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# Schwarzschild Radius and Black Hole Thermodynamics with $\alpha'$ Corrections from Simulations of SUSY Matrix Quantum Mechanics

Jun Nishimura (KEK)

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“Lattice Supersymmetry and Beyond”  
at The Niels Bohr International Academy, Nov.24, '08

Ref.) Anagnostopoulos-Hanada- J.N.-Takeuchi, PRL 100 ('08) 021601  
Hanada-Miwa-J.N.-Takeuchi, arXiv:0811.2081[hep-th]  
Hanada-Hyakutake-J.N.-Takeuchi,arXiv:0811.3102[hep-th]

- based on collaborations with

Konstantinos Anagnostopoulos  
(National Technical University, Athens, Greece)

Masanori Hanada (Weizmann Inst., Israel)

Yoshifumi Hyakutake (Osaka Univ.)

Akitsugu Miwa  
(U. of Tokyo, Komaba  
→ Harish-Chandra Research Inst., India)

Shingo Takeuchi (KEK → APCTP, Korea)

# 0. Introduction

# Supersymmetric large- $N$ gauge theories

suitable formulation for describing *superstrings* non-perturbatively

- **non-pert. formulation of superstring/M theories**

e.g.) Matrix Theory (Banks-Fischler-Shenker-Susskind '97)

IIB matrix model (Ishibashi-Kawai-Kitazawa-Tsuchiya '97)

➤ **dynamical origin of space-time dimensionality, gauge group, matters, etc.**

- **gauge/string duality**

e.g.) AdS/CFT Maldacena ('97)

➤ **quantum/stringy description of black holes in terms of gauge theories**

**Difficulty : Strongly coupled dynamics should be investigated !**

# Monte Carlo simulation

can be a powerful approach as in lattice QCD, but...

SUSY : broken on the lattice

$$\{Q, \bar{Q}\} \propto P_\mu \rightarrow \begin{array}{l} \text{translational symmetries} \\ \text{broken to discrete ones} \end{array}$$

- Lattice formulation preserving **a part of supersymmetry** by using various ideas such as twisting, orbifolding, ...

(Cohen-Kaplan-Katz-Unsal, Catterall, Sugino,, Kanamori-Suzuki,  
Arianos-D'Adda-Feo-Kawamoto-Saito, Nagata, Damgaard-Matsuura,...)

- Non-lattice approach respecting SUSY maximally  
systems with **16 supercharges** can be studied !

1d gauge theory (**SUSY matrix QM**)                    (This talk)

non-perturbative gauge fixing + Fourier mode cutoff

**$\mathcal{N}=4$  SYM** on  $R \times S^2$  and  $R \times S^3$   
**fuzzy sphere**

# Supersymmetric matrix quantum mechanics

1d U(N) gauge theory with 16 supercharges

$$S = \frac{N}{\lambda} \int_0^\beta dt \operatorname{tr} \left\{ \frac{1}{2} (DX_i(t))^2 - \frac{1}{4} [X_i(t), X_j(t)]^2 + (\text{fermionic part}) \right\}$$

non-perturbative formulation of M theory

BFSS conjecture

Banks-Fischler-Shenker-Susskind ('97)

low energy effective theory of  $N$  D0 branes

This talk

gauge/gravity correspondence (non-conformal ver.)

Itzhaki-Maldacena-Sonnenschein-Yankielowicz ('98)

SUSY mass deformation



plane-wave matrix model

Berenstein-Maldacena-Nastase ('02)

Expanding around a (multi-)fuzzy-sphere background

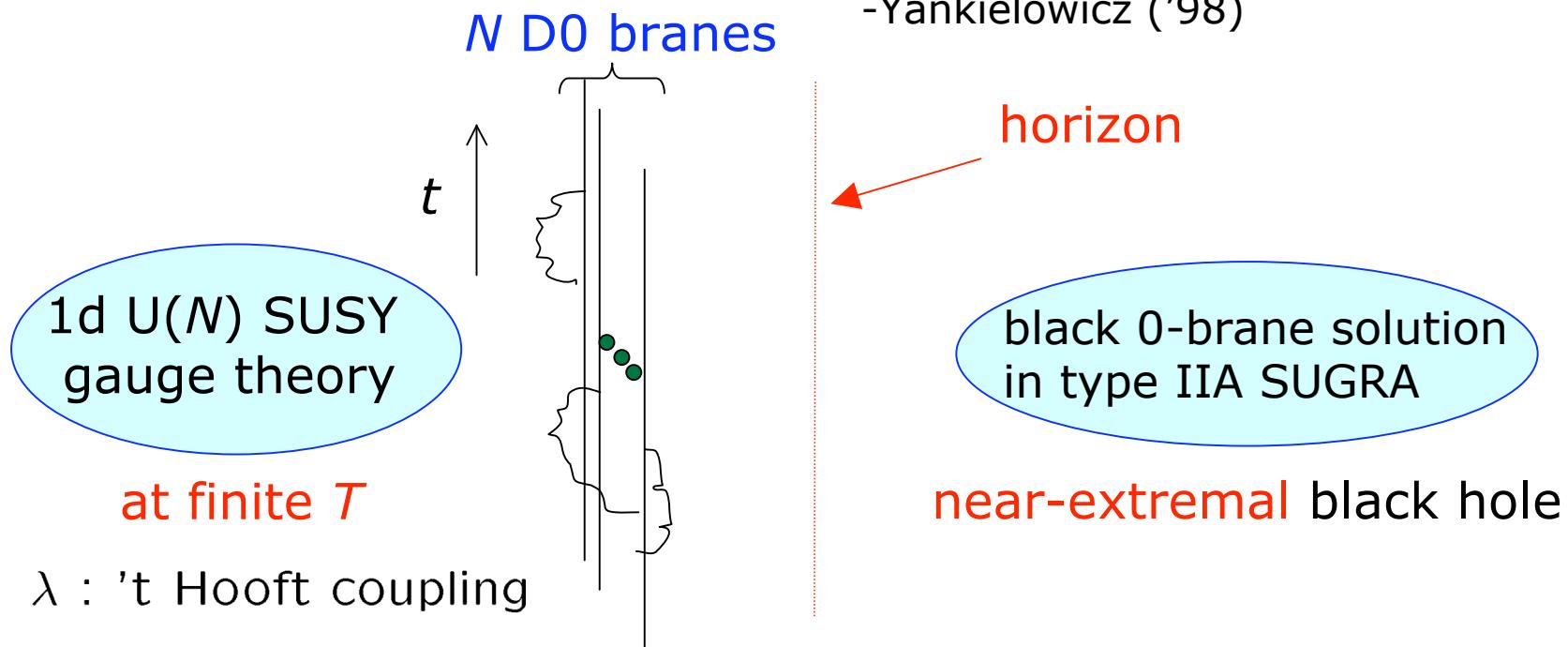


$\mathcal{N} = 4$  SYM on  $R \times S^2$  and  $R \times S^3$

Ishii-Ishiki-Shimasaki-Tsuchiya ('08)

# Gauge-gravity duality for D0-brane system

type IIA superstring



In the decoupling limit, the D0 brane system describes the black hole **microscopically**.

**large  $N$**  and **large  $\lambda$**   $\rightarrow$  SUGRA description : valid

# Simulating superstrings inside a black hole

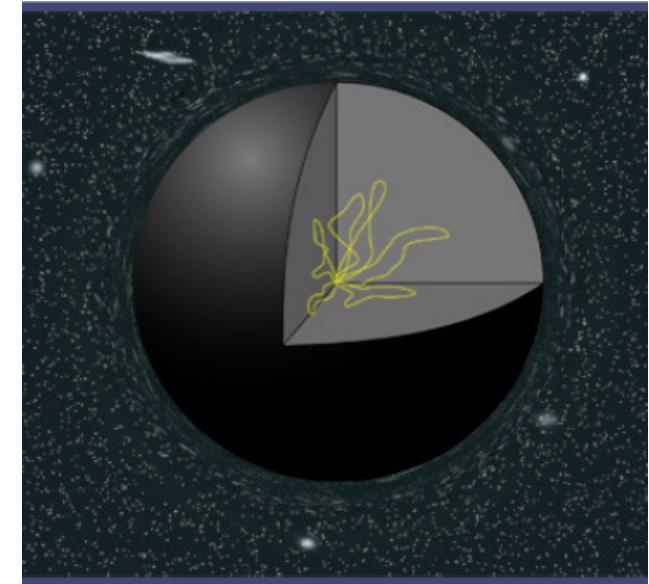
## ➡ black hole thermodynamics

Anagnostopoulos-Hanada-J.N.-Takeuchi ('08)

