$

$$\begin{aligned} G_{\omega_k \ell m}^{(1,0)}[h_{\omega_k}^{\mathcal{R}(2)}] &= -G_{\omega_k \ell m}^{(2,0)}[h_{\omega_{k'} \ell' m'}^{(1)}, h_{\omega_{k''} \ell'' m''}^{(1)}] - G_{\omega_k \ell m}^{(1,0)}[h_{\omega_k \ell m}^{\mathcal{P}(2)}] \\ &\quad - G_{\omega_k \ell m}^{(1,1)}[h_{\omega_k \ell m}^{(1)}] \end{aligned}$$

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## Necessary inputs:

- $h_{\omega_k \ell m}^{\mathcal{P}(2)} \rightarrow h_{\mu\nu}^{\text{R}(1)}(r_0), \nabla_\rho h_{\mu\nu}^{\text{R}(1)}(r_0)$
- $G_{\omega_k \ell m}^{(2,0)} \rightarrow h_{\omega_k \ell m}^{(1)}(r), h_{\omega_k \ell m}^{\mathcal{R}(1)}(r), \partial_r h_{\omega_k \ell m}^{\mathcal{R}(1)}(r), \partial_r^2 h_{\omega_k \ell m}^{\mathcal{R}(1)}(r)$
- $G_{\omega_k \ell m}^{(1,1)}[h_{\omega_k \ell m}^{(1)}] \rightarrow F_A^{(0)} \frac{\partial}{\partial \mathcal{J}_A} h_{\omega_k \ell m}^{(1)}$