

Heavy Neutral Leptons search with KM₃NeT-ORCA

NBIA Summer School 2025

9th July 2025

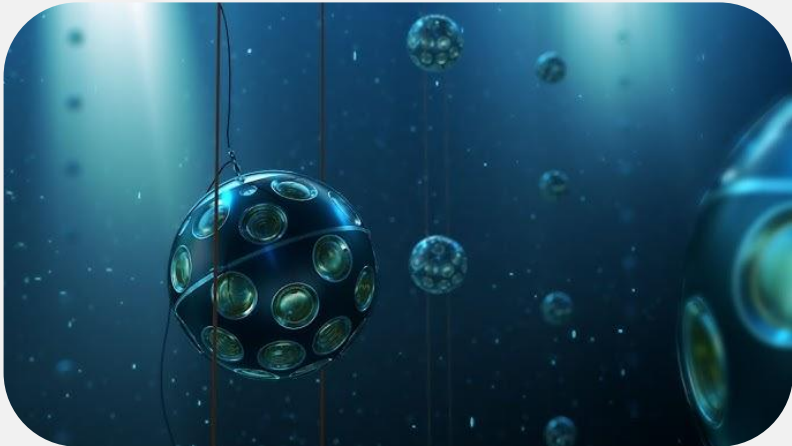
Jorge Prado González (jprado@km3net.de)

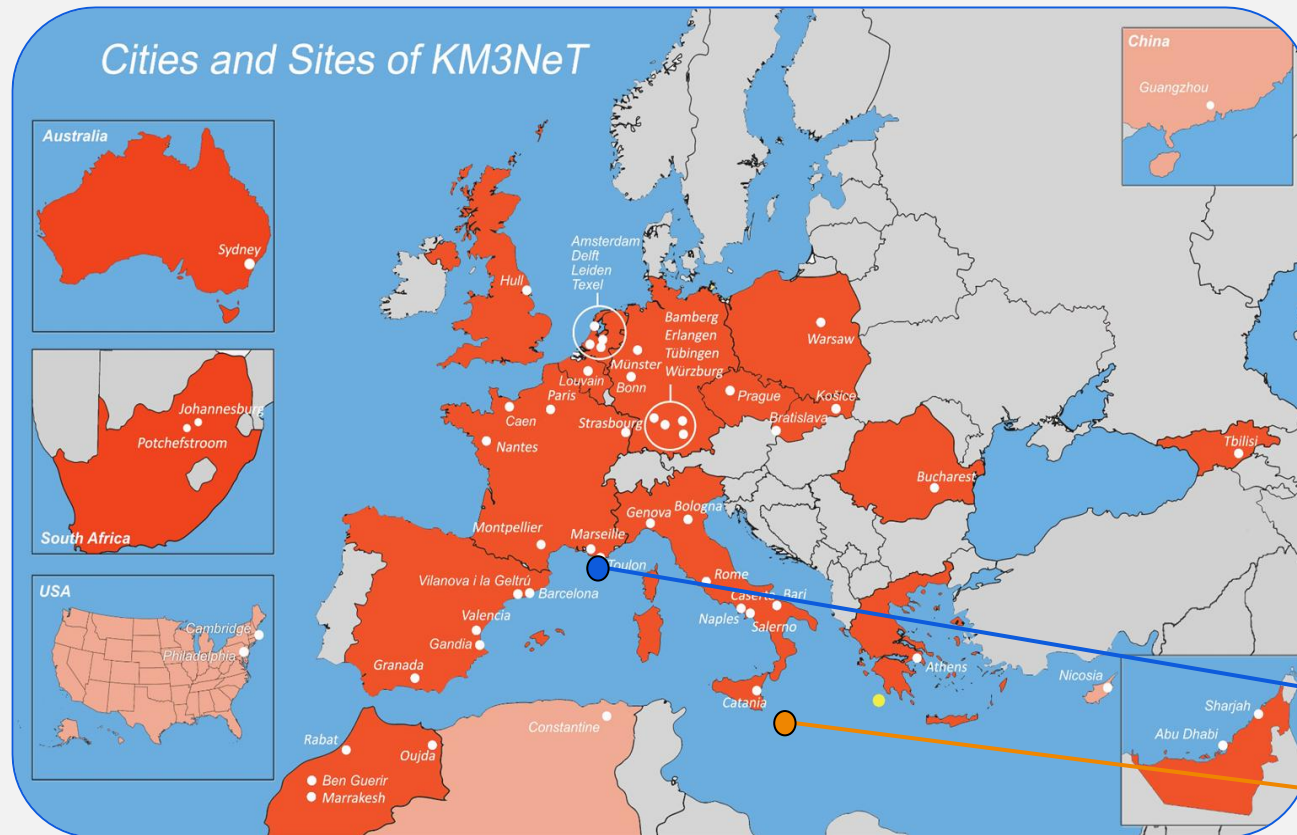
Instituto de Física Corpuscular, Valencia, Spain