$$\frac{1}{N^2} \left( \frac{E}{\lambda^{1/3}} \right) = \underbrace{\frac{9}{14} \left\{ 4^{13} 15^2 \left( \frac{\pi}{7} \right)^{14} \right\}^{1/5}}_{7.41} \left( \frac{T}{\lambda^{1/3}} \right)^{14/5}$$

including  $\alpha'$  corrections

Hanada-Hyakutake-J.N.-Takeuchi, arXiv:0811.3102



## ➡ Schwarzschild radius from Wilson loop

Hanada-Miwa-J.N.-Takeuchi, arXiv:0811.2081[hep-th]

$$W \equiv \text{tr } \mathcal{P} \exp \left[ i \int_0^\beta dt \{ A(t) + i X_9(t) \} \right] \sim \exp \left( \frac{\beta R_{\text{Sch}}}{2\pi\alpha'} \right)$$

$$\ln W = \frac{\beta R_{\text{Sch}}}{2\pi\alpha'} = \underbrace{\frac{1}{2\pi} \left\{ \frac{16\sqrt{15}\pi^{7/2}}{7} \right\}^{2/5}}_{1.89} \left( \frac{T}{\lambda^{1/3}} \right)^{-3/5}$$

# Plan of the talk

0. Introduction
1. Simulating SUSY matrix QM with 16 supercharges
2. Dual gravity description and black hole thermodynamics
3. Higher derivative corrections to black hole thermodynamics from SUSY QM
4. Schwarzschild radius from Wilson loop
5. Summary

# 1. Simulating SUSY QM with 16 supercharges

# SUSY matrix QM with 16 supercharges

$$S_b = \frac{N}{\lambda} \int_0^\beta dt \operatorname{tr} \left\{ \frac{1}{2} (DX_i(t))^2 - \frac{1}{4} [X_i(t), X_j(t)]^2 \right\}$$

$$S_f = \frac{N}{\lambda} \int_0^\beta dt \operatorname{tr} \left\{ \frac{1}{2} \psi_\alpha D\psi_\alpha - \frac{1}{2} \psi_\alpha (\gamma_i)_{\alpha\beta} [X_i, \psi_\beta] \right\}$$

1d gauge theory

$$D = \partial_t - i [A(t), \cdot]$$

$$\begin{cases} X_j(t) & (j = 1, \dots, 9) & \text{p.b.c.} \\ \psi_\alpha(t) & (\alpha = 1, \dots, 16) & \text{anti p.b.c.} \end{cases}$$

$T = \beta^{-1}$  temperature

$\lambda = g^2 N$  't Hooft coupling

$$\lambda_{\text{eff}} = \frac{\lambda}{T^3}$$

$\lambda = 1$  (without loss of generality)

$$\begin{cases} \text{low T} & \rightarrow \text{strongly coupled} & \text{dual gravity description} \\ \text{high T} & \rightarrow \text{non-zero modes : weakly coupled (high T exp.)} \\ & & \text{(zero modes : integrated non-perturbatively)} \end{cases}$$

Kawahara-J.N.-Takeuchi,  
JHEP 0712 (2007) 103, arXiv:0710.2188[hep-th]

# Fourier-mode simulation respecting SUSY maximally

Hanada-J.N.-Takeuchi, PRL 99 (07) 161602 [arXiv:0706.1647]

$$X_i(t) = \sum_{n=-\Lambda}^{\Lambda} \tilde{X}_{i,n} e^{i\omega n t} \quad \omega = \frac{2\pi}{\beta}$$

Note: Gauge symmetry can be fixed non-perturbatively in 1d.

- static diagonal gauge :

$$A(t) = \frac{1}{\beta} \text{diag}(\alpha_1, \dots, \alpha_N)$$

$$S_{\text{FP}} = - \sum_{a < b} 2 \ln \left| \sin \frac{\alpha_a - \alpha_b}{2} \right|$$

- residual gauge symmetry :  $g(t) = \text{diag}(e^{i\omega\nu_1 t}, \dots, e^{i\omega\nu_N t})$

$$\begin{cases} \tilde{X}_{i,n}^{ab} \mapsto \tilde{X}_{i,n-\nu_a+\nu_b}^{ab} \\ \alpha_a \mapsto \alpha_a + 2\pi\nu_a \end{cases} \quad \begin{aligned} X_i &\mapsto g X_i g^\dagger \\ A &\mapsto g A g^\dagger + i g \partial_t g^\dagger \end{aligned}$$

should be fixed by imposing  $-\pi < \alpha_a \leq \pi$

c.f.) lattice approach : Catterall-Wiseman, PRD78 (08) 041502

## 2. Dual gravity description and black hole thermodynamics

Anagnostopoulos-Hanada- J.N.-Takeuchi, PRL 100  
('08) 021601 [arXiv:0707.4454]

# Dual gravity description

After taking the decoupling limit :  $\alpha' \rightarrow 0$

$$U \equiv \frac{r}{\alpha'} , \quad \lambda \equiv g_s N \alpha'^{-3/2} \quad (\text{fixed}) \quad f(U) \equiv \frac{U^{7/2}}{\sqrt{d_0 \lambda}} \left\{ 1 - \left( \frac{U_0}{U} \right)^7 \right\}$$

$$ds^2 = \alpha' \left\{ f(U) dt^2 + \frac{1}{f(U)} dU^2 + \sqrt{d_0 \lambda} U^{-3/2} d\Omega_{(8)}^2 \right\}$$

range of validity:  $N^{-10/21} \ll \frac{T}{\lambda^{1/3}} \ll 1$

## Black hole thermodynamics

$$\begin{cases} \text{Hawking temperature :} & T = \frac{7}{16\sqrt{15}\pi^{7/2}} \left( \frac{U_0}{\lambda^{1/3}} \right)^{5/2} \\ \text{Bekenstein-Hawking entropy :} & S = \frac{1}{28\sqrt{15}\pi^{7/2}} N^2 \left( \frac{U_0}{\lambda^{1/3}} \right)^{9/2} \end{cases}$$

$$\rightarrow \frac{1}{N^2 \lambda^{1/3}} \frac{E}{= \frac{9}{14} \underbrace{\left\{ 4^{13} 15^2 \left( \frac{\pi}{7} \right)^{14} \right\}^{1/5}}_{7.41} \left( \frac{T}{\lambda^{1/3}} \right)^{14/5}}$$

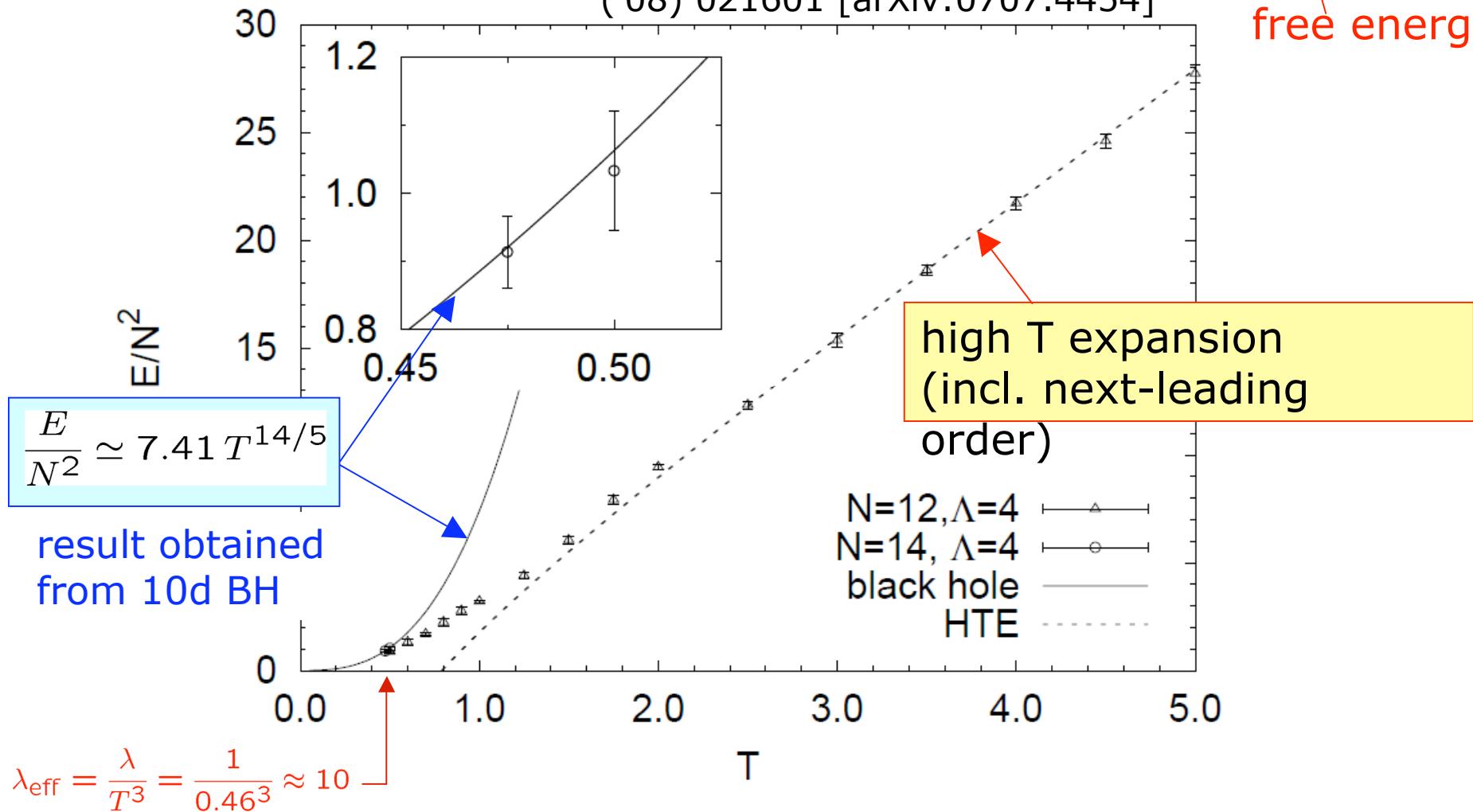
Klebanov-Tseytlin ('96)

