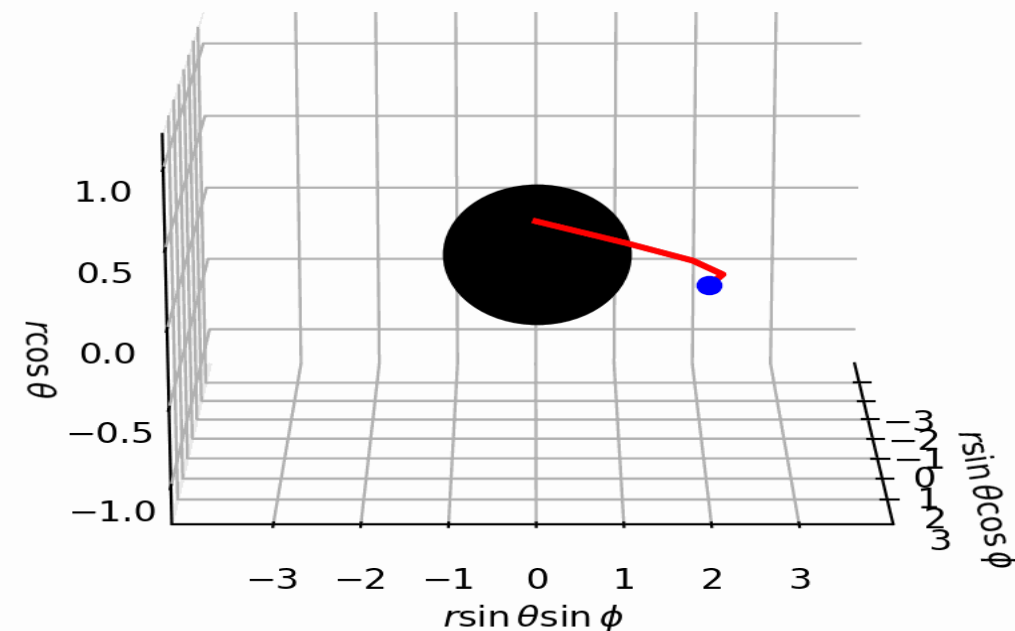


Adiabatic vs Post-adiabatic

Circular Schwarzschild Orbits

Ollie Burke, Chris Kavanagh, Niels Warburton
Philip Lynch, Barry Wardell, Lorenzo Speri

Near_Plunge: Eccentric orbit into a rotating black hole
 $M = 10^6 M_\odot$, $\mu = 10 M_\odot$, $a = 0.9$, $p_0 = 2.9$, $e_0 = 0.3$, $t_0 = 0.3$

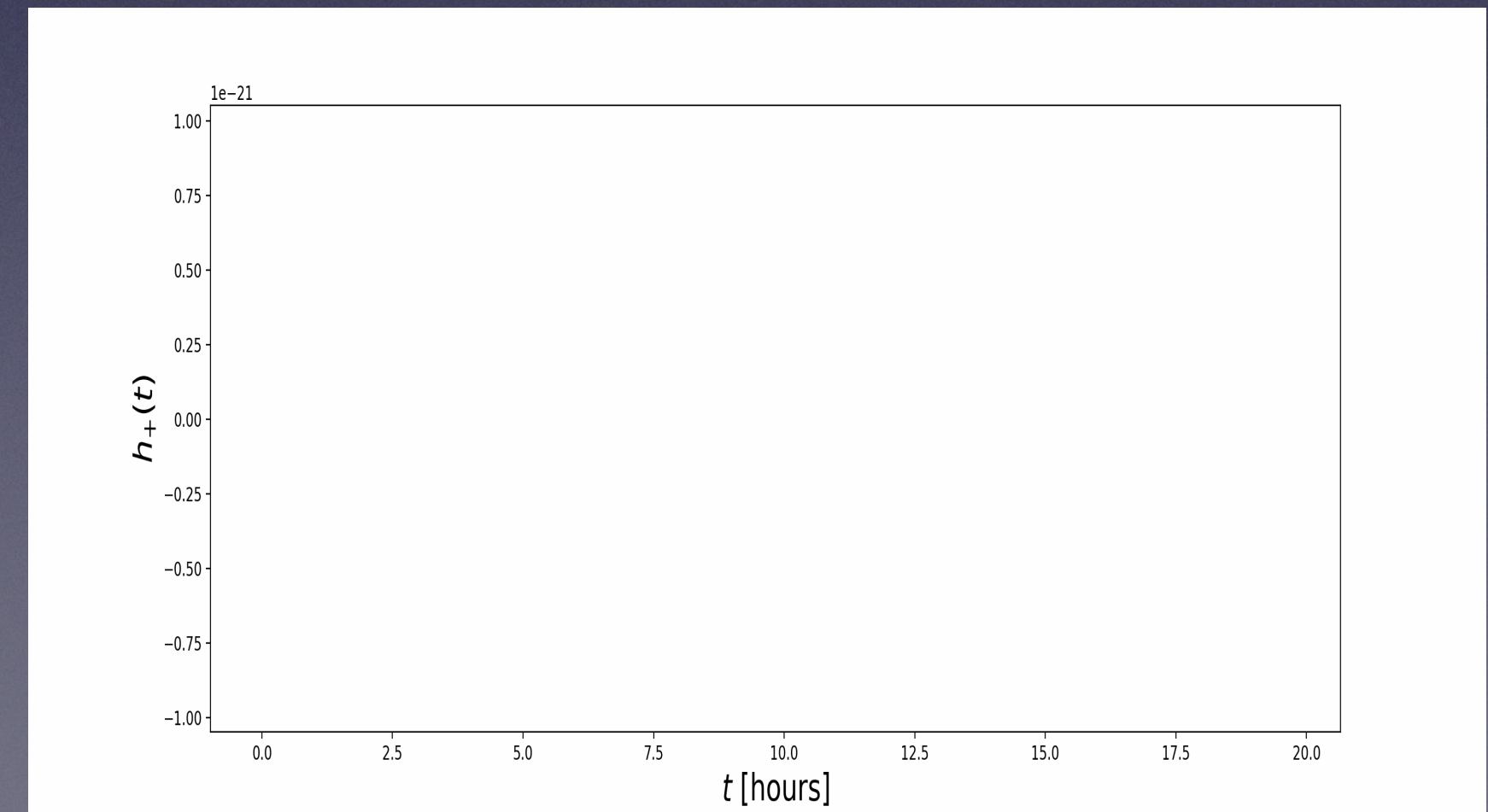


ollie.burke@l2it.in2p3.fr



<https://github.com/OllieBurke/animations>

Capra, 05/07/23
Copenhagen



The structure

- 1. Importance of accuracy:** The global picture
- 2. Circular Schwarzschild Orbits:** 0PA vs 1PA — How well can we do?
- 3. Concluding remarks :** Take home message + Discussion

Part 1: **Accuracy**

The global picture

Parameter estimation
what is important?

Cover parameter
space



Why?

Failure to claim
Detection

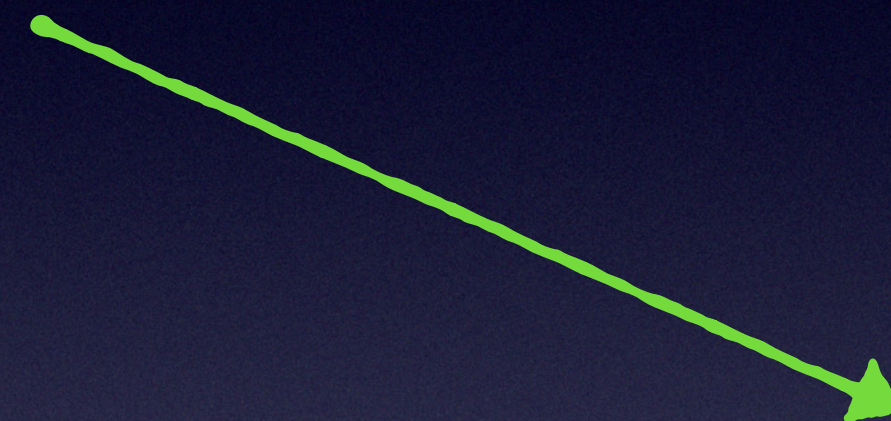


Fast to evaluate



Why?

