

WELCOME TO THE FIFTH GRAPHNeT WORKSHOP

Outline:

History of GraphNeT through Earlier Workshops

Example results with GraphNeT

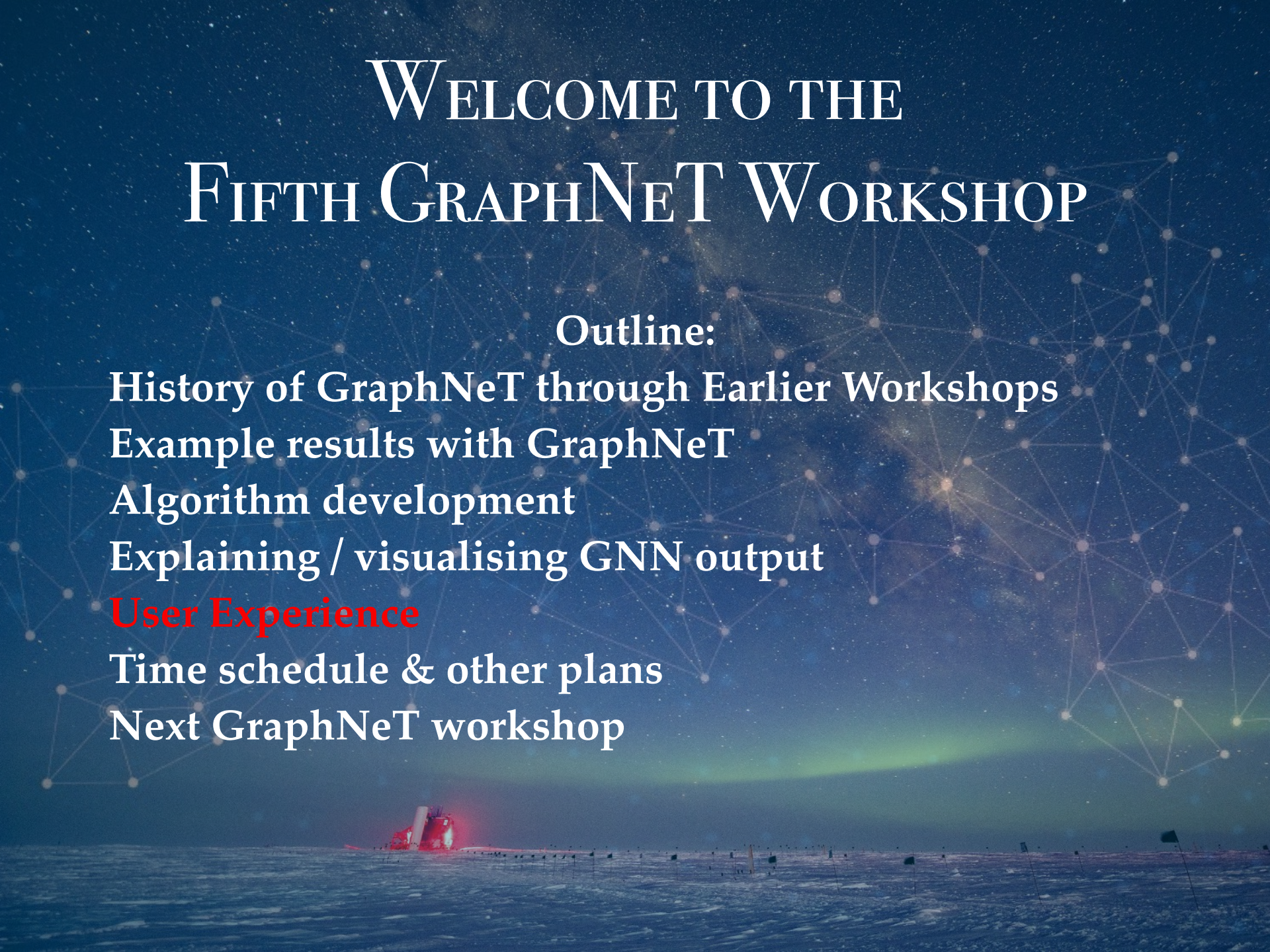
Algorithm development

Explaining / visualising GNN output

User Experience

Time schedule & other plans

Next GraphNeT workshop



INITIAL WORKSHOPS

IceCubeML Workshop

Niels Bohr Institute, 15th of September 2020

Preliminary program (TBC):

9:00 Breakfast

9:30 Status of ML methods (3 talks)

10:15 Discussion of ML methods

10:30 Break

10:45 Status on data formats (1 talk)

