

Dual-Comb Spectroscopy of Heavy Metals (New Zr Oscillator Strengths!)

Hope Danner¹, Andrew Jarymowycz¹, Kirsten Dowd²,
Padraig Dunne², and R. Jason Jones¹

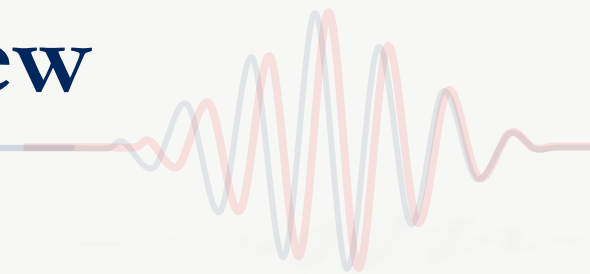
¹ College of Optical Sciences, University of Arizona, Tucson, AZ 85721, USA

² School of Physics, University College Dublin, Belfield, Dublin 4, Ireland



THE UNIVERSITY OF ARIZONA
Wyant College
of Optical Sciences





- **Experimental Setup**

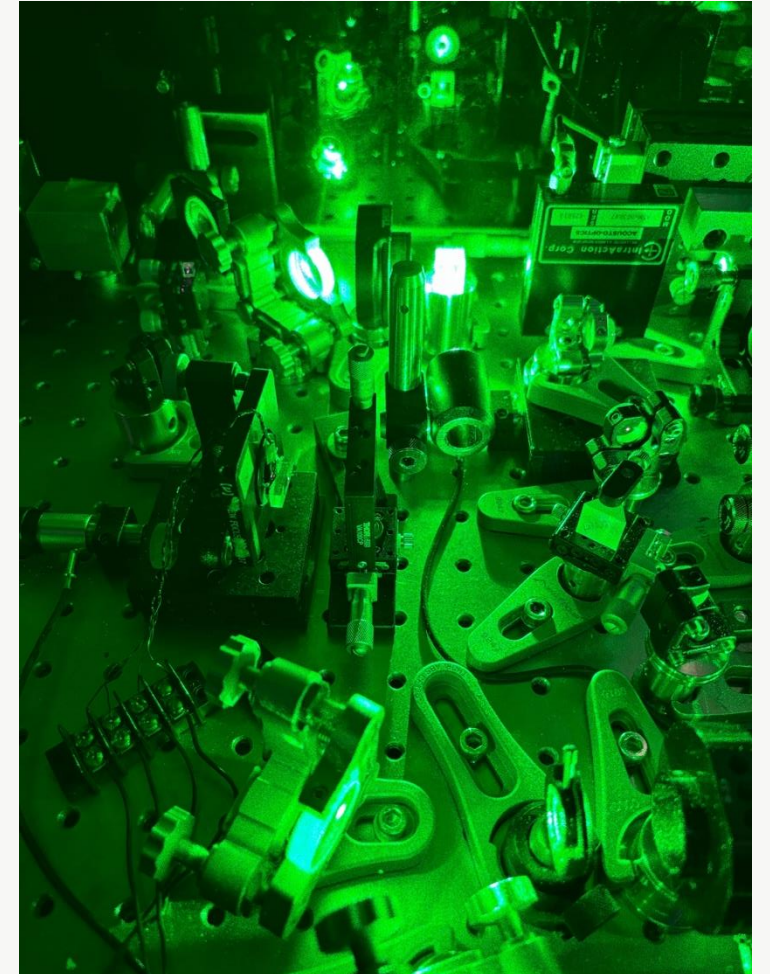
- Fundamental and second harmonic of Ti:s **simultaneously** probe solid samples using a laser produced plasma

- **Analysis**

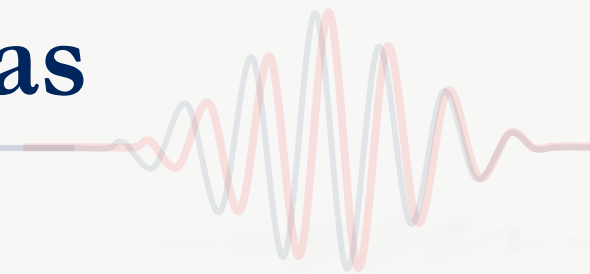
- Line Identifications (**tangent on databases**)
- Plasma temperature and column densities
- Extraction of **oscillator strengths**

- **Moving Forward**

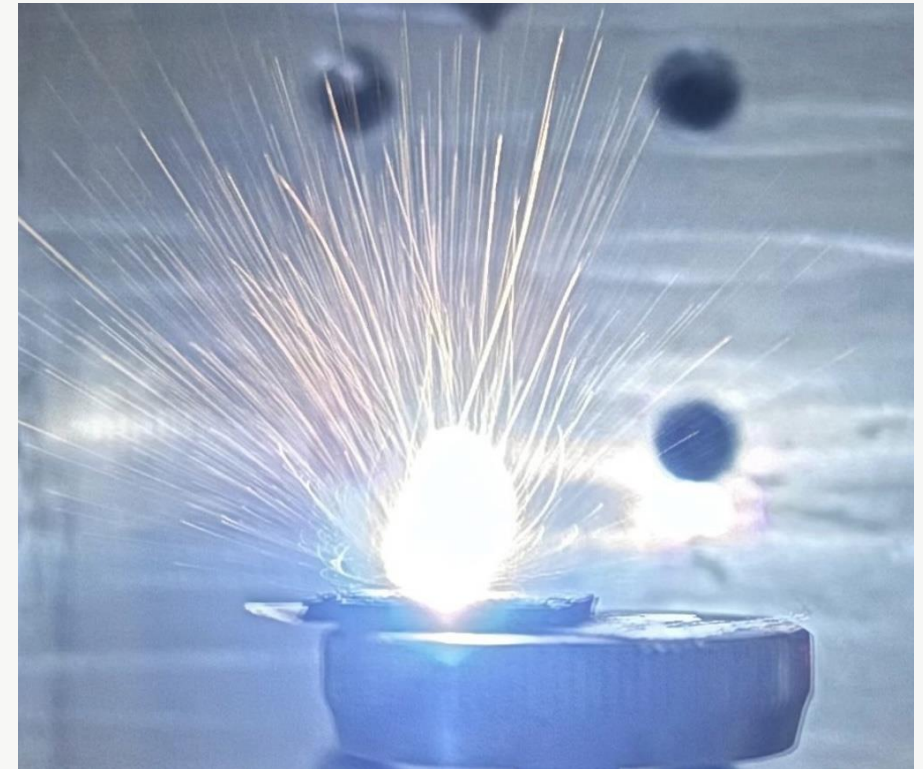
- Larger spectrum
- Alloys as an anchor



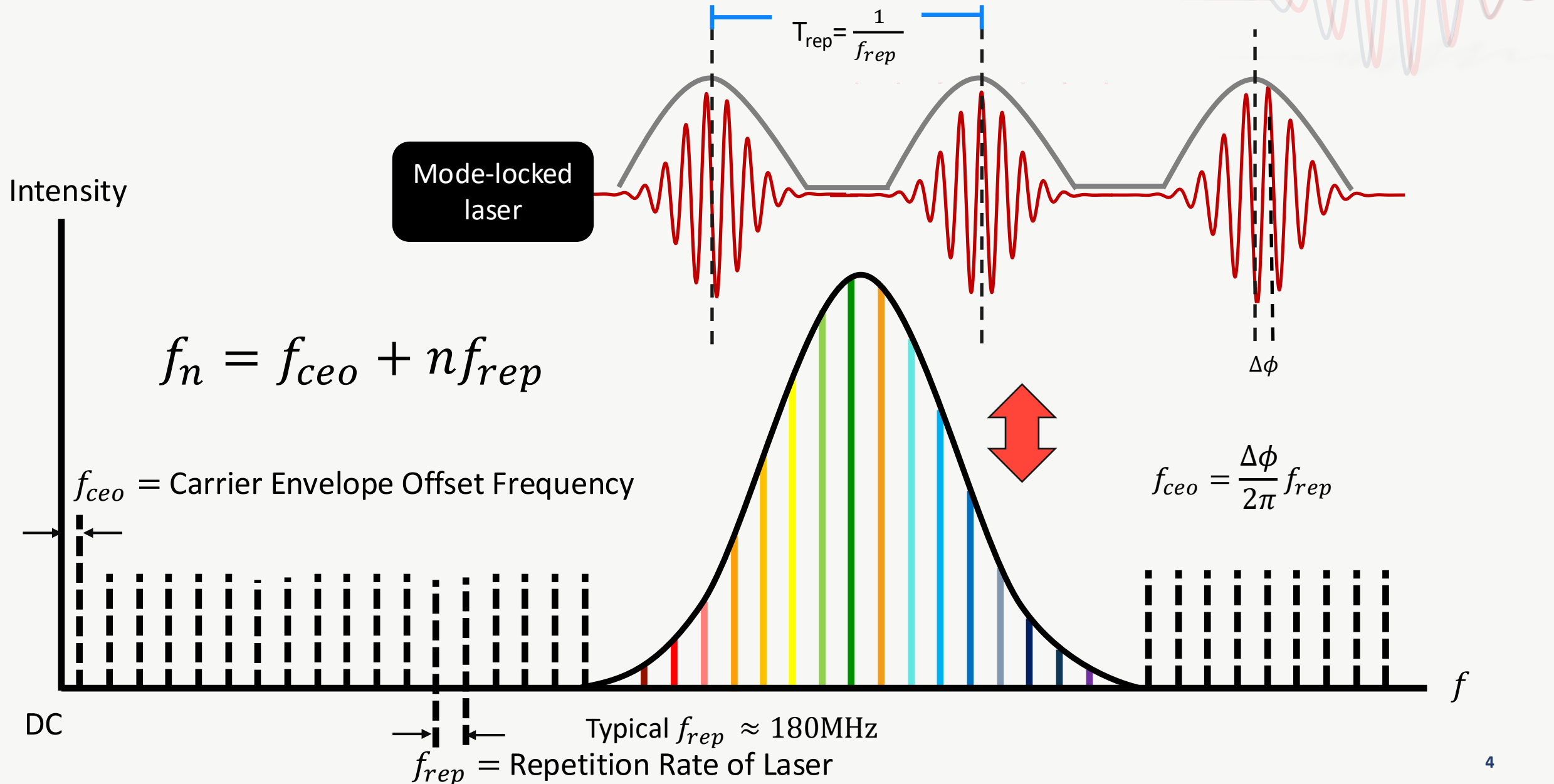
Laser Produced Plasmas



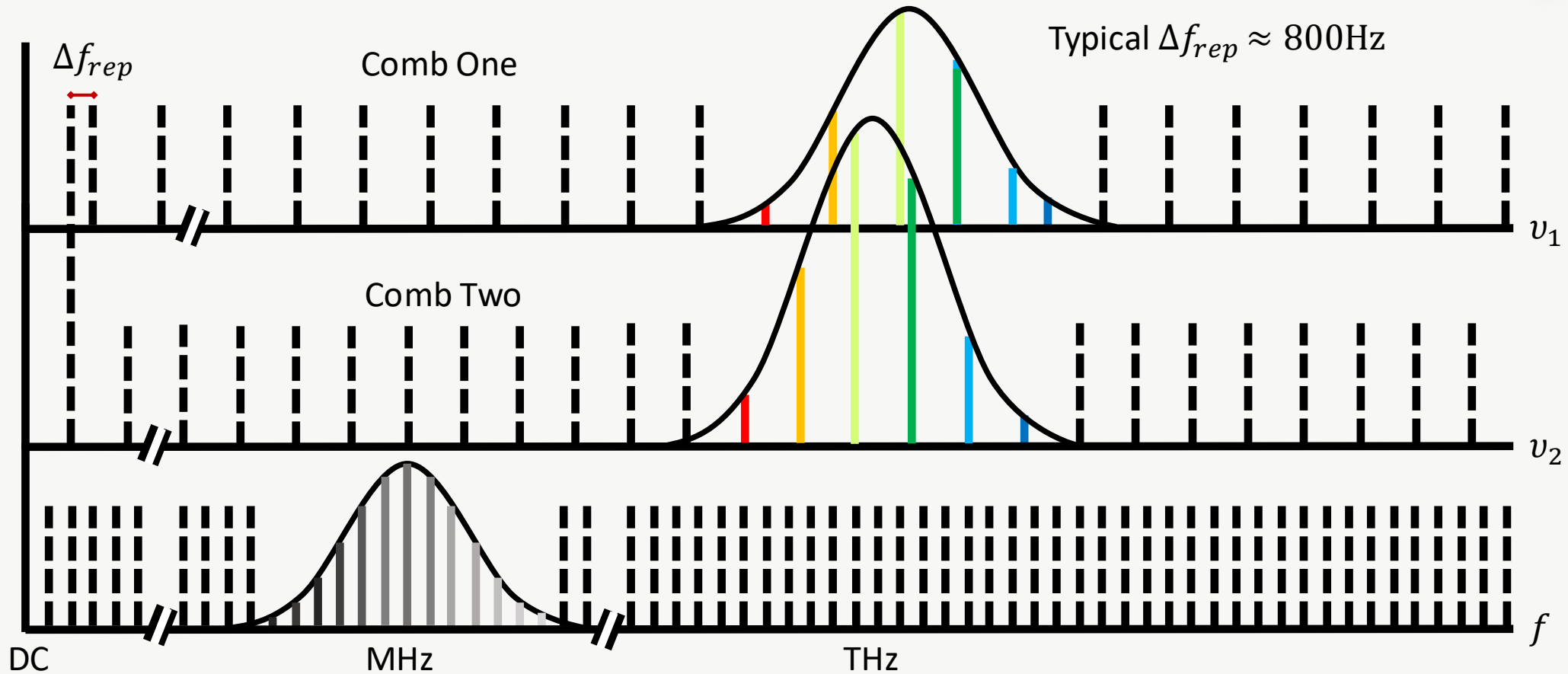
- Convenient sample preparation
- Fundamental atomic/optical properties
 - e.g. transitions and osc. strengths
- Laser-Plasma dynamics & fundamentals
 - ns-ms time scales
 - temperature $10,000+ \text{ K} \rightarrow 500 \text{ K}$
 - number densities $10^{21} \rightarrow 10^{13} \text{ cm}^{-3}$
 - species formation, plasma-chemistry
 - Tests of LTE conditions



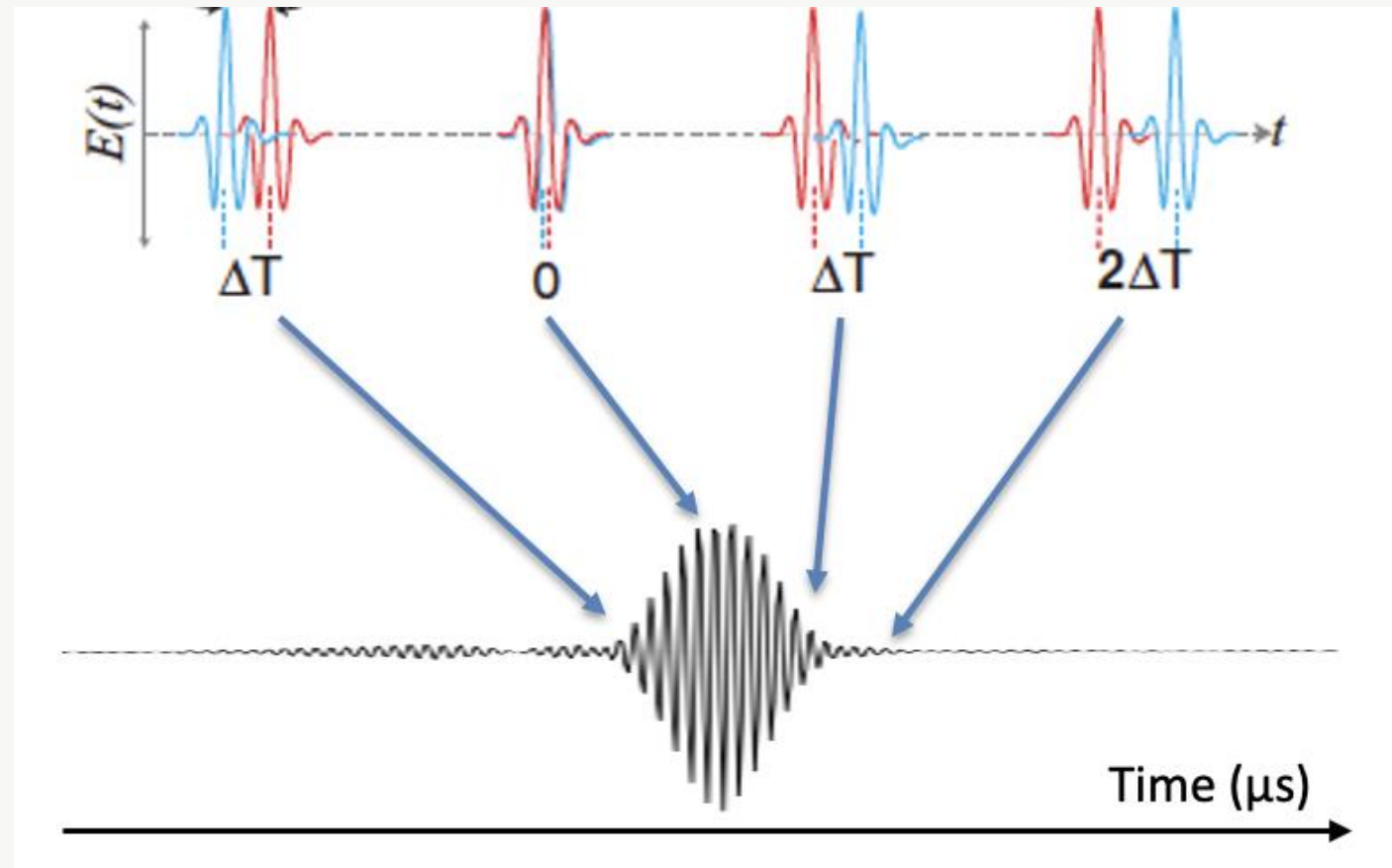
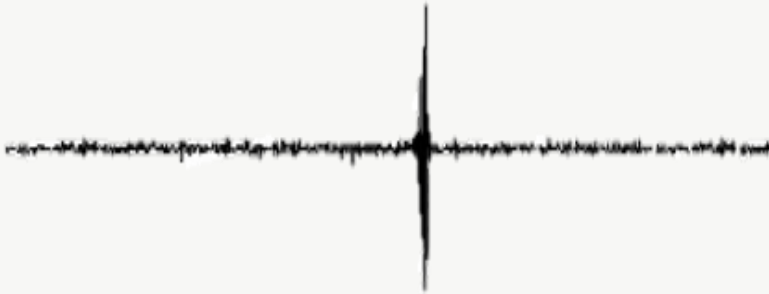
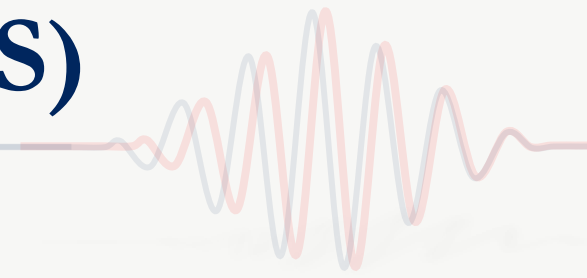
Frequency Combs



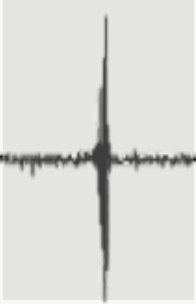
Dual-Comb Spectroscopy (DCS)