# Result: Internal energy

$$E = \frac{\partial}{\partial \beta} (\beta \mathcal{F})$$

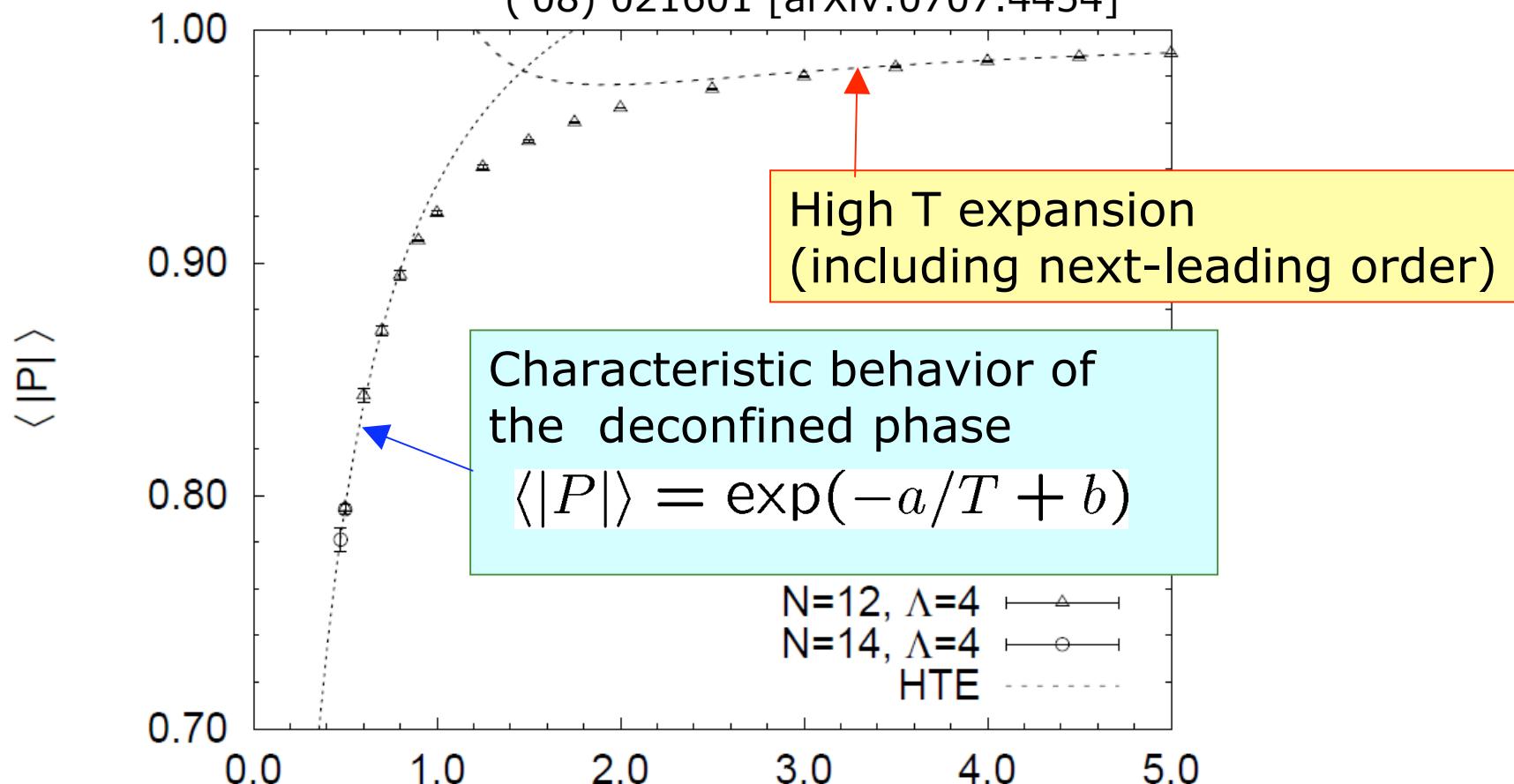
Anagnostopoulos-Hanada- J.N.-Takeuchi, PRL 100 ('08) 021601 [arXiv:0707.4454]

free energy



# Result: Polyakov line

Anagnostopoulos-Hanada- J.N.-Takeuchi, PRL 100 ('08) 021601 [arXiv:0707.4454]



no phase transition unlike in bosonic case  $T$

→ consistent with analyses on the gravity size (Barbon et al., Aharony et al.)

### 3. Higher derivative corrections to black hole thermodynamics from SUSY QM

Hanada-Hyakutake-J.N.-Takeuchi, arXiv:0811.3102[hep-th]

# $\alpha'$ corrections to type IIA SUGRA action

low energy effective action of type IIA superstring theory

← tree-level scattering amplitudes of the massless modes

leading term : type IIA SUGRA action

$$\mathcal{S}_{(0)} = \frac{1}{16\pi G_N} \int d^{10}x \sqrt{-g} \left\{ e^{-2\phi} (R + 4\partial_\mu\phi\partial^\mu\phi) - \frac{1}{4}G_{\mu\nu}G^{\mu\nu} \right\}$$
$$G_N \sim \alpha'^4 g_s^2$$

explicit calculations of 2-pt and 3-pt amplitudes

→  $\mathcal{S}_{(1)} = \mathcal{S}_{(2)} = 0$

4-pt amplitudes

→  $\mathcal{S}_{(3)} = \frac{\alpha'^3}{16\pi G_N} \int d^{10}x \sqrt{-g} \left\{ e^{-2\phi} \mathcal{R}^4 + \dots \right\}$

Complete form is yet to be determined,  
but we can still make a dimensional analysis.

