SCAPEZILLA : the backreaction of anti-branes in flux compactifications

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Metastable vacua

 Exist in gauge theories N=1SQCD Intriligator, Seiberg, Shih Lots of other theories everybody and their brother • No type IIA realizations of metastable vacua Bena, Gorbatov, Hellerman, Seiberg, Shih · Why?

No IIA brane realization

- N=1 engineered with D4 + NS5
- D4 ends on codimension 2 line inside NS5
- End of D4 branes sources log mode on NS5
- NS5 brane bending
 - ⇔ Log running of N=1 coupling constant Witten
- Tiny IR perturbation $\Rightarrow \log \Rightarrow UV$ messed up

different $UV \Leftrightarrow$ not vacua of the same theory

Bena, Gorbatov, Hellerman, Seiberg, Shih

What about AdS-CFT No asmpt-AdS₅ metastable solutions One candidate: Kachru Pearson Verlinde Antí-D3 branes in Klebanov Strassler • Codímension 6 $\xrightarrow{2}$ modes ~ $1/r^4$ • Normalizable \Rightarrow metastable vacuum Much used in string cosmology

Klebanov-Strassler



 $\frac{1}{4\pi^2 \alpha'} \int_{S^3} F^{(3)} = M$

r = 0IR
D3 charge dissolved in fluxes
H3 x F3 \rightarrow F5
F5 x F3 \rightarrow H3

UV



AdS-CFT modes

- Normalizable modes (NM)
 - dual to vevs
 - Fíníte energy, IR
- Non-normalizable (NNM)
 - deformations of Lagrangian
 - Infinite energy, UV



Energy

BDHM - BKLT

- Different NNM \Rightarrow different theories
- ◆ Same NNM ⇒ different vacua, same theory

metastable \Leftrightarrow NNM=0

Big Question

Antí-D3 \Rightarrow normalizable or non-normalizable modes?

- Fluxes ⇒ KS field ~ logr
- encodes log running of coupling constant $\frac{1}{g_1^2} - \frac{1}{g_2^2} \sim \int_{S^2} B_2 \sim \log r$
- Anti-D3 couple to this field
- IIA intuition: log messed up \Rightarrow non-normalizable
- every dual of non-conformal 4D theory ⇒ log modes

Big Implications if NNM

0

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- No AdS-CFT metastable 4D vacua
- String cosmology/landscap /:

anti-D3 down long KS throats → redshift → tunably-small energy → lift AdS to dS KKLT, etc. anti-D3 non-normalizable energy not tunably-small moduli stabilization messed up

$$V = \frac{aAe^{-a\sigma}}{2\sigma^2} \left(\frac{1}{3} \sigma aAe^{-a\sigma} + W_0 + Ae^{-a\sigma} \right) + \frac{D}{\sigma^3} \longrightarrow \frac{3 \times 10^{-9}}{\sim 1}$$

Scape-zílla

- 4D N=1 gauge theories log running generic phenomenon, not restricted to KS
- Same happens in LARGE volume scenarios
- No vacuum uplift by small-energy ! antí-D3 gíve O(1) contribution !
- Landscape of AdS vacua

Landscape of dS vacua





Smear anti-D3's $SU(2) \times SU(2) \times \mathbb{Z}_2$ Solution(T) Perturbation theory in anti-D3 number

8 modes satisfying second-order eqs.
16 integration constants
expanded around BPS solution ⇒ first-order system:

 $\frac{d\xi_a}{d\tau} + \xi_b M^b{}_a(\phi_0) = 0,$ $\frac{d\phi_1^a}{d\tau} - M^a{}_b(\phi_0)\phi_1^b = G^{ab}\xi_b$

Papadopoulos, Tseytlin 2000 Borokhov,Gubser 2002 Kuperstein, Sonnenschein 2003

The Hunting Method

- Solve first 8 equations for ξ . Integration constants X.
- Use ξ + other 8 eqs. to get ϕ . Integration constants Y

dim Δ	non-norm/norm	int. constant
8	r^4/r^{-8}	Y_4/X_1
7	r^{3}/r^{-7}	Y_5/X_6
6	r^2/r^{-6}	X_{3}/Y_{3}
- 5	r/r^{-5}	
4	r^{0}/r^{-4}	$Y_7, Y_8, Y_1/X_5, X_4, X_8$
3	r^{-1}/r^{-3}	$X_2, X_7/Y_6, Y_2$
2	r^{-2}/r^{-2}	

 X_2 and $X_7 \sim 1/r$

non-normalizable

The hard work

- Implicit solution 8 nested integrals
- Smart grad students → nested integrals
 can be simplified:
- ξ solved in terms of one integral !
- • 2 or 3 nested integrals !
- · Easy to find all mode profiles numerically

The silver bullet !!! 16 constants - 14 physical ones Probe D3 brane attracted by antí-D3's • Force is universal: KKLMMT $F_r \sim \frac{N_{\overline{D3}}}{m_5}$ • We get $F_r \sim \frac{X_1}{r^5} + \mathcal{O}\left(\frac{1}{r^{11}}\right)$ Only depends on 1 of the 14 constants !!! • Only force-mode is ξ_1

Look in the infrared

- Kill very divergent guys + ξ_1 must be nonzero !!!
- Physical divergence: antí-D3 smeared on S³
- Warp factor diverges $\sim \tau^{-1}$
- Curvature diverges: $R \sim F_{(5)}^2 \sim \tau^{-4}$
- Another divergence no obvious reason

• Subleading singularity ~ ξ_1

Everything depends on it !!!