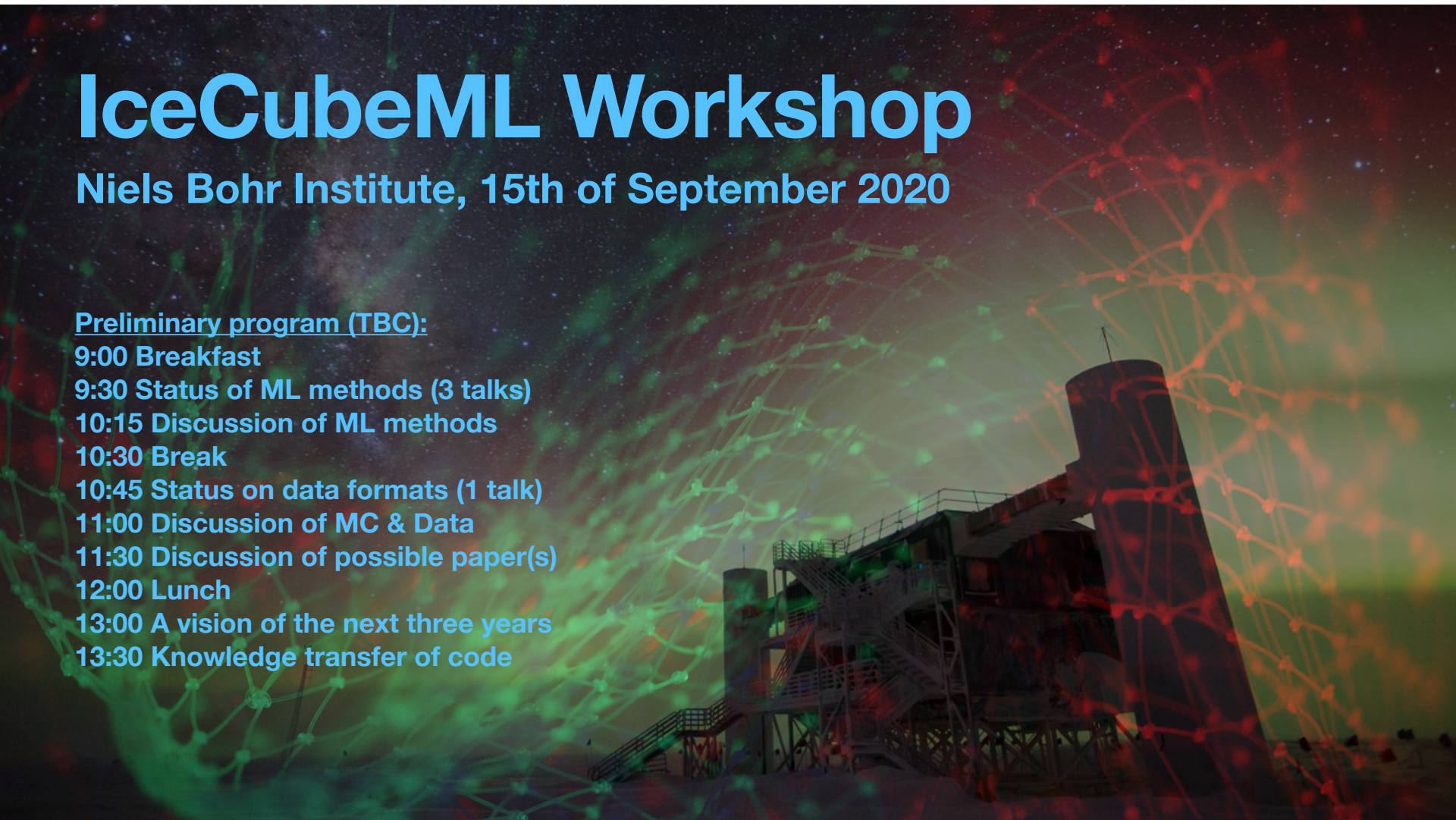
11:00 Discussion of MC & Data

11:30 Discussion of possible paper(s)

12:00 Lunch

13:00 A vision of the next three years

13:30 Knowledge transfer of code



INITIAL WORKSHOPS

IceCubeML Workshop

Niels Bohr Institute, 20th of October 2020

Preliminary program (TBC):

- 9:00 Breakfast
- 9:30 Discussion of MC & Data samples (All)
- 10:15 Chat about upgrade paper (Tom?)
- 10:30 Break
- 11:00 Hands on database format (Mads)
- 11:45 Knowledge transfer of code (Bjørn)

*Neutrinos induce courage in theoreticians,
and persererance in experimenters.*

[Maurice Goldhaber]



GraphNeT

Graph Neural Networks for
Neutrino Telescope Event Reconstruction

GraphNet is our attempt at putting GNN models for IceCube (and others) using the “DynEdge” architecture build in PyTorch Geometric into an easily available software package.

<https://github.com/icecube/graphnet/>

In the following, I will try to convince you, that GNNs “match” the IceCube data really well, and perhaps whet your appetite for using it.

Our results can be found in an IceCube paper, submitted to JINST in September.

GRAPHNET WORKSHOP II

MUNICH 2022

GraphNeT Workshop

from Thursday, September 15, 2022 (9:00 AM) to Saturday, September 17, 2022 (5:00 PM)

Garching (MDSI Meeting Room)



: Sessions / : Talks : Breaks

Sep 15, 2022		Sep 16, 2022	
AM	<p>9:00 AM Check-in at venue (until 9:45 AM) (MDSI Meeting Room) </p> <p>9:45 AM GraphNeT Road Map - Mr Andreas Sogaard (MDSI Meeting Room) 220915_GraphNeT_Workshop_AndreasSogaard.pdf </p> <p>10:30 AM --- Coffee ---</p> <p>10:45 AM Graphite & Bending Graphs - Mahdi Saleh (MDSI Meeting Room) GraphNet_talk_graphite.pdf </p> <p>11:30 AM High Energy Update - Martin Ha Minh Martin Ha Minh (MDSI Meeting Room) 22_09_gnn_HE.pdf </p>		<p>9:45 AM Impact So Far - Rasmus Orsøe (TUM) (MDSI Meeting Room) GraphNeT Workshop @ Garching.pdf </p> <p>10:30 AM --- Coffee ---</p> <p>10:45 AM Calibration & Other Ideas - Prof. Troels Petersen (MDSI Meeting Room) Munich_GraphNetWorkshop_AnalysisIdeas_16sep2022-1.pdf </p> <p>11:30 AM GNNs in P-ONE - Arturo Anaya (MDSI Meeting Room) gnns-for-event-classification-at-p-one-detector.pdf </p>
PM	<p>12:00 PM --- Lunch ---</p> <p>1:00 PM Hackathon (until 3:00 PM) (MDSI Meeting Room) </p> <p>3:00 PM --- Coffee ---</p> <p>3:15 PM Hackathon (until 5:00 PM) (MDSI Meeting Room) </p> <p>6:30 PM Workshop Dinner (until 8:30 PM) (MDSI Meeting Room) </p>		<p>12:00 PM --- Lunch ---</p> <p>1:00 PM Hackathon (until 3:00 PM) (MDSI Meeting Room) </p> <p>3:00 PM --- Coffee ---</p> <p>3:15 PM Hackathon (until 5:00 PM) (MDSI Meeting Room) </p>

GRAPHNET WORKSHOP III BORNHOLM 2023

3rd Workshop on Graph Neural Networks for Neutrino Telescope Event Reconstruction

GraphNeT III

1-4 May 2023

Bornholm, Denmark

Featured speakers

Matthias Fey

Technical University Dortmund and kumo.ai
Creator of PyG (Pytorch Geometric)

Raghav Selvan

University of Copenhagen

Local organiser
Andreas Søgaard (NBI)

Workshop committee
Rasmus Ørsoe (TUM)
Philipp Eller (TUM)
Troels Petersen (NBI)

Visit indico.nbi.ku.dk/e/graphnet2023

GRAPHNeT WORKSHOP IV TUM 2024

4th GraphNeT Workshop

“Graph Neural Networks and Beyond”

6th – 9th May, 2024

Institute for Advanced Study, Munich, Germany

Local organiser

Rasmus Ørsøe (TUM)

Workshop Committee

Rasmus Ørsøe (TUM)

Aske Rosted (CU)

Philipp Eller (TUM)

Troels Petersen (NBI)

Visit <https://indico.ph.tum.de/e/graphnet2024>



Organization



UNIVERSITY OF
COPENHAGEN



Financial Support



SFB 1258



GRAPHNET WORKSHOP V

THE NEXT STEPS FOR GRAPHNET

Scientific advisory committee:

Rasmus Ørsøe (TUM)
Jorge Prado (IFIC)
Philipp Eller (TUM)
Aske Rosted (Chiba Univ.)
Ivan Mozun (LPC Caen)
Troels C. Petersen (NBI)

Local organising committee:

Troels C. Petersen (NBI)
Rasmus Ørsøe (TUM)
Johann Ioannou-Nikolaides (NBI)
Marc Jacquart (NBI)

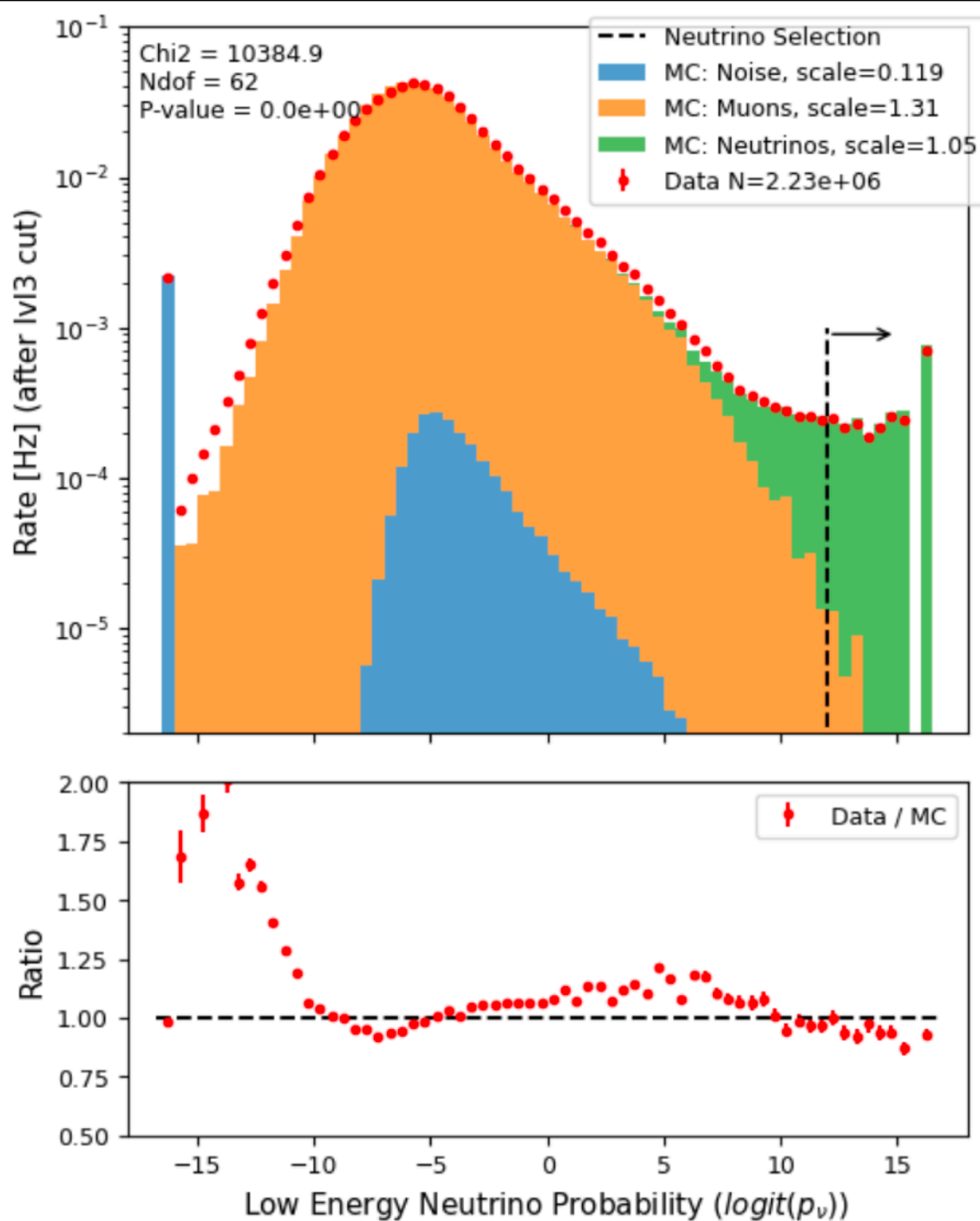
Conference Link:

<https://indico.nbi.ku.dk/event/2170/>

NIELS BOHR INSTITUTE 18TH - 19TH OF AUGUST 2025



A highlight



Classification of neutrinos, muons, and noise interacting in DeepCore.

Not only was the performance increased, but the data/MC correspondence remained good.

We would like something similar for High Energy!

ALGORITHMIC DEVELOPMENTS



Algorithmic Developments

1. GNN autoencoder (for general search)

If an AE learned to encode general IceCube events (say noise, muons, and neutrinos from simulation), it could be used to detect other objects.

2. Hierarchical Graph Pooling

Would this work better? Or better only at high energies? Or...?

3. GNN optimisation (architecture, training sample and reweighting, ensemble combination)

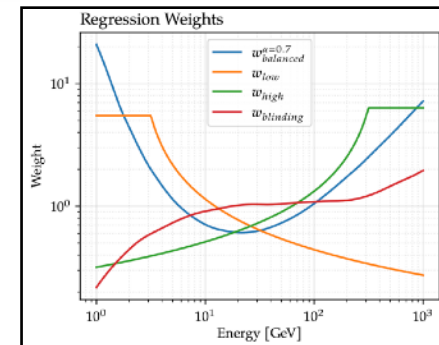
Much has been done, but we should push it to the limit :-)
Perhaps put in 3D-PCA transformation to begin with?

4. Using "non-signal" DOMs as input

Do these contain information to be incorporated, or not?

5. Adversarial training for better performance in data

Do common data+MC training. But technically how?



