

Strong field amplitudes and classical physics

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- ① **Eikonal amplitudes from curved backgrounds**
Tim Adamo, A.C. and Piotr Tourkine
SciPost Phys. 13 (2022)
- ② **Classical physics from amplitudes on curved backgrounds**
Tim Adamo, A.C. and Anton Ilderton
JHEP 08 (2022) 281
- ③ **All orders waveforms from amplitudes**
Tim Adamo, A.C., Anton Ilderton and Sonja Klisch
arXiv:2210.04696
- ④ **Large gauge transformations and the structure of amplitudes**
A.C., Asaad Elkhidir, Anton Ilderton and Donal O'Connell
arXiv:2211.16438

Motivation

- The post-Minkowskian approximation in general relativity
- On-shell data as natural building blocks (KMOC)

Classical observables on curved background

- The post-background approximation
- Strong field amplitudes as natural building blocks

Main results

- Recovering memory effects neglected perturbatively
- Relation between 3-points and large gauge transformations
- Self-force results on plane wave backgrounds

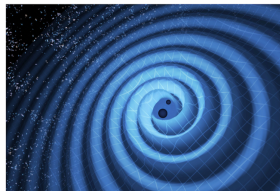
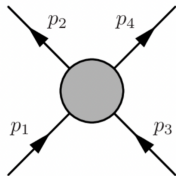
The two-body problem in GR

- Gravitational waves carry fingerprints of a **two-body dynamics**

$$\mathcal{R}_{\mu\nu} - \frac{1}{2}g_{\mu\nu}\mathcal{R} = \frac{8\pi G}{c^4}T_{\mu\nu} \quad , \quad \ddot{x}_a^\mu = -\Gamma_{\alpha\beta}^\mu \dot{x}_a^\alpha \dot{x}_a^\beta$$

... however, **no exact solution** is known!

- The **post-Minkowskian approximation** (PM) has gained a renewed attention after a remarkable state of the art calculation from scattering amplitudes (**Zvi Bern et al.**)



Credit: Tim Pyle

The post-Minkowskian approximation

- The change in momentum due to a **scattering** is

$$\Delta p_1^\mu = -\frac{1}{2} \int_{-\infty}^{+\infty} d\sigma \partial^\mu g_{\alpha\beta}(x_1(\sigma)) p_1^\alpha(\sigma) p_1^\beta(\sigma)$$

Expanding around straight trajectories in the **weak field** limit

$$x_a^\mu(\sigma) = x_{a,0}^\mu + \sigma p_a^\mu + \dots \quad ; \quad h^{\mu\nu}(x(\sigma)) = -16\pi G \mathcal{P}^{\mu\nu\alpha\beta} T_{\alpha\beta} + \dots$$

Classical result at 1PM $\sim G$ (Damour)

The Fourier domain computation contains a **scattering amplitude**

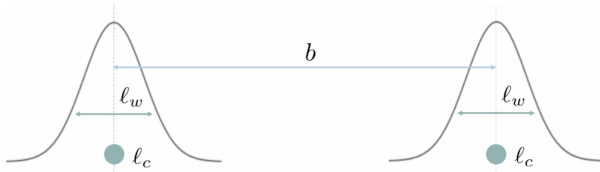
$$\Delta p_1^\mu = \int_q e^{iq \cdot b} \hat{\delta}(q \cdot p_1) \hat{\delta}(q \cdot p_2) i q^\mu \underbrace{8\pi G p_1^\alpha p_1^\beta \frac{\mathcal{P}_{\alpha\beta;\alpha'\beta'}}{q^2} p_2^{\alpha'} p_2^{\beta'}}_{\mathcal{A}_4^{\text{tree}}}$$

The KMOC formalism

- **Binary system** as superposition of single particle states

$$|\psi\rangle = \int d\Phi(p_1) d\Phi(p_2) \phi_1(p_1) \phi_2(p_2) e^{\frac{ib \cdot p_1}{\hbar}} |p_1 p_2\rangle$$

- **Classical limit** \leftrightarrow Goldilocks relations $l_c \ll l_w \ll l_s$



Credit: Ben Maybee, 2105.10268

Main idea (Kosower, Maybee, O'Connell)