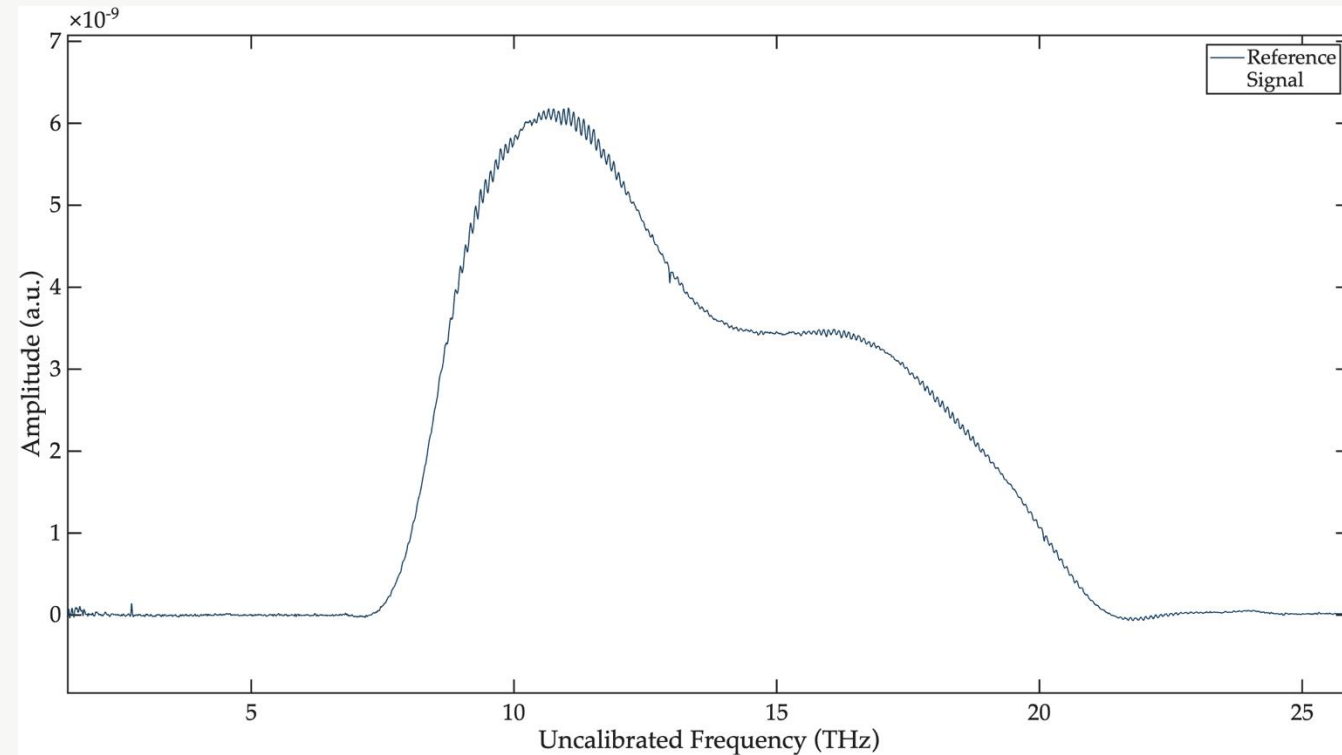
Dual-Comb Spectroscopy (DCS)



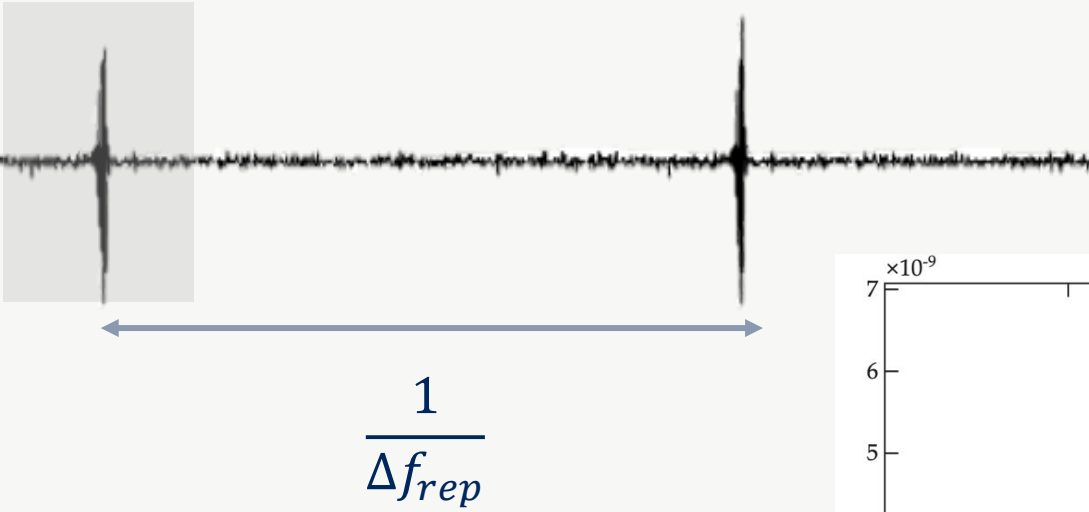
Fourier Transform spectroscopy



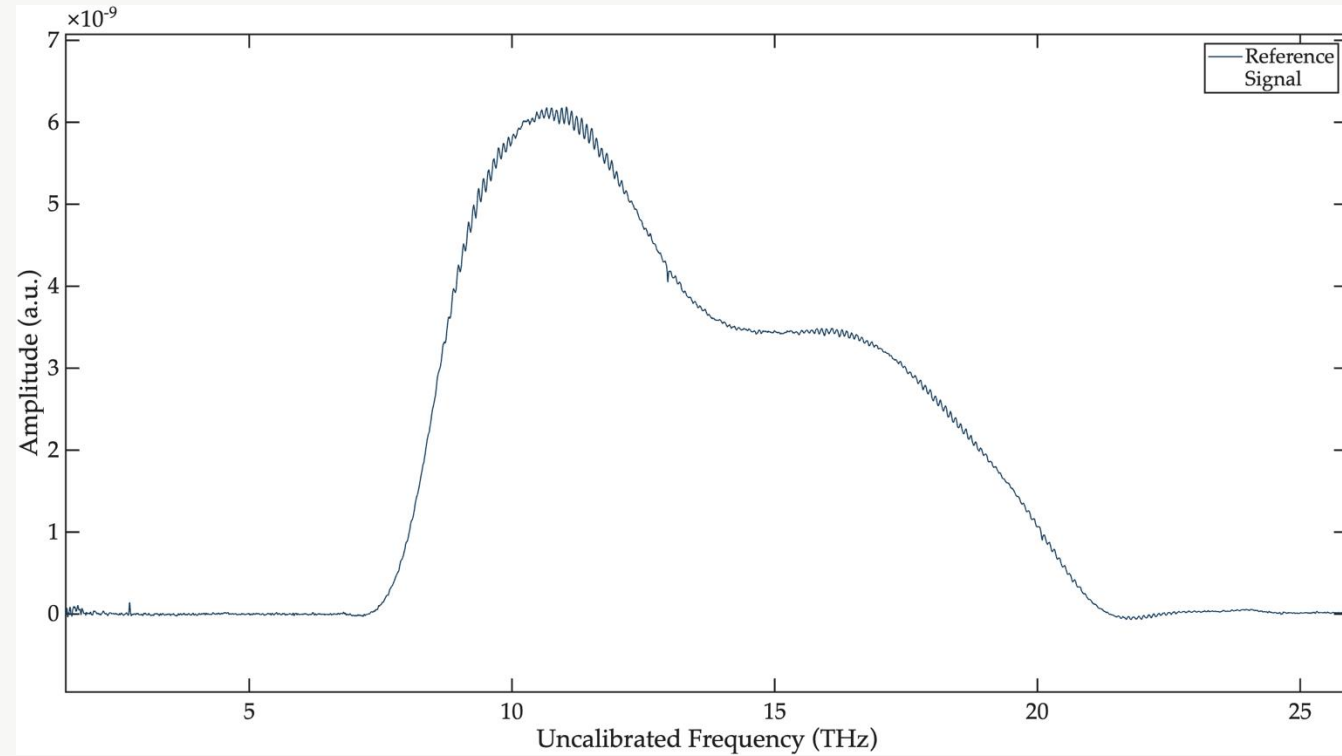
- Similar to Fourier Transform spectroscopy but no longer limited by a mechanical scanning arm!



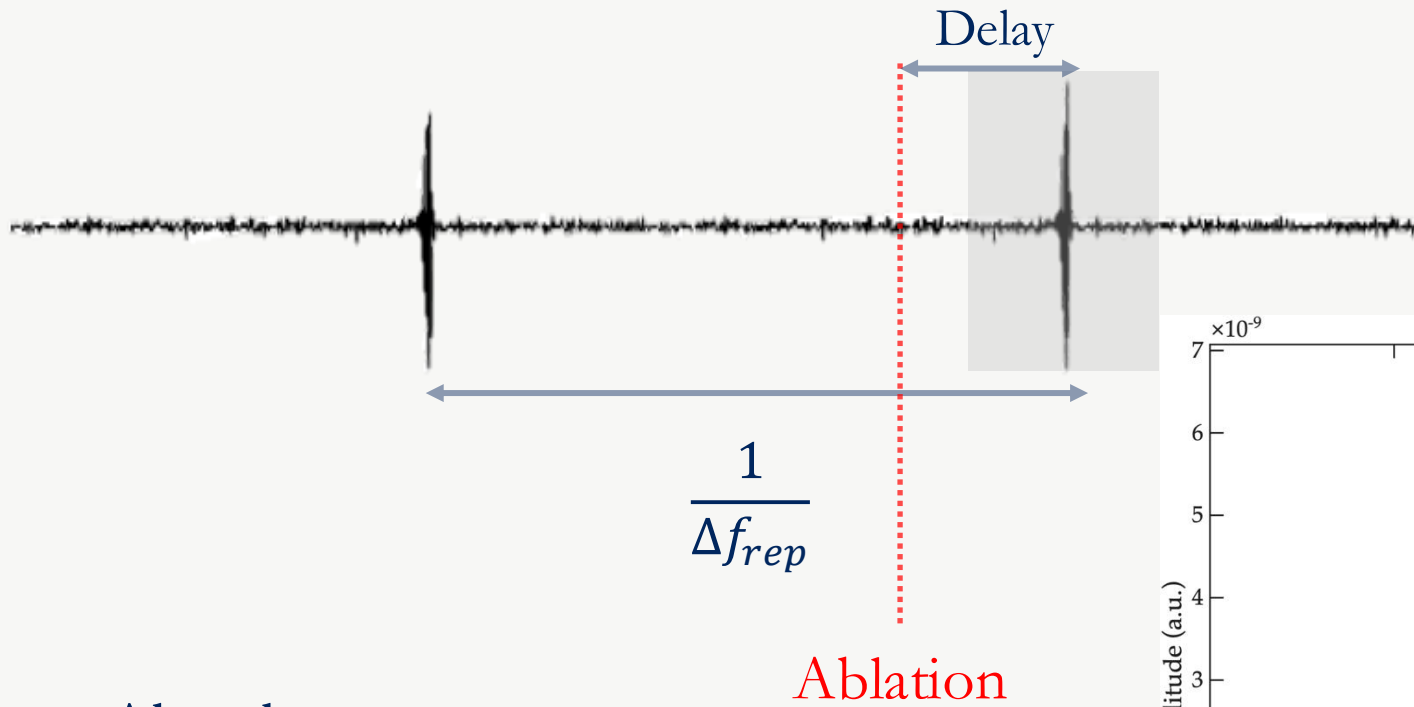
DCS Signal



- $\Delta f_{rep} = 800$ Hz
- Able to collect both reference and signal

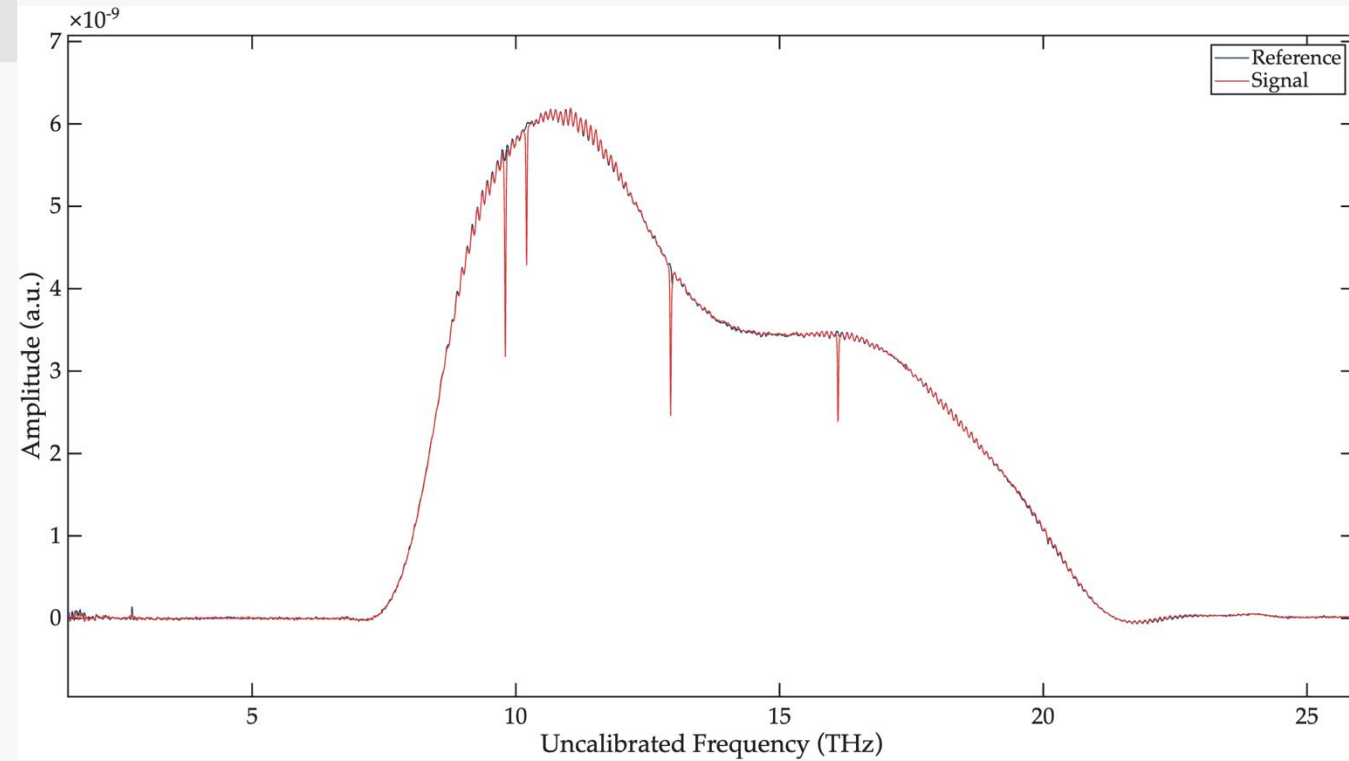


DCS Reference

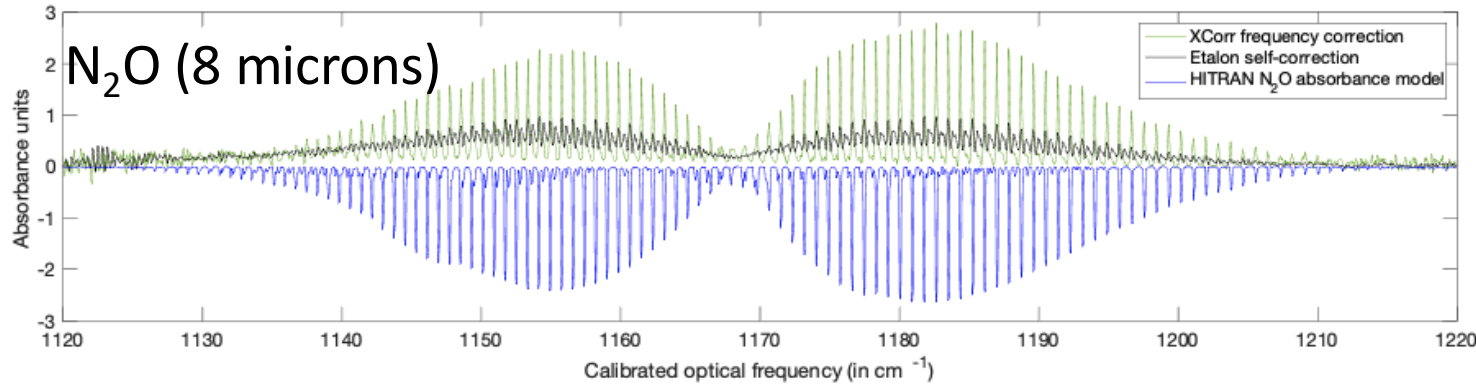


Absorbance:

$$A(\nu) = -\ln\left(\frac{I}{I_o}\right)$$

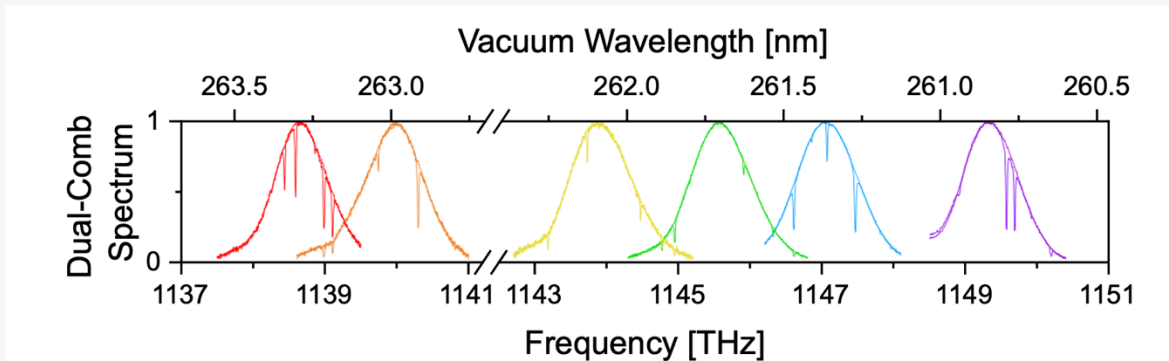


Dual Comb Spectroscopy: MIR to the VUV

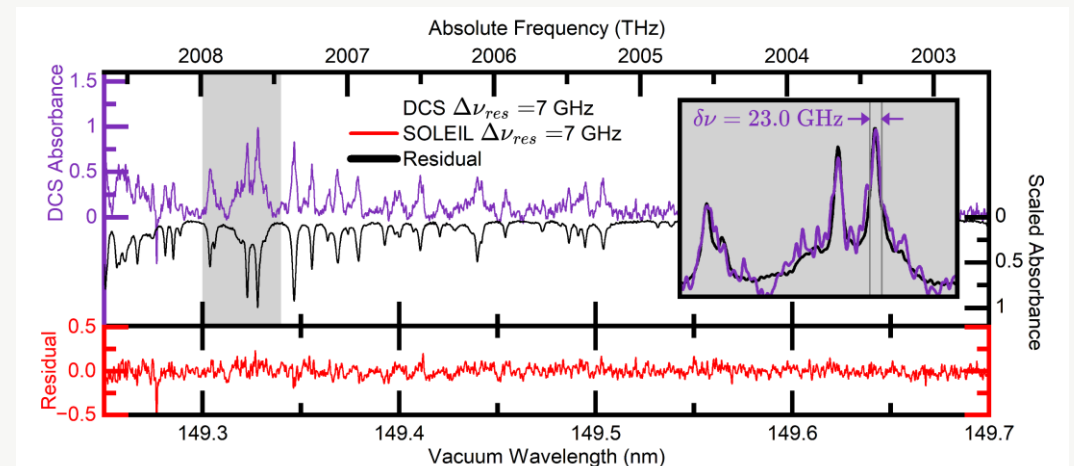


Resolving Power $> 16,000,000$

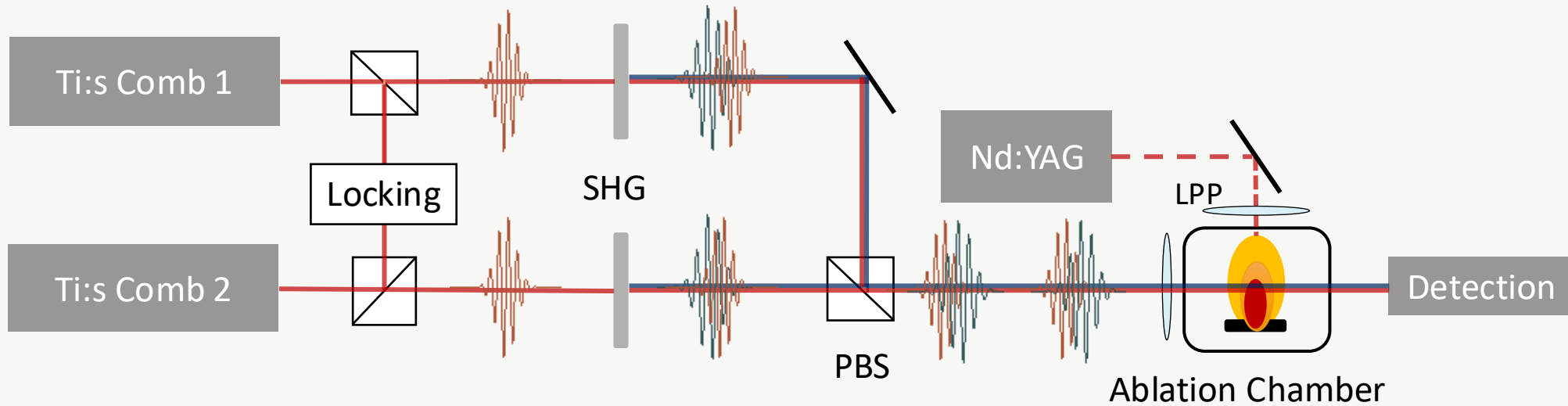
Fe in the DUV (260 nm)



ammonia VUV spectroscopy (148nm)