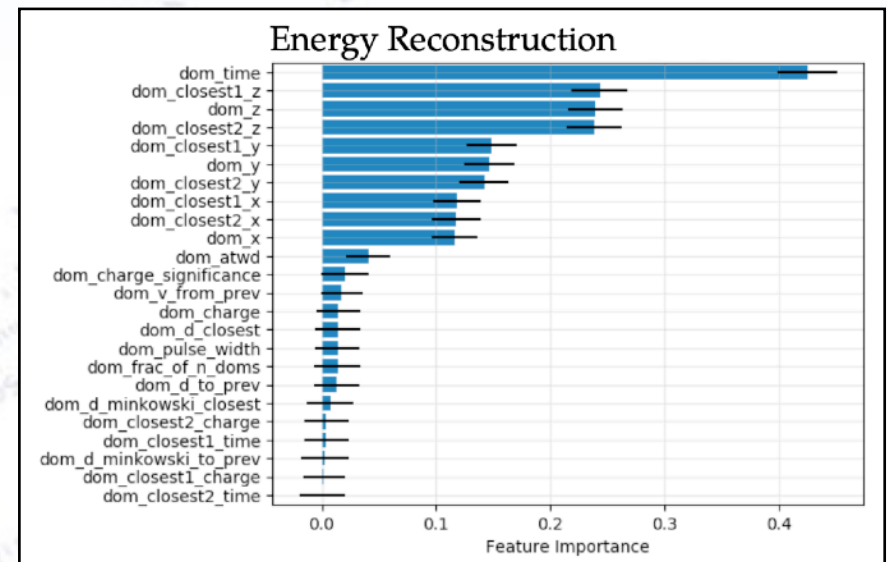
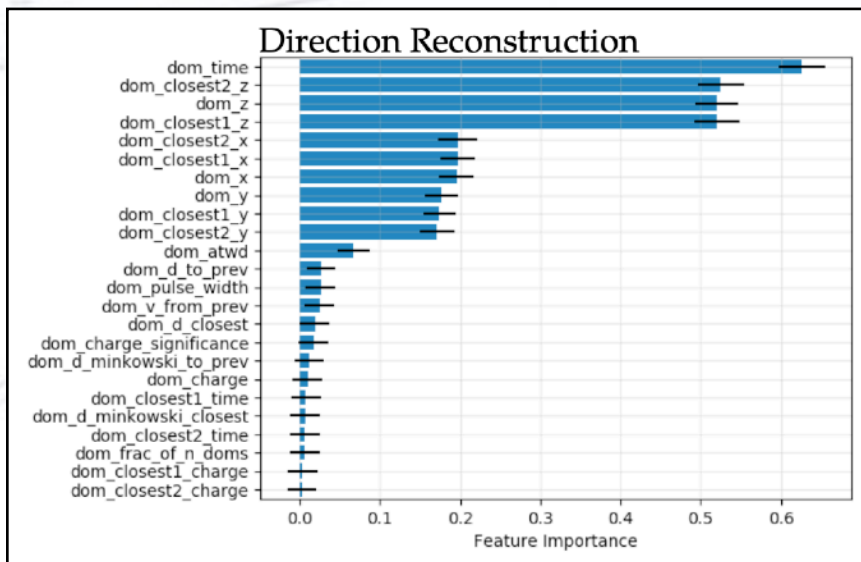
EXPLAINING/VISUALISING THE GRAPHNET ANSWERS



Explaining / Visualising GraphNeT output

1. Overall feature ranking for each task

Which are the important features for energy regression and directional regression? Are they the same or do some cases stand out?



Results shown are for a GRU from 2021

USER EXPERIENCE

... OR LACK OF THE SAME!



LOCAL INFORMATION

Niels Bohr Institute Locations for Workshop

Possible Hackathon location

Possible parking

See Troels for parking vouchers!

PLEASE SHOW UP HERE!
Entrance for Auditorium B

Adresse: Blegdamsvej 17
Nearest station: Østerport (1.5km)
Nearest metro: Trianglen (0.5km)
Problems: +45 26 28 37 39 (Troels)

Blegdamsvej

INDICO TIMETABLE - MONDAY

09:00	Beginning with Breakfast and Badges	
	<i>Aud. B, Niels Bohr Institute</i>	09:00 - 09:30 📍
	1 - Welcome and Framing of workshop	<i>Troels Petersen</i> 📍
10:00	2 - GraphNeT: Status and Plans	<i>Rasmus Ørsøe</i>
	<i>Aud. B, Niels Bohr Institute</i>	09:45 - 10:15 📍
	3 - GraphNeT: Transformers and next steps	<i>Iván Mozún Mateo</i>
11:00	<i>Aud. B, Niels Bohr Institute</i>	10:15 - 10:45 📍
	Break	
	<i>Aud. B, Niels Bohr Institute</i>	10:45 - 11:15 📍
12:00	4 - Technical requirements to reconstruction software in IceCube	<i>Philipp Soldin</i>
	<i>Aud. B, Niels Bohr Institute</i>	11:15 - 11:45 📍
	5 - GraphNeT as seen from a Magic point of view	<i>Jarred Green</i>
13:00	<i>Aud. B, Niels Bohr Institute</i>	11:45 - 12:15 📍
	Lunch Break	
	<i>Aud. B, Niels Bohr Institute</i>	12:15 - 13:15 📍
14:00	8 - Takeaways from training a foundation model for IceCube	<i>Inar Timiryasov et al.</i>
	<i>Aud. B, Niels Bohr Institute</i>	13:15 - 13:45 📍
	9 - All the technicalities: Databases, MC labels, and such stuff	<i>Aske Rosted</i>
15:00	<i>Aud. B, Niels Bohr Institute</i>	13:45 - 14:15 📍
	18 - Introduction to Hackathon & Implementation discussion	<i>Troels Petersen</i> 📍
	Break	
	<i>Aud. B, Niels Bohr Institute</i>	14:30 - 15:00 📍
15:00	7 - Workshop Hackathon I	

INDICO TIMETABLE - TUESDAY

09:00	Breakfast	
	<i>Aud. B, Niels Bohr Institute</i>	09:00 - 09:30
	10 - State of the Art in AstroPhysics research	<i>Markus Ahlers</i>
	<i>Aud. B, Niels Bohr Institute</i>	09:30 - 10:00
10:00	12 - GraphNeT as seen from an analyser and software perspective	<i>Tom Stuttard</i>
	<i>Aud. B, Niels Bohr Institute</i>	10:00 - 10:30
	Break	
	<i>Aud. B, Niels Bohr Institute</i>	10:30 - 11:00
11:00	13 - GraphNeT from a student user point of view	<i>Cyan Yong Ho Jo</i>
	<i>Aud. B, Niels Bohr Institute</i>	11:00 - 11:30
	14 - GraphNeT from an experienced outsider view	<i>Antonin Vacheret</i>
	<i>Aud. B, Niels Bohr Institute</i>	11:30 - 12:00
12:00	Lunch Break	
	<i>Aud. B, Niels Bohr Institute</i>	12:00 - 13:00
13:00	19 - GraphNeT Issues and Wishes	<i>Jorge Prado González</i>
	<i>Aud. B, Niels Bohr Institute</i>	13:00 - 13:30
	20 - GraphNeT Community Discussion	<i>Iván Mozún Mateo et al.</i>
	<i>Aud. B, Niels Bohr Institute</i>	13:30 - 14:30
14:00	Break	
	<i>Aud. B, Niels Bohr Institute</i>	14:30 - 15:00
15:00	15 - Workshop Hackathon II	
	<i>Aud. B, Niels Bohr Institute</i>	15:00 - 15:45
16:00	16 - Summary and Organisation of Plans	
	<i>Aud. B, Niels Bohr Institute</i>	15:45 - 16:15

Directions to Troels' place



GRAPHNET WORKSHOP VI

GERMANY/DENMARK/???

XTH TO YTH OF Z 2026?