PART I

Neutrino Telescope: KM3NeT





International collaboration:

- ~250 members
- 65 partner institutes
- Over 22 countries

Two detectors in two different sites:

- Same technology
- Same data processing
- Shared software and same data formats
- Different size and granularity.

KM3NeT/ORCA

KM3NeT/ARCA

KM3NeT-ARCA and KM3NeT-ORCA

KM3NeT-ORCA:

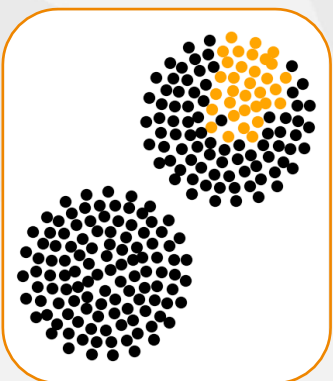
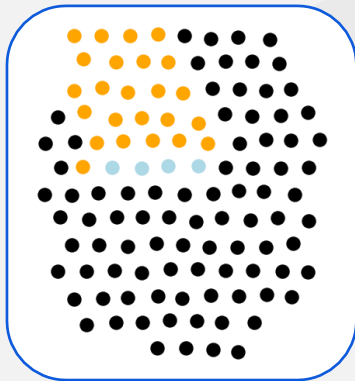
- Low energy (few to hundreds of GeV)
- Fundamental neutrino properties (mainly).

KM3NeT-ARCA:

- High energy (sub TeV to few PeV)
- Astrophysical studies (mainly).

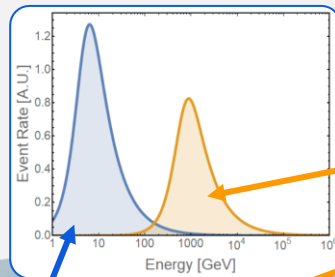
ORCA has 28 DUs

ARCA has 33 DUs



DU: Detection Unit. String of 18 DOMs.

DOM: Digital Optical Module.