We want results within
a Hubble time



Faithful

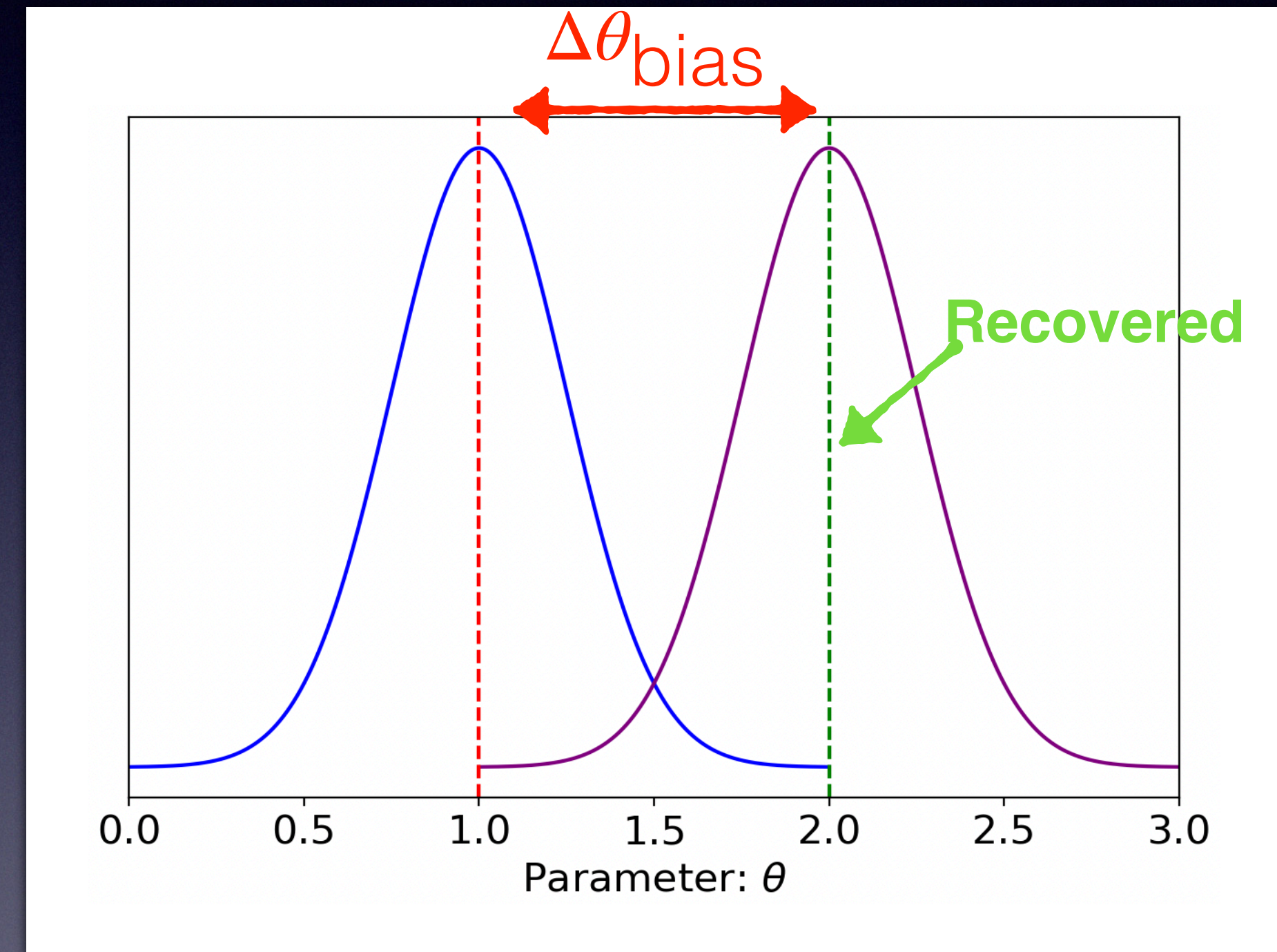
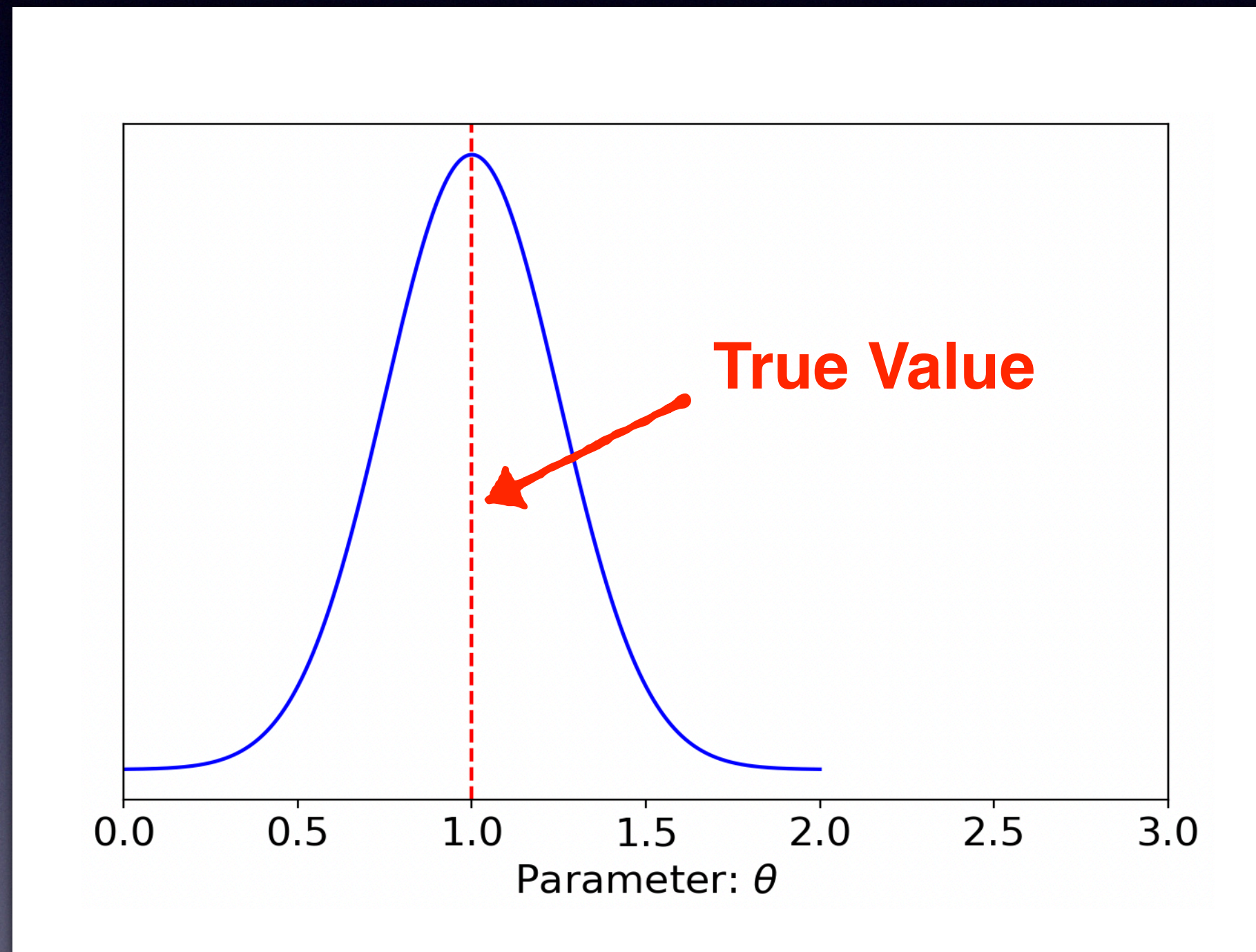


Why?