# Black hole thermodynamics with $\alpha'$ corrections

curvature radius of the dual geometry

$$\rho^2 \sim \left( \frac{\lambda^{1/3}}{U_0} \right)^{3/2} \alpha'$$

$\alpha'$  corrections

$$\rightarrow \frac{\alpha'}{\rho^2} \sim \left( \frac{U_0}{\lambda^{1/3}} \right)^{3/2} \sim \left( \frac{T}{\lambda^{1/3}} \right)^{3/5} \quad \frac{T}{\lambda^{1/3}} \sim \left( \frac{U_0}{\lambda^{1/3}} \right)^{5/2}$$

corrections at  $\alpha'^3$  order gives

$$\frac{1}{N^2} \frac{E}{\lambda^{1/3}} = 7.41 \left( \frac{T}{\lambda^{1/3}} \right)^{14/5} \left\{ 1 + \boxed{a \left( \frac{T}{\lambda^{1/3}} \right)^{9/5}} \right\}$$

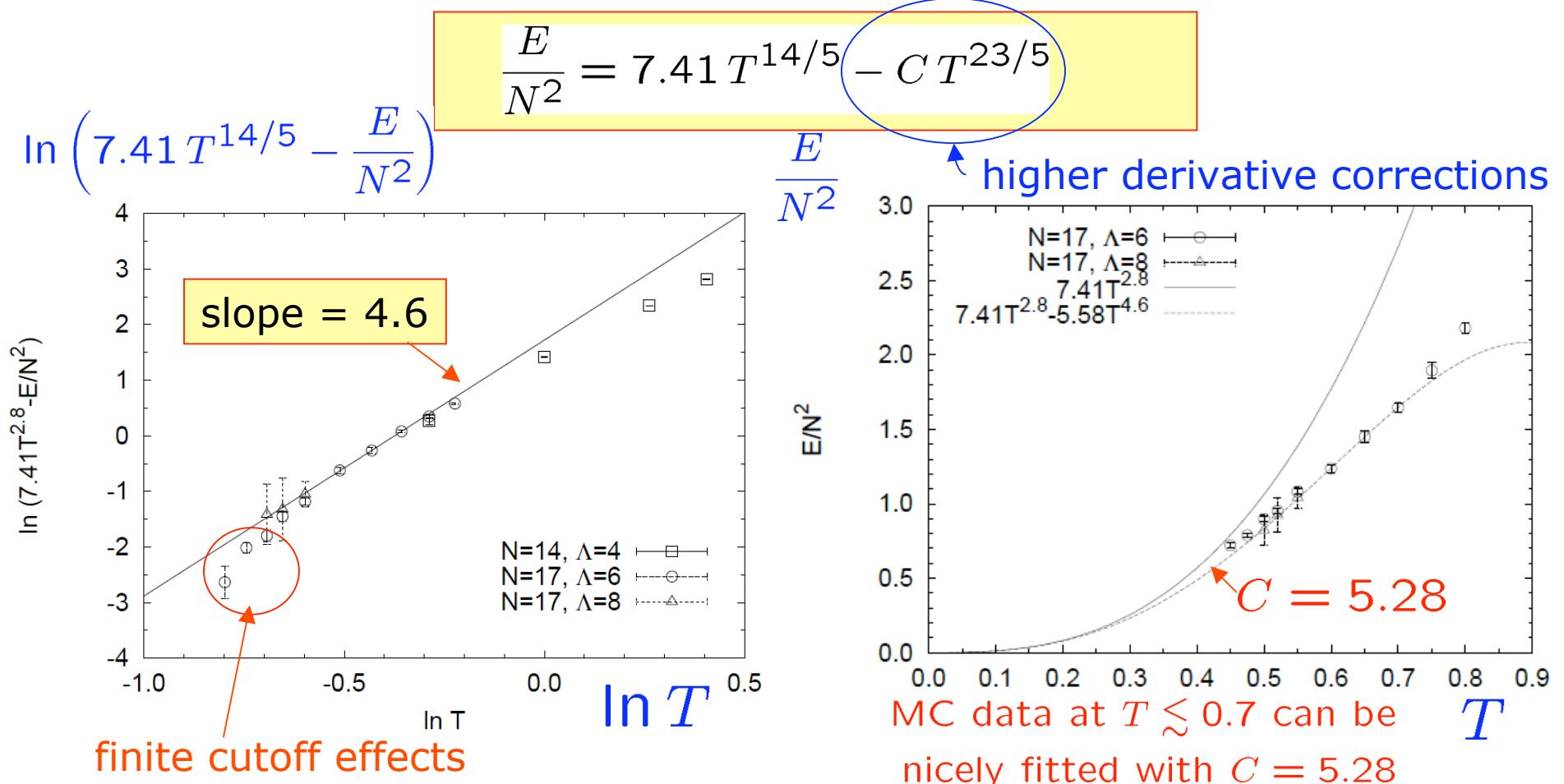
More careful treatment leads to the same conclusion.  
(Hanada-Hyakutake-J.N.-Takeuchi,arXiv:0811.3102)

Setting  $\lambda = 1$ ,

$$\frac{E}{N^2} = 7.41 T^{14/5} - C T^{23/5}$$

# Higher derivative corrections to black hole thermodynamics from SUSY QM

Hanada-Hyakutake-J.N.-Takeuchi, arXiv:0811.3102[hep-th]

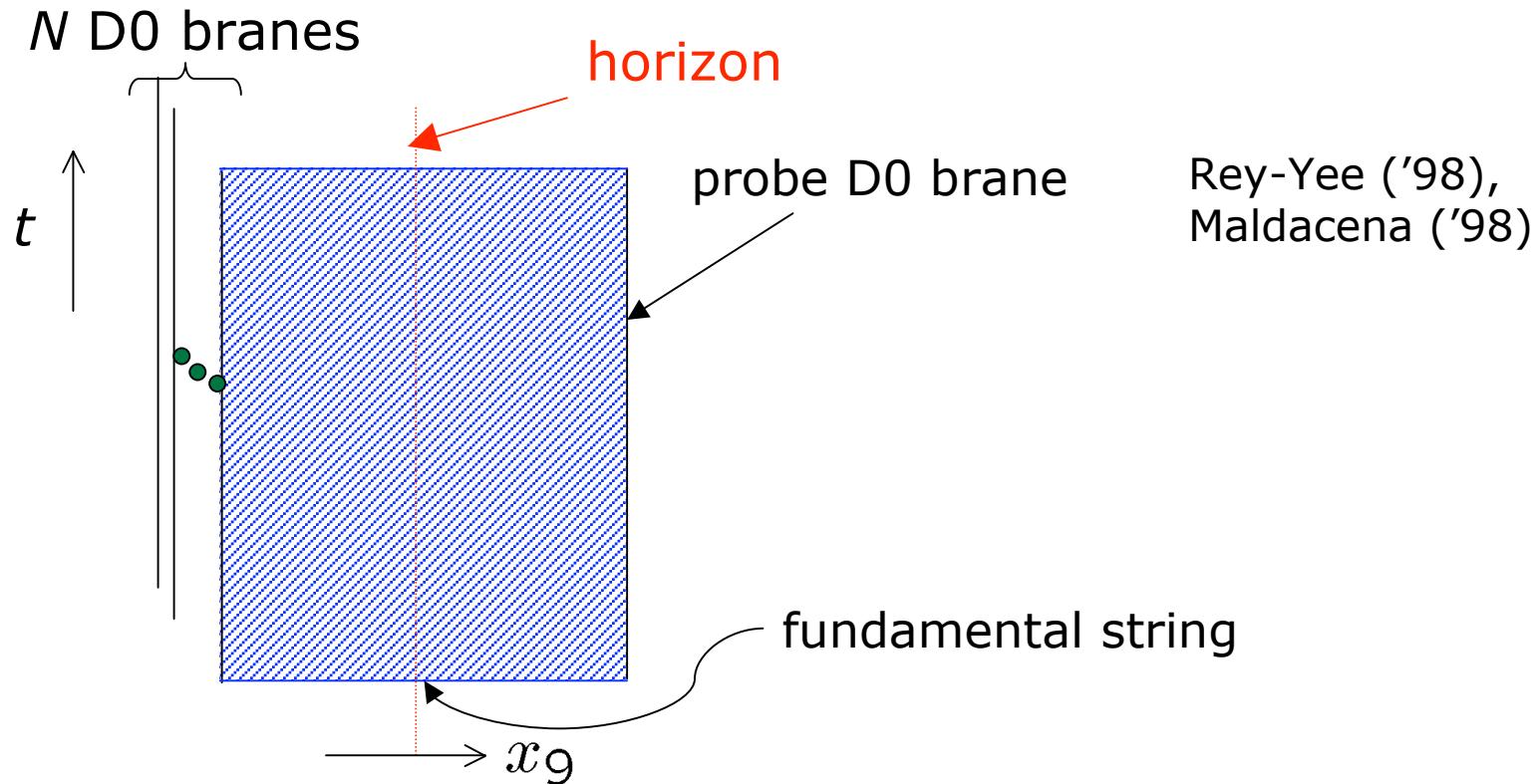


### 3. Schwarzschild radius from Wilson loop

Hanada-Miwa-J.N.-Takeuchi , arXiv:0811.2081[hep-th]

# Calculation of Wilson loop

Hanada-Miwa-J.N.-Takeuchi, arXiv:0811.2081[hep-th]