ARCA

36m vert. separat
90m horiz. Sepa.
1 Gton detector

DOM

31x3" PMTs



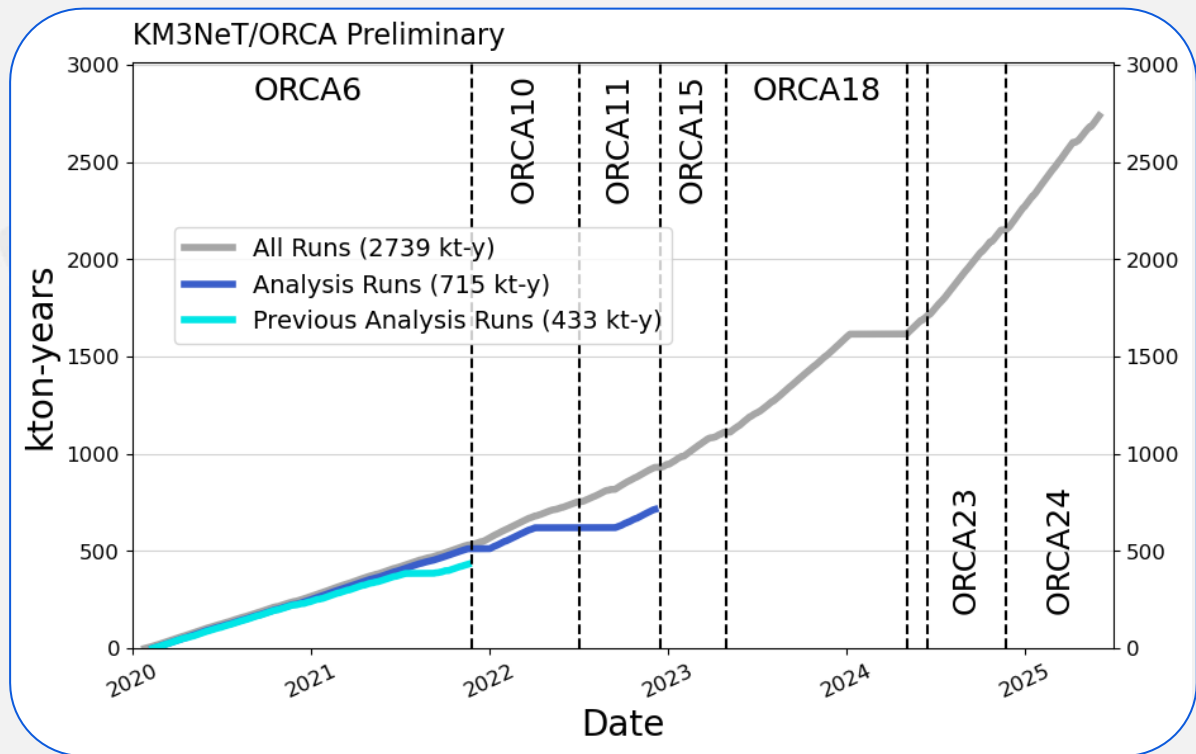
ORCA

9m vert. separat
20m horiz. Sepa.
7 Mton detector



KM3NeT-ORCA

- In this study we will focus on **KM3NeT-ORCA**.
- Already collected **more than 2.7 Mt-y of data**.
- The goal is to study the presence of Heavy Neutral Leptons in ORCA-18.



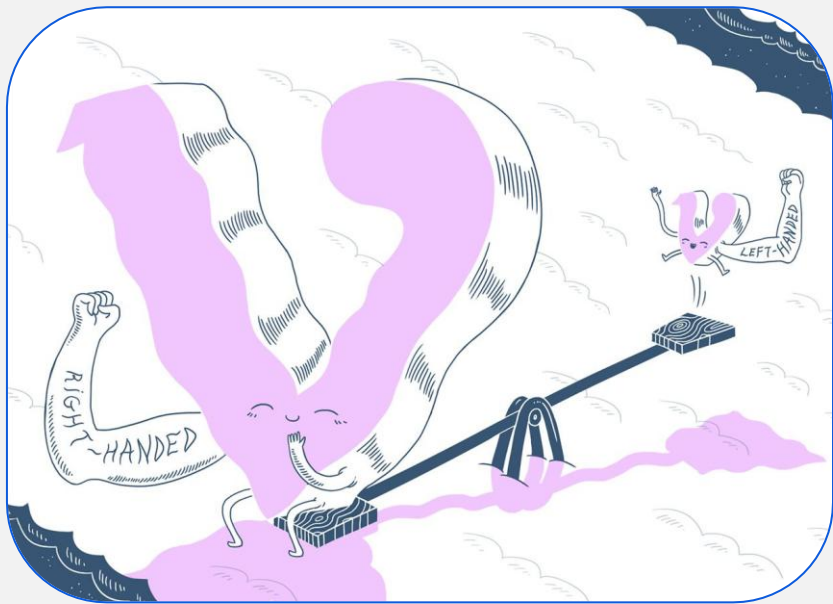
PART II

Heavy Neutral Leptons



Physics motivation

- Neutrino oscillations -> Neutrino have masses.
- No right-handed neutrinos in the standard model means no neutrino mass. **Let's add them!**