Potentially wrong
parameters recovered

The importance of accuracy

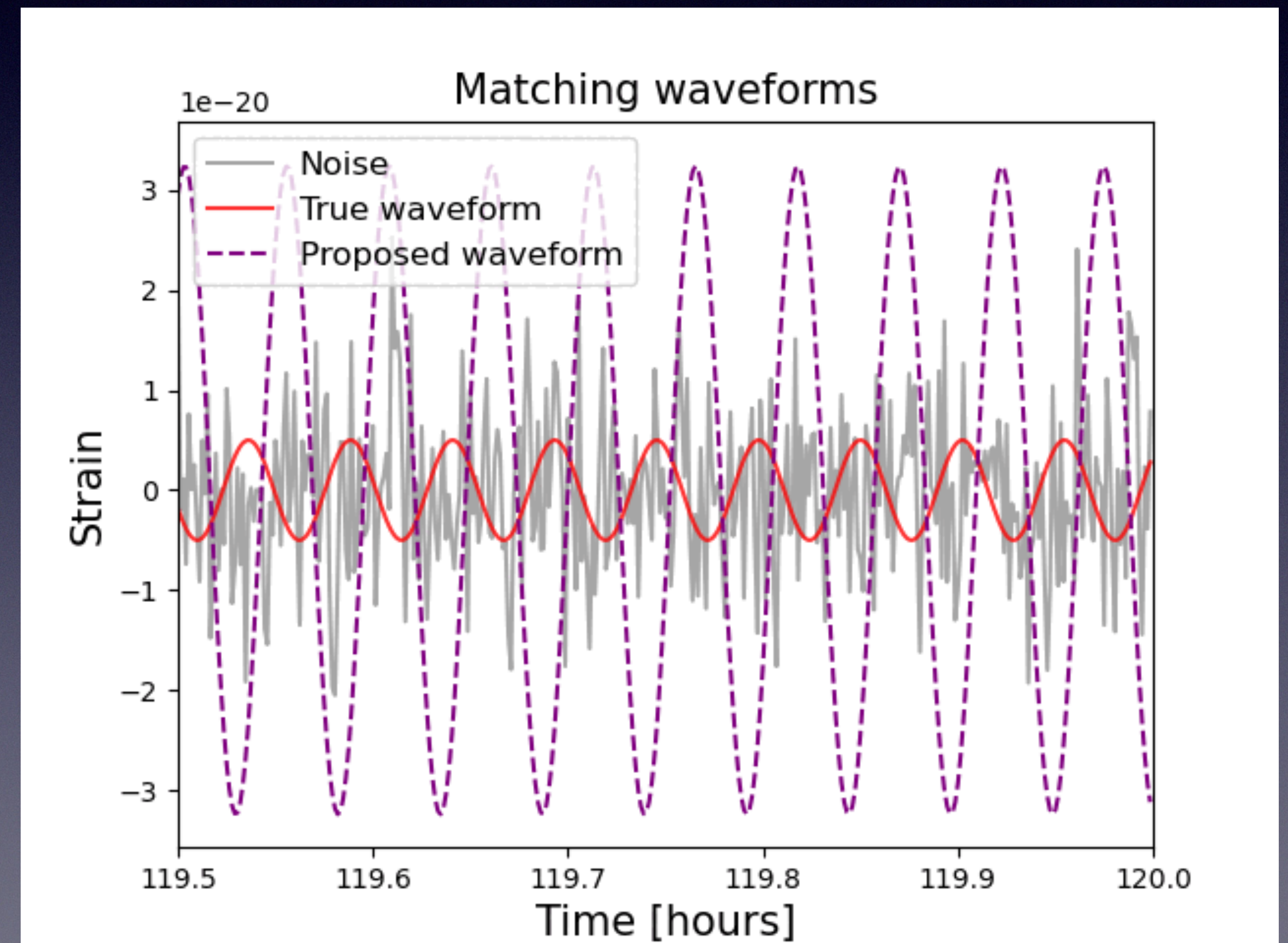
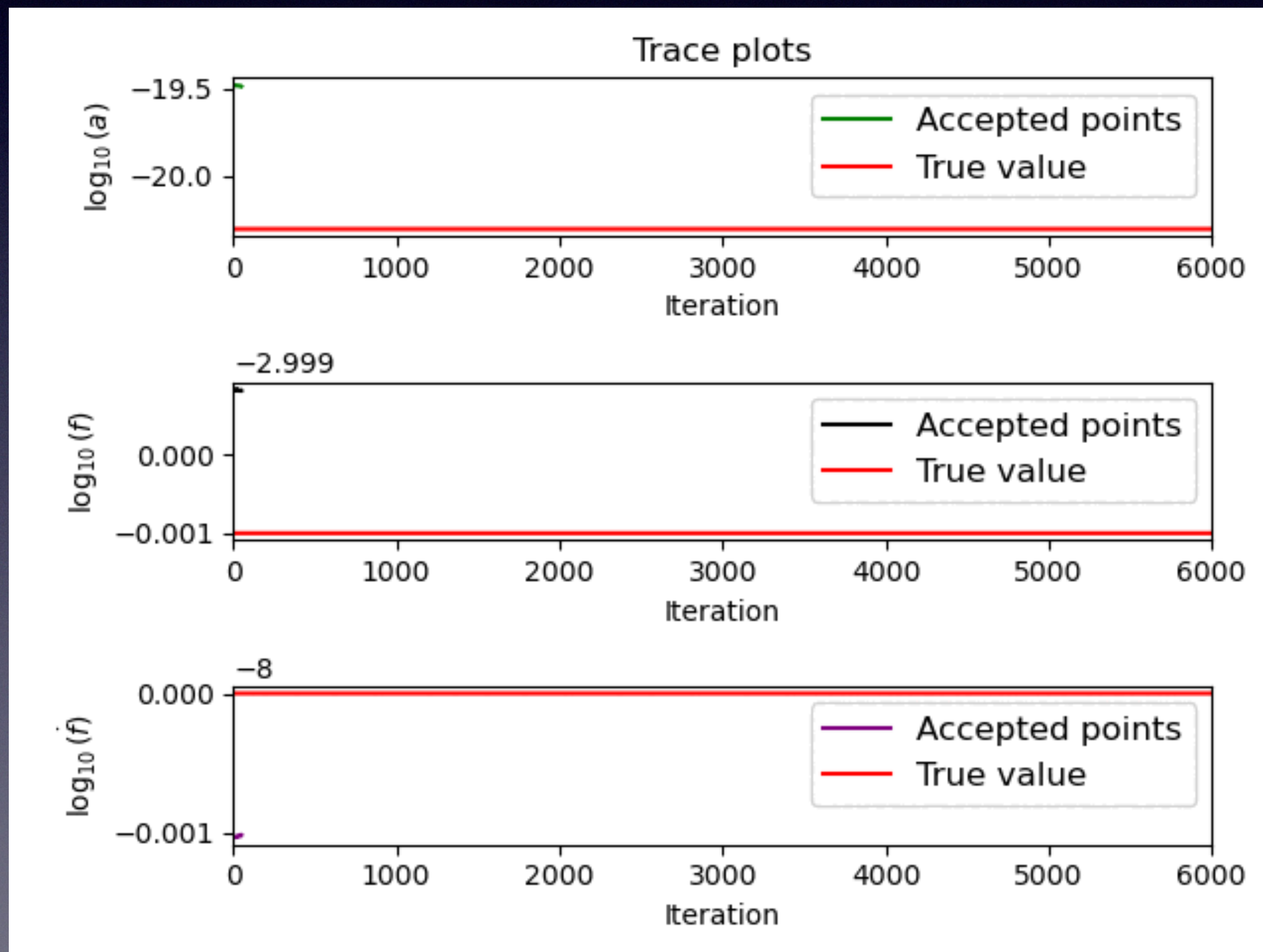
Inaccurate Waveforms \implies Biases in Parameters



Goal: Recovered parameter within 1σ width of true distribution

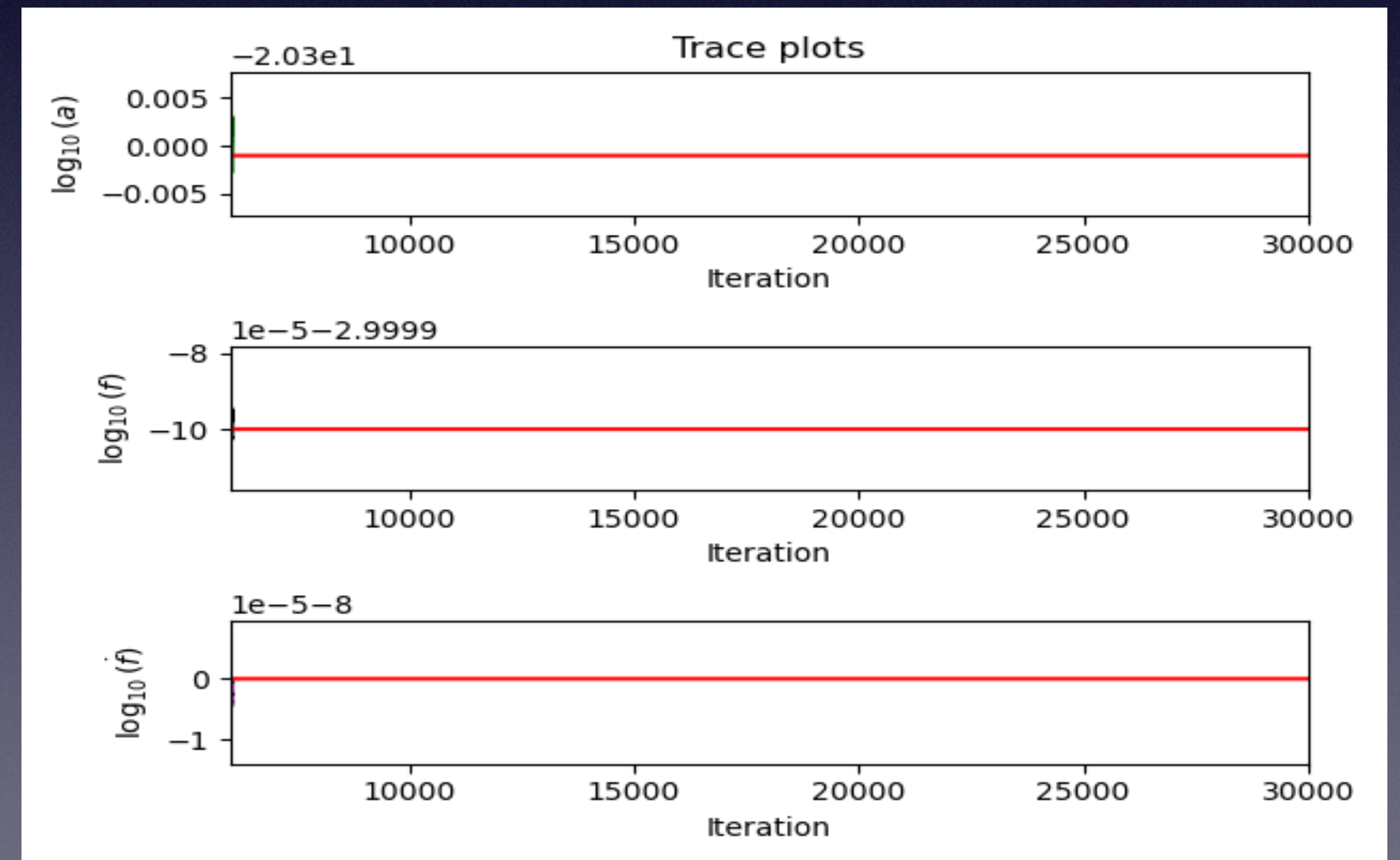
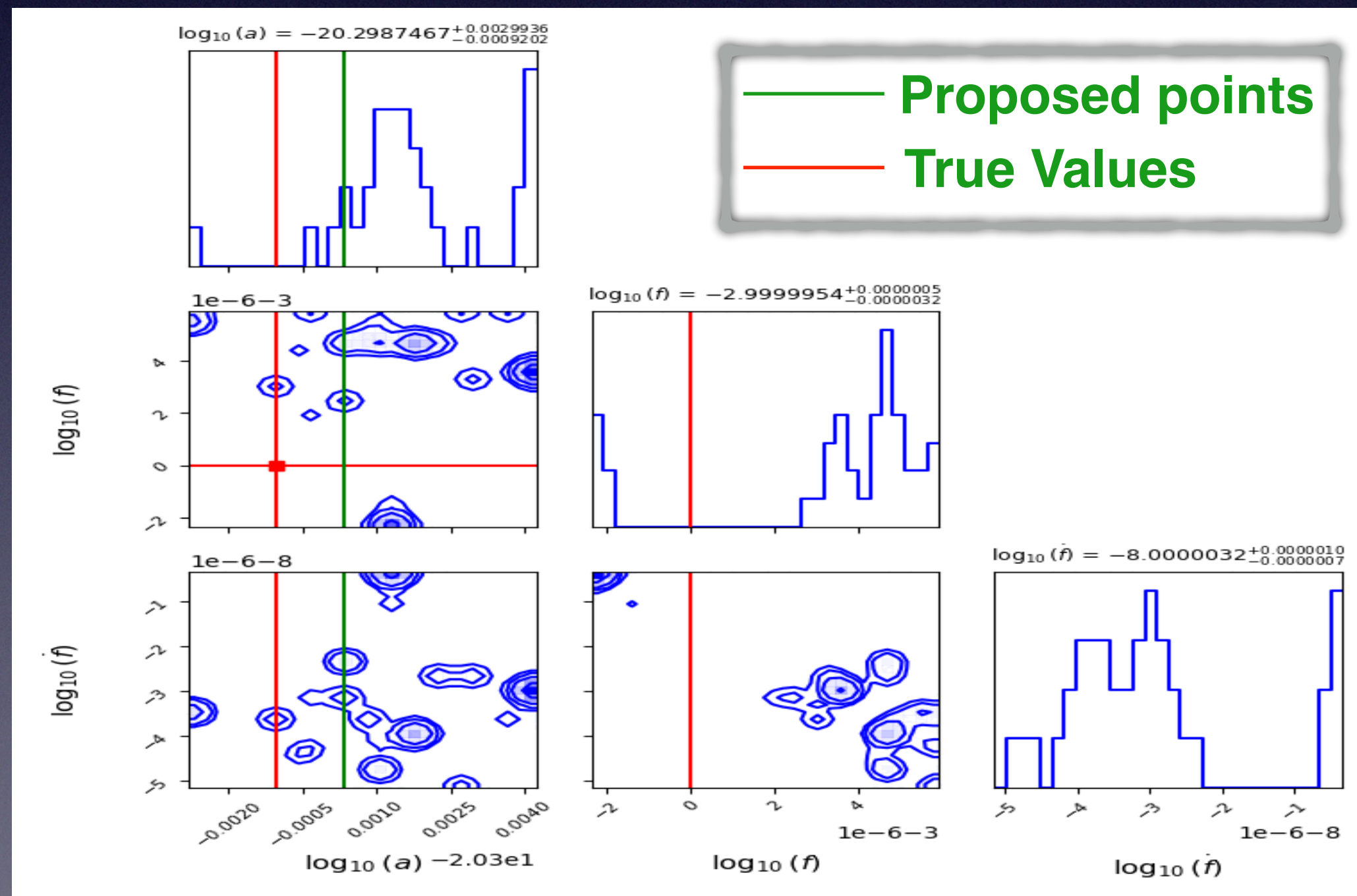
Monte Carlo: finding the “best” signal