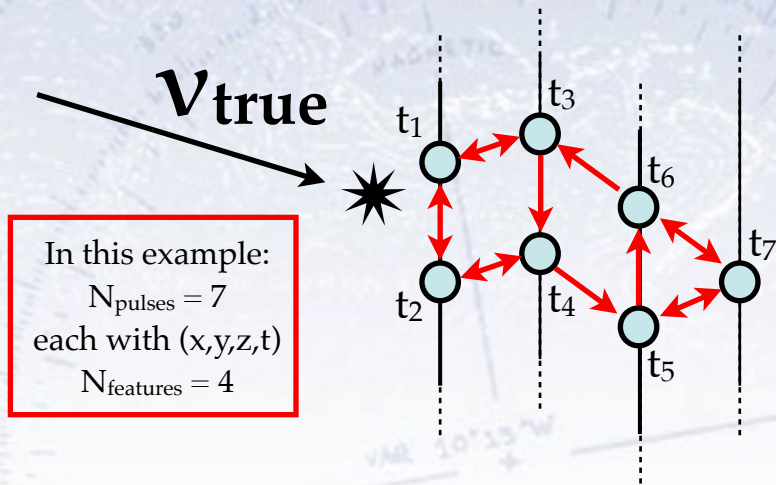




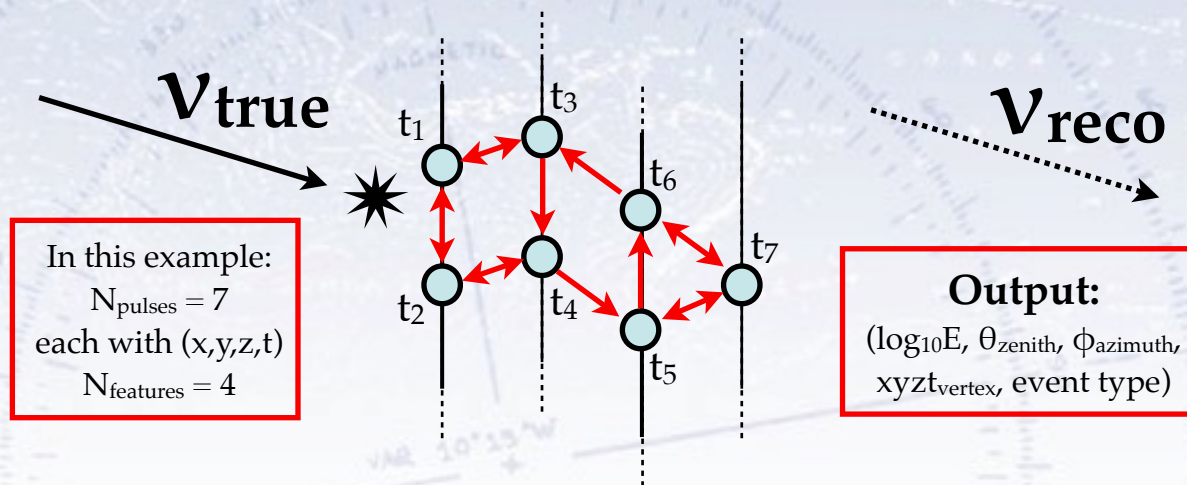
The background is a faded nautical chart. It features a compass rose with a vertical line pointing to 0 degrees, labeled 'MAGNETIC'. Concentric curved lines represent magnetic isogons, with labels such as 10, 20, 30, 40, 50, 60, 70, 80, 90, 100, 110, 120, 130, 140, 150, 160, 170, 180, 190, 200, 210, 220, 230, 240, 250, 260, 270, 280, 290, and 300. A specific line is labeled 'VAR 10° 13' W'. In the upper right, there is a label '1000 YARDS' and '1000 FATHOMS'. The text 'Bonus slides' is centered over the chart.

