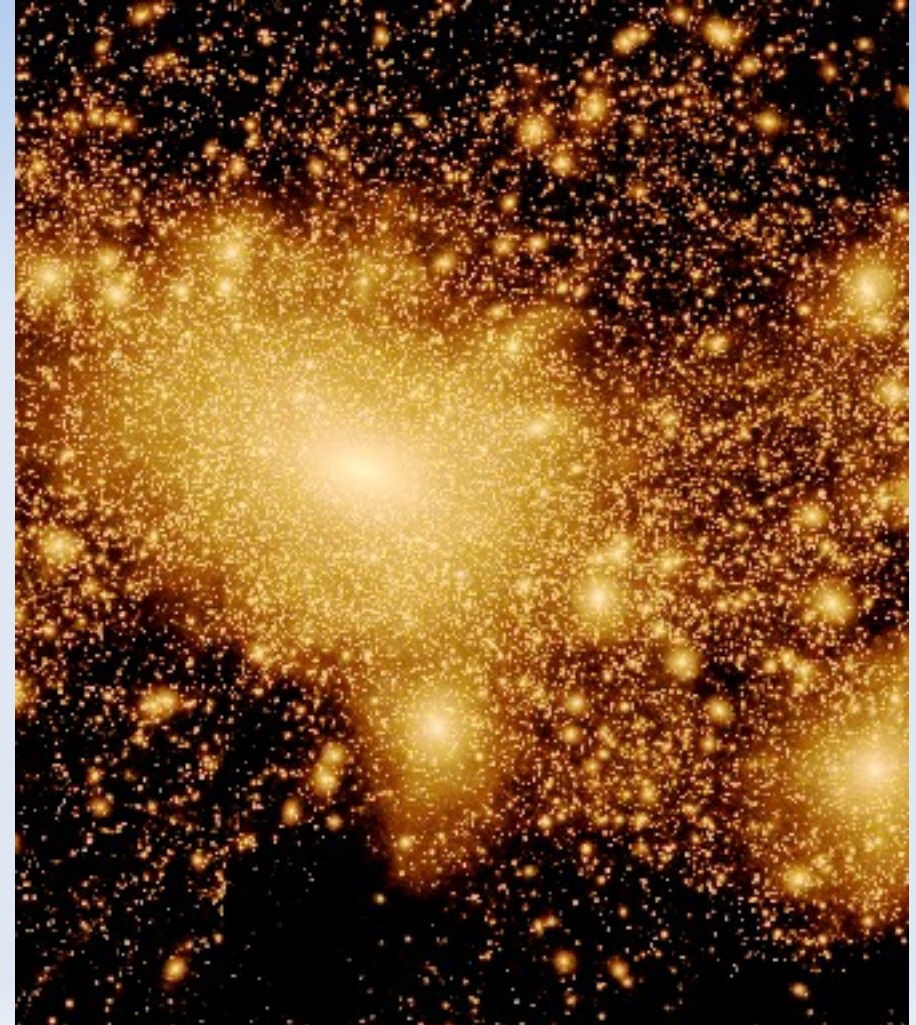


Astrophysical Dark Matter Structures

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Gausdal 2012

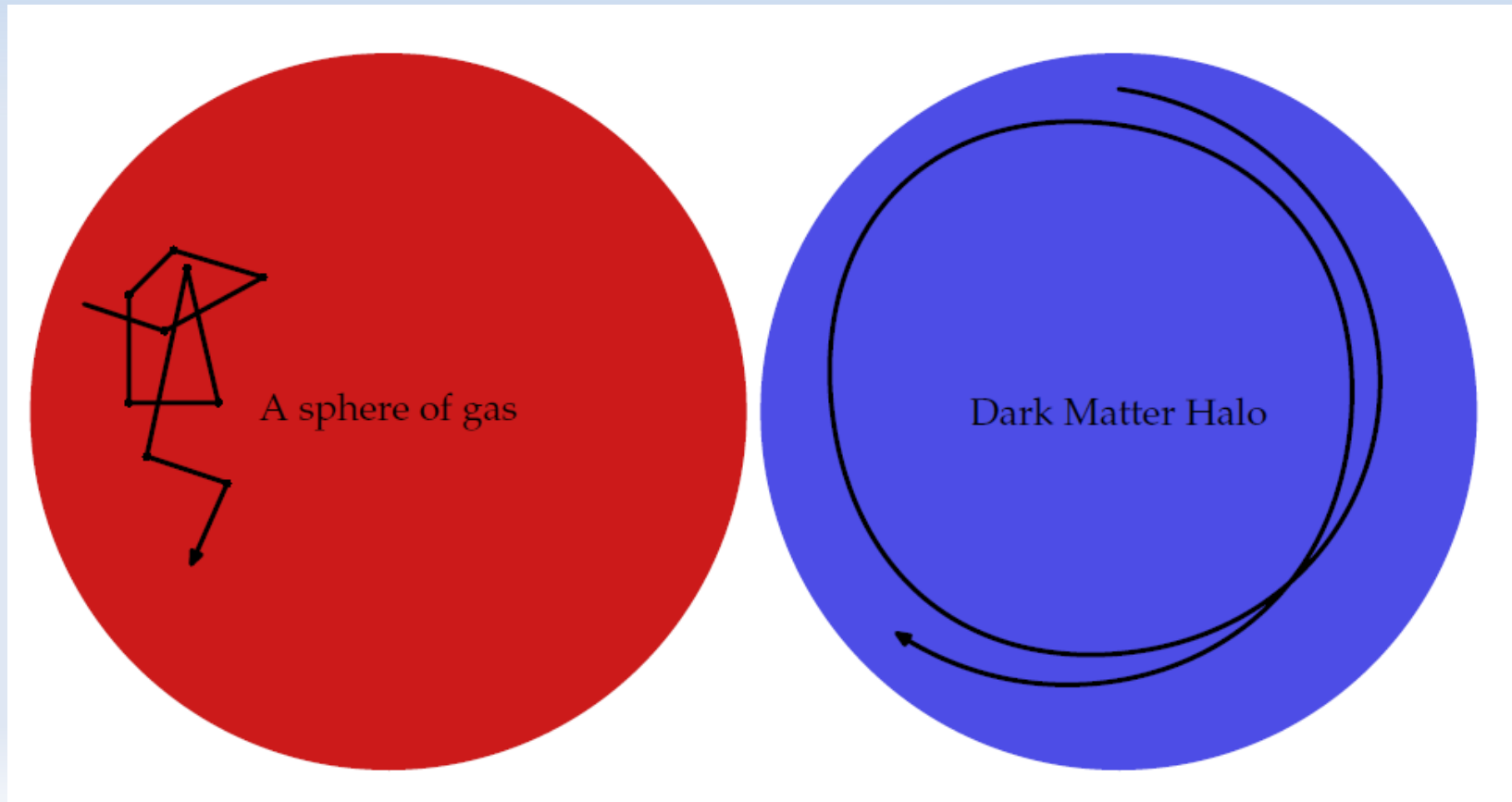


Simulating the Formation of Dark Matter Structures in the Universe

- The initial distribution of matter comes from inflation (Harrison-Zeldovich power spectrum).
- The growth of these perturbations is calculated with
 - 1) Linear perturbation theory (until $t \sim 50$ Myr).
 - 2) Numerical simulations, where matter densities and velocities are represented by particles.

Collisionless Particles

- Dark matter particles are collisionless unlike normal gas particles:



Cosmological Simulation Movie

- The formation of structure from $z=12$ to $z=0$ ($t=0.375$ Gyr to today):

Credit:

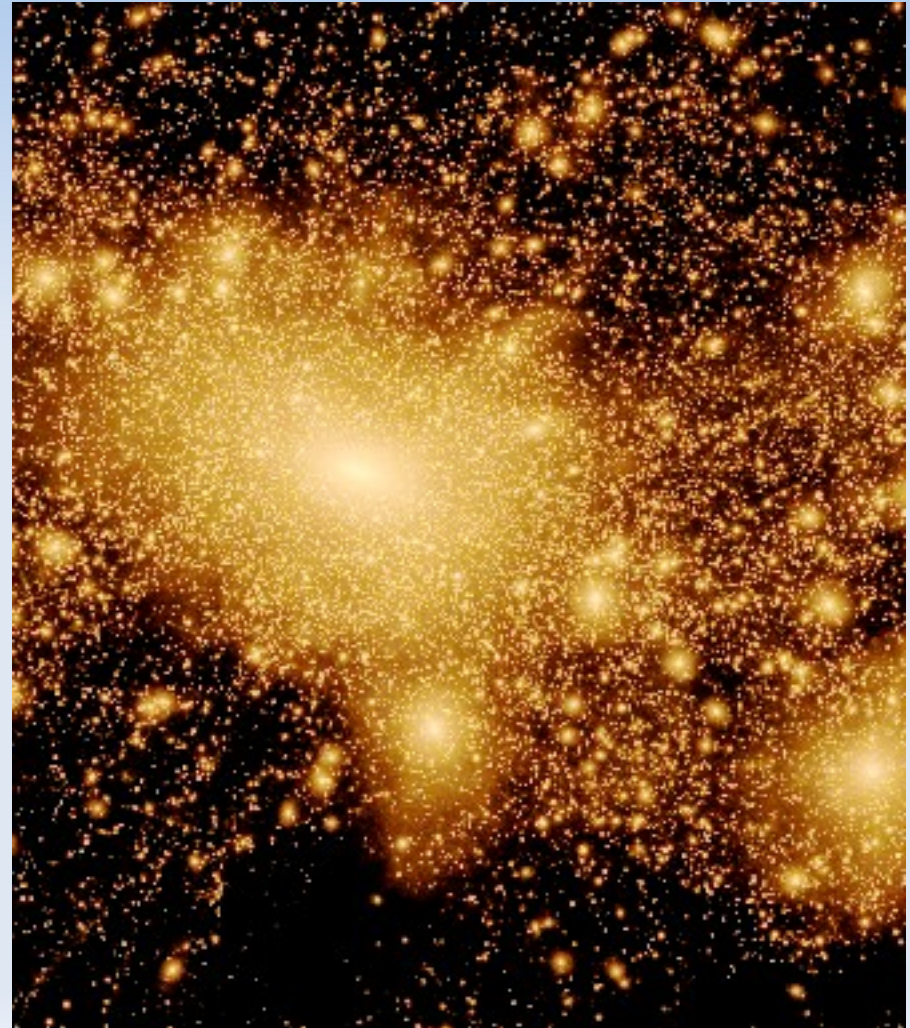
Diemand et al.

The Via Lactea Project

Results from simulations

- Halos are complicated.
- Universal density profile ("NFW-profile"):