I		II		III	
2.3 MeV left u up right		1.28 GeV left c charm right		173.2 GeV left t top right	
4.8 MeV left d down right		95 MeV left s strange right		4.7 GeV left b bottom right	
511 keV left e electron right		105.7 MeV left μ muon right		1.777 GeV left τ tau right	
< 2 eV left ν_e right		< 190 keV left ν_μ right		< 18.2 MeV left ν_τ right	

CERN Courier

→

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CERN Courier

N_1 N_2 N_3

Heavy Neutral Leptons – A bit of theory

- Heavy Neutral Leptons famously appear in See-saw type-I mechanism to explain the tiny neutrino masses.

$$\mathcal{L}_{\text{see-saw}}^{\text{mass}} = -\frac{1}{2}(\bar{\Phi}_L, \bar{\Phi}_R) \begin{pmatrix} 0 & m_D \\ m_D & M \end{pmatrix} \begin{pmatrix} \Phi_L \\ \Phi_R \end{pmatrix}$$

$$\text{Diagonalizing mass matrix: } m_{\text{light}} \simeq \frac{m_D^2}{M}, m_{\text{heavy}} \simeq M$$

- Models with $M \sim \text{KeV}$ are good dark matter candidates and models with $M \sim \text{GeV}$ can successfully generate matter-antimatter asymmetry.

**HNLs
are/have:**

**Right-handed
neutrino partners**

$$M \gg \gg \text{eV}$$

**Feeble interactions
with SM neutrinos**

Portals to Heavy Neutral Leptons

There are many proposed portals between HNLs and SM-neutrinos:

In this study only the **Dipole-Portal model** will be explored

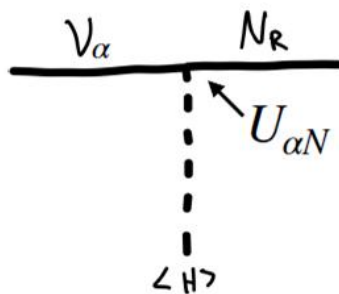
See N.Kamp+ (2025)