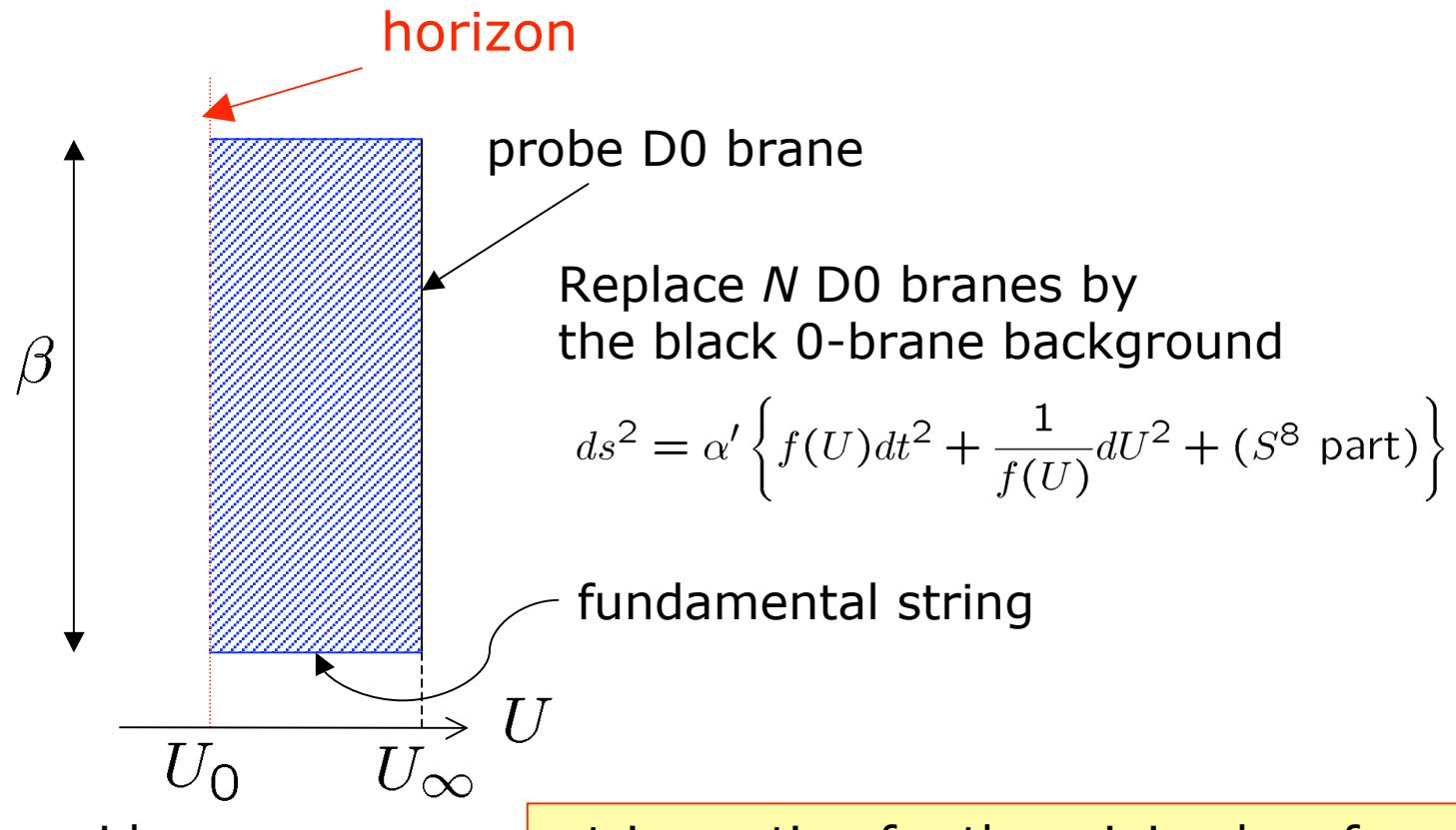


gauge theory side :

propagation of a test particle  
coupled to  $A(t)$  and  $X_9(t)$

$$W = \text{tr} \mathcal{P} \exp \left[ i \int_0^\beta dt \{ A(t) + iX_9(t) \} \right]$$

# Calculation of Wilson loop (cont'd)



gravity theory side :

propagation of the string  
in the b.g. geometry

string action for the minimal surface :

$$S_{\text{string}} = \frac{1}{2\pi} \beta (U_\infty - U_0)$$

## Calculation of Wilson loop (cont'd)

$$W e^{-M\beta} = e^{-S_{\text{string}}} \quad \text{at large } N \text{ and large } \lambda$$

perimeter-law suppression factor  
due to propagation of a particle with mass  $M$

$$S_{\text{string}} = \frac{1}{2\pi} \beta (U_\infty - U_0)$$

$$\log W - \boxed{\beta M} = \frac{\beta U_0}{2\pi} - \boxed{\frac{\beta U_\infty}{2\pi}}$$

natural to identify

more sophisticated justification  
a la Drukker-Gross-Ooguri ('99)

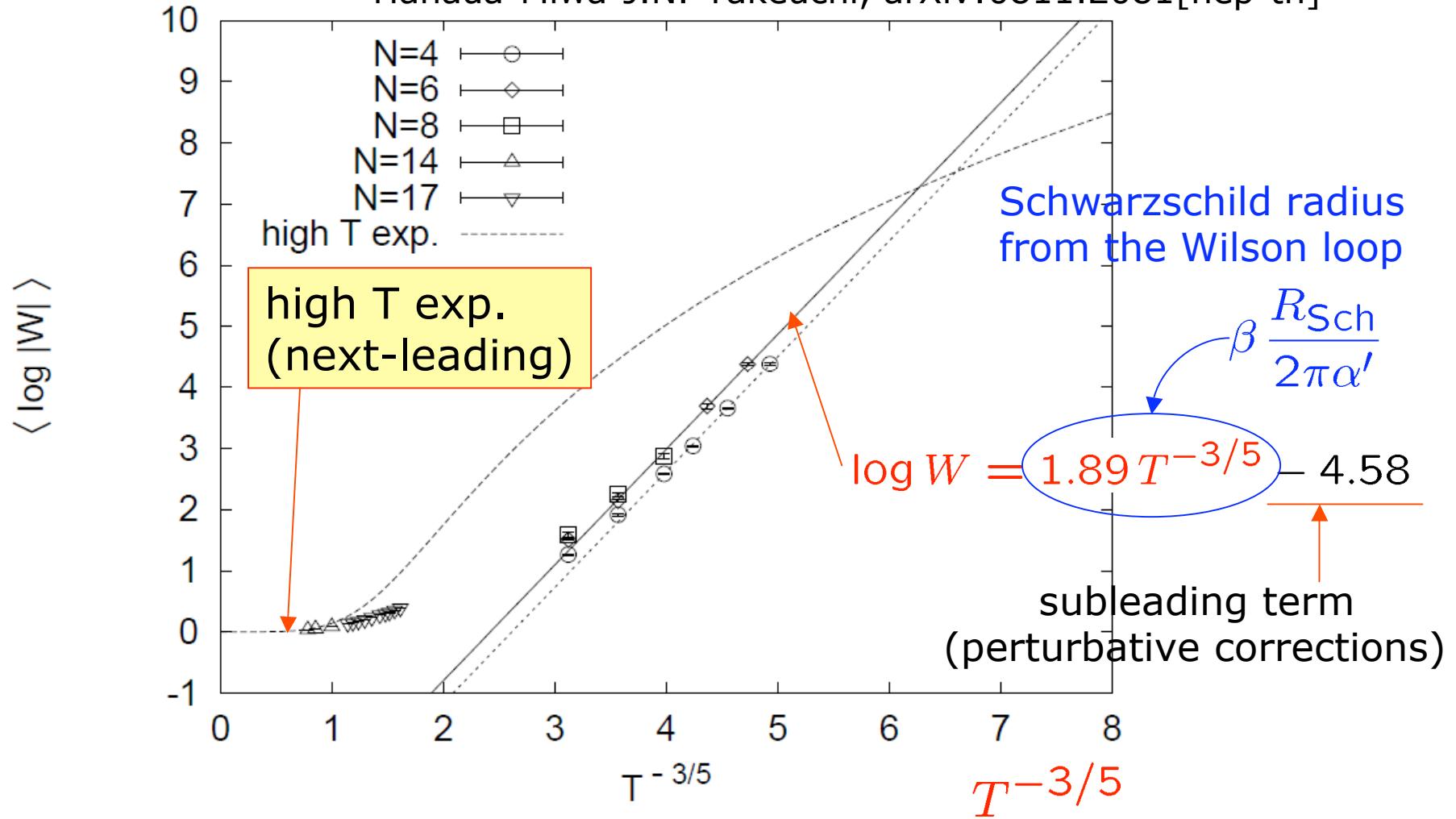


$$\log W = \frac{\beta U_0}{2\pi} = \frac{\beta R_{\text{Sch}}}{2\pi\alpha'} = \underbrace{\frac{1}{2\pi} \left\{ \frac{16\sqrt{15}\pi^{7/2}}{7} \right\}^{2/5}}_{1.89} \left( \frac{T}{\lambda^{1/3}} \right)^{-3/5}$$

# Results: Wilson loop

$$W = \text{tr } \mathcal{P} \exp \left[ i \int_0^\beta dt \{ A(t) + iX_9(t) \} \right]$$

Hanada-Miwa-J.N.-Takeuchi, arXiv:0811.2081[hep-th]



## 5. Summary

# Summary

- Monte Carlo studies of supersymmetric large N gauge theories  
→ powerful method for superstring theory
- simulating superstrings inside a black hole based on gauge/string duality
  - { Black hole thermodynamics (E v.s. T relation)  
Schwarzschild radius reproduced from Wilson loop
- a highly nontrivial check of the duality  
microscopic origin of the black hole thermodynamics  
including higher derivative corrections !

# Future prospects

- extension to lower SUSY case  
(Hanada-Matsuura-J.N., in progress)  
easier, and many things to explore
- higher dimensional case  
possible using mass deformation  
(Ishiki-Kim-J.N.-Tsuchiya, in progress)
- so far, planar limit  
important next step is to study non-planar limit  
Matrix theory, IIB matrix model
  - sign problem has to be treated carefully
    - toy model (Anagnostopoulos-Azuma-J.N., in prep.)
    - 6d IKKT model (Aoyama-Azuma-Hanada-J.N., in progress)
  - $\text{SO}(6) \rightarrow \text{SO}(3)$  from Gaussian expansion (Aoyama-J.N.-Okubo, in prep.)

## 6. Related on-going projects

SYM on  $R \times S^2$ ,  $R \times S^3$  from SUSY matrix QM

4d universe from 10d(11d) space-time ?

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# SYM on $R \times S^2$ , $R \times S^3$ from SUSY matrix QM

respect SUSY maximally

c.f.) lattice approach

1)  $U(k)$  SYM on  $R \times S^2$  in the planar limit



plane wave matrix model

(matrix QM with mass terms & CS terms)

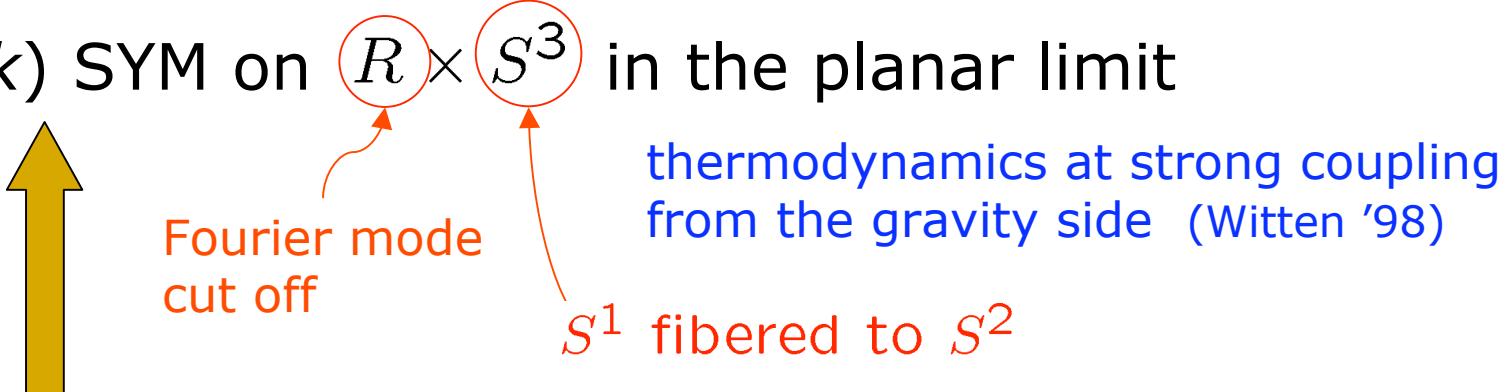
around  $k$  copies of the fuzzy sphere solution

$$X_a = L_a^{(n)} \otimes \mathbf{1}_k \quad (a = 1, 2, 3)$$

$k \rightarrow \infty$  limit removes fuzzyness

## SYM on $R \times S^2$ , $R \times S^3$ from SUSY matrix QM (cont'd)

2) U( $k$ ) SYM on  $R \times S^3$  in the planar limit



plane wave matrix model  
around  $k$  copies of the multi-fuzzy-sphere solution

$$X_a = L_a \otimes \mathbf{1}_k \quad (a = 1, 2, 3) \quad \text{Ishii-Ishiki-Shimasaki-Tsuchiya ('08)}$$

$$L_a = \begin{pmatrix} L_a^{(n)} & & & \\ & L_a^{(n+1)} & & \\ & & \ddots & \\ & & & L_a^{(n+s)} \end{pmatrix}$$

$k \rightarrow \infty$  limit removes fuzziness

# Agreement at weak coupling

Ishiki-Kim-J.N.-Tsuchiya, arXiv:0810.2884[hep-th]

$$S_{\text{PWMM}} = N \int_0^\beta dt \text{tr} \left\{ \frac{1}{2} (DX_i)^2 + \frac{1}{2} \mu^2 (X_a)^2 + \frac{1}{8} \mu^2 (X_m)^2 \right. \\ \left. + \frac{1}{2} i \mu \epsilon_{abc} X_a [X_b, X_c] - \frac{1}{4} [X_i, X_j]^2 \right. \\ \left. + \frac{1}{2} \psi^\dagger D\psi + \frac{3}{8} i \mu \psi^\dagger \gamma^{123} \psi - \frac{1}{2} \psi^\dagger \gamma^i [X_i, \psi] \right\}$$

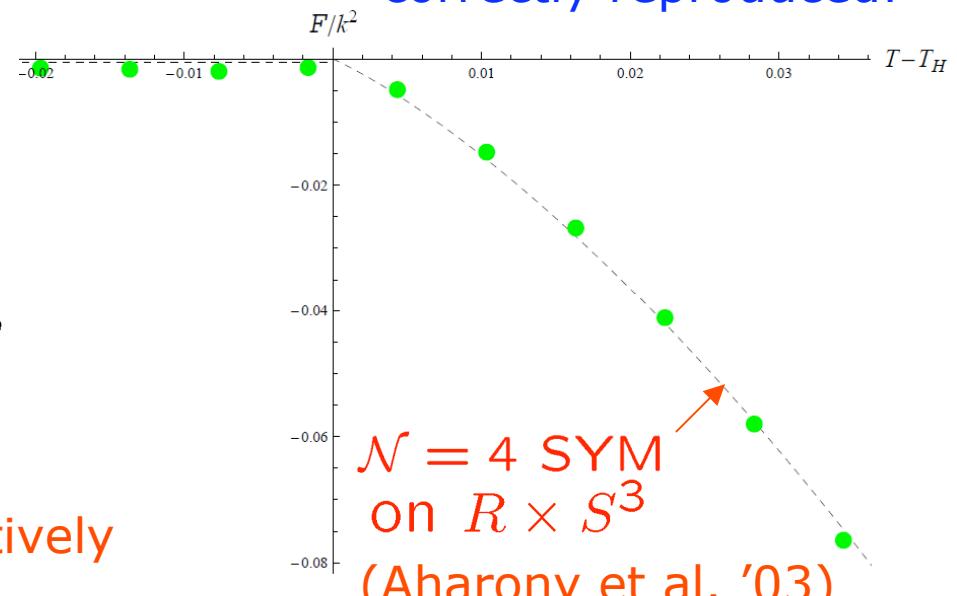
$$X_a = L_a \otimes \mathbf{1}_k \quad (a = 1, 2, 3)$$

$$L_a = \begin{pmatrix} L_a^{(n)} & & & \\ & L_a^{(n+1)} & & \\ & & \ddots & \\ & & & L_a^{(n+s)} \end{pmatrix}$$

For  $\mu \gg 1$  (weak coupling),

- { fluctuations of  $X_i$  and  $\psi$   
integrated out at **1-loop**
- gauge field  $A(t)$  : moduli  
integrated **non-perturbatively**  
by **MC sim.**

deconfinement transition  
correctly reproduced!



# Simulating Quantum Universe

Another interpretation of D0 brane quantum mechanics  
(different large- $N$  limit)

## Matrix Theory

Banks-Fischler-Shenker-Susskind ('97)

microscopic description of M Theory

c.f.) Matrix cosmology  
Freedman-Gibbons-Schnabl (hep-th/0411119)

How does our 4d space-time appear  
from 10d (11d) space-time ?

$$\text{e.g.) } \text{SO}(9) \xrightarrow{\text{SSB}} \text{SO}(3) ?$$

Gaussian expansion method (GEM)

Aoyama-J.N.-Okubo-Takeuchi, in progress

application to matrix QM  
Kabat-Lifschytz ('99)

BH thermodynamics  
Kabat-Lifschytz-Lowe('00)

# Gaussian Expansion Method

$$F = -\log \left( \int d\phi e^{-S} \right)$$

$$\tilde{S}(\epsilon, t) \equiv S_G(t) + \epsilon \{S - S_G(t)\}$$

$$\tilde{F}(\epsilon, t) \equiv -\log \left( \int d\phi e^{-\tilde{S}(\epsilon, t)} \right)$$

$$\tilde{F}(\epsilon, t) = \sum_{k=0}^{\infty} \epsilon^k f_k(t)$$

$$\tilde{F}_n(t) \equiv \sum_{k=0}^n f_k(t)$$

How to identify the plateau ?  
if  $S_G$  contains many param.

self-consistency eq.:

$$\frac{\partial}{\partial t} \tilde{F}_n(t) = 0$$

Search for concentration of solutions

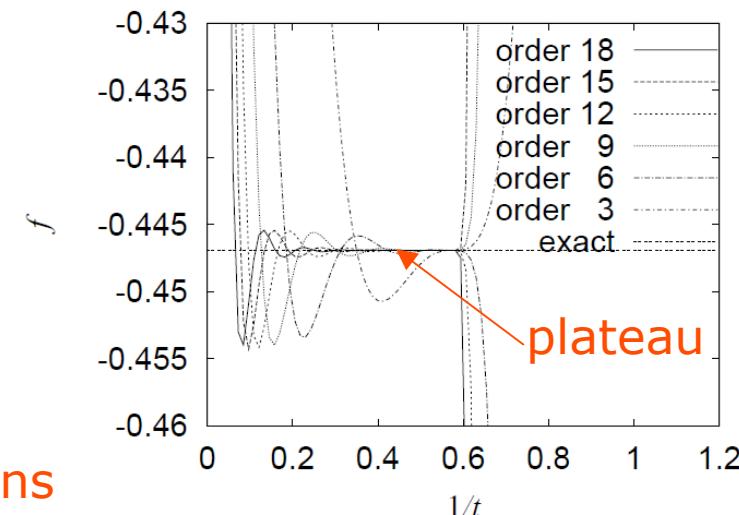
$$S = \frac{1}{4} N \text{tr} \phi^4$$

$$S_G(t) = \frac{1}{2} N t \text{tr} \phi^2$$

$$F = \tilde{F}(1, t)$$

indep. of  $t$

Truncate the series at  $k = n$   
and set  $\epsilon = 1$ .



# Dynamical generation of space-time in type IIB matrix model

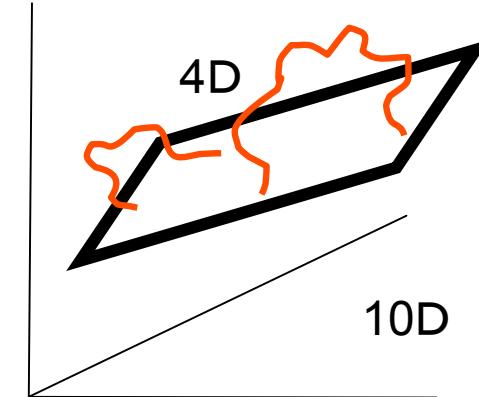
$$S = N \text{tr} \left\{ -\frac{1}{4} [X_\mu, X_\nu]^2 + \frac{1}{2} \psi_\alpha (\Gamma_\mu)_{\alpha\beta} [X_\mu, \psi_\beta] \right\}$$

Gaussian expansion method

J.N.-Sugino ('01),  
Kawai et al. ('01),...



Eigenvalue distribution of  $X_\mu$



Analogous studies  
in the D=6 model (less SUSY)  
Aoyama-J.N.-Okubo, in prep.

more systematic studies of SSB patterns

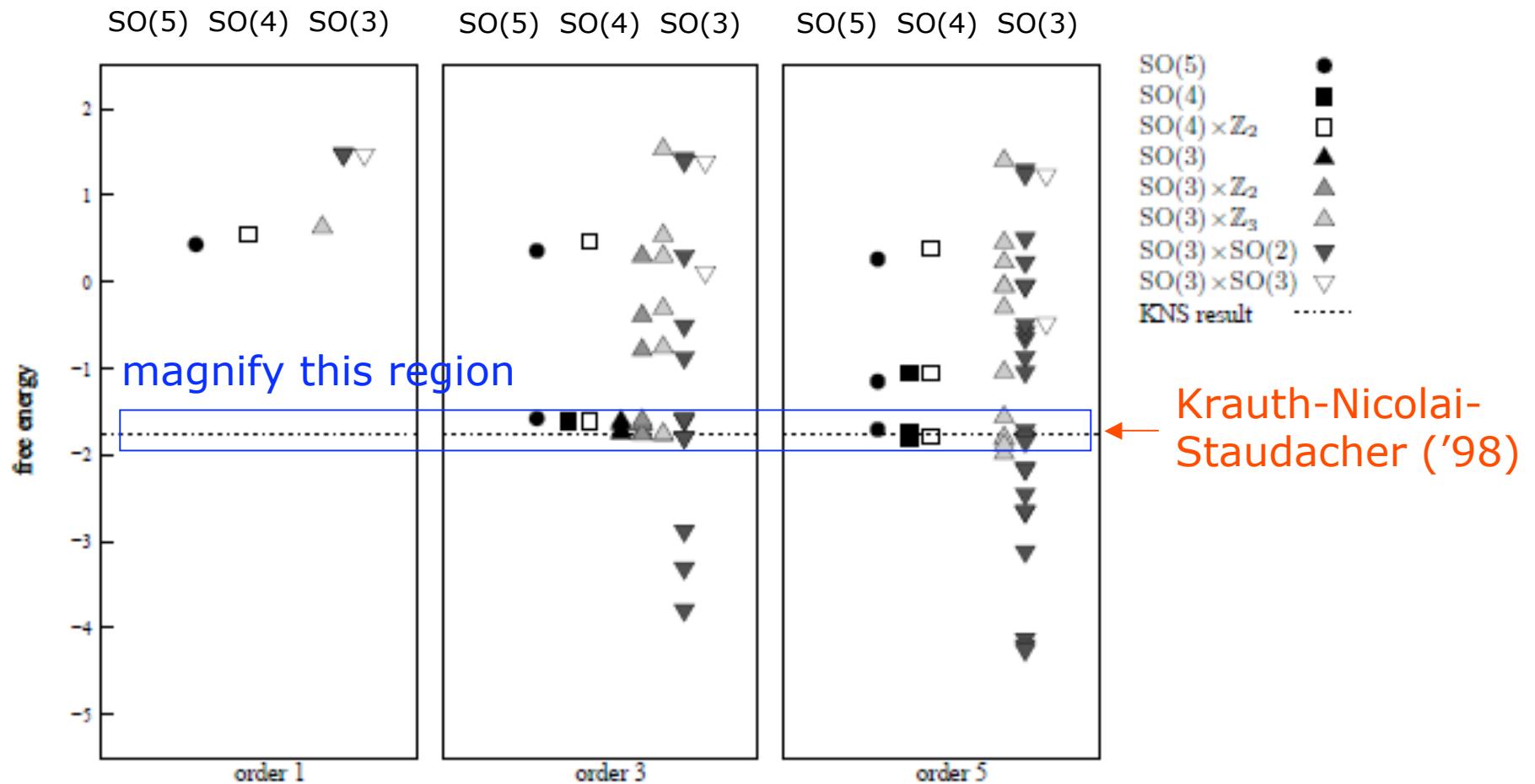
$$\text{SO}(6) \xrightarrow{\text{SSB}} \text{SO}(3)$$

finite extra dimensions !