Goal: Identify parameters θ that best match the waveform



Parameter Estimation

- Identified “best” location in parameter space
- Now sample $\theta \sim p(\theta | d)$ and explore posterior!
- Make statements on θ given observed data d



Part 2: **Search and Destroy**

Infrastructure

- **FastEMRIWaveforms** — GPU accelerated waveforms — **Fast** (Katz, et al 2021)
- **Circular Schwarzschild Waveform Models with 0PA + (full) 1PA.** (Pound, et al 2019)
- Full LISA response (TDI) including latest (mission requirements) noise model. (Katz, et al 2021)

Procedure: Injection with Accurate waveform then attempt recovery with innacurate waveform

1PA vs 1PA waveform

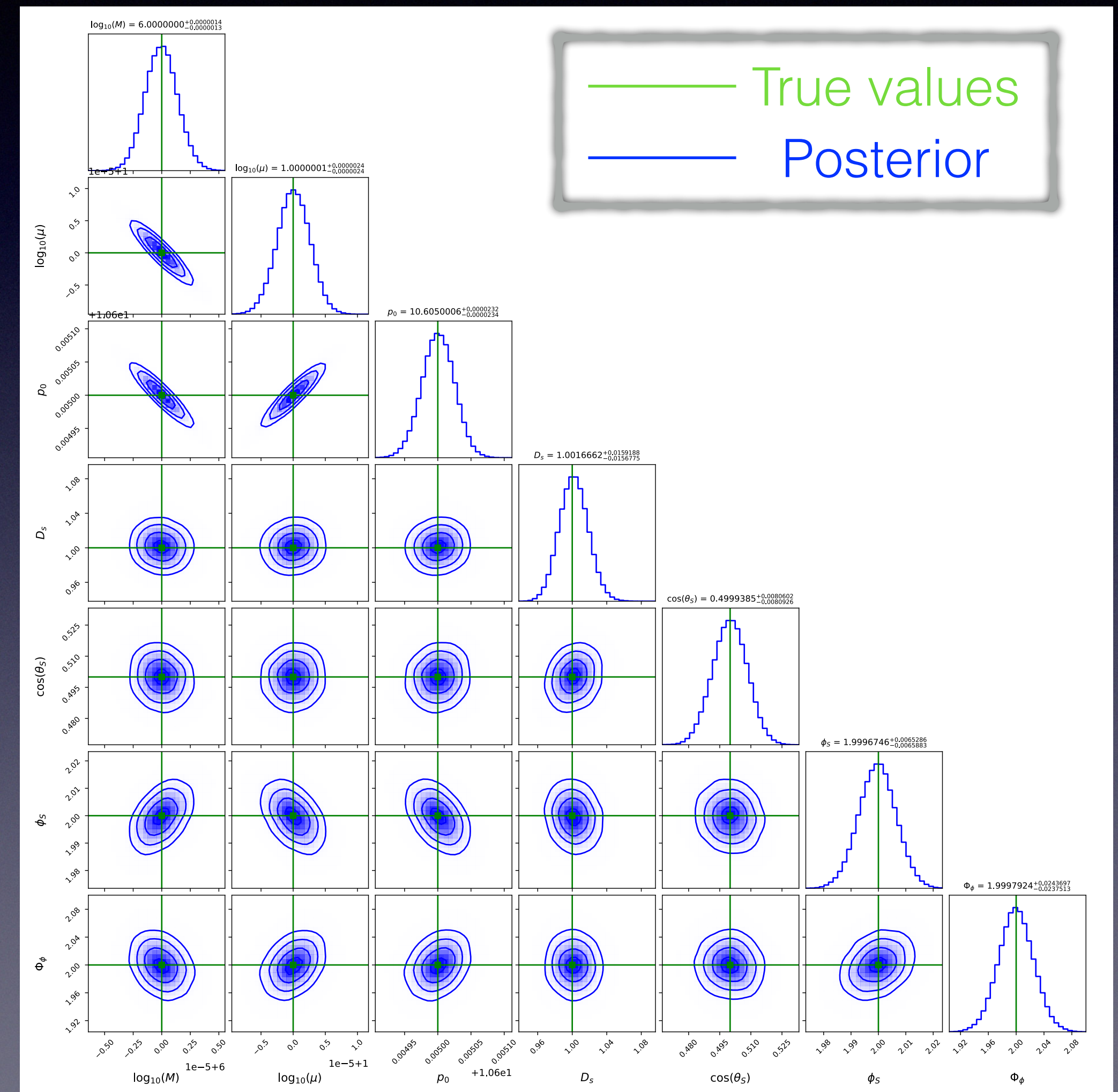
Inject waveform 1PA, **Recover with 1PA**

Quite loud: SNR ~ 62 ,

Strong field: $r - r_{\text{isco}} \sim 0.3M$

Parameters: $M = 10^6 M_{\odot}$, $\mu = 10 M_{\odot}$, $r_0 = 10.605M$

Unbiased. As expected.



0PA vs 1PA waveform

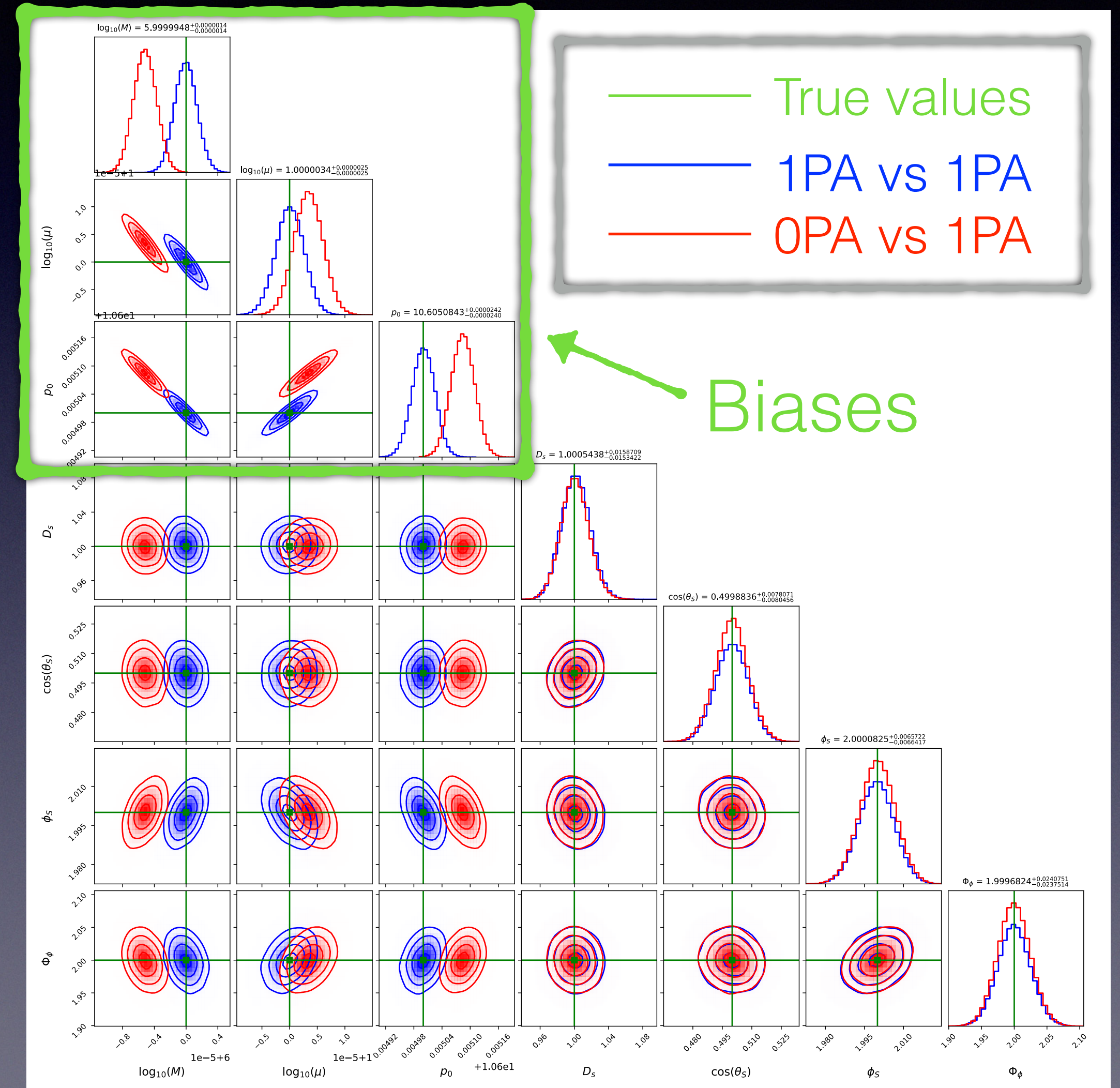
Inject waveform 1PA, **Recover with 0PA**