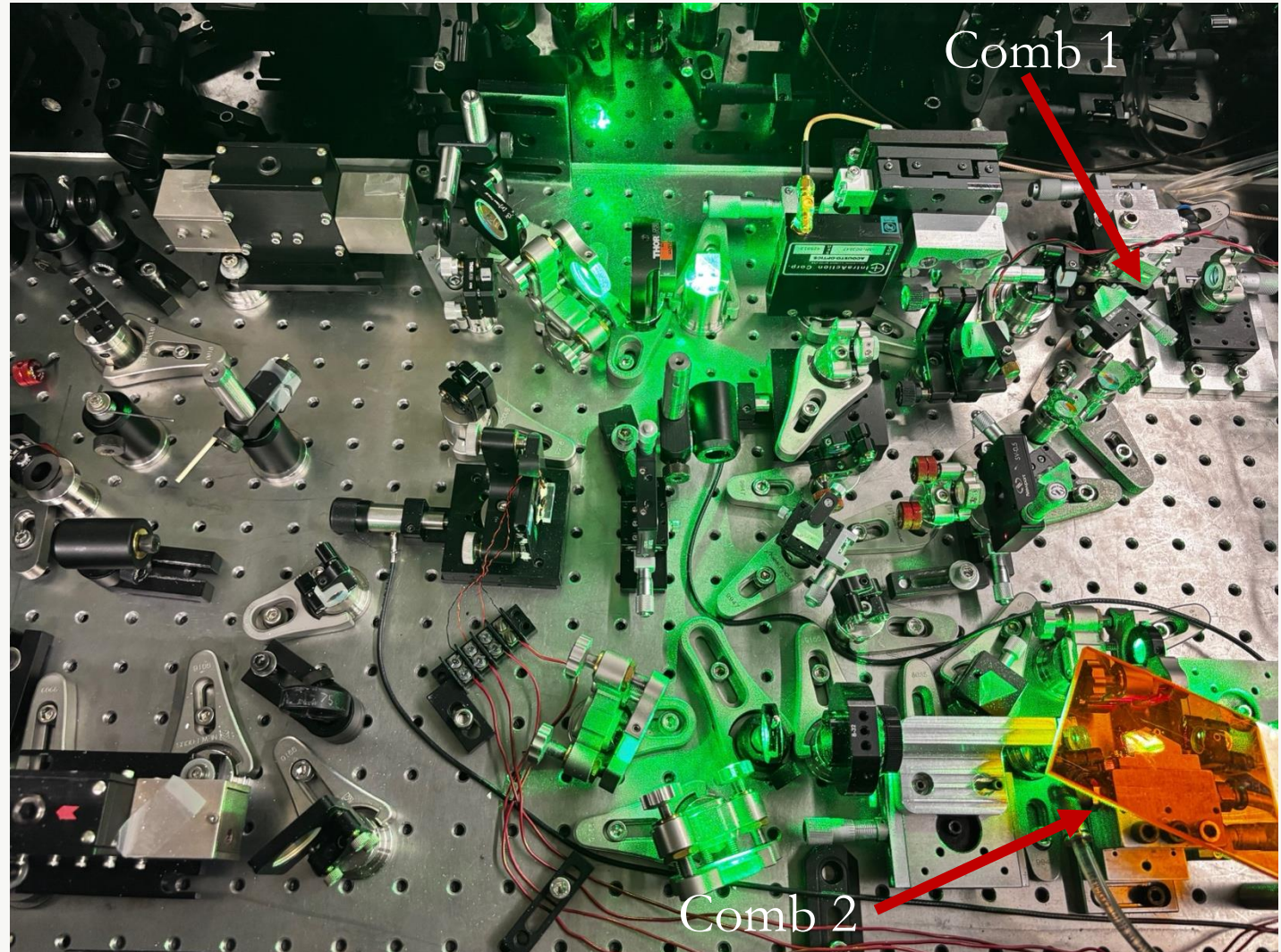


Experimental Setup

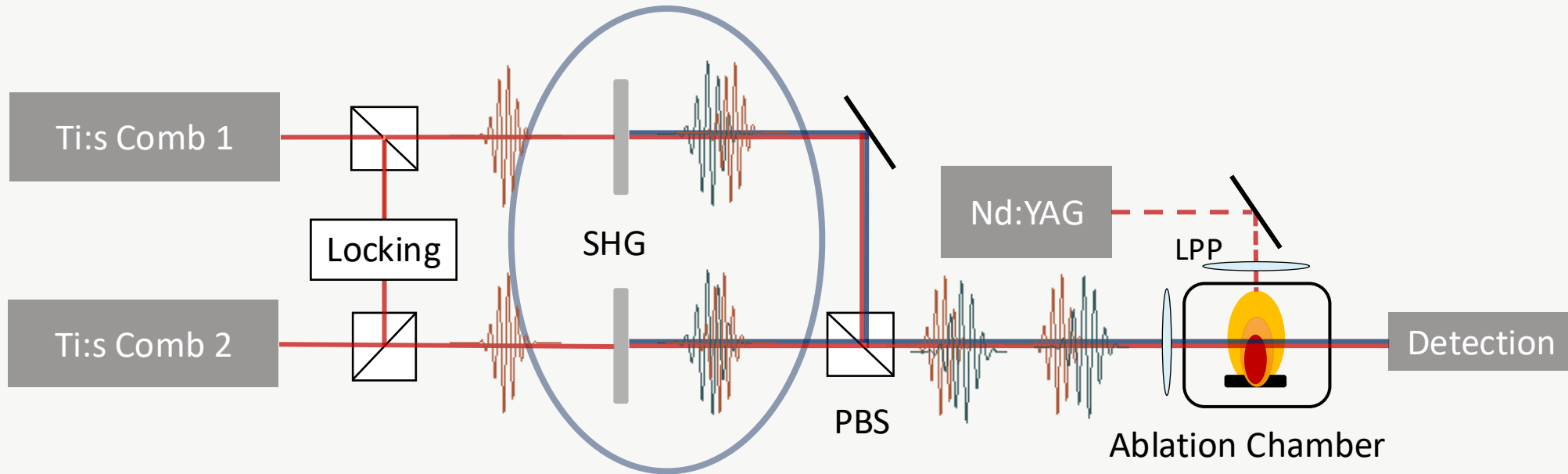


Home-Built Ti:s Combs

- 800nm center wavelength
- FWHM $\sim 20\text{nm}$
- Tunable over 50nm
- Rep Rate: 180 MHz

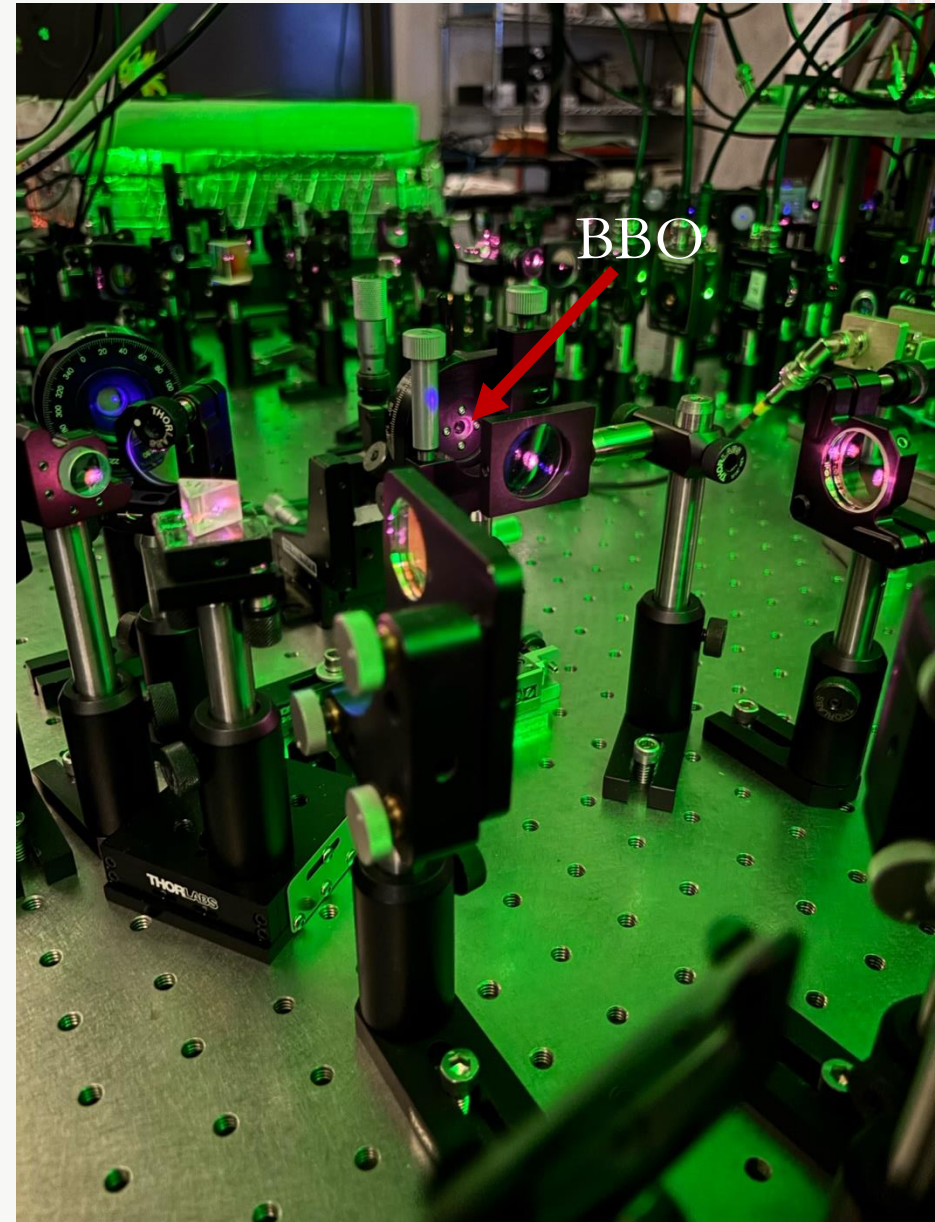


Second-Harmonic Generation

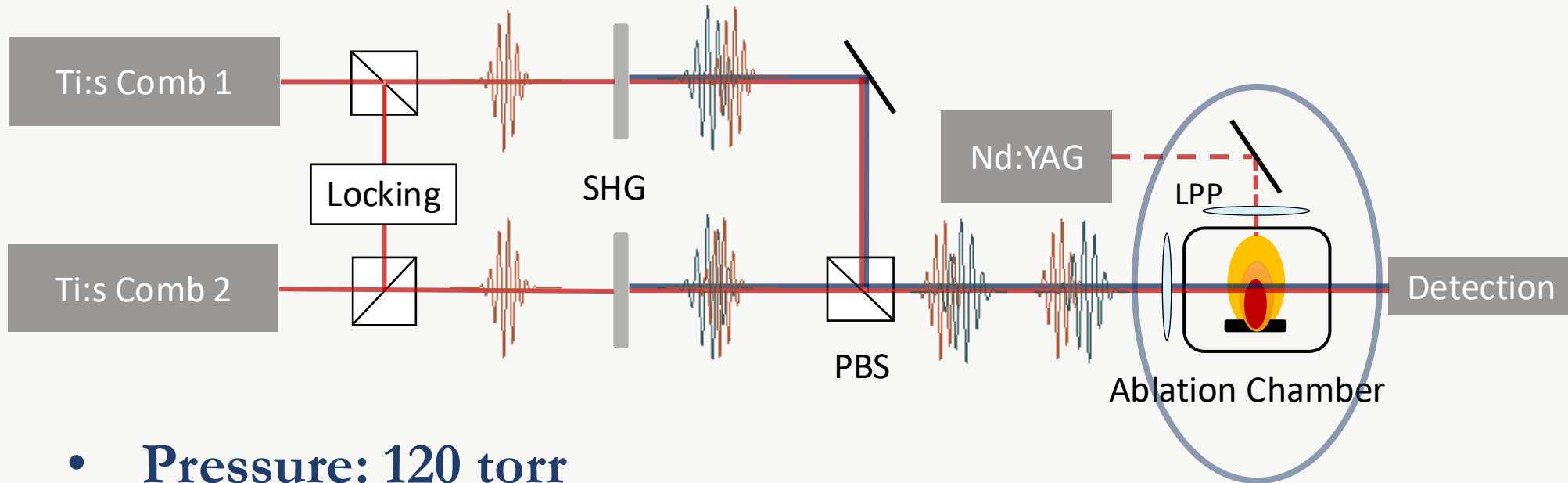


Second-Harmonic Generation

- Two BBOs used to frequency double the Ti:s combs from 800nm to 400nm
- FWHM $\sim 10\text{nm}$



Laser Produced Plasma Conditions



- **Pressure: 120 torr**
- **Beam Height: 2.5mm**
- **Flowing with 35 lpm of Ar**