confirmation by MC sim.  
Aoyama-Azuma-Hanada-J.N.,  
in progress

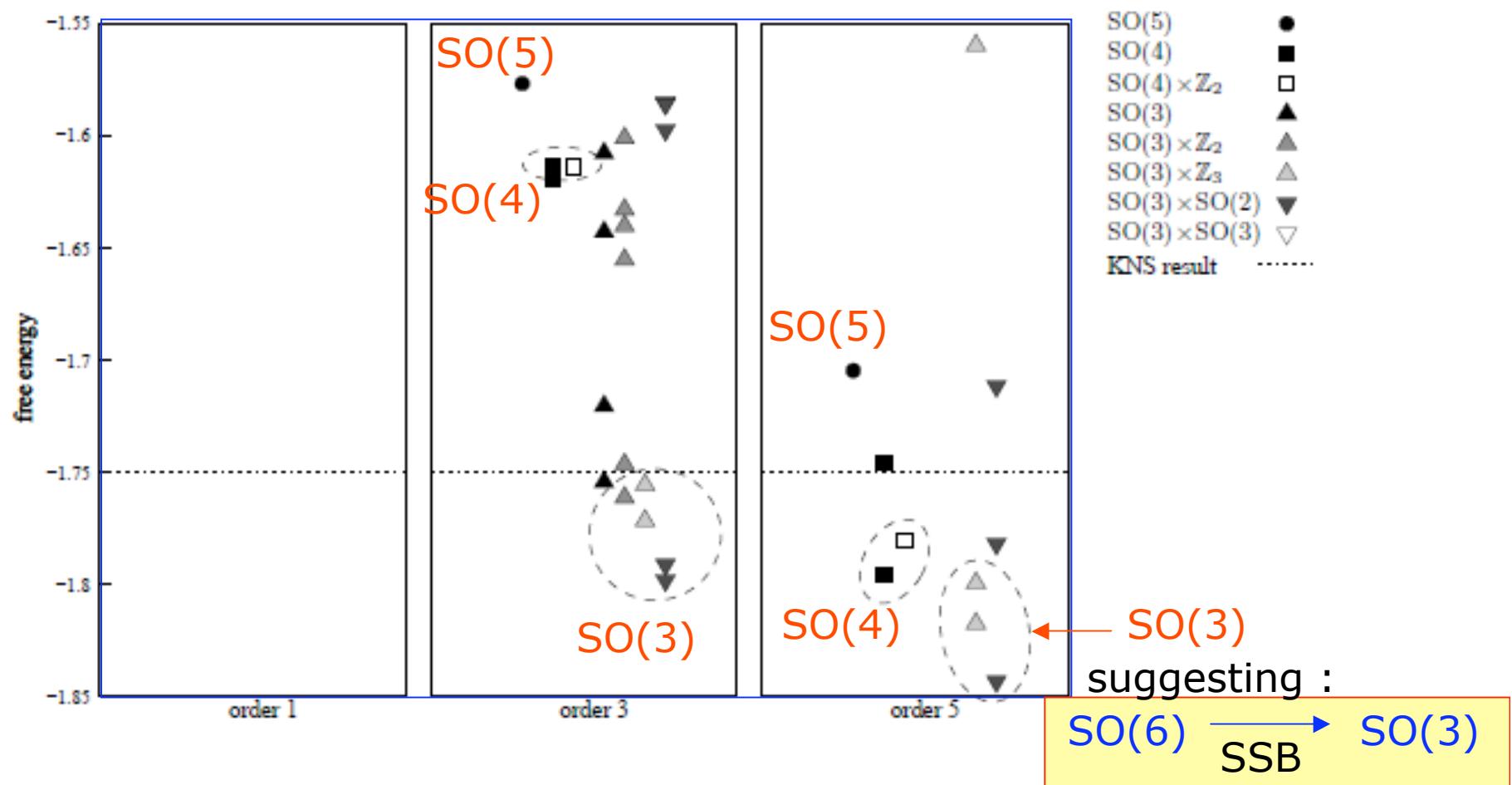
# Results of GEM for the little IIB matrix model

Aoyama-J.N.-Okubo, in prep.



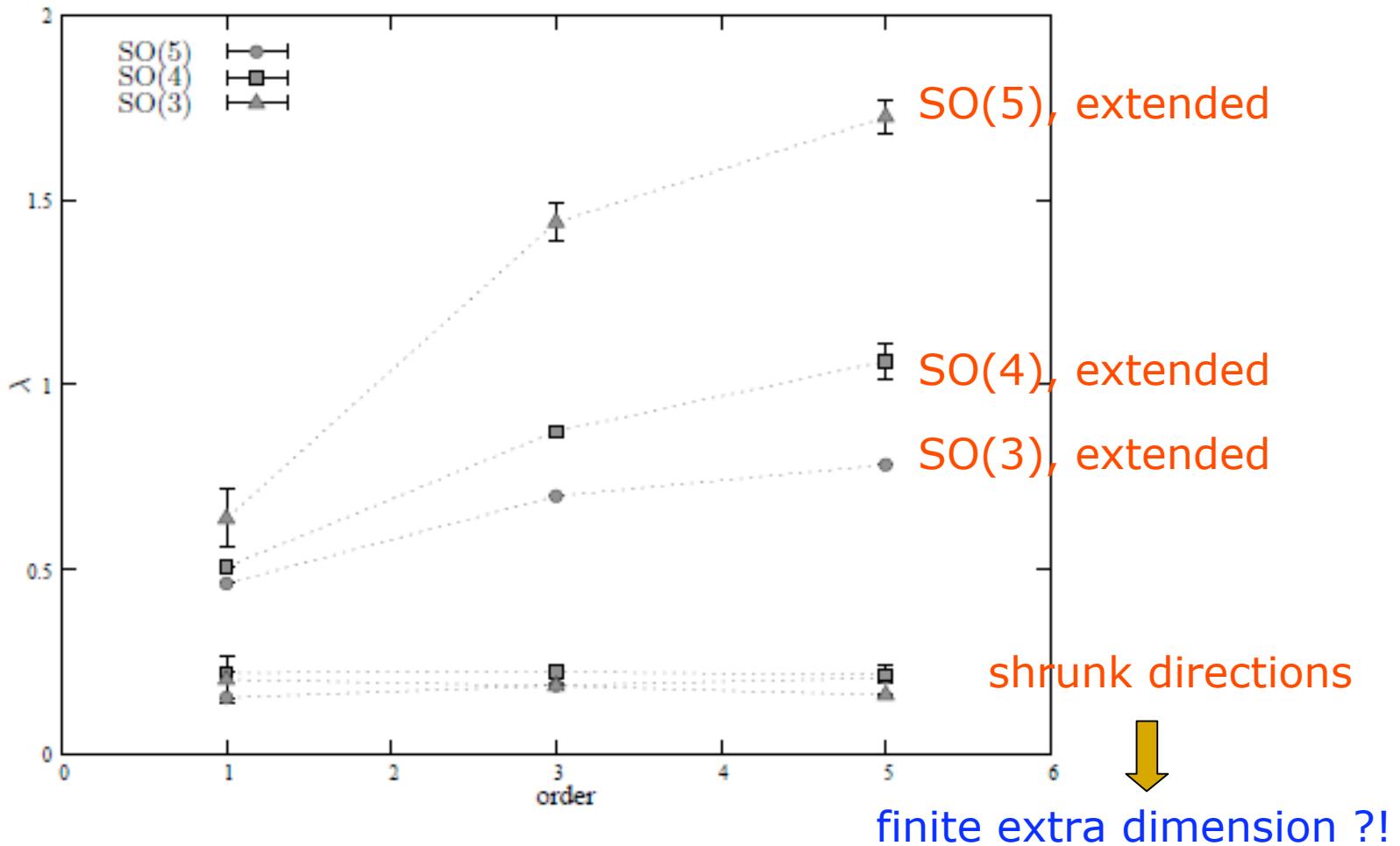
# Results of GEM for the little IIB matrix model (cont'd)

concentration of solutions identified



# Results of GEM for the little IIB matrix model (cont'd)

extent of the eigenvalue distribution  
in the extended/shrunk direction



# Future prospects

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easier, and many things to explore
- higher dimensional case  
possible using mass deformation  
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  - SO(6)  $\rightarrow$  SO(3) from Gaussian expansion (Aoyama-J.N.-Okubo, in prep.)