Bonus slides

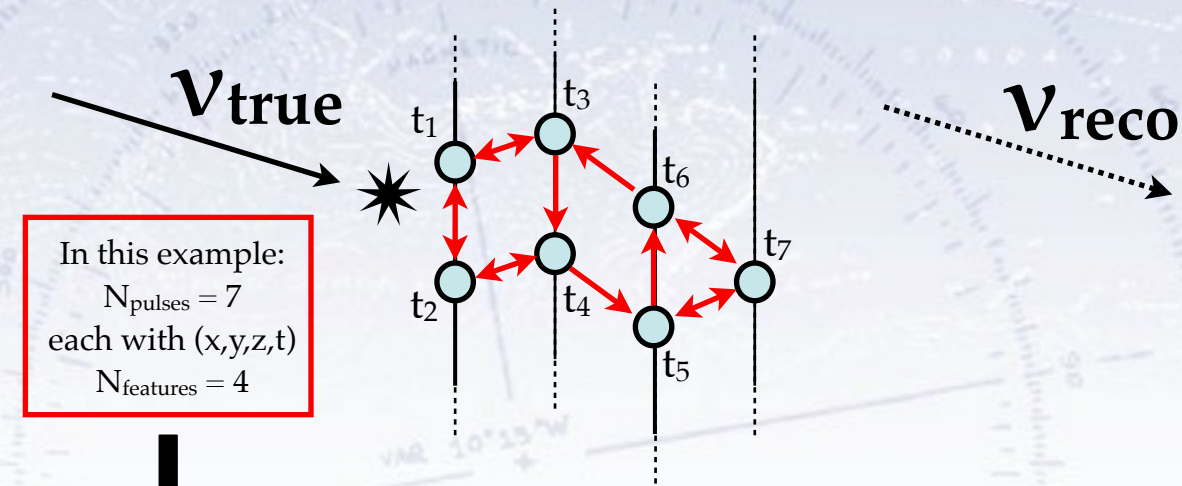
Details of GNN reconstruction



Details of GNN reconstruction



Details of GNN reconstruction



$$\vec{v}_1 = [x_1 \ y_1 \ z_1 \ t_1]$$

$$\vec{v}_2 = [x_2 \ y_2 \ z_2 \ t_2]$$

\vdots

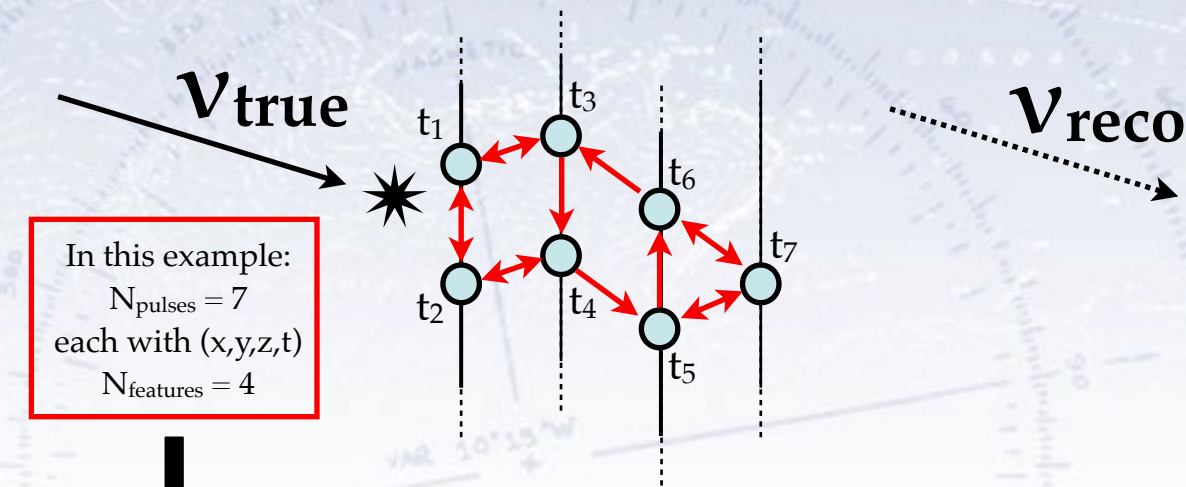
$$\vec{v}_7 = [x_7 \ y_7 \ z_7 \ t_7]$$

Input:

$$N = N_{\text{pulses}} \times N_{\text{features}}$$

The input features of a node are combined with that of $N (=2)$ nearby nodes

Details of GNN reconstruction



$$\begin{aligned}
 \vec{v}_1 &= [x_1 \ y_1 \ z_1 \ t_1] \xrightarrow{EC(\vec{v}_1, \vec{v}_2, \vec{v}_3)} [g_{11} \dots g_{1N_1}] \\
 \vec{v}_2 &= [x_2 \ y_2 \ z_2 \ t_2] \xrightarrow{\hspace{1.5cm}} [g_{21} \dots g_{2N_1}] \\
 &\vdots \xrightarrow{EC(\vec{v}_4, \vec{v}_5, \vec{v}_6)} \vdots \\
 \vec{v}_7 &= [x_7 \ y_7 \ z_7 \ t_7] \xrightarrow{\hspace{1.5cm}} [g_{71} \dots g_{7N_1}]
 \end{aligned}$$

Input:

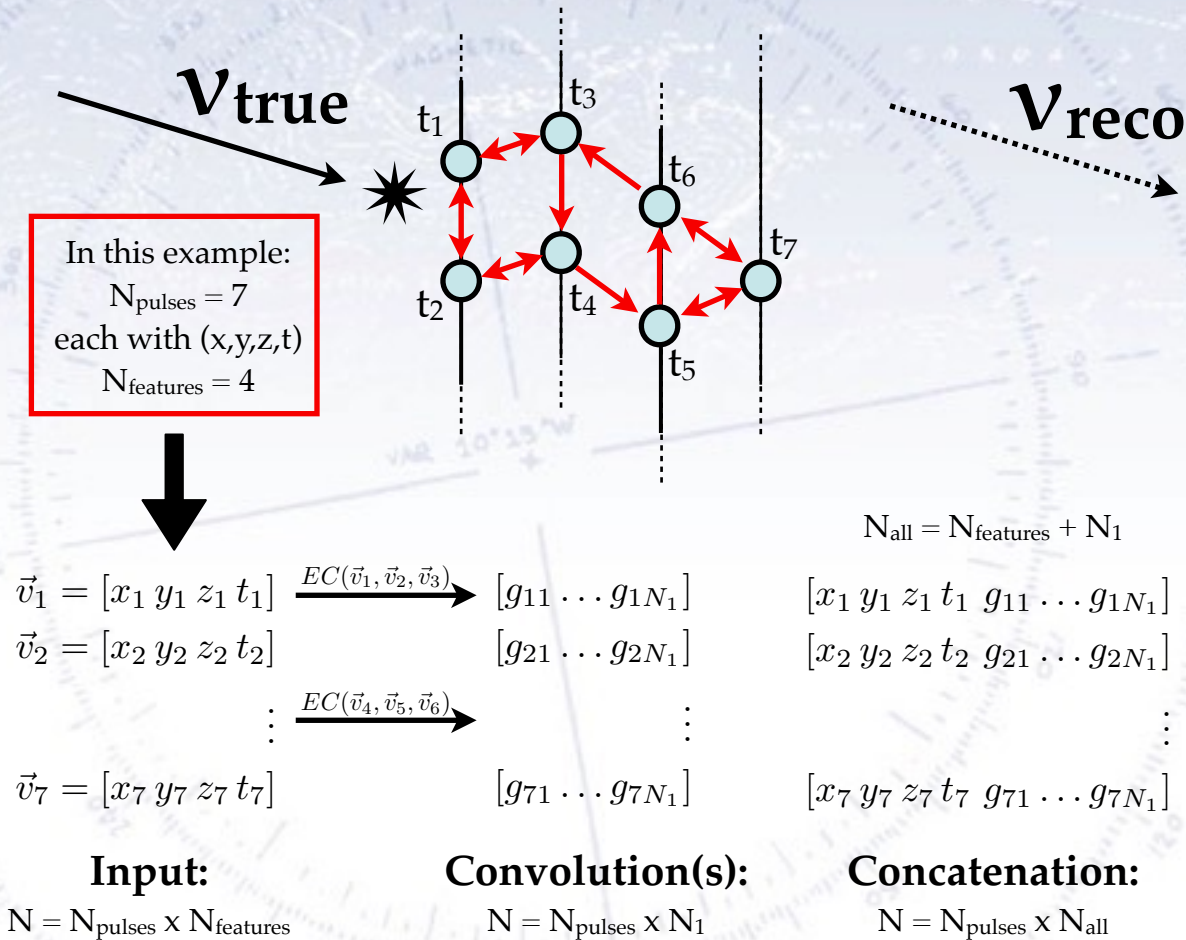
$$N = N_{\text{pulses}} \times N_{\text{features}}$$