Quite loud: SNR ~ 62 ,

Strong field: $r - r_{\text{isco}} \sim 0.3M$

Parameters: $M = 10^6 M_{\odot}$, $\mu = 10 M_{\odot}$, $r_0 = 10.605M$

Largest bias $\sim 4\sigma$ from truth



Teeny-tiny perturbations to parameters

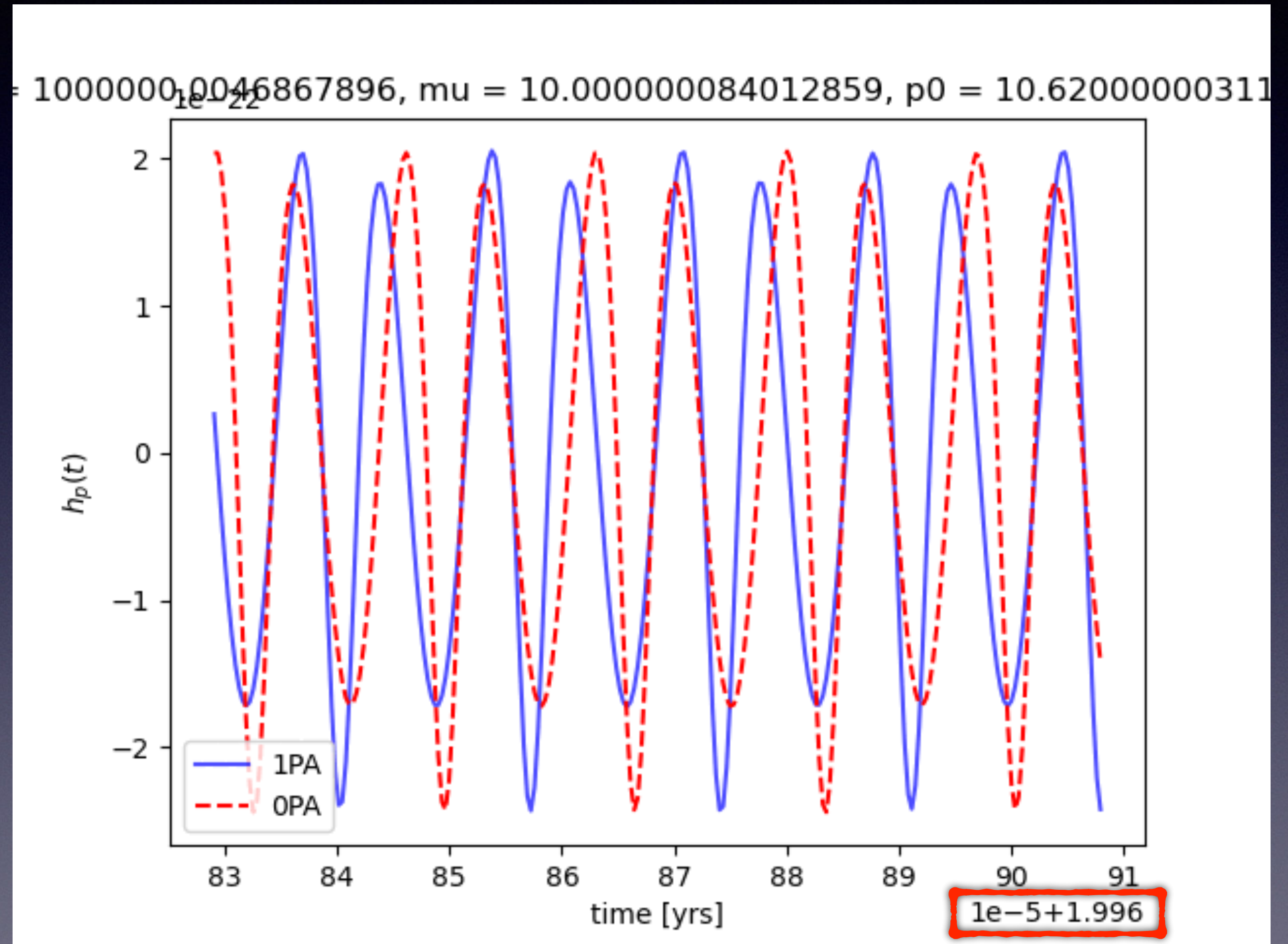
$$\mathcal{O} \left(h^{1PA}(\theta_{\text{true}}), h^{OPA}(\theta_{\text{true}}) \right) = 0.203$$

$$\Delta\phi = \max \left| \phi_{\text{true}}^{(1PA)} - \phi_{\text{true}}^{(OPA)} \right| \approx 3 \text{ rads}$$

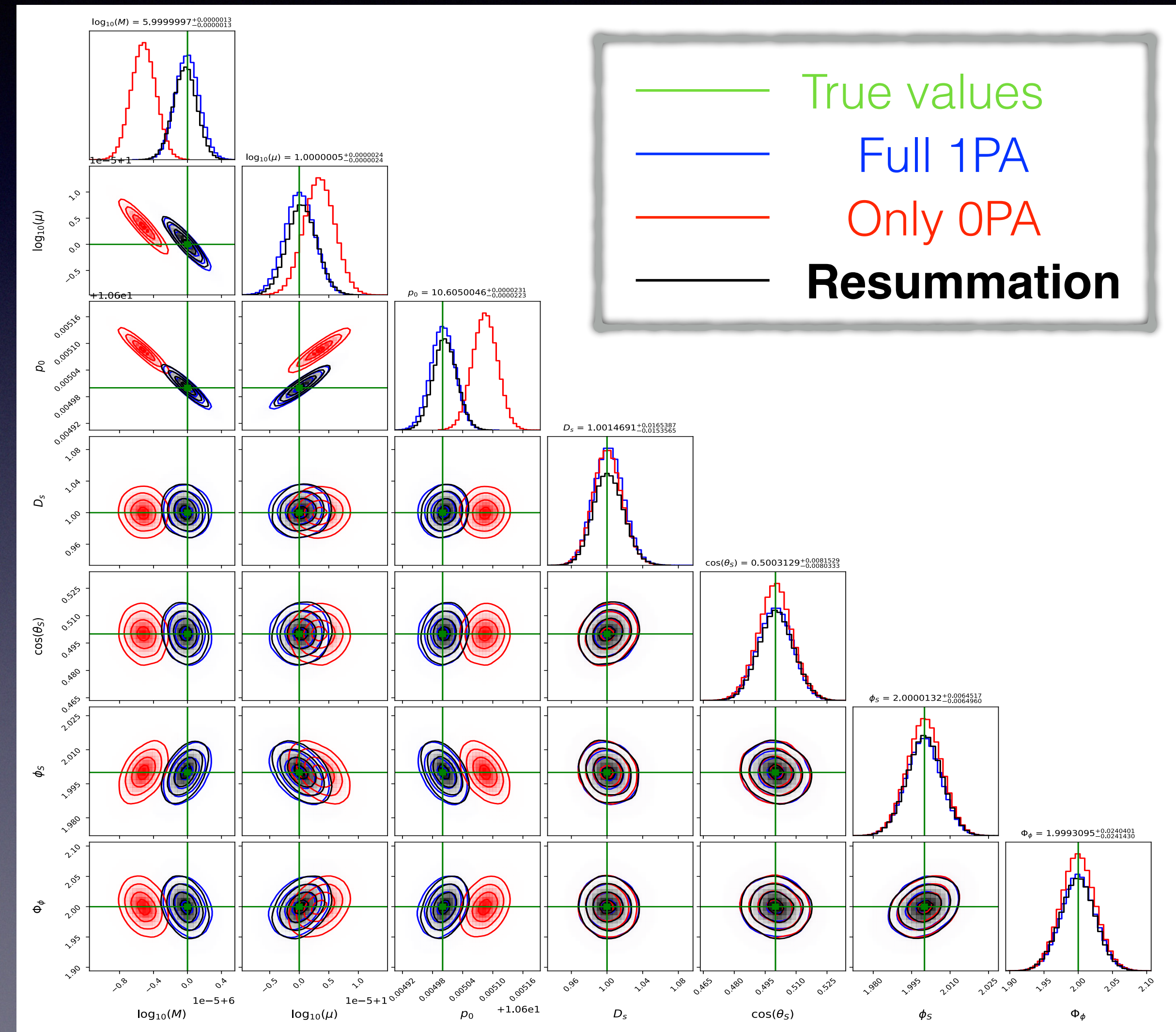
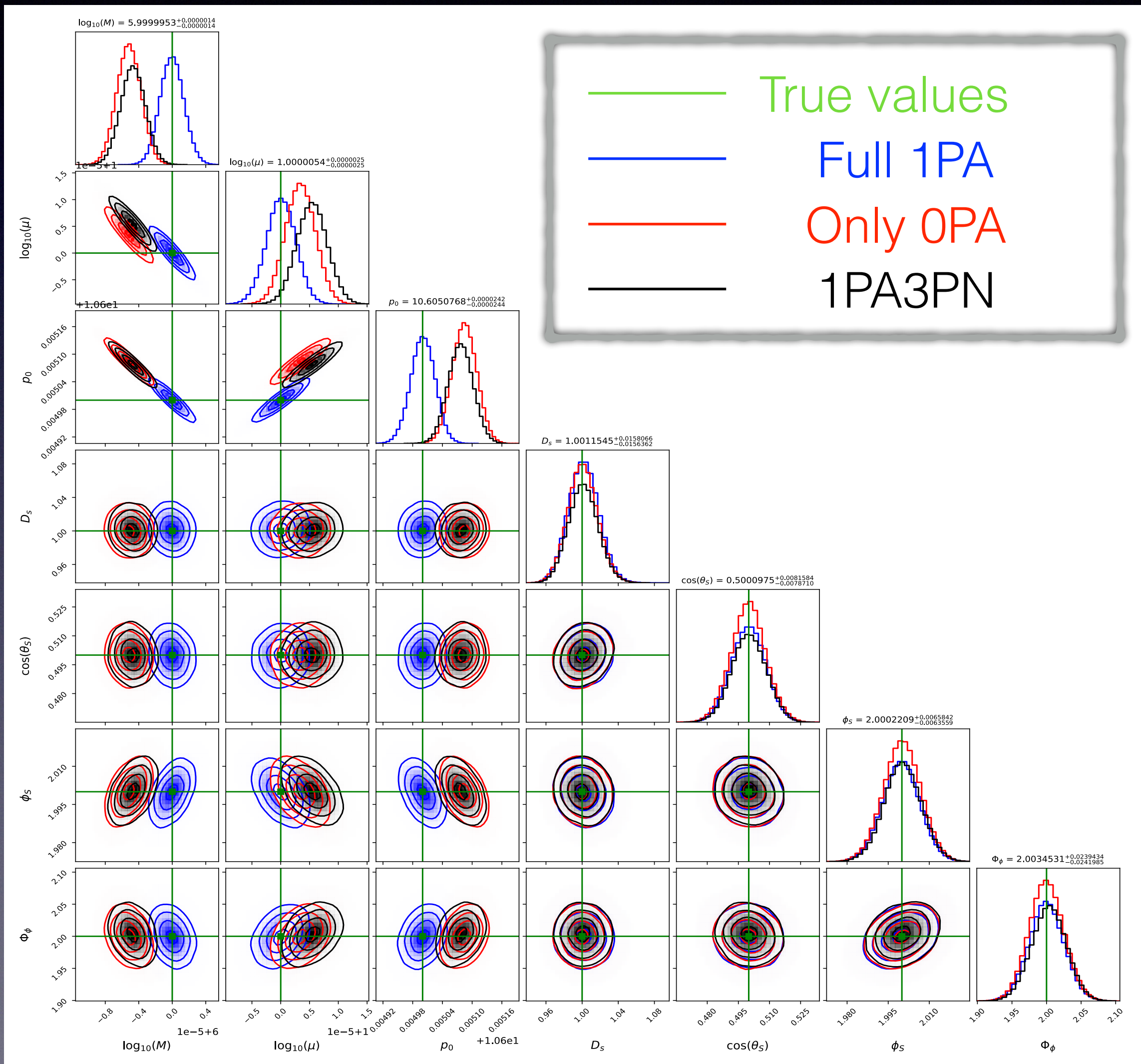
EMRIs are extremely sensitive to minor perturbations in parameters