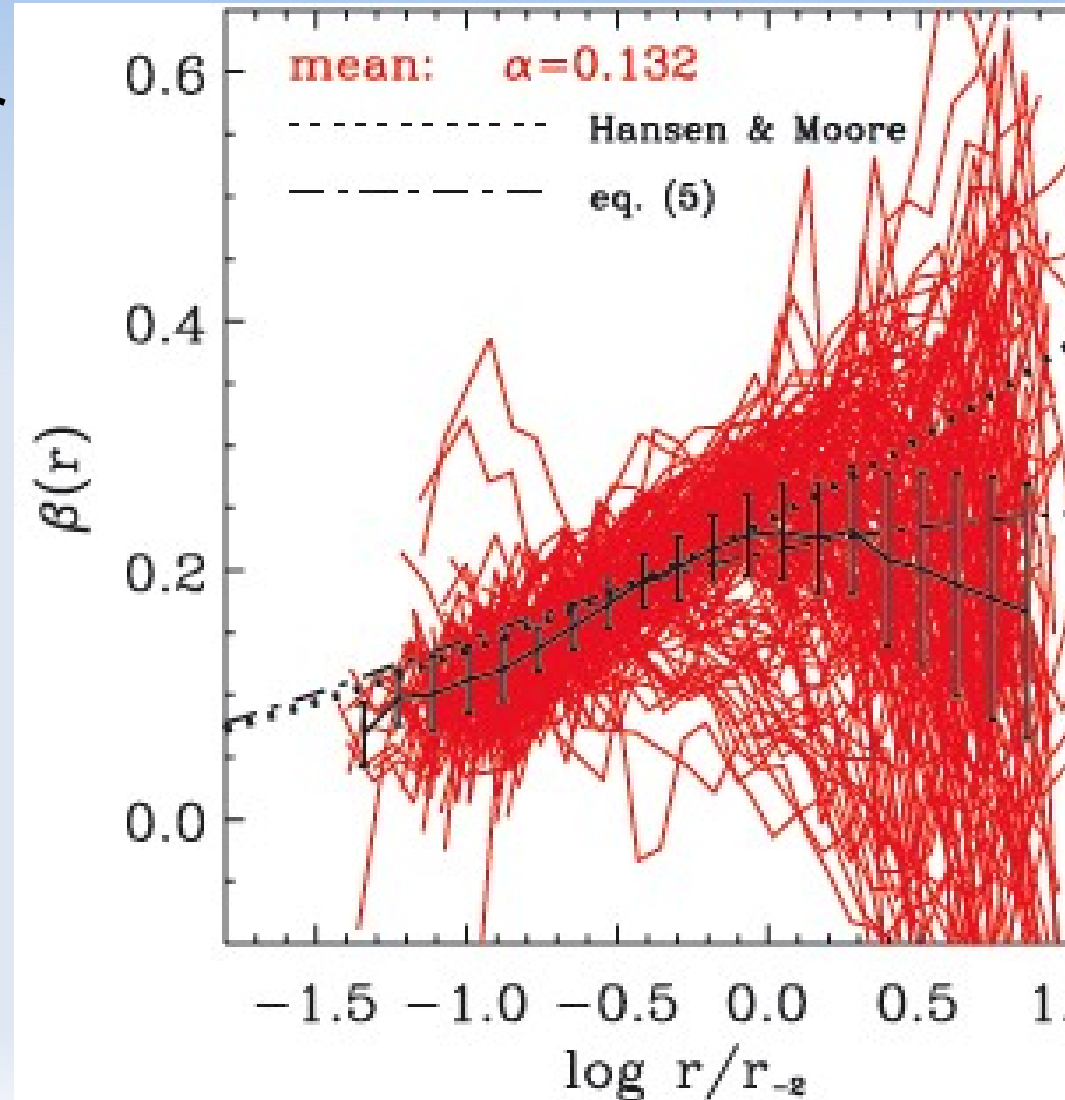
$$\rho(r) \propto \frac{1}{r} \frac{1}{(1+r)^2}$$



Anisotropic velocity distributions

- Velocity dispersion is larger in the radial direction than in a tangential direction.
- Definition:

$$\beta(r) \equiv 1 - \frac{\sigma_{\text{tangential}}^2(r)}{\sigma_{\text{radial}}^2(r)}$$



Ludlow et al. (2011) ->

My own work

- Controlled simulations....

Collaborators:

Steen Hansen, Diana Juncher, Johan Samsing

Simulations

- We generated many different initial DM structures.
- The particle-velocities were multiplied by different random numbers.
- The systems was then evolved in time (with a N-body simulation code...)
- ... and then perturbed again.

Hansen, Juncher & Sparre (2010)

Simulations

- beta, the velocity anisotropy parameter.

- Density slope, $\frac{d \log \rho(r)}{d \log r}$

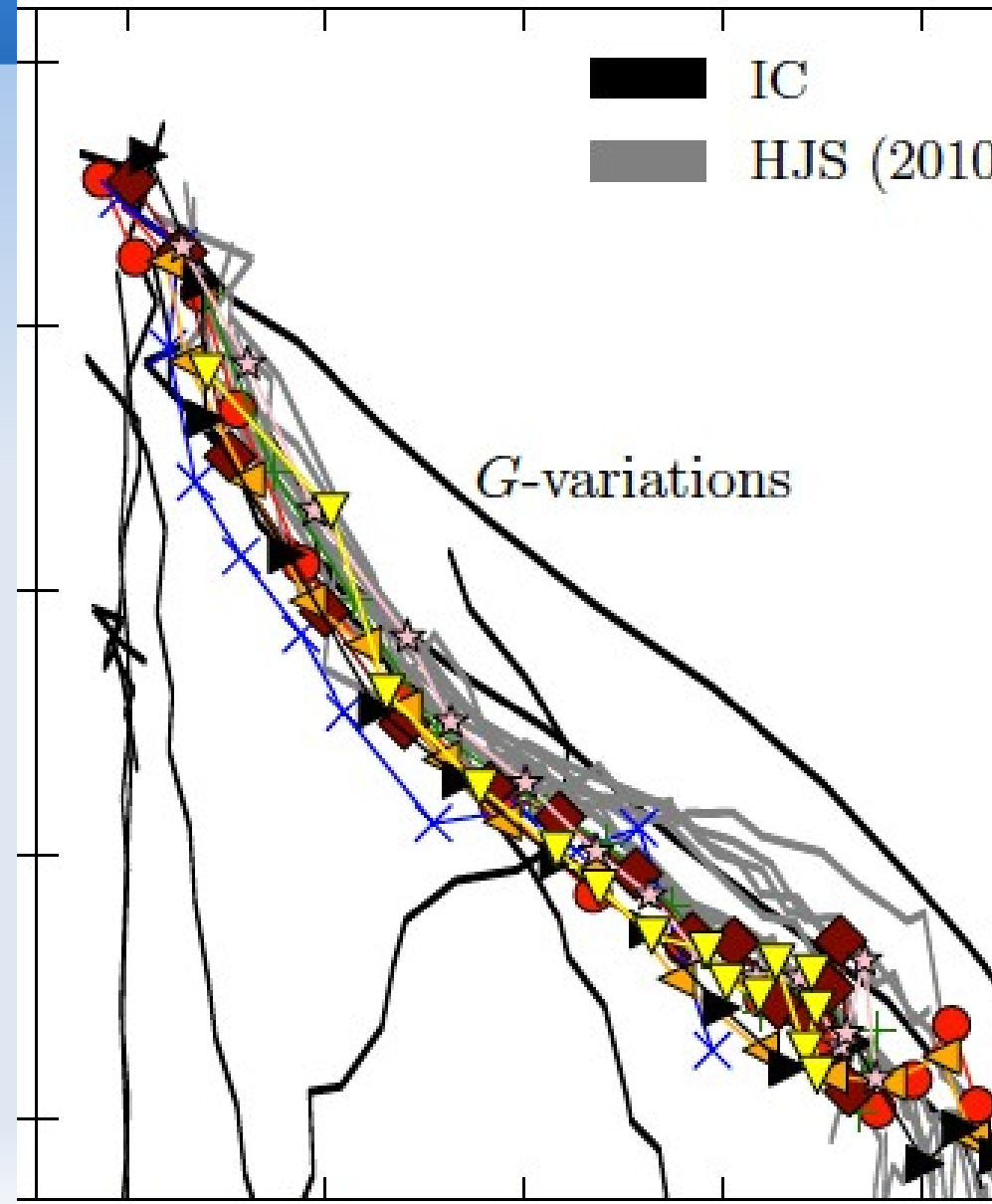
- Velocity dispersion slope $\frac{d \log \sigma_r^2(r)}{d \log r}$

Show animation...

Another Experiment

- The acceleration of the particles were changed instantaneously.
- The structures ended on "the attractor".

Sparre & Hansen
(2012, submitted)



Conclusions

- A 1-dimensional attractor for dark matter systems has been identified in a set of simulations.
- The attractor determines the velocity distribution and the velocity anisotropy for dark matter systems.
- Halos in the universe are expected to be on this attractor.

See more in **Hansen, Juncher & Sparre (2010)**, and **Sparre & Hansen (2012, submitted)**.

What determines the end product of structures?

- The equilibrium state of (ideal) gas particles in a box can be calculated from the maximum entropy principle.
- This method does not work for dark matter structures.

Example: Maximizing the entropy of a self-gravitating structures consisting of DM-particles leads to structures with infinite mass.