Mass-Mixed HNLs

Coupling via mass mixing
as in Seesaw Type I

Parameters

- M_N : HNL mass
- $U_{\alpha N}$: mixing matrix elements

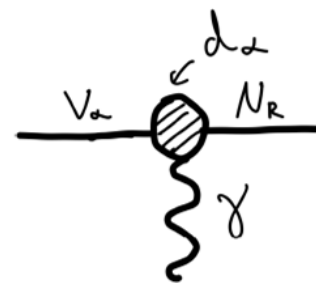


Dipole-Portal HNLs

Coupling via an effective
transition magnetic moment

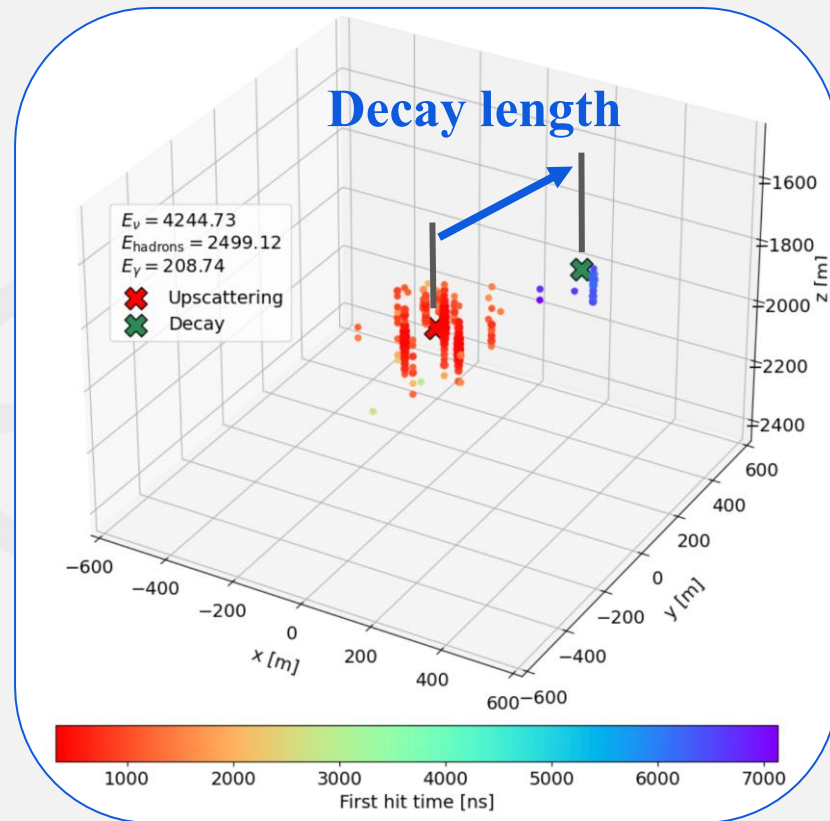
Parameters

- M_N : HNL mass
- $d_{\alpha N}$: effective dipole moment



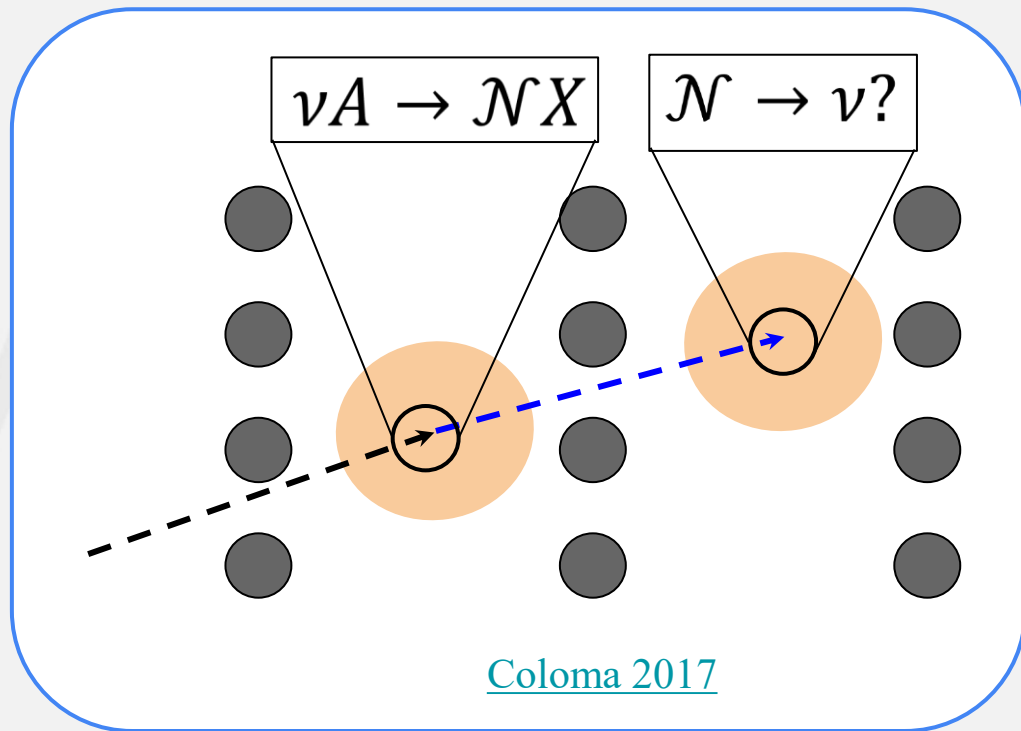
Heavy Neutral Leptons

The presence of HNLs could leave
a signal in KM3NeT-ORCA!