$$\mathcal{O} \left(h^{1PA}(\theta_{\text{true}}), h^{OPA}(\theta_{\text{bf}}) \right) = 0.9999\dots$$

$$\Delta\phi = \max \left| \phi_{\text{true}}^{(1PA)} - \phi_{\text{bf}}^{(OPA)} \right| \approx 0.003 \text{ rads}$$



Can we do better with PN?

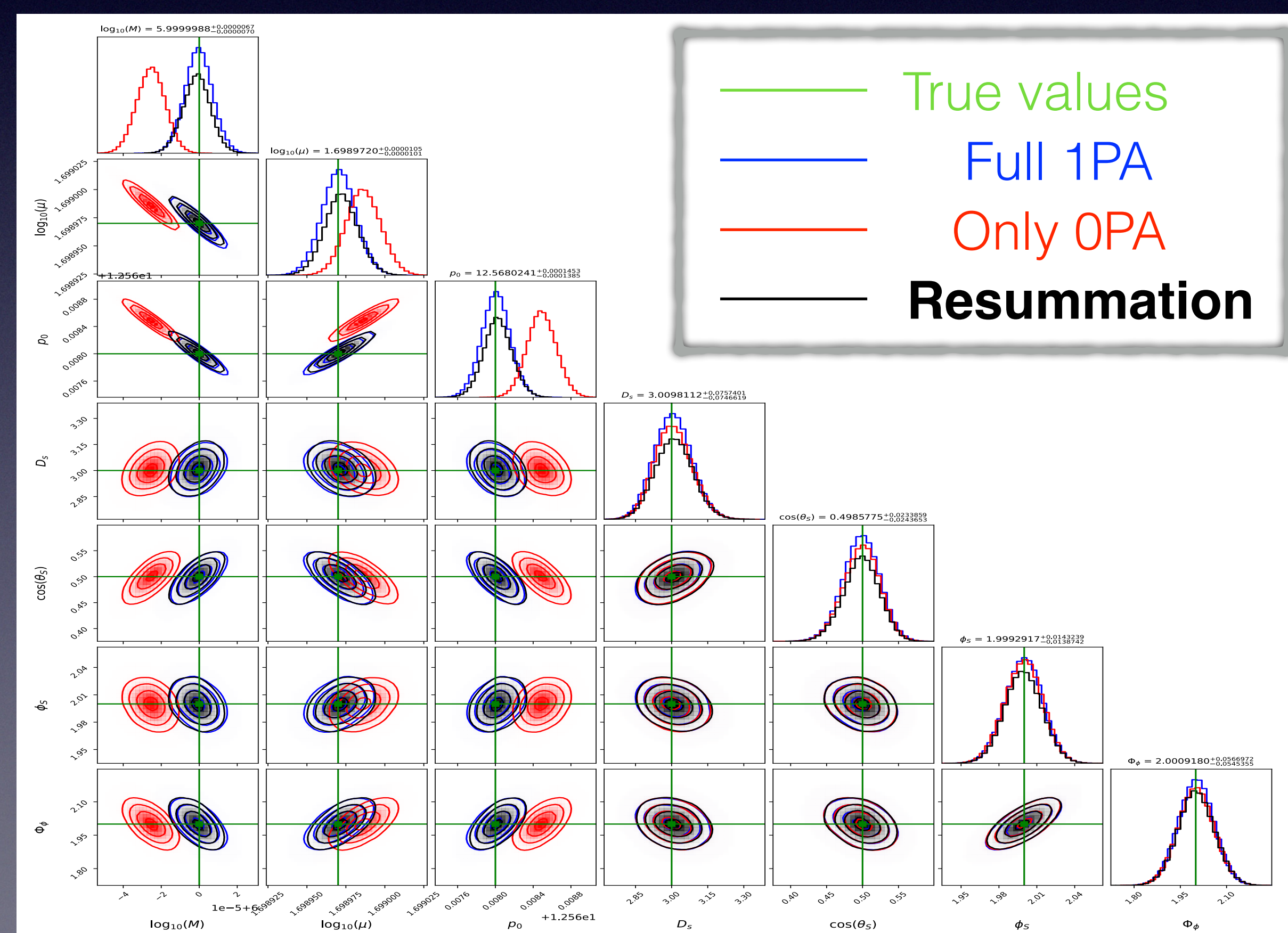
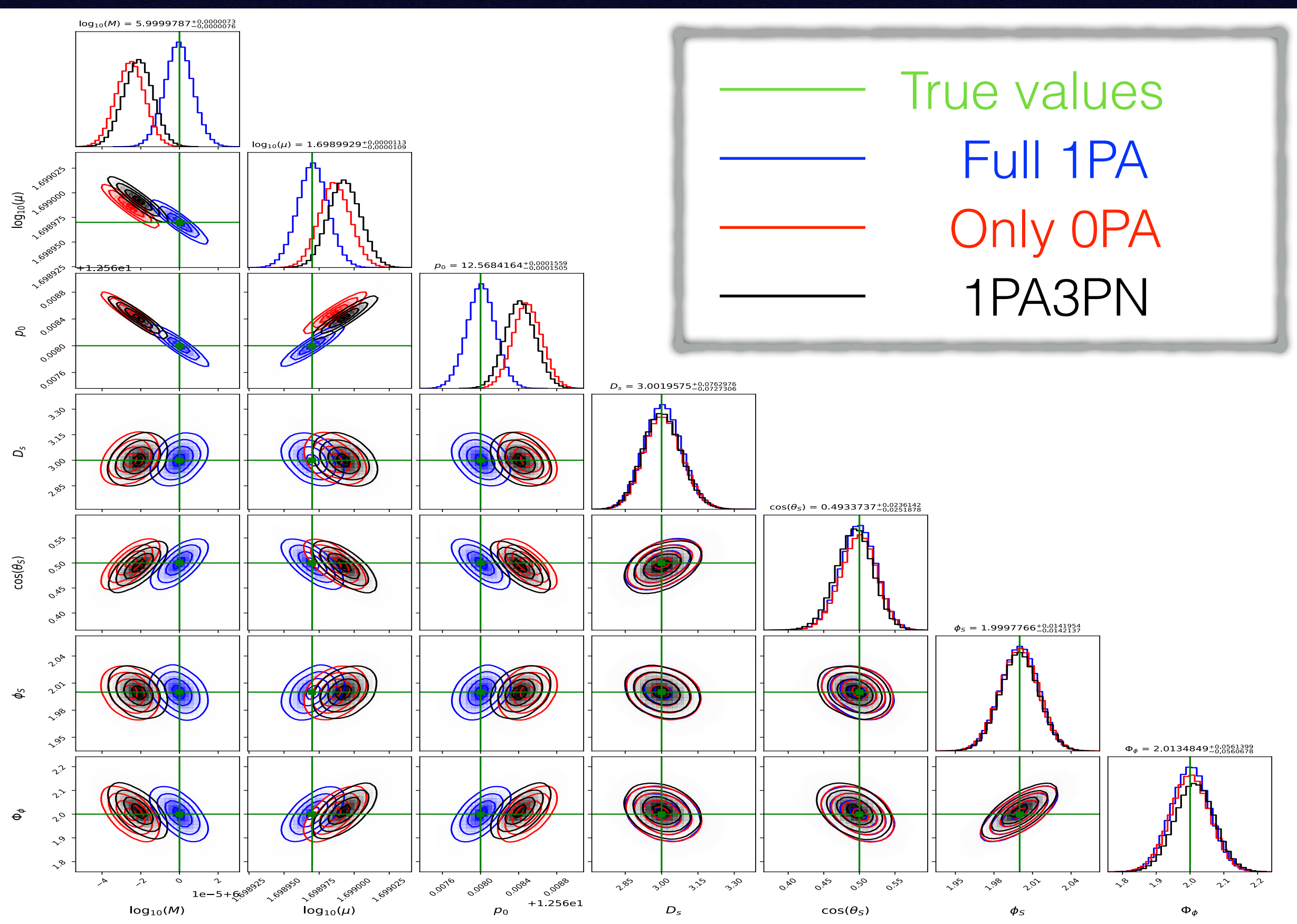


