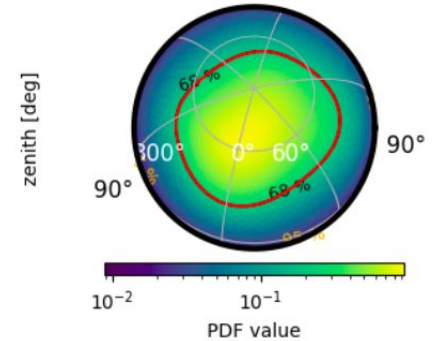
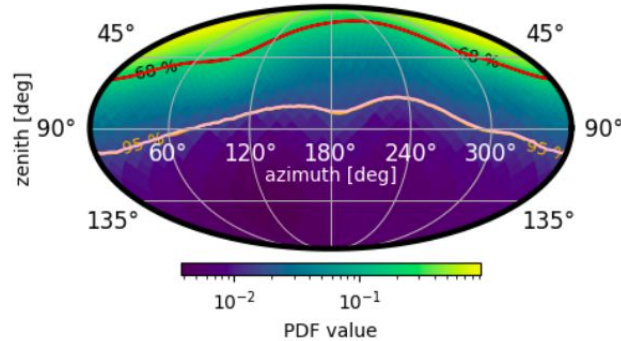


Transformer-encoded normalizing flows for IceCube reconstruction

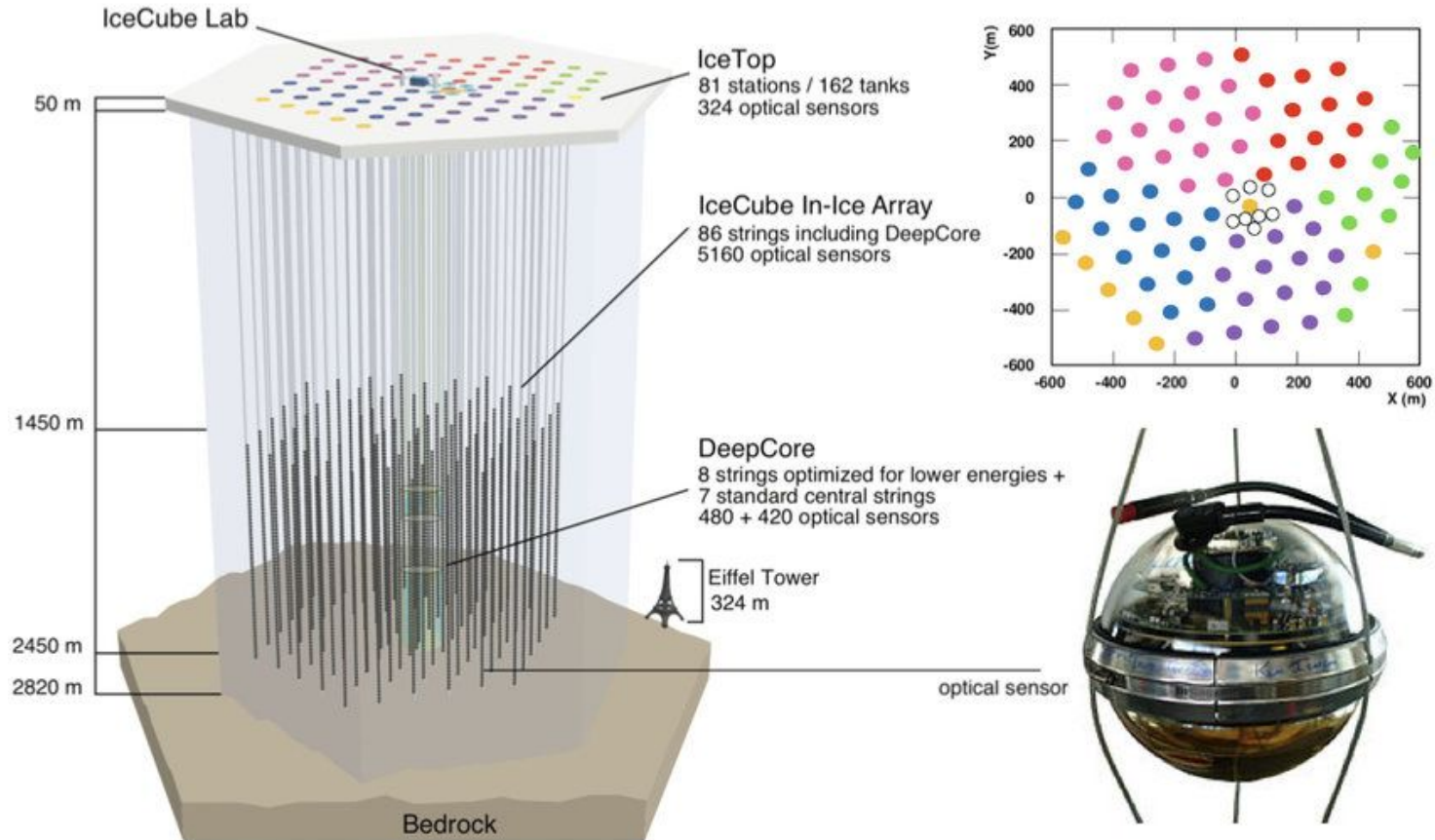


Stockholm
University

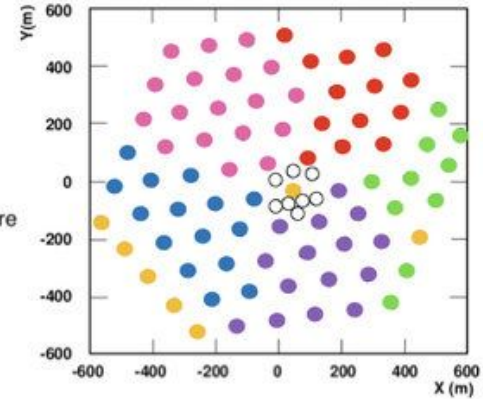
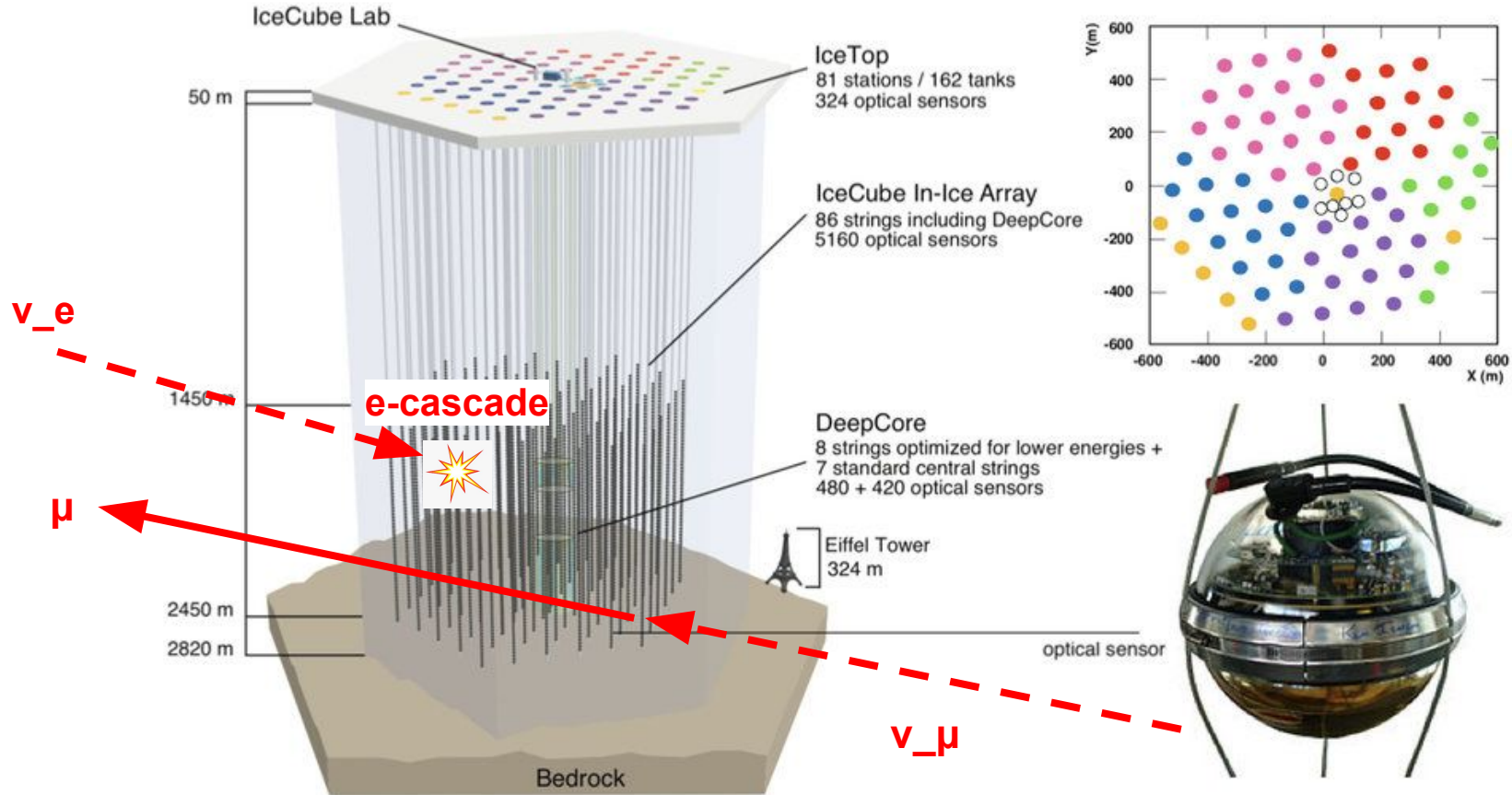


HAMLET-physics meeting, Aug 20th, 2026,
Thorsten Glüsenskamp

IceCube Detector



IceCube Detector



Digital
Optical
Module

IceCube Detector

Cherenkov light cascade



ν_e

e-cascade



μ

1450 m

2450 m

2820 m

Bedrock

IceTop

81 stations / 162 tanks
324 optical sensors

IceCube In-Ice Array

86 strings including DeepCore
5160 optical sensors

DeepCore

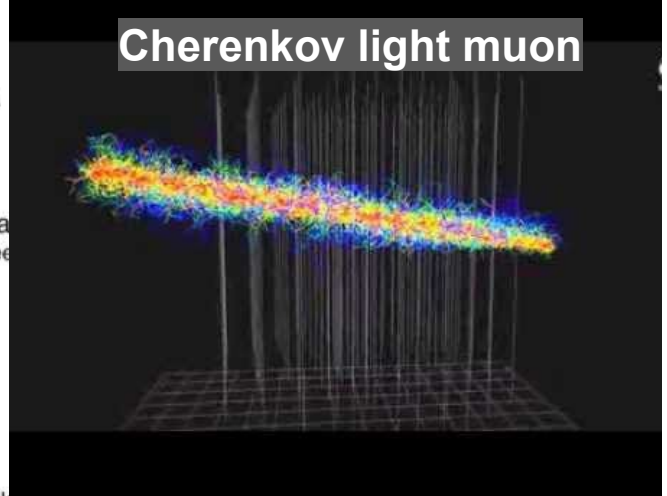
8 strings optimized for lower energies +
7 standard central strings
480 + 420 optical sensors

Eiffel Tower
324 m

optical sensor

ν_μ

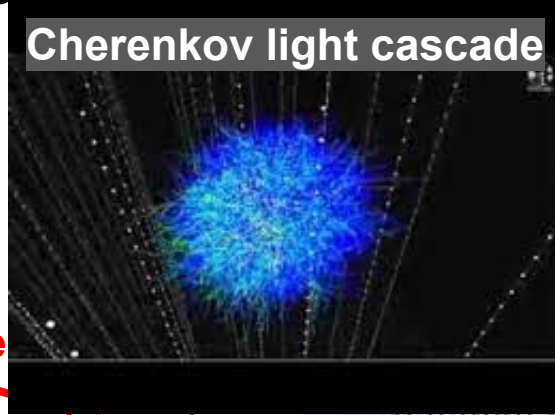
Cherenkov light muon



Digital
Optical
Module

IceCube Detector

Cherenkov light cascade



ν_e

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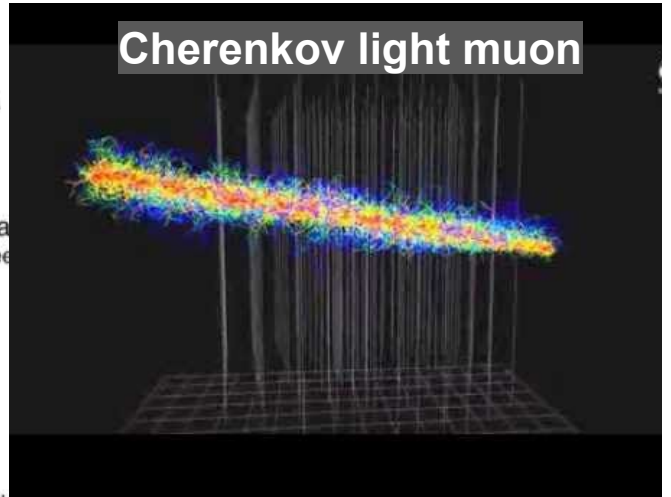
8 strings optimized for lower energies +
7 standard central strings
480 + 420 optical sensors

Eiffel Tower

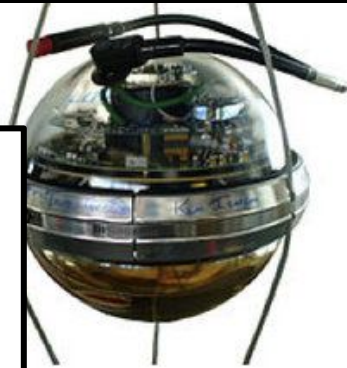
Interested in few degrees of freedom of these events:

- Energy
- Direction
- Position
- ...

Cherenkov light muon



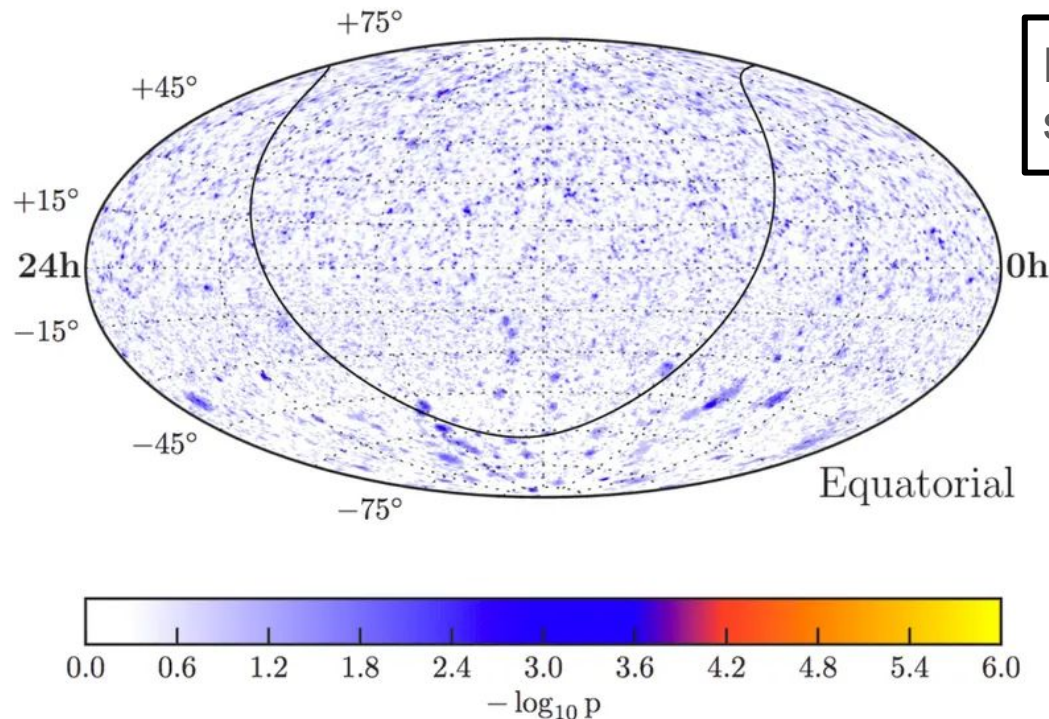
Digital
Optical
Module



h

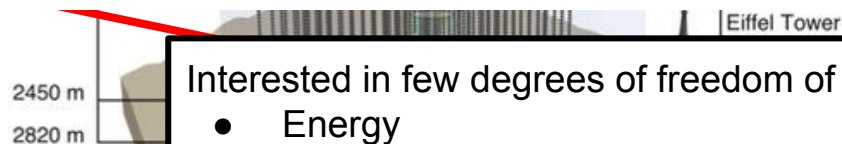
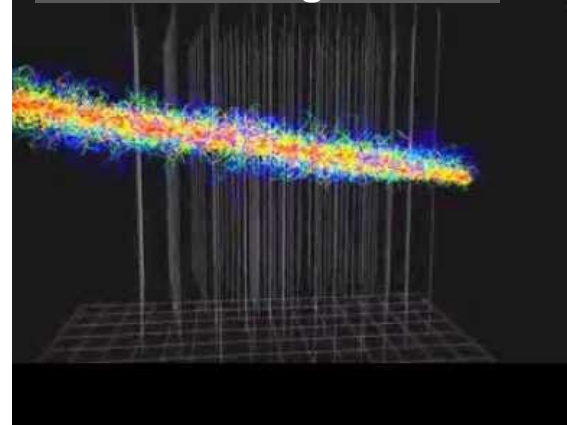
v

l



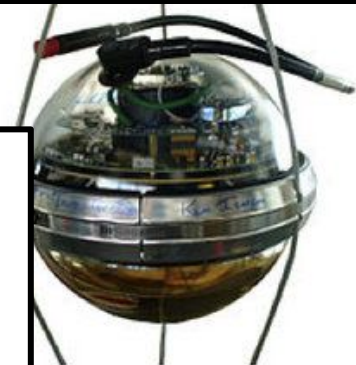
E.g. direction for point source searches

Cherenkov light muon



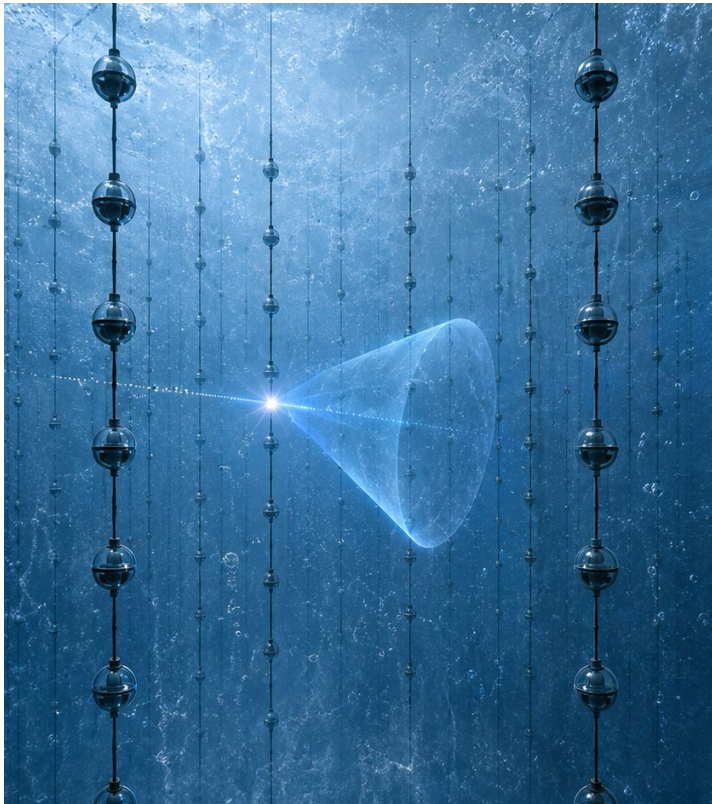
Interested in few degrees of freedom of these events:

- Energy
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- ...