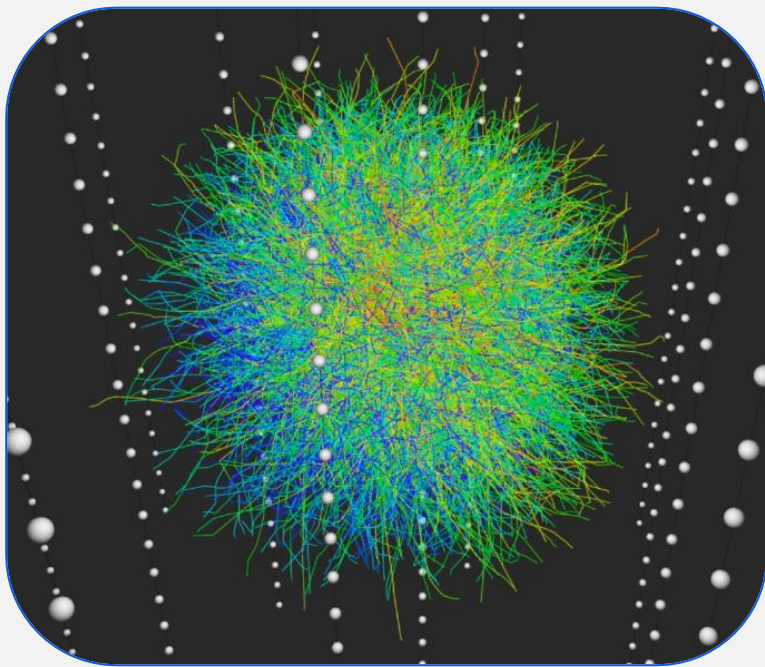


Heavy Neutral Leptons

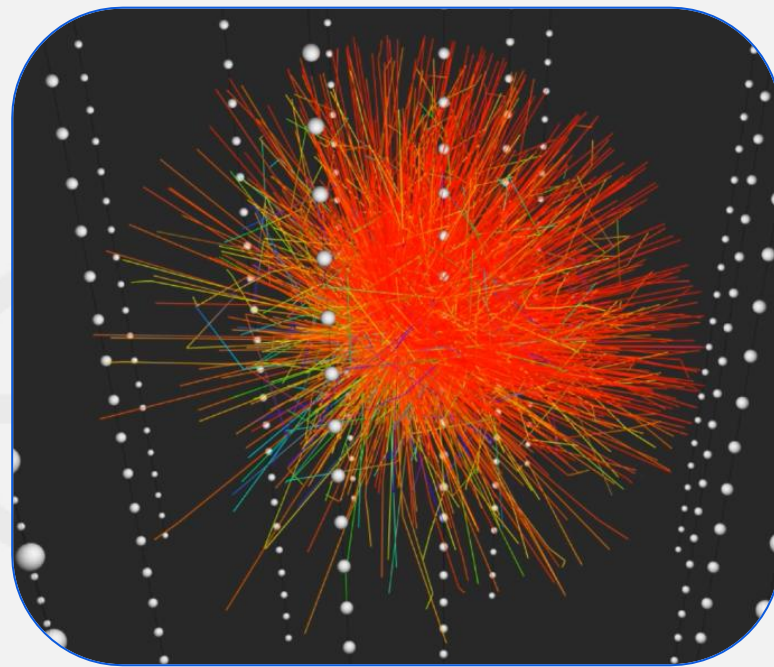
- The signature of this events in ORCA would be the one of **two showers separated a certain distance** at low energy.
- Very unique signal as tau-neutrinos double bang at GeV energies generate showers \sim micrometers apart from each other.



Why in KM3NeT-ORCA-18?