Current Wavelength Range

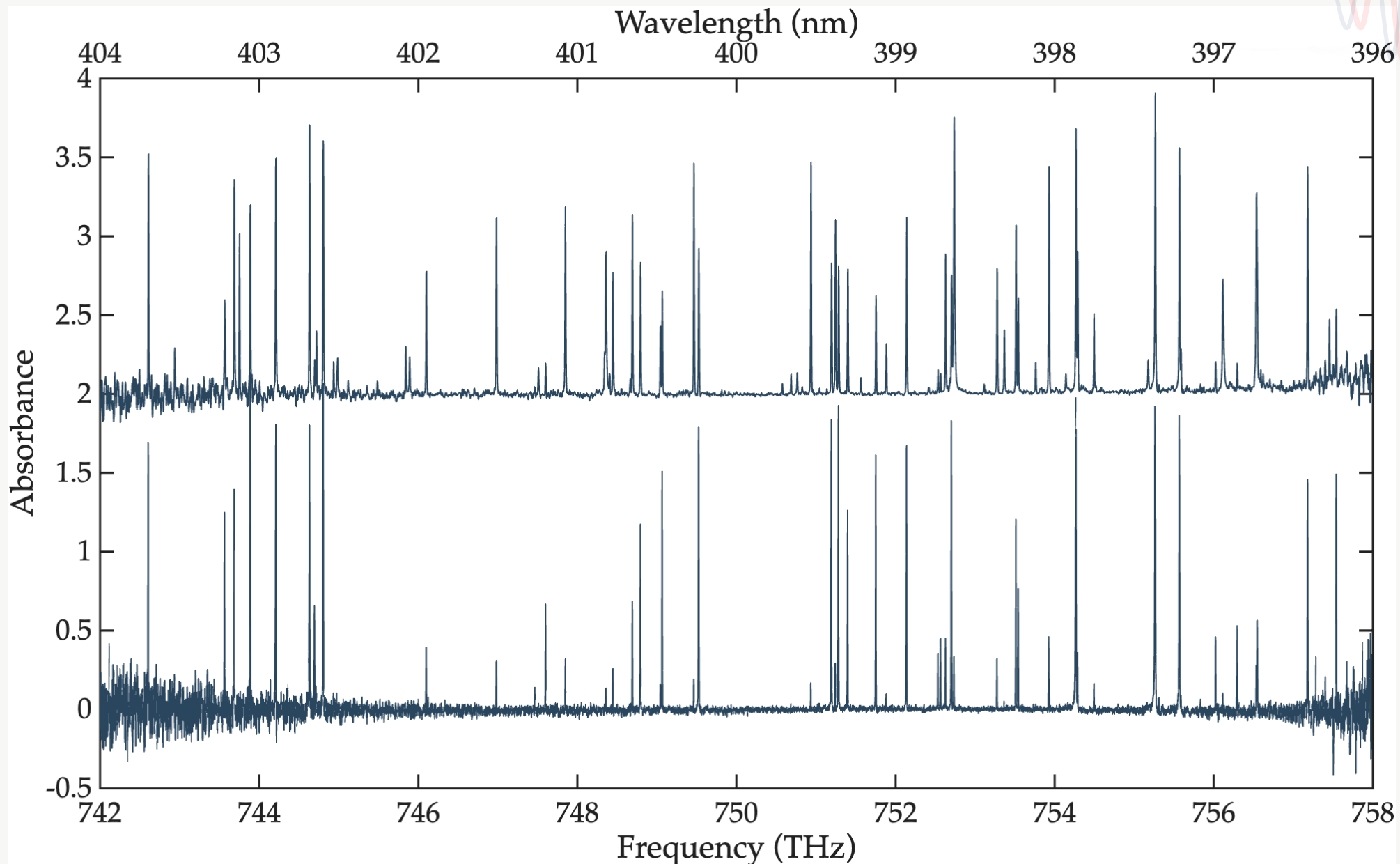


Data I Will Be Showing Today



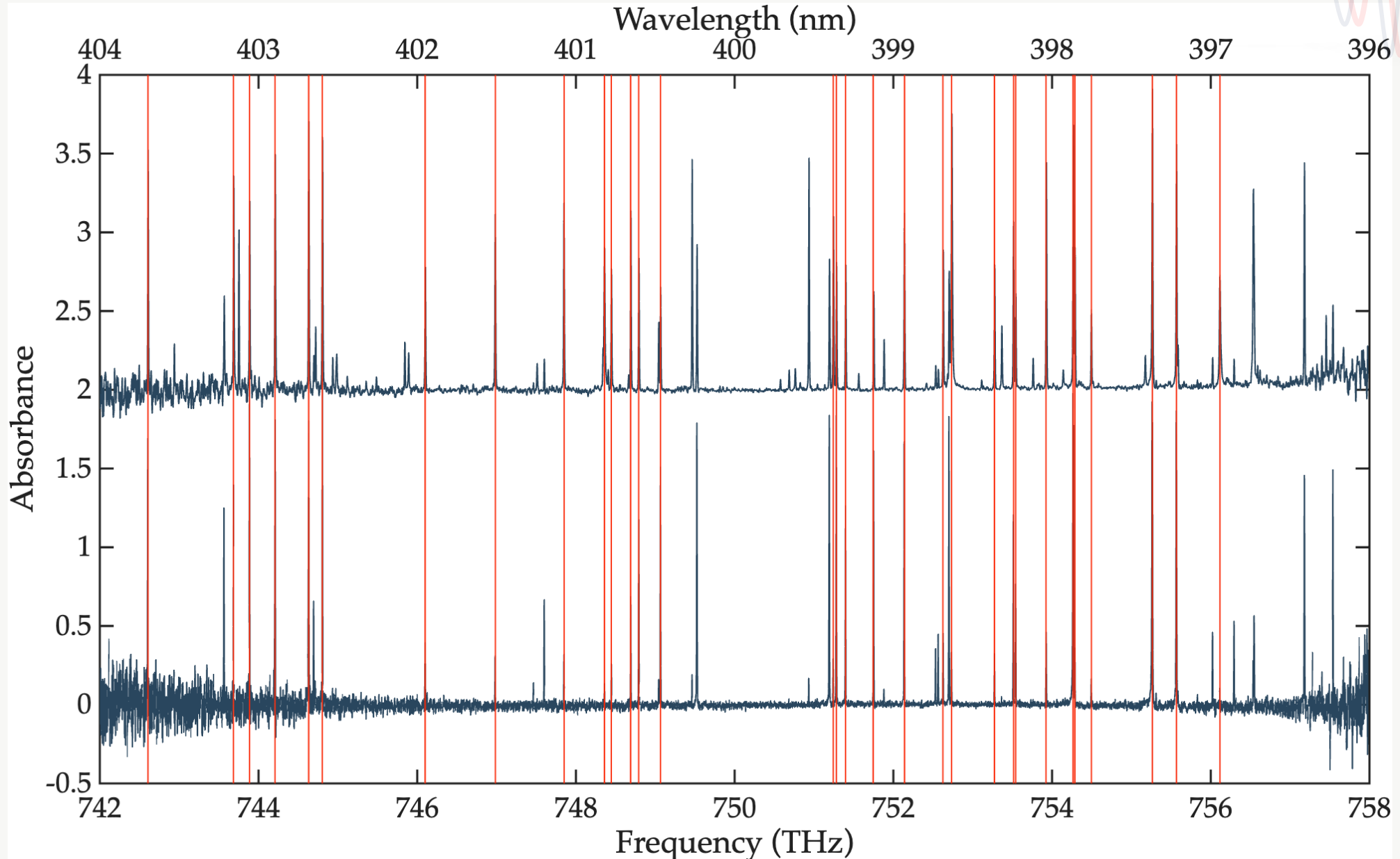
Measured Zr absorbance

Resolution
~ 0.5 pm



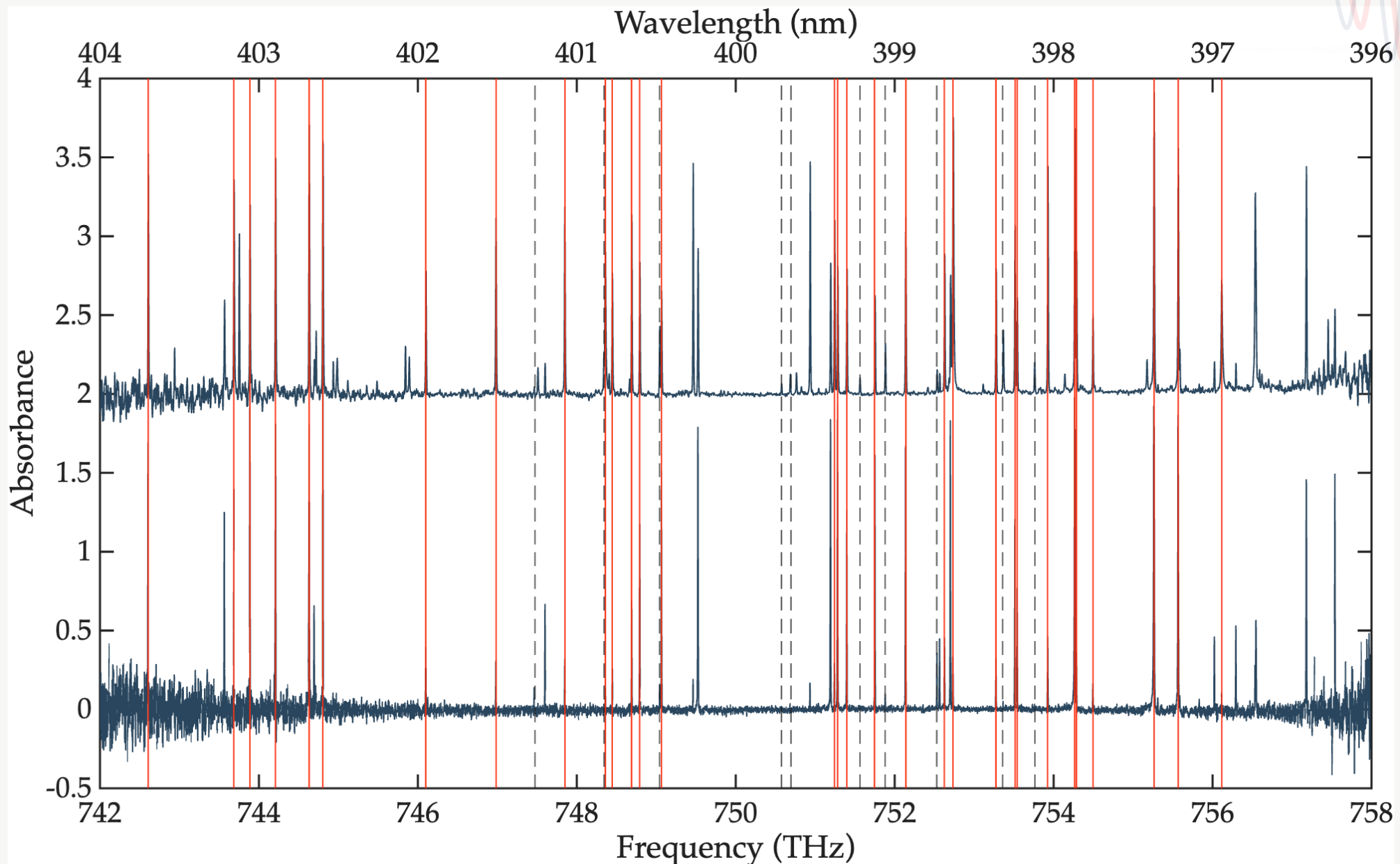
Red lines – Previously Identified Transitions

Resolution
~ 0.5 pm



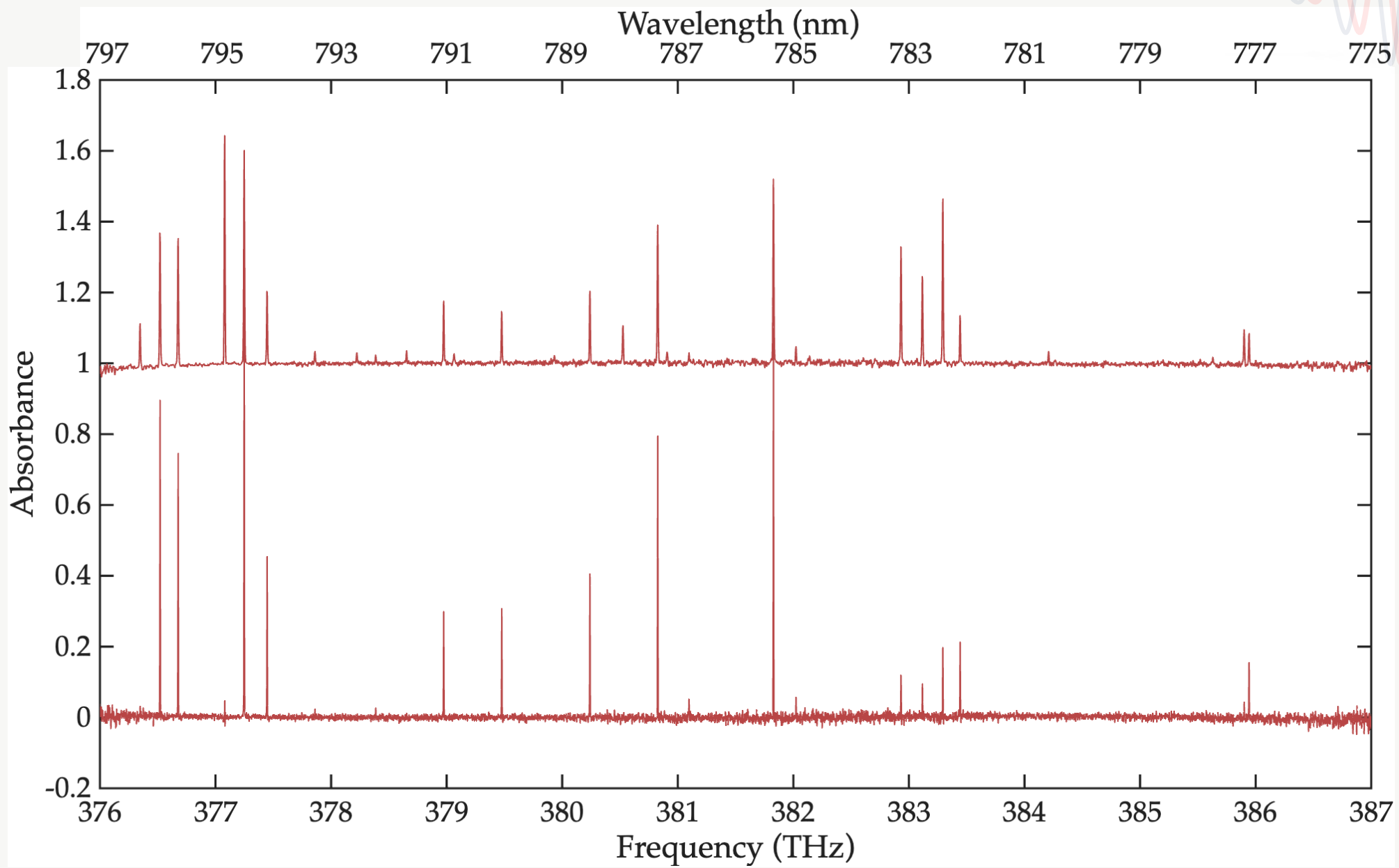
Dashed lines – Newly Identified Transitions

Resolution
~ 0.5 pm



Measured Zr absorbance

Resolution
~ 0.5 pm

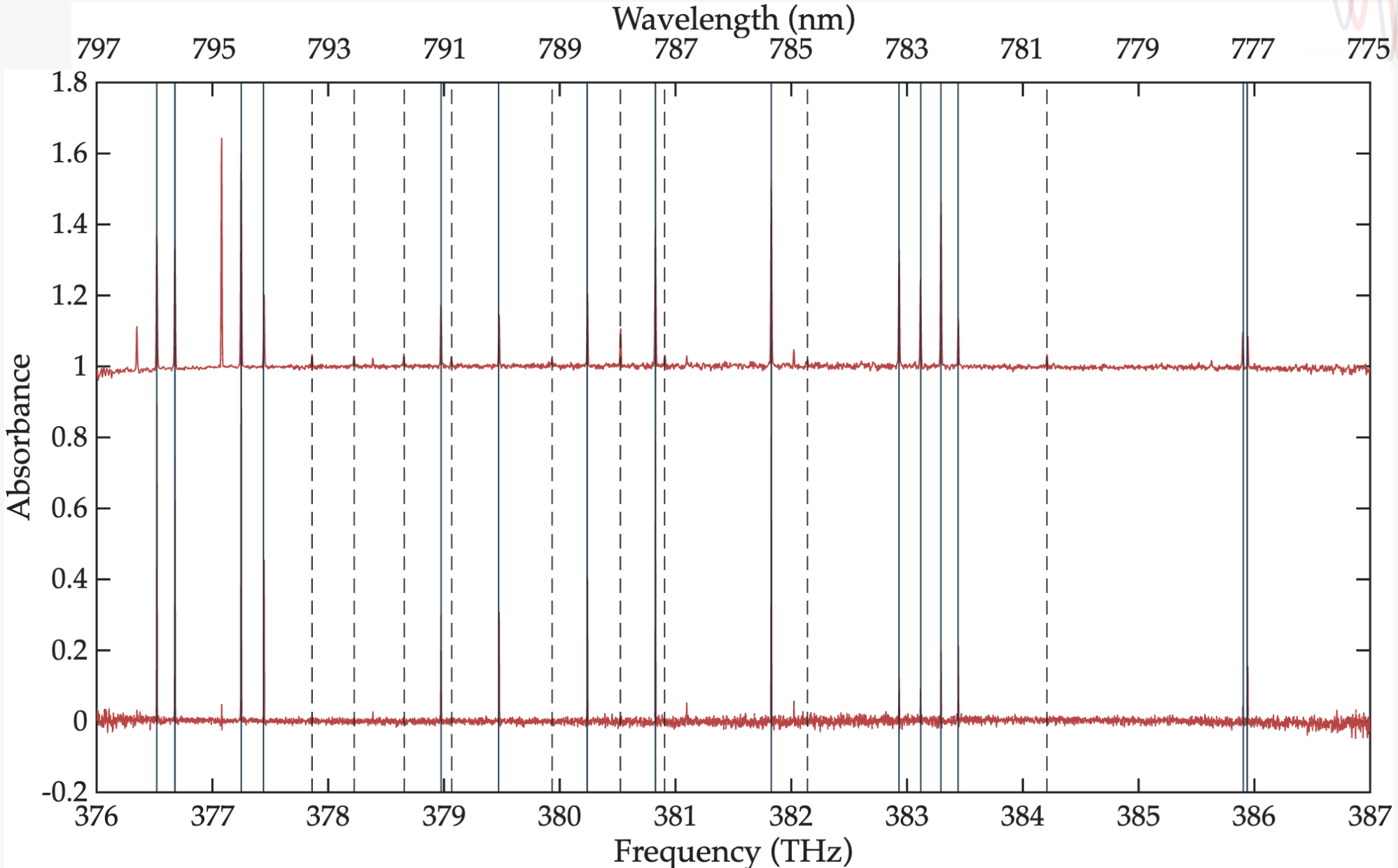


$50\mu s$

$150\mu s$

Identified Transitions

Resolution
~ 0.5 pm



50 μ s

150 μ s

Timeline of Key Zr Spectroscopy Studies

THE STRUCTURE OF THE ARC SPECTRUM OF ZIRCONIUM

By C. C. Kiess and Harriet K. Kiess

1930

1931

THE STRUCTURE OF THE SPECTRUM OF SINGLY IONIZED ZIRCONIUM

By C. C. Kiess and Harriet K. Kiess

ATOMIC ENERGY LEVELS

As Derived From the Analyses of Optical Spectra

Volume II

The Spectra of Chromium, Manganese, Iron, Cobalt, Nickel, Copper, Zinc, Gallium, Germanium, Arsenic, Selenium, Bromine, Krypton, Rubidium, Strontium, Yttrium, Zirconium, and Niobium

BY CHARLOTTE E. MOORE

1952

Tables of Spectral-Line Intensities Part I - Arranged by Elements

Second Edition

William F. Meggers, Charles H. Corliss and
Bourdon F. Scribner

Institute for Basic Standards
National Bureau of Standards
Washington, D.C. 20234

1961

Expanded - 1975

1962

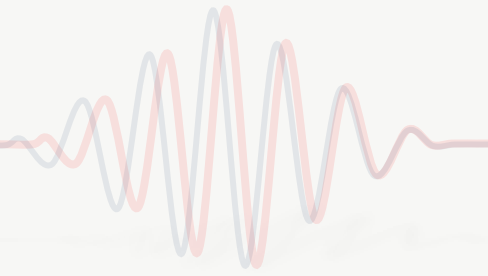
Experimental Transition Probabilities for Spectral Lines of Seventy Elements

Derived From the
NBS Tables of Spectral-Line Intensities

The wavelength, energy levels,
transition probability, and oscillator strength
of 25,000 lines between 2000 and 9000 Å
for 112 spectra of 70 elements.

Charles H. Corliss and William R. Bozman

Key Lessons



- **NIST does not have everything (this is known)**
 - Surprise: data from the 30s (where the energy levels were originally derived from - Moore) are left out of NIST.
- **Kurucz wavelengths are theory**
 - There are papers that are listed but this is **just for the oscillator strengths**.
 - Kurucz calculates energy levels and gives the Ritz transitions for lines that have oscillator strengths from experiments.
- **A lot of Zr I & II oscillator strengths are derived from relative intensities.**
 - There was a large dataset (Meggers) that had a known number density and arc temperature. The intensities were well characterized.
 - Then Corliss and Bozman calculated the oscillator strengths.
 - **These have been shown to have a large wavelength and intensity dependence.**



Boltzmann Plot Analysis

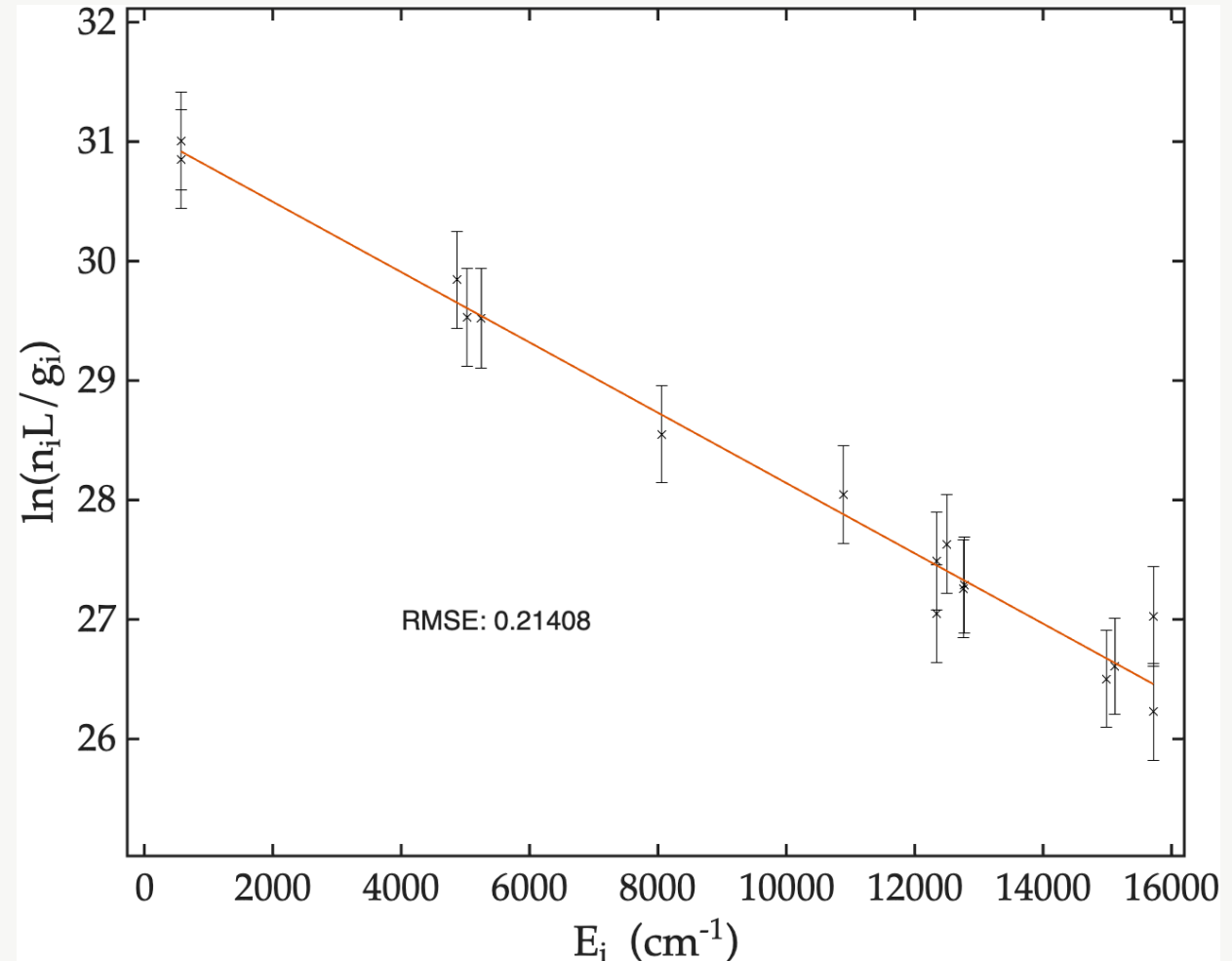
- Calculate individually for each time delay
- Assumes LTE

