

# An on-shell approach to radiation reaction

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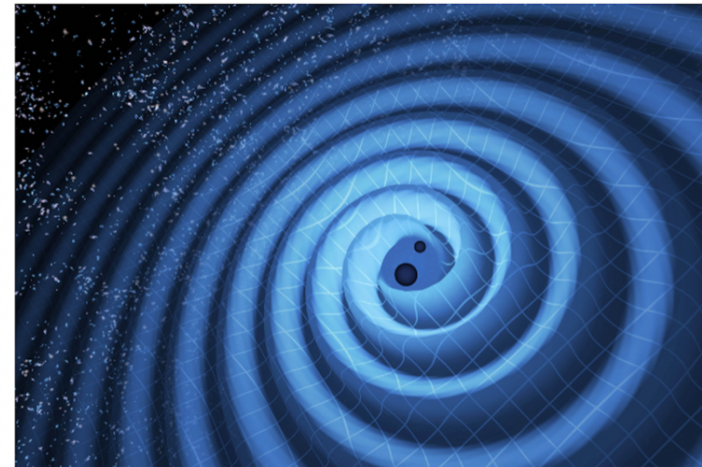
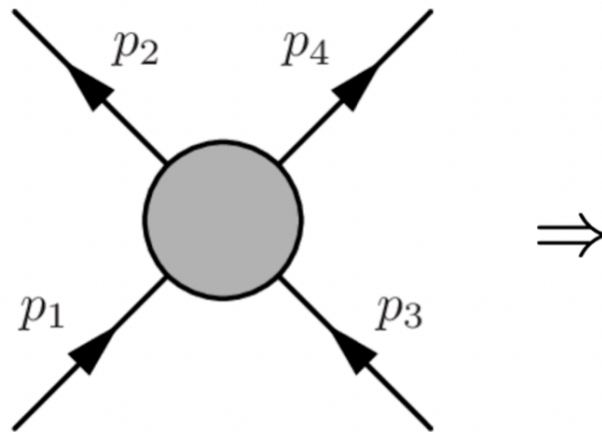
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# The two-body problem

- 2019: State-of-the-art calculation in the [post-Minkowskian](#) approximation using only the classical limit of [scattering amplitudes](#) ([Bern](#), [Cheung](#), [Roiban](#), [Shen](#), [Solon](#), [Zeng](#))



Credit: Tim Pyle

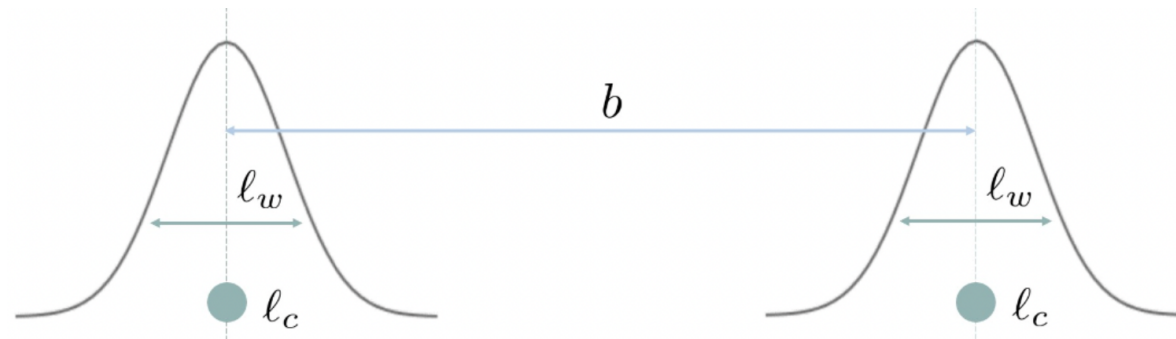
Main message of this talk

We can use on-shell amplitudes to revisit the classical two-body problem

# The KMOC formalism

- **Binary system** as superposition of single particle states

$$|\psi\rangle = \int_{p_1, p_2} \phi_1(p_1) \phi_2(p_2) e^{\frac{ib \cdot p_1}{\hbar}} |p_1 p_2\rangle$$