Classical observables from **on-shell amplitudes** to all PM orders

$$\langle \psi | S^\dagger \mathbb{P}_1^\mu S | \psi \rangle = p_1^\mu + \int_q e^{iq \cdot b} \hat{\delta}(q \cdot p_1) \hat{\delta}(q \cdot p_2) i q^\mu \mathcal{A}_4^{\text{tree}} + \dots$$

The post-background approximation

- We have defined a classical observable around [Minkowski](#)

$$\Delta p^\mu = \int_{-\infty}^{+\infty} d\sigma \partial^\mu g_{\alpha\beta}(x(\sigma)) p^\alpha(\sigma) p^\beta(\sigma) \quad , \quad g_{\alpha\beta} = \eta_{\alpha\beta} + h_{\alpha\beta}$$

...however, we could have chosen any [curved spacetime](#)

$$g_{\alpha\beta} = g_{\alpha\beta}^0 + h_{\alpha\beta}$$

where $g_{\alpha\beta}^0$ is an exact solution to Einstein field equations

Question

Can we recover these expansions around g^0 from KMOC?

QFT on curved backgrounds

- Consider QFT in presence of a non trivial background, where the \mathcal{S} -matrix carries information on the background g^0

$$\mathcal{S}|\psi\rangle = \int d\Phi(p', p)\phi(p) \underbrace{\langle p' | \mathcal{S} | p \rangle}_{2\text{-point}} |p'\rangle + \dots$$

We call the building blocks "strong field amplitudes"

$$\langle p' | \mathcal{S} | p \rangle \quad , \quad \langle p' | \mathcal{S} | p, k^\eta \rangle \quad , \quad \langle p' | \mathcal{S} | p, k_1^{\eta_1}, k_2^{\eta_2} \rangle \quad \dots$$

KMOC on curved backgrounds

Observables on g^0 can be computed from strong field amplitudes

$$\langle \Delta \mathcal{O} \rangle = \lim_{\hbar \rightarrow 0} \left[\langle \psi | \mathcal{S}^\dagger \hat{\mathcal{O}} \mathcal{S} | \psi \rangle - \langle \psi | \hat{\mathcal{O}} | \psi \rangle \right]$$

Strong field amplitudes

- The simplest **strong field amplitude** is a scalar 2-point given by the quadratic part of the action $S[\Phi]$ on $\Phi = \epsilon_1 \Phi_{in} + \epsilon_2 \Phi_{out}$

$$S[\Phi] = \int d^4x \sqrt{-g} \left(g^{\mu\nu}(x) \partial_\mu \Phi(x) \partial_\nu \Phi^*(x) + m^2 |\Phi(x)|^2 \right)$$

$$\langle p' | \mathcal{S} | p \rangle := \left. \frac{\partial^2 S[\Phi]}{\partial \epsilon_1 \partial \epsilon_2} \right|_{\epsilon_1 = \epsilon_2 = 0}$$

- N -points are defined by the **multilinear part of the action**.
Hard to compute as they resum infinite amplitudes on η

$$\Rightarrow = \text{[wavy red line]} + \text{[red line with loop]} + \dots$$

2-points on stationary backgrounds

- Consider the KG equation on a stationary background

$$(\square + m^2)\Phi(x) = h^{\mu\nu}(x)\partial_\mu\partial_\nu\Phi(x) + \dots$$

We can apply a WKB approximation (Kol, O'Connell, Telem)

$$\Phi_{in/out} \rightarrow \chi(x_\perp) := M \int_q \hat{\delta}(P \cdot q) \hat{\delta}(p \cdot q) e^{-iq_\perp \cdot x_\perp} \tilde{h}^{\mu\nu}(q) p_\mu p_\nu$$

Relation with traditional amplitudes (Adamo, C., Tourkine)

2-points on g^0 as **eikonal amplitudes**. The quadratic part of the action resums an infinite number of amplitudes in Minkowski

$$\langle p' | \mathcal{S} | p \rangle = N \hat{\delta}(p'_0 - p_0) \int_{x_\perp} e^{-iq_\perp \cdot x_\perp} \left(e^{i\chi(x_\perp)/\hbar} - 1 \right)$$