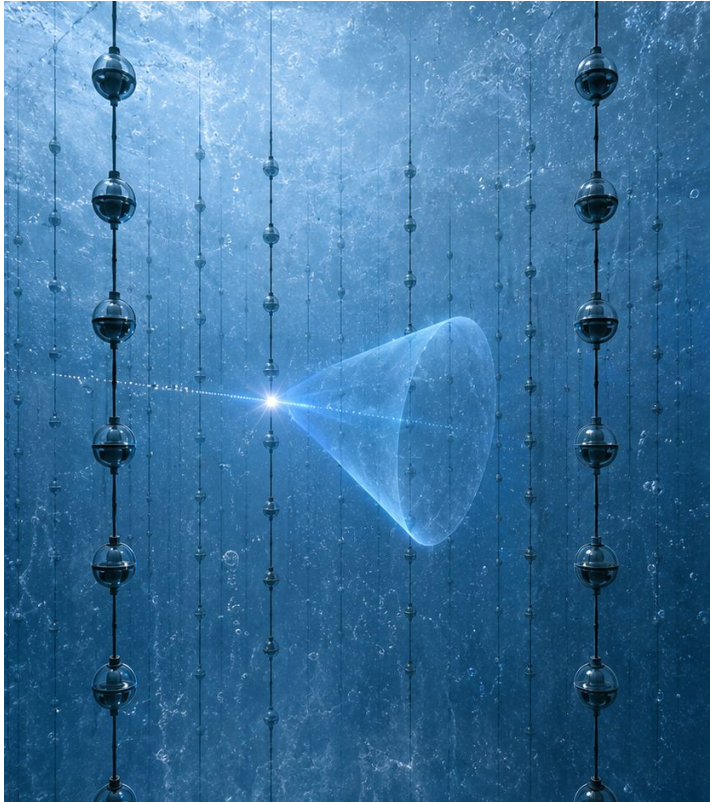


Digital
Optical
Module

Interaction happens...



Interaction happens...



A few nanoseconds later..



Interaction happens...

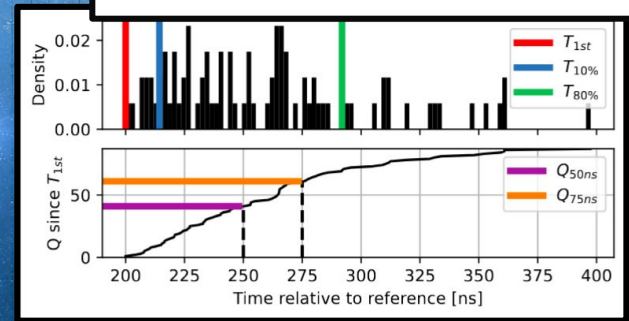
Forward (likelihood) models

$\theta = \{\text{Energy, Direction, ...}\}$

A few nanoseconds later..

Data "x" (hits, bin counts)

Data distribution $p(x; \theta)$

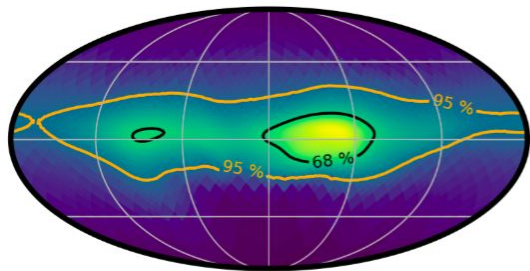


Interaction happens...

Forward (likelihood) models

$\theta = \{\text{Energy, Direction, ...}\}$

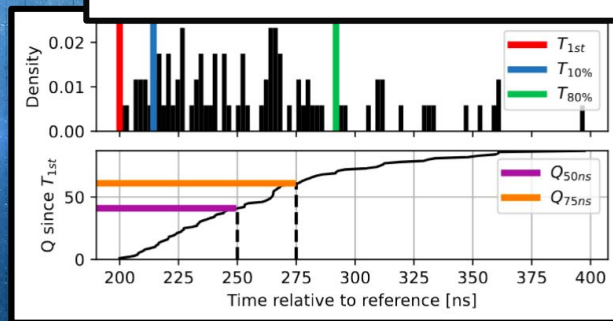
Posterior $p(\theta; x)$



A few nanoseconds later..

Data "x" (hits, bin counts)

Data distribution $p(x; \theta)$



Inverse (likelihood-free) modelling

(amortized neural posterior estimation, NN-Regression / Classification, ...)

Interaction happens...

A few nanoseconds later..

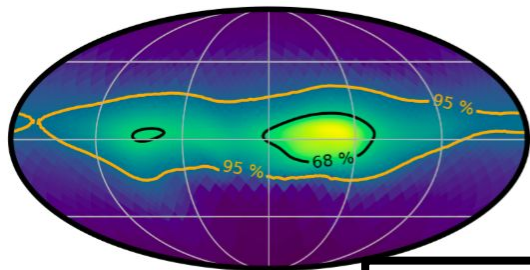
Forward (likelihood) models

“Fit all data generation distributions at once” (obtain the likelihood)

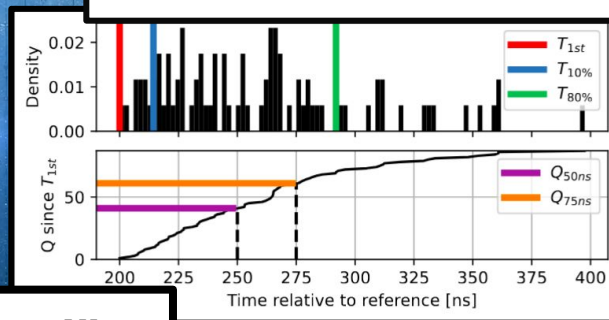
$\theta = \{\text{Energy, Direction, ...}\}$

Data “x” (hits, bin counts)

Posterior $p(\theta; x)$



Data distribution $p(x; \theta)$



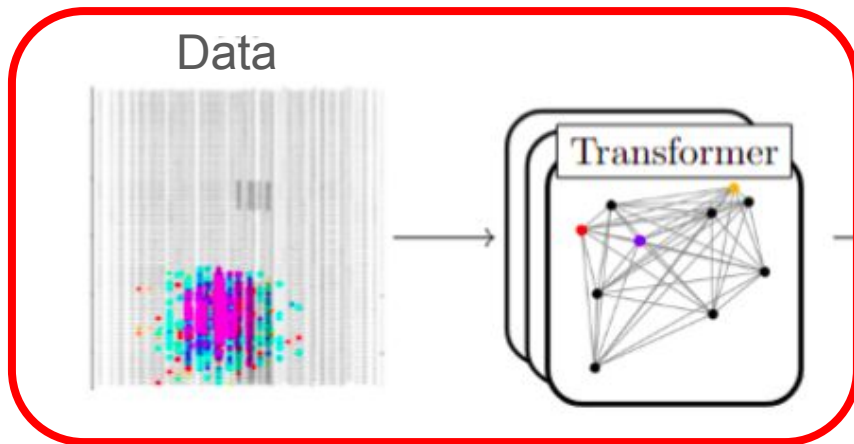
“Fit all posterior distributions at once!”

Inverse (likelihood-free) modelling

(amortized neural posterior estimation, NN-Regression / Classification, ...)

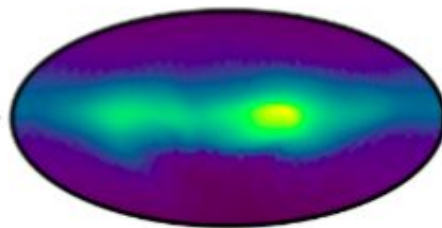
The model: a posterior PDF prediction machine

STEP 1



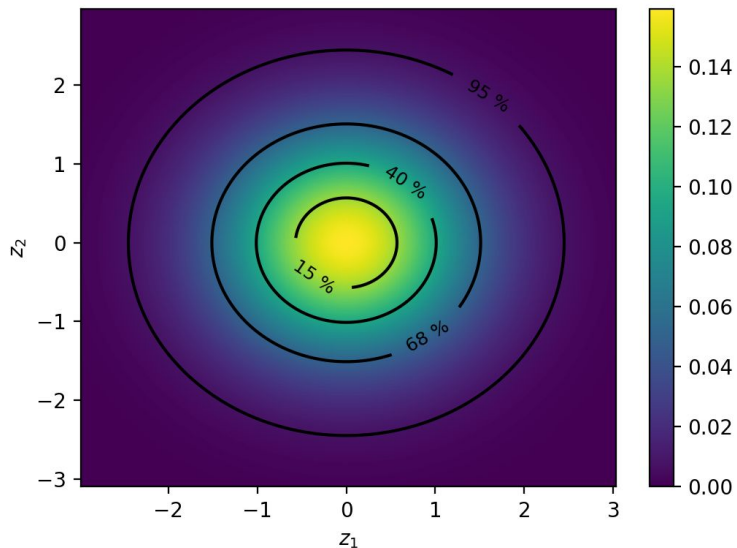
Normalizing-flow PDF

STEP 2



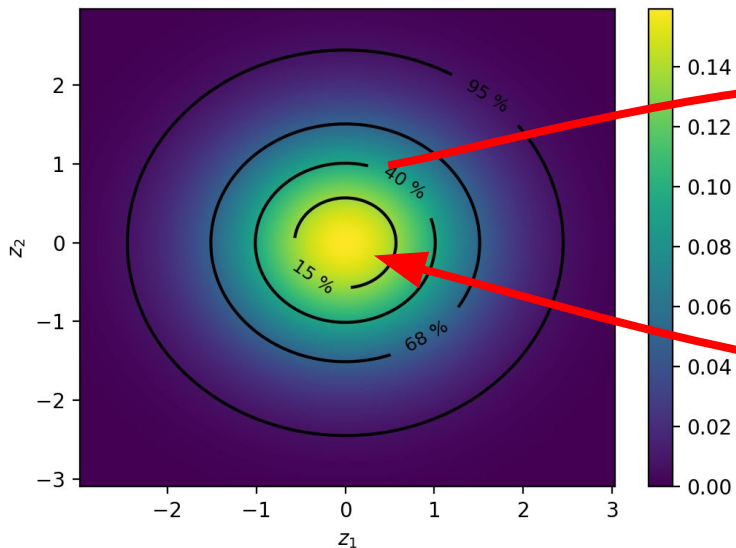
Normalizing flows: flexible PDFs

Auxiliary base

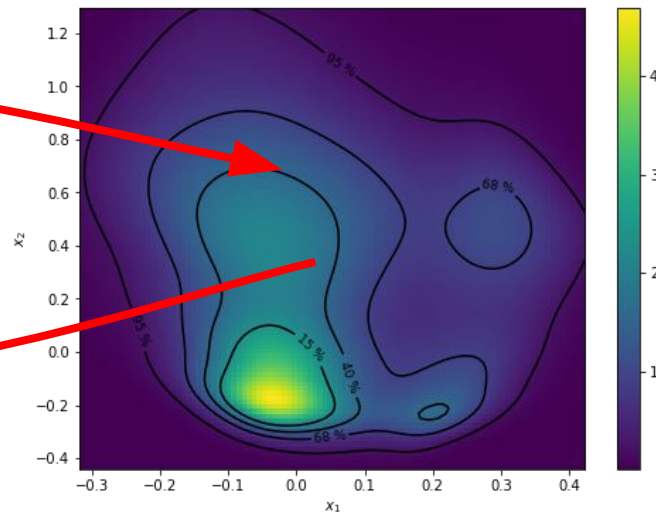


Normalizing flows: flexible PDFs

Auxiliary base



Actual PDF

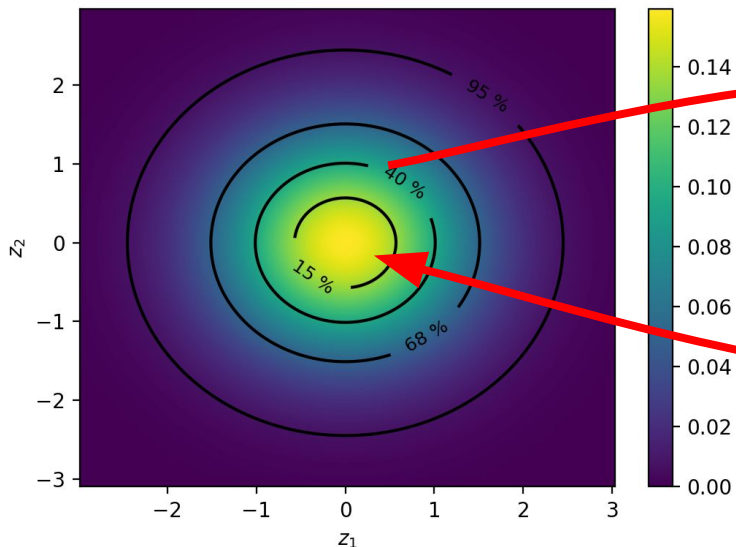


$$f(\vec{z})$$

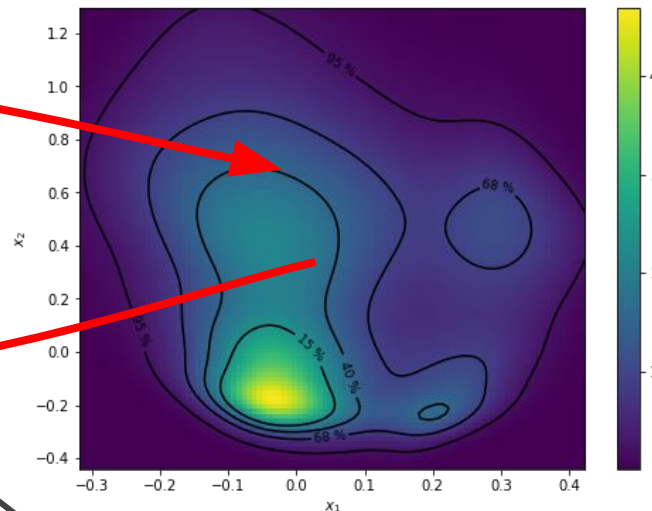
$$f^{-1}(\vec{x})$$

Normalizing flows: flexible PDFs

Auxiliary base



Actual PDF



$f(\vec{z})$

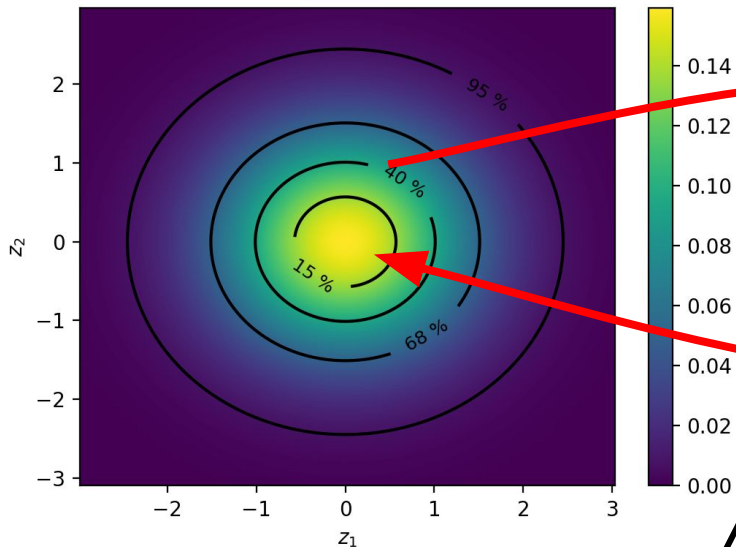
$f^{-1}(\vec{x})$

exact log-probability using inverse

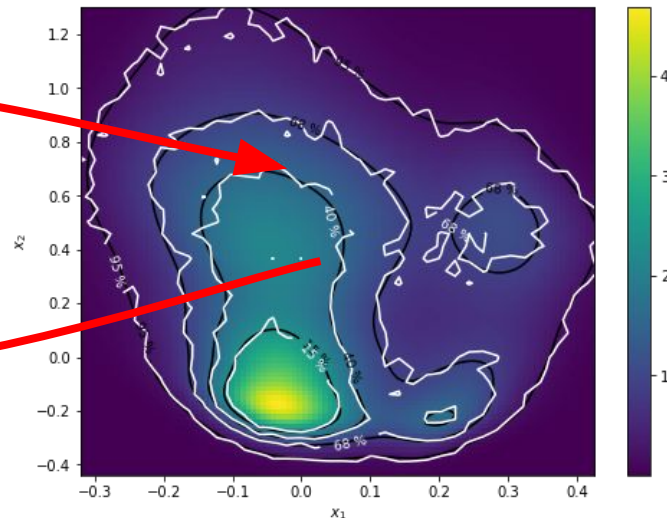
$$p(\vec{x}) = p_0(f^{-1}(\vec{x})) \cdot \left| \det \frac{\partial f^{-1}(\vec{x})}{\partial \vec{x}} \right|$$

Normalizing flows: flexible PDFs

Auxiliary base



Actual PDF



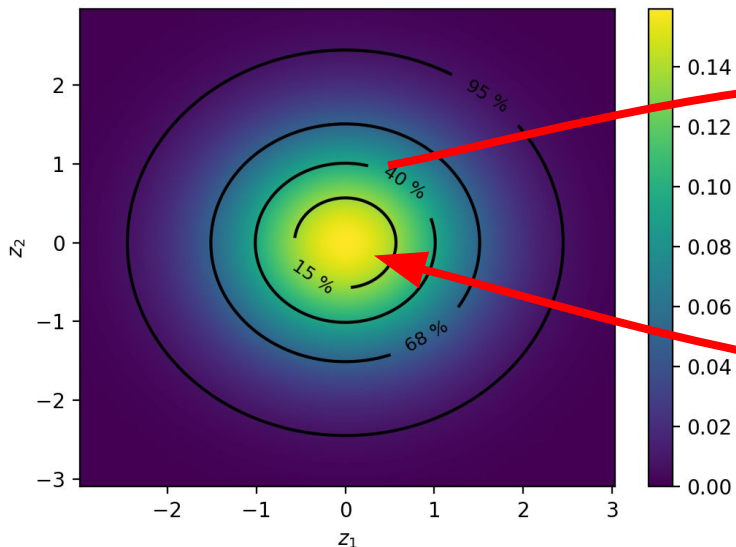
$f(\vec{z})$

$f^{-1}(\vec{x})$

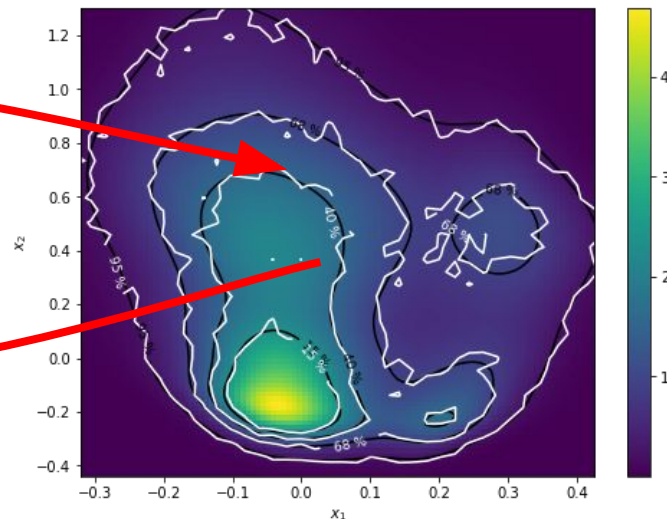
efficient sampling via forward $f(\vec{z})$

Normalizing flows: flexible PDFs

Auxiliary base



Actual PDF



$$f(\vec{z})$$

$$f^{-1}(\vec{x})$$

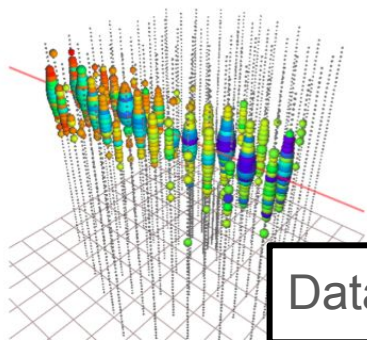
PDF with **very efficient sampling + analytic log-probability**
... no deep learning yet!

