

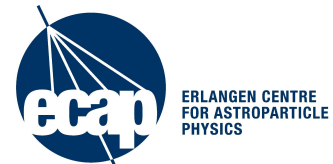
Anomalies in the Radio Neutrino Observatory Greenland

NBIA Neutrino Summer School

Zack Meyers

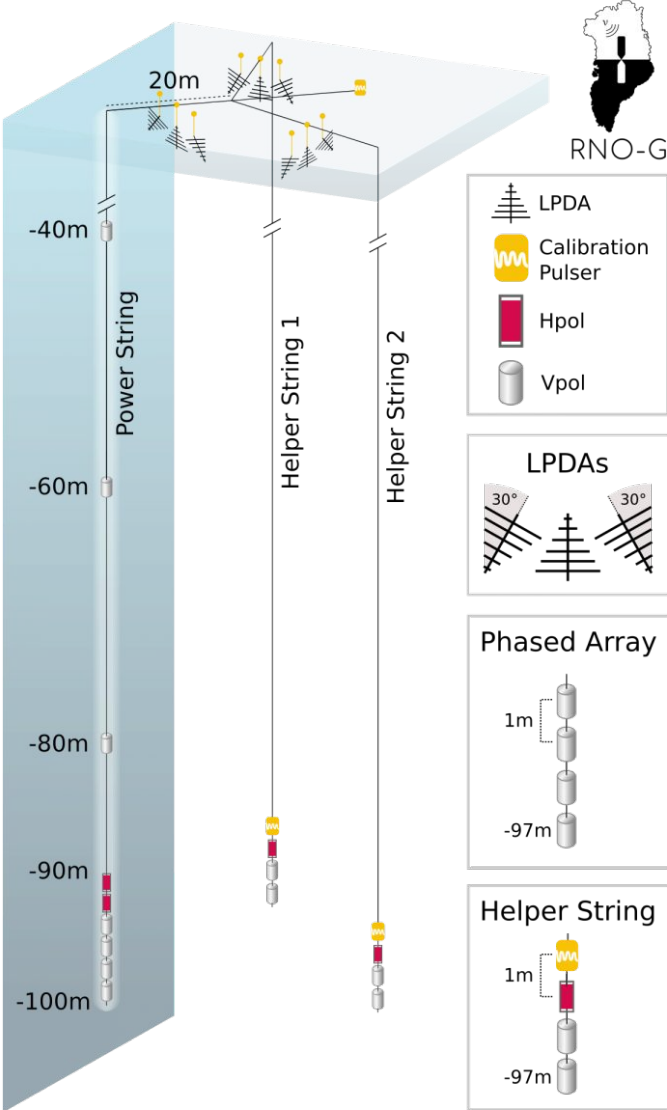
Copenhagen, July 20, 2023

HELMHOLTZ RESEARCH FOR
GRAND CHALLENGES

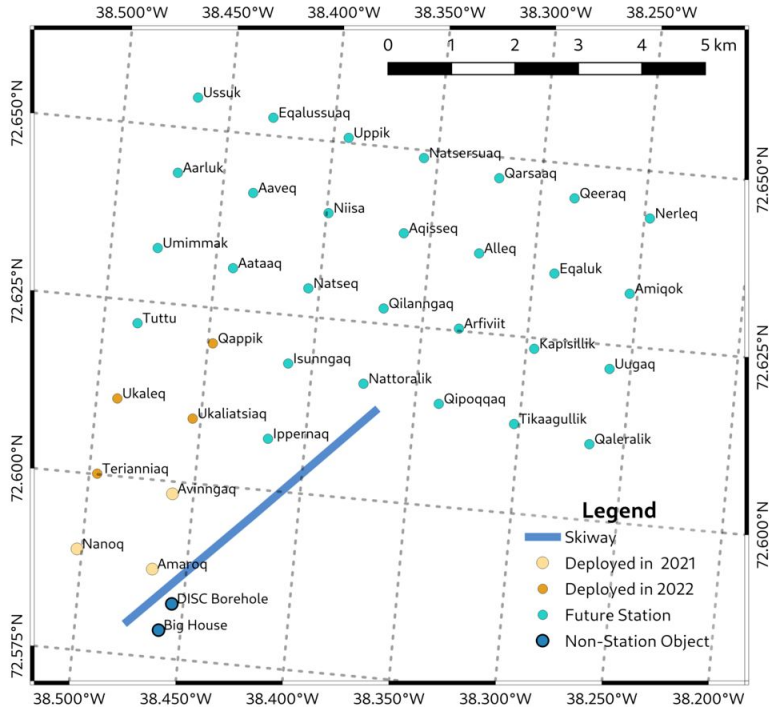


The Radio Neutrino Observatory Greenland

RNO-G

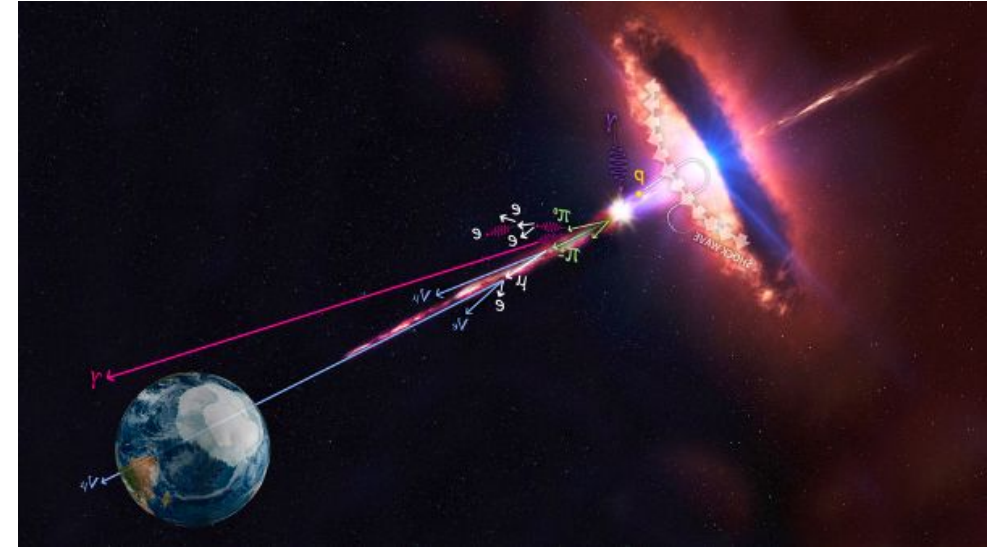


RNO-G Planned Layout



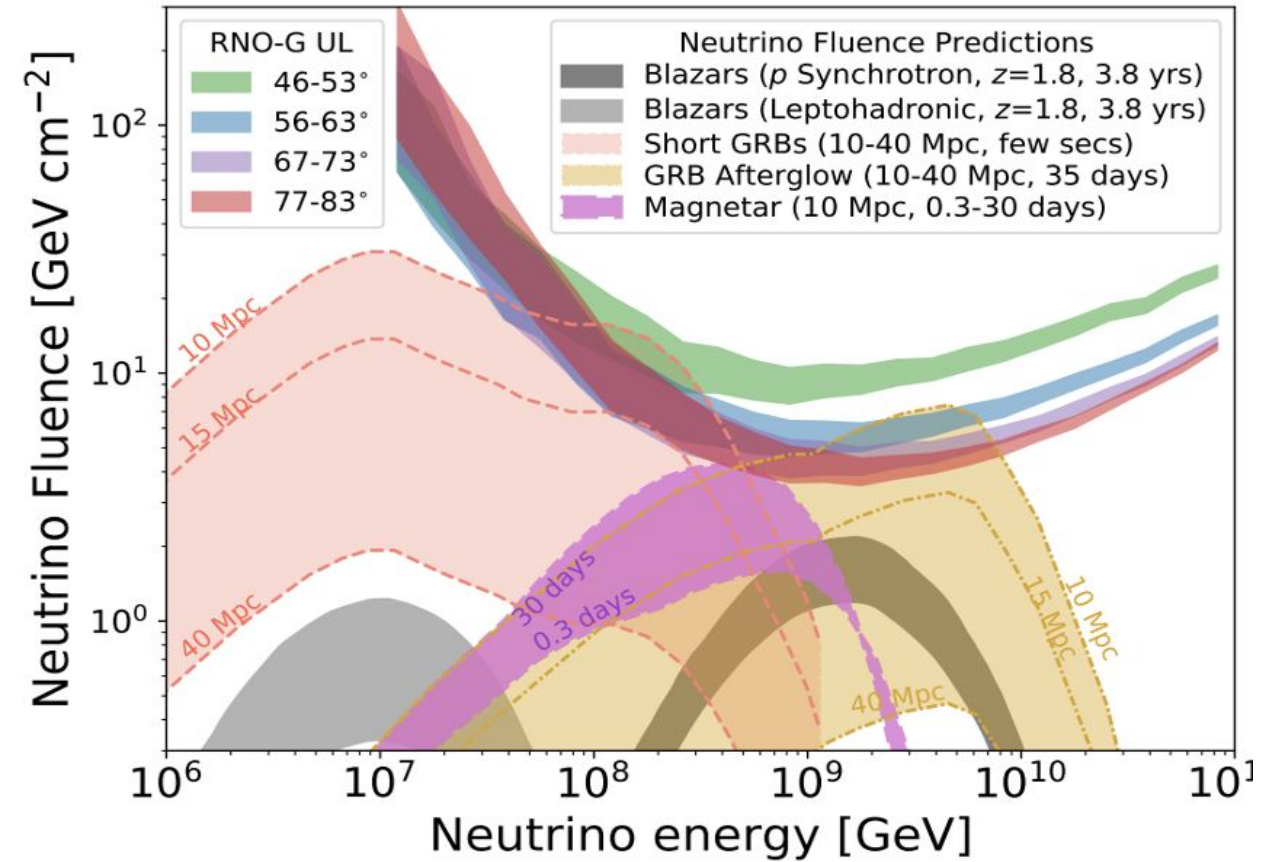
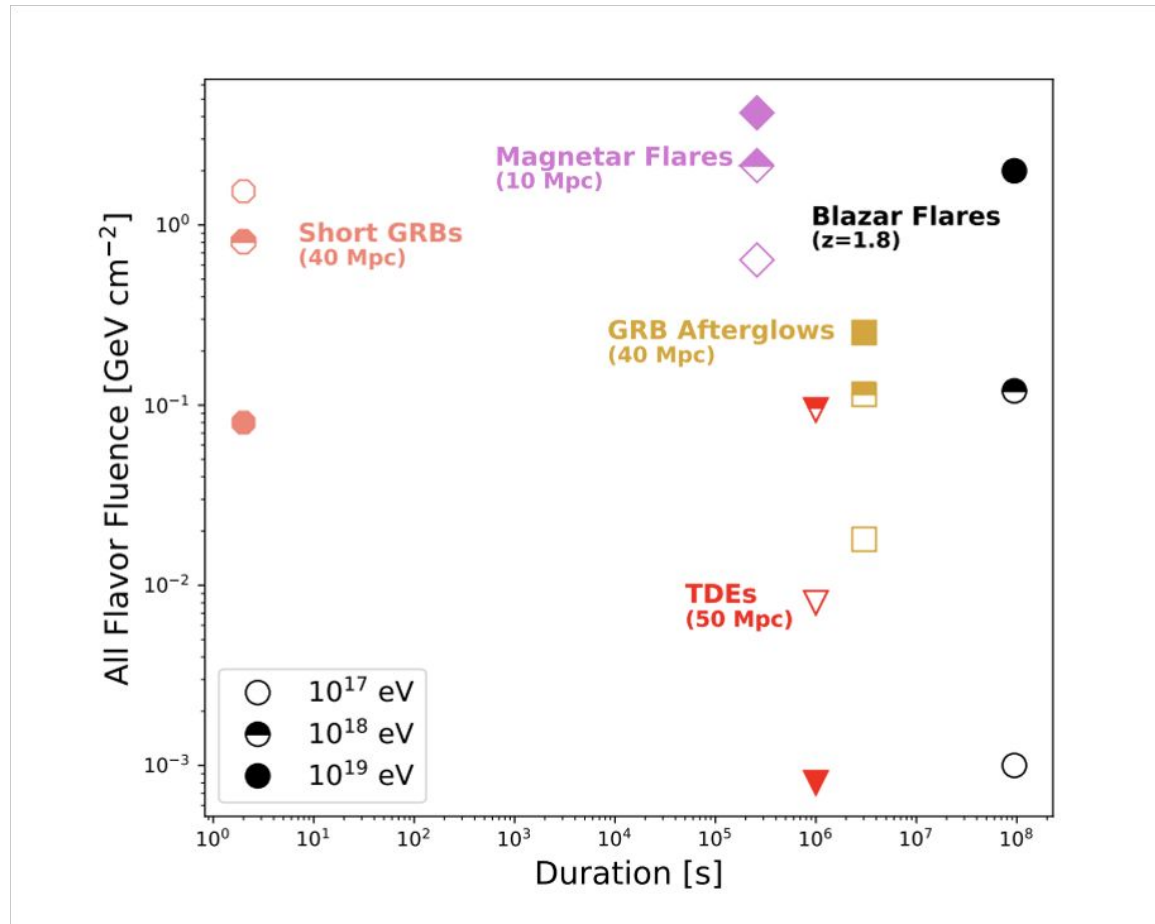
The Radio Neutrino Observatory Greenland

Why do we like Neutrinos



The Radio Neutrino Observatory Greenland

Progenitors and Sensitivities

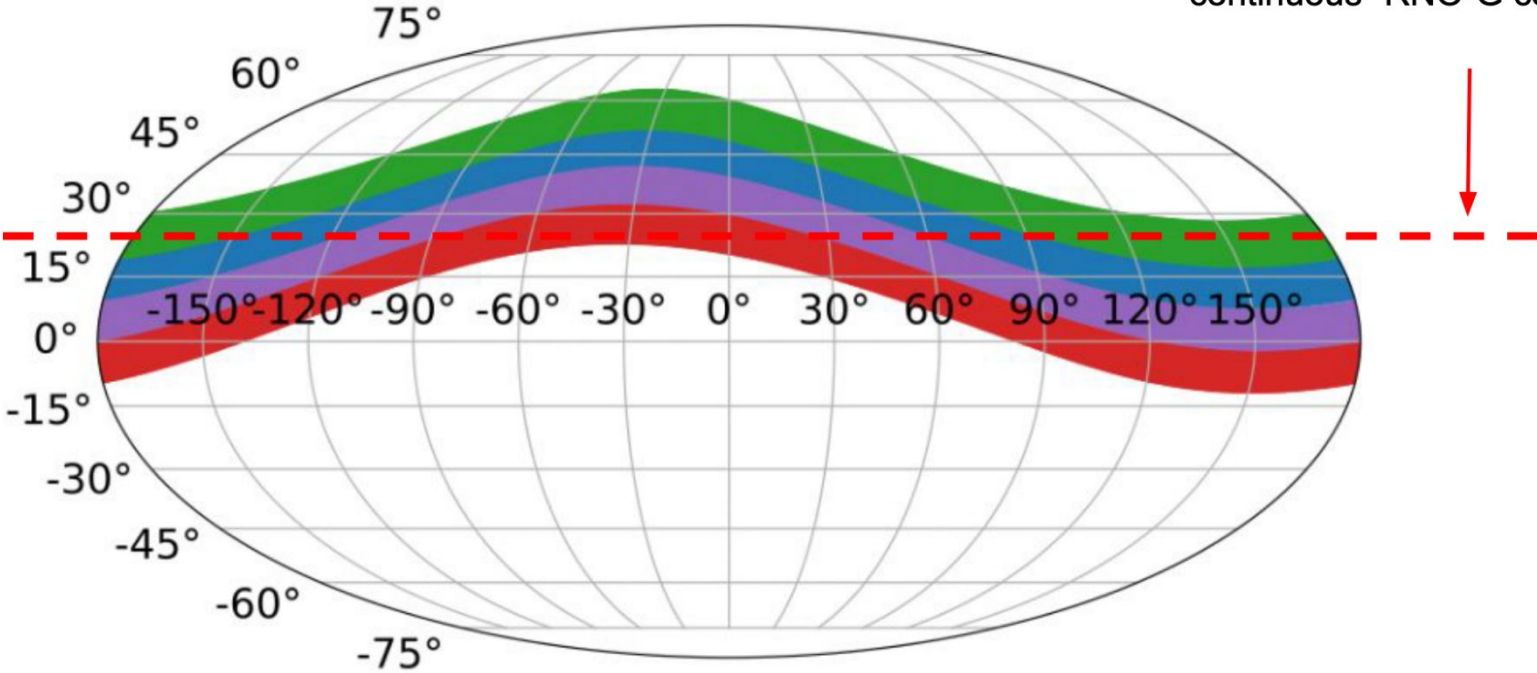


GRB 221009A

(non-)detection

GRB declination (19 deg)

One of the only dec bands with "continuous" RNO-G coverage

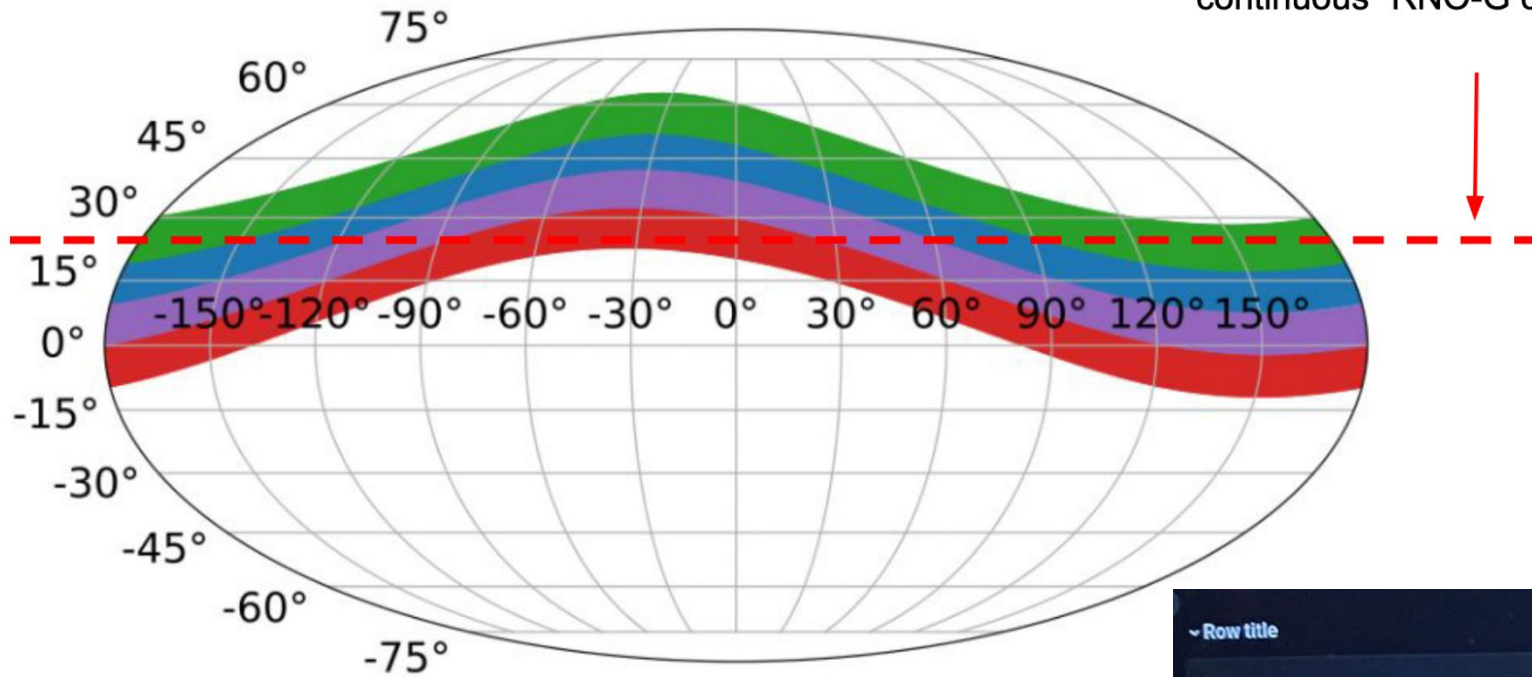


GRB 221009A

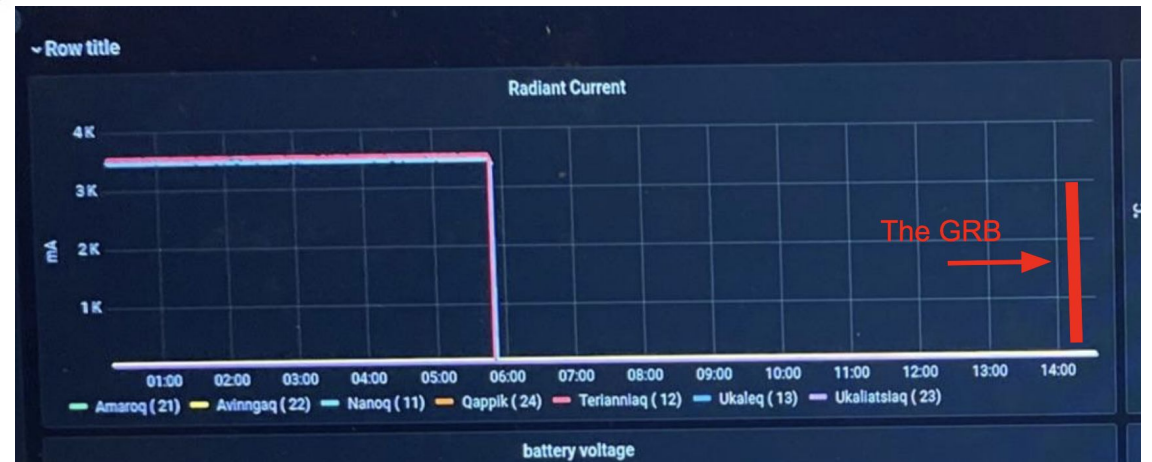
(non-)detection

GRB declination (19 deg)

One of the only dec bands with "continuous" RNO-G coverage

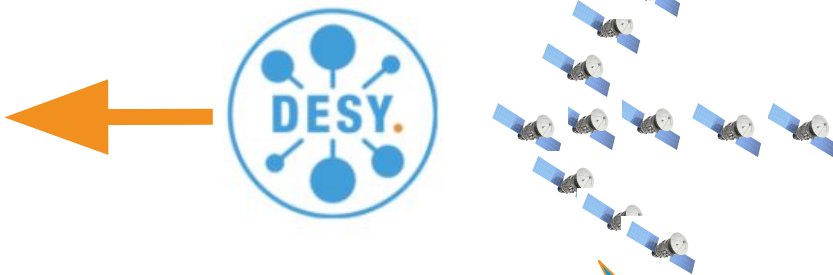
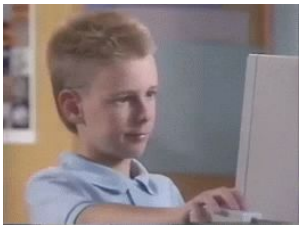
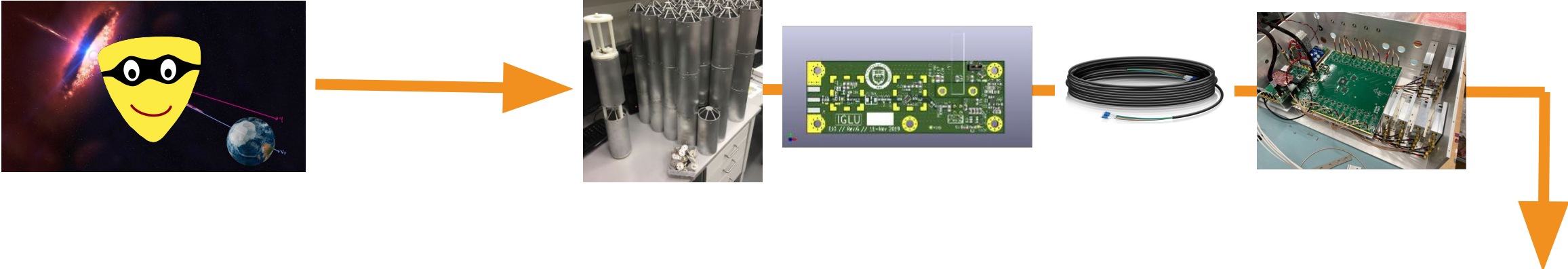


From B.Clark, 2022



Signal Chain

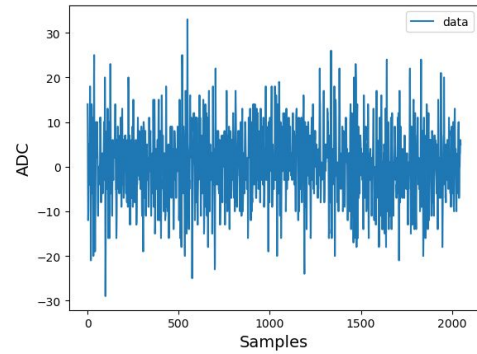
From the cosmos to your computer



10%
1% to 5%?