Example: 2-point on Kerr

- $\langle p' | \mathcal{S} | p \rangle$ on **Kerr** depends on the following eikonal amplitude

$$I_a(q_\perp) = \int d^2 x_\perp e^{-i q_\perp \cdot x_\perp / \hbar} |x_\perp - a_\perp|^\alpha |x_\perp + a_\perp|^{2\beta}$$

- **Analytic continuation** provides a KLT-like factorization

$$I_a(q_\perp) = - \left(\tilde{l}_1, \tilde{l}_2 \right) \mathcal{K}(l_1, l_2)^T$$

where

$$\mathcal{K} := \frac{i}{2} \begin{pmatrix} 1 - e^{2i\pi\beta} & -e^{i\pi\alpha} (-1 + e^{2i\pi\beta}) \\ -e^{i\pi\alpha} (-1 + e^{2i\pi\beta}) & 1 - e^{2i\pi(\alpha+\beta)} \end{pmatrix}$$

- **Complex poles** at $i\alpha_\pm(s) = n$, $n \in \mathbb{N}$ (**Adamo, C., Tourkine**)

2-point on plane wave backgrounds

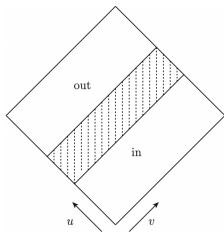
- We can also consider "strong field amplitudes" on non stationary backgrounds like **gravitational plane waves**

$$ds^2 = 2 dudv - H_{ab}(u) x^a x^b (du)^2 - dx^\perp dx^\perp$$

- A scalar 2-point is given by

$$\langle p' | \mathcal{S} | p \rangle = \frac{4\pi \hat{\delta}(p'_+ - p_+)}{\sqrt{|\det(\mathbf{c})| \hbar}} e^{-\frac{i}{2p_+ \hbar} \mathbf{q}_\perp \cdot \mathbf{c}^{-1} \cdot \mathbf{q}_\perp}$$

where \mathbf{c} is a 2×2 matrix encoding classical **memory effects**



Observables from 2-points

- 2-points can be used to construct a semiclassical final state

$$\mathcal{S}|\psi\rangle = \int d\Phi(p', p) \phi_b(p) \langle p' | \mathcal{S} | p \rangle | p' \rangle$$

- For **Schwarzschild** and **Kerr**, stationary phase arguments gives

$$\mathcal{S}|\psi\rangle = \int d^4 p \phi_b(p - \partial_b \chi(b)) | p \rangle \Rightarrow \Delta p^\mu = \partial_b^\mu \chi(b) + \dots$$

- For **plane waves**, memory effects appear (**Adamo, C., Ilderton**)

$$\Delta p^i = \partial_{x^-} E_a^i(x_f) z^a + \dots \quad , \quad E_a^i = b_a^i + \sqrt{G} c_a^i x^-$$

Comparison with on-shell amplitudes

- The 2-point on a plane wave background shows that the impulse has a linear term in $\kappa \sim \sqrt{G}$ (Adamo, C., Ilderton)

$$\Delta p^\mu = \sqrt{G} c_a^\mu z^a + \dots$$

- From a perturbative approach, the leading term should be a 4-point Compton amplitude... but this scale as $\kappa^2 \sim G$

$$\begin{aligned} \Delta p^\mu &= \int_q d\Phi(k) \hat{\delta}(2q \cdot p_1) \hat{\delta}^+(2q \cdot k - q^2) \\ &\times \alpha(k - q) \alpha(k) e^{-iz \cdot q} i q^\mu \mathcal{A}_4^{\text{tree}} \sim G \end{aligned}$$

Solution (C., Ilderton, Elkhidir, O'Connell)

3-point amplitudes on Minkowski are actually non vanishing when large gauge transformations are included in the LSZ reduction