Boltzmann:

$$n_i = g_i \cdot n_{total} \cdot \frac{e^{\frac{-E_i}{k_B T}}}{Z(T)}$$

$$\ln\left(\frac{n_i L}{g_i}\right) = \ln\left(\frac{Area_{ij}}{\tilde{\sigma}_0 g_i f_{ij}}\right) = \ln\left(\frac{n_{tot} L}{Z(T)}\right) - \frac{E_i}{k_B T}$$

$$y = mx + b$$



Obtaining New Oscillator Strengths

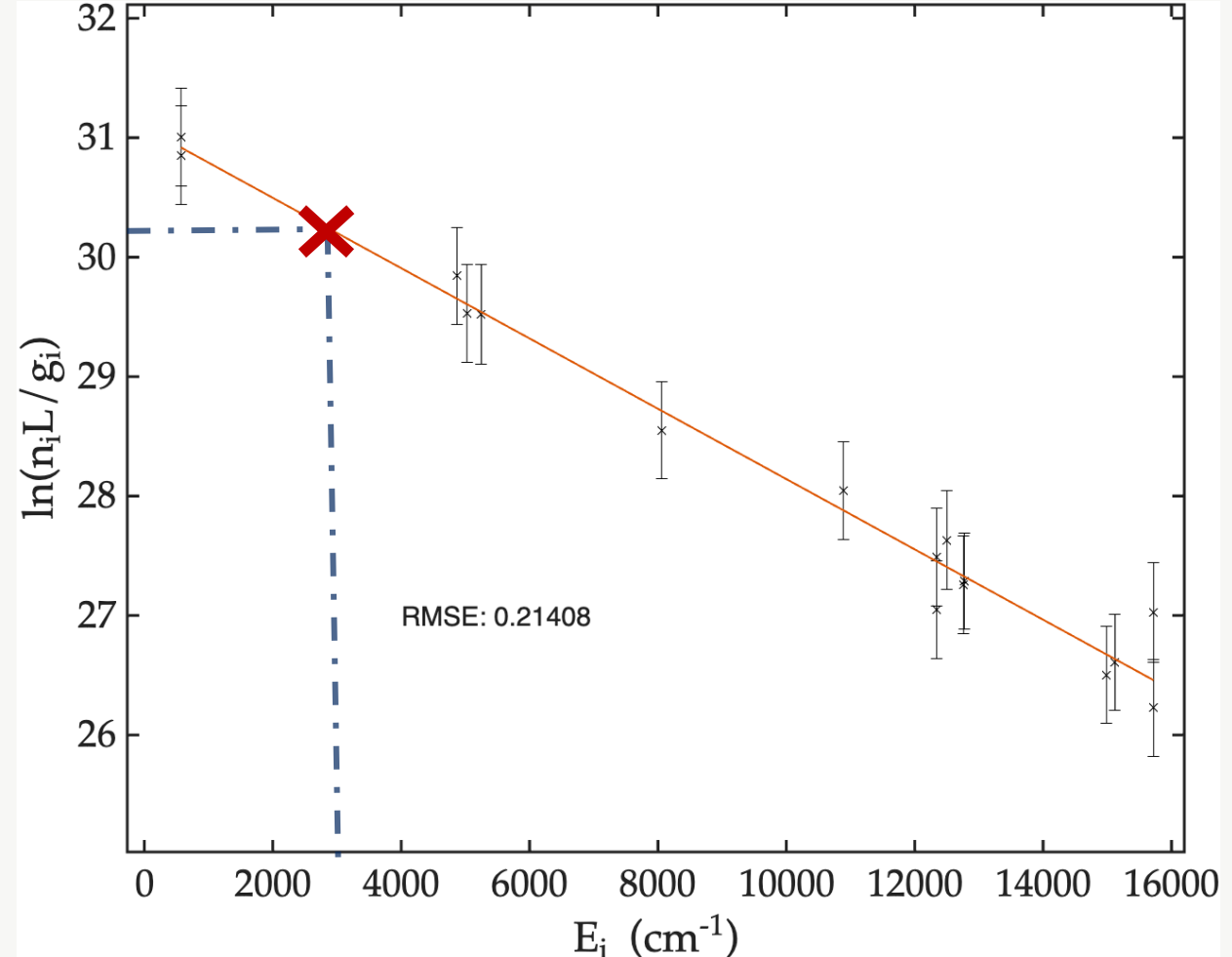
- T and $n_{tot}L$ now known
- $\rightarrow$ solve for f_{ij}

Boltzmann:

$$n_i = g_i \cdot n_{total} \cdot \frac{e^{\frac{-E_i}{k_B T}}}{Z(T)}$$

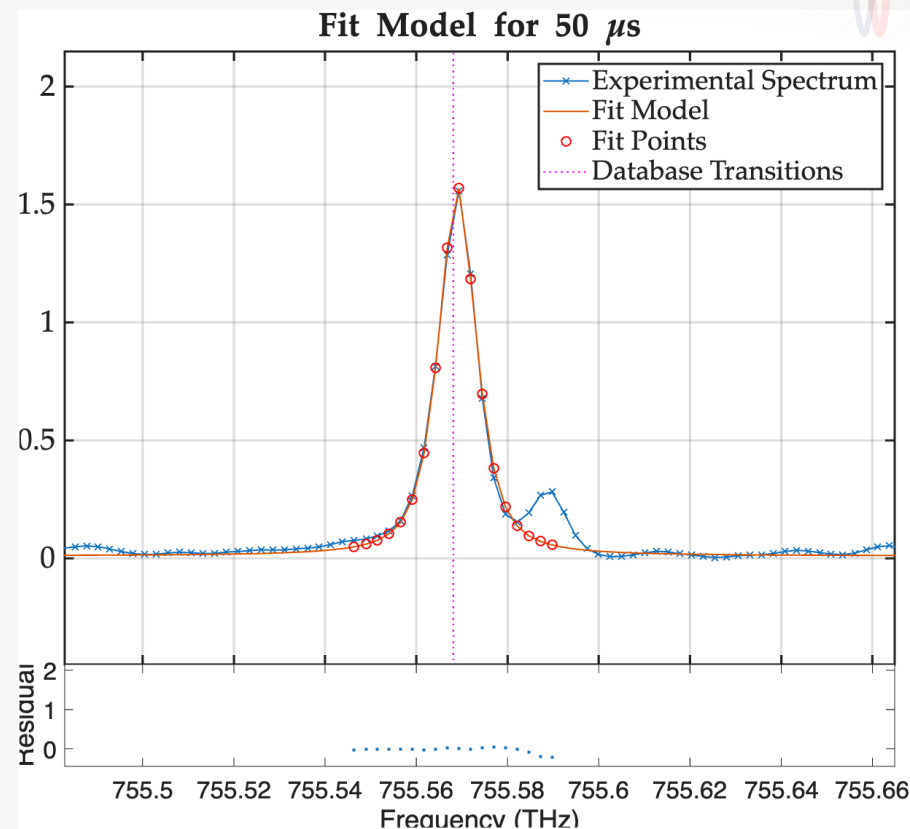
$$\ln\left(\frac{n_i L}{g_i}\right) = \ln\left(\frac{Area_{ij}}{\tilde{\sigma}_0 g_i f_{ij}}\right) = \ln\left(\frac{n_{tot} L}{Z(T)}\right) - \frac{E_i}{k_B T}$$

$$y = mx + b$$



Spectral Fitting

- Each transition fit independently (e.g., peak area, Gaussian & Lorentzian widths)
- Fit includes instrument lineshape broadening
- Uncertainties calculated from fit statistics



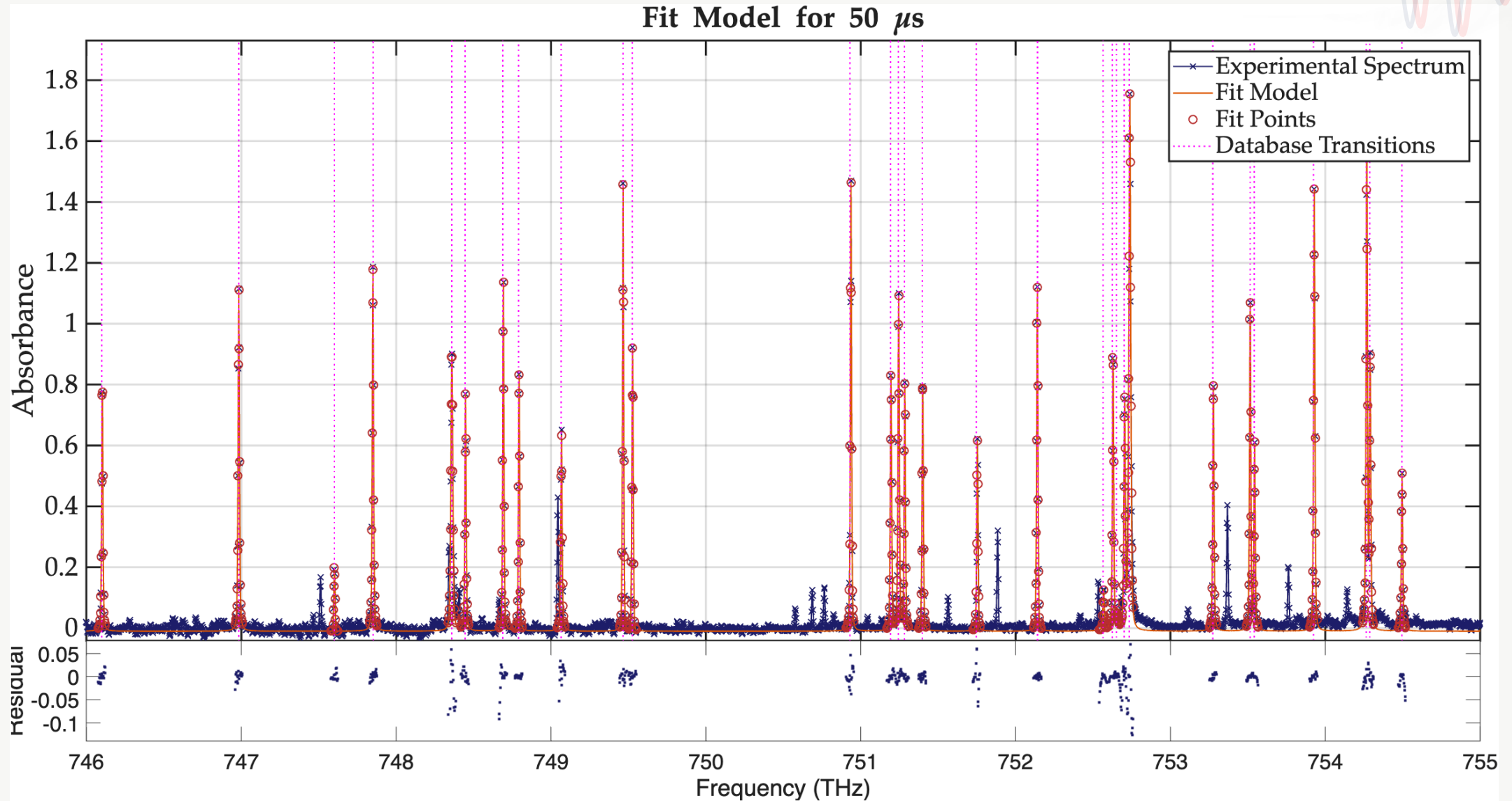
- Directly obtain **state resolved column densities** for each transition

$$A_i = \int -\ln\left(\frac{I(\nu)}{I_0(\nu)}\right) \cdot d\nu$$

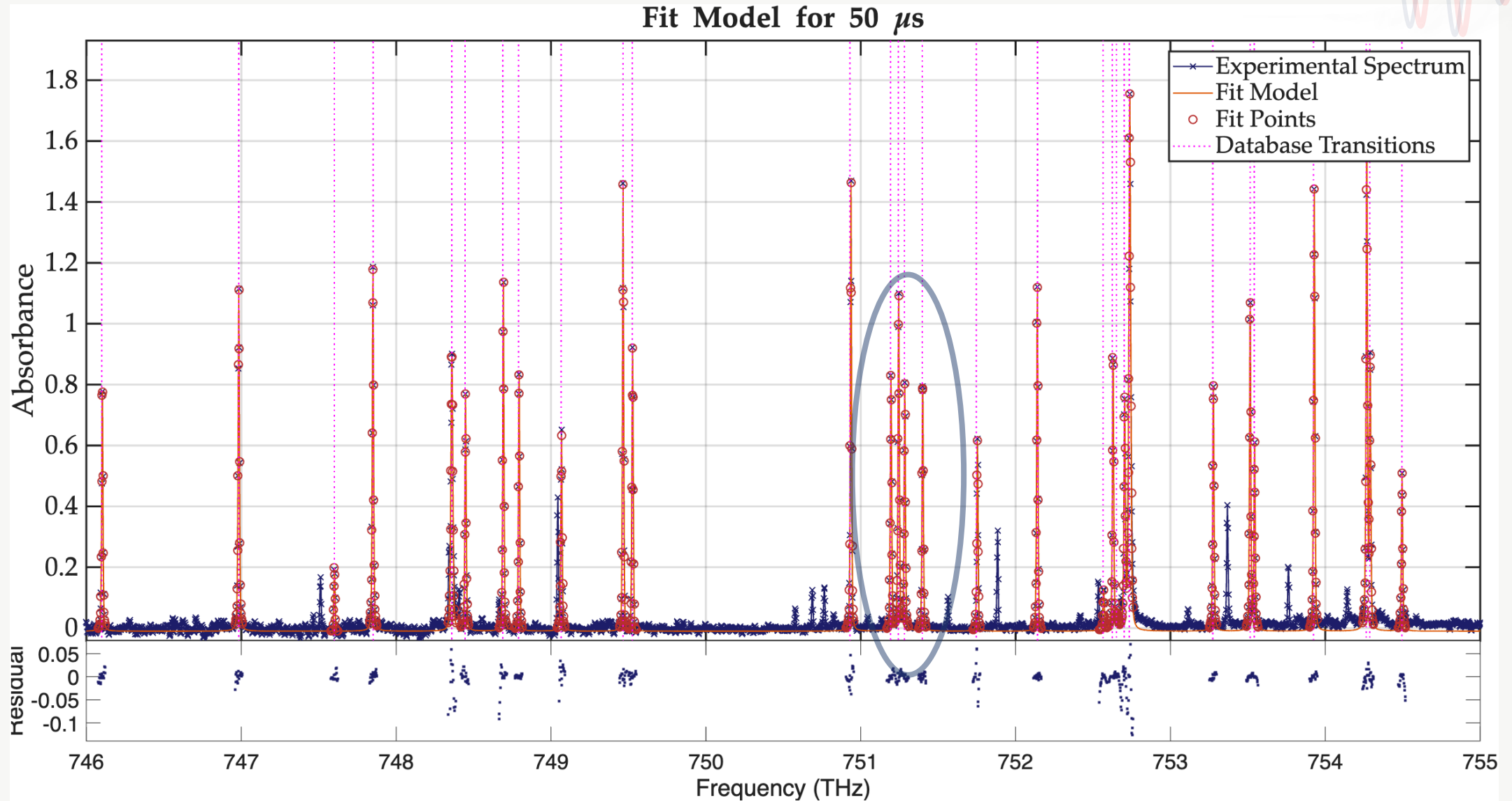
$$= \int \alpha_i(\nu) \cdot L \cdot d\nu \approx \sigma_0 \cdot g_i \cdot f_{ij} \left(\frac{n_i}{g_i} - \frac{n_j}{g_j} \right) L$$