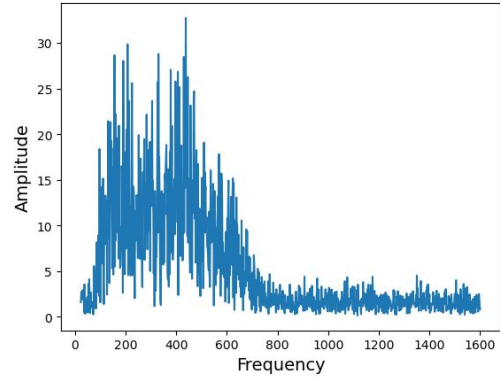
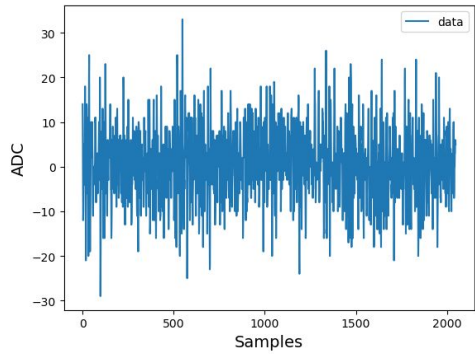
Anomaly Detection

Using STFT Mosaic of Phased Array



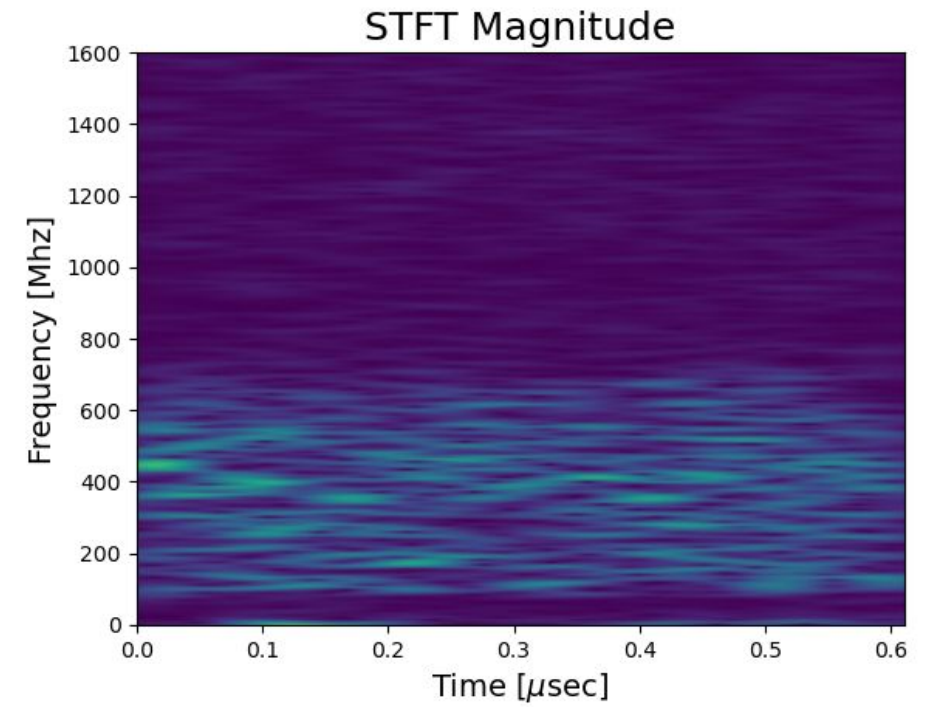
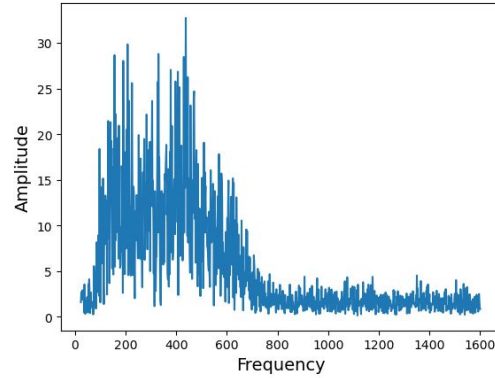
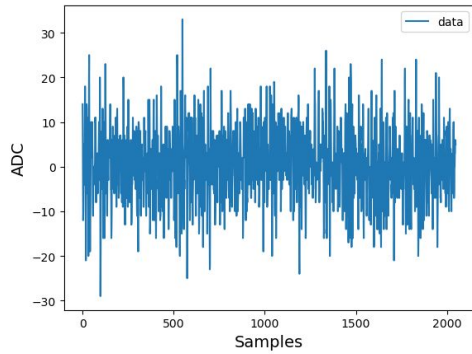
Anomaly Detection

Using STFT Mosaic of Phased Array



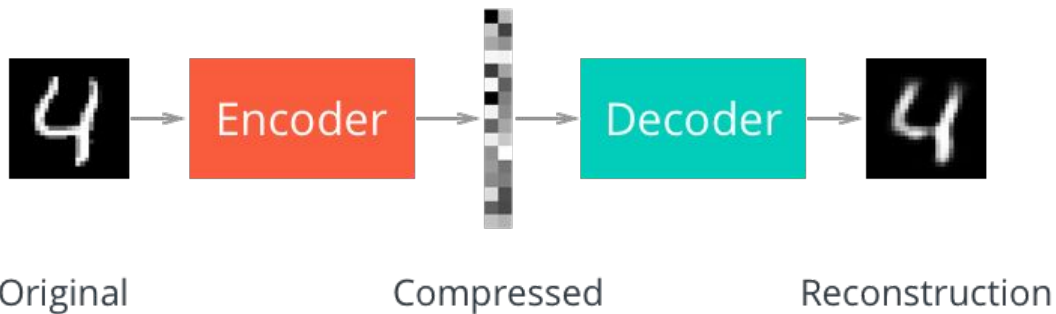
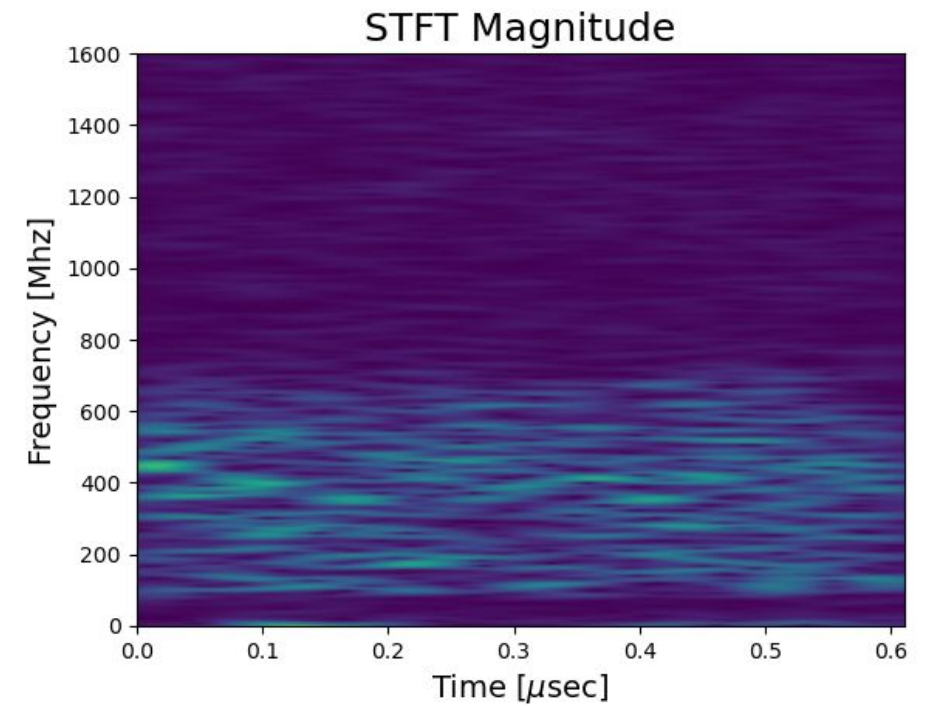
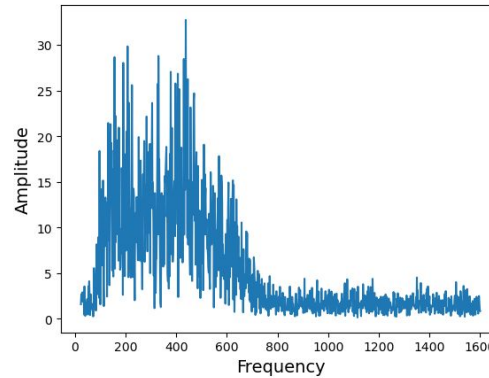
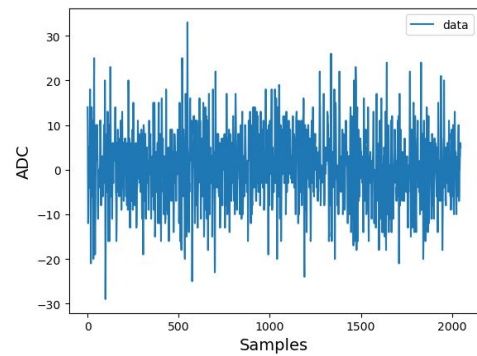
Anomaly Detection

Using STFT Mosaic of Phased Array



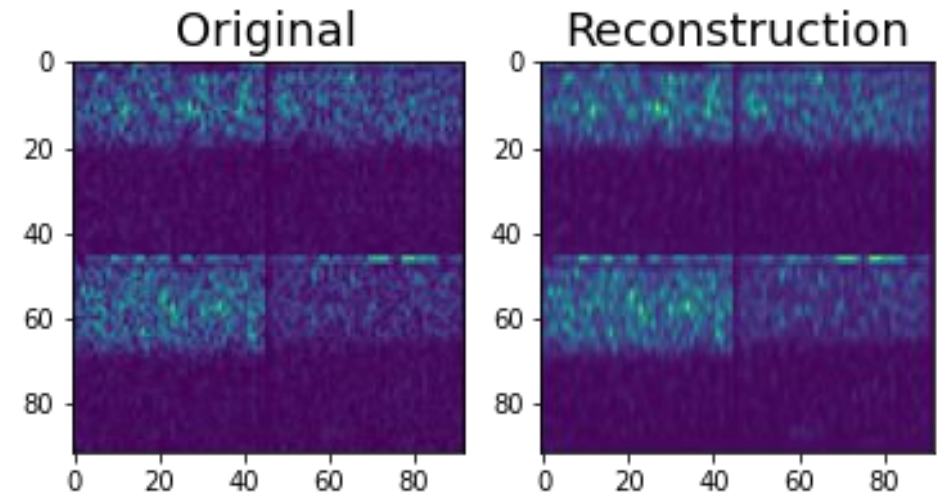
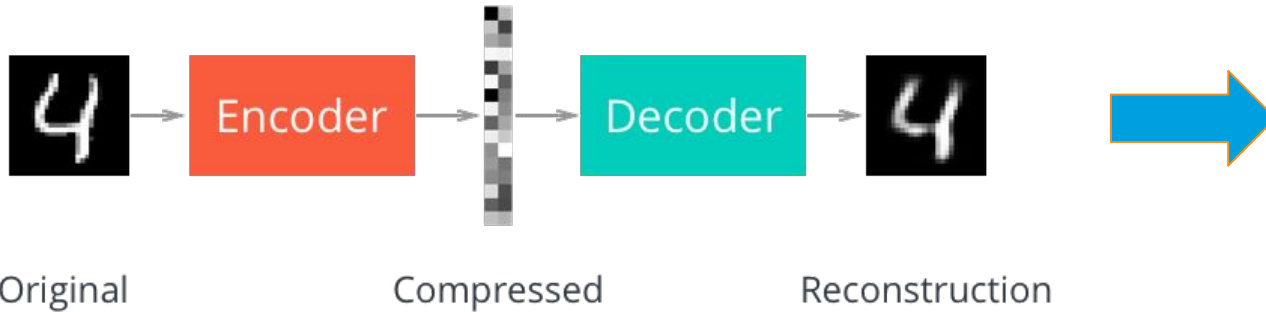
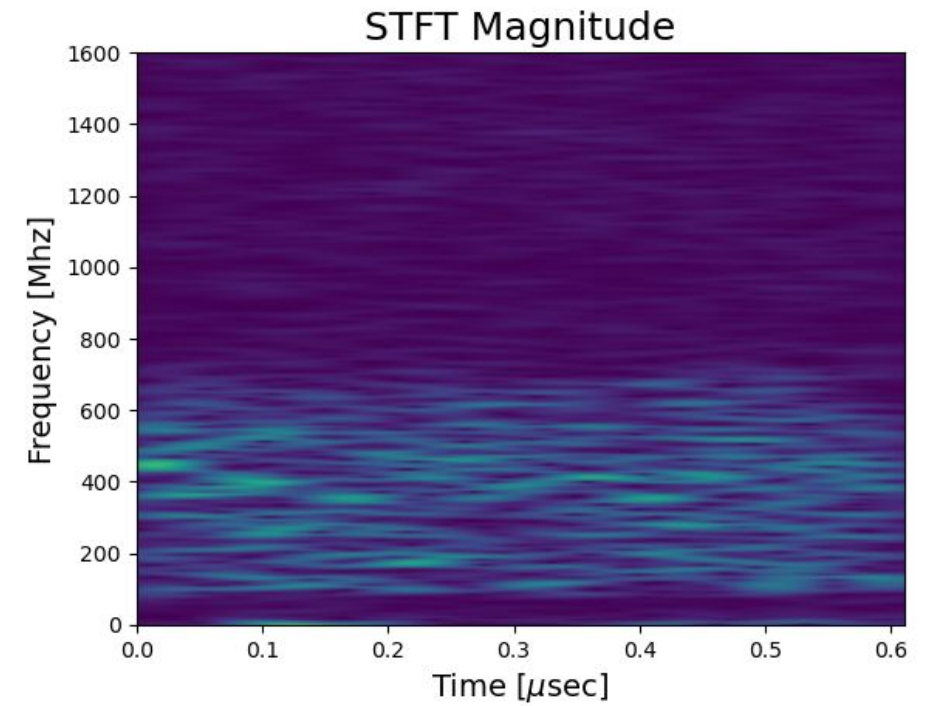
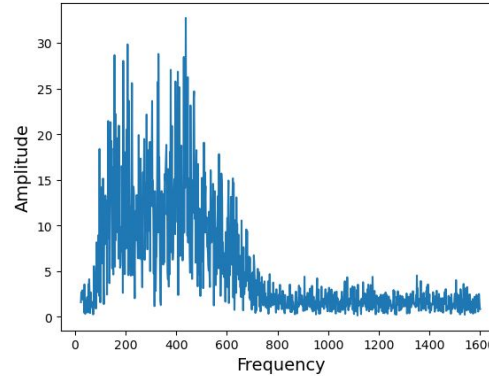
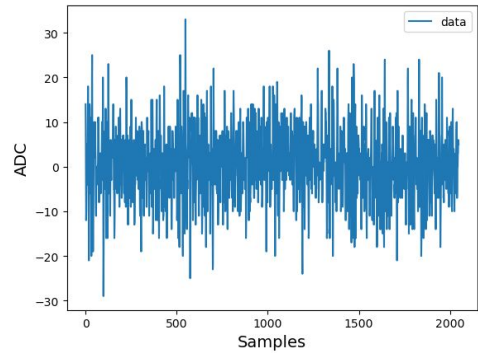
Anomaly Detection

Using STFT Mosaic of Phased Array



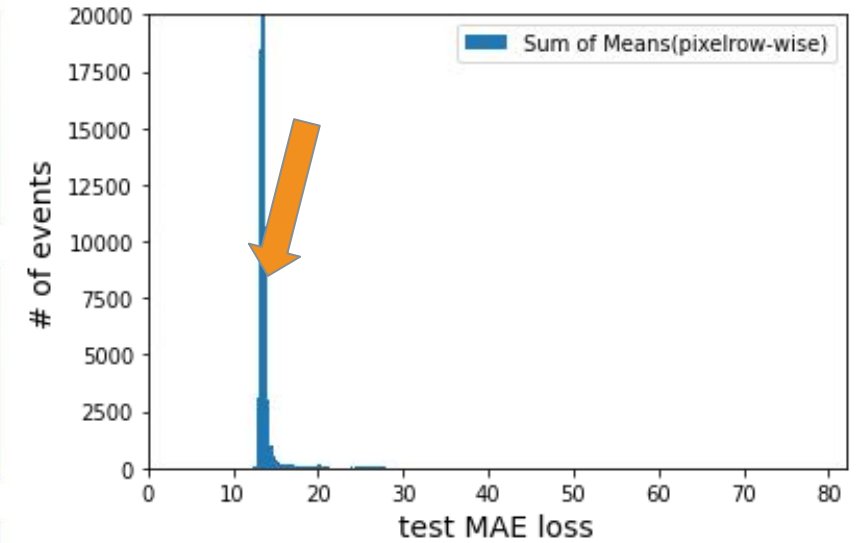
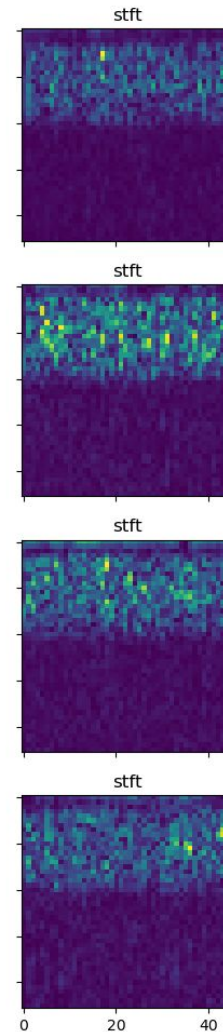
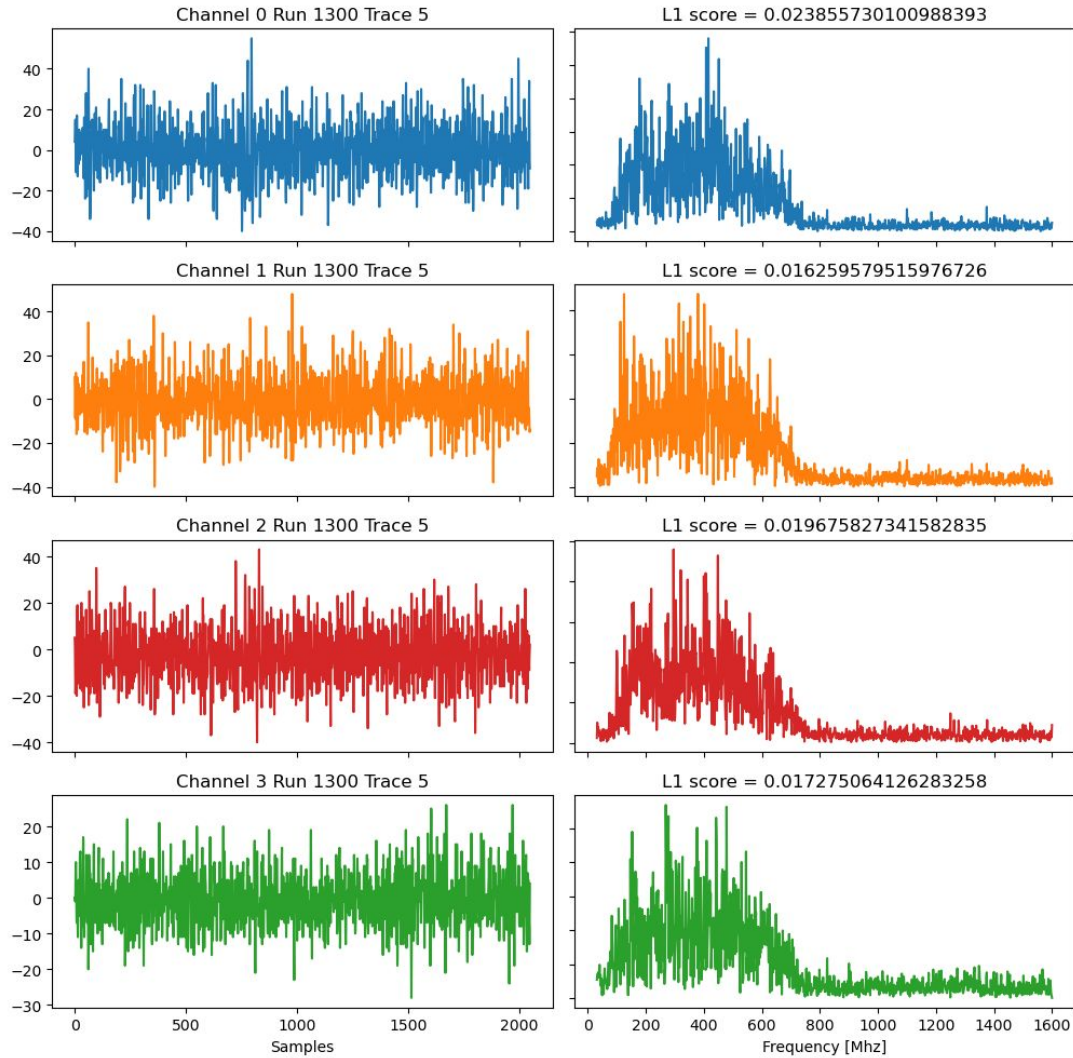
Anomaly Detection

Using STFT Mosaic of Phased Array



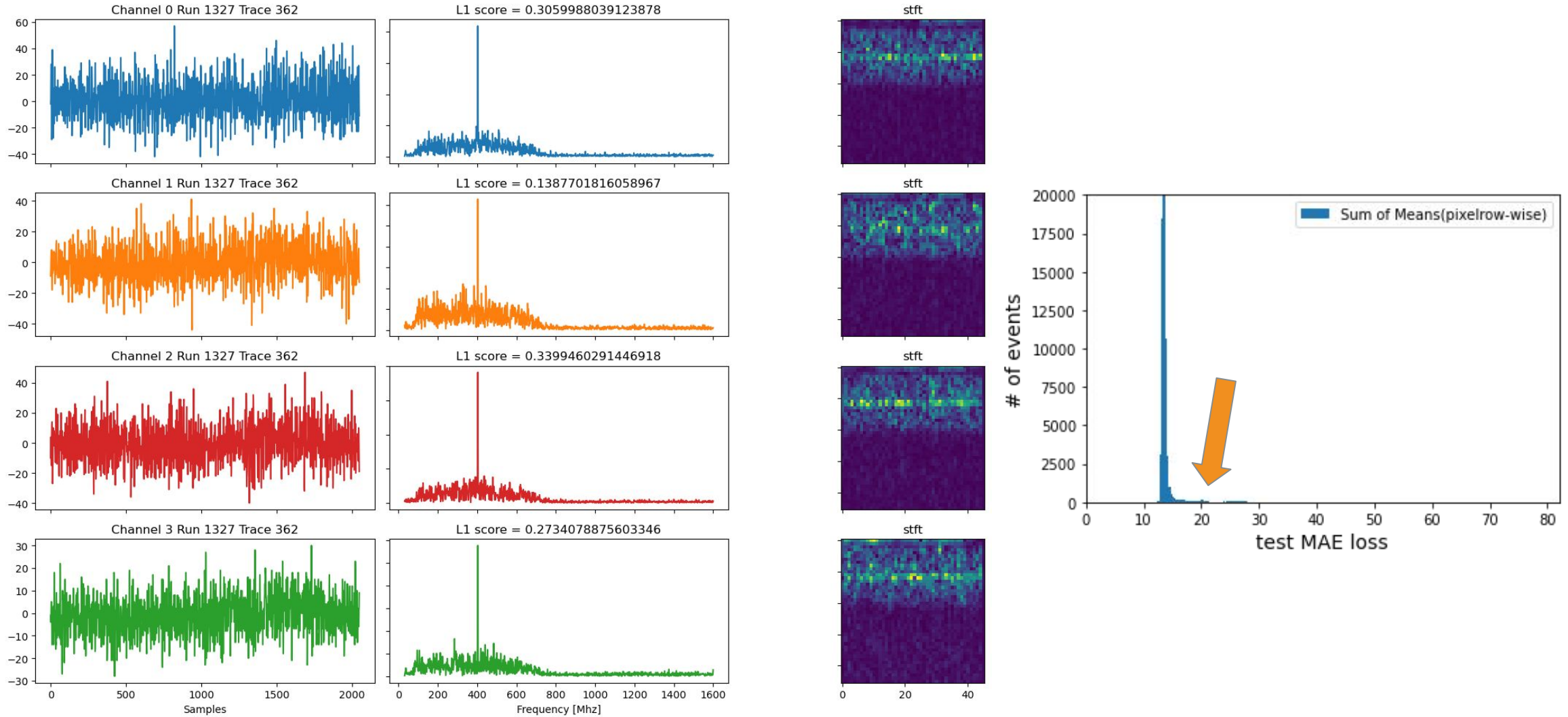
Anomaly Detection

Thermal Noise



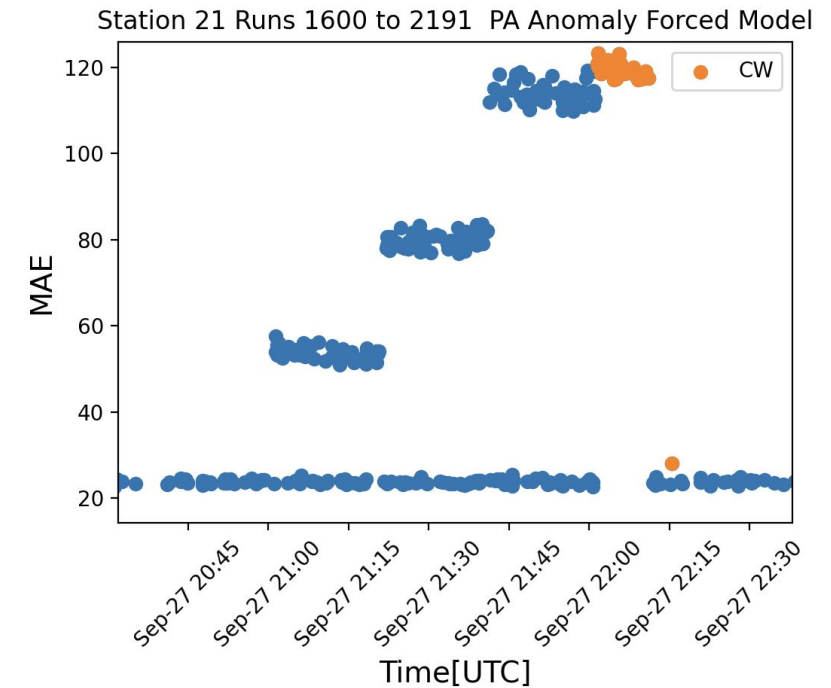
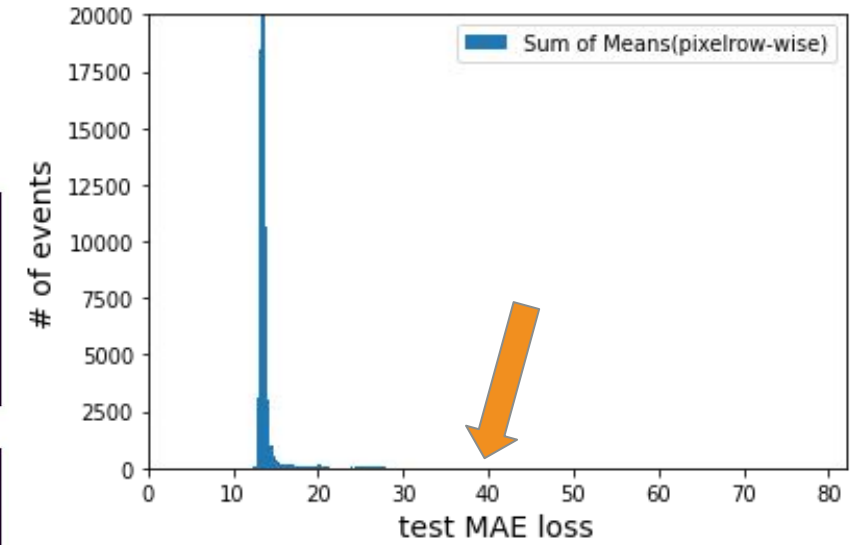
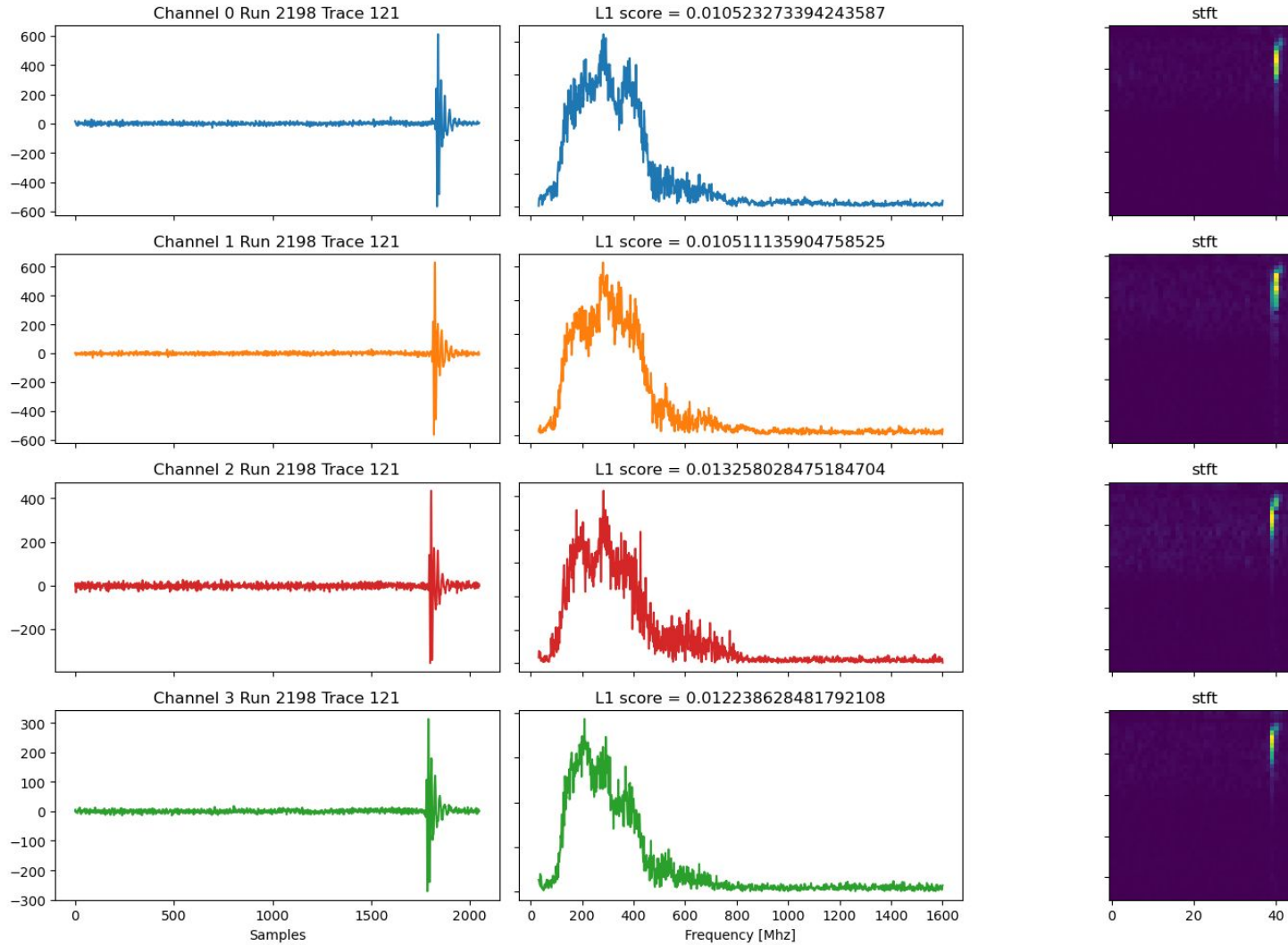
Anomaly Detection