Conditioning via NN

-> conditional PDF

$f(z)$ depends on parameters “ u ”: $f_u(z)$

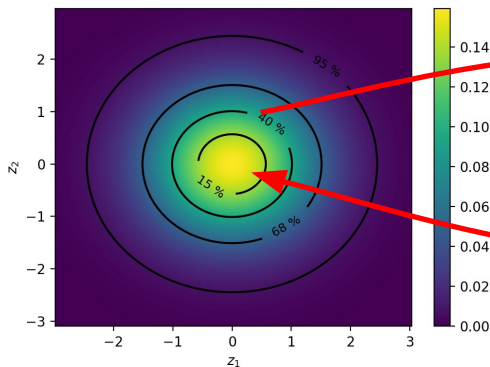
Predict u / diffeomorphism via “ X ”



Data “ X ”

$$u = \text{NN}_{\theta}(X)$$

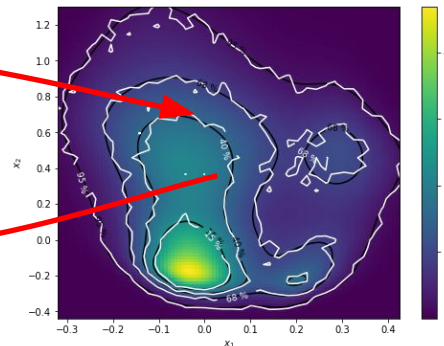
Auxiliary base



$f(\vec{z})$

$f^{-1}(\vec{x})$

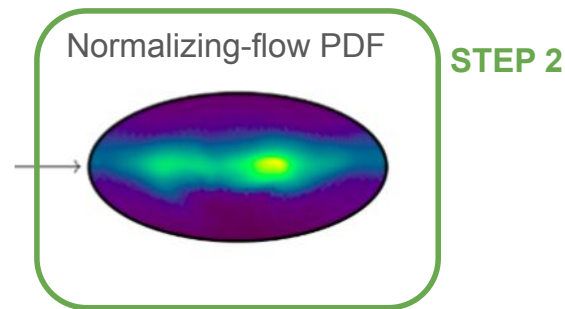
Actual PDF



The PDF - conditional normalizing flow on sphere

The specific normalizing-flow recipe:

Final PDF:
$$p_u(\theta) = p_0(f_u^{-1}(\theta)) \cdot |\det J_u^{-1}(\theta)|$$

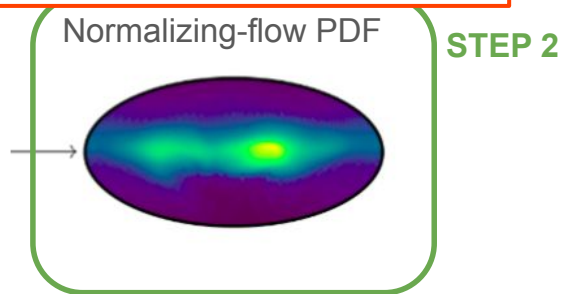


The PDF - conditional normalizing flow on sphere

The specific normalizing-flow recipe:

Final PDF: $p_u(\theta) = p_0(f_u^{-1}(\theta)) \cdot |\det J_u^{-1}(\theta)|$

modified Jacobian for manifold



The PDF - conditional normalizing flow on sphere

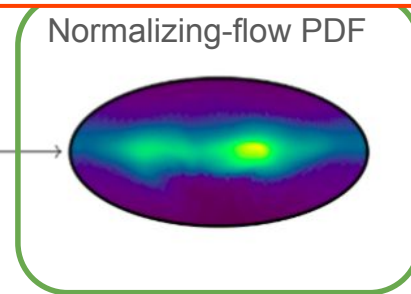
The specific normalizing-flow recipe:

Final PDF: $p_u(\theta) = p_0(f_u^{-1}(\theta)) \cdot |\det J_u^{-1}(\theta)|$

modified Jacobian for manifold

Building blocks of PDF are 15 invertible functions f_{ϕ_i}

$$f_u(\theta) = [f_{\phi_{15}} \circ \dots \circ f_{\phi_1}](\theta)$$



STEP 2

The PDF - conditional normalizing flow on sphere

The specific normalizing-flow recipe:

Final PDF: $p_u(\theta) = p_0(f_u^{-1}(\theta)) \cdot |\det J_u^{-1}(\theta)|$

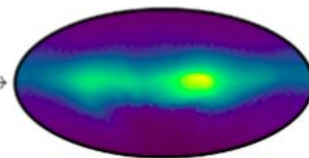
modified Jacobian for manifold

Building blocks of PDF are 15 invertible functions f_{ϕ_i}

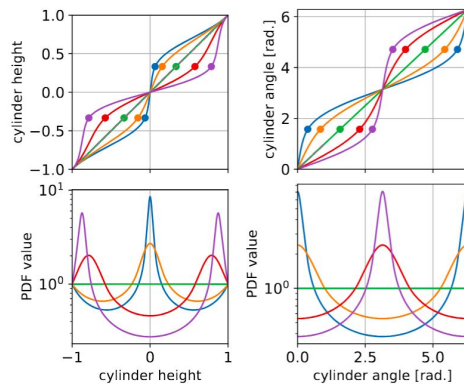
$$f_u(\theta) = [f_{\phi_{15}} \circ \dots \circ f_{\phi_1}](\theta)$$

Normalizing-flow PDF

STEP 2



Each function block f_{ϕ_i}



1) "Smooth RQ Spline"

The PDF - conditional normalizing flow on sphere

The specific normalizing-flow recipe:

Final PDF: $p_u(\theta) = p_0(f_u^{-1}(\theta)) \cdot |\det J_u^{-1}(\theta)|$

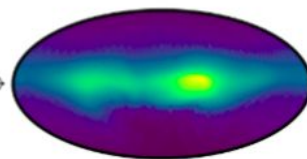
modified Jacobian for manifold

Building blocks of PDF are 15 invertible functions f_{ϕ_i}

$$f_u(\theta) = [f_{\phi_{15}} \circ \dots \circ f_{\phi_1}](\theta)$$

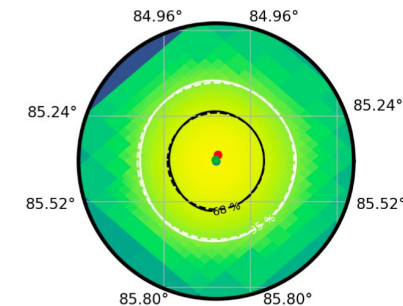
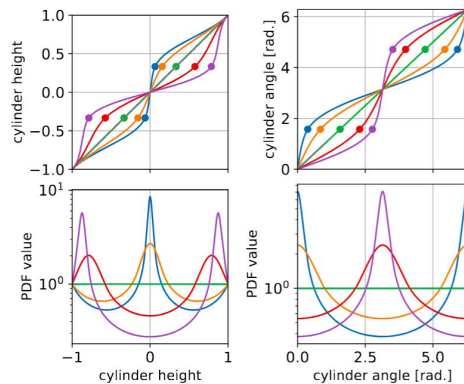
Normalizing-flow PDF

STEP 2



Each function block f_{ϕ_i}

“Fisher zoom”



$$F(z) = 1 + (1/\kappa) \cdot \ln \left(\frac{1+z}{2} + \left(1 - \frac{1+z}{2} \right) \cdot e^{-2\kappa} \right)$$

1) “Smooth RQ Spline”

2) von-Mises Fisher

The PDF - conditional normalizing flow on sphere

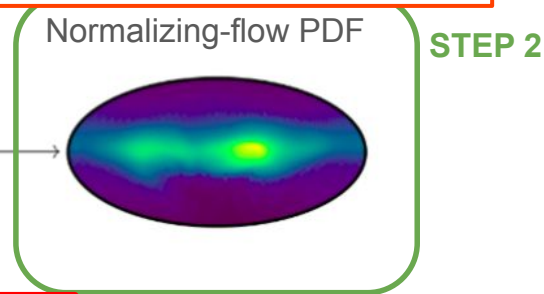
The specific normalizing-flow recipe:

Final PDF: $p_u(\theta) = p_0(f_u^{-1}(\theta)) \cdot |\det J_u^{-1}(\theta)|$

modified Jacobian for manifold

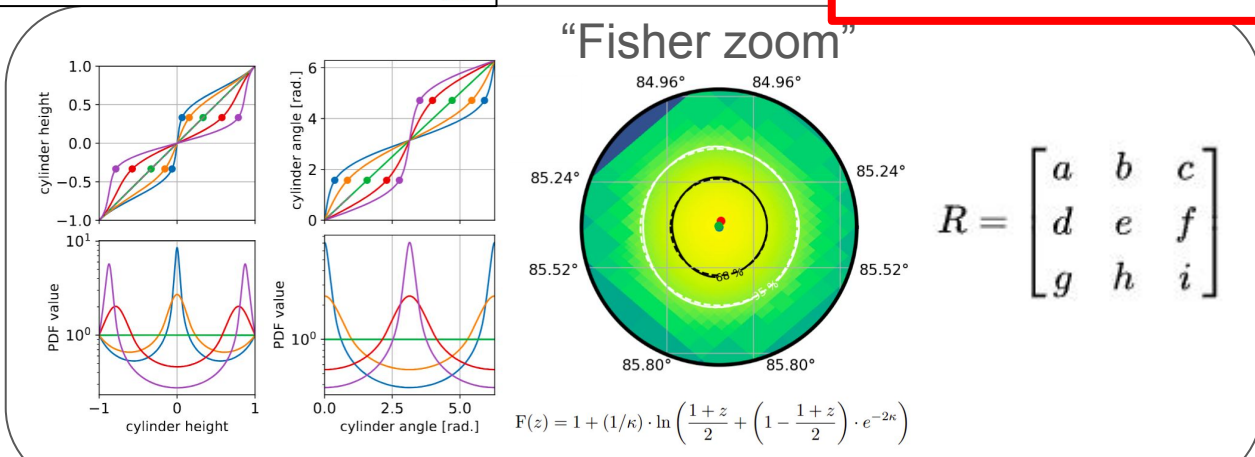
Building blocks of PDF are 15 invertible functions f_{ϕ_i}

$$f_u(\theta) = [f_{\phi_{15}} \circ \dots \circ f_{\phi_1}](\theta)$$



Each function block f_{ϕ_i}

~ 220 params



- “Mechanistic” way of PDF creation
- 15 blocks seem sufficient for all PDFs in IceCube
- von-Mises Fisher is special case

1) “Smooth RQ Spline”

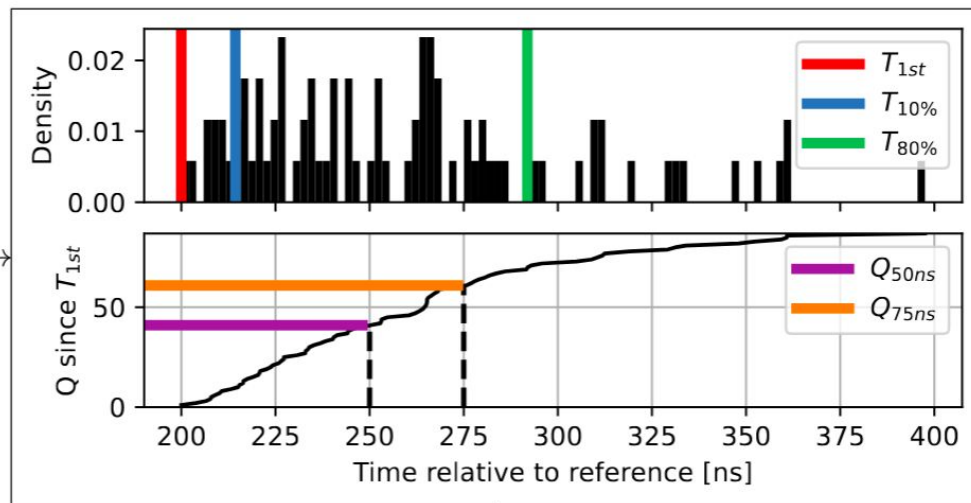
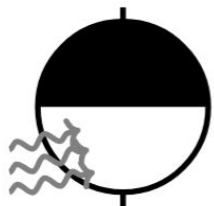
2) von-Mises Fisher

3) Rotation

Encoding

Step 1 - encoding the data

Photons hit DOM



Can include all dom types

Normal DOM $\rightarrow T_i = (\tilde{x}, \tilde{y}, \tilde{z}, \tilde{c}_x, \tilde{c}_y, \tilde{c}_z, \tilde{\vec{Q}}, \tilde{T}_{1st}, \tilde{\vec{T}}_o, \text{PMT}_{\text{type}}, 1, 0, 0)$

Saturated DOM $\rightarrow T_i = (\tilde{x}, \tilde{y}, \tilde{z}, \tilde{c}_x, \tilde{c}_y, \tilde{c}_z, \vec{0}, \tilde{T}_{1st}, \vec{0}, \text{PMT}_{\text{type}}, 0, 1, 0)$

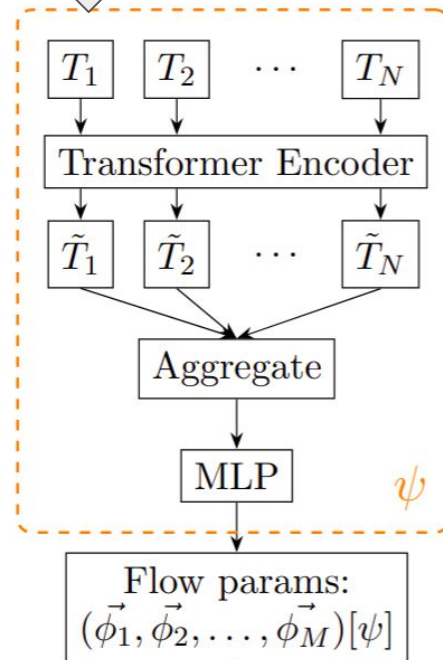
Empty DOM $\rightarrow T_i = (\tilde{x}, \tilde{y}, \tilde{z}, \tilde{c}_x, \tilde{c}_y, \tilde{c}_z, \vec{0}, 0, \vec{0}, \text{PMT}_{\text{type}}, 0, 0, 1)$

Step 1 - encoding the data

Normal DOM $\rightarrow T_i = (\tilde{x}, \tilde{y}, \tilde{z}, \tilde{c}_x, \tilde{c}_y, \tilde{c}_z, \tilde{\mathbf{Q}}, \tilde{T}_{1st}, \tilde{\mathbf{T}}_o, \text{PMT}_{\text{type}}, 1, 0, 0)$

Saturated DOM $\rightarrow T_i = (\tilde{x}, \tilde{y}, \tilde{z}, \tilde{c}_x, \tilde{c}_y, \tilde{c}_z, \mathbf{0}, \tilde{T}_{1st}, \mathbf{0}, \text{PMT}_{\text{type}}, 0, 1, 0)$

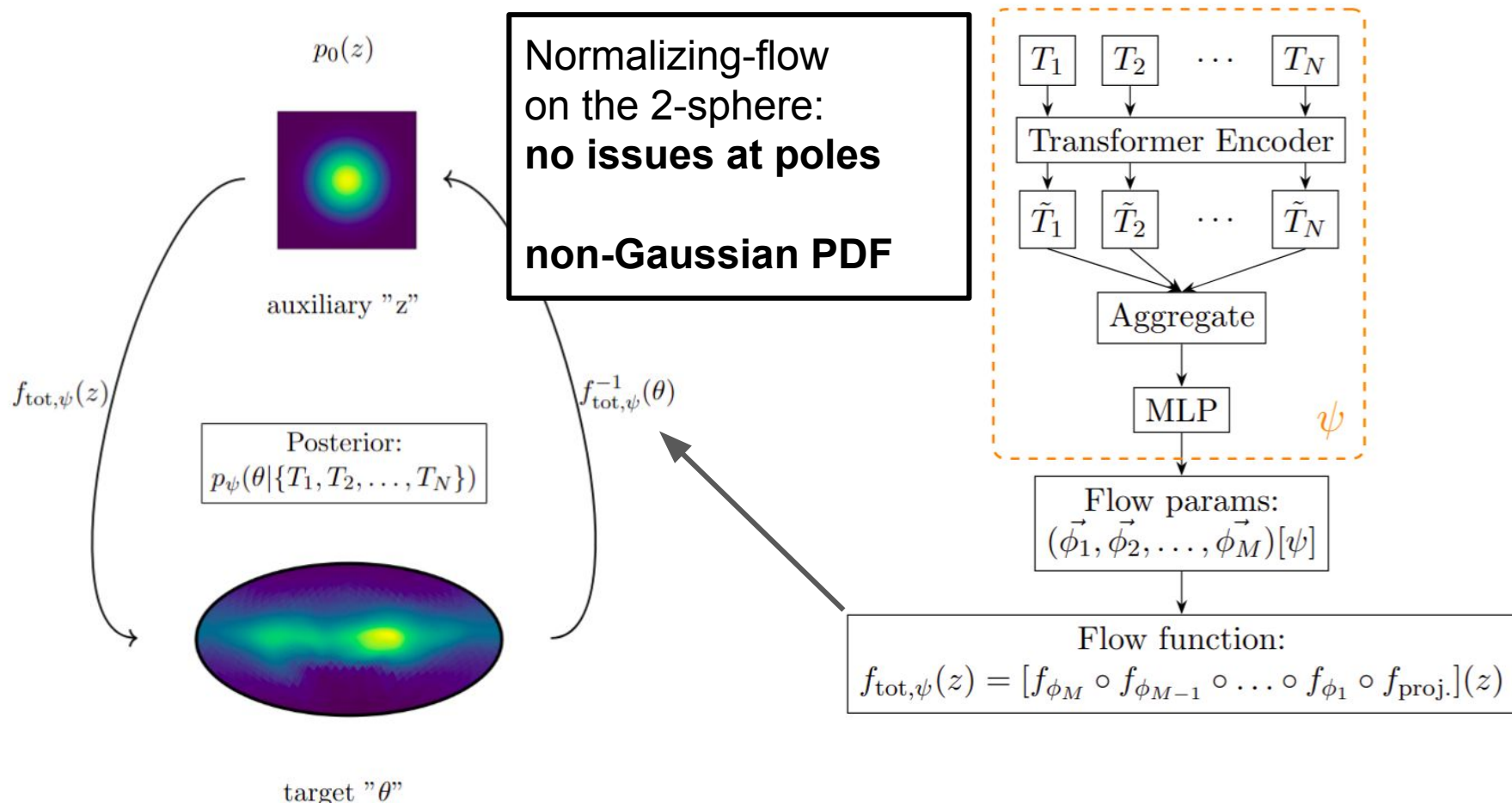
Empty DOM $\rightarrow T_i = (\tilde{x}, \tilde{y}, \tilde{z}, \tilde{c}_x, \tilde{c}_y, \tilde{c}_z, \mathbf{0}, 0, \mathbf{0}, \text{PMT}_{\text{type}}, 0, 0, 1)$