Convolution(s):

$$N = N_{\text{pulses}} \times N_1$$

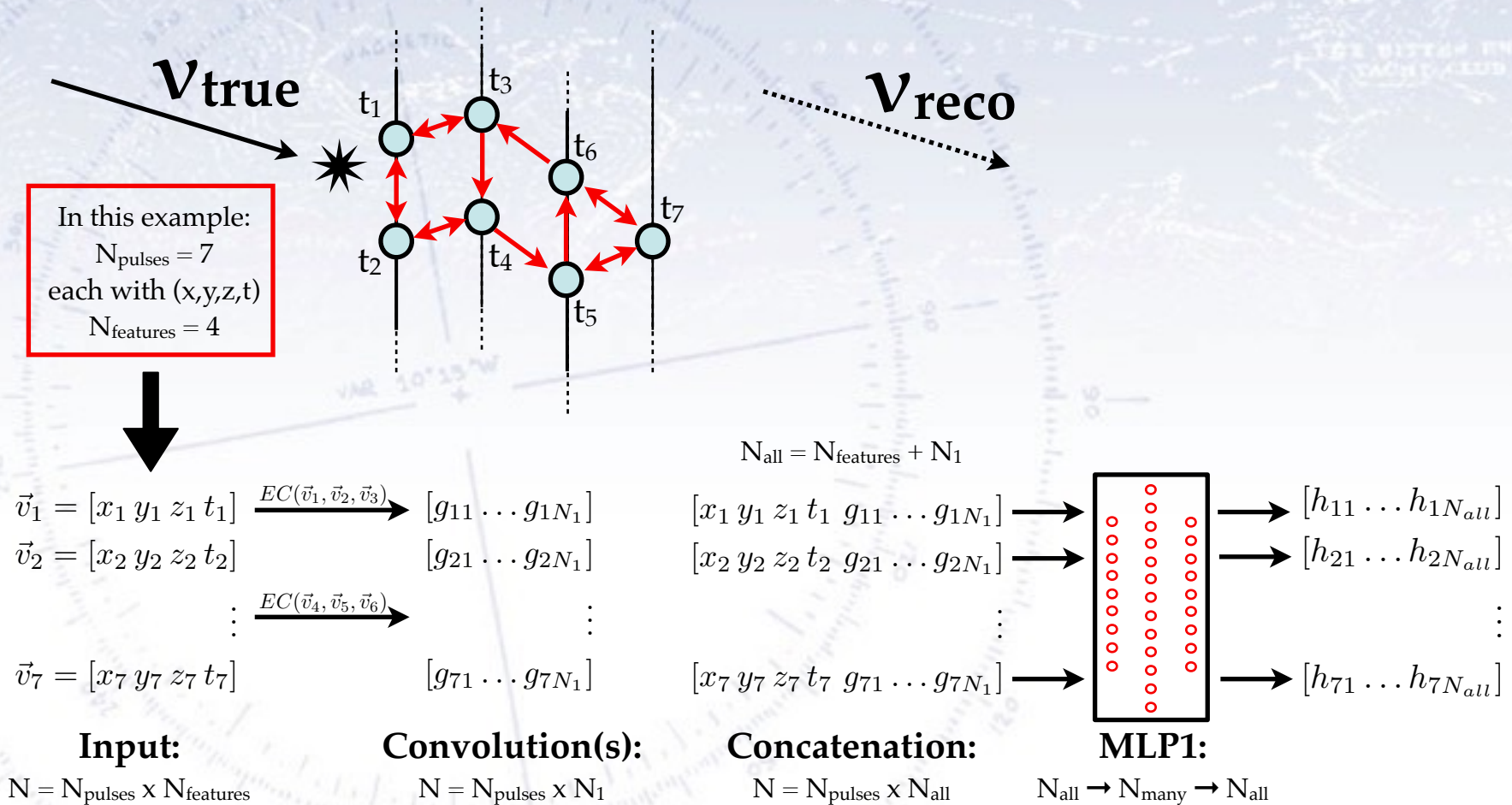
The input features of a node are combined with that of N ($=2$) nearby nodes through an NN (MLP0) function, yielding an (abstract) vector for each node. This can be repeated (not shown).

Details of GNN reconstruction



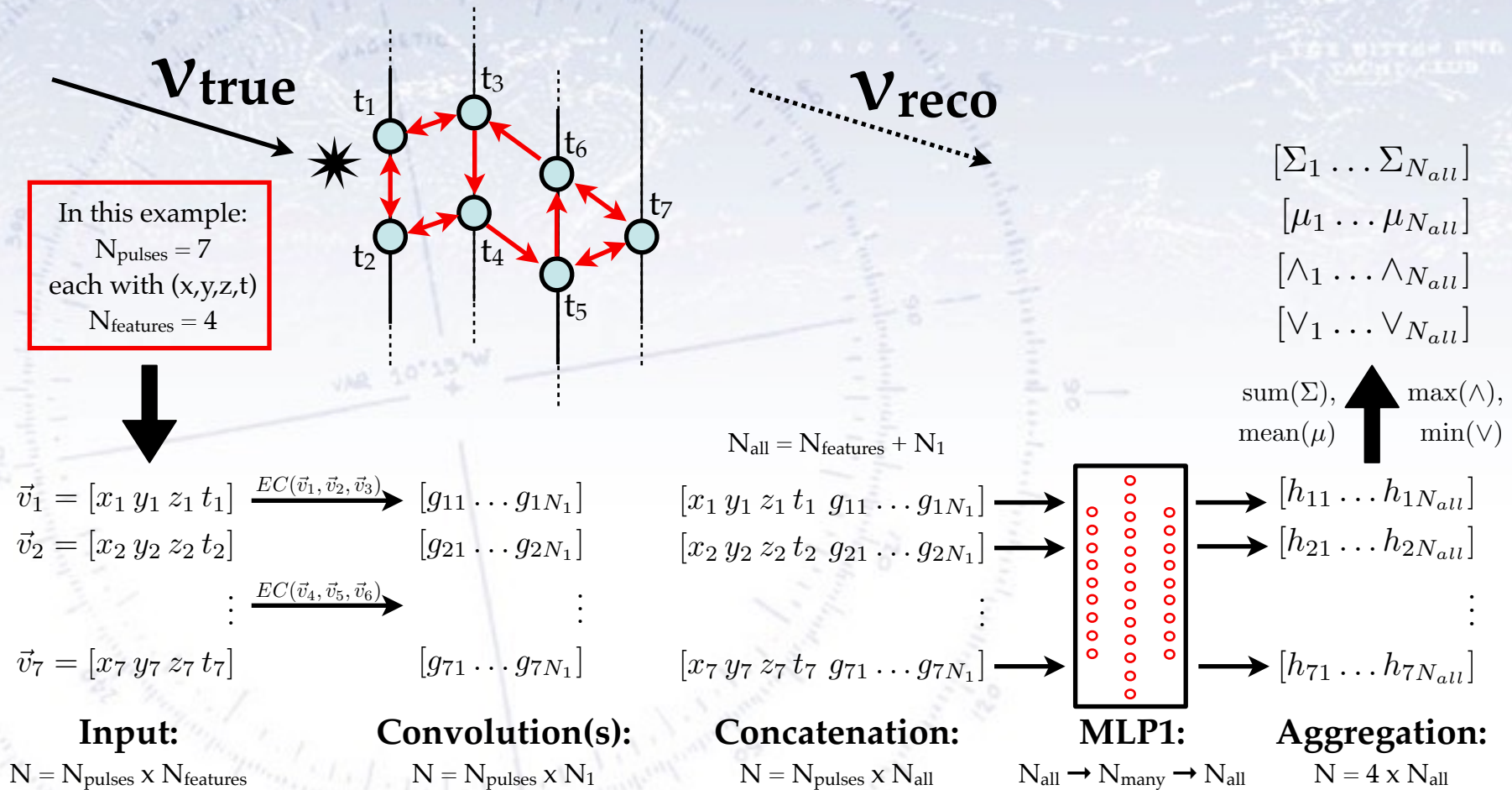
The input features of a node are combined with that of N ($=2$) nearby nodes through an NN (MLP0) function, yielding an (abstract) vector for each node. This can be repeated (not shown). All the features are then combined (concatenated) into long vectors,