Continuous Wave (CW)



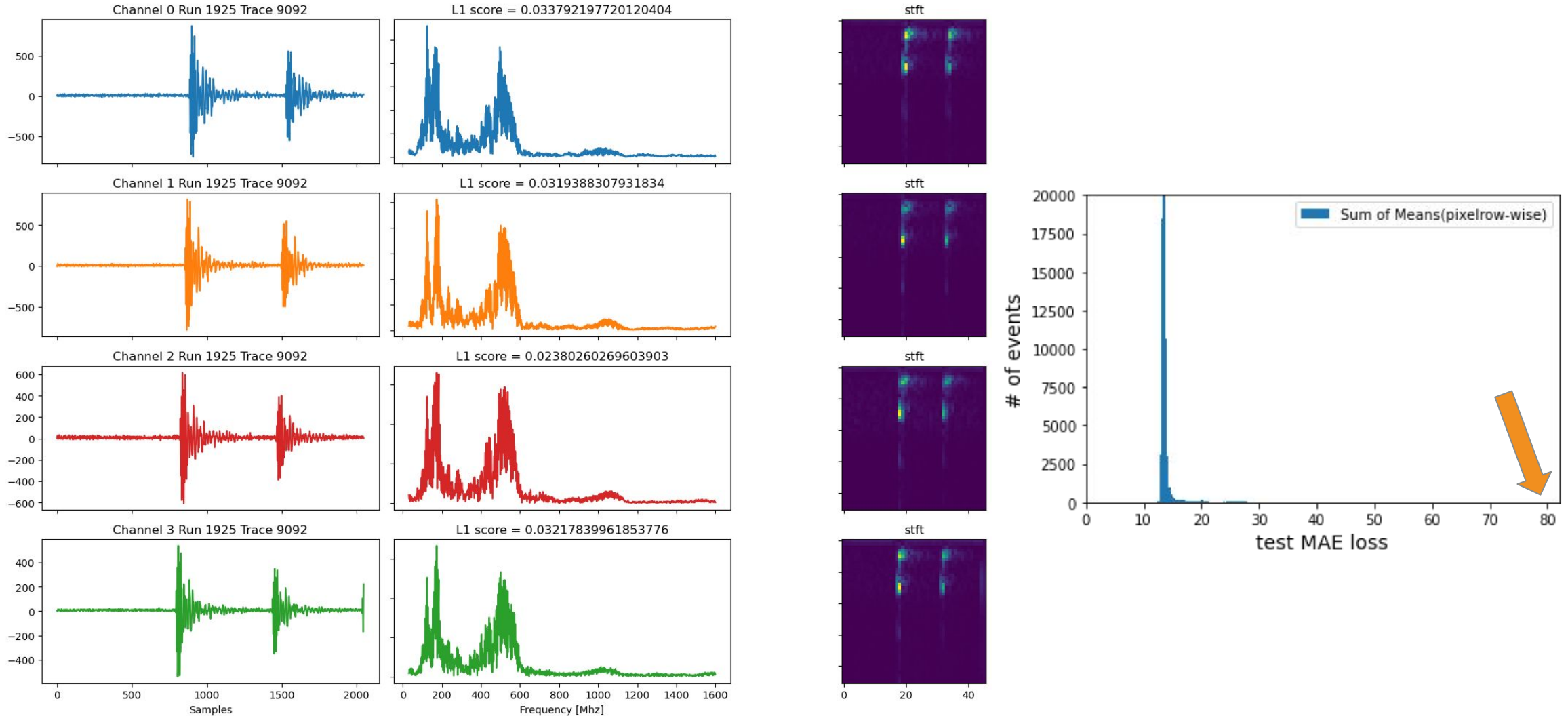
Anomaly Detection

Calibration Pulsing



Anomaly Detection

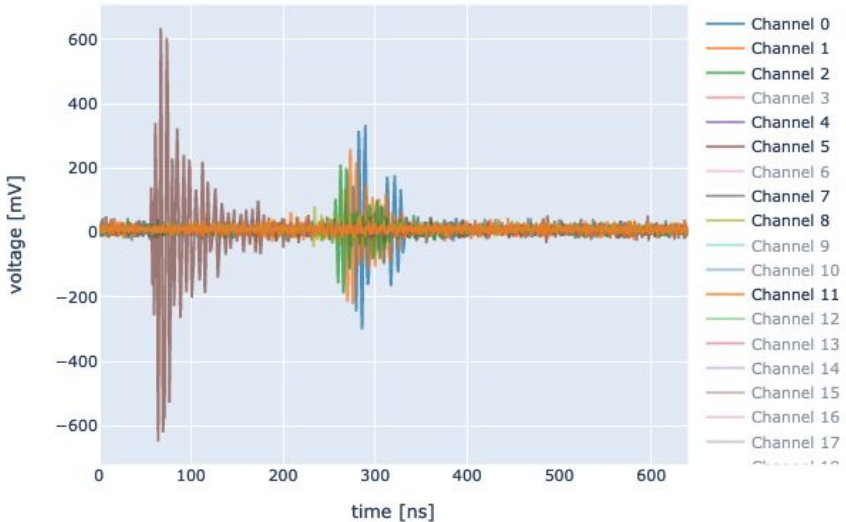
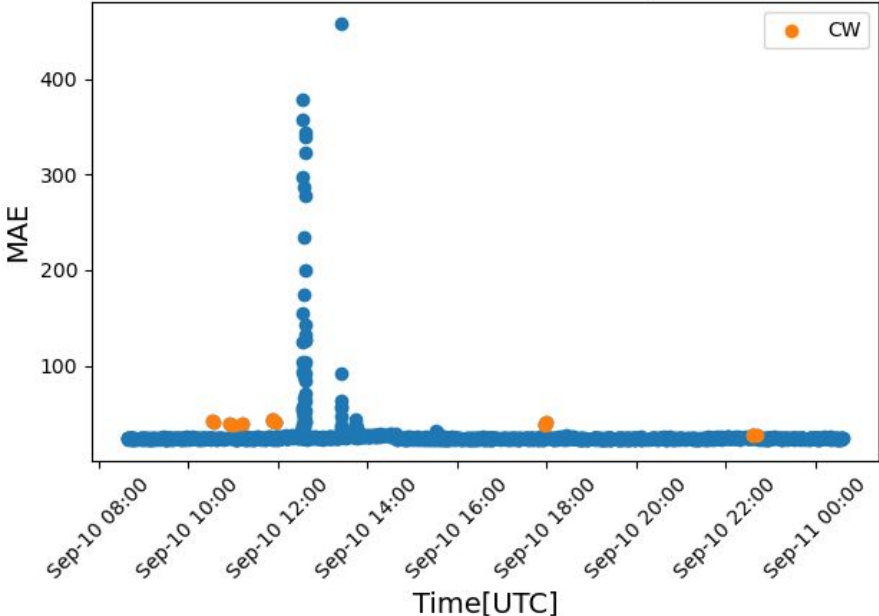
Storm? Ice Quakes?



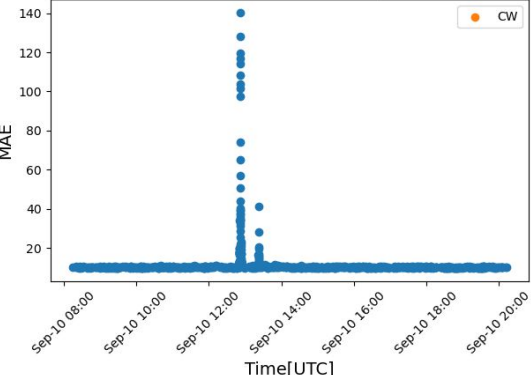
Snowmobiles?

September 10th – Snowmobile?

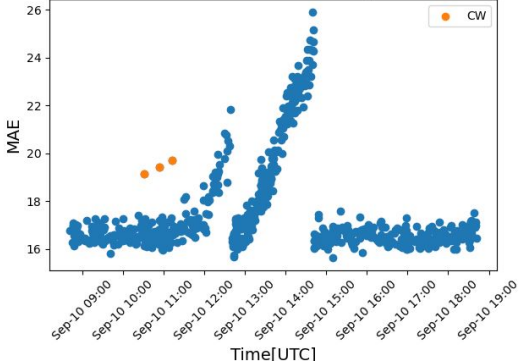
Station 21 Runs 1924 to 1931 PA Anomaly Forced Model



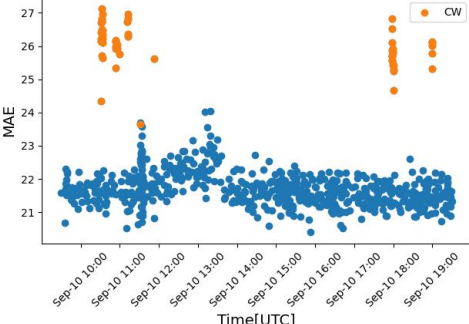
Station 13 Runs 817 to 822 PA Anomaly Forced Model



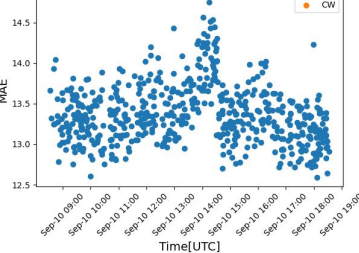
Station 22 Runs 1849 to 1853 PA Anomaly Forced Model



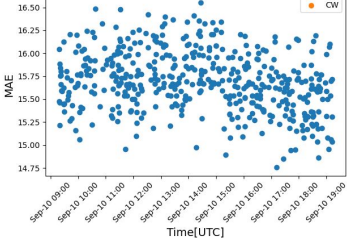
Station 11 Runs 1527 to 1531 PA Anomaly Forced Model



Station 23 Runs 835 to 839 PA Anomaly Forced Model

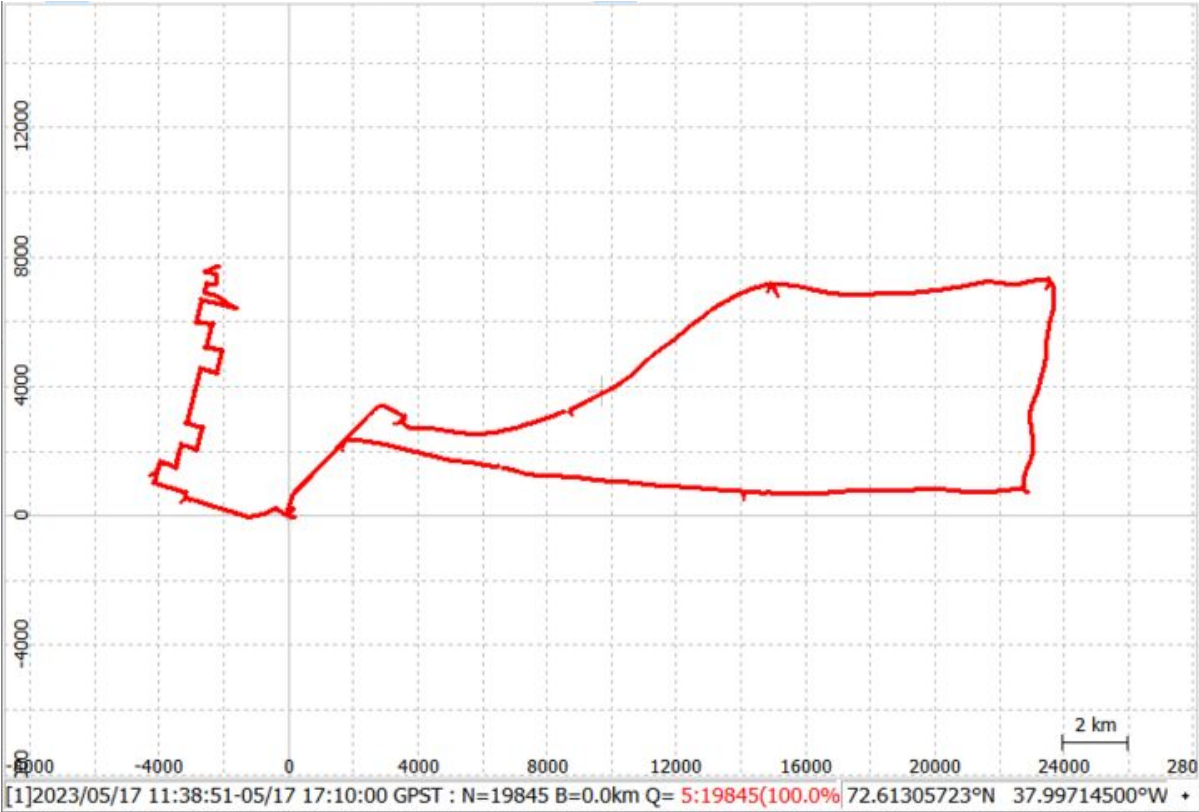
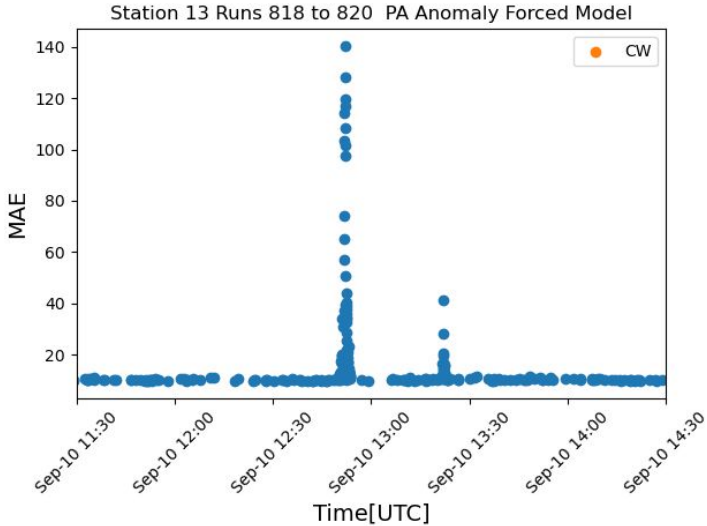
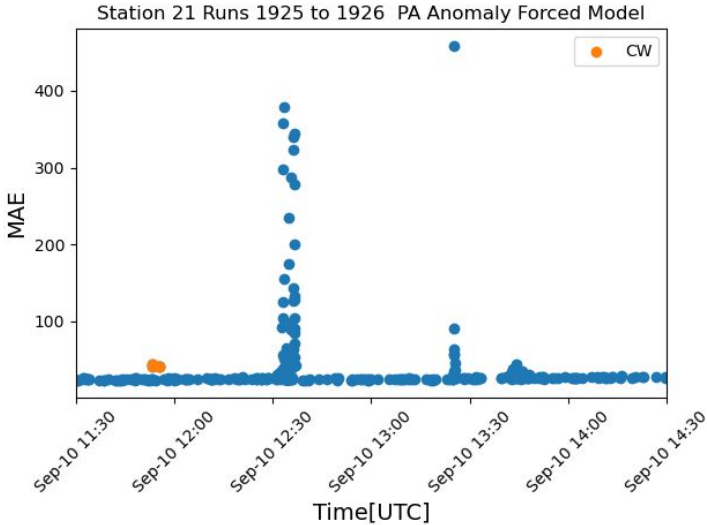


Station 24 Runs 675 to 679 PA Anomaly Forced Model



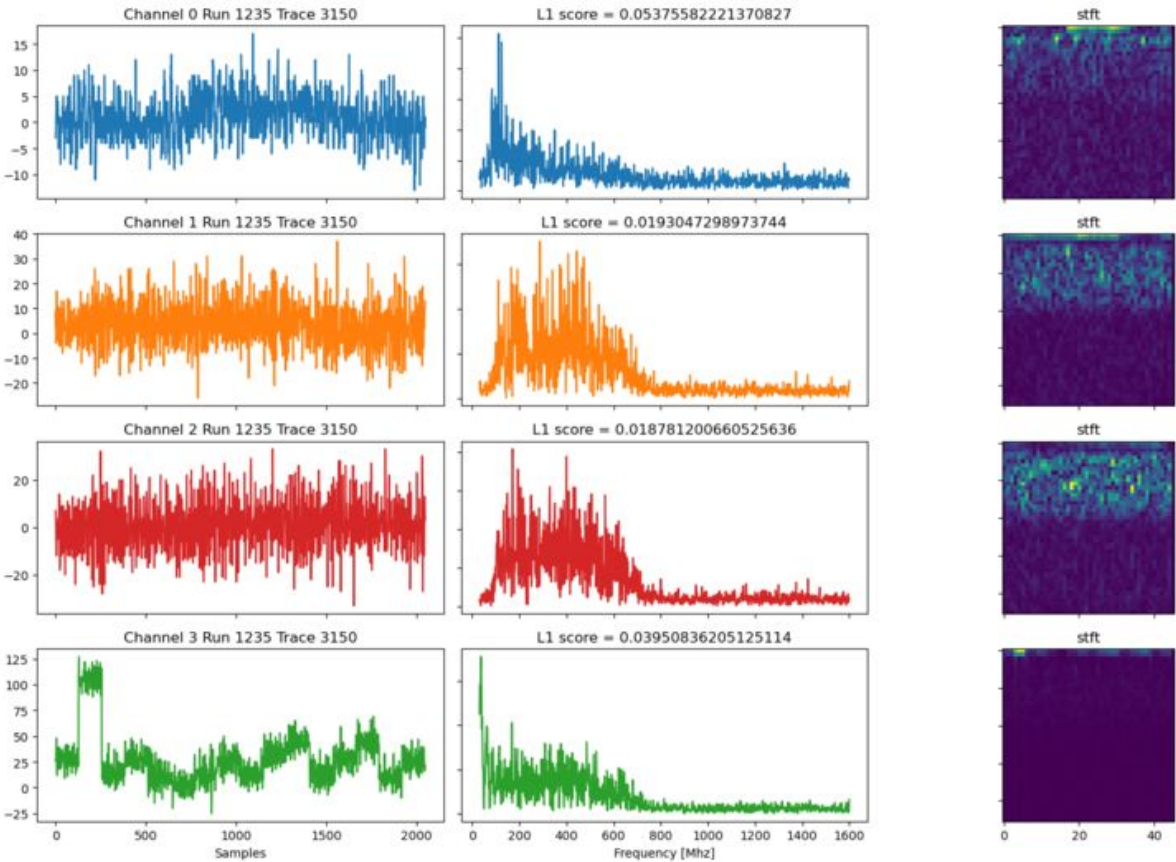
Snowmobiles!

Ice Sat Traverse



Block Offsets

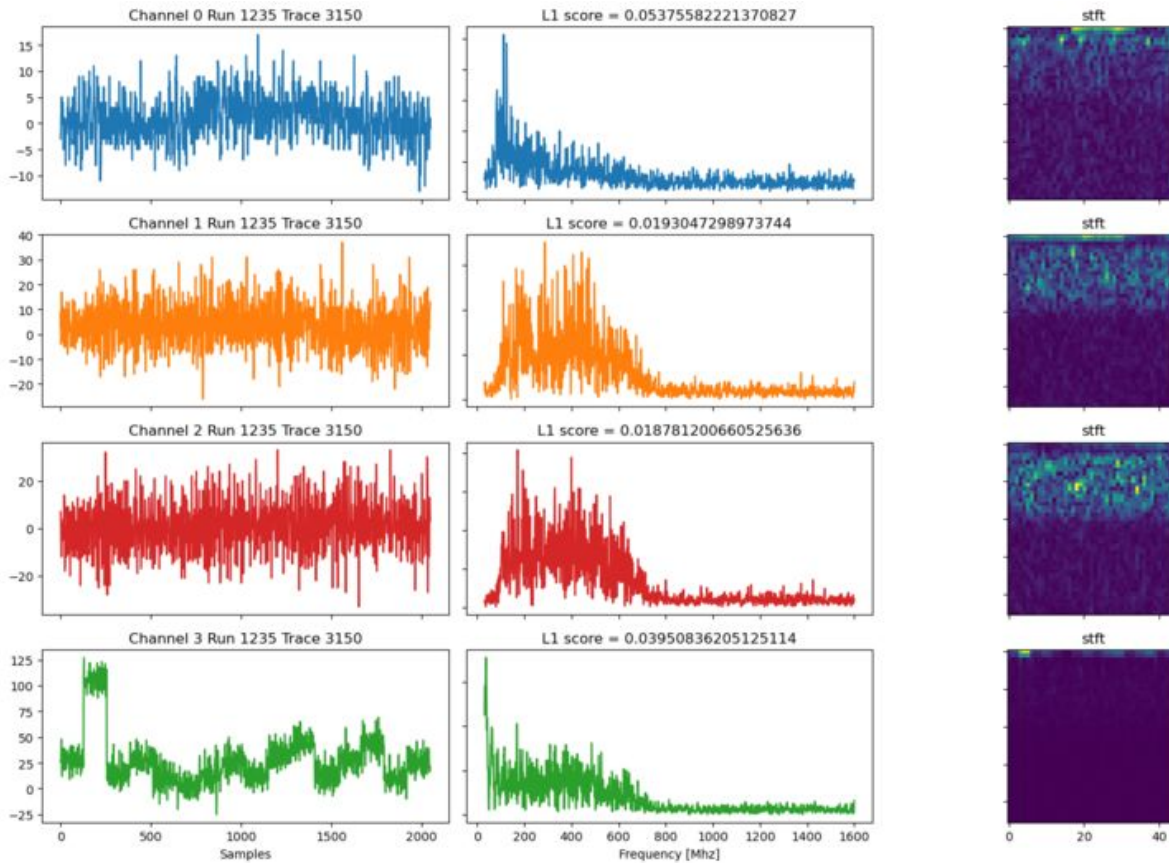
Station 13 channel 3 suffers from frequent block offsets



- LAB4D digitizes in 16 blocks of 128 samples
- Probably due to unstable bias in the supplied voltage
- Fixable in level 1 calibration

Block Offsets

Station 13 channel 3 suffers from frequent block offsets

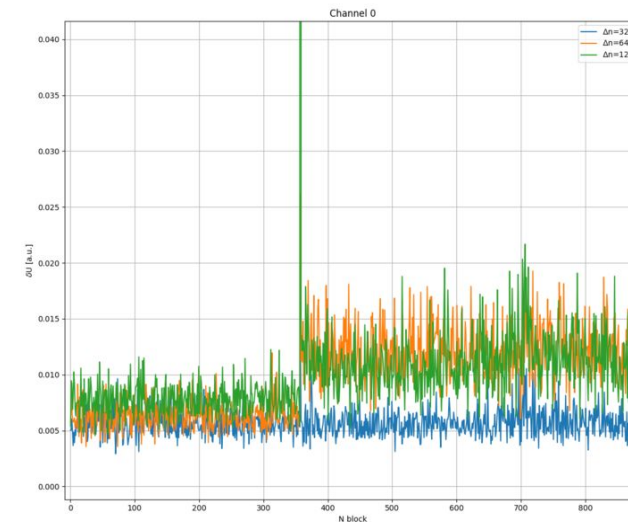
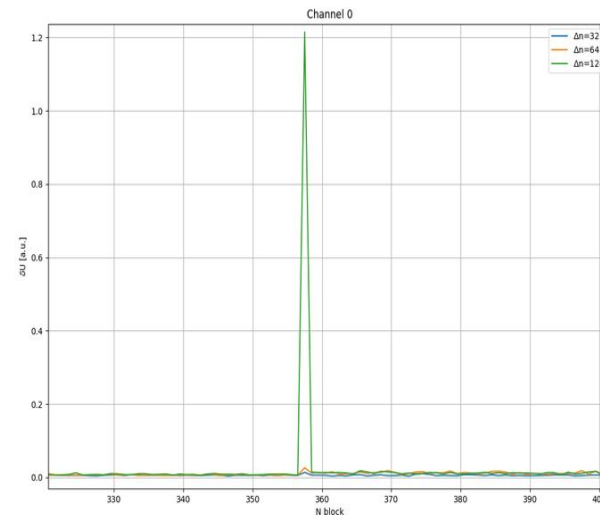
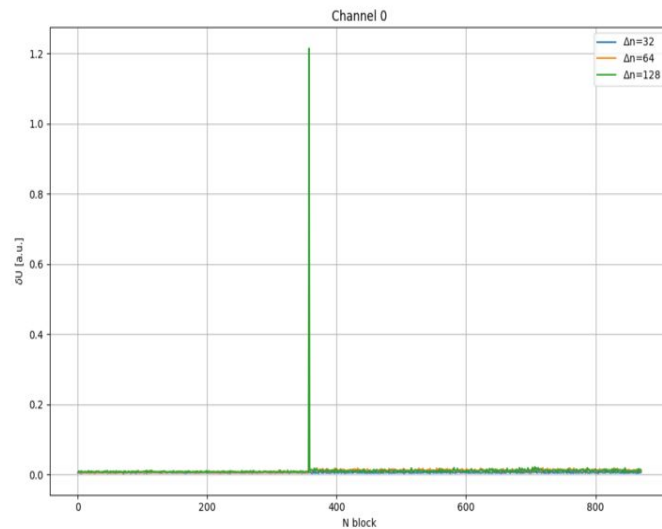
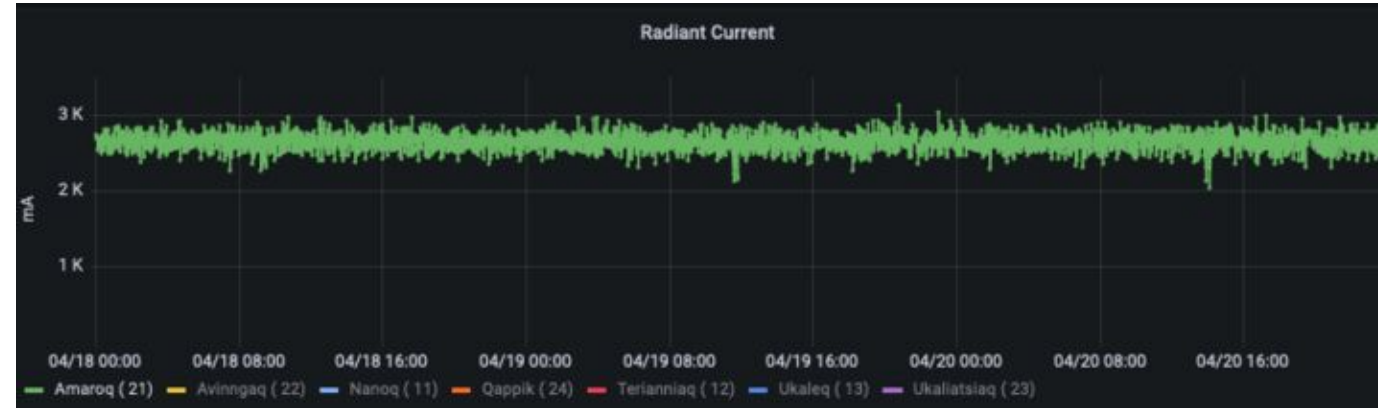
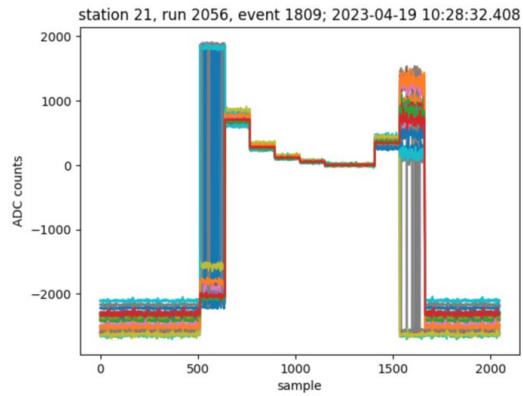


- LAB4D digitizes in 16 blocks of 128 samples
- Probably due to unstable bias in the supplied voltage
- Fixable in level 1 calibration