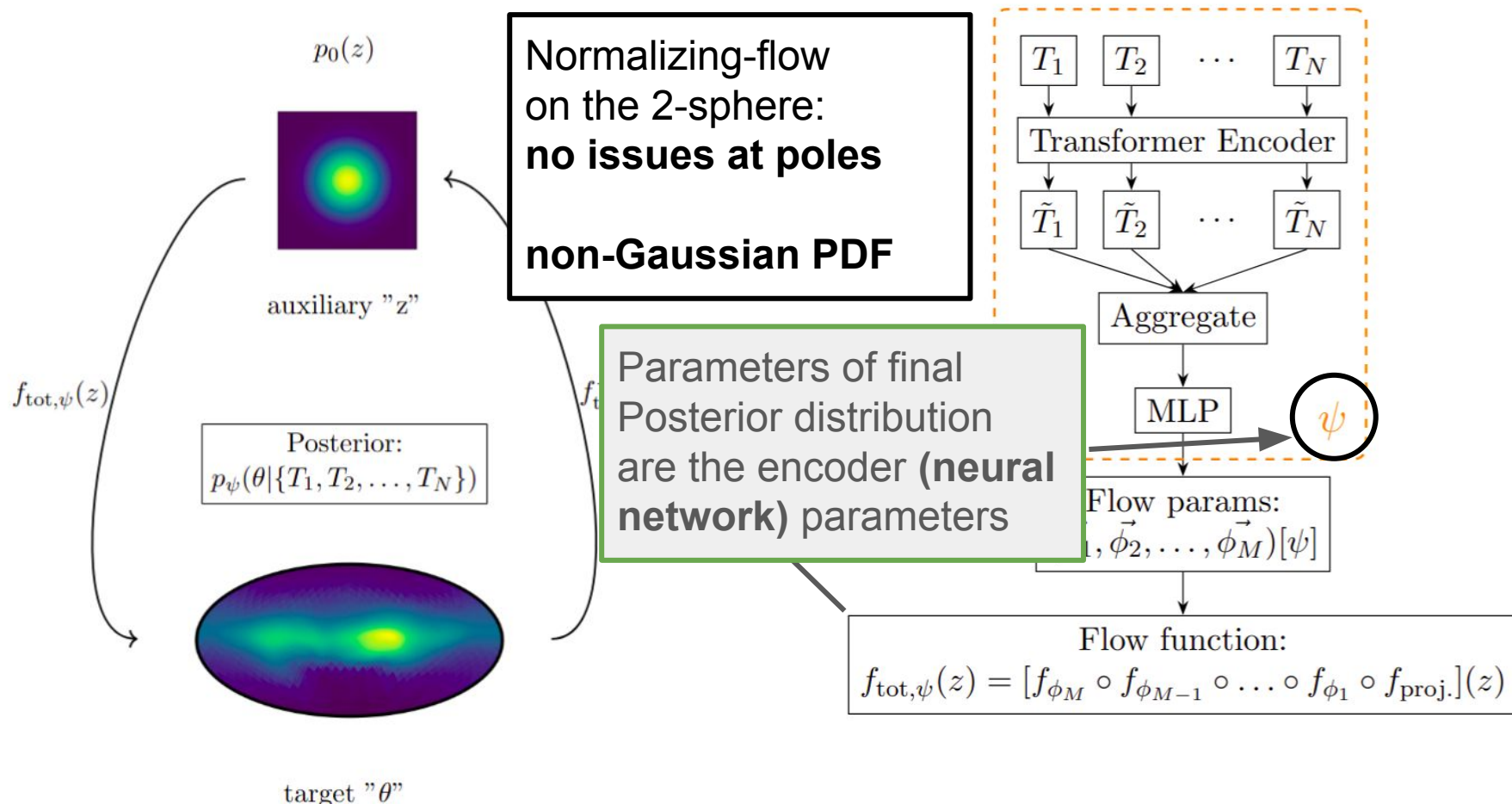
Flow params:
 $(\vec{\phi}_1, \vec{\phi}_2, \dots, \vec{\phi}_M)[\psi]$

Flow function:
 $f_{\text{tot}, \psi}(z) = [f_{\phi_M} \circ f_{\phi_{M-1}} \circ \dots \circ f_{\phi_1} \circ f_{\text{proj}}.](z)$

Step 2 - predicting the posterior normalizing flow

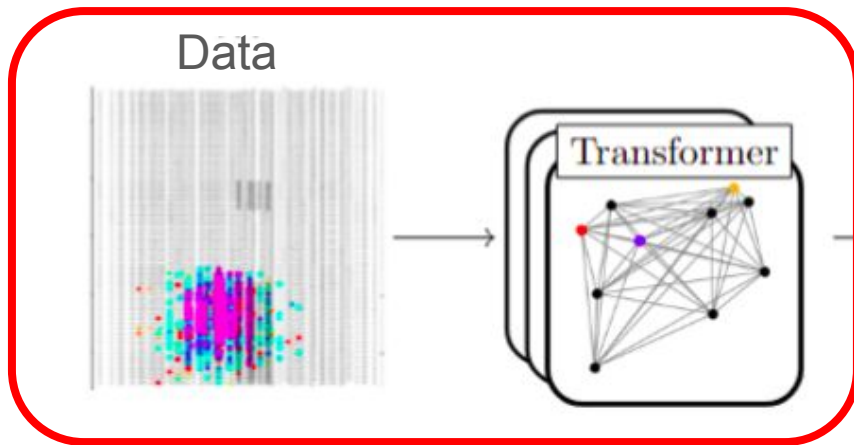


Step 2 - predicting the posterior normalizing flow



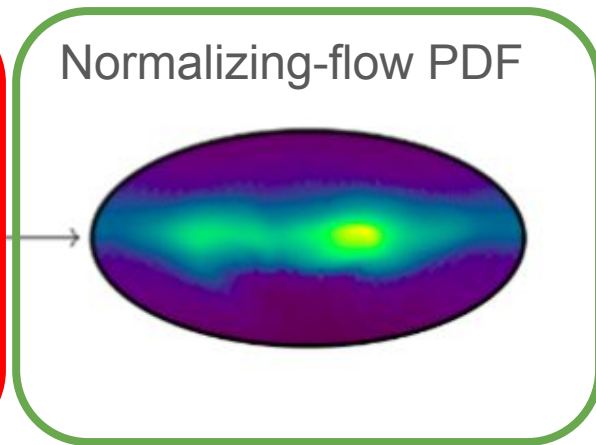
The model: a posterior PDF prediction machine

STEP 1



Normalizing-flow PDF

STEP 2



data encoding / compression

PDF prediction from
compressed data

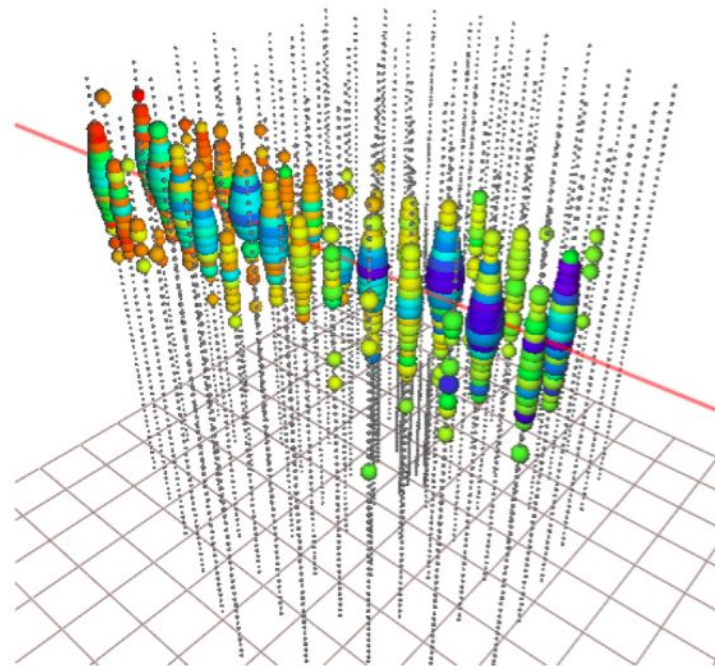
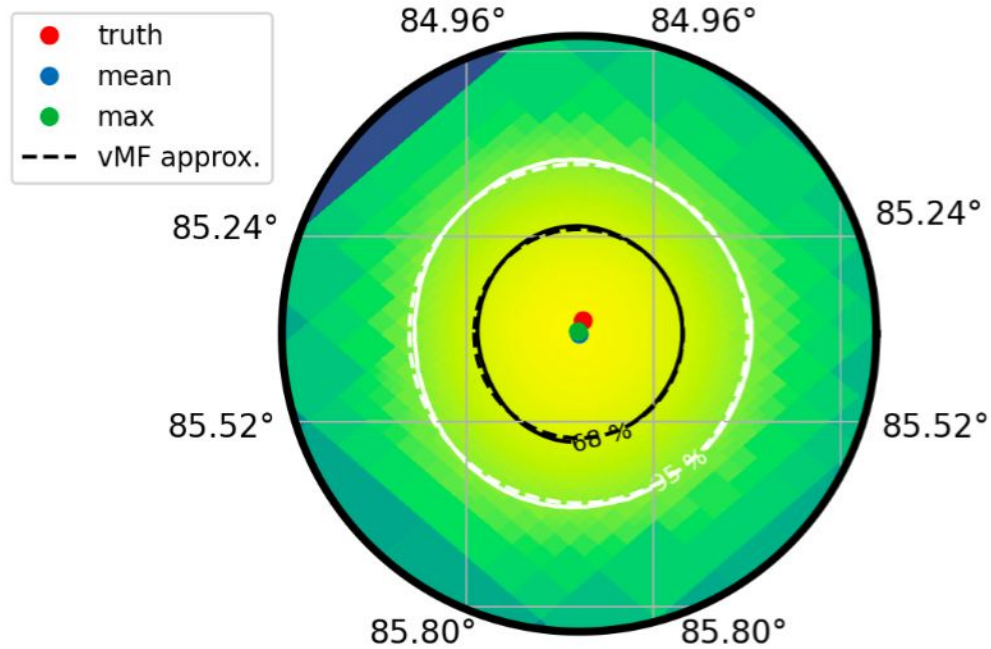
Target: 2-dim (sphere)

Posterior:

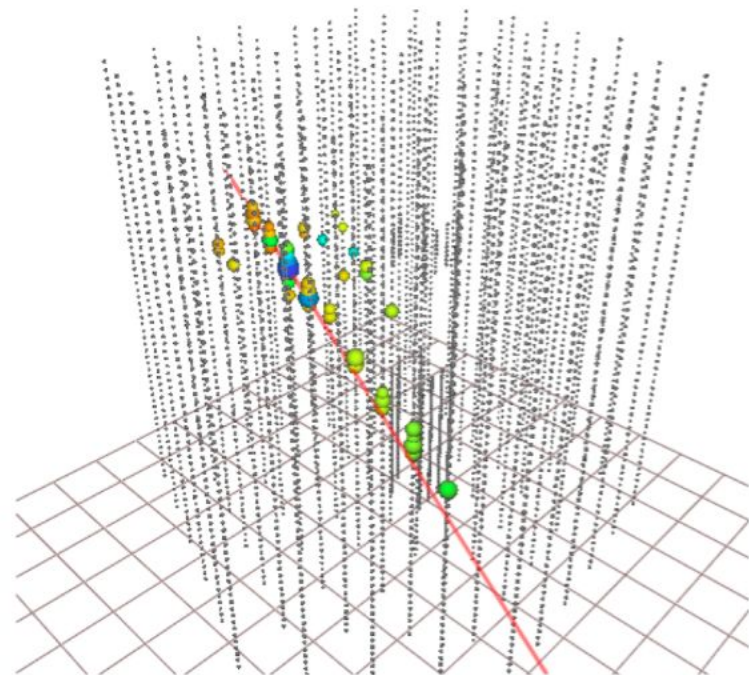
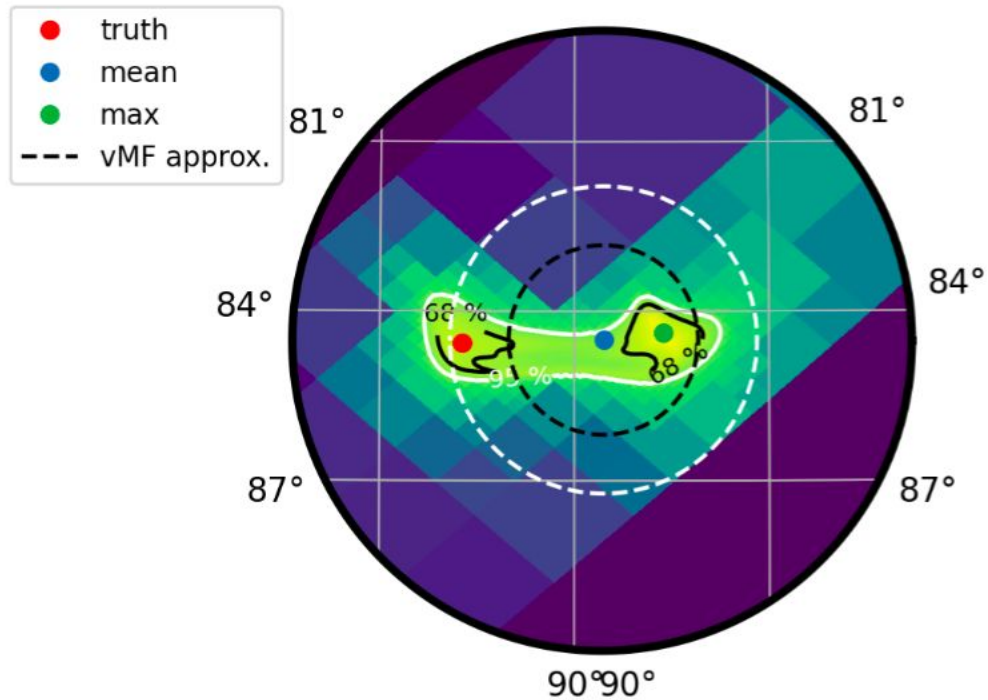
5-10 mio parameters

$$p_{\psi}(\theta | \{T_1, T_2, \dots, T_N\})$$

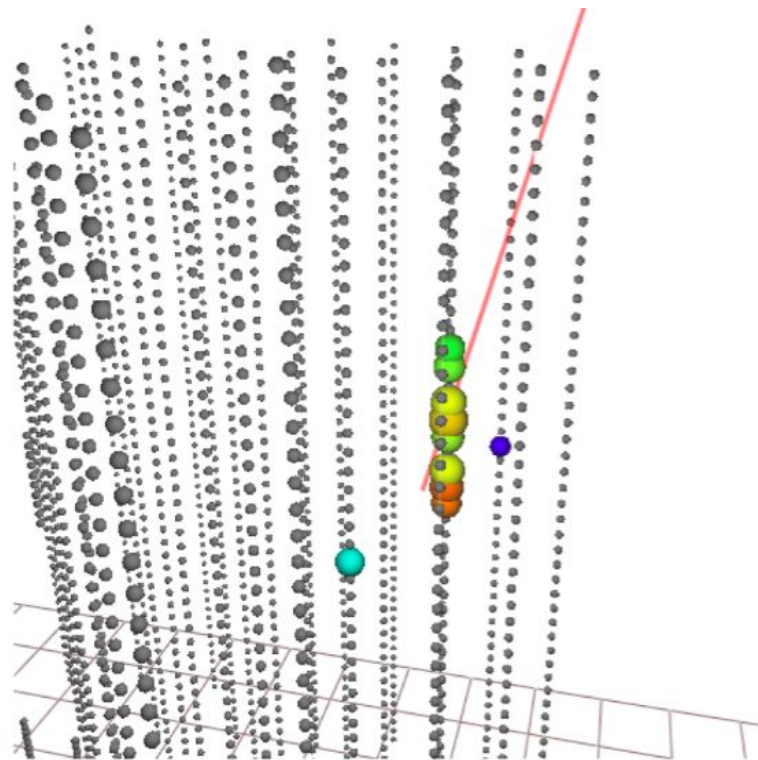
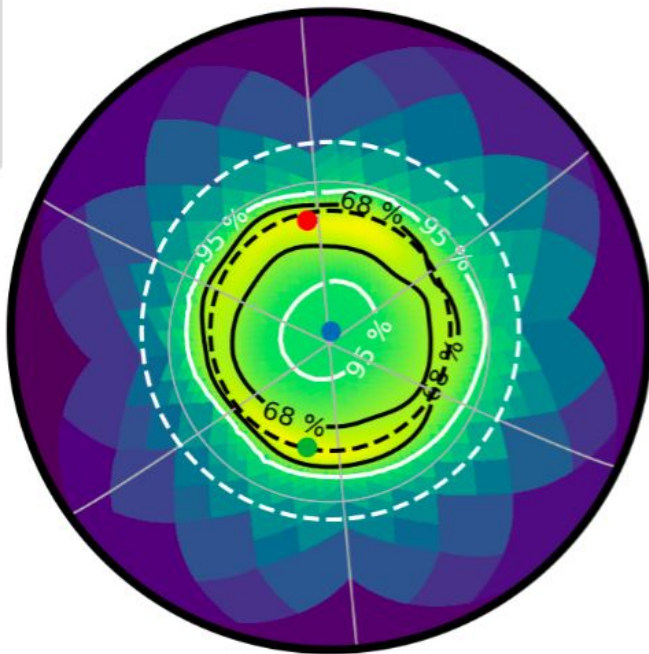
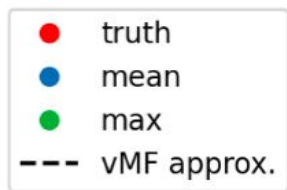
Input: ~ 100-100k dim
(IceCube data)



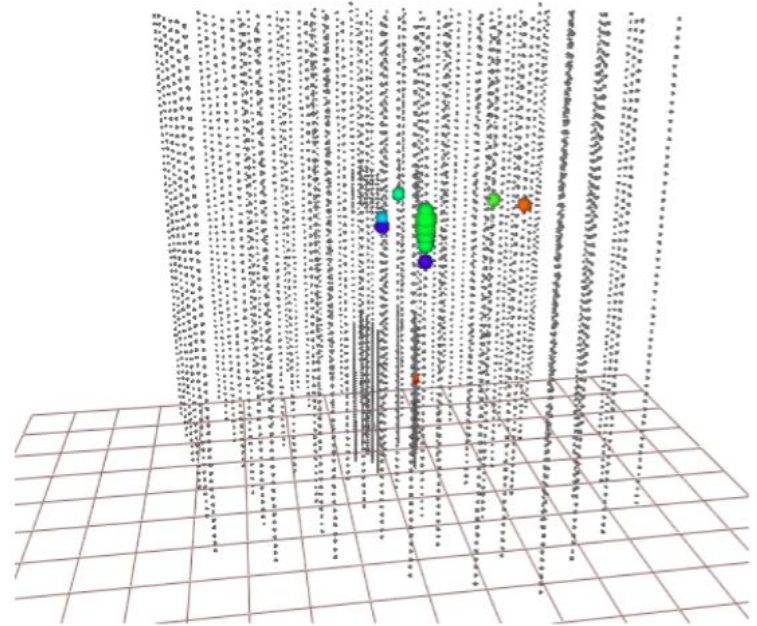
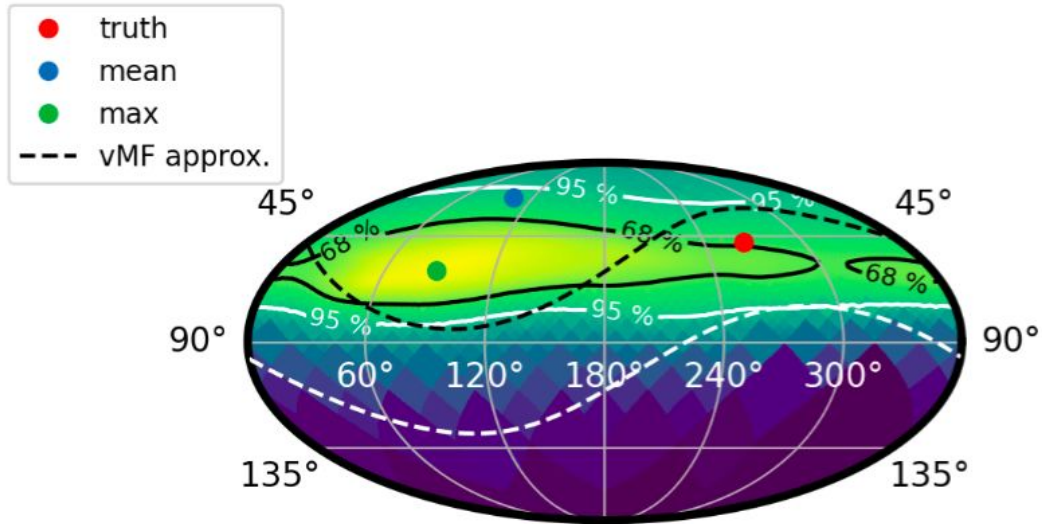
(a) Bright horizontal track.