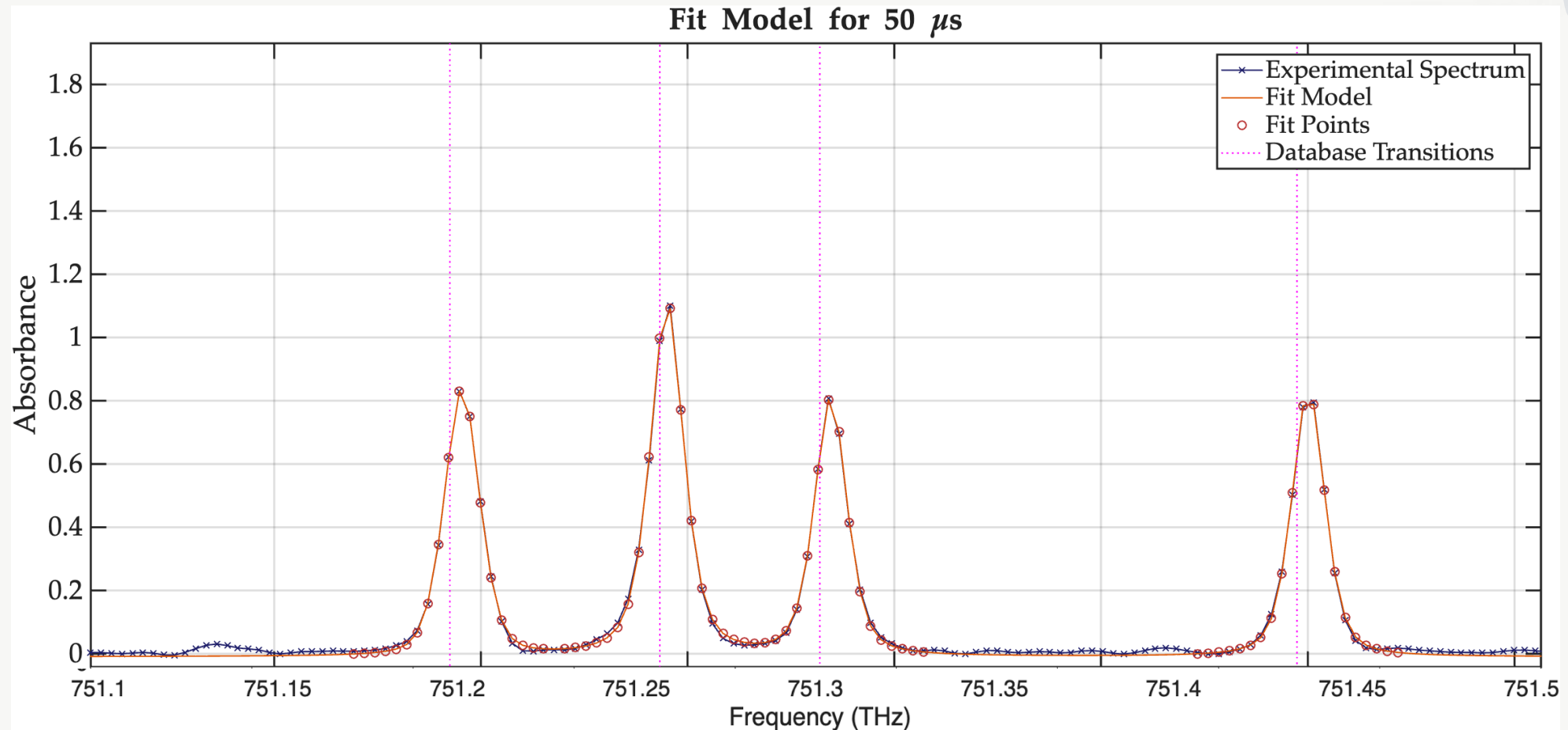
Fitting the Blue Region



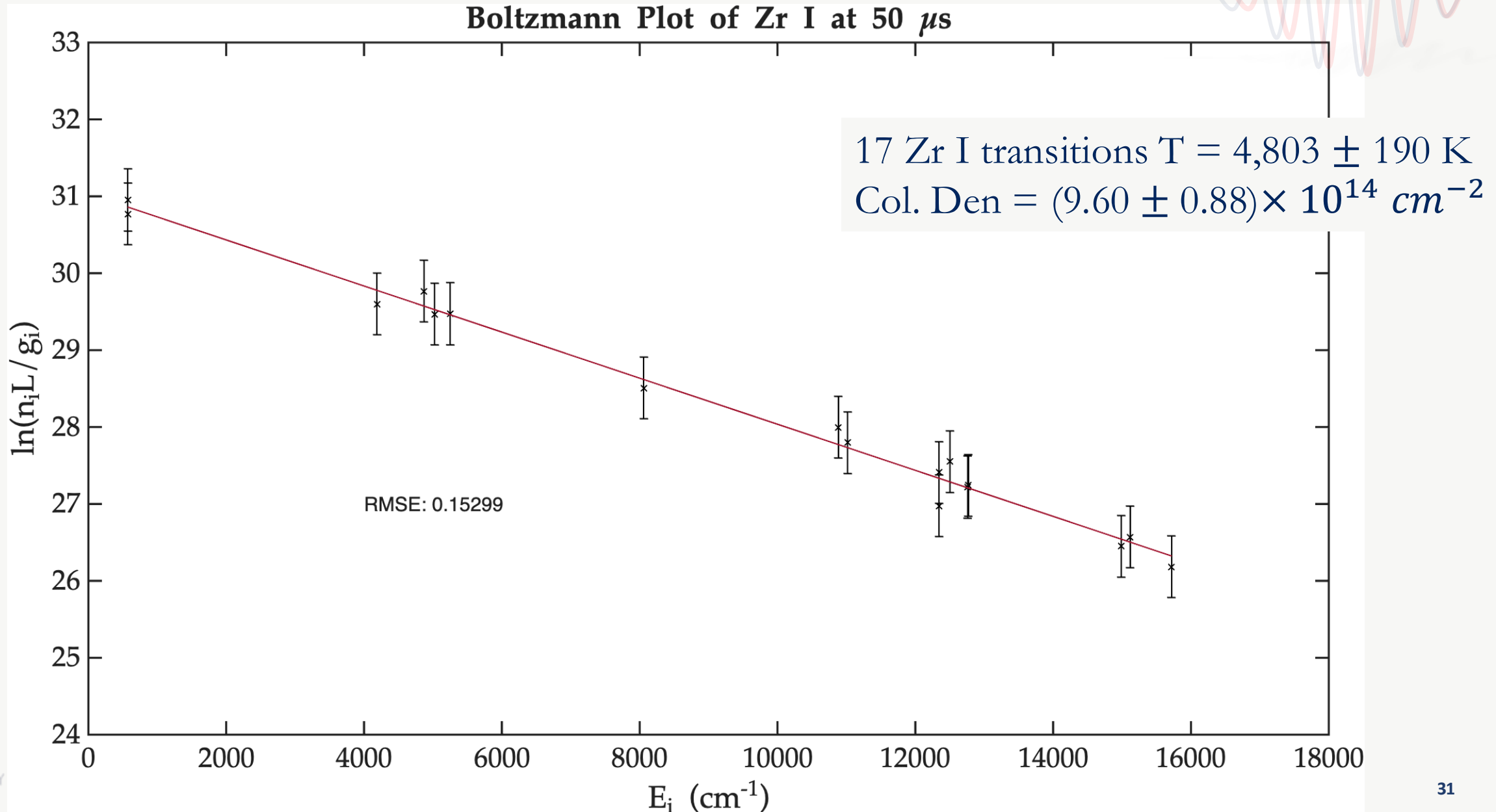
Fitting the Blue Region



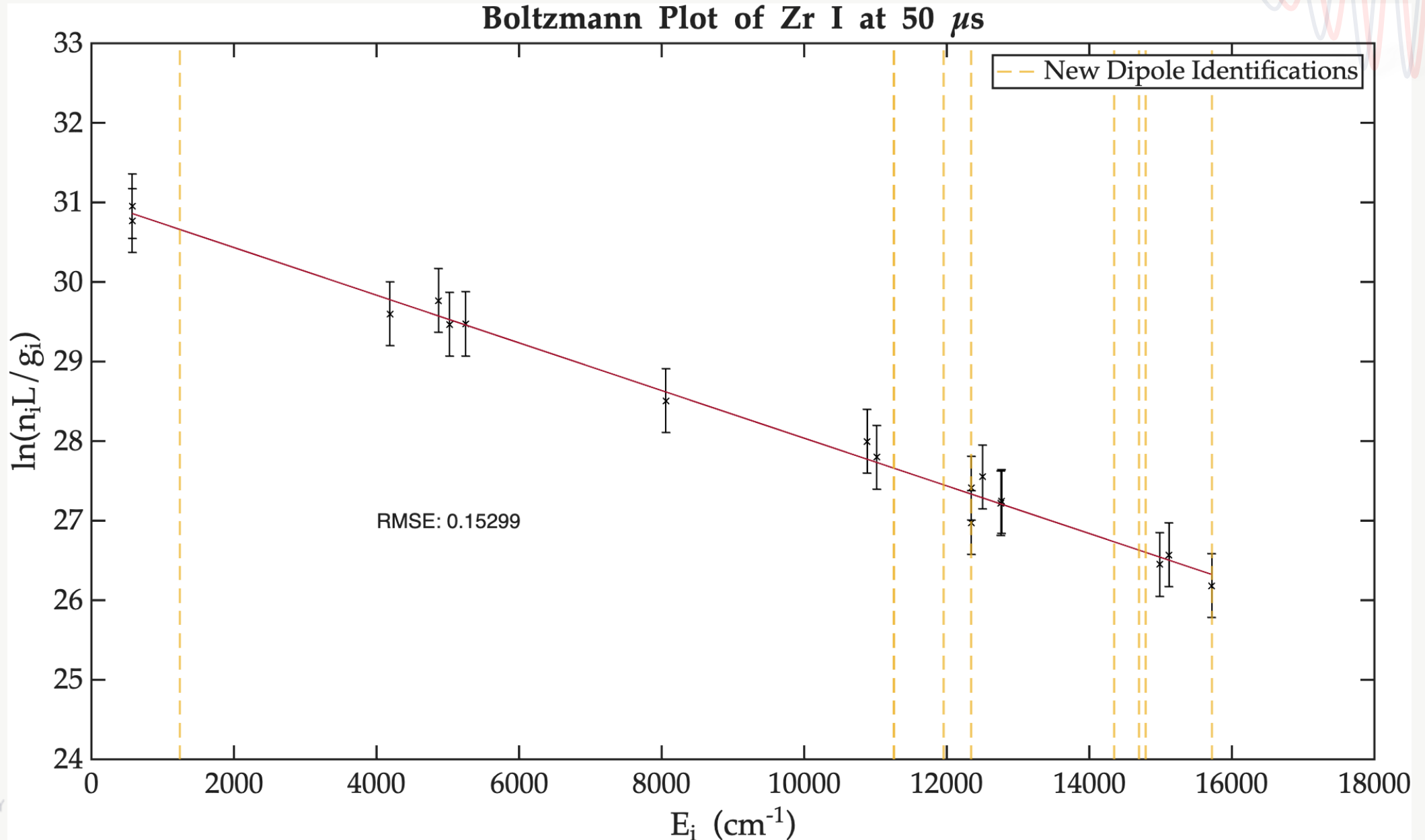
Fitting the Blue Region



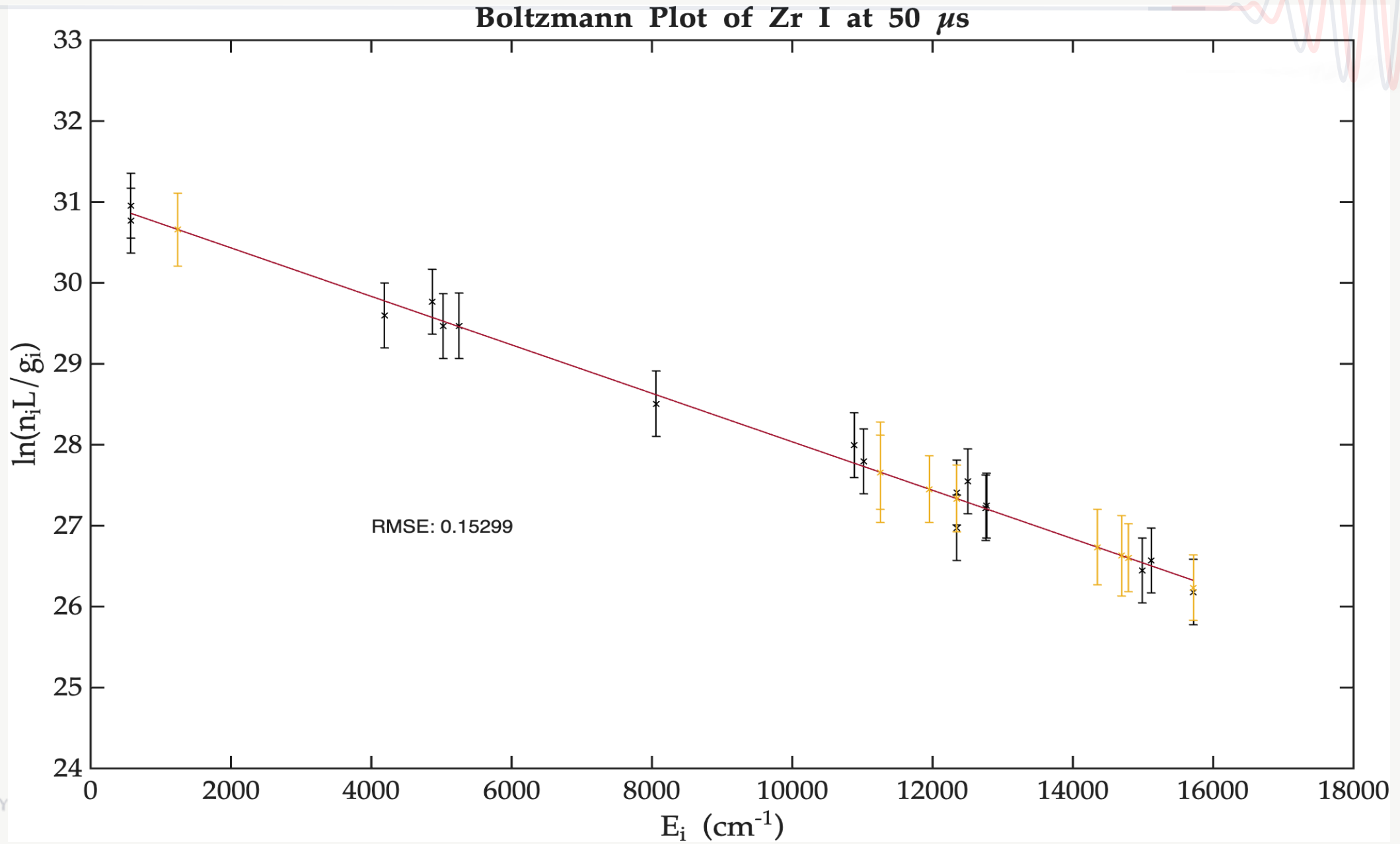
50 μ s Blue - Boltzmann Plot Analysis



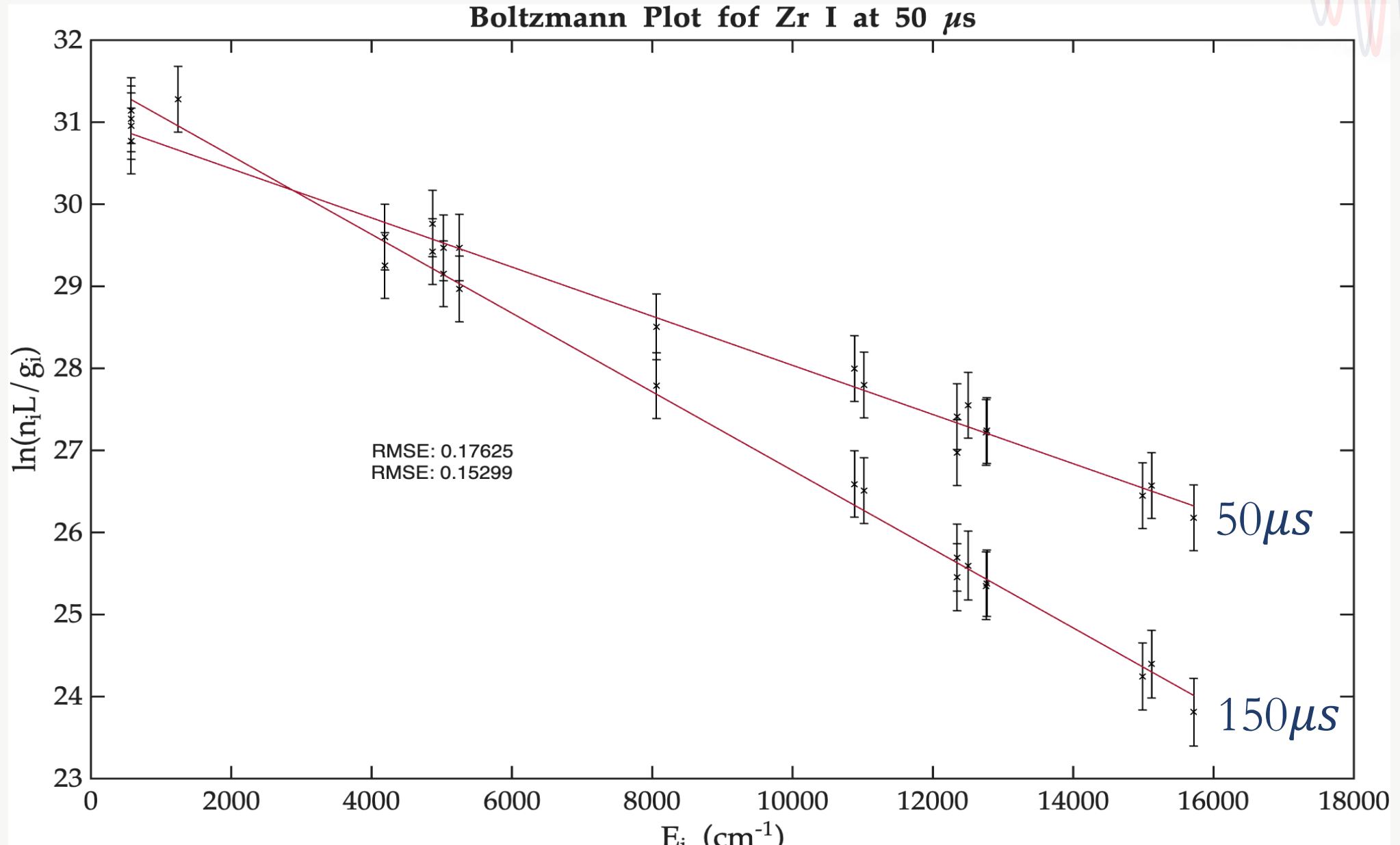
Unidentified Oscillator Strengths



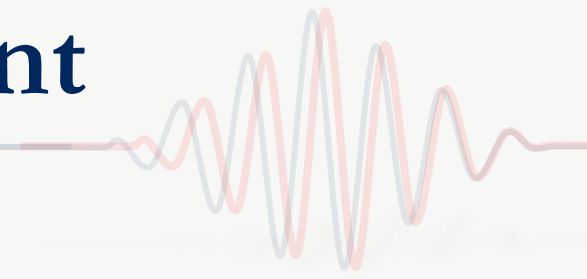
9 New Oscillator Strengths – $50\mu s$



150 and 50 μs



Oscillator Strength Agreement

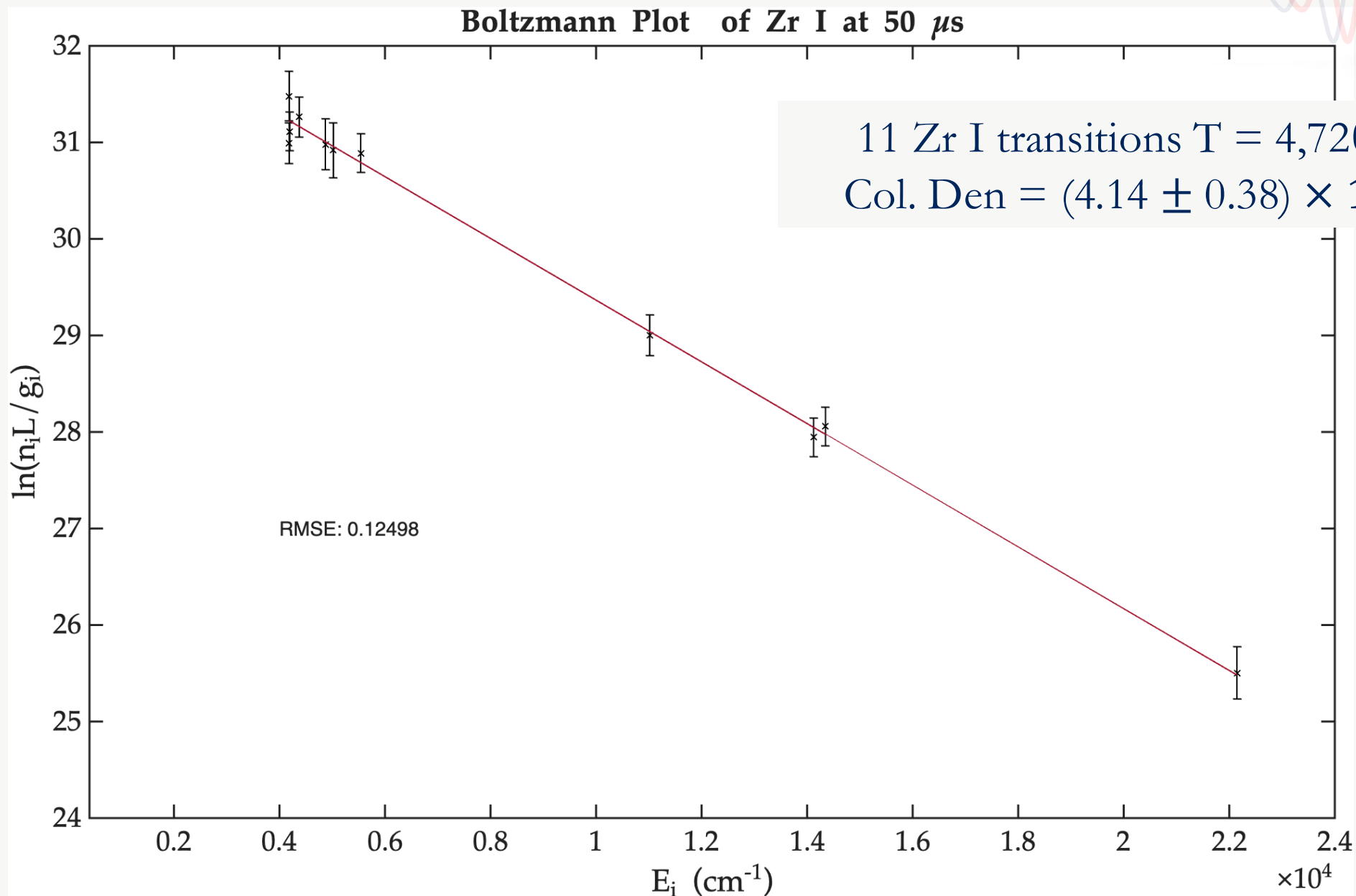


Frequency (THz)	$\log(gf)$ ($50\mu s$)	Relative Uncertainty (dex)	$\log(gf)$ ($150\mu s$)	Relative Uncertainty (dex)
751.879	-0.75	0.05	-0.68	0.11
749.041	-0.69	0.05	-0.61	0.08
752.530	-2.45	0.10	-2.55	0.05

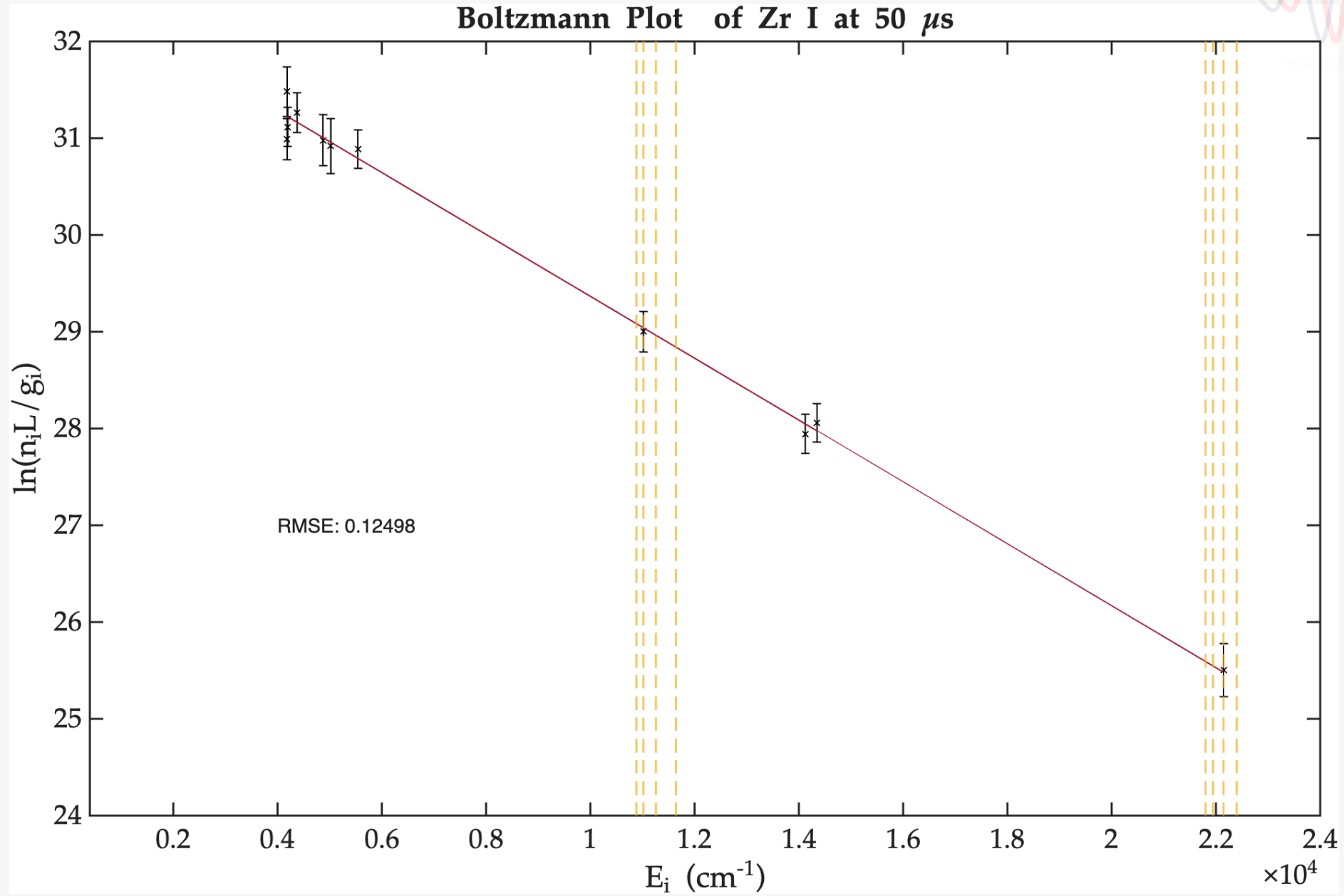




Red -Boltzmann Plot Analysis

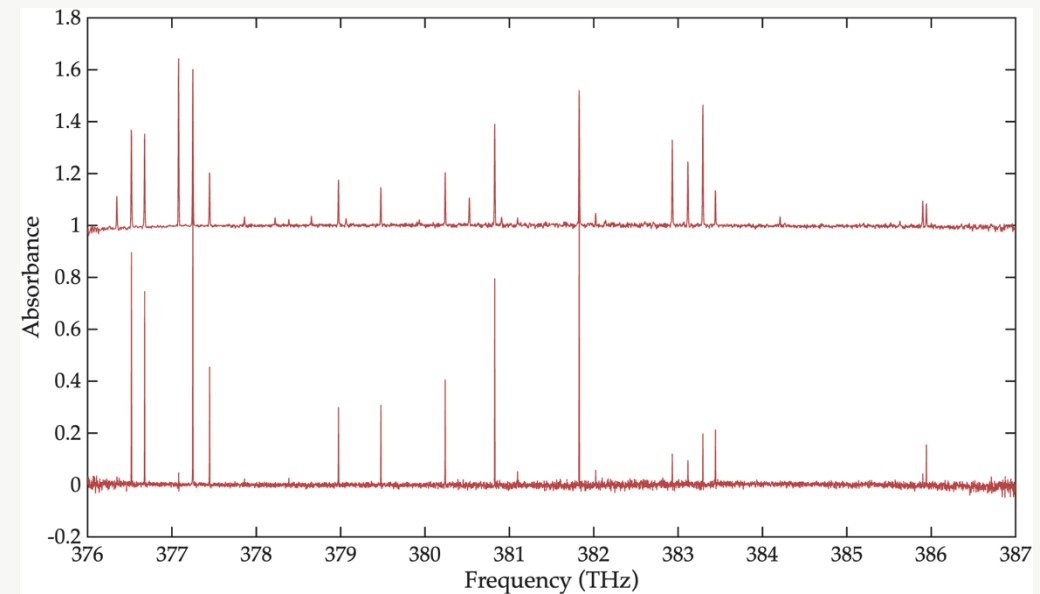
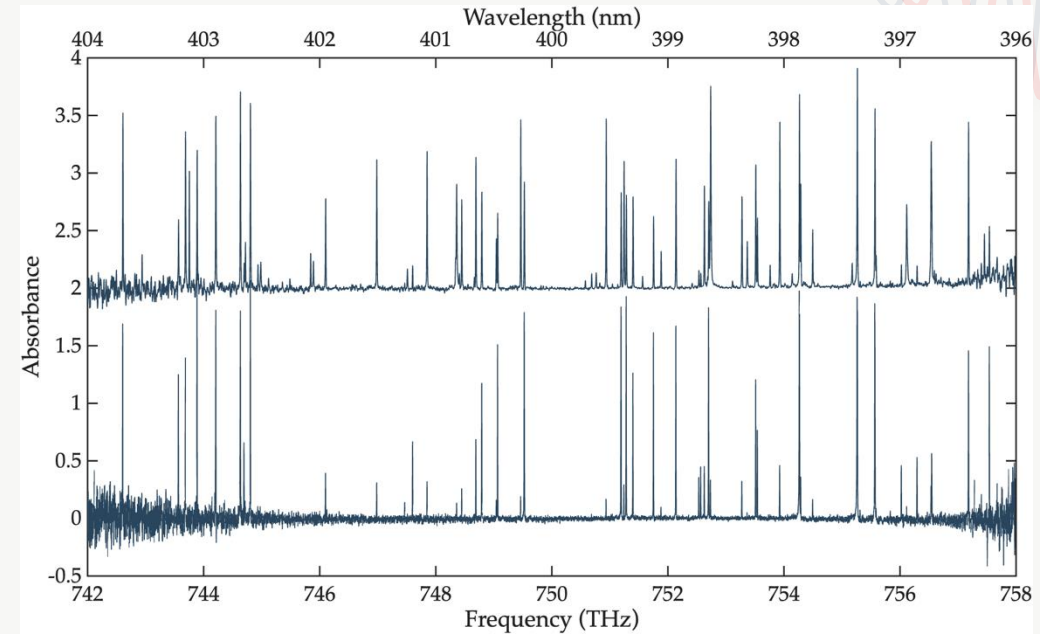


Potential Oscillator Strengths

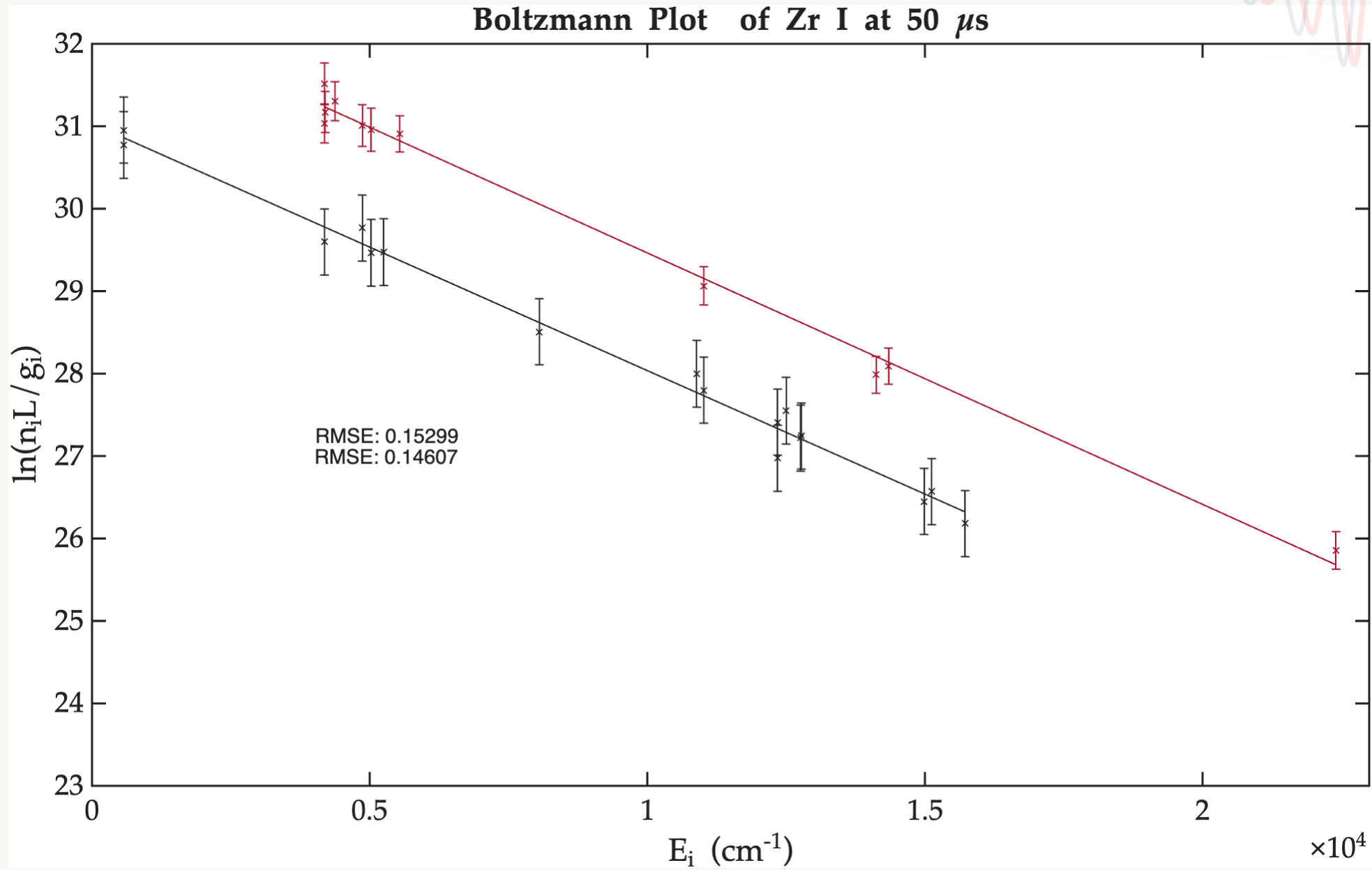


15 New Preliminary Oscillator Strengths !

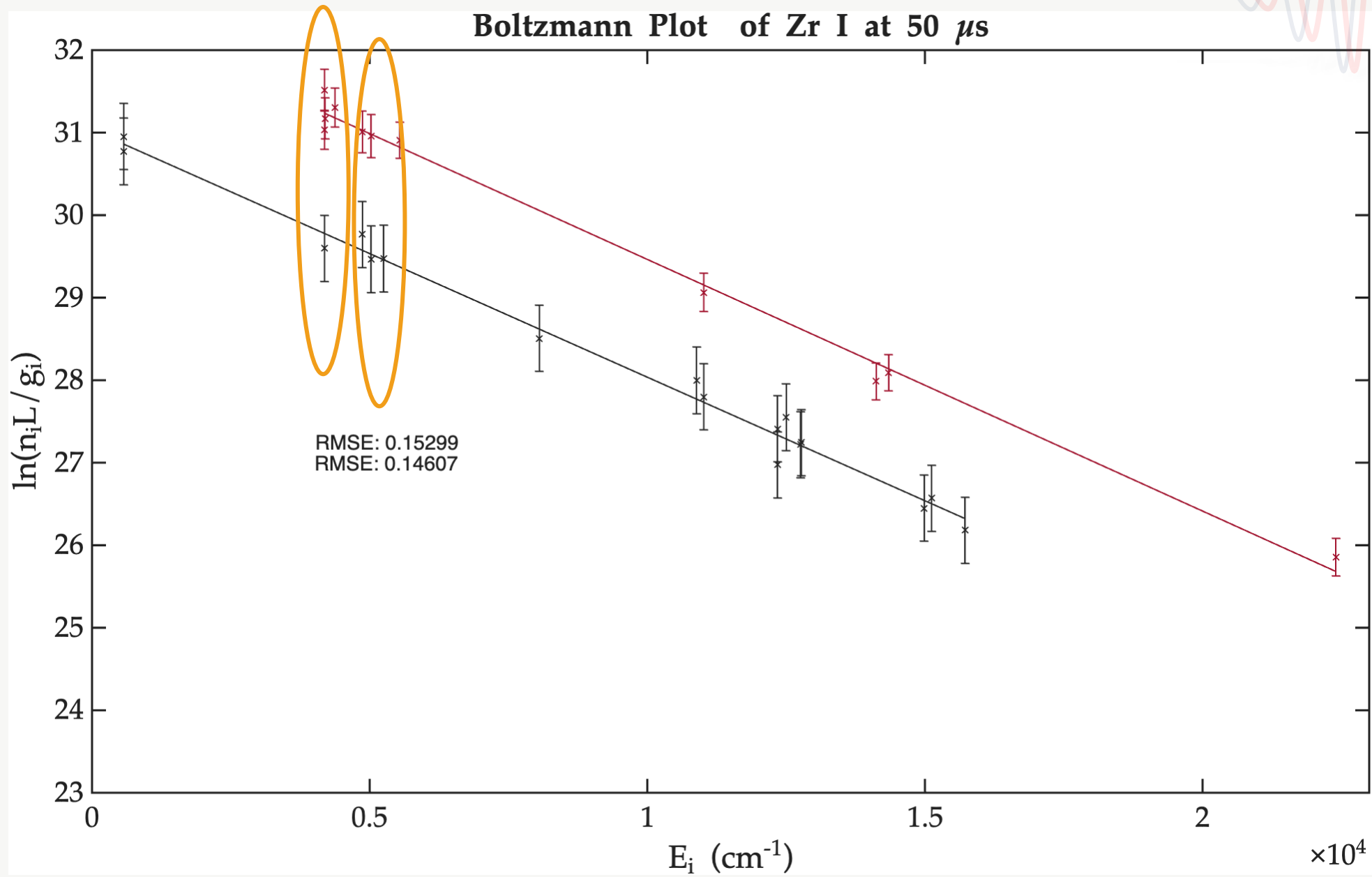
Wavelength (nm)	Frequency (THz)	E_i (cm^{-1})	Log(gf)	Relative Uncertainty (dex)
400.6078	748.3440	11258	-0.98	0.10
400.2350	749.0410	11956	-0.69	0.05
399.4170	750.5751	11258	-1.63	0.21
399.3529	750.6956	14349	-0.93	0.11
398.8920	751.5629	14697	-0.96	0.13
398.7242	751.8792	12342	-0.75	0.05
398.3796	752.5297	1241	-2.45	0.10
397.9412	753.3587	15720	-0.06	0.04
397.7272	753.7640	14791	-0.56	0.07
787.0514	380.9058	22145	-0.89	0.08
789.0629	379.9348	21944	-1.00	0.10
790.8689	379.0672	11017	-2.50	0.09
791.7248	378.6574	11258	-2.43	0.08
793.3945	377.8605	10885	-2.52	0.08
784.5105	382.1395	11641	-2.42	0.11



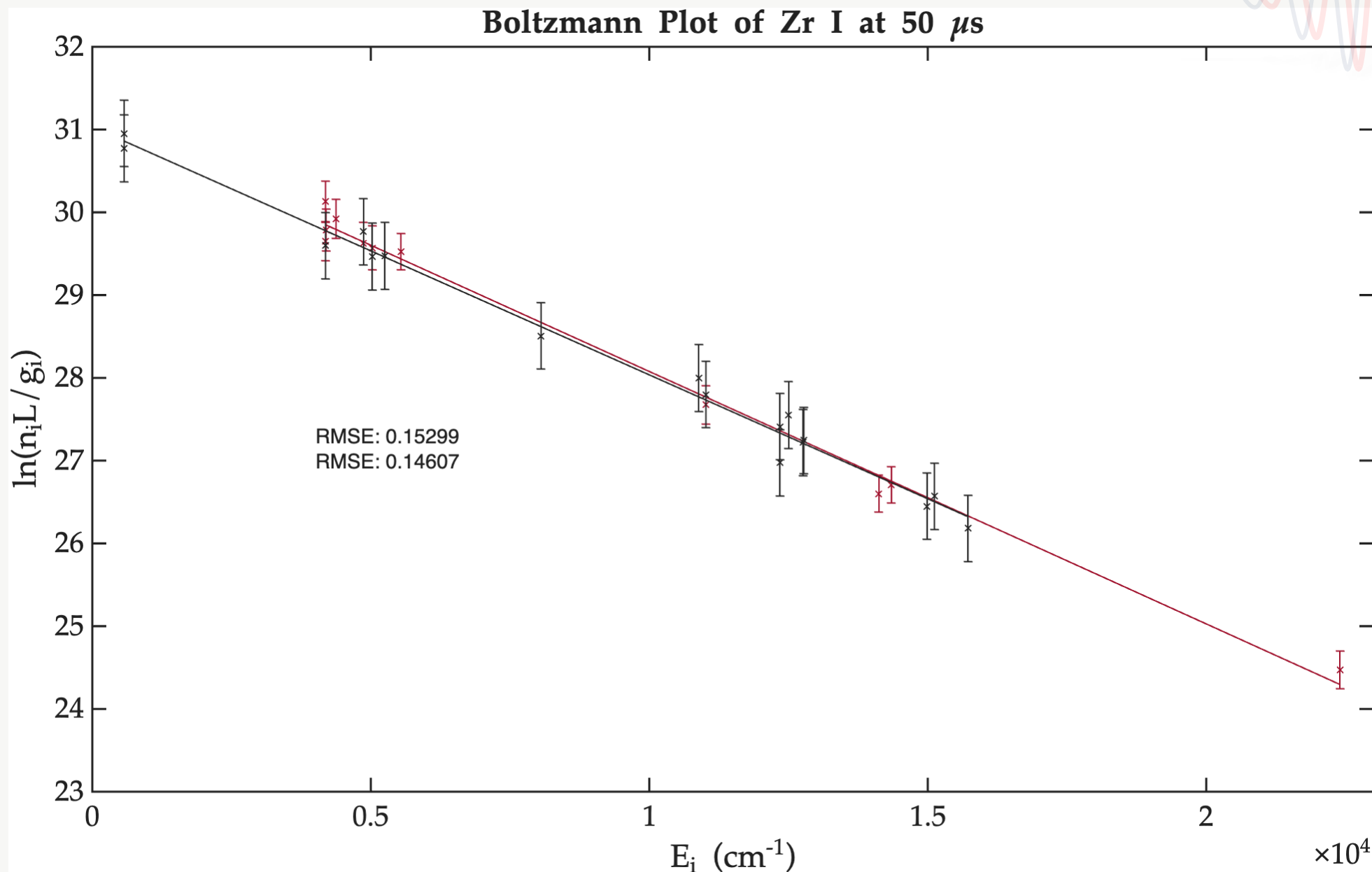
? The mystery ?



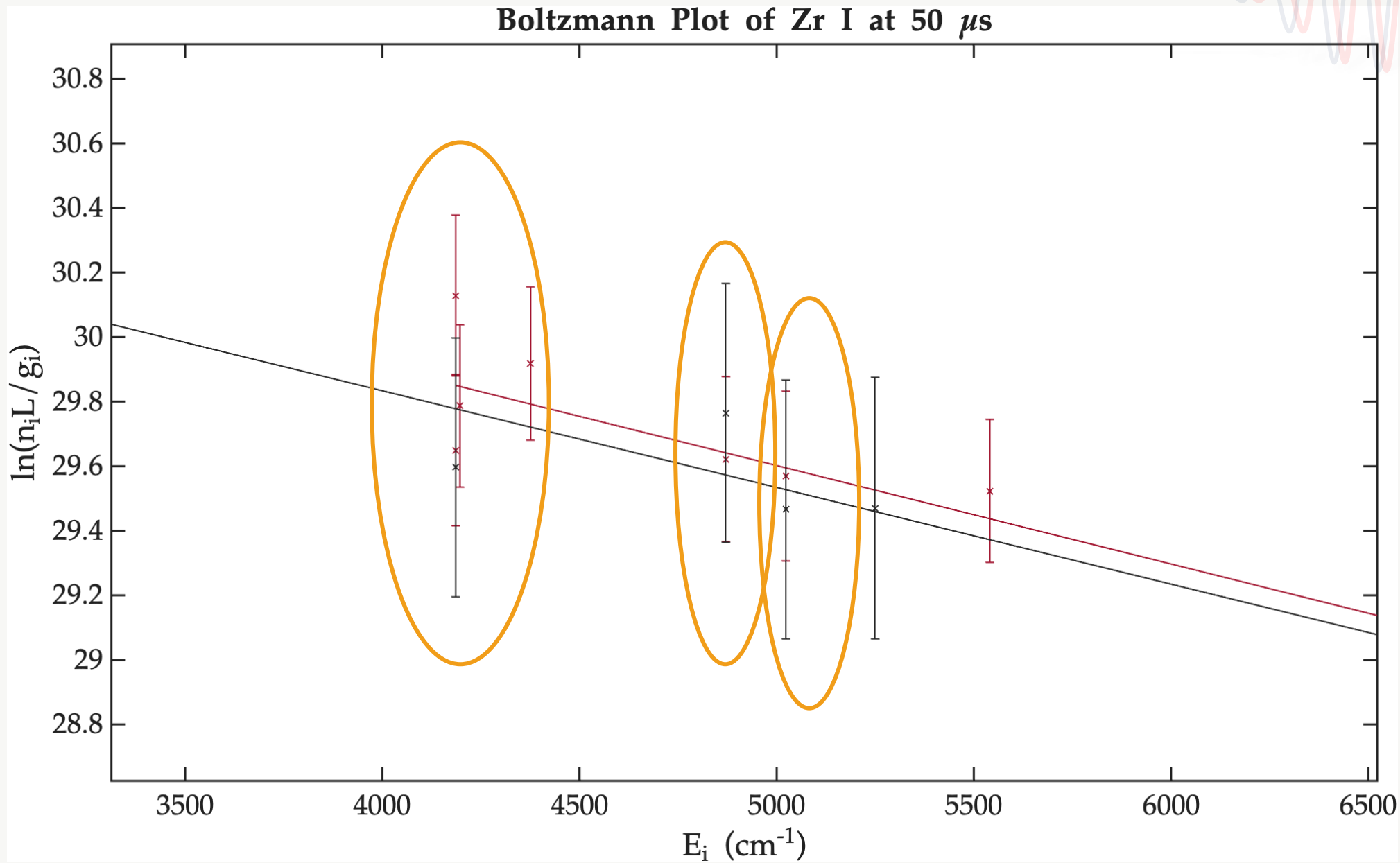
? The mystery ?



Accounting for the Factor of 4



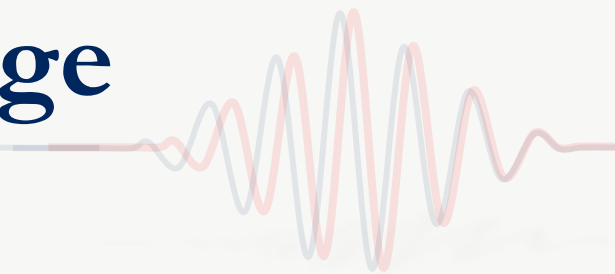
Originating from the Same Energy Level



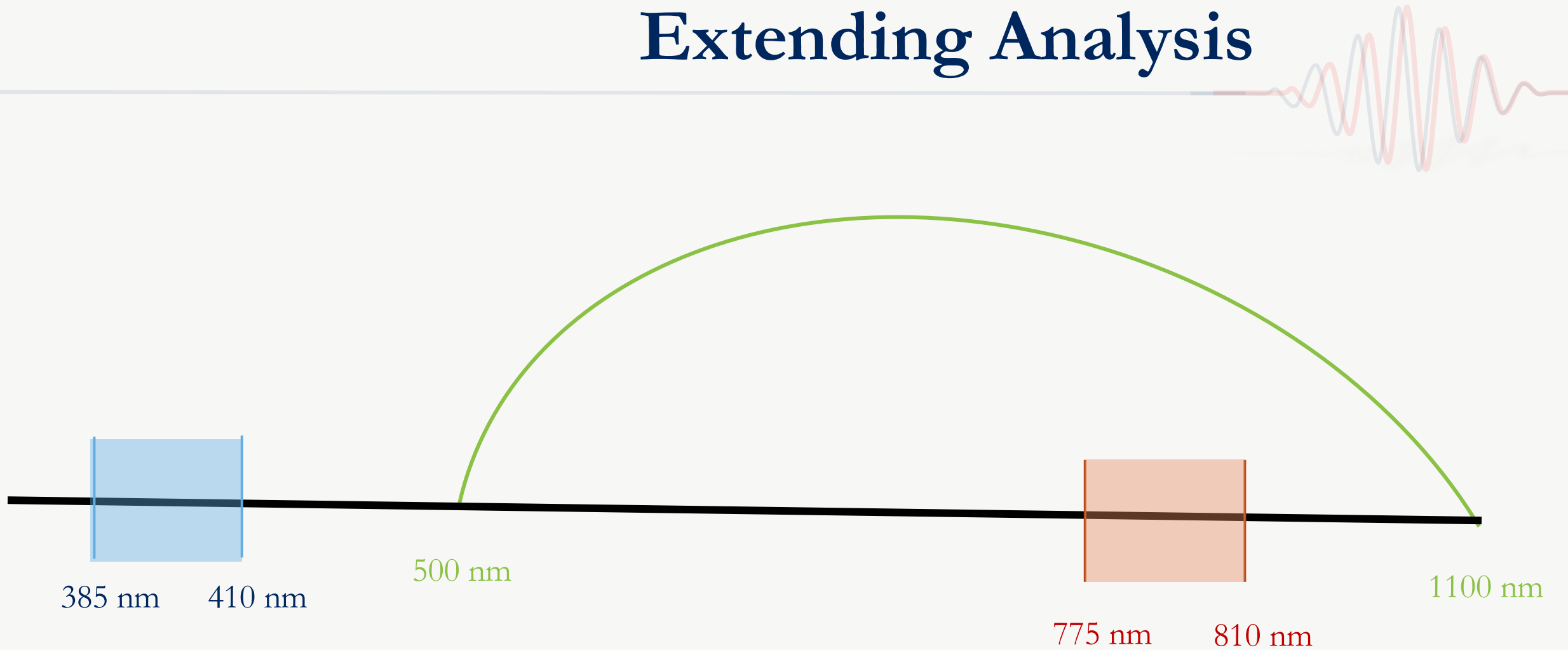
Current Wavelength Range



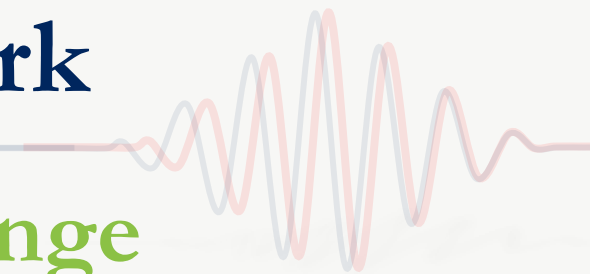
Extended Wavelength Range



Extending Analysis



Ongoing Work



Fundamental (780-810nm)

- 27 Zr transitions
- 3 new Zr I transitions

2nd harmonic (392-403nm)

- 70 Zr transitions
- 26 new Zr I transitions
- 4 new Zr II transitions

Total Range (500-1100nm)

- 217 New Transitions
- 193 Zr I
- 24 Zr II

Paper In Review:

Valence Photoabsorption Spectroscopy of Zr and Zr⁺ for Astrophysical Applications

Kirsten Dowd,^{1★} Hope Dannar,² Michael McCann,³ R. Jason Jones,² Catherine A. Ramsbottom,³
Andrew Jarymowycz,² Connor P. Ballance,³ Padraig Dunne¹

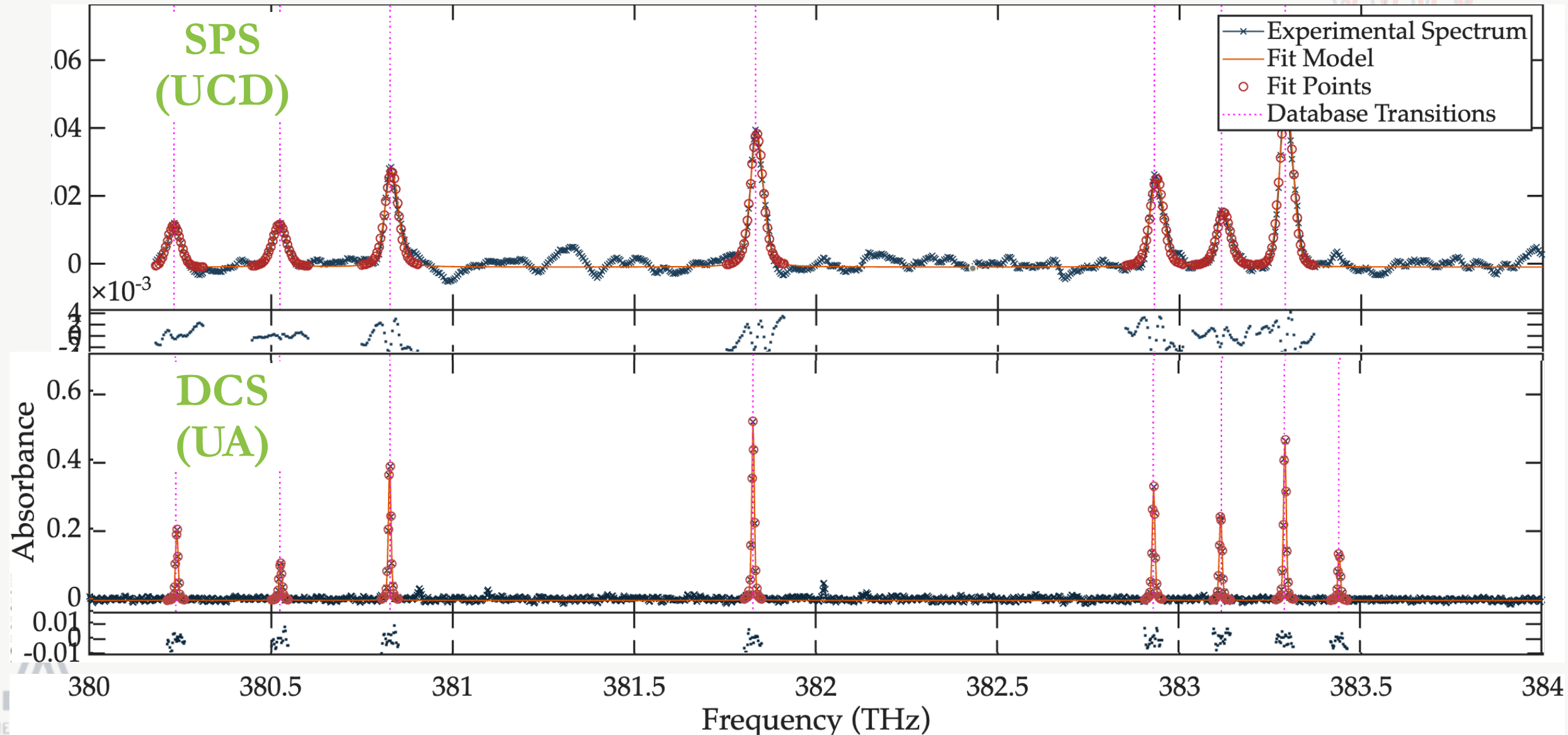
¹*School of Physics, University College Dublin, Belfield, Dublin 4, Ireland*

²*Wyant College of Optical Sciences, University of Arizona, 1630 E. University Blvd., Tucson, AZ 85721, USA*

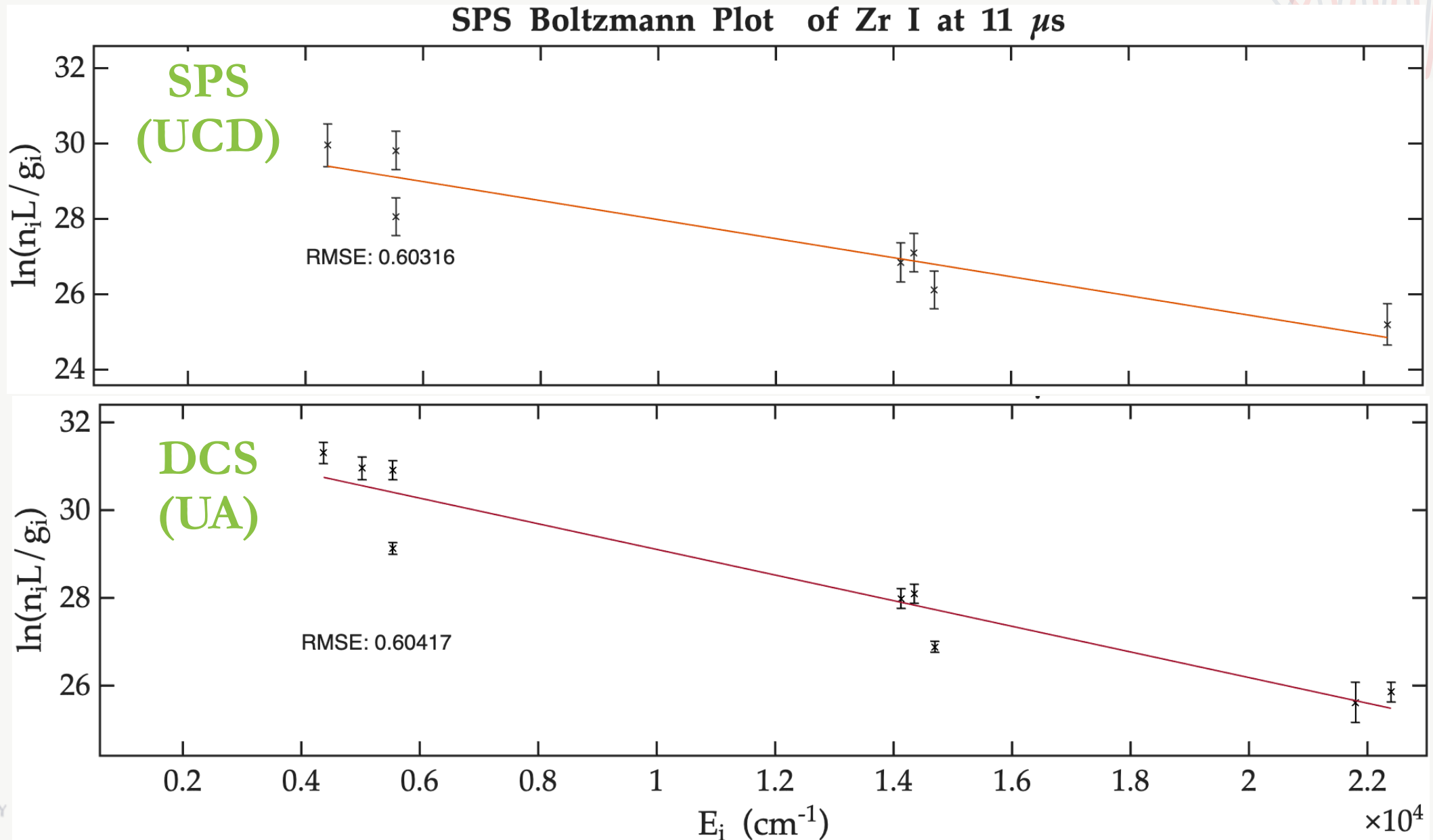
³*Astrophysics Research Centre, School of Mathematics & Physics, Queens University Belfast, BT7 1NN, Northern Ireland.*



Area Fit Comparison

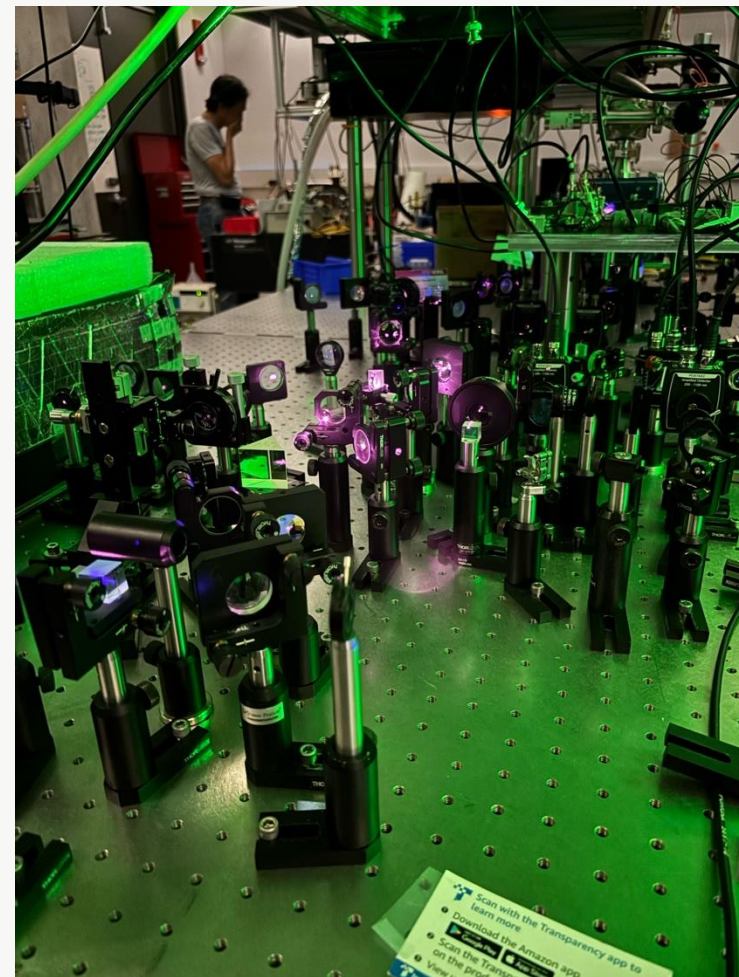


Boltzmann Fit Comparison



Frequency Combs for Astrophysics

- Approach-identifies new transitions, energy levels, and oscillator strengths, and can help reduce uncertainties in current data base.
- Using two spectral regions simultaneously provides greater coverage of neutral and ionic transitions, additional fitting constraints, improving the reliability of the resulting oscillator strengths.
- Going Forward: Alloys, broader bandwidth, and more species.



Thank you

PI: Dr. R. Jason Jones

Dr. Mark C. Phillips

PhD Students:

- Ryland G. Wala
- John J. McCauley
- Dylan Tooley
- **Andrew Jarymowycz**
- Charlie Chisholm
- Liam Frank
- Kelby Todd



THE UNIVERSITY OF ARIZONA
Wyant College
of Optical Sciences

UA Jones Group
Hope Dannar

hdannar@arizona.edu