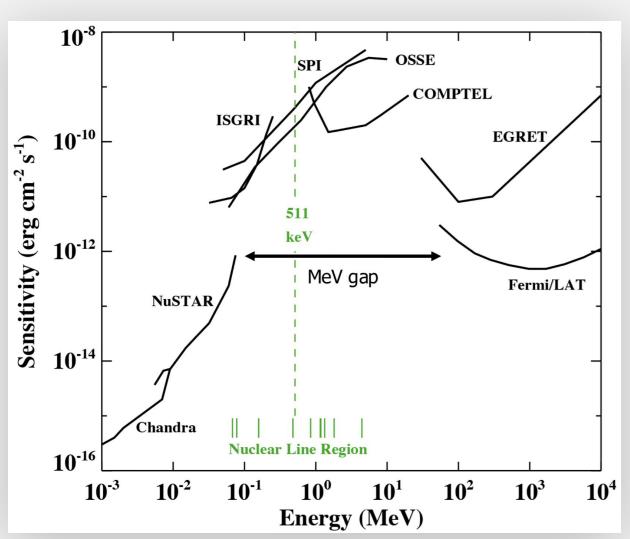


# Unlocking the MeV Spectrum: Advancing Detector Technology with Al

Supernovae



The MeV spectrum



Plot from John Tomsick, "Compton Polarimetry and the Compton

Spectrometer and Imager", FiXPAcademy 2025

- Medium-energy X- and gamma-ray (~0.1-100 MeV) emitted by cosmic events such as gamma-ray bursts or Supernovae
- Difficult to detect due to low flux, low photon interaction probability and complex energy-loss processes.
- The MeV Gap a significant gap in sensitivity due to limited detector capabilities compared to adjacent energy bands.

Electron

Gamma-ray bursts

Kilonovae

Nearby galaxies

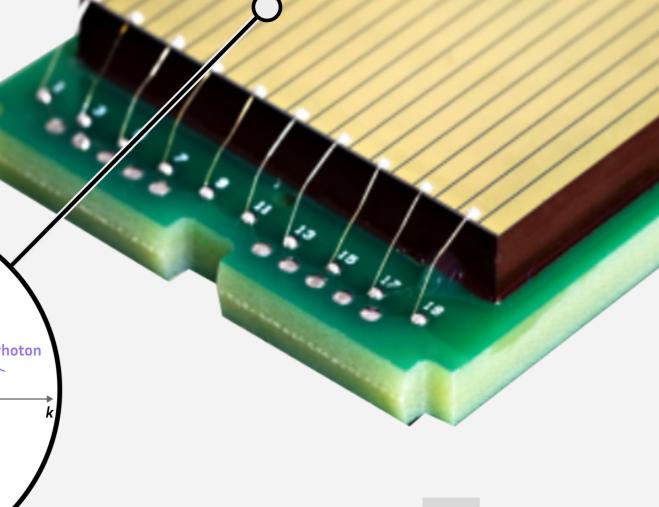
Magnetars

# Radiation detector

Advanced detector technology such as the 3D CZT (Cadmium Zinc Telluride) drift strip detector, developed at DTU Space, offers high-resolution X-ray and gamma-ray detection.

The detector's **compact size and high sensitivity** make it promising for applications future space mission and medical imaging.

Each photon interaction generates complex, **high-dimensional data** (pulse shapes), encoding the radiation trace through underlying physical reactions.



Key takeaways 5



AI-driven signal processing enables compact, high-performance instrumentation for future space missions and MeV astrophysics.

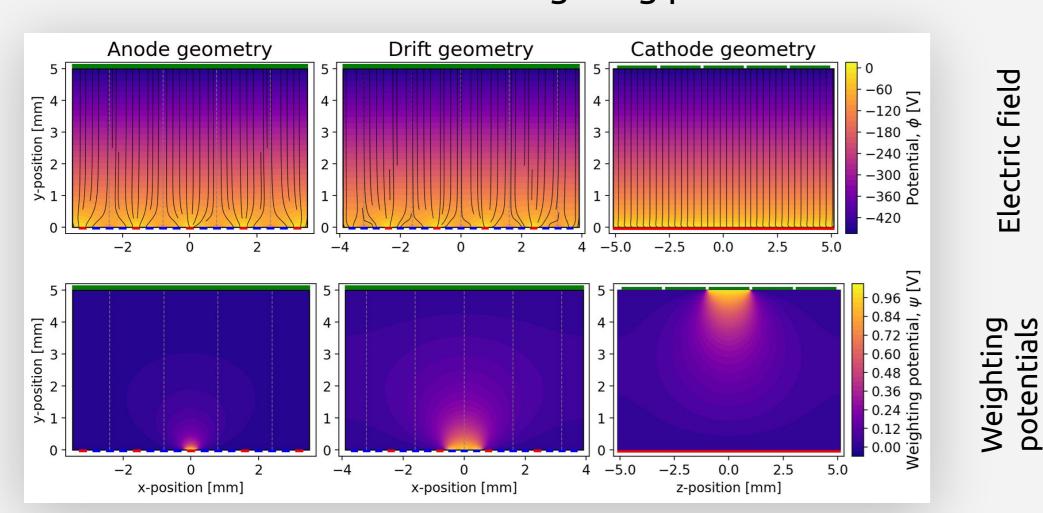
MeV photon

Neural networks trained on simulated data can outperform conventional algorithms especially near the boundary regions of the detector allowing more accurate reconstruction of photon interaction events.

# Simulations

Detailed modeling of the detector's internal physics, including electric fields and charge transport, is used to generate synthetic detector output (pulse-shaped signals). This, in turn, is used to train Neural Networks, ensuring data-efficient learning.

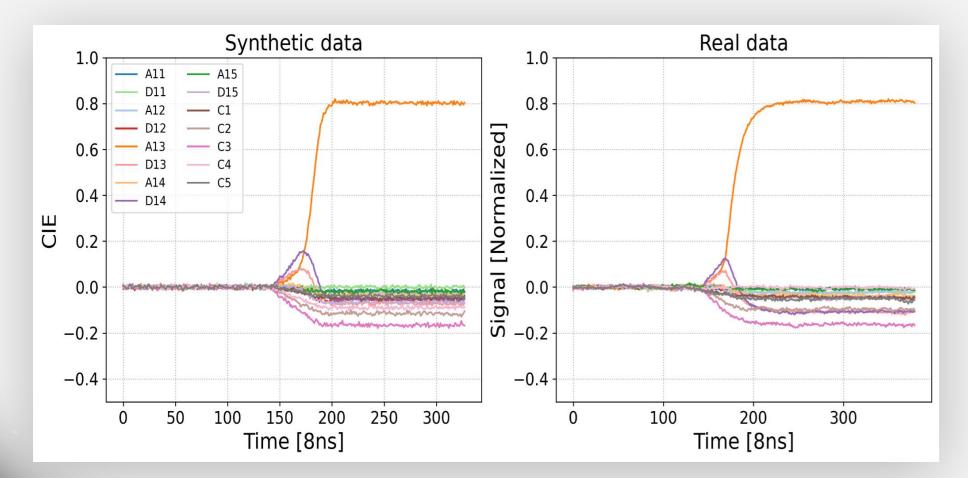
Simulated electric field and weighting potentials



Charge transport model

$$\mu_n 
abla \phi \cdot n^+ + 
abla \cdot (D_n 
abla n^+) - rac{n^+}{ au_n} + \mu_n 
abla \phi \cdot 
abla \psi_k = rac{\partial n^+}{\partial t}$$
 Charge Induction Efficiency (CIE) trapping

Generated detector output

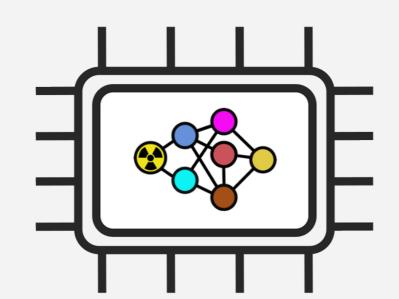


The better the match, the better the performance of Neural Network models.

#### Signal converters



Real-time signal processing

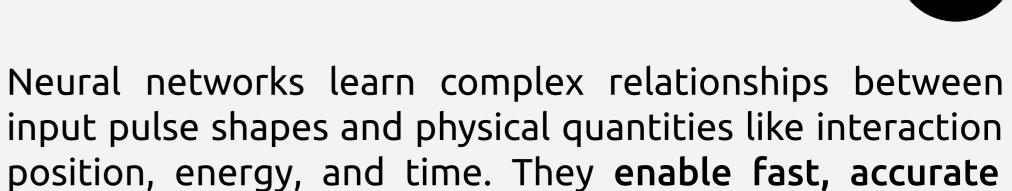


Incident radiation information (photon-by-photon)

Time

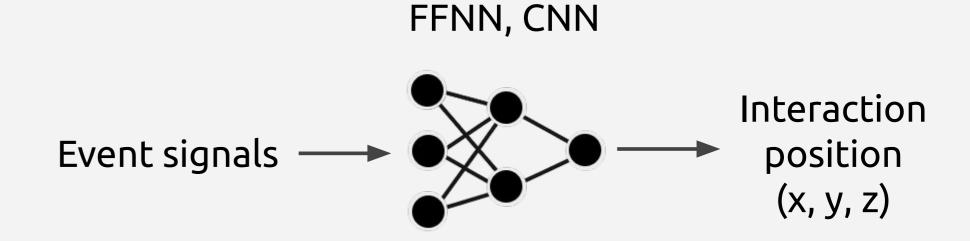
Energy **Position** 

#### Neural Networks 4

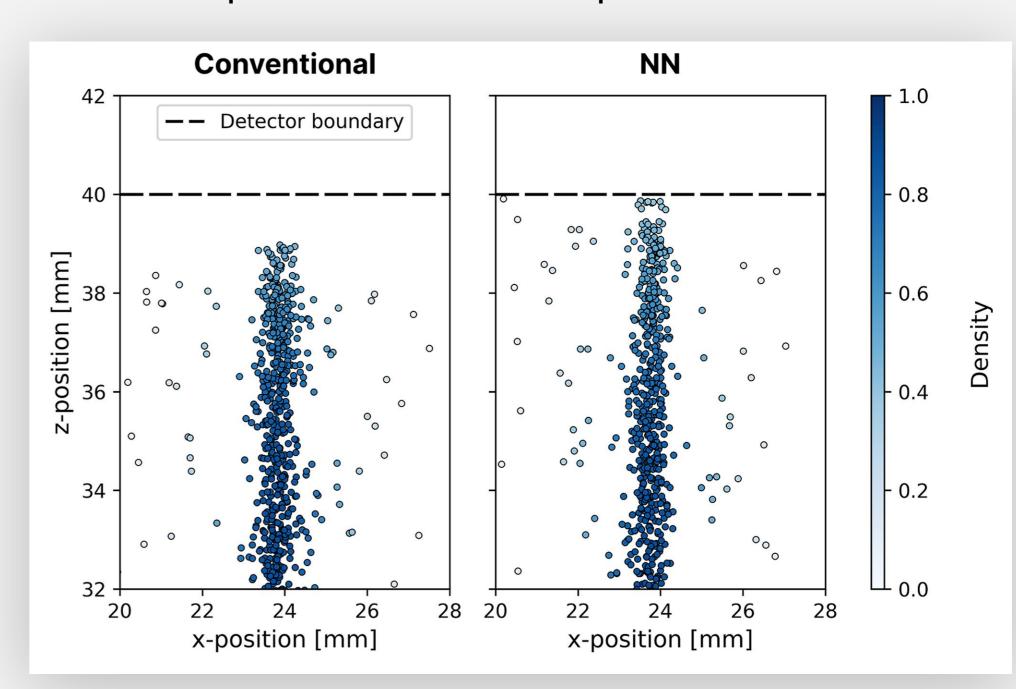


• Training: supervised learning on simulated data

**reconstruction** of photon events.

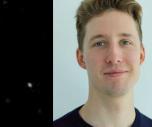


**Results:** predicted interaction positions on real data



Neural Networks show increased performance near the edges of the detector, thanks to learning from data.

Want to learn more about the project?







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Type