MITIGATED

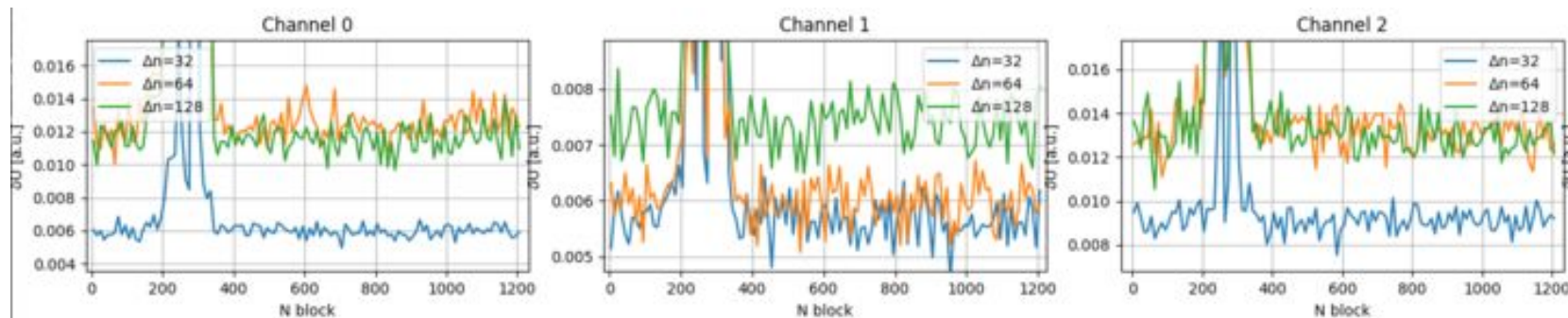
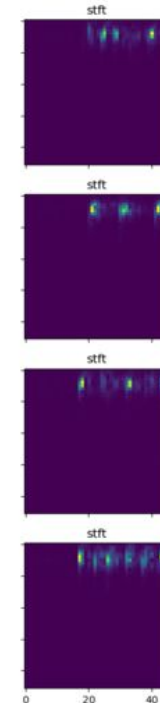
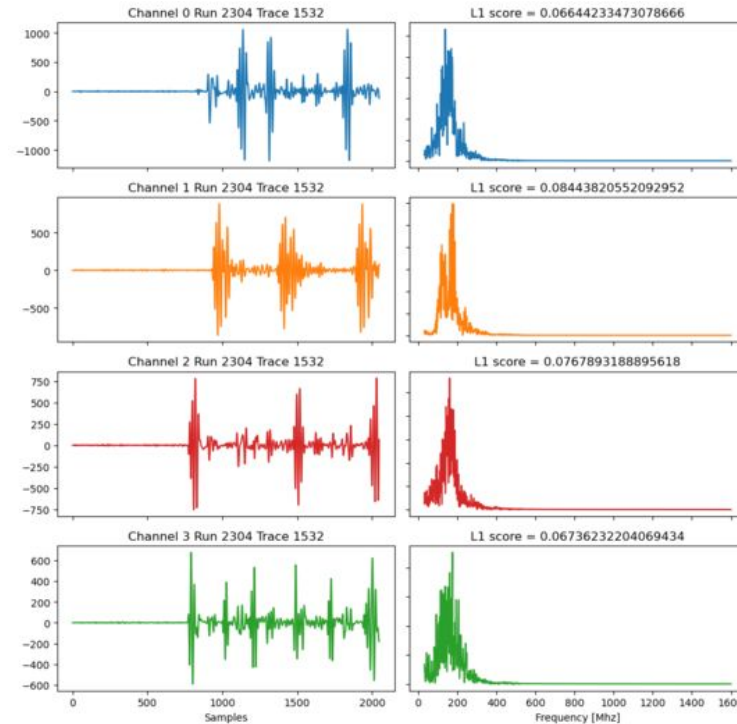
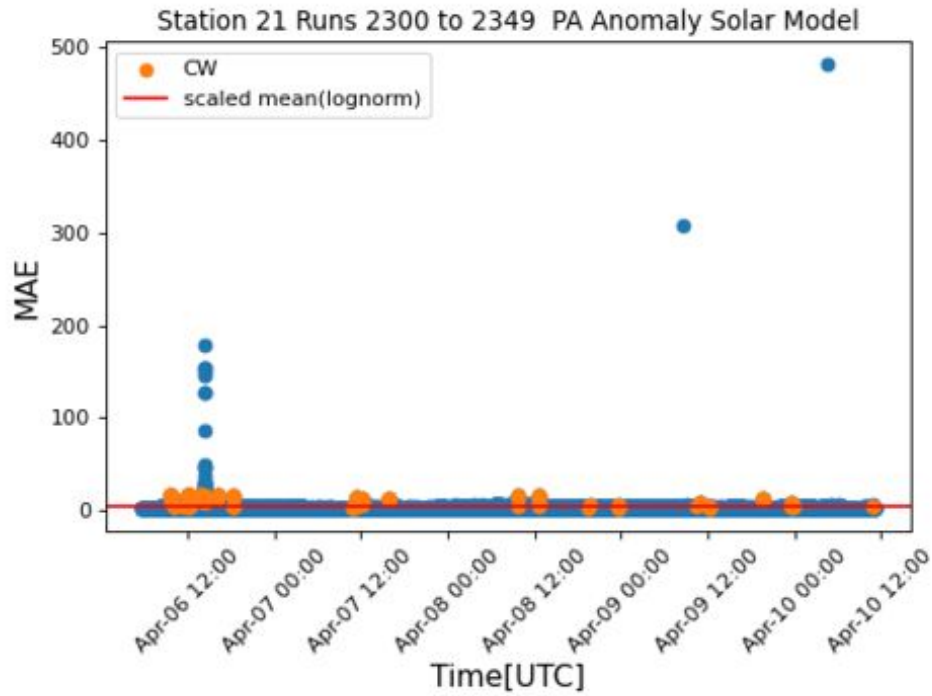
Another Anomalous Event in 21 on April 19

MEGA Block offsets. Glitching can occur



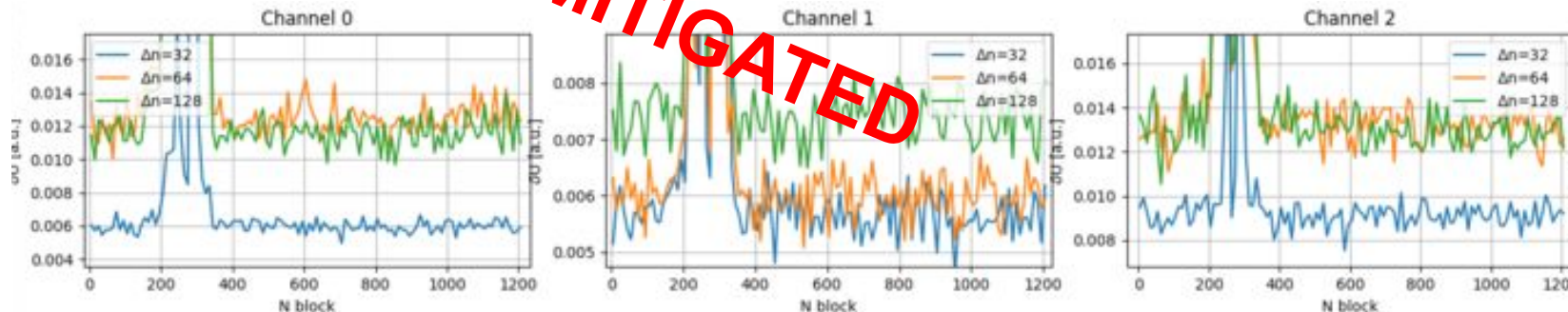
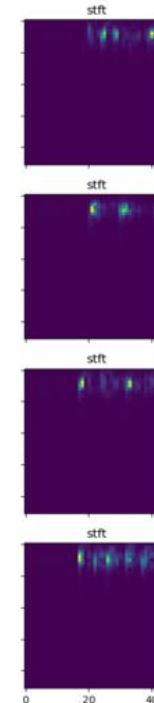
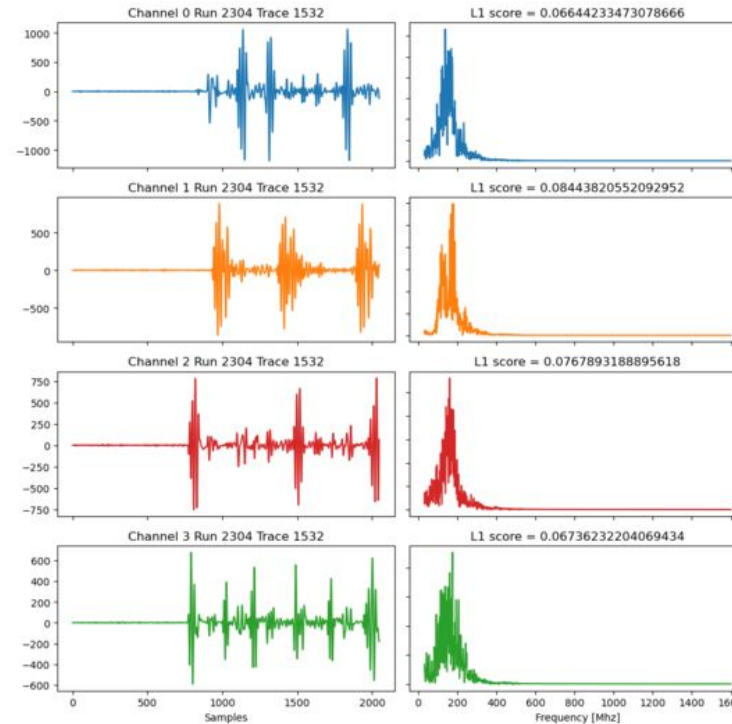
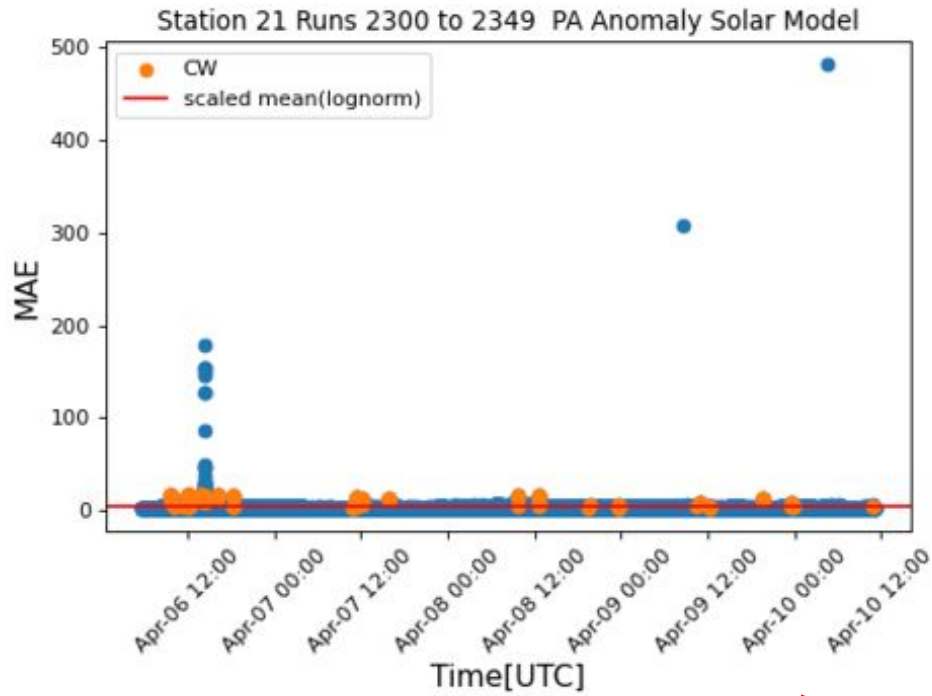
Another Anomalous Event in 21 on April 19

MEGA Block offsets. Caused by loss of power?



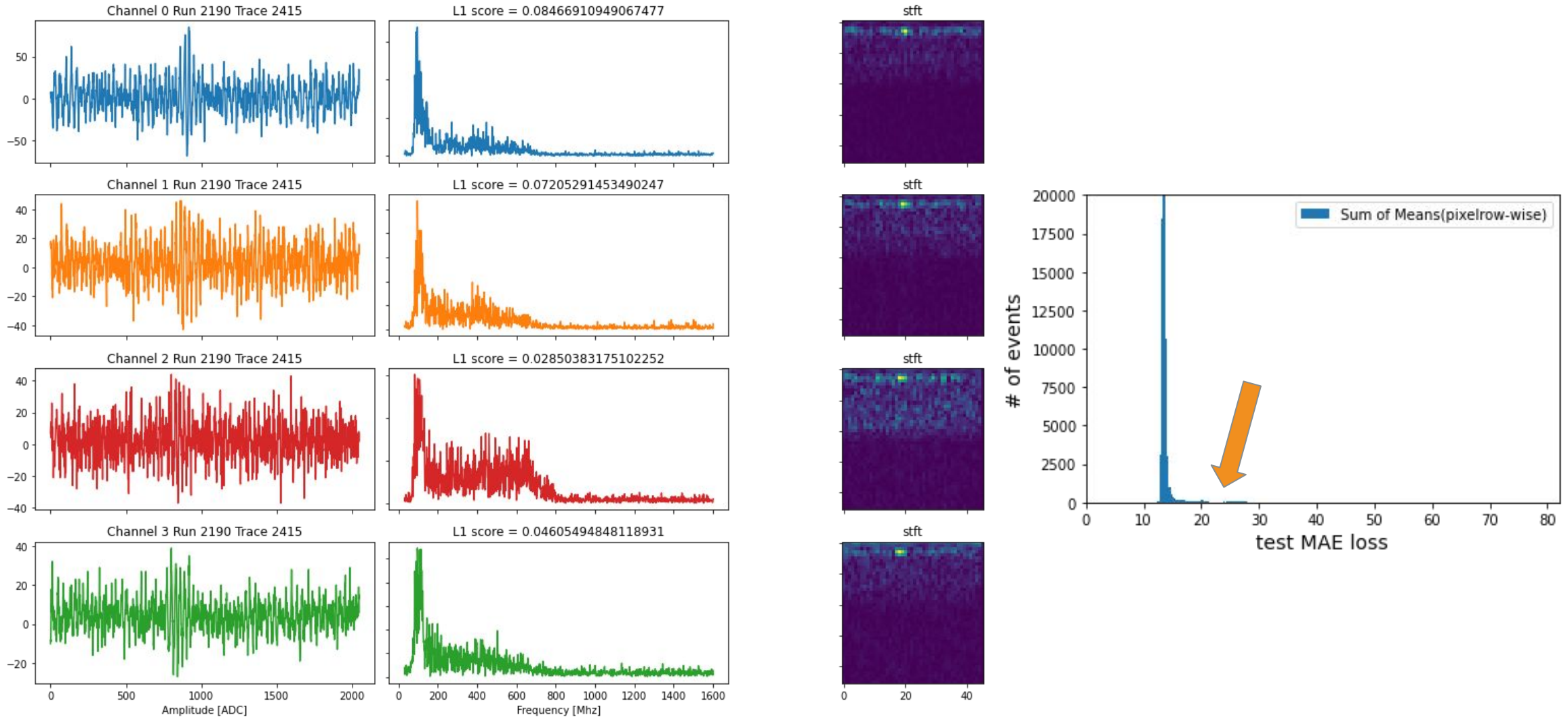
Another Anomalous Event in 21 on April 19

MEGA Block offsets. Caused by loss of power?



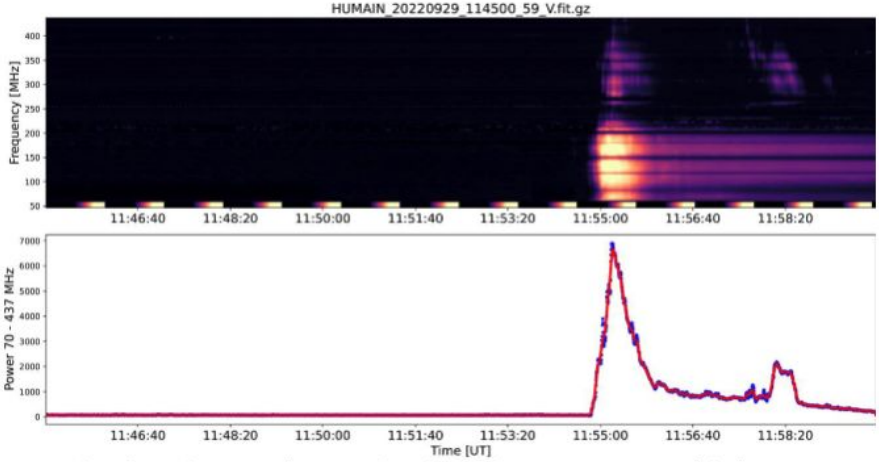
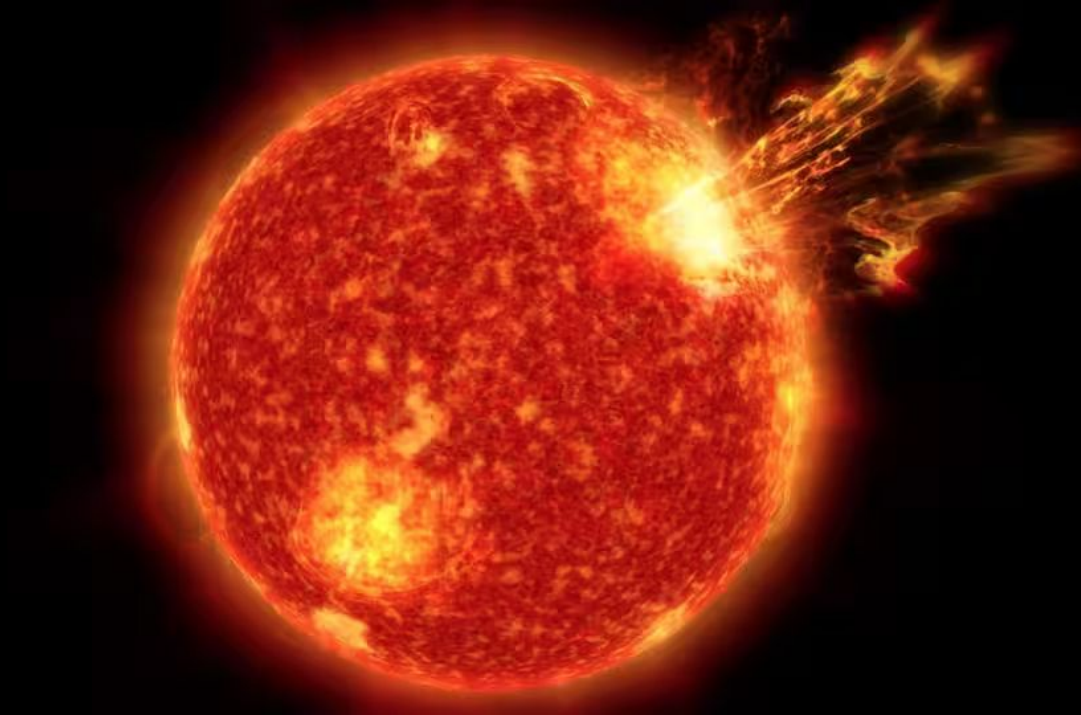
Anomaly Detection

Anomalous Low Threshold Event

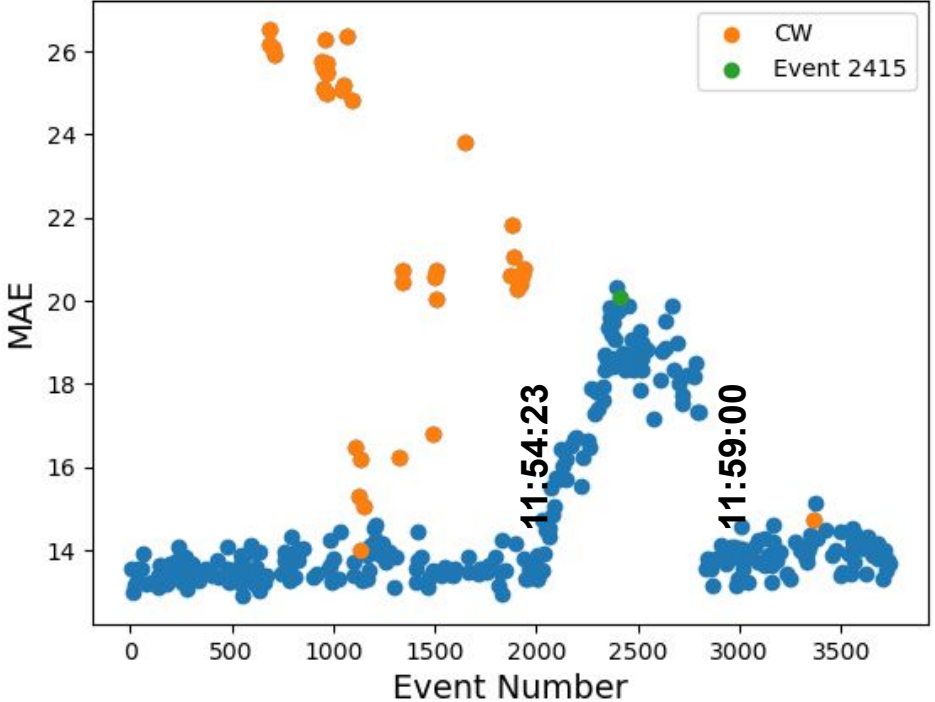


“Anomalous Low Threshold event”

Type II Solar Flare



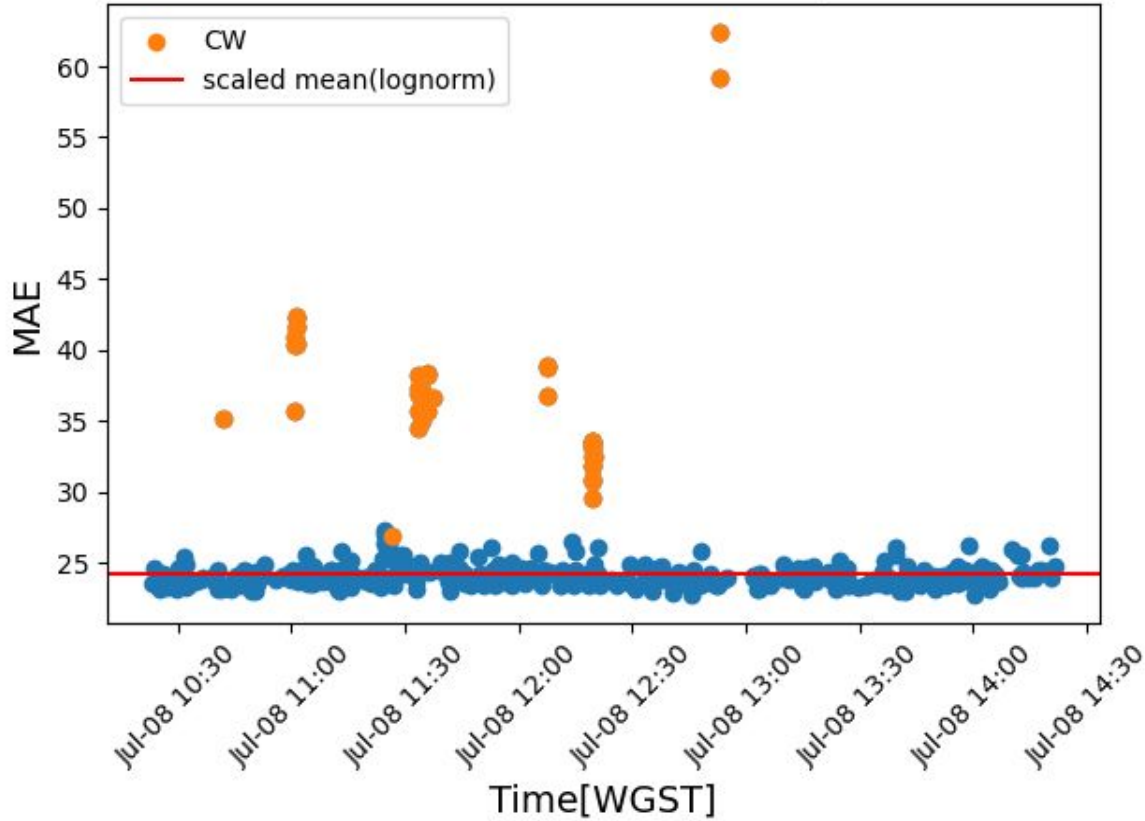
Station 21 Run 2190 Phased Array Anomaly Scores



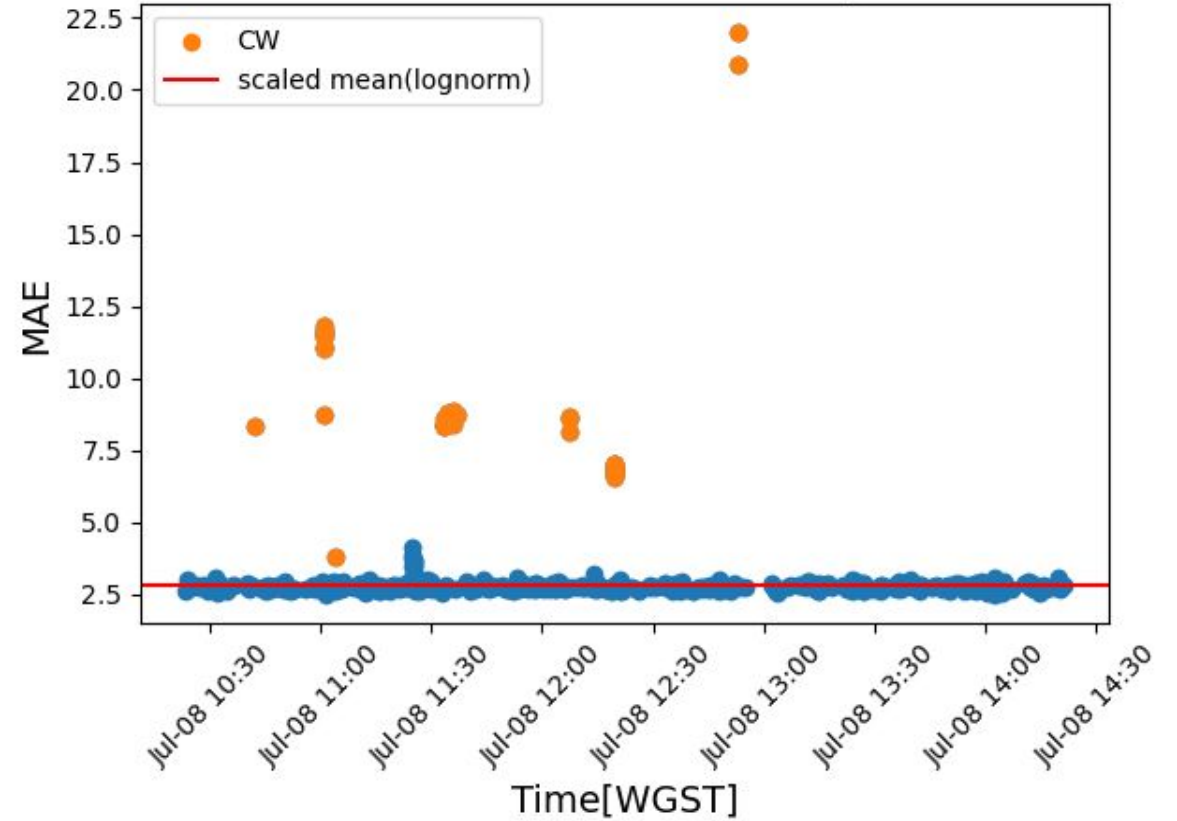
“Solar Model”

Butterworth filter 25 - 250 MHz

Station 21 Runs 1219 to 1220 PA Anomaly Forced Model

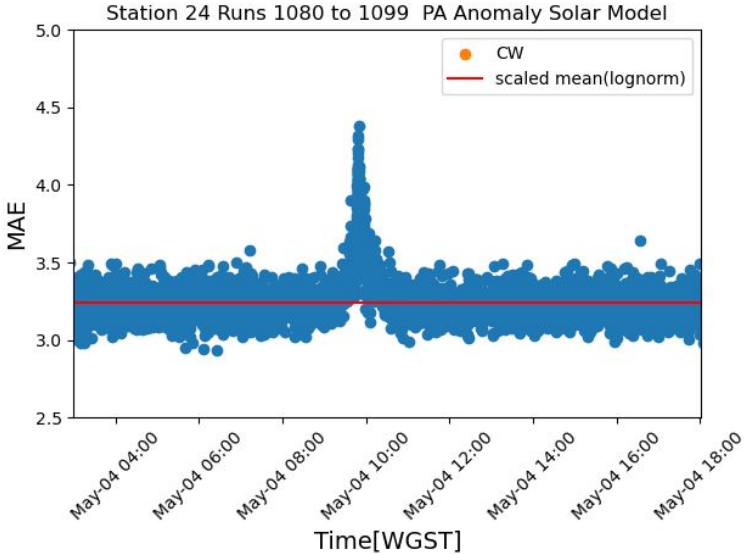
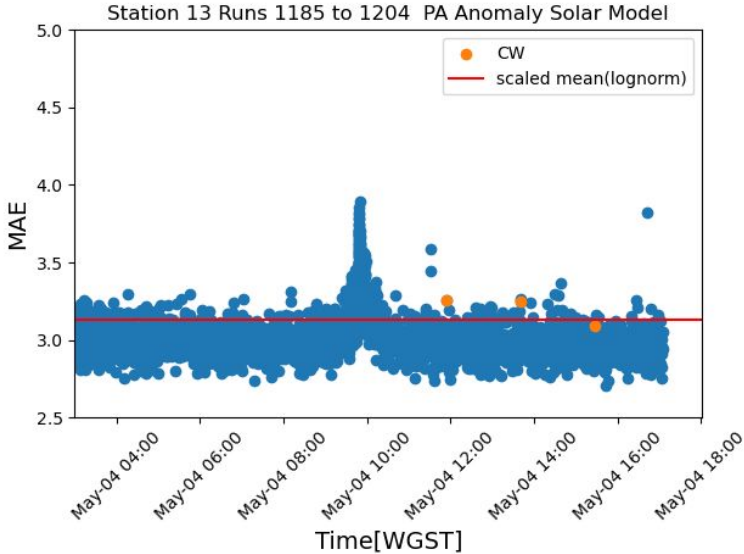
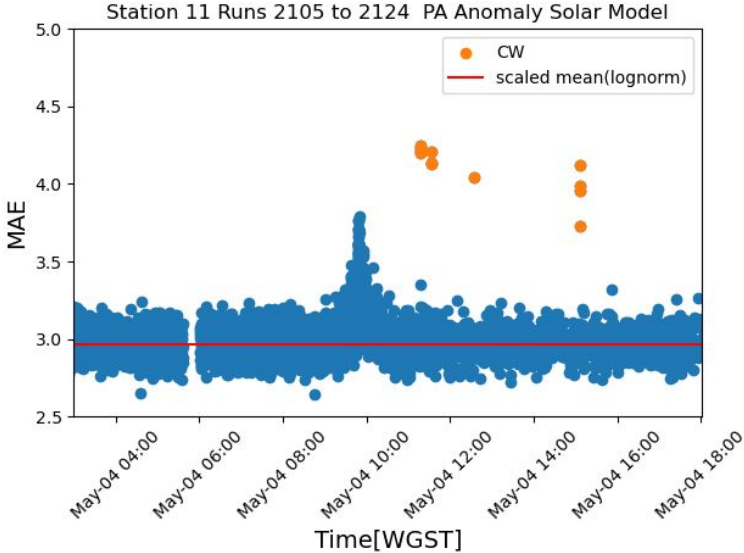
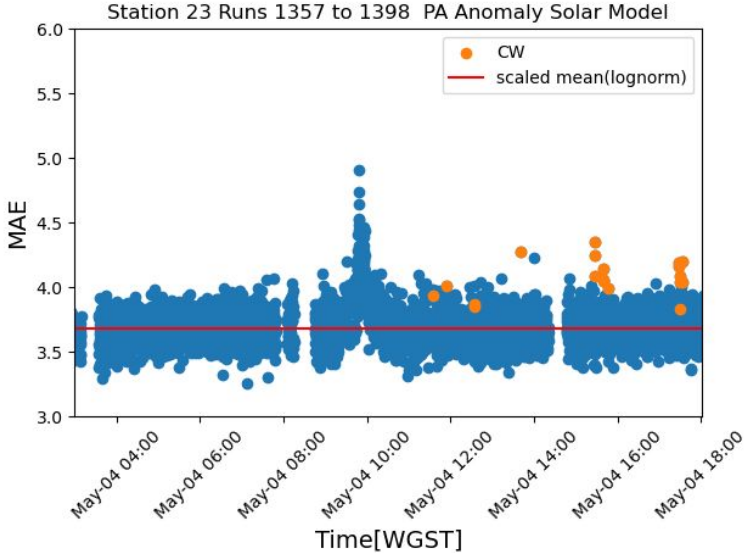


Station 21 Runs 1219 to 1220 PA Anomaly Solar Model



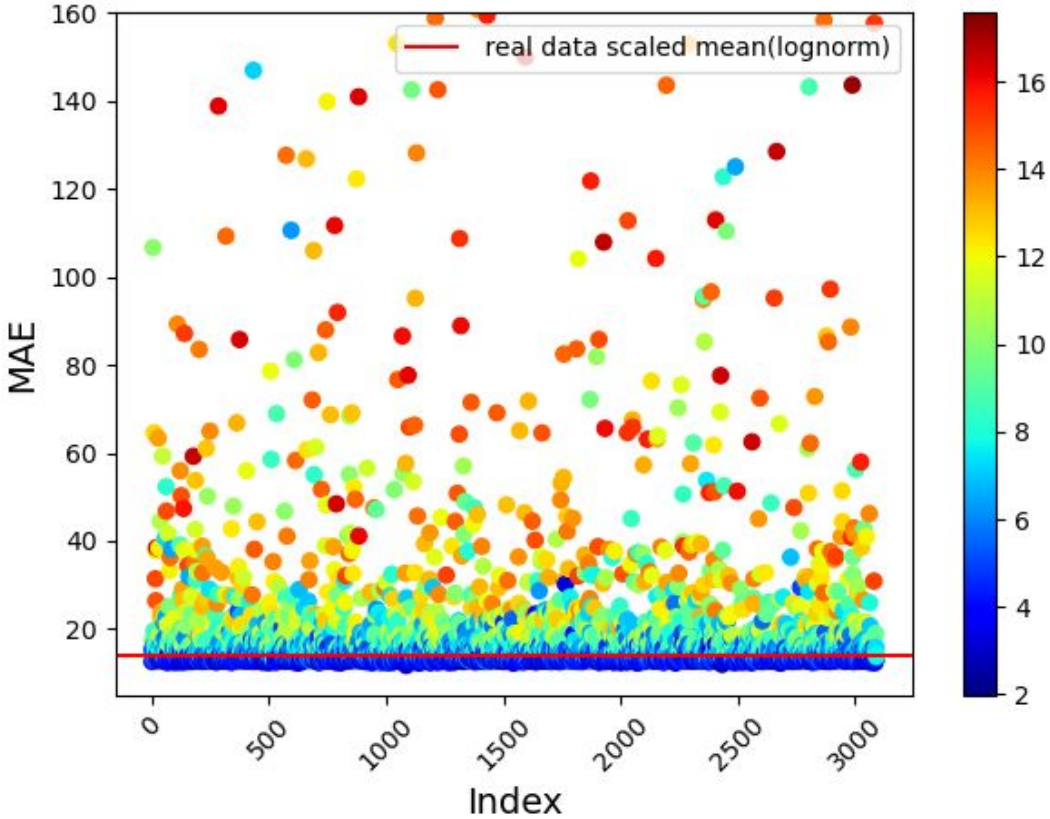
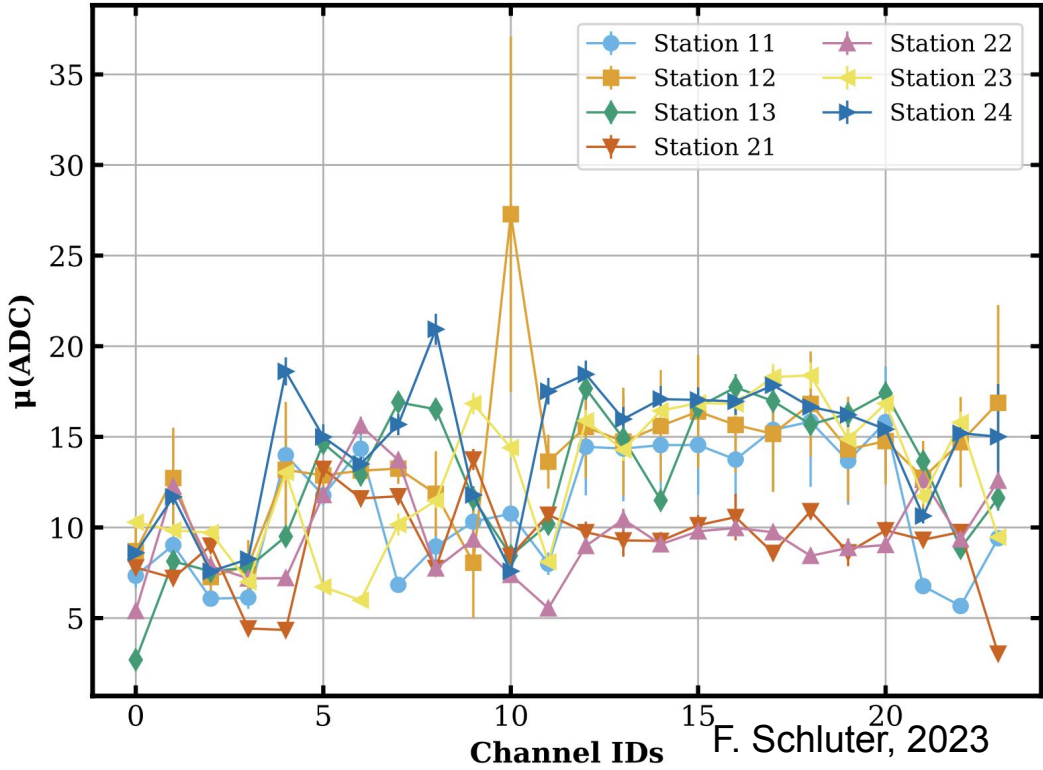
More Solar Bursts

Another strong event



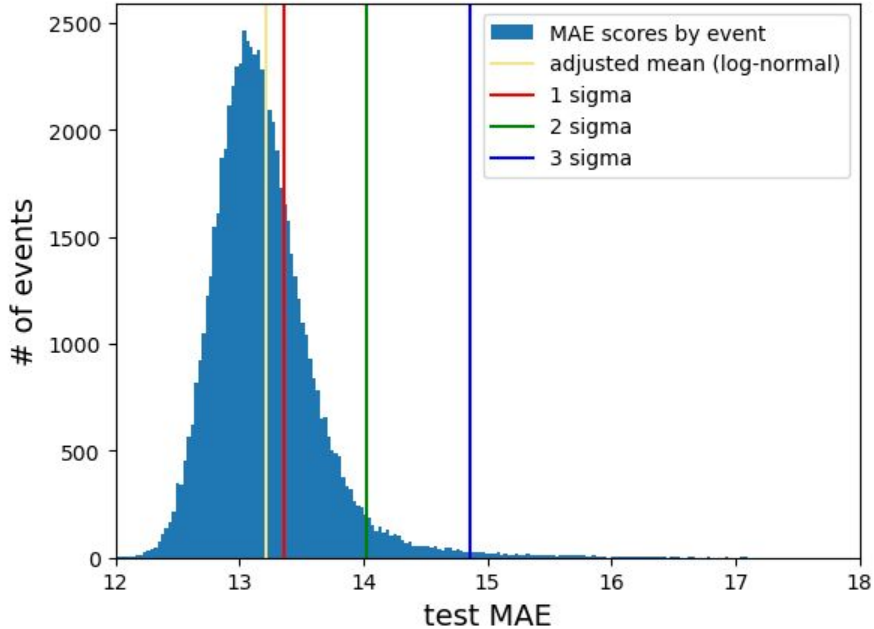
Sim Station 23

Are Neutrinos Anomalous



Prioritization

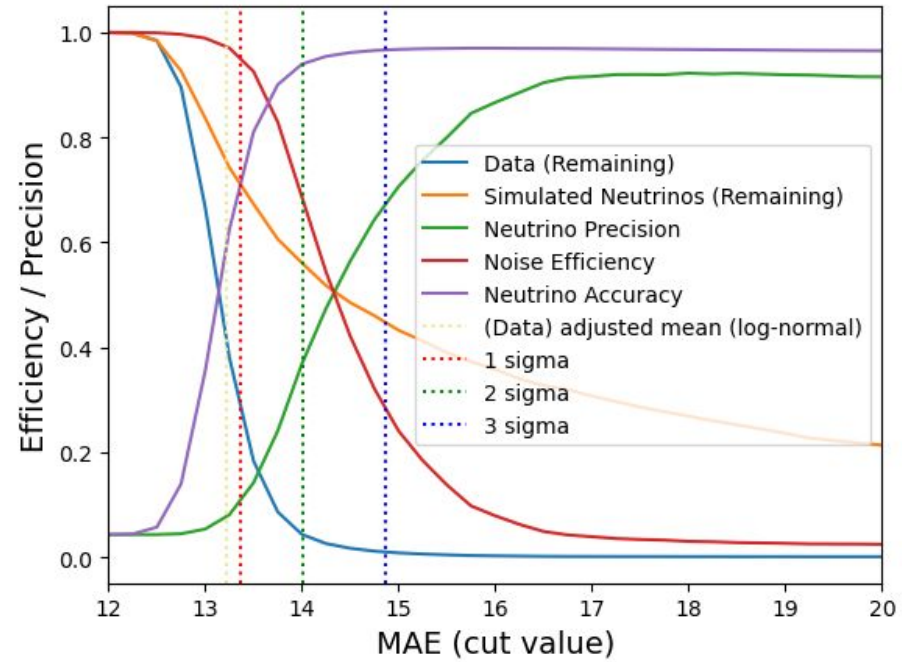
Neutrinos left over in the sample



	# of events	# of simulated Neutrinos	Neutrino Purity
Full Sample	68596	3092	0.043
MAE cut mean	28590	2338	.077
MAE cut aggressive	90	977	0.44

Prioritization

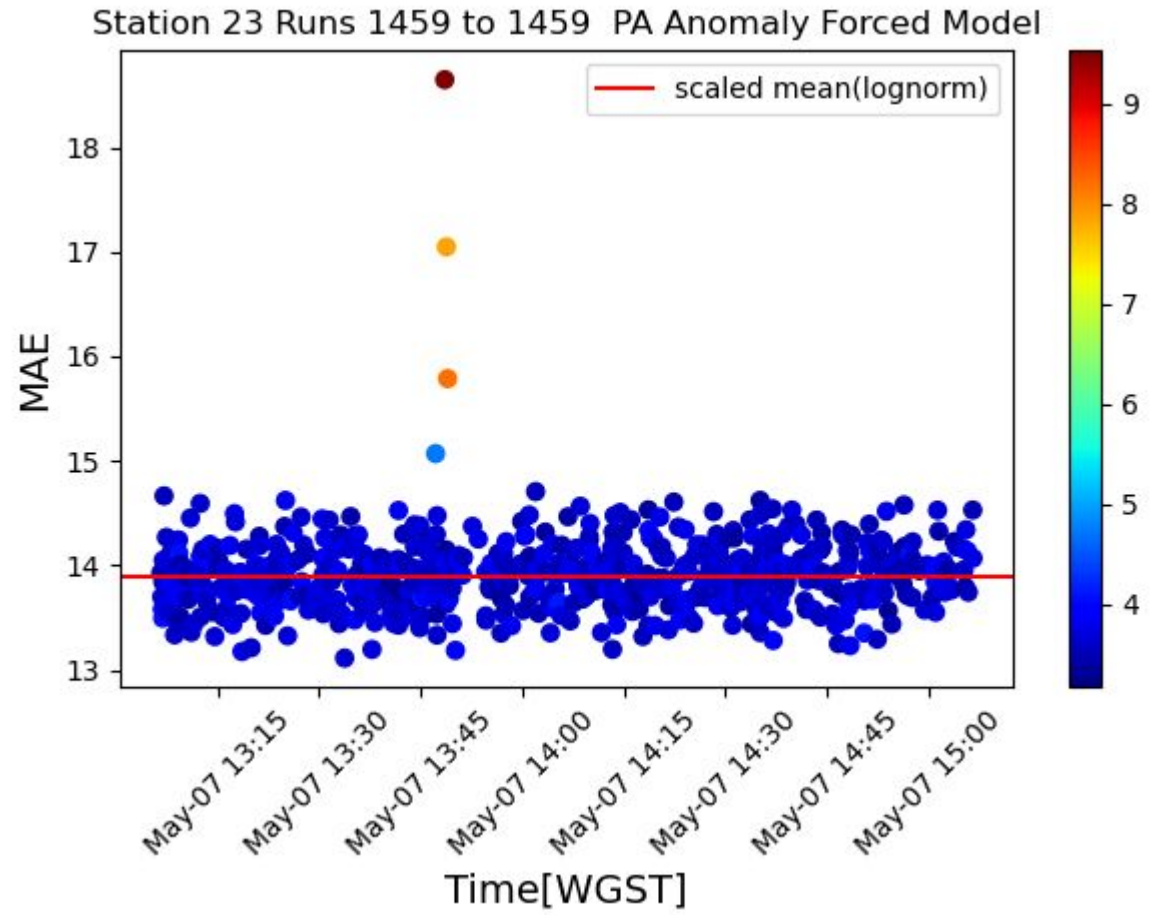
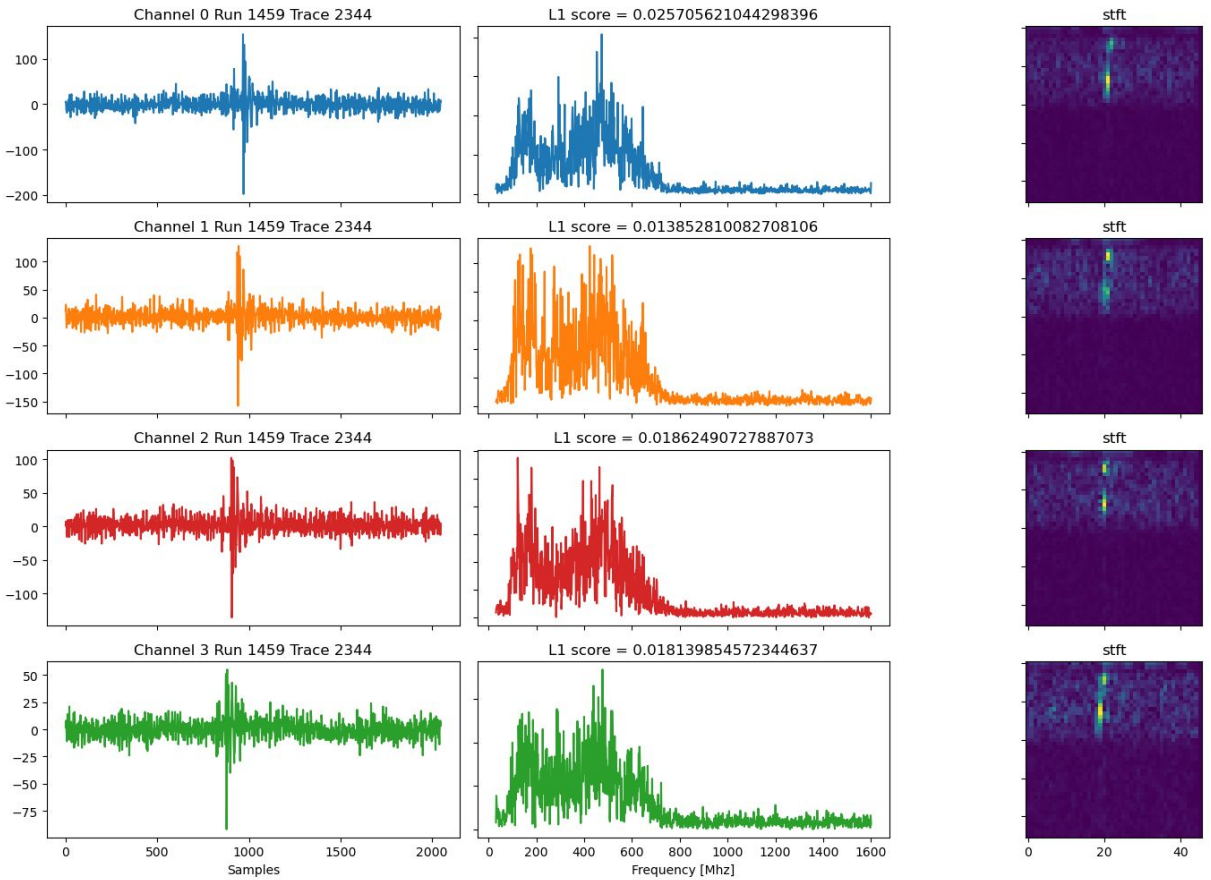
Neutrinos left over in the sample



	# of events	# of simulated Neutrinos	Neutrino Purity
Full Sample	68596	3092	0.043
MAE cut mean	33083	2407	.077
MAE cut aggressive	90	977	0.44

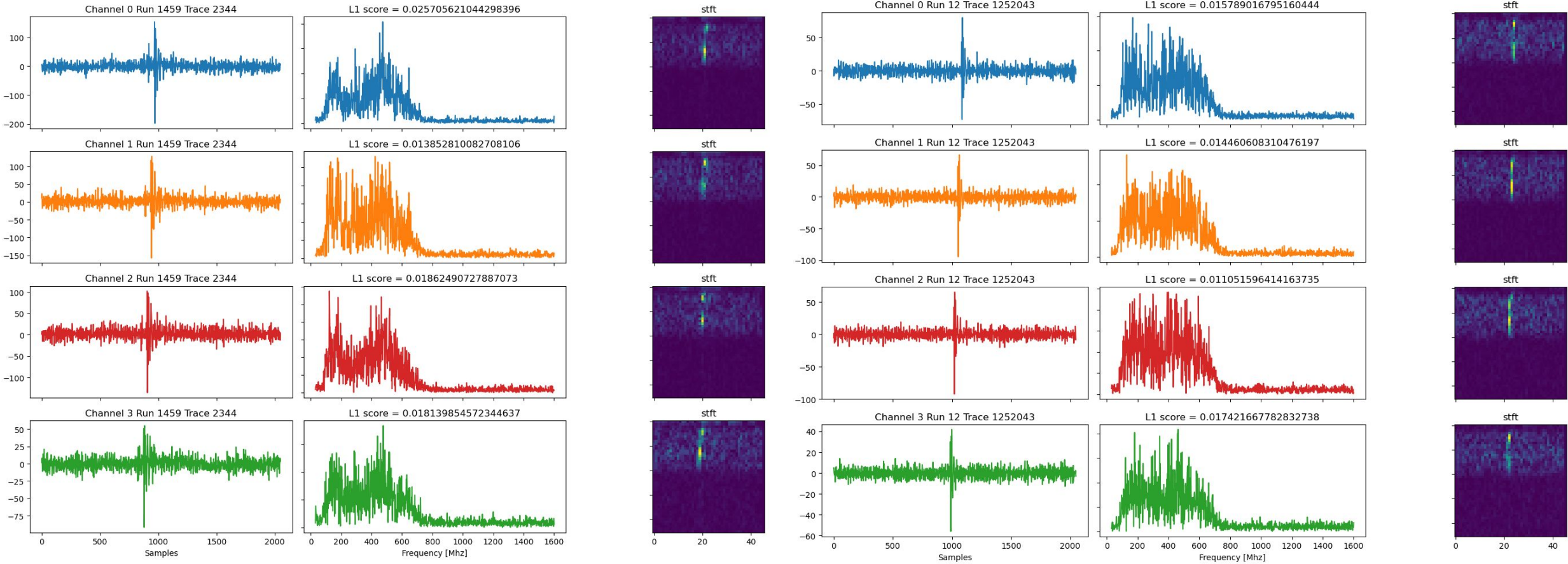
Neutrino-ish Signal

THIS IS NOT A NEUTRINO



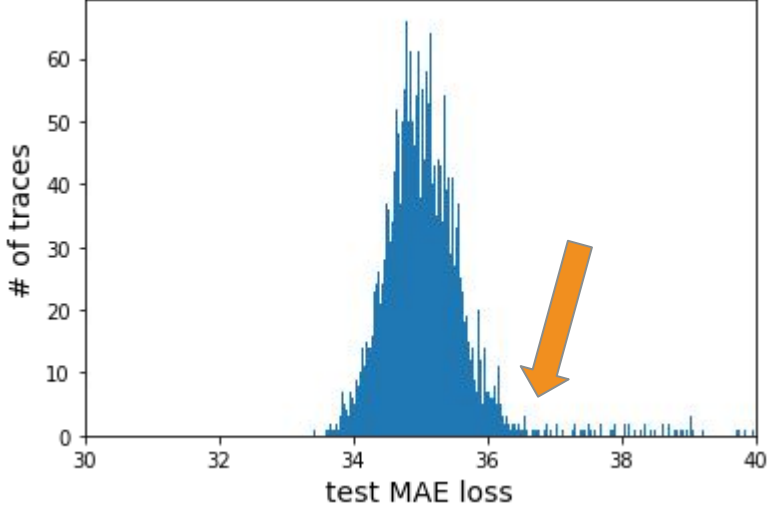
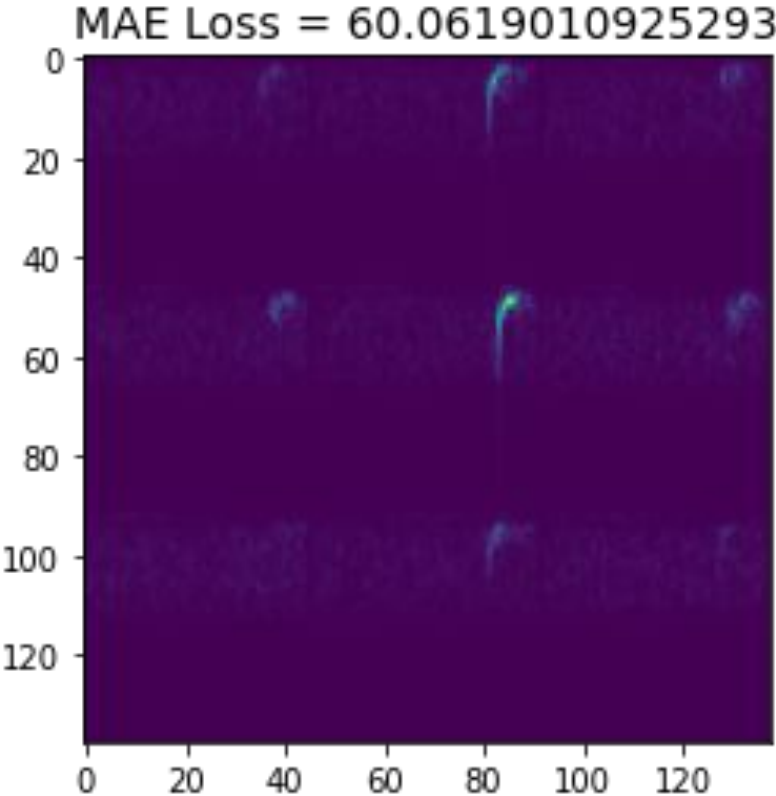
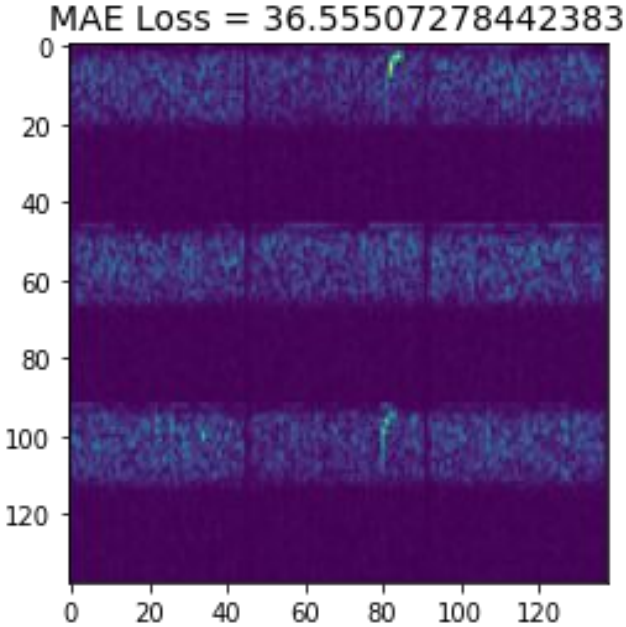
Neutrino-ish Signal

....but it does kind of look like one



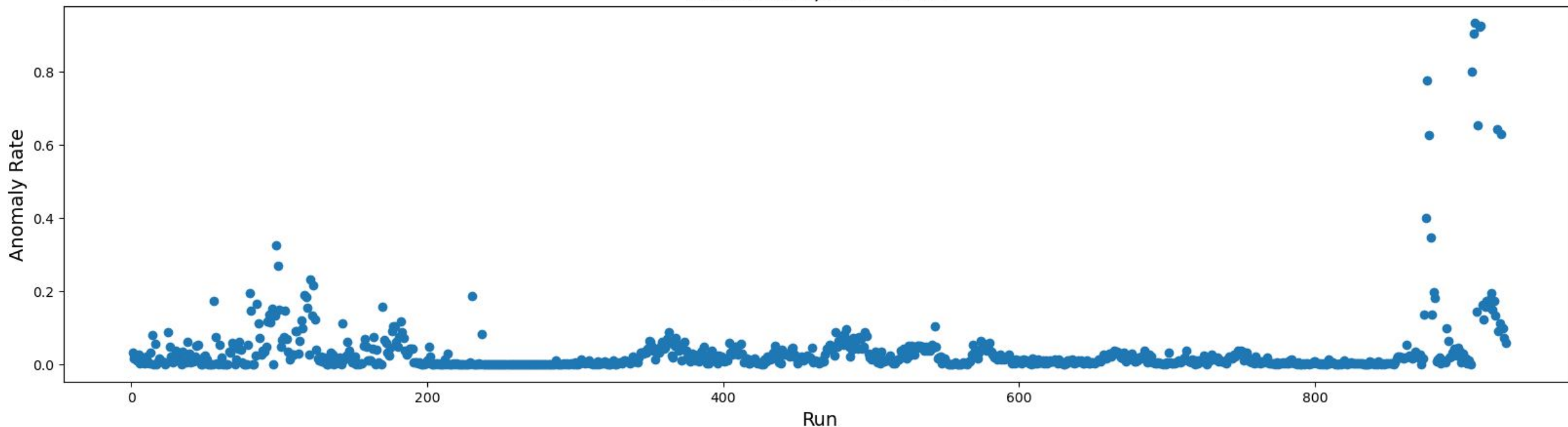
Cosmic Ray Candidates

Surface Candidates (748 run 793(*Aldrin* .69)) (1718 run 1091(*Armstrong* .87))



Summary

Anomalies, station 24



- **Neutrinos (simulated) and Cosmic Ray Candidates are highly anomalous**
- **New noise classes / glitches discovered and mitigated**
- **Active Sun can be used as calibration source, characteristic burst shapes**
- **Only one component of filtering and Multi-Messenger Response**
- **Models are simple – more complex 3D and impulsivity parameters for better efficiency**

Bonus Slides

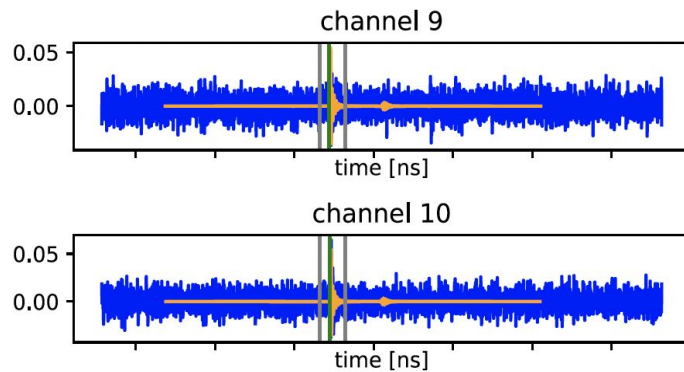
Data Overview

Data rates

- Most triggered events are noise:
 - Expect ~ 1 neutrino / full RNO-G / year
 - Expect $O(1)$ cosmic ray events in surface component / day
 - other physics backgrounds: Sun, Galaxy
 - Thermal noise fluctuations
 - Anthropogenic and hardware induced noise (intermittent, up to ~ 10 Hz)

Rate

Ability to simulate

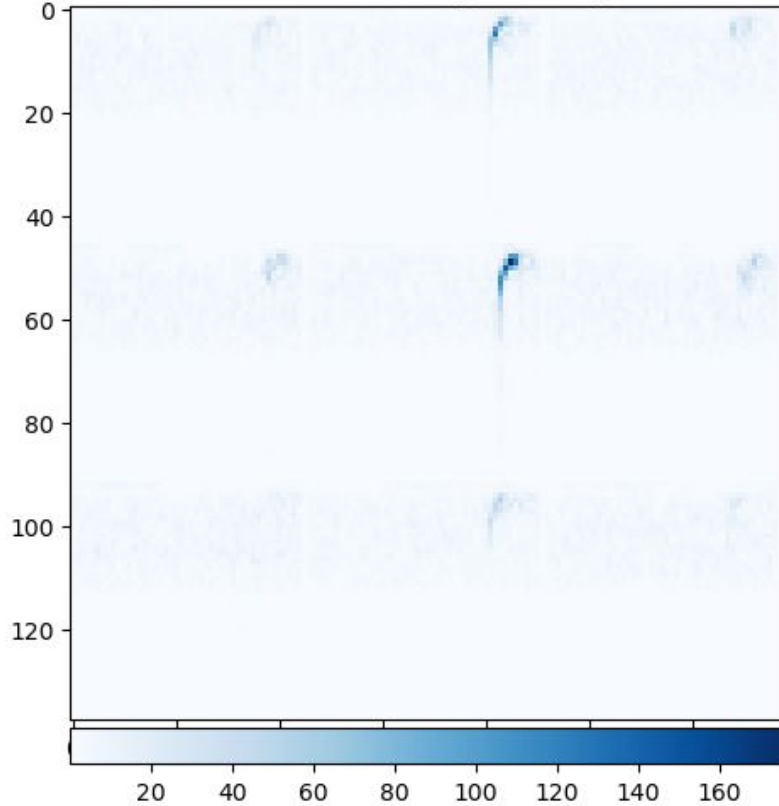


- Signal pulses near threshold
- Characteristic polarized bipolar pulses convolved with hardware response

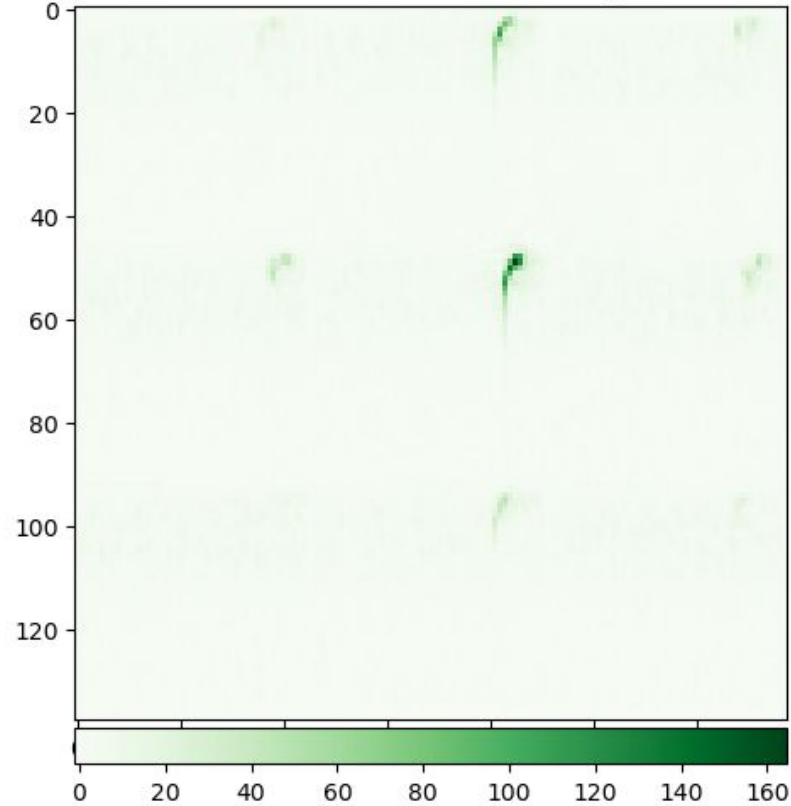
MAE calculation

Sum of mean pixel difference per time slice

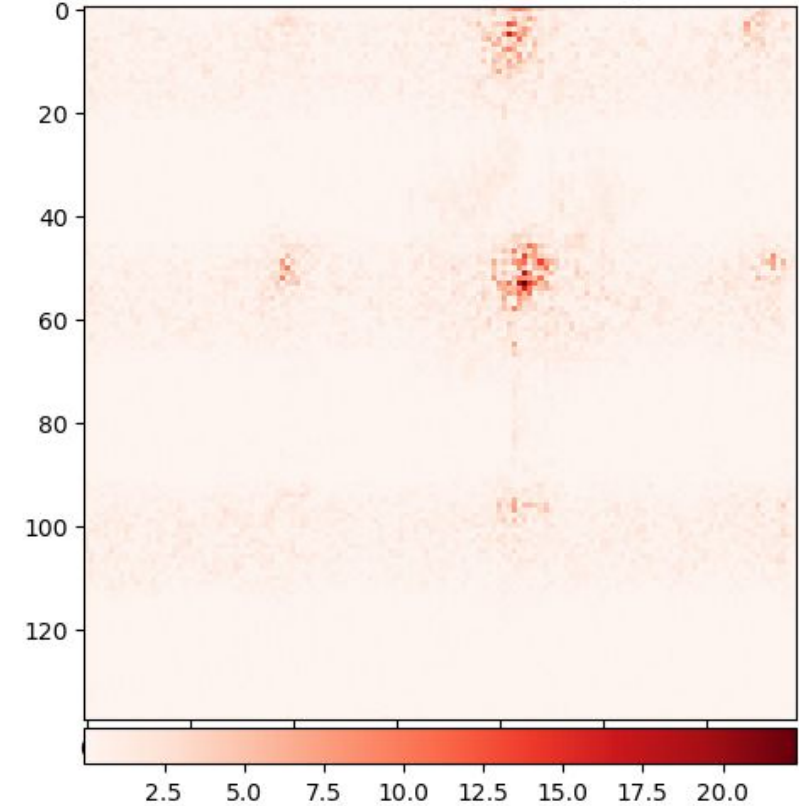
ARMSTRONG Original Image



ARMSTRONG Predicted Image

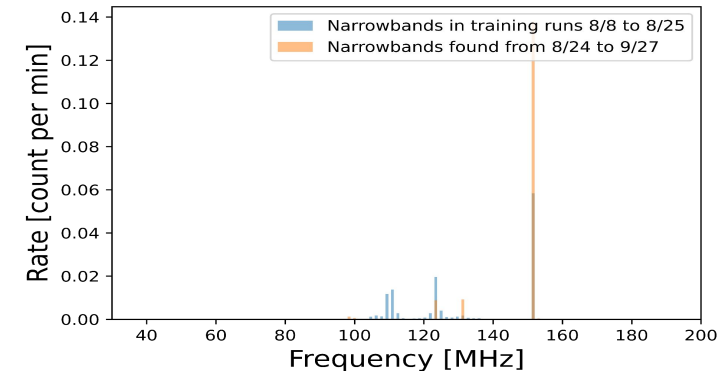
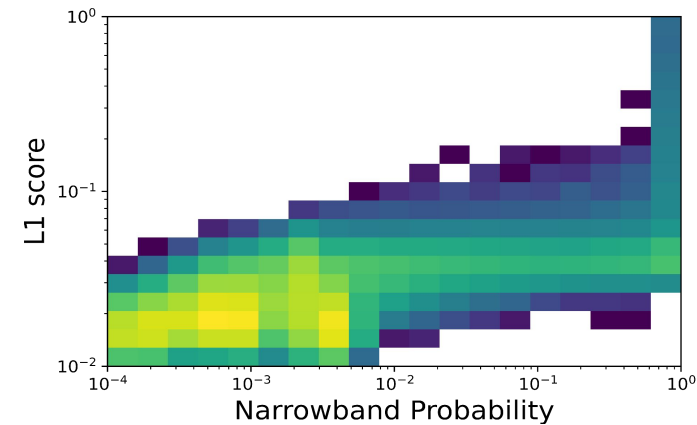
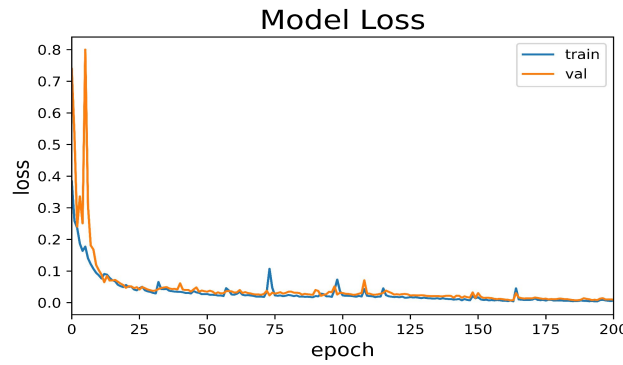
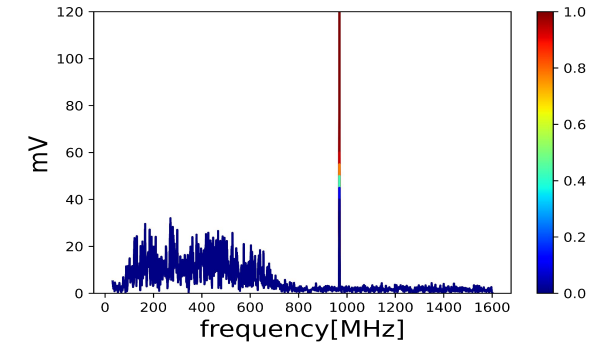
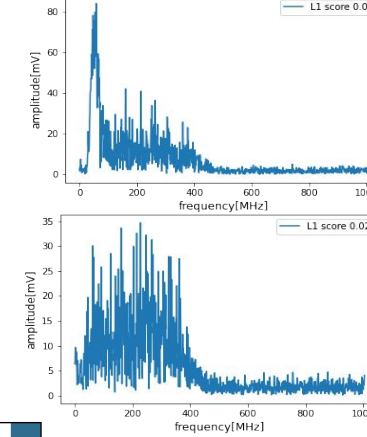
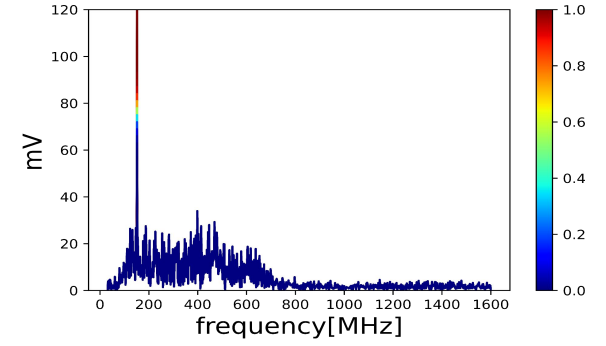
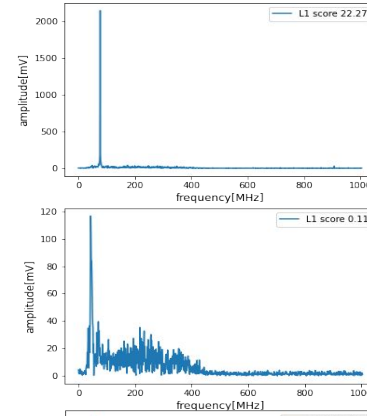
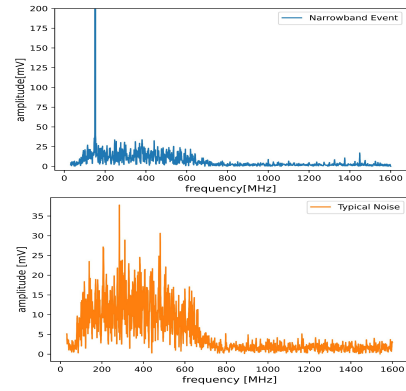
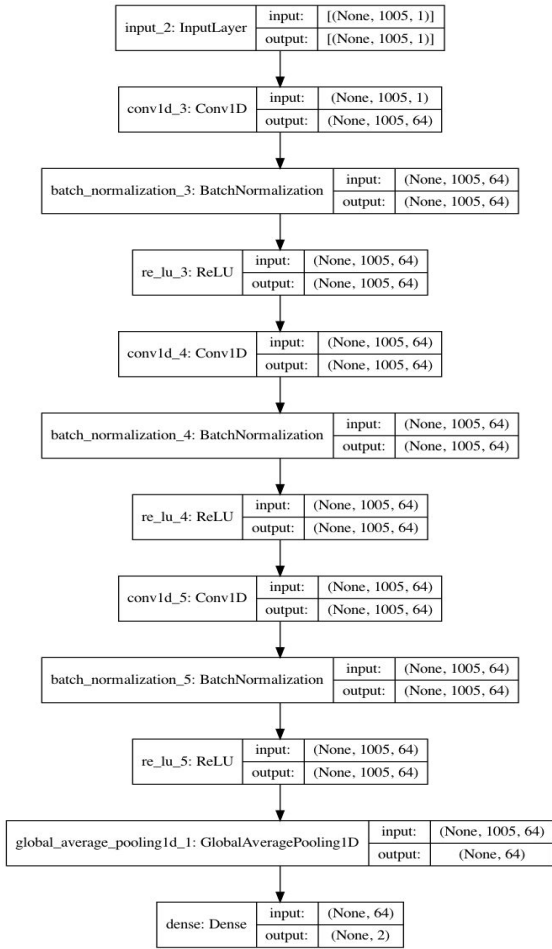


ARMSTRONG Difference Image



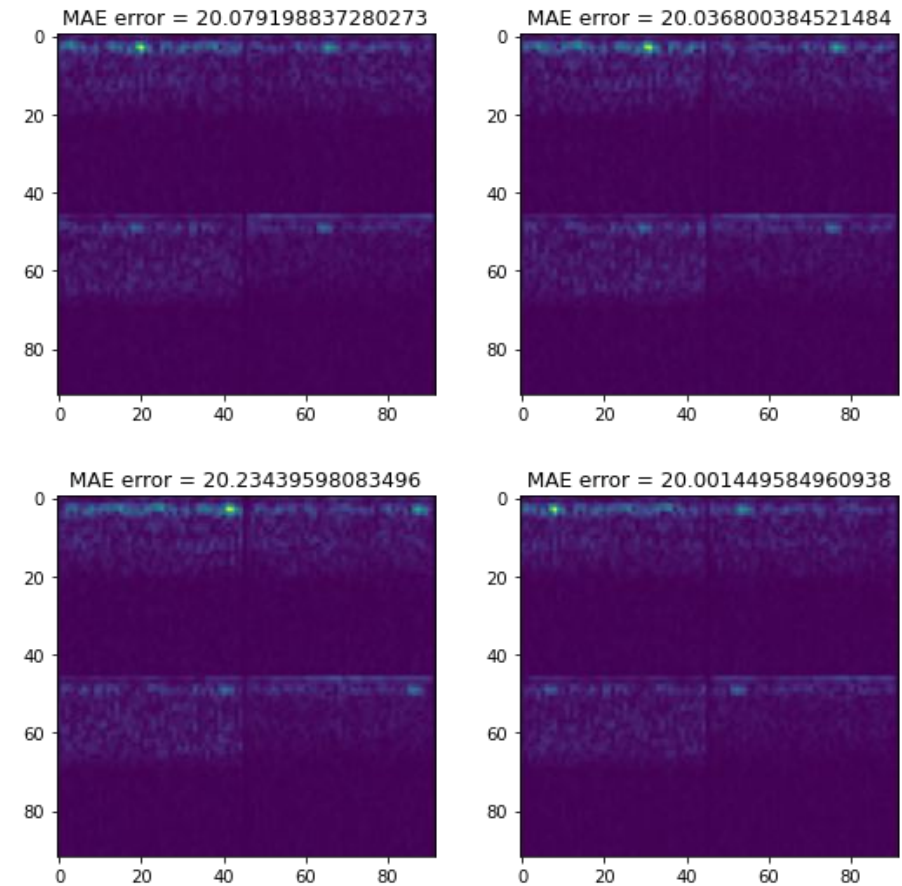
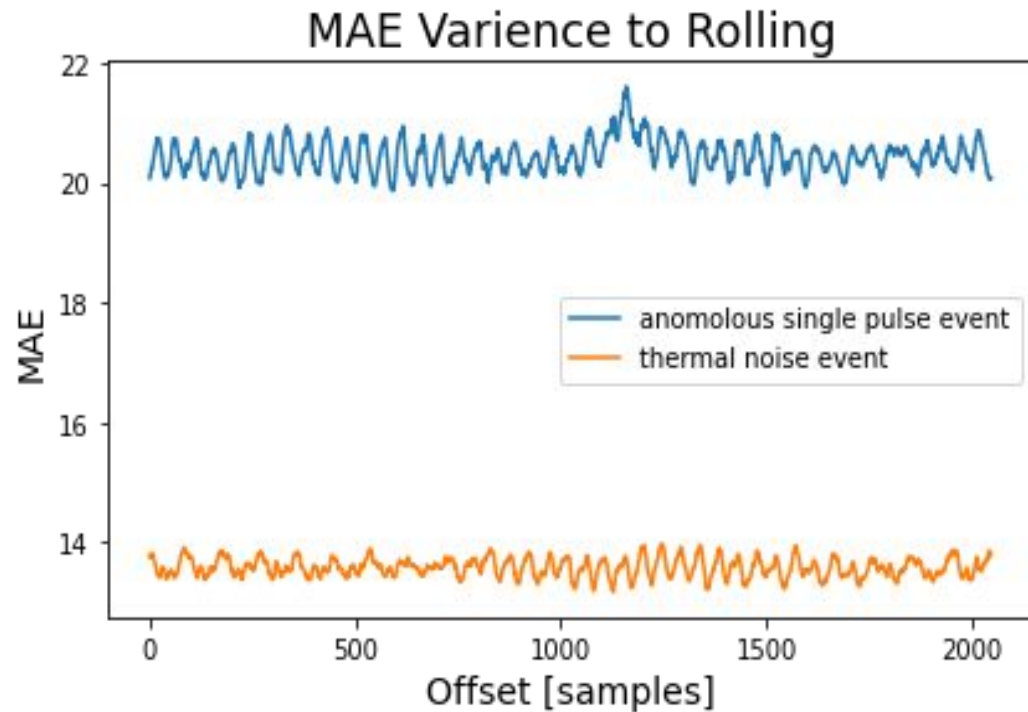
CW Noise

Simple CNN Classification



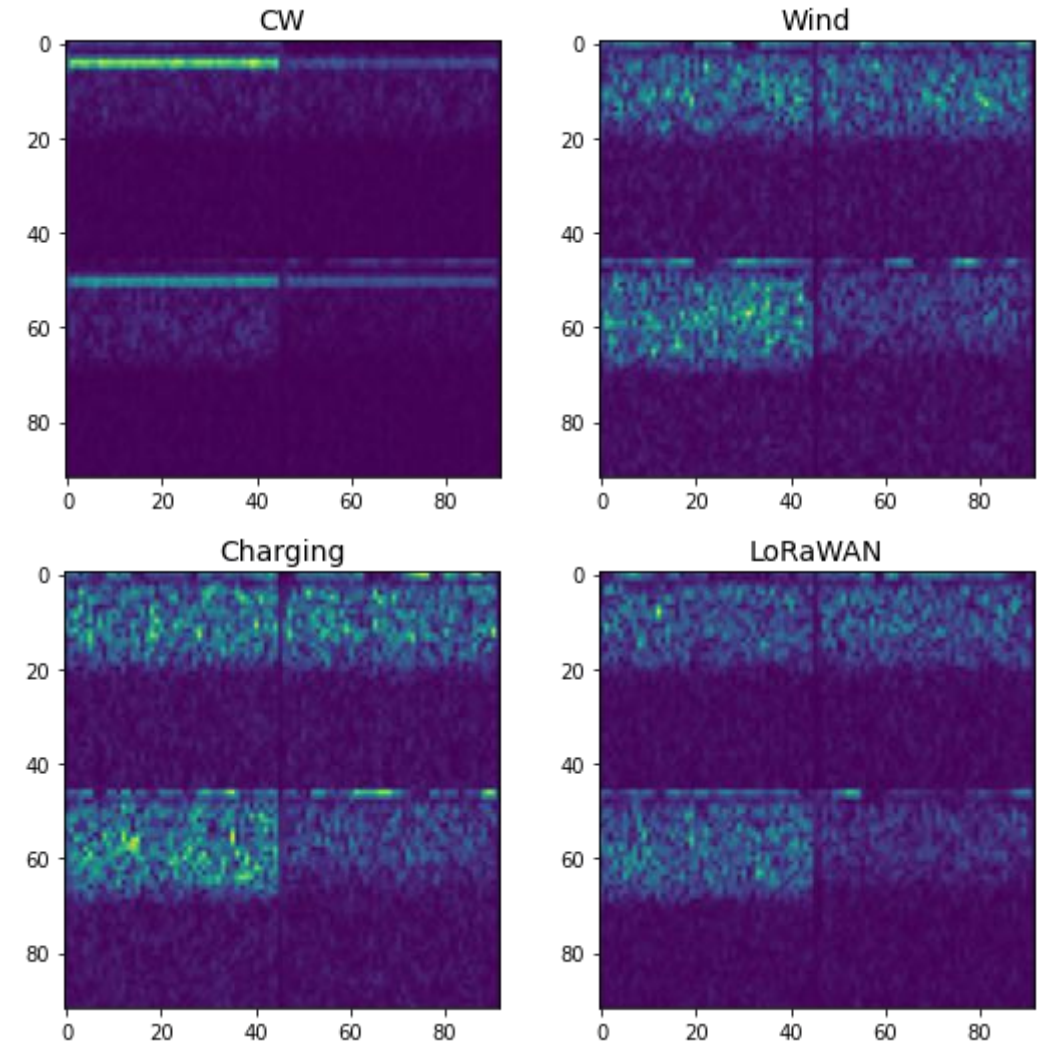
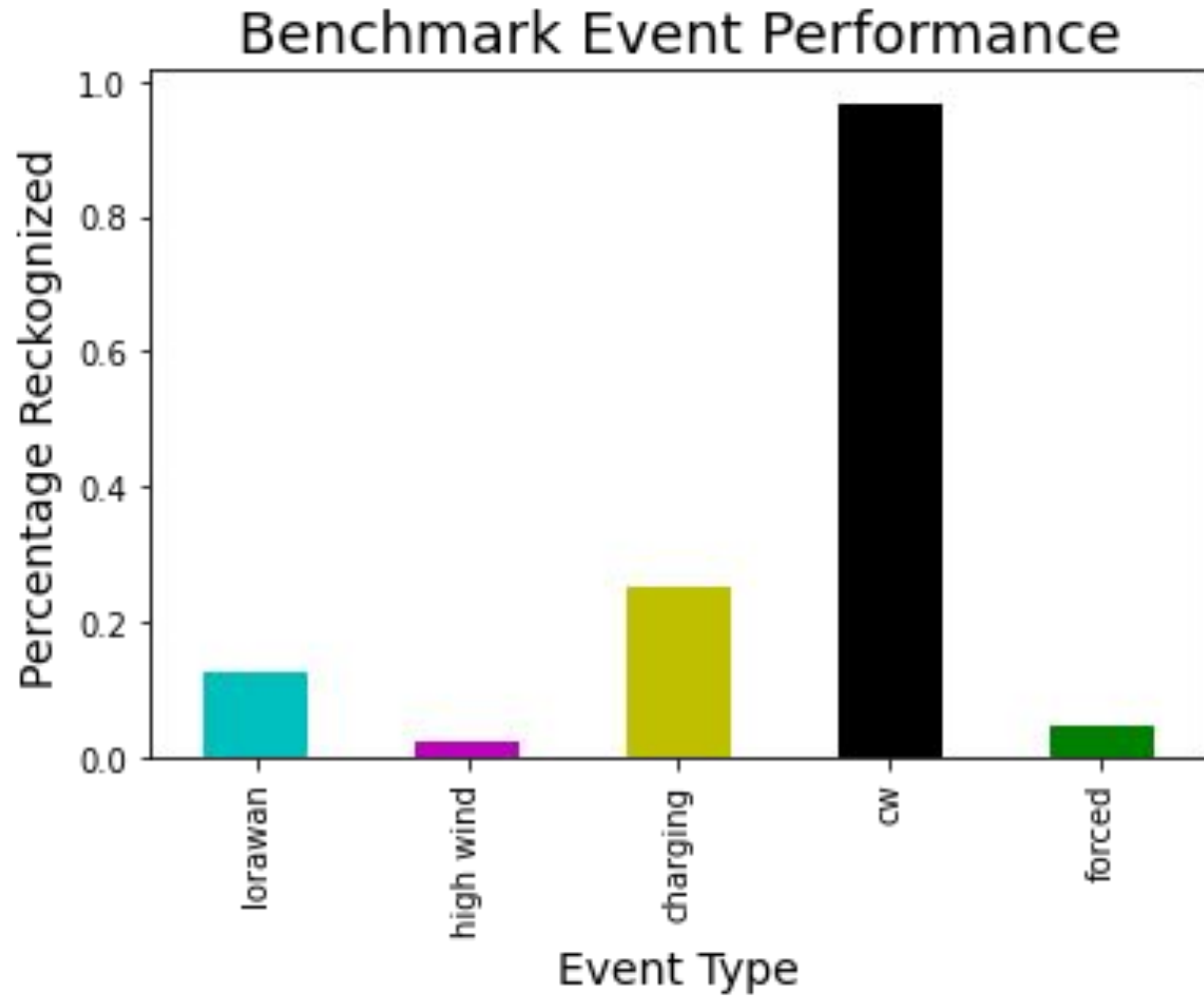
Anomaly Detection Update

Anomalous Event Rolling Invariance



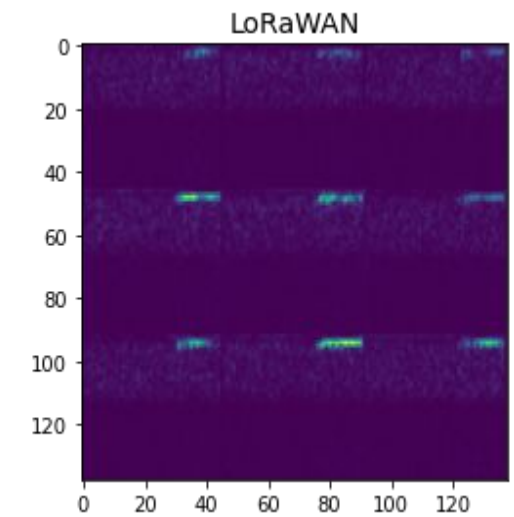
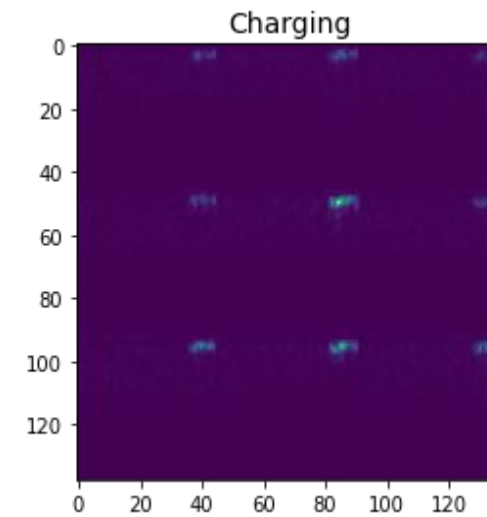
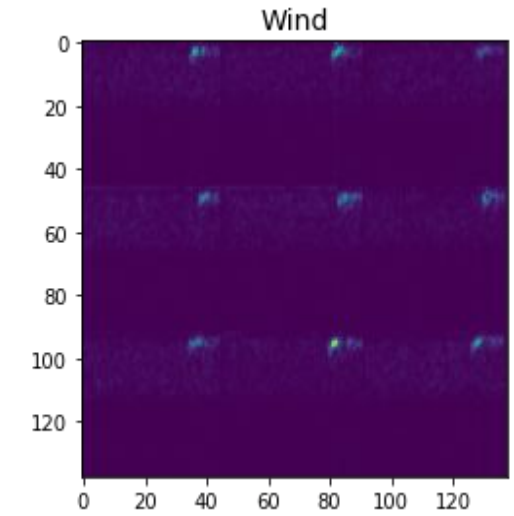
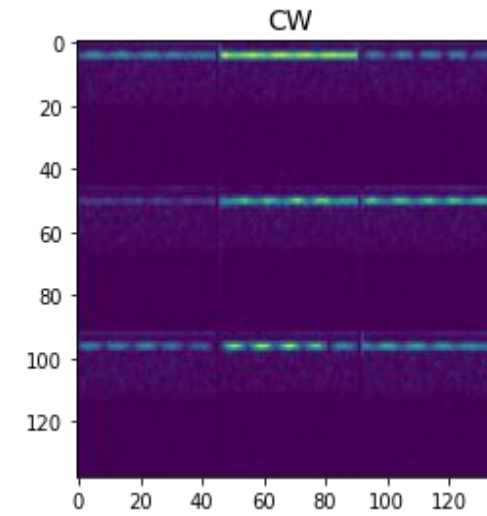
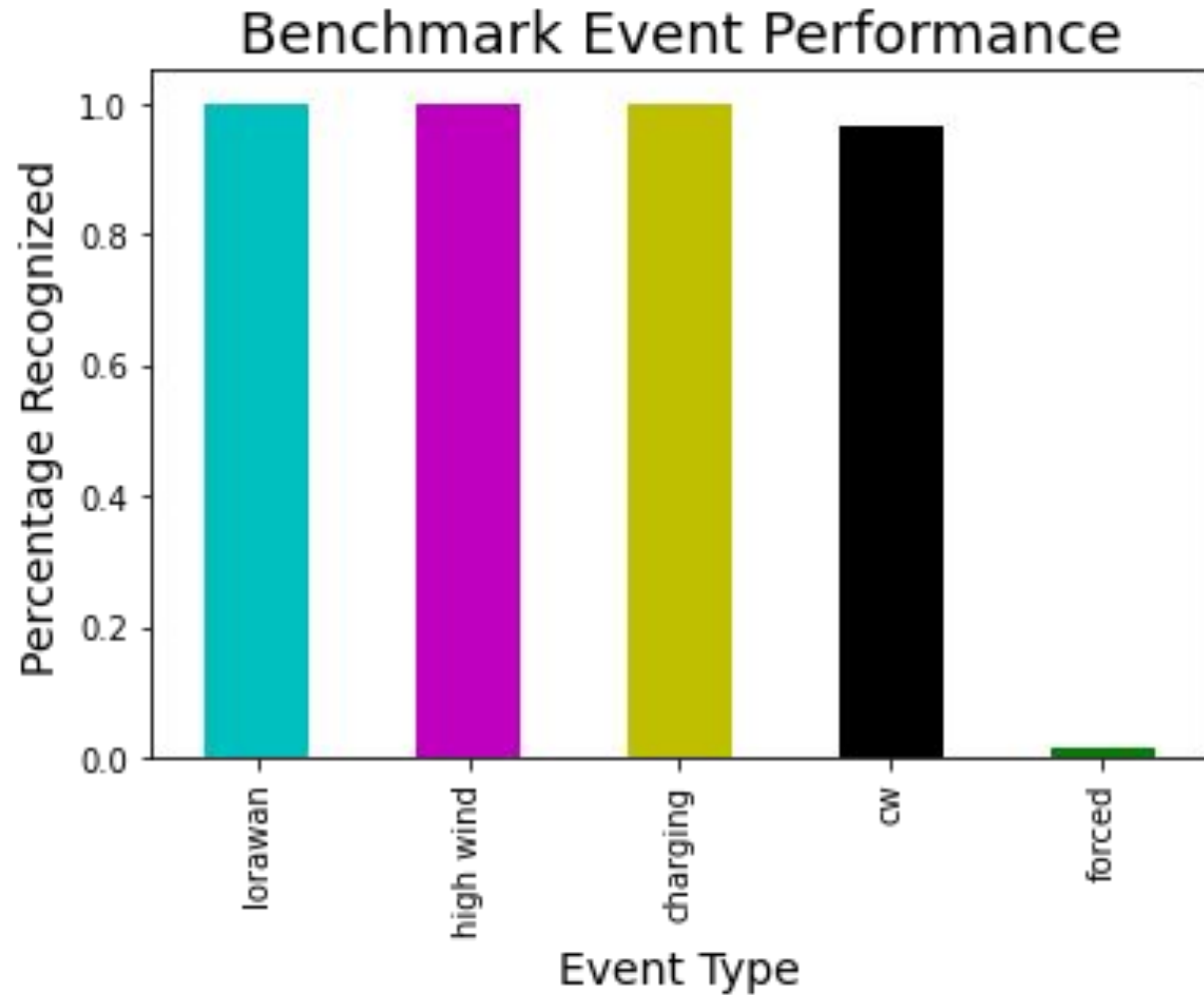
Anomaly Detection

Deep Benchmarks



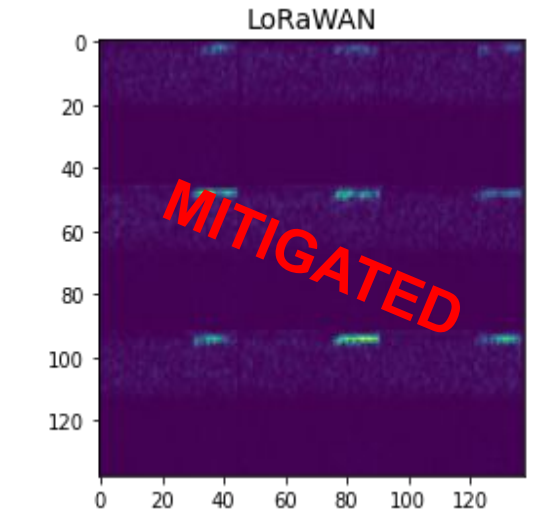
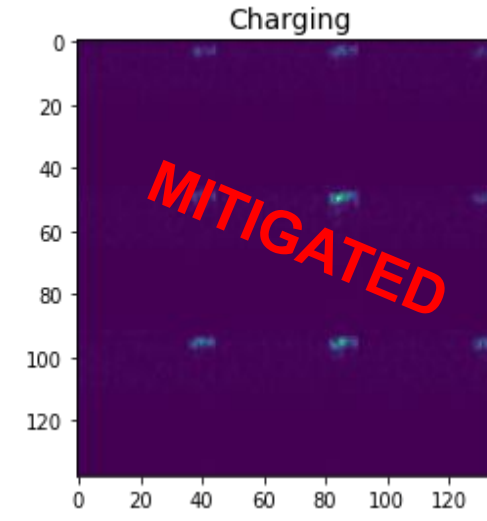
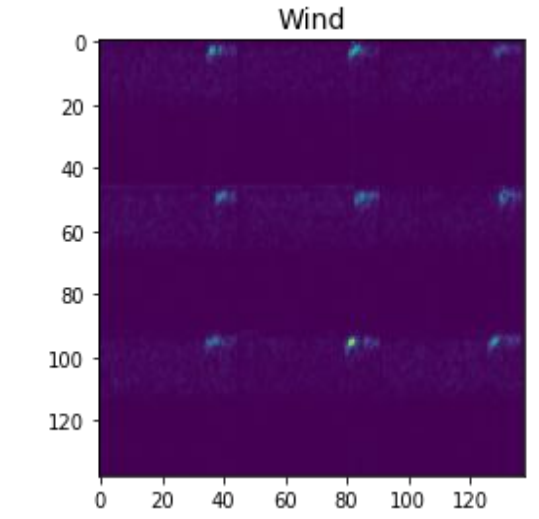
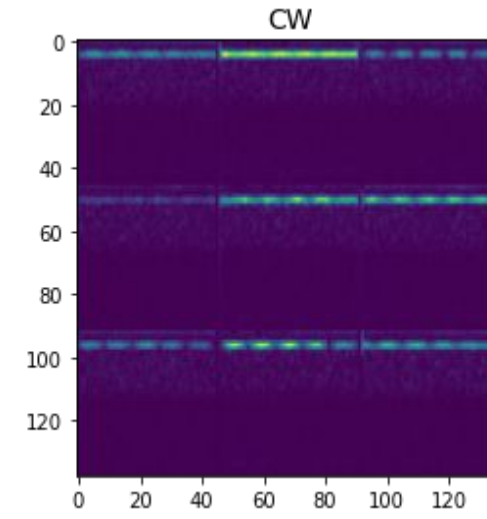
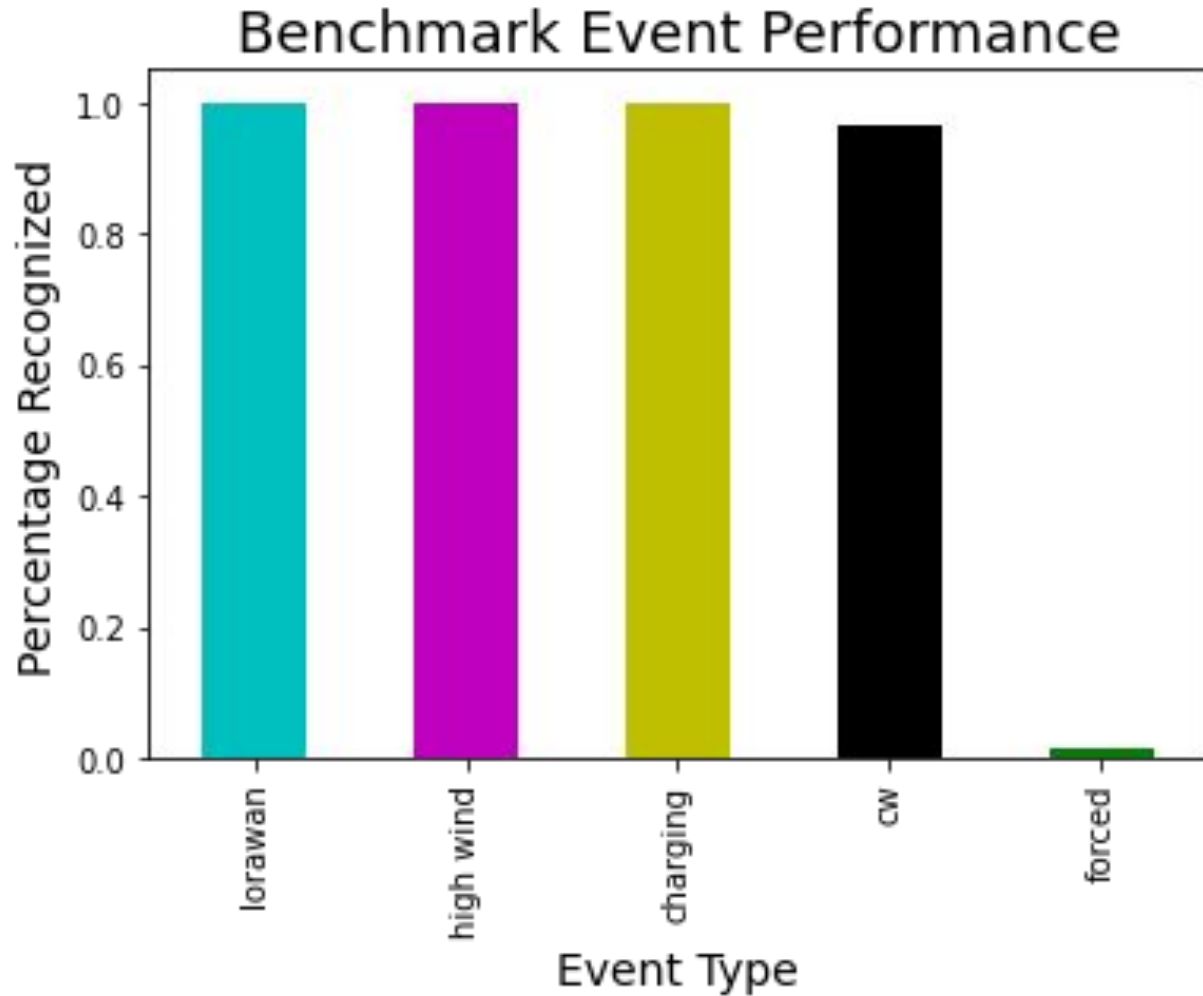
Anomaly Detection

Surface Noise Benchmarks



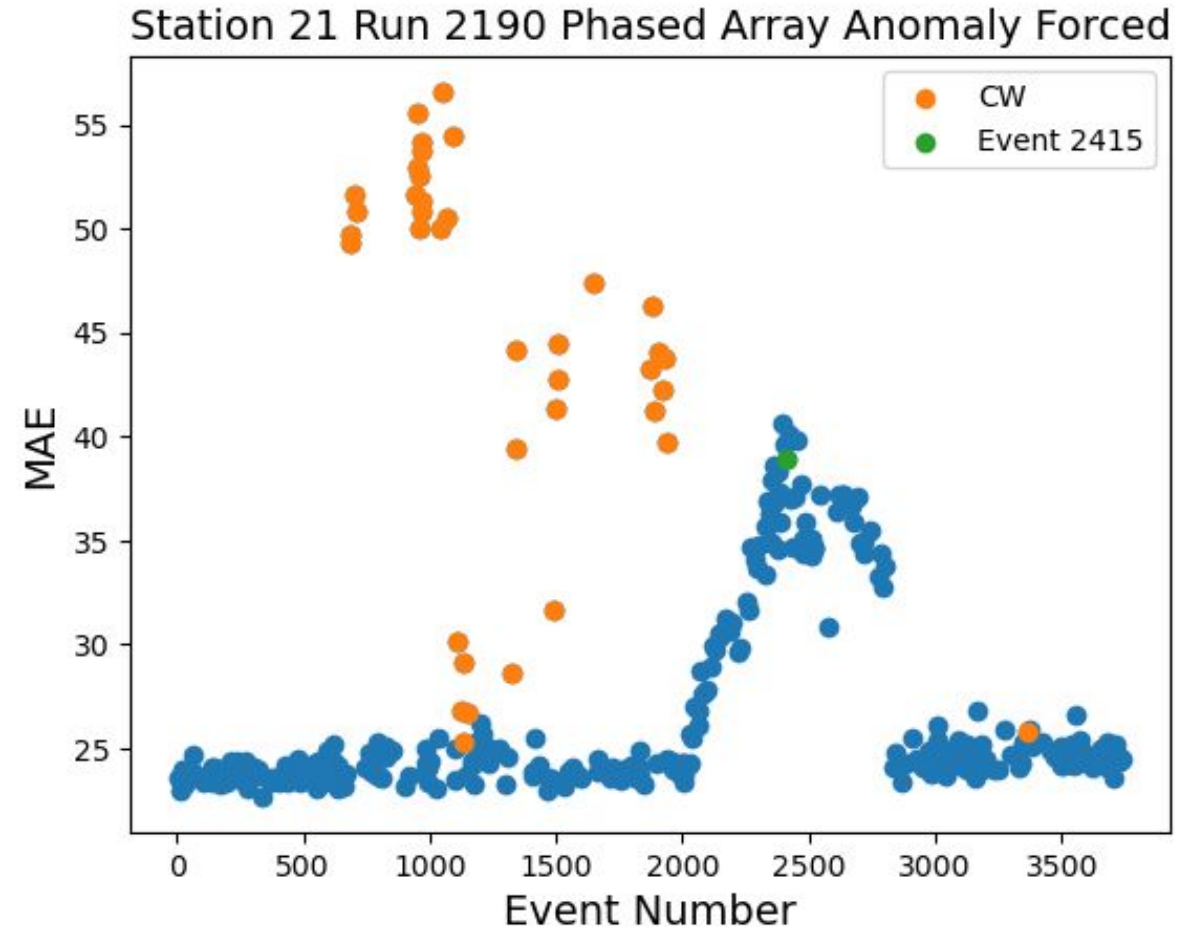
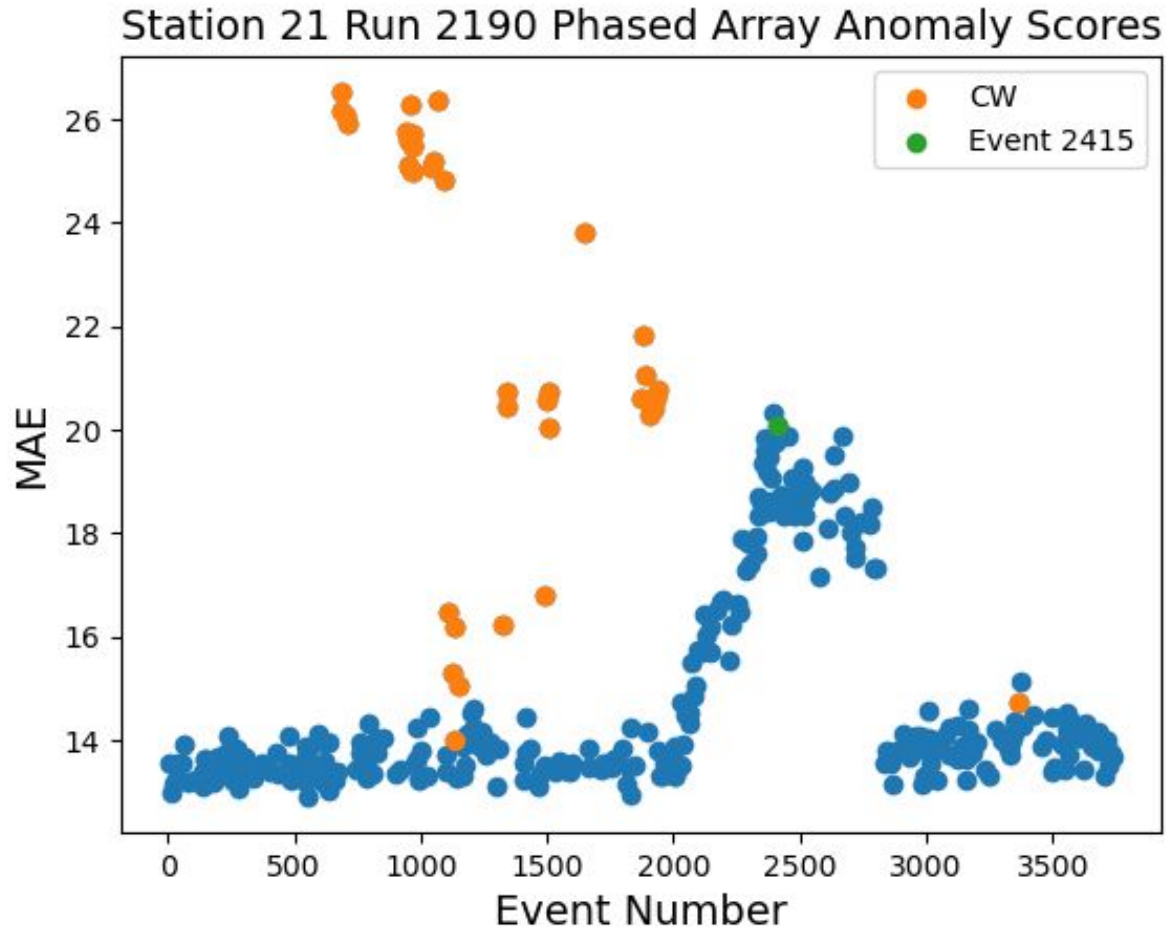
Anomaly Detection

Surface Noise Benchmarks

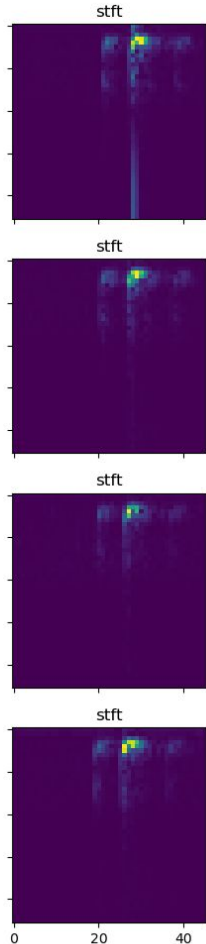
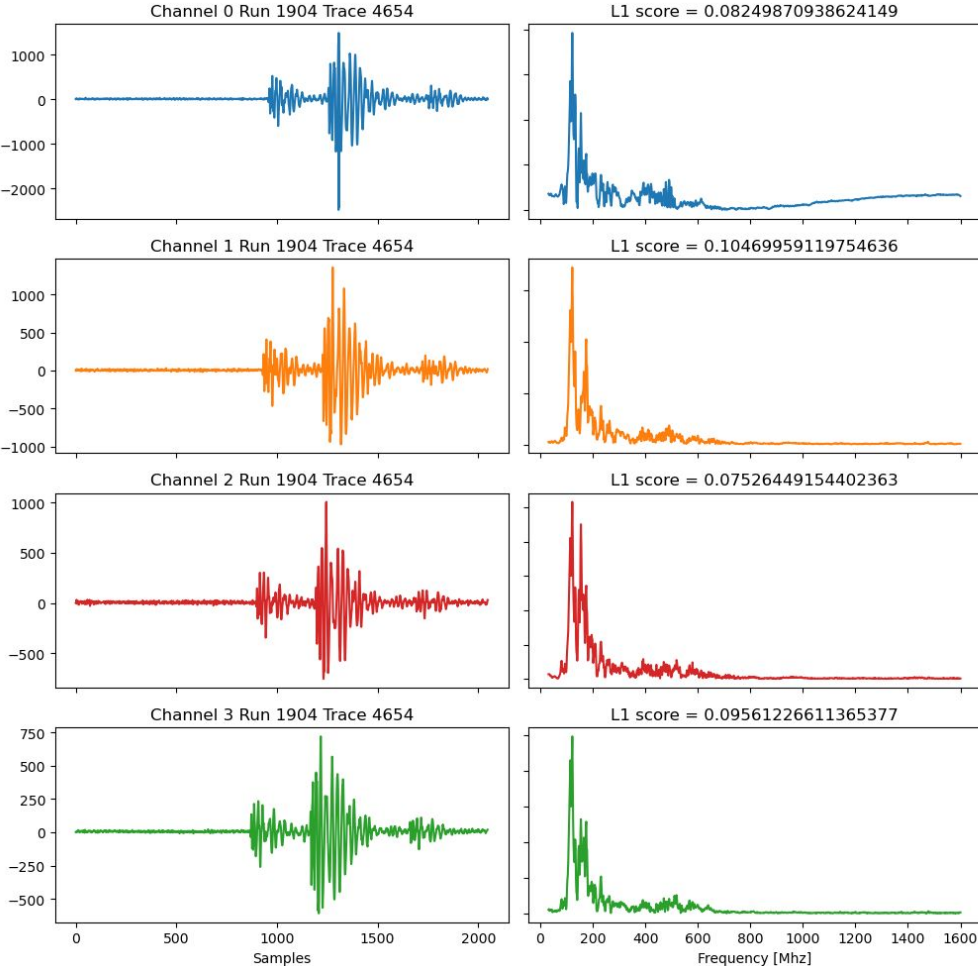
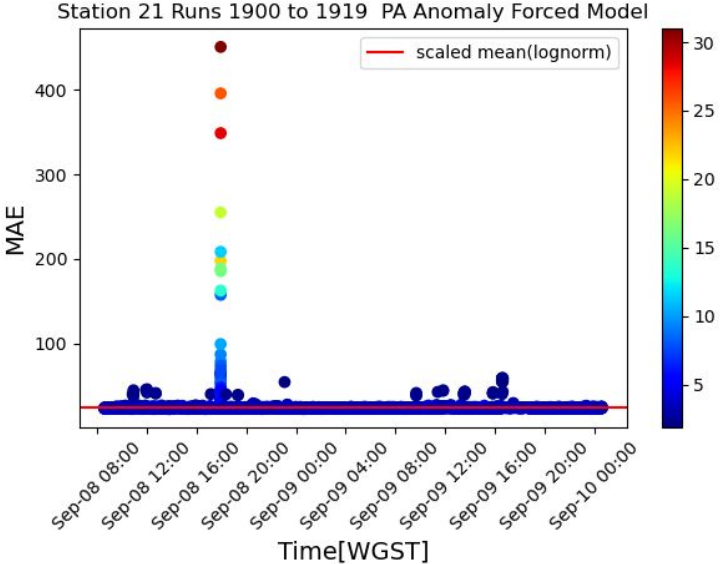


“Anomalous Low Threshold Event”

Forced triggers vs “Quiet Periods”



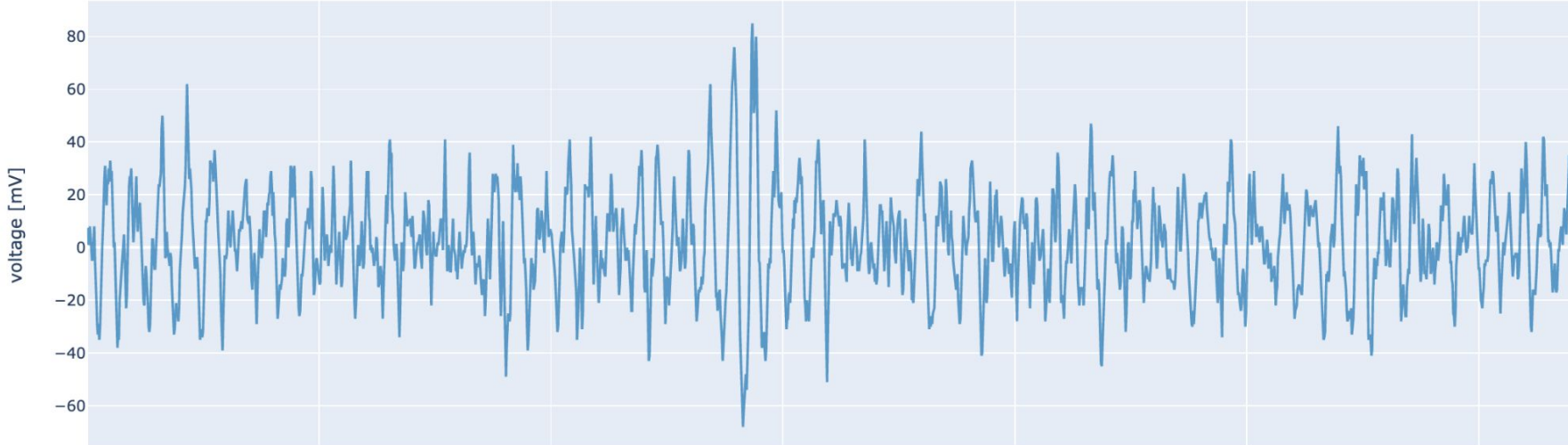
Snowmobiles / Ice Sat Traverse



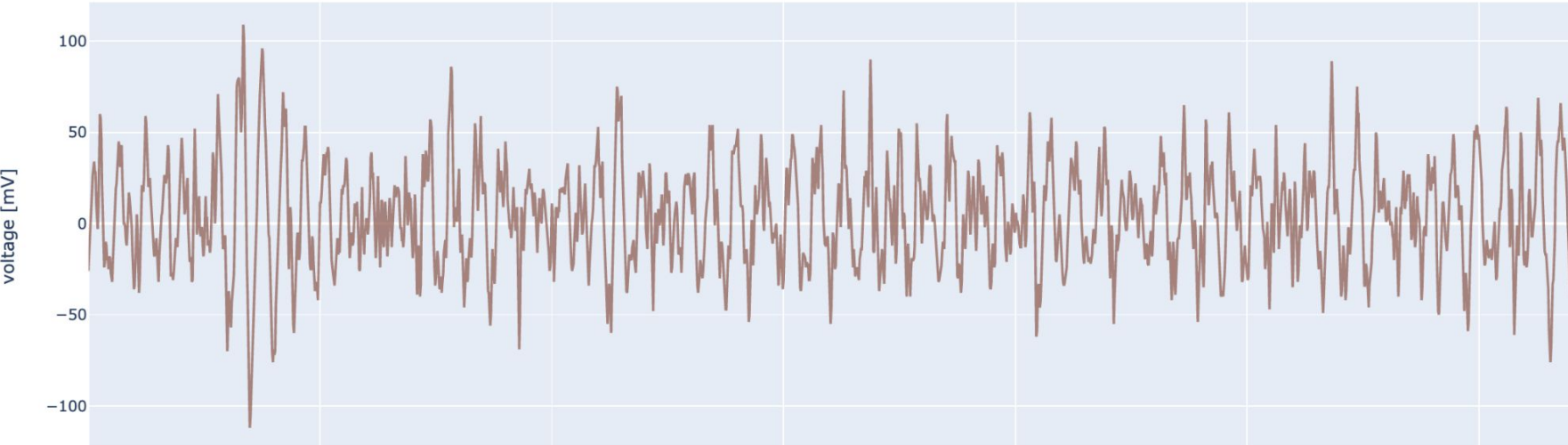
“Anomalous Low Threshold Event”

Reconstruction / Coming from above

Channel 0

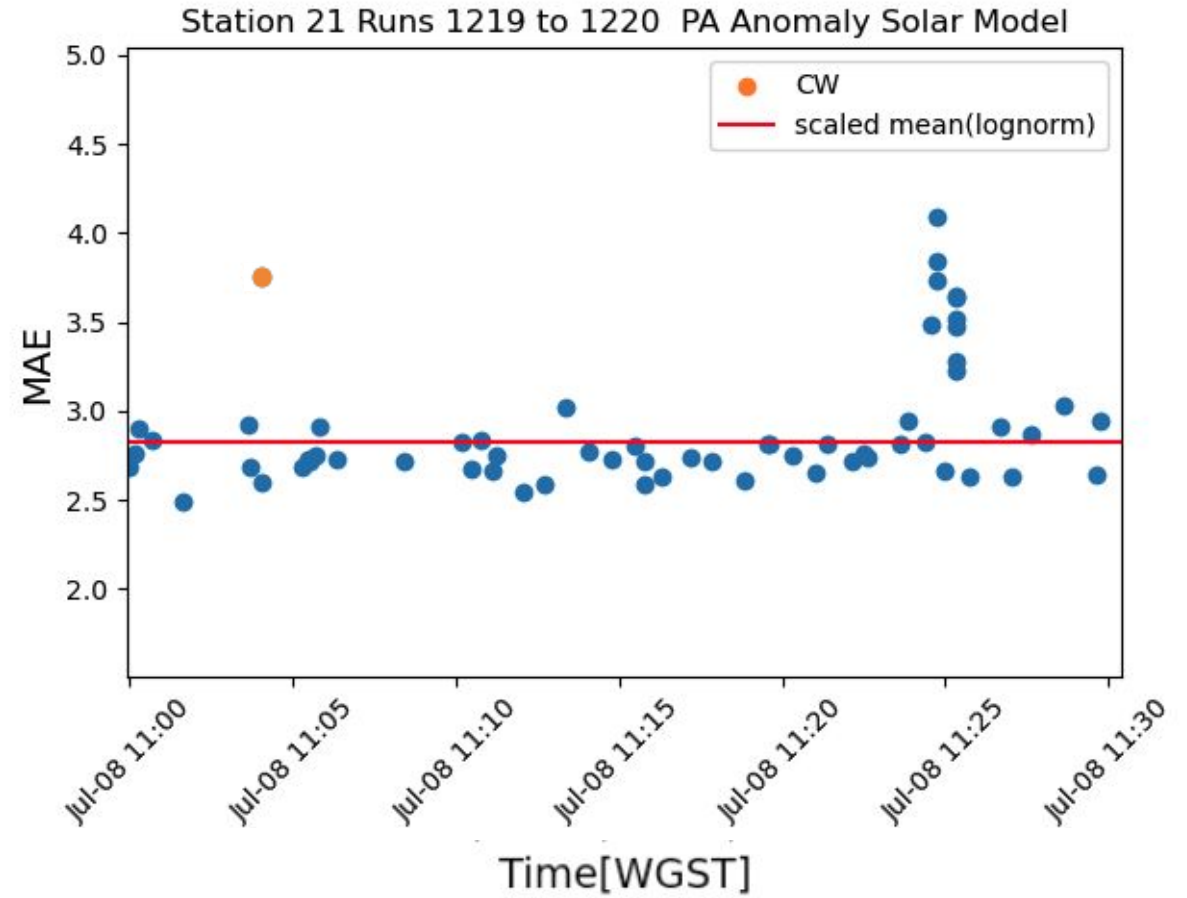
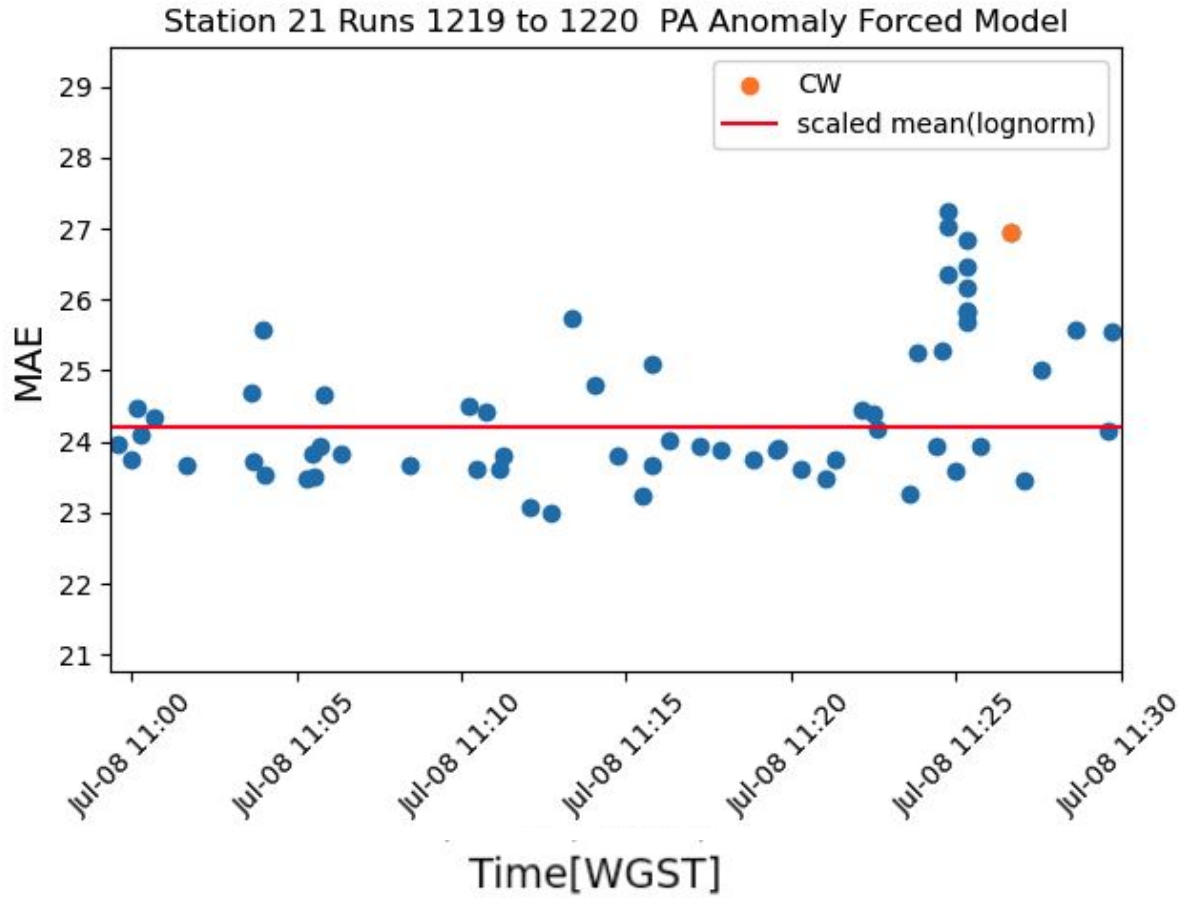


Channel 5



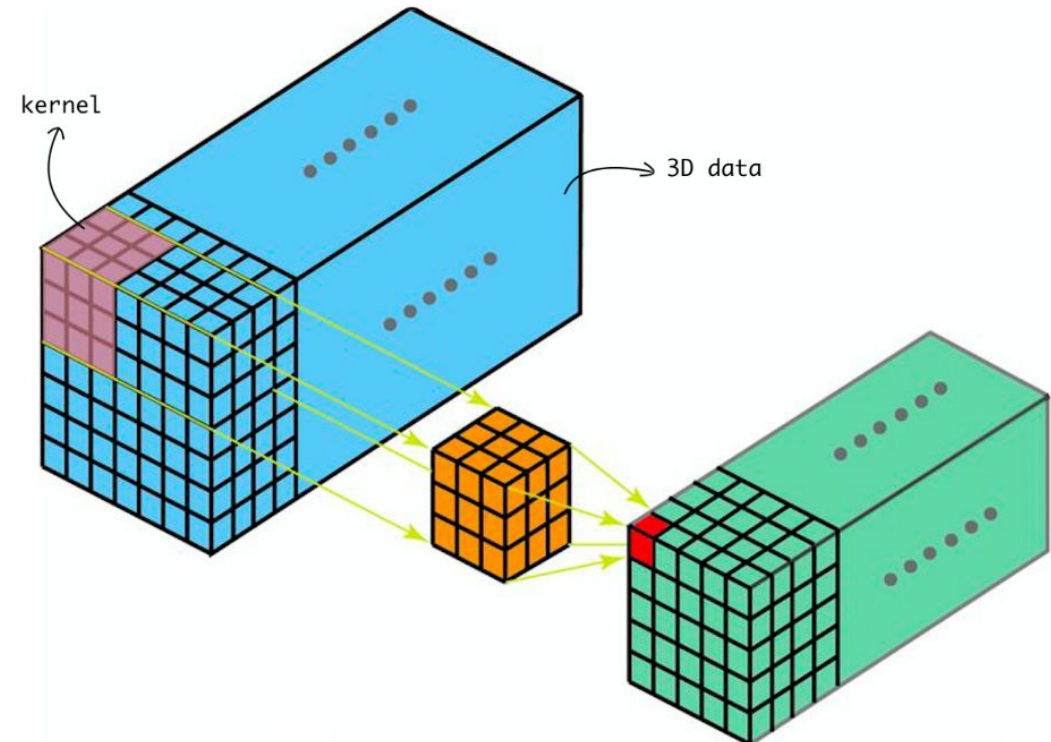
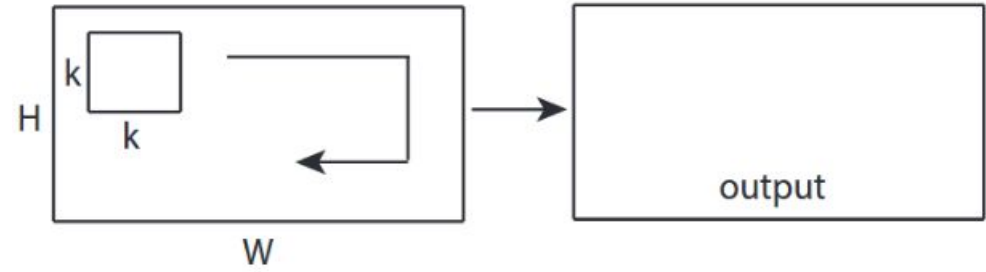
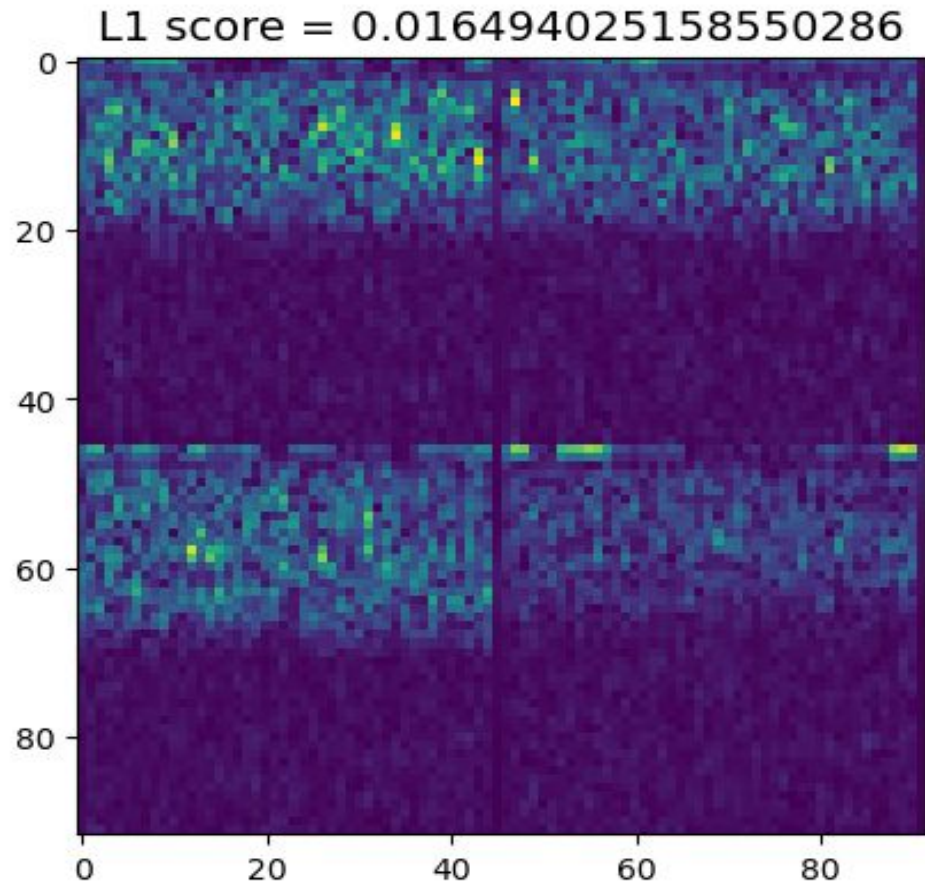
“Solar Model”

Butterworth filter 25 - 250 MHz



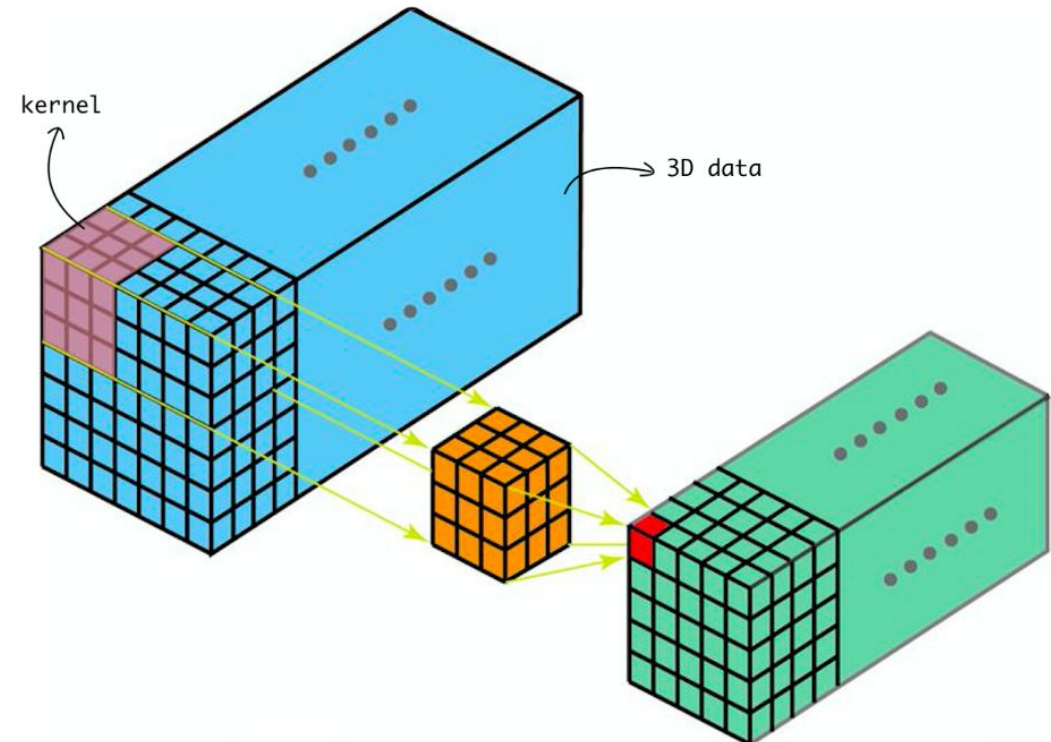
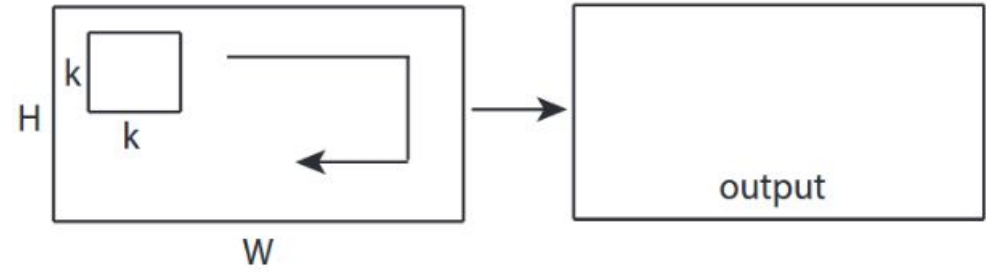
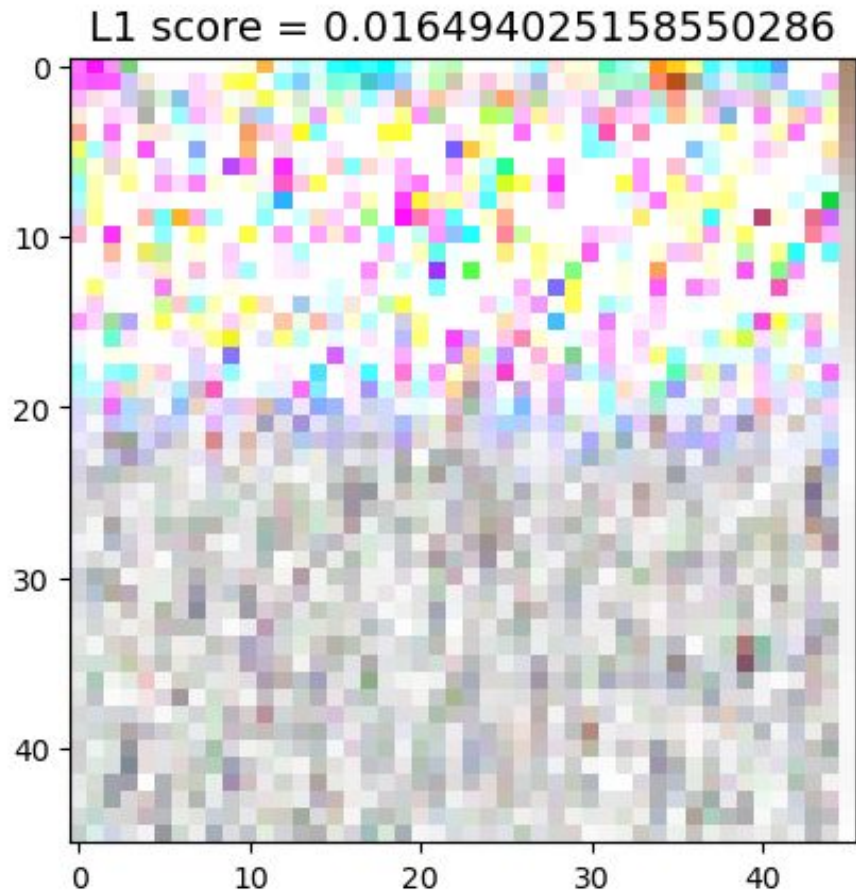
Anomaly Detection

Going 3D



Anomaly Detection

Going 3D



Anomaly Detection

Going 3D

Model: "sequential"

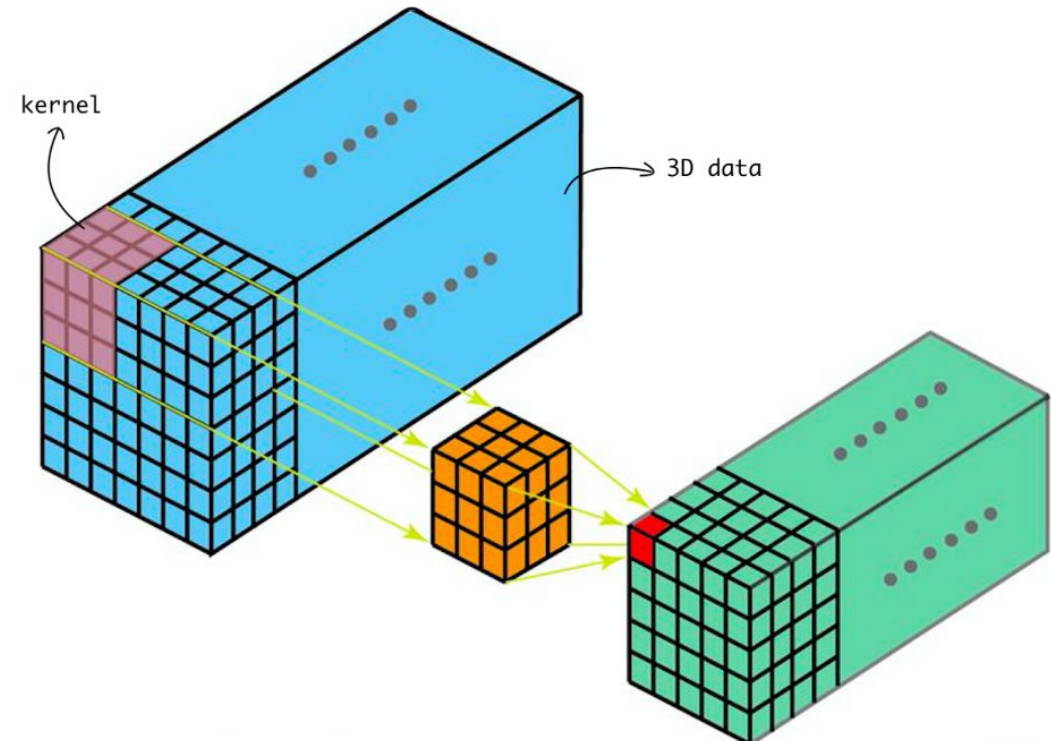
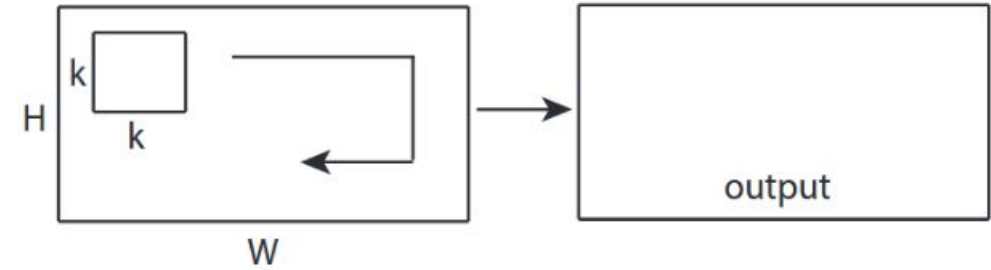
Layer (type)	Output Shape	Param #
conv2d (Conv2D)	(None, 92, 92, 32)	2080
dropout (Dropout)	(None, 92, 92, 32)	0
conv2d_1 (Conv2D)	(None, 46, 46, 16)	131088
conv2d_transpose (Conv2DTran	(None, 46, 46, 16)	65552
dropout_1 (Dropout)	(None, 46, 46, 16)	0
conv2d_transpose_1 (Conv2DTr	(None, 92, 92, 32)	32800
conv2d_transpose_2 (Conv2DTr	(None, 92, 92, 1)	33

Total params: 231,553
Trainable params: 231,553
Non-trainable params: 0

Model: "sequential_7"

Layer (type)	Output Shape	Param #
conv3d_10 (Conv3D)	(None, 46, 46, 4, 32)	16416
dropout_14 (Dropout)	(None, 46, 46, 4, 32)	0
conv3d_11 (Conv3D)	(None, 23, 23, 2, 16)	2097168
conv3d_transpose_14 (Conv3DT	(None, 46, 46, 4, 16)	1048592
dropout_15 (Dropout)	(None, 46, 46, 4, 16)	0
conv3d_transpose_15 (Conv3DT	(None, 46, 46, 4, 32)	262176
conv3d_transpose_16 (Conv3DT	(None, 46, 46, 4, 1)	2049

Total params: 3,426,401
Trainable params: 3,426,401
Non-trainable params: 0



Anomaly Detection Update

Going 3D