Increase mass-ratio?

Parameters: $M = 10^6 M_\odot$, $\mu = 50 M_\odot$, $r_0 = 12.568M$, SNR ~ 48

One year long inspiral

Strong field: $r - r_{\text{isco}} \sim 0.2M$

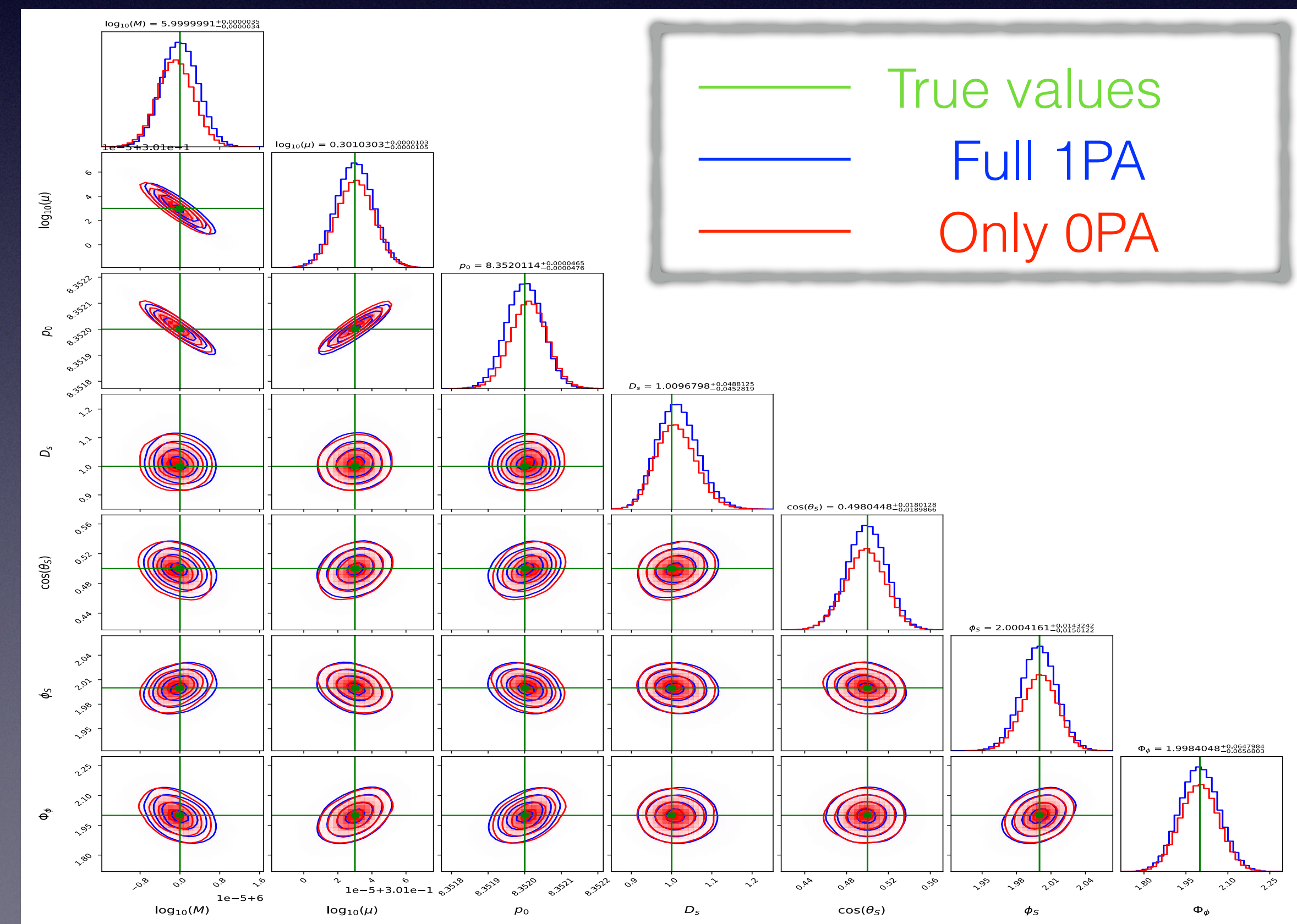
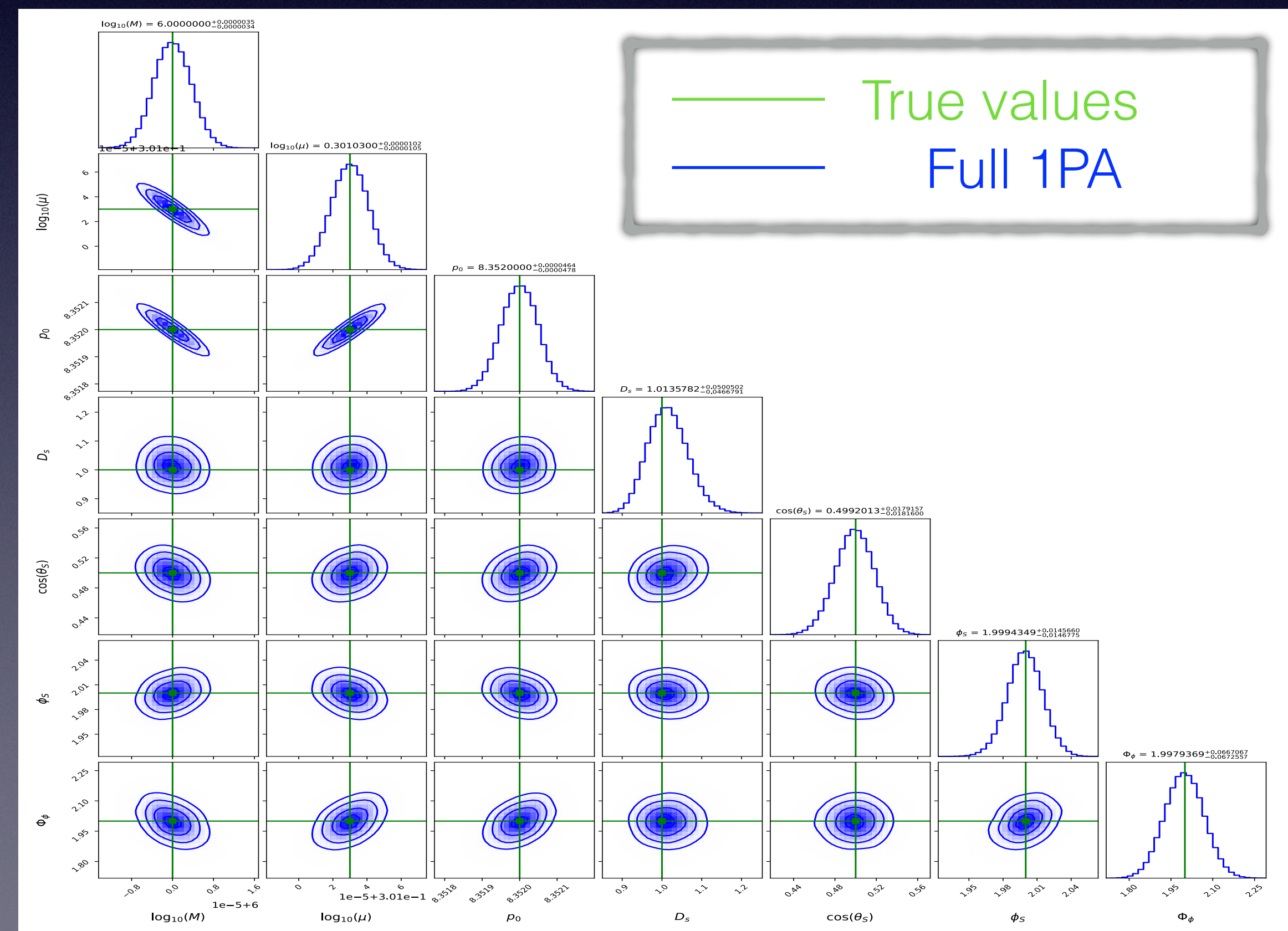


Decrease mass-ratio?