10 TeV cascade ice

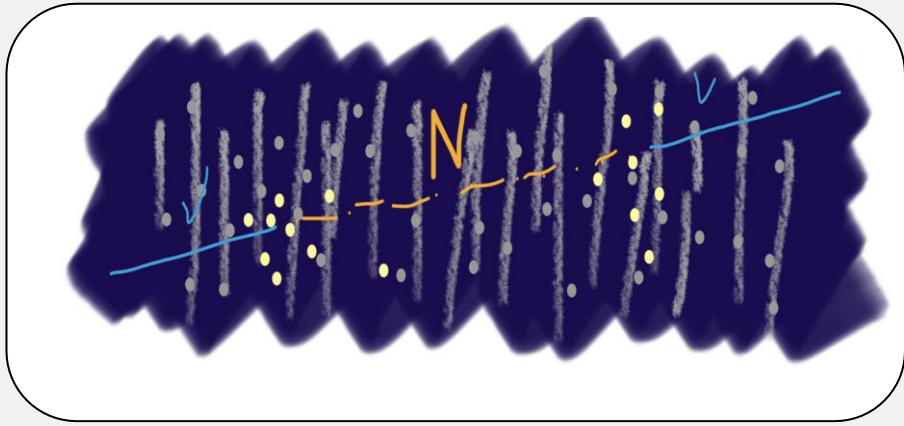


10 TeV in water

Water detectors are expected to have better particle identification capability

PART III

Some preliminary results



Software

- HNL-Events simulated with **SIREN** (Sampling and Injection for Rare EveNts) [[GITHUB](#), [2406.01745](#)]

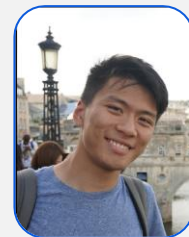
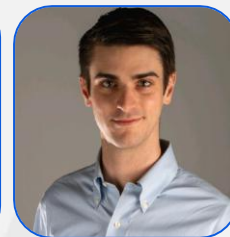
SIREN: An Open Source Neutrino Injection Toolkit

Austin Schneider^{a,b,*}, Nicholas W. Kamp^{c,*} and Alex Y. Wen^c

^aLos Alamos National Laboratory, Los Alamos , NM, United States

^bMassachusetts Institute of Technology, Cambridge , MA, United States

^cDepartment of Physics & Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge 02138, MA, United States

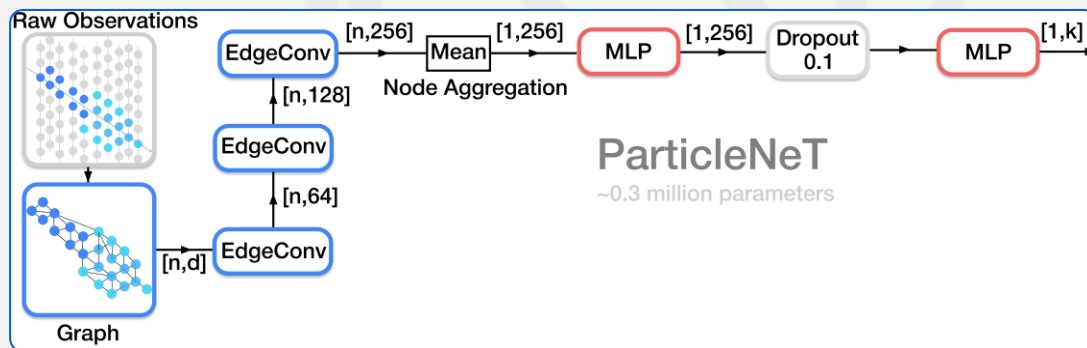


- A slight modification of **ParticleNeT** model (GNN) used to reconstruct this events.