Details of GNN reconstruction



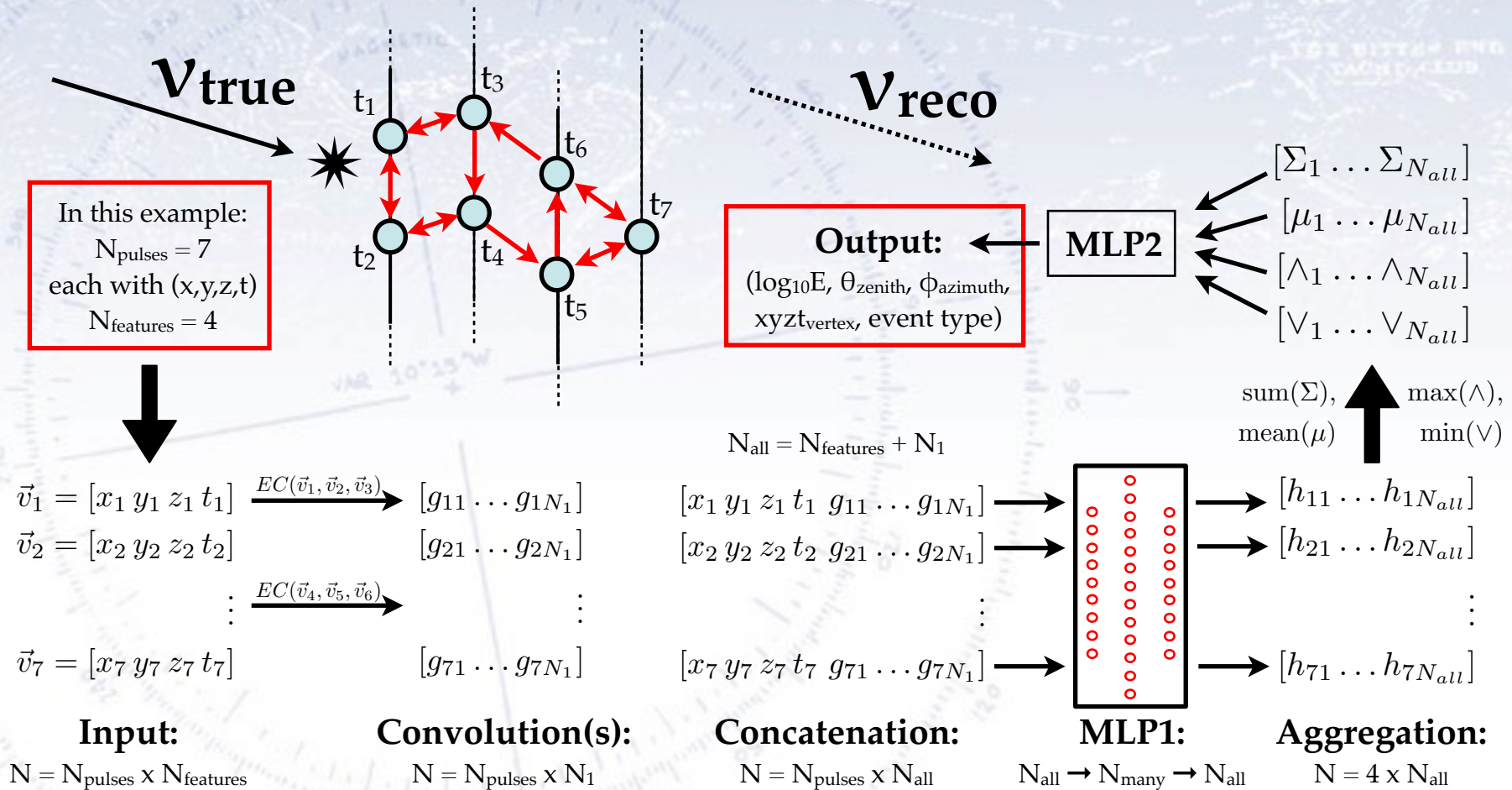
The input features of a node are combined with that of N ($=2$) nearby nodes through an NN (MLP0) function, yielding an (abstract) vector for each node. This can be repeated (not shown). All the features are then combined (concatenated) into long vectors, which are again put through an NN (MLP1) function with a large hidden layer.

Details of GNN reconstruction



The input features of a node are combined with that of N ($=2$) nearby nodes through an NN (MLP0) function, yielding an (abstract) vector for each node. This can be repeated (not shown). All the features are then combined (concatenated) into long vectors, which are again put through an NN (MLP1) function with a large hidden layer. The outputs are aggregated in four ways: Min, Max, Sum & Mean, breaking the variation with number of nodes.

Details of GNN reconstruction



The input features of a node are combined with that of N ($=2$) nearby nodes through an NN (MLP0) function, yielding an (abstract) vector for each node. This can be repeated (not shown). All the features from all the convolutions are then combined (concatenated) into long vectors, which are again put through an NN (MLP1) function with a large hidden layer. The outputs are aggregated in four ways: Min, Max, Sum & Mean, breaking the variation with number of nodes. These are then fed into a final NN (MLP2), which outputs the estimated type(s) and parameters of the event.

Further specifics of DynEdge

In DynEdge, there are several “enlargements” compared to the previous illustration of the GNN architecture. These are essentially:

- We use 6 input features: x , y , z , t , charge, and Quantum Efficiency.
- We convolute each node with the nearest 8 nodes (not two).
- We do 4 (not 1) convolutions, each with 192 inputs and outputs.
- The concatenation is of all convolution layers and the original input.
- In the results to be shown, we have trained separate GNNs for each output.

The repeated convolutions allows all signal parts to be connected.

The EdgeConv convolution operator ensures permutation invariance.

The number of model parameters is about 750.000 for the angular regressions, while the energy only requires 150.000. In principle one can go down to 70.000 parameters, but there is no reason for this - it is already a “small” ML model.

GraphNet

The GNN model is outlined more simply below, which is also the (current) figure for the GNN paper in process.