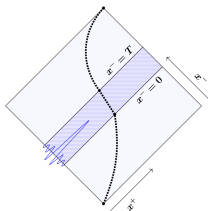
Beyond geodesics motion

- Consider a **3-point amplitude** for a scalar particle emitting a graviton on a plane wave background

$$\langle p', k^\eta | \mathcal{S} | p \rangle \sim \frac{2i\kappa}{\hbar^{3/2}} \int_x \frac{e^{i\mathcal{V}(x)}}{\sqrt{|E(x)|}} \mathcal{E}_{\mu\nu}^\eta(k; x) P^\mu(x) P'^\nu(x)$$

$$\mathcal{V}(x) := \int_y \frac{\theta(x-y) P(y) \cdot \bar{K}(y)}{p_+ - k_+}$$

- If we use this strong field amplitude with KMOC we obtain observables containing a series of **all order PM contributions**.



- **Radiation** emitted on \mathcal{I}^+ is controlled by

$$\mathbb{O}_{\vec{\mu}}(u, r, \hat{x}) = - \frac{i\hbar^2}{4\pi r} \int_0^\infty \hat{d}\omega e^{-i\omega u} \underbrace{C_{\vec{\mu}}^\eta(k)}_{\text{helicity}} a_\eta(k) \Big|_{k=\hbar\omega\hat{x}} + \text{c.c.}$$

- The **waveform** $W_{\vec{\mu}}$ is the leading coefficient in $1/r$

$$W_{\mu\nu\sigma\rho}(u, \hat{x}) = \frac{i\kappa}{2\pi\hbar^{\frac{1}{2}}} \int_0^\infty \hat{d}\omega e^{-i\omega u} k_{[\mu}\varepsilon_{\nu]}^{-\eta} k_{[\sigma}\varepsilon_{\rho]}^{-\eta}$$

$$\times \int d\Phi(p') \underbrace{\langle \psi | \mathcal{S}^\dagger | p' \rangle \langle p', k^\eta | \mathcal{S} | \psi \rangle}_{LO} \Big|_{k=\hbar\omega\hat{x}} + \text{c.c.}$$

Strong field waveform

General result - 1PB (post-background)

$$W_{\mu\nu\sigma\rho} = -\frac{\kappa^2}{\pi} \hat{x}_{[\mu} \hat{x}_{[\sigma} \int_y \delta(u - \bar{V}(y)) \left[\mathcal{D}^2 T_{\rho] \nu]}^0(\hat{x}, y) - \mathcal{D} T_{\rho] \nu]}^1(\hat{x}, y) \right]$$

$$T_{\nu\rho}^0(\hat{x}, y) := \frac{\mathbb{P}_{\nu\alpha}(\hat{x}, y) \mathbb{P}_{\rho\beta}(\hat{x}, y) P^\alpha(y) P^\beta(y) - \frac{1}{2} \eta_{\nu\rho} m^2}{\sqrt{|E(y)|}}$$

$$T_{\nu\rho}^1(\hat{x}, y) := \frac{\sigma_{\nu\rho}(y)}{\hat{x}_+ \sqrt{|E(y)|}} p_+^2$$

Impulsive wave for $\nu \sim \sqrt{G} \lambda |u|$ (Adamo, C., Ilderton, Klisch)

$$W_{\mu\nu\sigma\rho} = -\frac{\kappa^2 p_+}{\pi^2 \sqrt{8}} \delta_{[\mu}^+ \delta_{[\sigma}^+ (-1)^{(a)} \delta_{\rho]}^a \delta_{\nu]}^a \frac{\partial^2}{\partial u^2} \left(\frac{\nu \log(\nu + \sqrt{\nu^2 - 1})}{\sqrt{\nu^2 - 1}} \right)$$

Summary

- **Strong field amplitudes** are the natural building blocks to study perturbation theory around non trivial backgrounds
- **Observables** from the classical limit of strong field amplitudes

Main results

- Recovering **memory effects** which were neglected
- **3-point** amplitudes are non vanishing and related to memory
- **Self-force** results from strong field amplitudes

Main message

We can gain a deeper understanding of perturbation theory on a flat spacetime by studying amplitudes on strong backgrounds