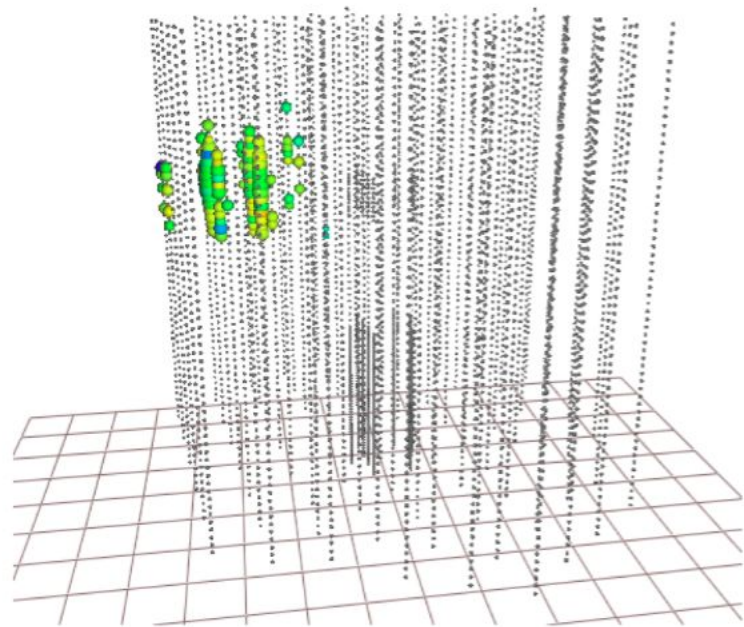
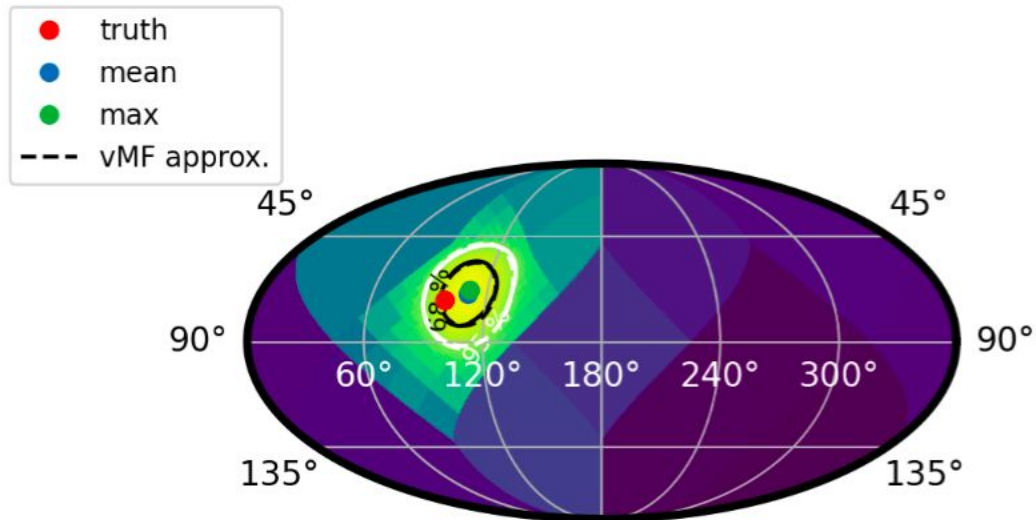
(b) Muon track along string lines.



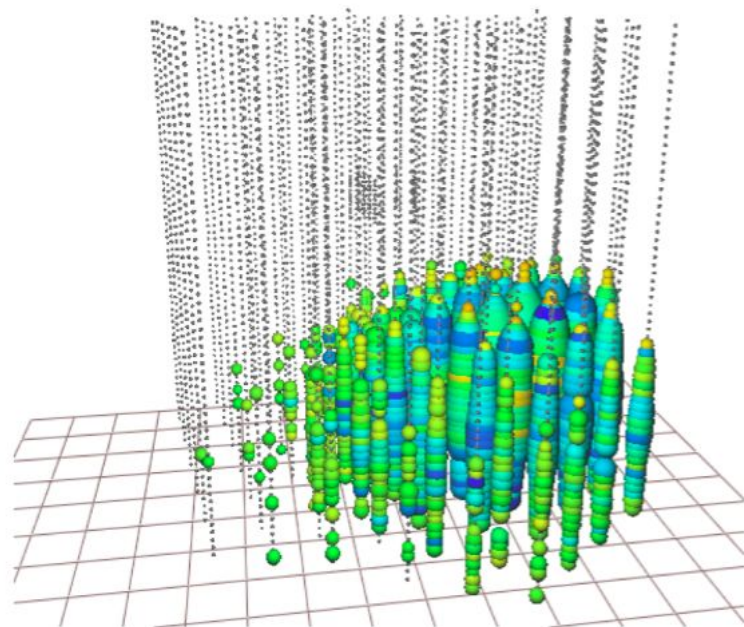
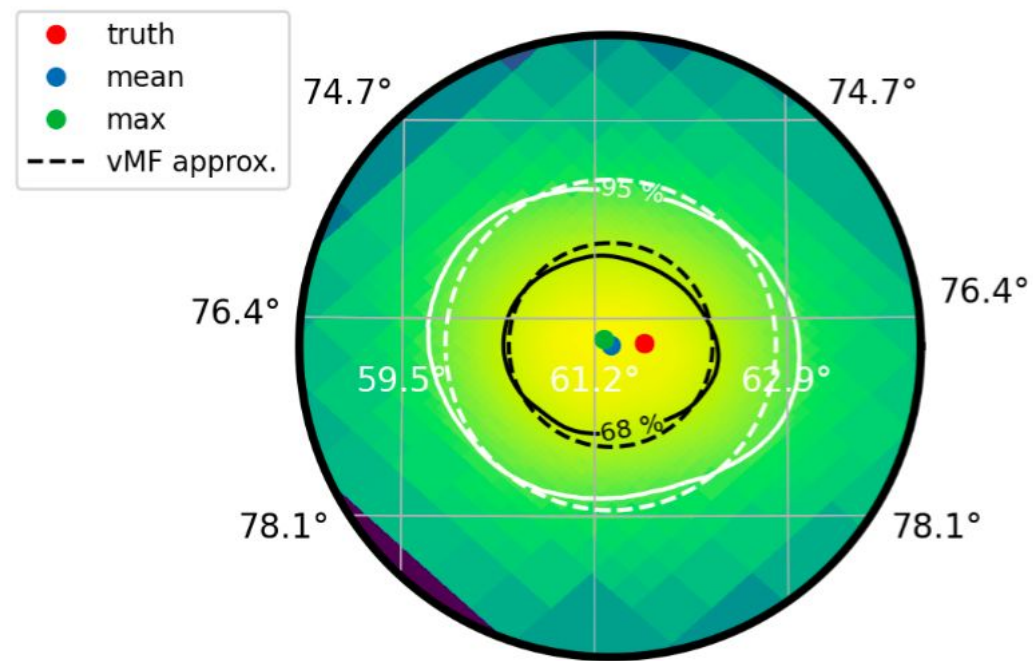
(c) Low energy contained upgoing track along a single string line.



(a) A low-energy shower that is mostly visible in a single string.



(b) A shower which is visible in several neighboring strings.



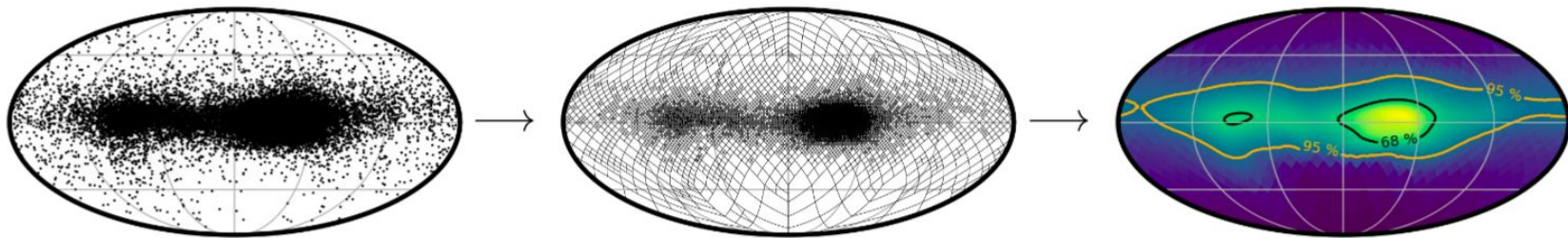
(c) A multi-PeV shower in the clear ice.

Adaptive grid in constant time (extended “all-sky” posterior)

samples

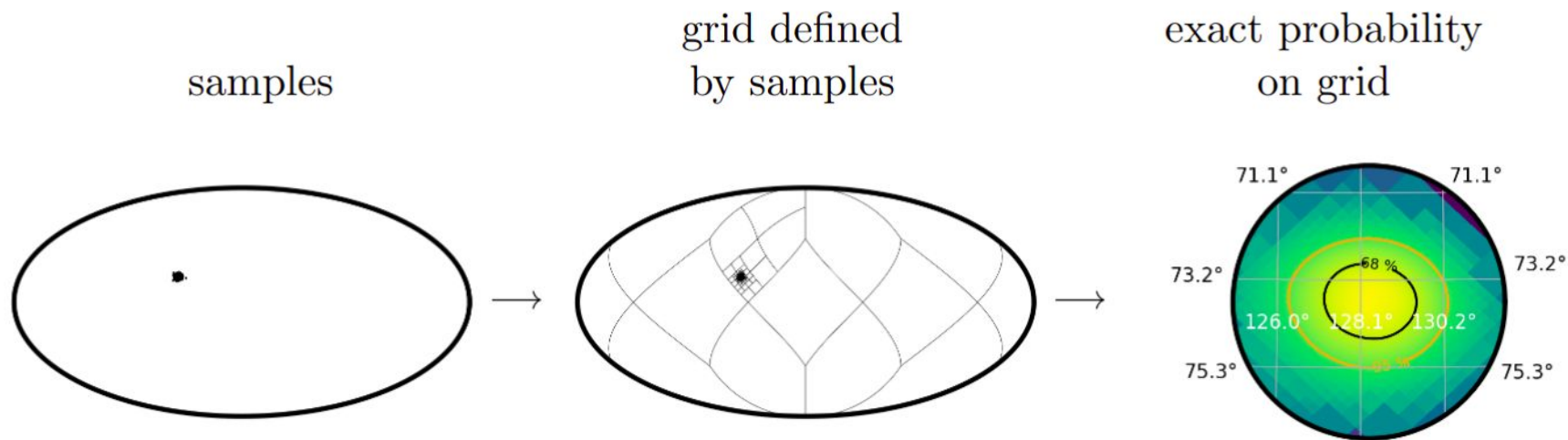
grid defined
by samples

exact probability
on grid



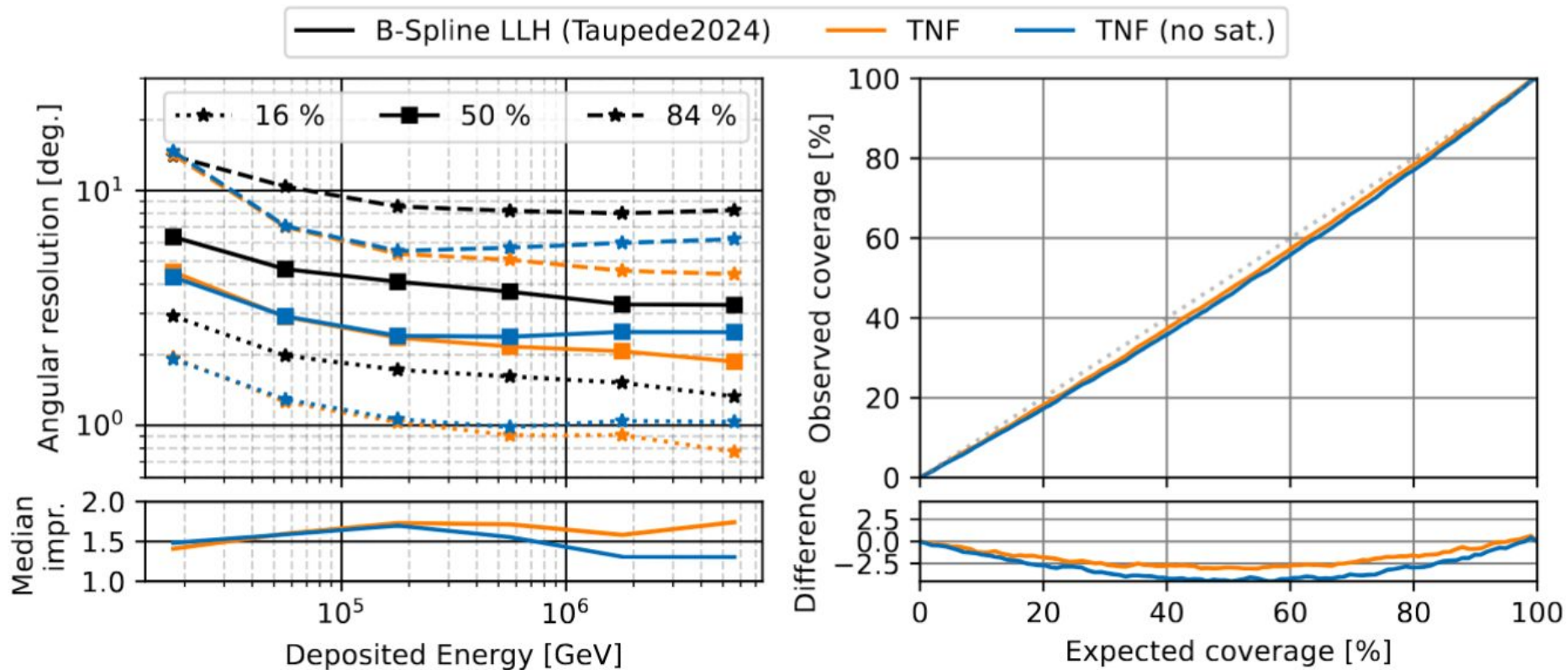
Normalizing flows: sampling + exact log-pdf evaluation **simultaneously!**

Adaptive grid in constant time (small localized posterior)



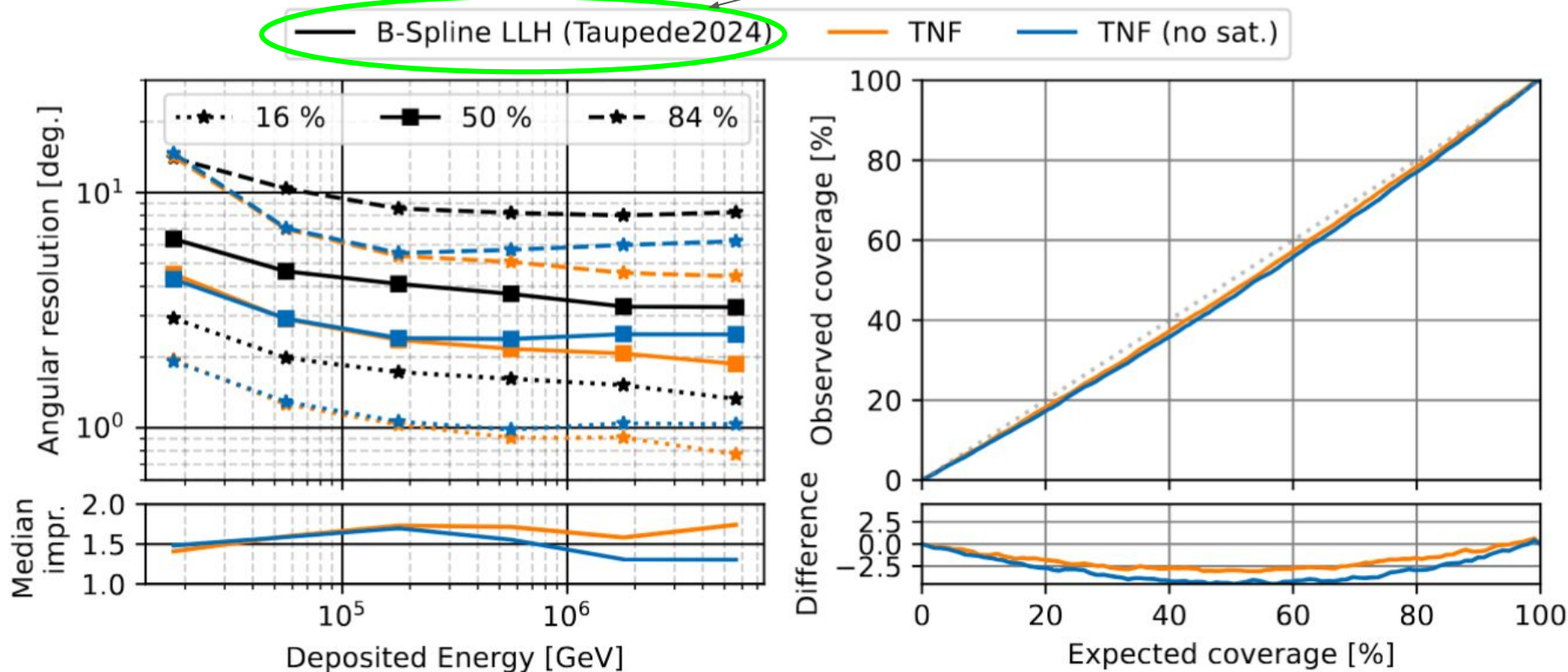
Normalizing flows: sampling + exact log-pdf evaluation **simultaneously!**