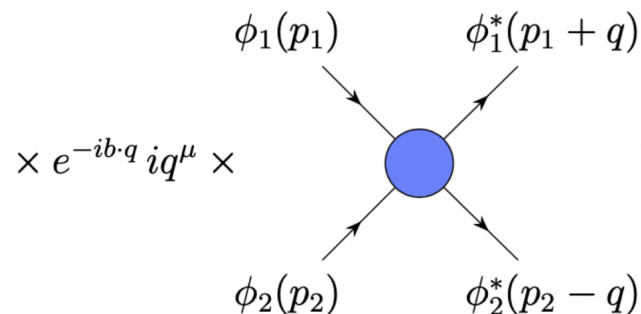
Credit: Ben Maybee, 2105.10268

- **Interactions** are expressed in terms of **scattering amplitudes**

$$S|\psi\rangle = \int_{p_1, p'_1, p_2, p'_2} \phi_1(p_1) \phi_2(p_2) e^{\frac{ib \cdot p_1}{\hbar}} \underbrace{\langle p'_1, p'_2 | S | p_1 p_2 \rangle}_{S=1+iT} |p'_1, p'_2\rangle + \dots$$

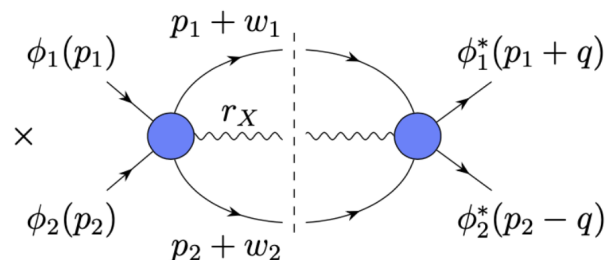
- The change in momentum  $\langle \psi | S^\dagger \mathbb{P}^\mu S | \psi \rangle - \langle \psi | \mathbb{P}^\mu | \psi \rangle$  is the sum of two contributions. One is linear in the  $T$ -matrix

$$I_{(1)}^\mu = \int d\Phi(p_1) d\Phi(p_2) d^4q \hat{\delta}(2p_1 \cdot q + q^2) \hat{\delta}(2p_2 \cdot q - q^2) \Theta(p_1^0 + q^0) \Theta(p_2^0 - q^0)$$



- The remaining one is quadratic in the  $T$ -matrix

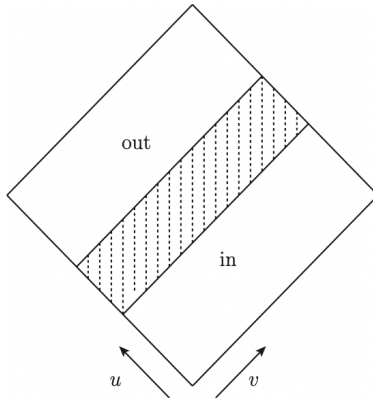
$$I_{(2)}^\mu = \sum_X \int \prod_{i=1,2} d\Phi(p_i) d^4w_i d^4q \hat{\delta}(2p_i \cdot w_i + w_i^2) \Theta(p_i^0 + w_i^0) \times \hat{\delta}(2p_1 \cdot q + q^2) \hat{\delta}(2p_2 \cdot q - q^2) \Theta(p_1^0 + q^0) \Theta(p_2^0 - q^0) \times e^{-ib \cdot q / \hbar} w_1^\mu \hat{\delta}^{(4)}(w_1 + w_2 + r_X)$$



Credit: Kosower, Maybee and O'Connell, 1811.10950

# KMOC on curved backgrounds

- The same logic applies for QFT on a **gravitational plane wave**



$$ds^2 = 2 dudv - H_{ab}(u) x^a x^b du^2 - dx^\perp dx^\perp$$

- The **geodesic motion** is captured by a 2-point amplitude (memory effects related to  $E_{ai}(u) \sim_{u \rightarrow +\infty} \delta_{ia} + u C_{ai}$ )

$$\underbrace{\mathcal{S}_\alpha}_{S \text{ on } \bar{g}} |\psi\rangle = \int d\Phi(p, p') \phi(p) \underbrace{\langle p' | \mathcal{S}_\alpha | p \rangle}_{2\text{-point}} |p'\rangle + \dots$$

$$\langle p' | \mathcal{S}_\alpha | p \rangle = \hat{\delta}^+(p' - p) \frac{4\pi e^{-\frac{i}{2\sqrt{G}p_+}(p'-p)_\perp} c^{-1}(p'-p)_\perp}}{\sqrt{\det(G \ c)} \hbar}$$

- **2-points** on the background  $\Rightarrow$  **geodesic motion**

$$\langle \psi | \mathcal{S}_\alpha^\dagger \mathbb{P}^\mu \mathcal{S}_\alpha | \psi \rangle = p^\mu + \sqrt{G} p_+ \delta_a^\mu c_i^a b^i - n^\mu \left( \frac{2\sqrt{G} p_+ p_a c_i^a b^i - G p_+^2 c_i^a b^i c_j^b b^j \delta_{ab}}{2p_+} \right)$$

- **3-points** on the background  $\Rightarrow$  **radiative observables.**

**Deviations** from geodesic motion appear in powers of  $m/\Lambda$

$$\langle \psi | \mathcal{S}_\alpha^\dagger \mathbb{W}_{\mu\nu\sigma\rho}(u, \hat{x}) \mathcal{S}_\alpha | \psi \rangle = \frac{i\kappa}{2\pi\hbar^{\frac{1}{2}}} \int_0^\infty \hat{d}\omega e^{-i\omega u} k_{[\mu} \varepsilon_{\nu]}^{-\eta} k_{[\sigma} \varepsilon_{\rho]}^{-\eta} \times \int d\Phi(p) \underbrace{\langle \psi | \mathcal{S}^\dagger | p \rangle \langle p, k^\eta | \mathcal{S} | \psi \rangle}_{2\text{-point} \times 3\text{-point}} \Big|_{k=\hbar\omega\hat{x}} + \text{c.c.}$$

All order waveforms from amplitudes (Adamo, Cristofoli, Klisch, Ilderton)

SF approximation on a plane wave from amplitudes on a background

# Example: electromagnetism

- **3-points** define scattering waveforms and radiation reaction

$$\underbrace{A_\mu(x)}_{\text{Background}} = -x^\perp E_\perp(x^-) n_\mu \quad , \quad \underbrace{a_\perp(u)}_{U(1) \text{ memory}} := \int_{-\infty}^u ds E_\perp(s)$$

$$\underbrace{\langle p', k^\eta | S | p \rangle}_{\text{3-point}} = \delta^{+,\perp}(q - k - ea(\infty)) \int_y \varepsilon^\eta \cdot P(y) e^{i \int_{-\infty}^y dz k \cdot P(z)}$$

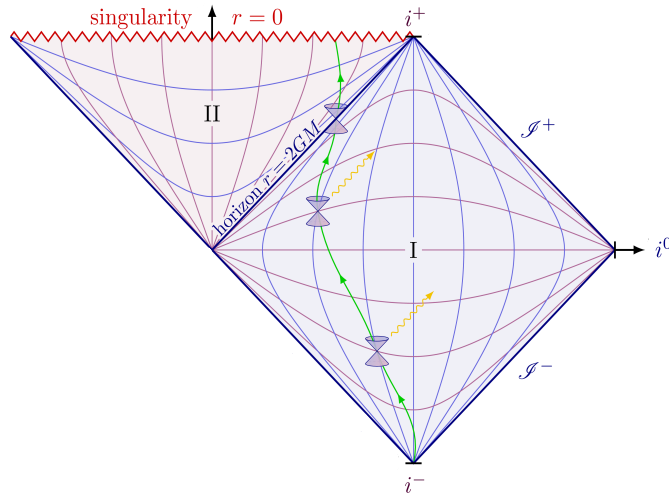
- **Scattering waveform** from a frequency integral over a 3-point

$$W_{\mu\nu}(u, \hat{x}) = -\frac{ie}{4\pi^2 p_+} \int_{y,\omega} \omega e^{-i\omega(u - \hat{x} \cdot X(y))} \hat{x}_{[\mu} P_{\nu]}(y)$$

- **Radiation reaction** from the square of a 3-point

$$\Delta p^\mu = -\frac{2}{3} \frac{e^4 p_+}{4\pi m^4} \int_{-\infty}^{+\infty} du E_\perp^2(u) P^\mu(u) - n^\mu \frac{P(u) \cdot P(+\infty)}{p_+}$$

- Consider QFT on a **static background** (e.g. Schwarzschild)



$$ds^2 = \left( \eta_{\mu\nu} + k_\mu k_\nu \frac{2GM}{r} \right) dx^\mu dx^\nu$$

- The **geodesic motion** is captured by the **classical limit** of a 2-point (equivalent to an eikonal amplitude in vacuum)

$$\underbrace{\mathcal{S}_\alpha}_{S \text{ on } \bar{g}} |\psi\rangle = \int d\Phi (p, p') \phi(p) \underbrace{\langle p' | \mathcal{S}_\alpha | p \rangle}_{2\text{-point}} |p'\rangle + \dots$$

$$\langle p' | \mathcal{S}_\alpha | p \rangle = \delta(E' - E) \sum_{\ell=0}^{\infty} (2\ell + 1) P_\ell(\hat{p} \cdot \hat{p}') e^{\underbrace{2i l_\ell(r = \infty)}_{\text{radial action}} - i\pi\ell}$$



- **2-points** on Schwarzschild are function of the **radial action**  $I(r)$

$$\underbrace{\langle p' | \mathcal{S}_\alpha | p \rangle}_{\text{geodesic motion}} = \delta(E' - E) N \int d^2 x^\perp e^{i x^\perp \cdot (\hat{p} - \hat{p}')} (2I(|x^\perp|) - \pi |\vec{p}'| |x^\perp|)$$

$$I(r) = \int_{r_{\text{turn}}}^r ds \sqrt{\frac{p^2 s^2 + 2GMm s - \frac{s-2GM}{s} \ell^2}{(s - 2GM)^2}}$$

- The classical limit of a **3-point**, on the other hand, is determined by **Hamilton's principal function**  $S(t, r, \phi, \theta)$

$$\underbrace{\langle p', k^\eta | \mathcal{S}_\alpha | p \rangle}_{\text{radiation}} = -\kappa \int_{\mathbb{R}^{1,3} \setminus \overline{B(r_s)}} d^4 x \sqrt{-|g|} \int_{I, I', k'} \Lambda^{*,p}(I) \Lambda^{p'}(I')$$

$$\times \Lambda_{\mu\nu}^k{}^{\rho\sigma}(k') \mathcal{E}_{\rho\sigma}^\eta \left[ 2\partial^\mu S_{I'} \partial^\nu S_I - g^{\mu\nu} \left( \partial S_{I'} \cdot \partial S_I + \frac{m^2}{2} \right) \right] e^{i(S_{k'} + S_{I'} - S_I)}$$

- We can take the **weak field limit** by considering a linearized **black hole** (e.g. Schwarzschild in de Donder gauge at order  $G$ )

$$\Lambda^P(l) = \delta(p_0 - l_0) \underbrace{\int d^2 x^\perp e^{i x^\perp \cdot (\hat{p} - \hat{l})} (2l_{lin.}(|x^\perp|) - \pi |\vec{p}| |x^\perp|)}_{f^{eik.}(p,l)}$$

$$\Lambda_{\mu\nu}^{k' \rho\sigma}(k) \Big|_{\text{lin. Schw.}} = -2 \delta_{(\mu}^{\rho} \delta_{\nu)}^{\sigma} \delta(k'_0 - k_0) f^{eik.}(k, k')$$

- The weak field limit of a 3-point on **Schwarzschild** reproduces the probe limit of a **5-point amplitude** on a **flat background**

$$\underbrace{\langle p', k | S_\alpha | p \rangle}_{S\text{-matrix on } \bar{g}} = \frac{\hat{\delta}(U \cdot (q + k))}{2M} \underbrace{\mathcal{A}_5(p, P \rightarrow p + q, P - q - k, k)}_{S\text{-matrix on } \eta} + \dots$$

Main message (Adamo, Cristofoli, Klisch, Ilderton)

On-shell amplitudes on a background are the building blocks for self-force