Parameters: $M = 10^6 M_\odot$, $\mu = 2M_\odot$, $r_0 = 8.32M$, SNR ~ 21

Two year long inspiral

Strong field: $r - r_{\text{isco}} \sim 0.22M$

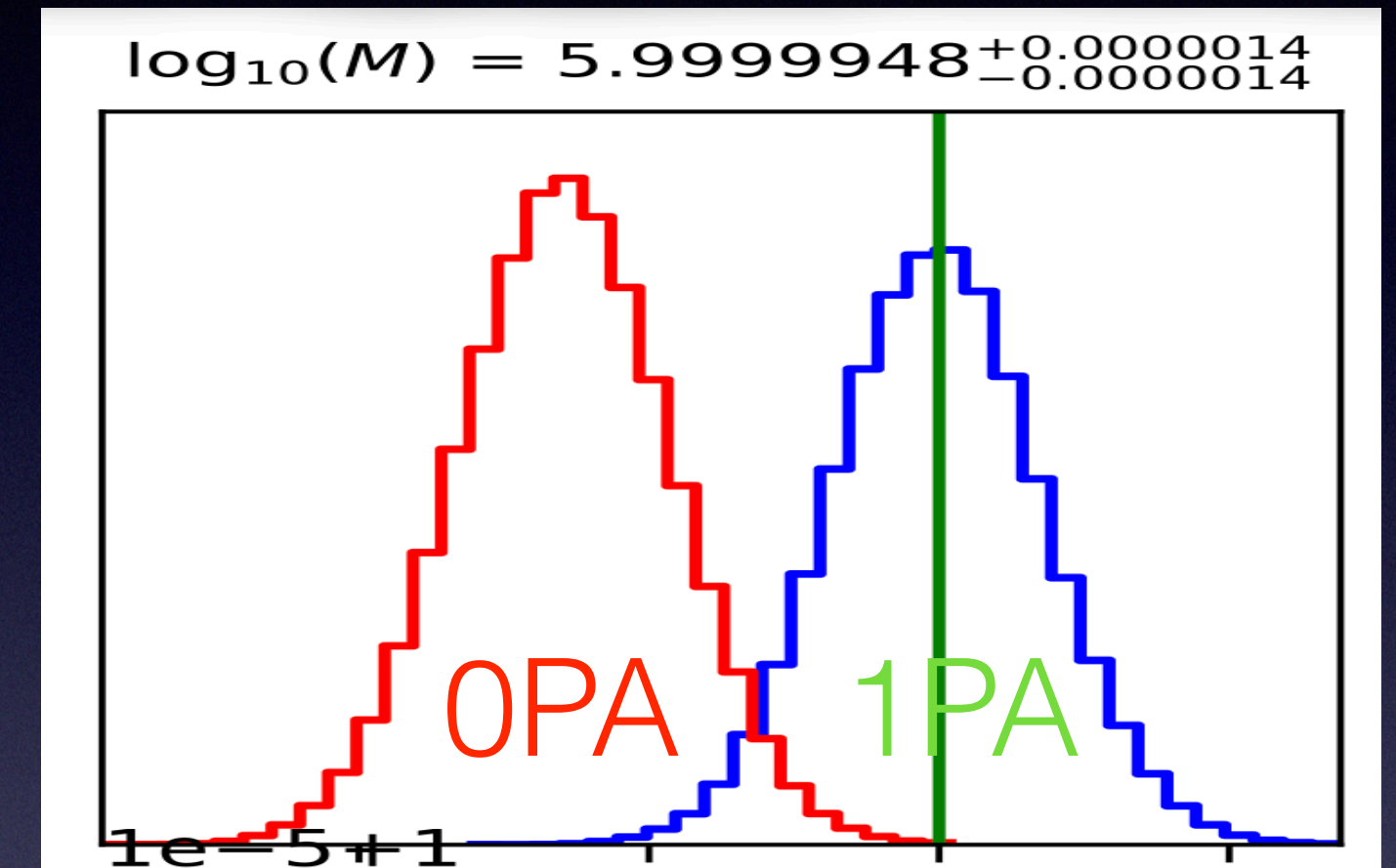


Final thoughts + musings

Neglecting post-adiabatic \implies Biases!

Biases “not as bad” as expected.

PN “tricks” could help!

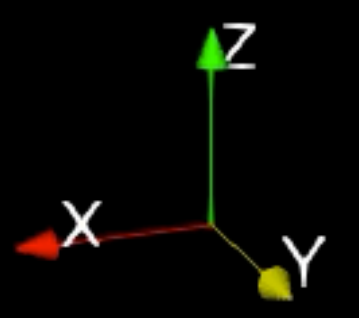
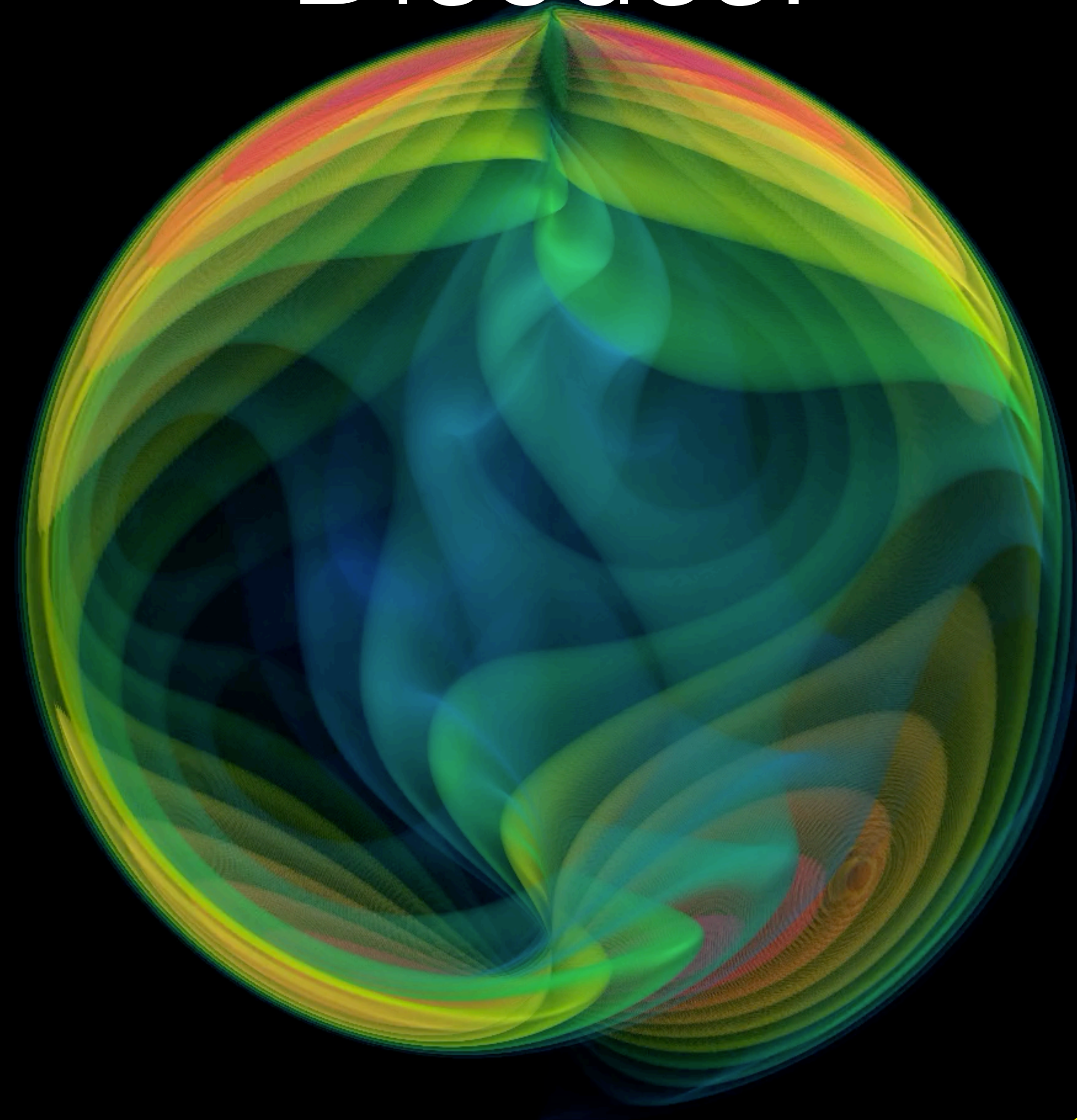


- Impact on **tests of GR + astrophysics?**
- Work only possible through **Self-force community** + **FEW**

Results may not extend for generic orbits / Kerr

Time: 0s

Discuss.



(Credit: Nils Fischer)

Why?

$$\mathcal{O} \left(h^{1PA}(\theta_{\text{true}}), h^{0PA}(\theta_{\text{true}}) \right) = 0.203$$

$$\Delta\phi = \max | \phi^{(1PA)} - \phi^{(0PA)} | \approx 3 \text{ rads}$$

$$\text{Lindblom} = \left(h^{(1PA)} - h^{(0PA)} \mid h^{(1PA)} - h^{(0PA)} \right) \approx 3900$$

$$\mathcal{O} \left(h^{1PA}(\theta_{\text{true}}), h^{0PA}(\theta_{\text{recovered}}) \right) = 0.999990043$$

$$\Delta\phi = \max | \phi^{(1PA)} - \phi^{(0PA)} | \approx 3 \cdot 10^{-3} \text{ rads}$$

$$\text{Lindblom} = \left(h^{(1PA)} - h^{(0PA)} \mid h^{(1PA)} - h^{(0PA)} \right) \approx 0.17$$