Angular resolution - showers

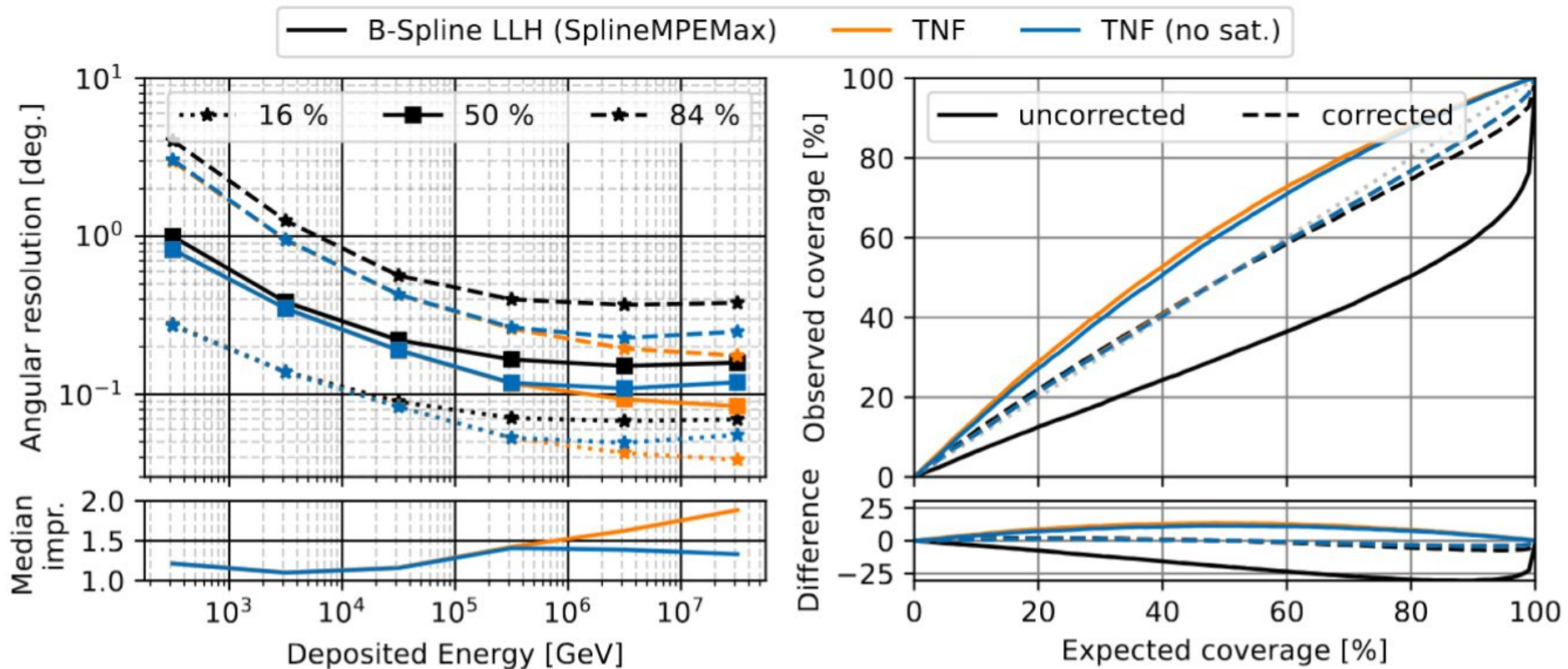


Angular resolution - sho

Used in most recent shower sample for Galactic Plane analysis (PoS-ICRC2025-1193 / 2507.08753)



Throughgoing tracks (>300 m)



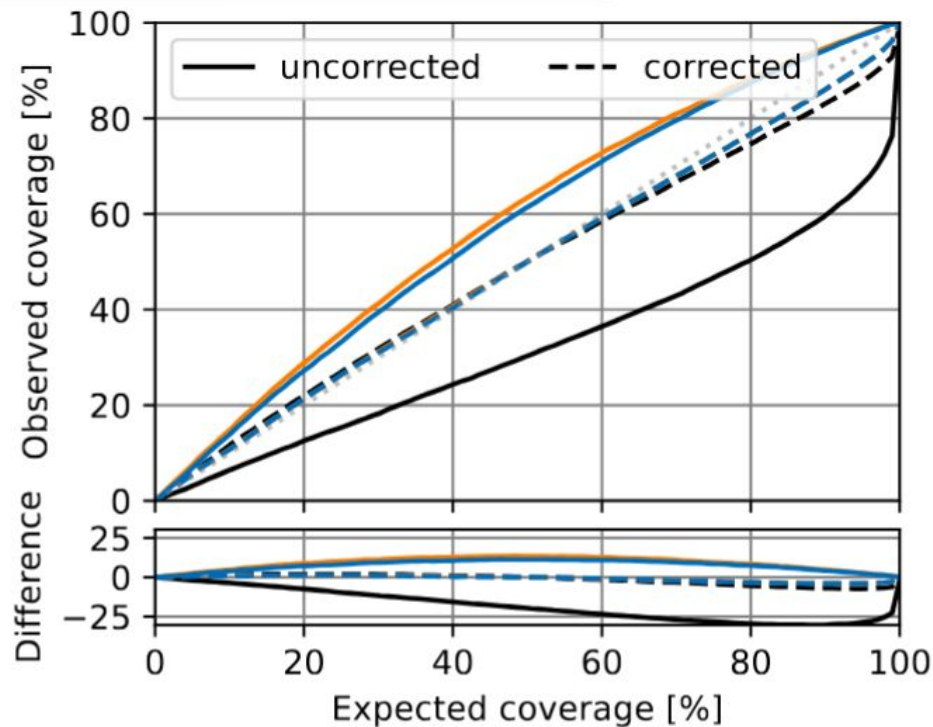
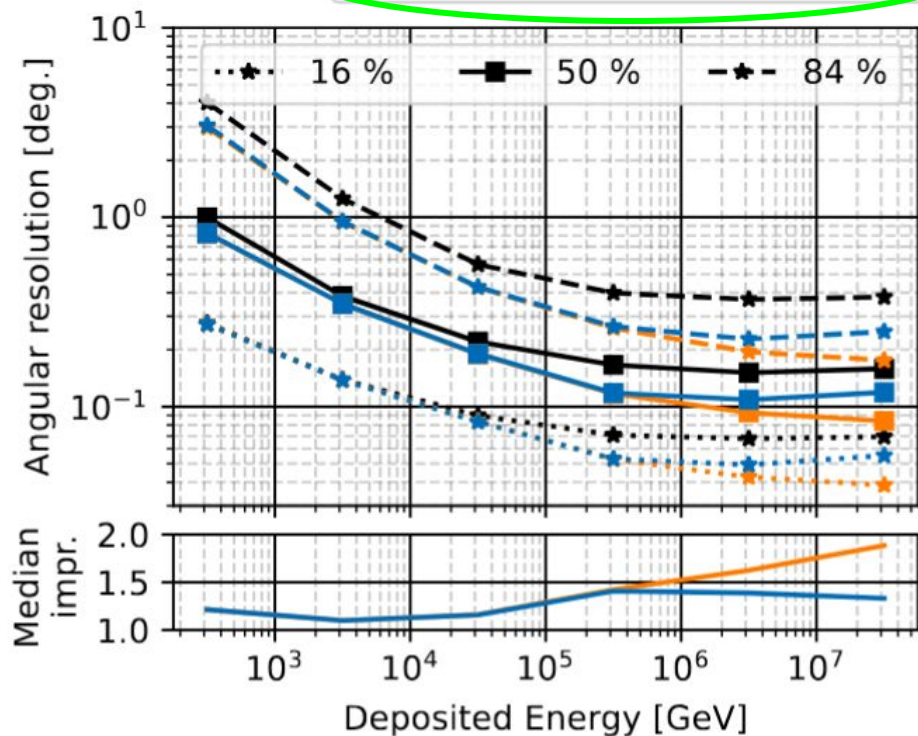
Throughgoing tracks (>300 m)

Used in all Point source results since 2013

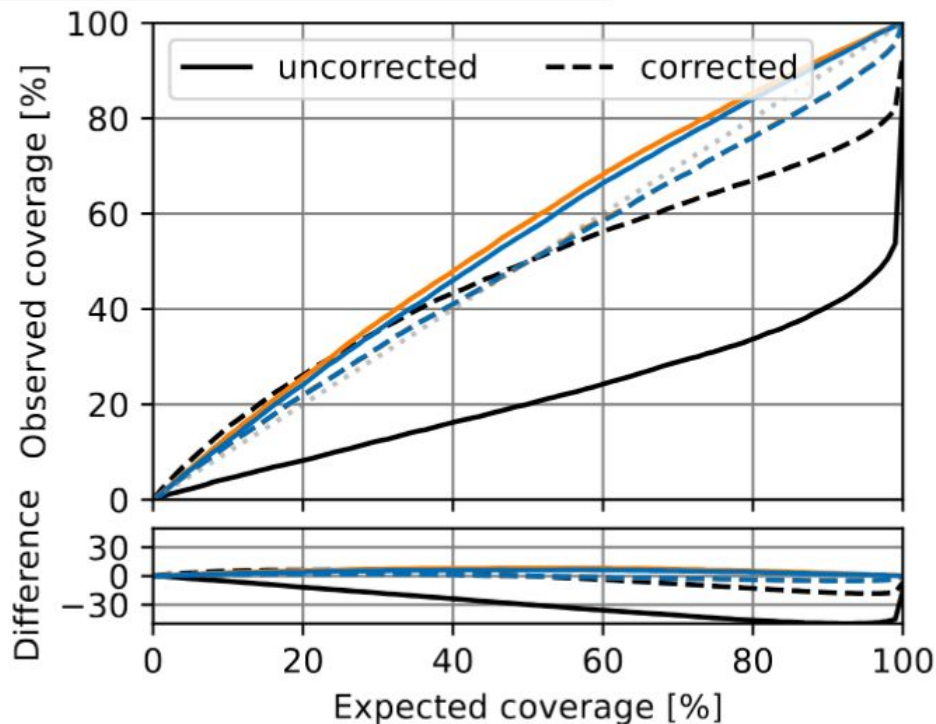
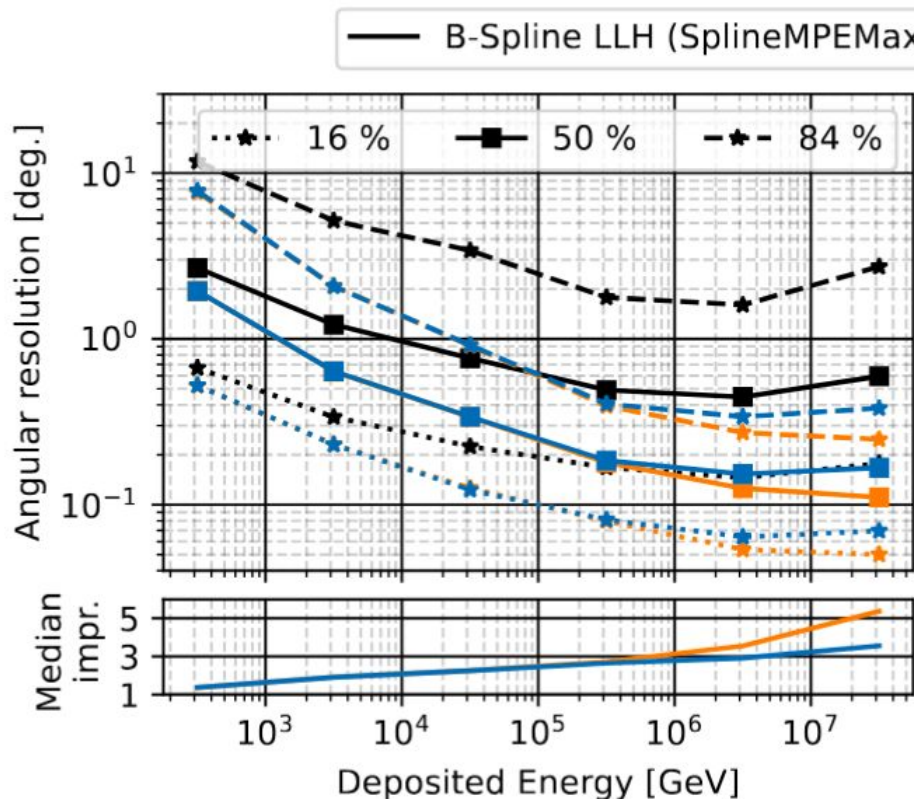
B-Spline LLH (SplineMPEMax)

TNF

TNF (no sat.)



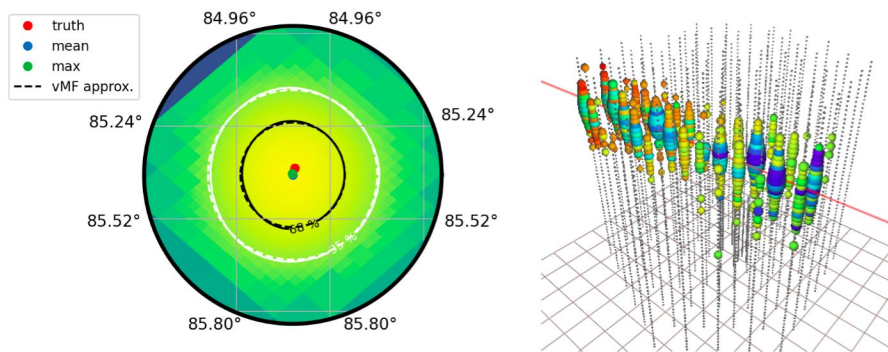
Starting tracks (>300 m)



Substantial improvement for starting tracks! Relevant for Galactic plane!

Summary

- Transformer Normalizing Flows (TNF) are a suitable model for neutrino detectors (arXiv:2604.19846)
- Explicitly learn full posterior from data: $p_{\psi}(\theta|\{T_1, T_2, \dots, T_N\})$



(a) Bright horizontal track.

Summary

- Transformer Normalizing Flows (TNF) are a suitable model for neutrino detectors (arXiv:2604.19846)
- Explicitly learn full posterior from data: $p_{\psi}(\theta|\{T_1, T_2, \dots, T_N\})$

Transformer -> Can respect permutation invariance (Bayesian symmetry)

Normalizing flow -> Flexible PDF (generalizes standard regression)

$$p(\theta|\mathbf{x}) = \frac{\overset{\text{permute}}{\overbrace{p(x_1;\theta) \cdot p(x_2;\theta) \cdot p(x_3;\theta) \cdot \dots \cdot p(\theta)}}}{p(\mathbf{x})} = \frac{p(x_1;\theta) \cdot p(x_3;\theta) \cdot p(x_2;\theta) \cdot \dots \cdot p(\theta)}{p(\mathbf{x})}$$

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- New state of the art in angular resolution: showers and tracks
- Full posterior skymap for astronomy - irrespective of shape+size in 1s

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- Normalizing flow implemented in:

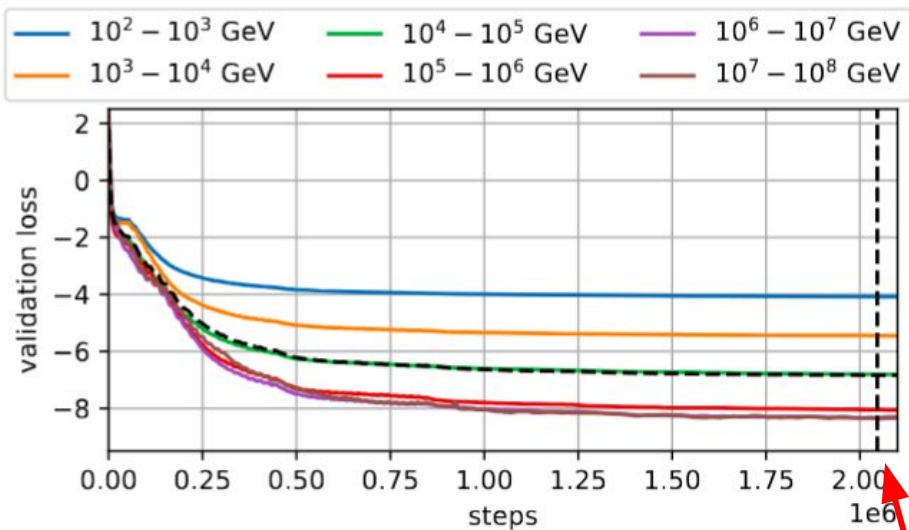
https://github.com/thoglu/jammy_flows/

$$f_u(\theta) = [f_{\phi_{15}} \circ \dots \circ f_{\phi_1}](\theta)$$

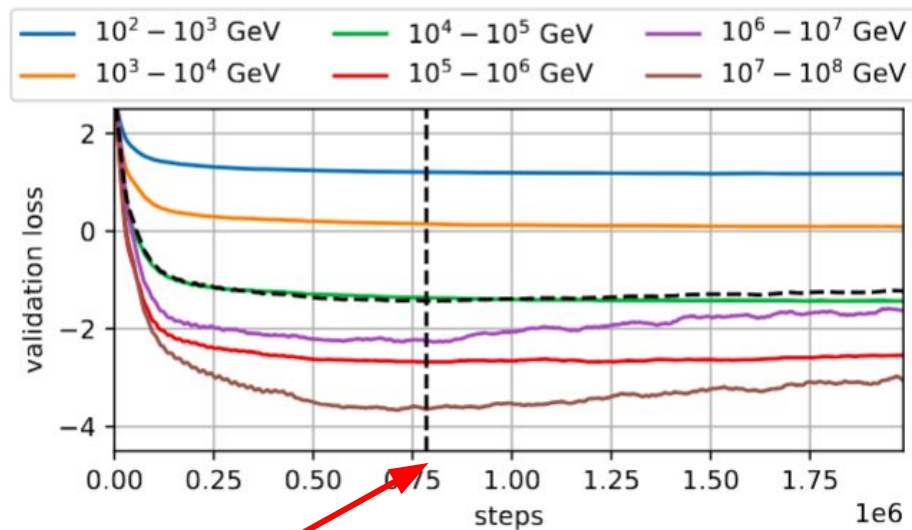
```
pdf=jammy_flows.pdf("s2", "ffffffffffffffff", options_overwrite=opt_dict)
```