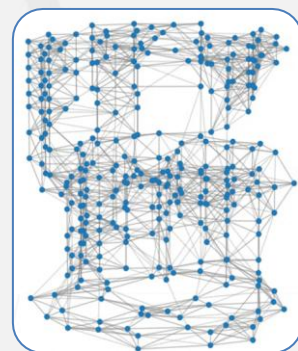
GraphNeT

Deep Learning for Neutrino Telescopes



ParticleNeT

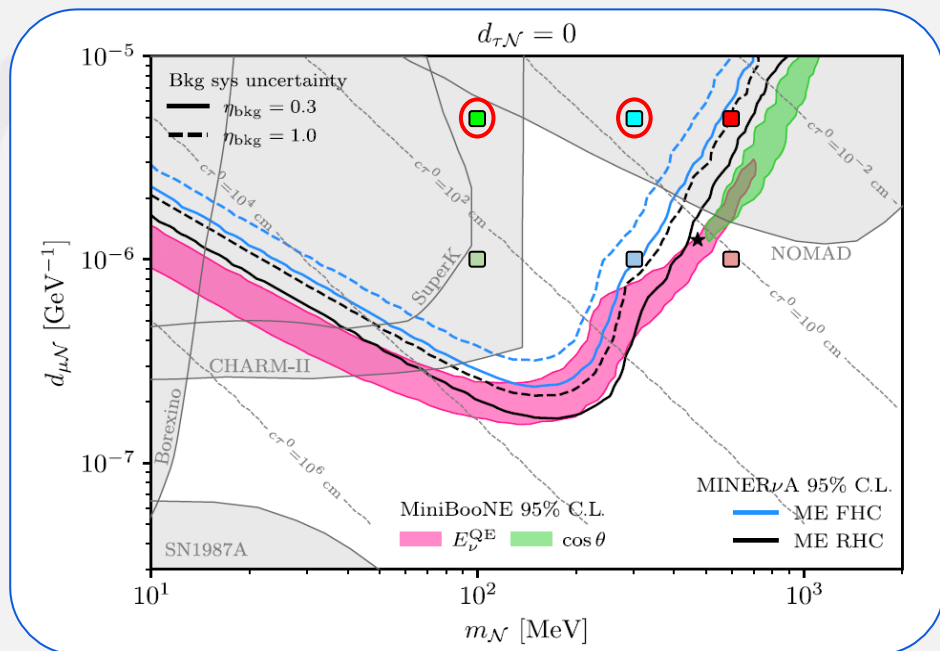
~0.3 million parameters



Test

- Test on a SIREN driven simulation on **ORCA-15**.
- Using models with $M_N = 100, 300$ and 600 MeV and $d_{\mu N} = 10^{-6}, 5 \times 10^{-6}$

	Total Events (per year)	Total Events with $d > 10m$ (per year)
$m=0.1, d=1e-6$	0.12	0.11
$m=0.1, d=5e-6$	35.3	33.7
$m=0.3, d=1e-6$	2.27	1.82
$m=0.3, d=5e-6$	80	14.8
$m=0.6, d=1e-6$	3.0	0.81
$m=0.6, d=5e-6$	78.5	0.27



Event selection

Steps to get a pure HNL sample.

- **L0 – Cut on variables:** Upgoing events with more than six hits in total.
- **L1 – Antinoise BDT:** Cut on score so that all noise is rejected.
- **L2 – Antimuon BDT:** Cut on score so that all the muons are rejected.
- **L3 – Antineutrino cut:** Cut on the reconstructed distance between vertices and a classifier both predicted by ParticleNeT.

Event selections

The BDTs are quite efficient rejecting muons and noise.



KM3NeT-Work in Progress

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Event selections

In this first study we were capable of
isolating an HNL sample with no noise,
muons or neutrino **contamination**.

KM3NeT-Work in Progress

**Very pure sample with ~XX events per
year!**

KM3NeT-Work in Progress

- First studies showing **very promising results**, even if the study is very preliminary.
- Already working with the ORCA-18 simulation. Exciting results son!

Thank you!

Models trained in:

