

## Classification of ESSnuSB WC Near Detector Events Using Graph Neural Networks

 $v_{\mu}$ - and  $v_{e}$ -events for neutrino oscillation studies

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## Outline

- Neutrino Physics
- Experiment and Motivation
- Performance on Charged Lepton Simulations
- Performance on Full Neutrino Simulations
- Investigation of Performance Differences





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 $\mathcal{V}$ 

 $v_{\tau}$ 

 $\nu_{\rm e}$ 

μ

• Neutrinos oscillate!





$$P_{\nu_{\mu} \to \nu_{e}(\overline{\nu_{\mu}} \to \overline{\nu_{e}})} \simeq 4s_{23}^{2}s_{13}^{2}\frac{1}{(1-r_{A})^{2}}\sin^{2}\frac{(1-r_{A})\Delta L}{2} +8J_{r}\frac{r_{\Delta}}{r_{A}(1-r_{A})}\cos\left(\delta_{CP}-\frac{\Delta L}{2}\right)\sin\frac{r_{A}\Delta L}{2}\sin\frac{(1-r_{A})\Delta L}{2} +4c_{23}^{2}c_{12}^{2}s_{12}^{2}\left(\frac{r_{\Delta}}{r_{A}}\right)^{2}\sin^{2}\frac{r_{A}\Delta L}{2}$$

$$J_{r} = c_{12}s_{12}c_{23}s_{23}s_{13}, \Delta = \frac{\Delta m_{31}^{2}}{2E_{v}}, r_{A} = \frac{a}{\Delta m_{31}^{2}}, r_{\Delta} = \frac{\Delta m_{21}^{2}}{\Delta m_{31}^{2}}, a = 2\sqrt{2}G_{F}N_{e}E_{v}$$



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 "atmospheric"  
+8 $J_{r}\frac{r_{\Delta}}{r_{A}(1-r_{A})}\cos\left(\delta_{CP}-\frac{\Delta L}{2}\right)\sin\frac{r_{A}\Delta L}{2}\sin\frac{(1-r_{A})\Delta L}{2}$  "interference"  
+4 $c_{23}^{2}c_{12}^{2}s_{12}^{2}\left(\frac{r_{\Delta}}{r_{A}}\right)^{2}\sin^{2}\frac{r_{A}\Delta L}{2}$  "solar"

$$J_{r} = c_{12}s_{12}c_{23}s_{23}s_{13}, \Delta = \frac{\Delta m_{31}^{2}}{2E_{v}}, r_{A} = \frac{a}{\Delta m_{31}^{2}}, r_{\Delta} = \frac{\Delta m_{21}^{2}}{\Delta m_{31}^{2}}, a = 2\sqrt{2}G_{F}N_{e}E_{v}$$



 $\begin{aligned} & \uparrow \\ & \varsigma^{T} \\ 1^{\text{st}} \text{ oscillation max: } A=0.3 \sin \delta_{CP} \\ 2^{\text{nd}} \text{ oscillation max: } A=0.75 \sin \delta_{CP} \end{aligned}$ 



M. Dracos, NuFact2022



- Greater sensitivity to  $\delta_{CP}$  at the second oscillation peak
- Within reach with the 5 MW proton beam produced at ESS



M. Dracos, NuFact2022



# **Experiment and Motivation**



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Luckily, you are already Cherenkov detector experts!



#### Current Framework

Charged Lepton Simulations

https://github.com/WCSim/WCSim

#### Neutrino Interaction Simulations

GENIE Generator. Nucl. Instrum. Meth. A 614:87–104, 2010

WCSIM https://github.com/WCSim/WCSim LLH Based Reconstruction

fiTQun J. Phys.: Conf. Ser. 888 012066, 2017



#### **Current Framework**

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#### Challenges

- Likelihood reconstruction takes ~1 min/event
- To explore different detector proposals, fast reconstruction is crucial



# Why Do We Need GNN Reconstruction?

- Fast and reliable event reconstruction enables testing of different detector layouts
- LLH-based methods are accurate, but reconstruction is **slow**
- ML methods are **fast once trained**, GNNs are well suited for sparse events with irregular geometry
- Multiple reconstruction methods provide a way to **cross check and find systematic errors**







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# Graph Neural Networks (GNNs)

Node = Data point In our case a DOM hit

- Based on graph theory
- Each graph is a neutrino event
- Each data point is a node
- A node has features like xyz, time, charge
- Suited for non-euclidian data





# Graph Neural Networks - Framework



**GraphNeT** - Graph Neural Networks for Neutrino Telescopes <a href="https://github.com/graphnet-team/graphnet">https://github.com/graphnet-team/graphnet</a>

See plenary talk by Rasmus on Monday



Pytorch Geometric - GNN framework for Pytorch

Model architecture: DynEdge



# Graph Neural Networks - Architecture





# Graph Neural Networks - Architecture



ESS neutrino Super Beam plus

LINIVERSITY

## Data Processing and Performance Measures



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### Data processing

- Cuts based on reconstructed variables
- Removes events that are hard to classify
- Reduces events by a factor ~2





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#### **False Positive Rate**

Due to the beam composition, we select samples that have:

- 1 % FPR for muon neutrinos
- 0.1 % FPR for electron neutrinos





# Charged Lepton Performance



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#### Charged lepton simulations - with cuts





#### Charged lepton simulations - with cuts





Charged lepton simulations - with cuts (electron neutrino events)





Charged lepton simulations - with cuts







Charged lepton simulations - with cuts

• For pure charged lepton simulations with filtering of difficult events, the GNN is on par with the fiTQun LLH method.

However:

- Event filter relies on fiTQun reconstructed variables
- Full neutrino events can contain more than single charged leptons (pions, double-decays etc.)



## Neutrino Event Performance



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Neutrino event simulations - without data cut





Neutrino event simulations - without data cut





Neutrino event simulations - without data cut (electron neutrino events)



ROC Curve



ROC Curve with logarithmic x axis

Neutrino event simulations - without data cut

- The GNN has acceptable performance even on the full events
- Using the GNN, the data cuts can be made obsolete

Further investigations

- Look at performance differences on an event basis
- Make a GNN-filter for good/bad events



Neutrino event simulations - without data cut

Data extraction	~10 <sup>-4</sup> mins/event
Training	~10 <sup>-3</sup> mins/event
Reconstruction	~10 <sup>-4</sup> mins/event
fiTQun Reconstruction	~1 min/event
Improvement	$10^3$ (w/ training) / $10^4$ (w/o training)



## **Performance** Investigations



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#### Factors impacting performance - multiple charged lepton signatures Neutrino event simulations - with data cut



#### Factors impacting performance - multiple charged lepton signatures Neutrino event simulations - with data cut





#### Events for which GraphNeT performs significantly better than fiTQun



#### Events for which GraphNeT performs significantly better than fiTQun



#### Factors impacting performance - pion creation Neutrino event simulations - with data cut



#### fiTQun performance with logarithmic x axis





#### Factors impacting performance - pion creation Neutrino event simulations - with data cut







#### Pion production classifier Neutrino event simulations - with data cut





## Factors impacting performance

Neutrino event simulations - with data cut

The GNN is able to identify the characteristics of both

- Events with two Cherenkov rings due to decaying muons
- Events with pion production

Filtering these types of events and treating them separately could be beneficial



# Thank you!

## Additional Slides



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- Example: The **FPR** for **electron neutrino** events is the number of **muon neutrino** events identified as **electron neutrinos**, divided by the total number of **muon neutrino** events



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We select samples that yield FPRs of:

- 1 % for muon neutrinos
- 0.1 % for electron neutrinos



#### Relation between interaction position and reconstruction performance Neutrino event simulations - with data cut

#### Event position distributions





#### Relation between interaction position and reconstruction performance Neutrino event simulations - with data cut