Backup

Example validation curves



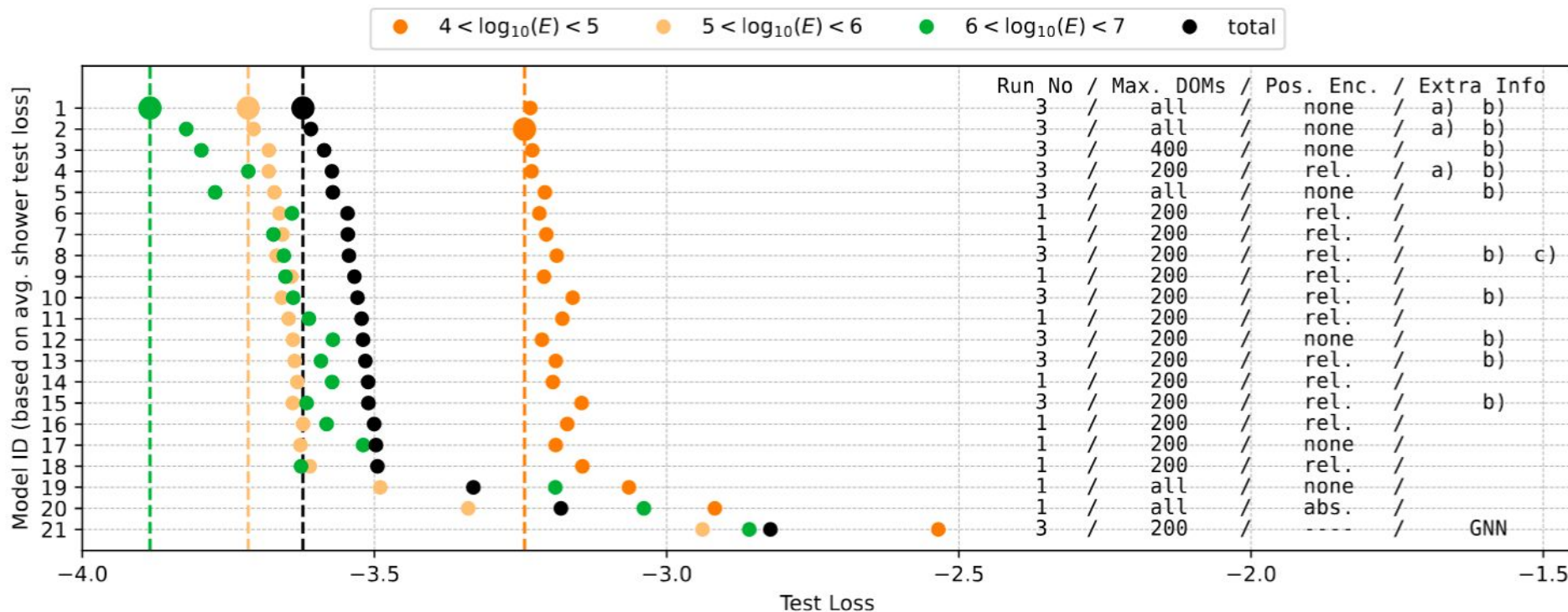
(a) Best track model.



(b) Best shower model

Model frozen for testing

Showers: test results on ~20 different shower model trainings



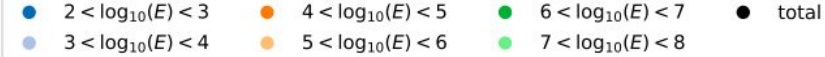
“Total” = average over individual log-energy results

Best run 3

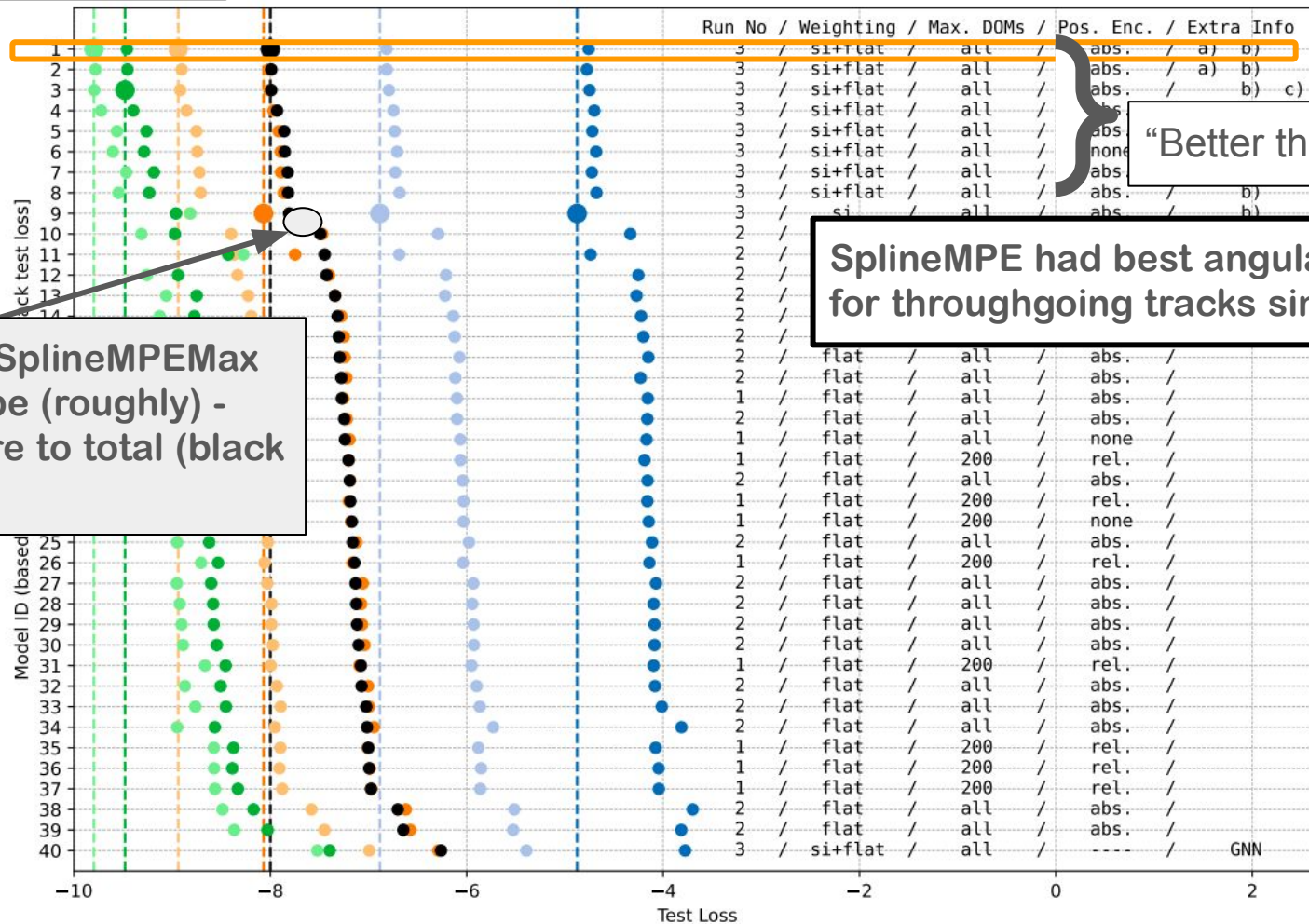


51

Throughgoing tracks



Test loss

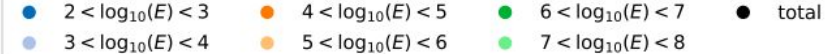


"Better than SplineMPE"

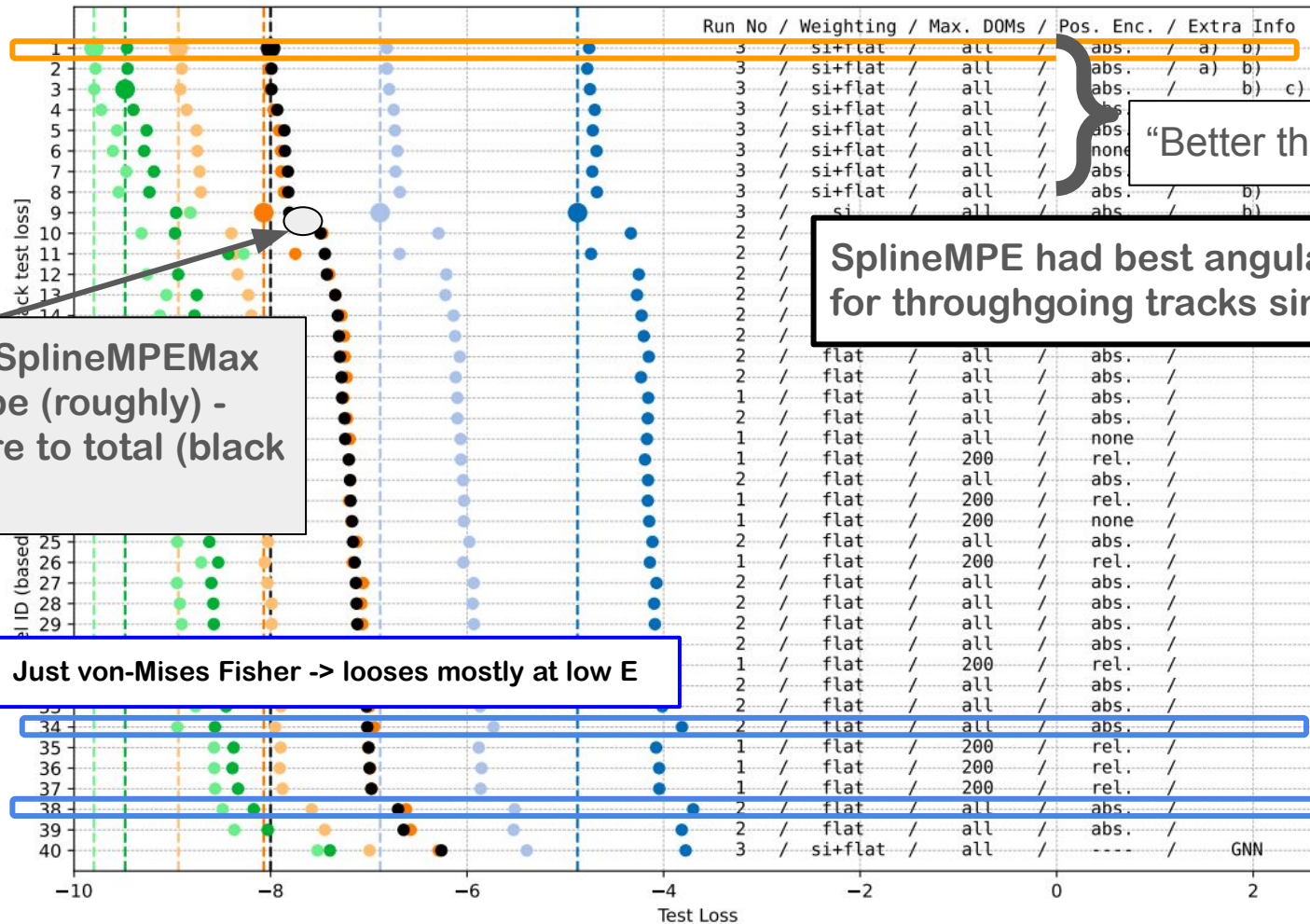
SplineMPE had best angular resolution for throughgoing tracks since 2013!

Where SplineMPEMax would be (roughly) - compare to total (black dots)

Throughgoing tracks



Test loss



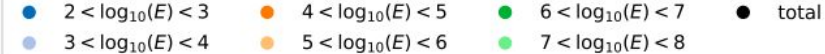
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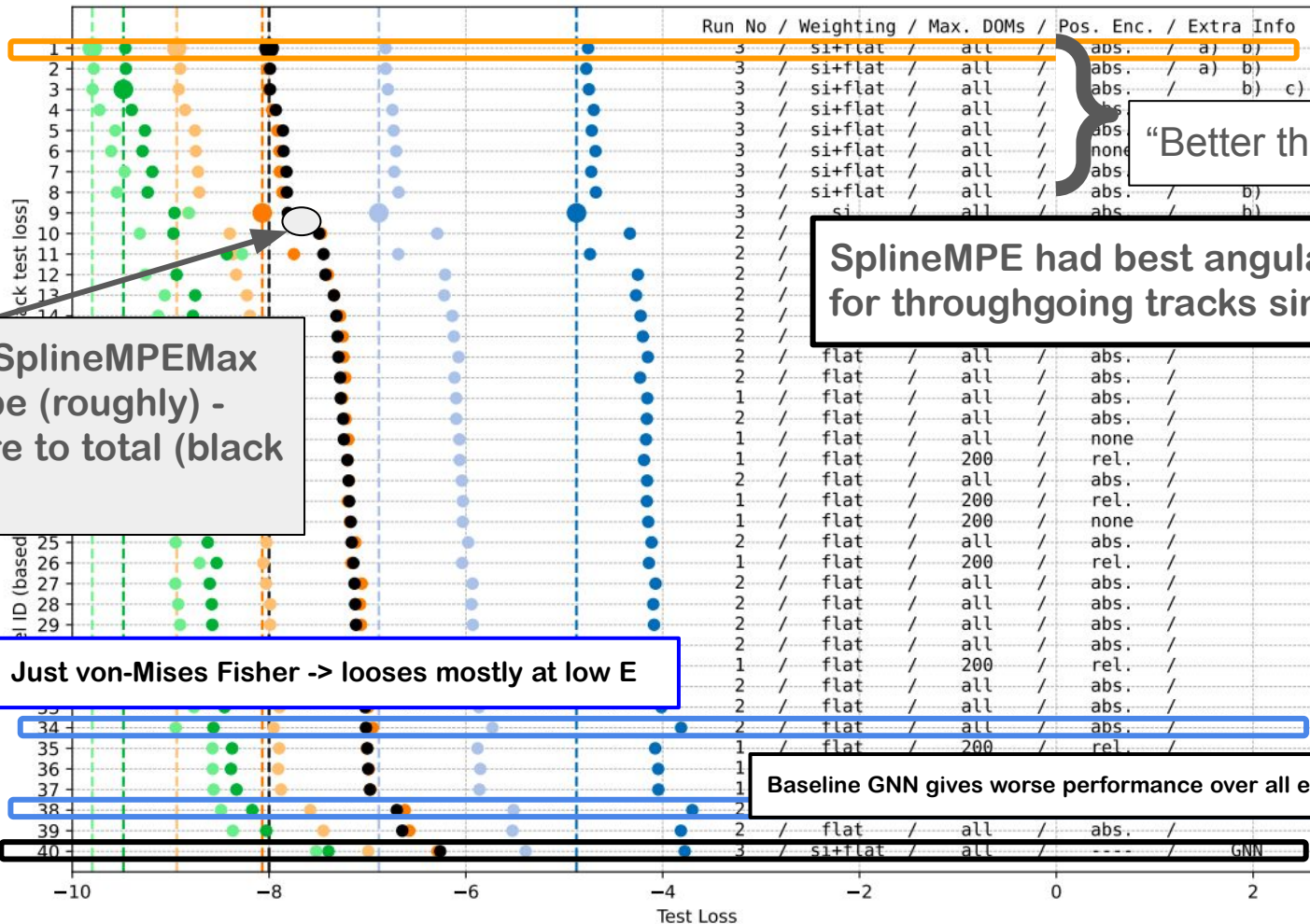
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Just von-Mises Fisher -> loses mostly at low E

Throughgoing tracks



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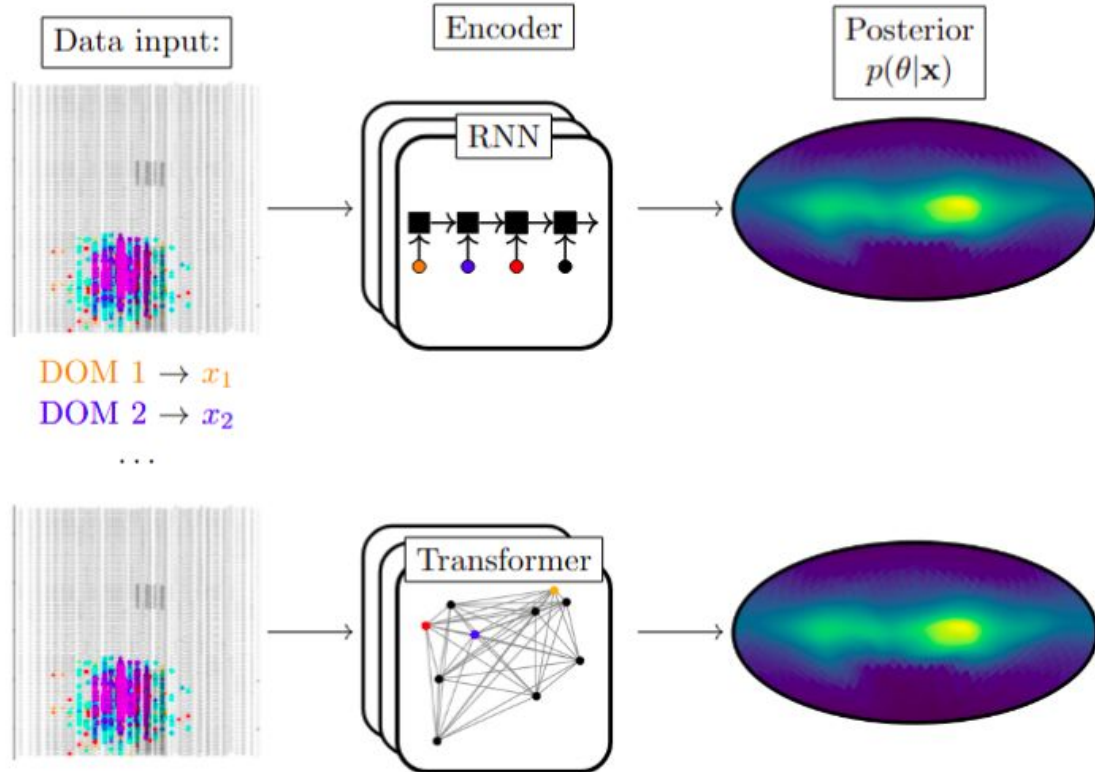
Baseline GNN gives worse performance over all energies

Bayes theorem with I.I.D. data

permute

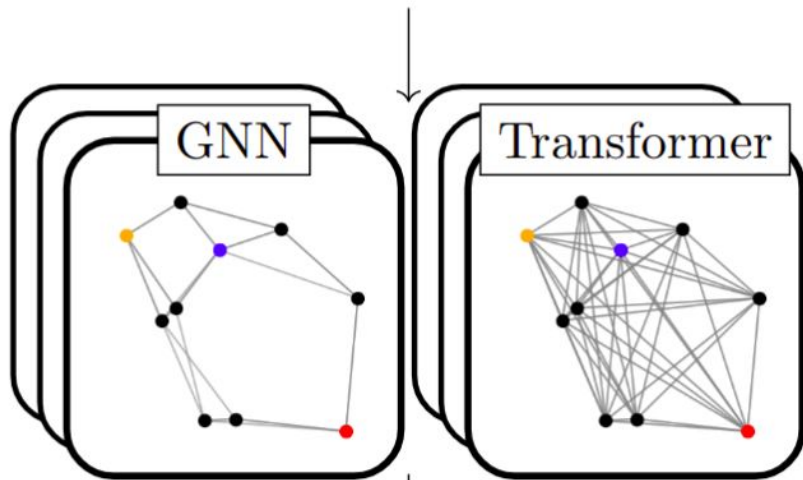
$$p(\theta|\mathbf{x}) = \frac{p(x_1;\theta) \cdot p(x_2;\theta) \cdot p(x_3;\theta) \cdot \dots \cdot p(\theta)}{p(\mathbf{x})} = \frac{p(x_1;\theta) \cdot p(x_3;\theta) \cdot p(x_2;\theta) \cdot \dots \cdot p(\theta)}{p(\mathbf{x})}$$

“Inductive bias”



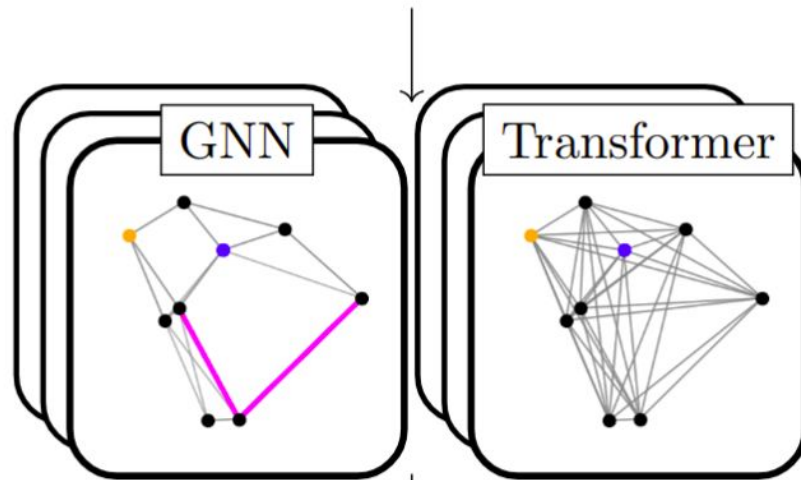
“Inductive bias”

Data factors: $x_1, x_2, x_3, x_4, \dots$

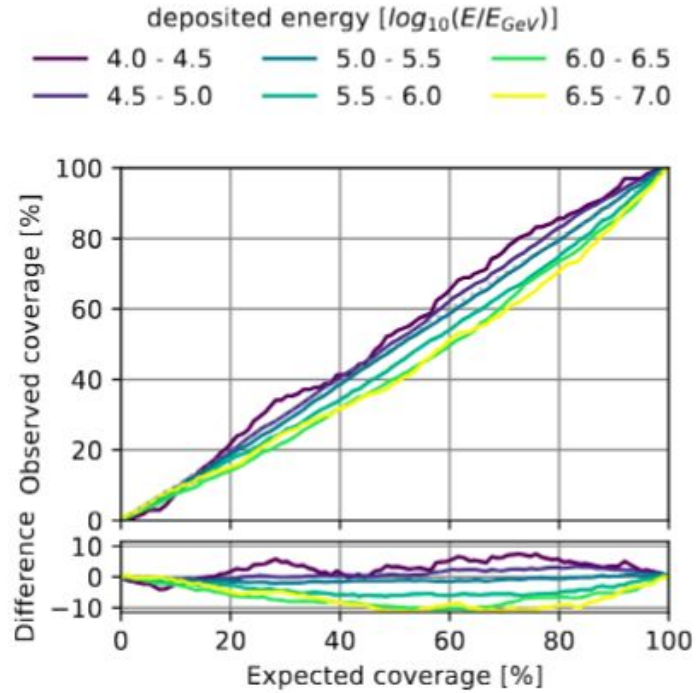


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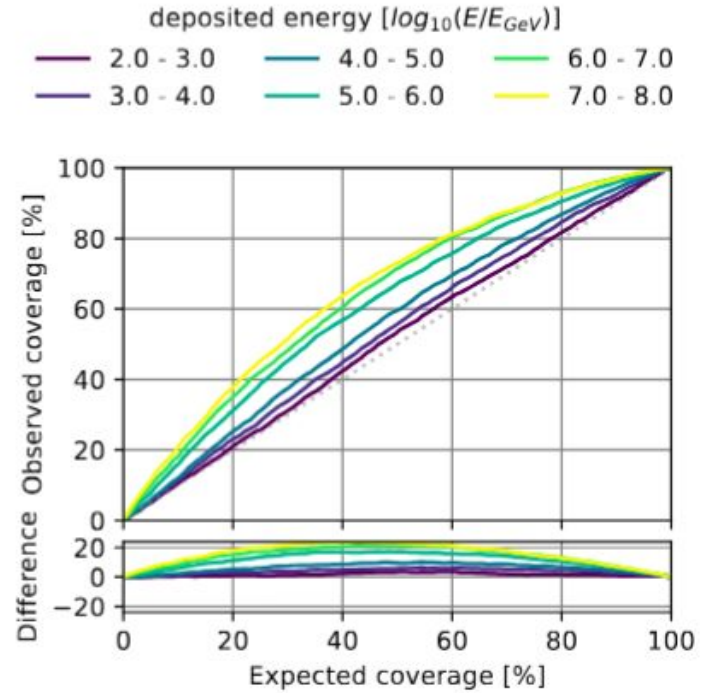
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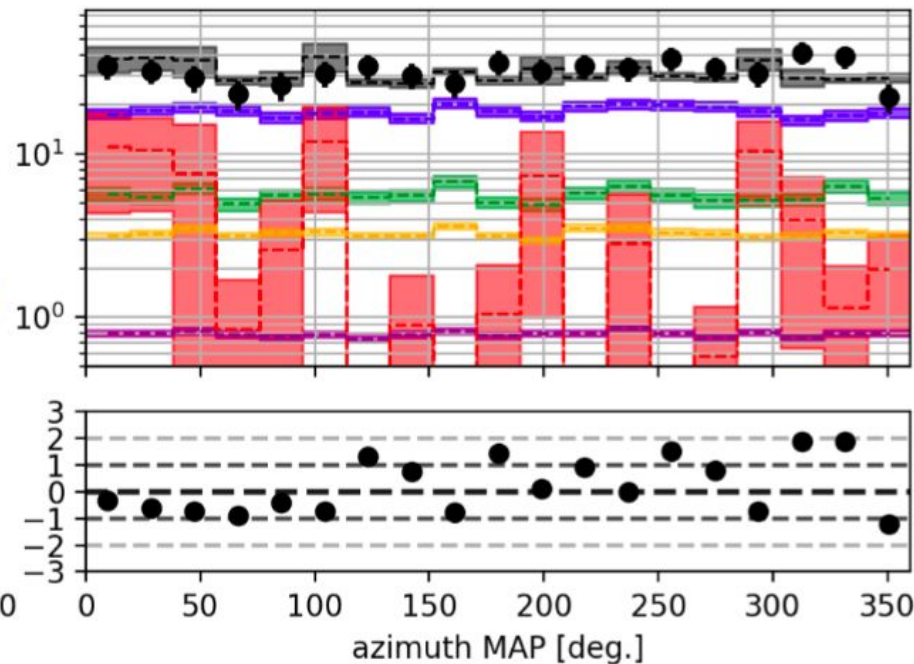
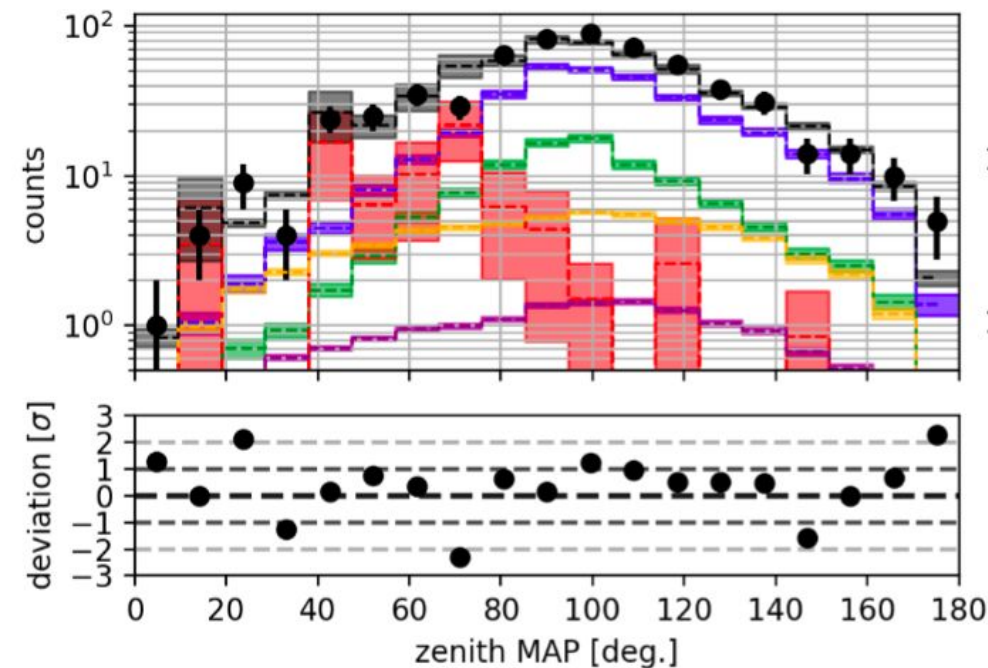
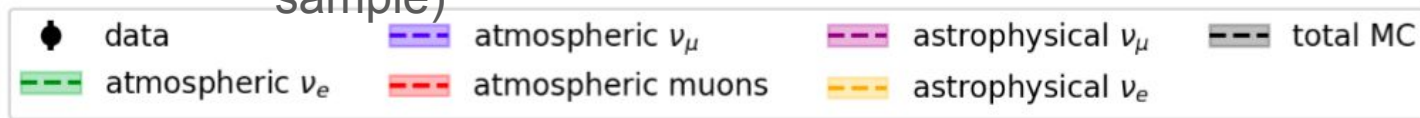
(a) Neutrino-induced showers



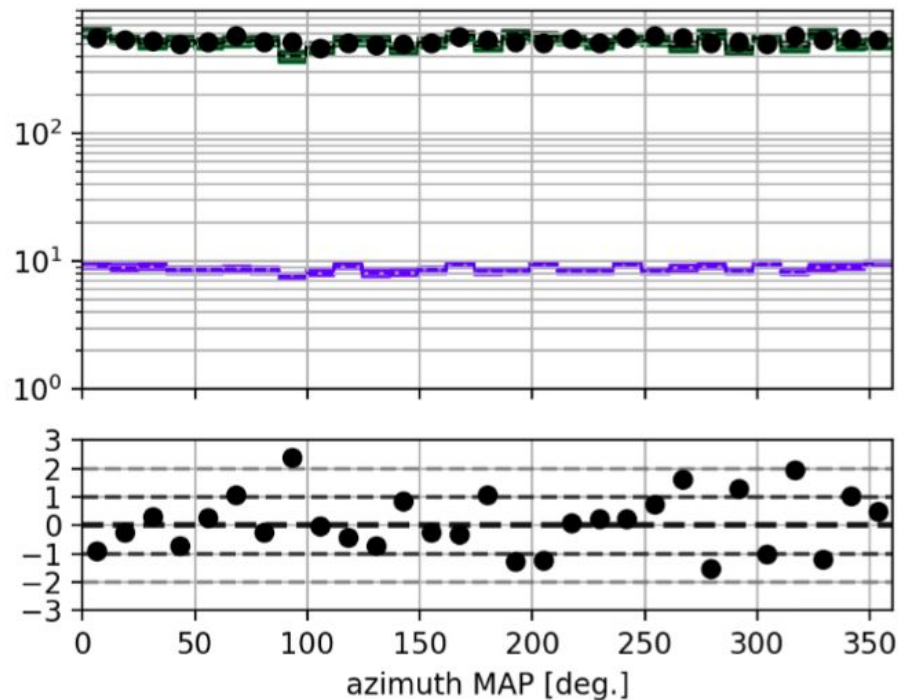
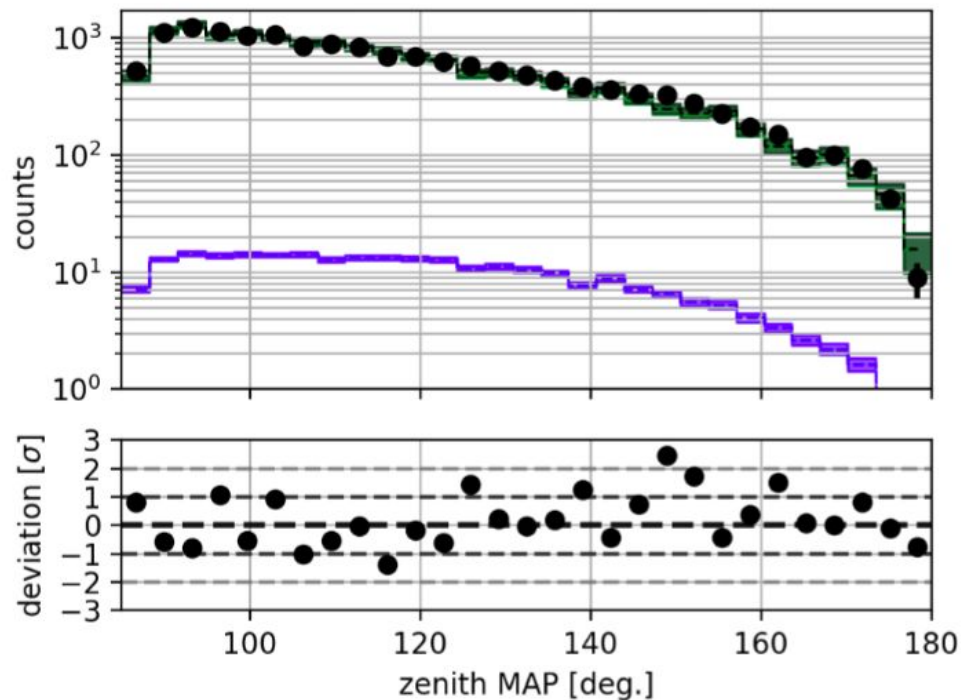
(b) Throughgoing tracks (at least 300m distance within detector)

Figure 9: Coverage split up in different deposited energy bands.

Shower zenith / azimuth (Zoe's shower sample)

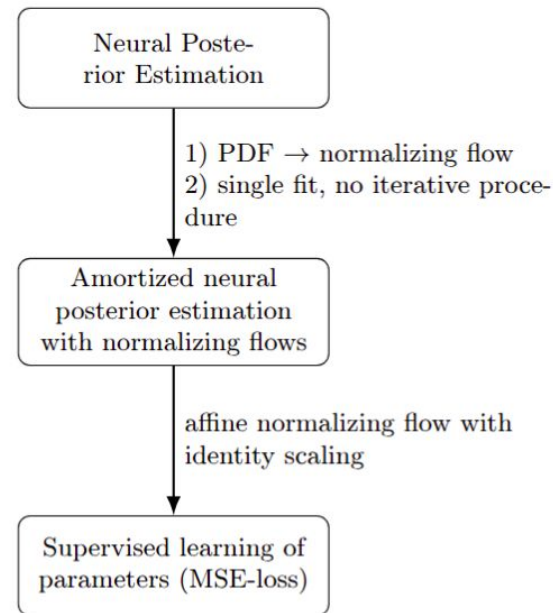


Track zenith / azimuth (northern track sample)

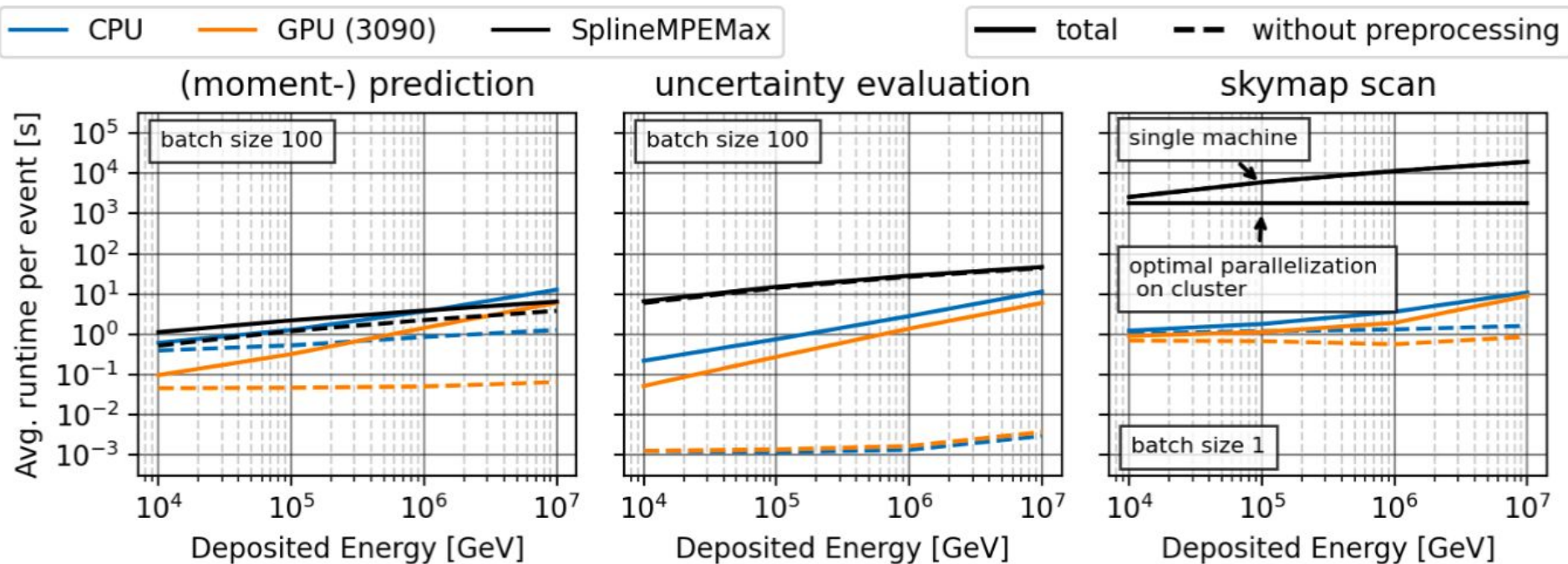


Loss function - standard cross entropy of the posterior

$$\begin{aligned}\psi^\star &= \arg \min_{\psi} D_{\text{KL}}\left(\tilde{p}(x, \theta) \parallel q(x) q_{\psi}(\theta|x)\right) \\&= \arg \min_{\psi} \mathbb{E}_{\tilde{p}(x, \theta)} \left[\log \frac{\tilde{p}(x, \theta)}{q(x) q_{\psi}(\theta|x)} \right] \\&= \arg \min_{\psi} \mathbb{E}_{\tilde{p}(x, \theta)} \left[D_{\text{KL}}\left(\tilde{p}(\theta|x) \parallel q_{\psi}(\theta|x)\right) \right] + \text{const} \\&= \arg \min_{\psi} \mathbb{E}_{\tilde{p}(x, \theta)} [-\log q_{\psi}(\theta|x)] \quad (\text{drop terms constant in } \psi) \\&\approx \arg \min_{\psi} \frac{1}{N} \sum^N -\log q_{\psi}(\theta_i|x_i) \equiv \arg \min_{\psi} \mathcal{L}_{\text{NPE}}(\psi).\end{aligned}$$



Timing (Tracks)



Timing (Showers)