 $H_{(3)}^2 \sim F_{(3)}^2 \sim \tau^{-2}$ Must be there !!!

Antí-D3 in KS is normalizable

- Dual to gauge theory metastable vacuum
- Nice physics vev's etc.
- Hunt for gauge theory dual

Dymarsky Klebanov Seiberg

AdS can be uplifted to dS
Landscape of dS vacua alive and frisky
No Scapezilla

If singularity unphysical:

- antí-D3 sources non-normalizable modes
- IR couplings to log mode (H₃) mess up UV
- No more dS landscape SCAPEZILLA
- Reminder BPS solution:
- $F_5 \times F_3 \rightarrow H_3$
- $H_3 \times F_3 \rightarrow F_5$



If singularity unphysical:

- $(-F_5) \times F_3 \rightarrow -H_3$
- $(-H_3) \times F_3 \rightarrow -F_5$
- Sign of D3 charge dissolved in flux not fixed !!!
- Only F₃ flux on S³ fixed.

anti-D3 dissolved in flux

Only physical solution with anti-D3 is anti-KS !!!

Is this generic?

- Do anti-branes always hate charge dissolved in flux ?
- I hope not ...
- M-theory version of Klebanov-Strassler CGLP
 Cvetic, Gibbons, Lu, Pope
- M2 + transverse 8D Stenzel Space, magnetic F4 + F4
- M2 charge in fluxes
- add antí-M2 → metastable
 Klebanov, Pufu
- Perturbative solution = singular !
- Idem for anti-D2 in CGLP, A8
- Insane antíbranes Giecold, Orsi, Puhm



What about non-extremal fuzzballs ? We have many many many BPS or extremal horízonless mícrostate geometríes (fuzzballs):



Bena, Bobev, Bossard, Dall'Agata, deBoer, Giusto, Niehoff, Ruef, Shigemori, Vasilakis, Warner & friends



Non-extremal microstates ? Add metastable supertube wrapping GH fiber: Bena, Puhm, Vercnocke





Decays via brane-flux annihilation May be only way to construct stationary nonextremal microstate geometries Gibbons, Warner



• If not physical:

- antibranes cannot coexist with charge in fluxes
- maybe no more dS landscape ☺
- maybe no systematic way to build non-extremal stationary microstate geometries (fuzzballs) ©
- brane of codimension 6 + fluxes $\rightarrow \log \mod es$

So it must be physical !!!



Incorrect AdS-CFT

- One should a-priori take only normalizable modes in UV, and accept whatever exists in the IR
- Maybe, but not in AdS-CFT
- IR regularity crucial to relate NNM with NM. Otherwise get wrong physics:
 - AdS-QCD-CMT without incoming b.c. at black hole
 - Confinement from Klebanov-Tseytlin

Scapezilla not easy to kill

Antí-D3 síngularity @ fírst-order backreaction

May go away at full backreaction Dymarsky

No intention: Bena, Grana, Kuperstein, Massai
 1. Eliminate IR singularity
 2a. Find full solution in an IR expansion to order T¹⁰
 2b. Examine r.h.s. of nonlinear er Antí-M2's as well
 Only possible solution with anti-

Anti-D3's are singular to the bitter end

ivergent energy density is finite!

- We can be by ast about origin of singularity
- · ccept Everything with the IR action
- After al Ads PT relates bulk and boundary actions

Counter-argument:

Klebanov

(Dymarsky)

- Negative-mass Sci warz ching
 Horowitz-Myers
- Integral of divergent energy dentity is finite/
- Must be eliminated if AdS-CFT. +8 make any sens
- Furthermore, antí-M2 and antí-D2 síngue rities have dívergent IR action

- Singularity indicates new physics
 - Instabilities
 - Polarization:
- Probe antí-D3's polarize into NS5 branes/S² ⊂ S³
 this could resolve singularity à la Polchinski-Strassler
- Smearing wipes out this polarization channel:
- PS has many channels: D5 branes/S² ⊂ T^{1,1} survive smearing
 No smeared anti-D3+D5 → no localized anti-D3+D5 ⇒
 no localized anti-D3+NS5 branes either !!!

Why Polchinski-Strassler does not save the landscape revenge on Bousso-Polchínskí © $V(\tau) \sim (2\pi n) a_2 \tau^2 - a_3 \tau^3 + \frac{1}{2\pi n} a_4 \tau^4$ Good intuition Same potential terms as in PS! No polarization if: $(a_3)^2 < \frac{32}{9}a_2a_4$ Long calculation: $a_2 = \frac{1}{3p^2} \left(4\lambda_f^2 + 3\lambda_F^2 \right), \quad a_3 = \frac{2}{3p} \lambda_f, \quad a_4 = \frac{1}{8}$ Could have worked, but it does not !!!

Maybe we are not smart-enough to understand resolution

- "Good, Bad, Ugly" criterion: Gubser
 Good singularities can be cloaked by horizon
- If physical $\Rightarrow \exists BH in KS/KT$ with negative charge

All KS/KT black holes must have positive charge: Bena, Buchel, Días

Black hole in Klebanov-Strassler/Tseytlin

Aharony, Buchel, Kerner; Buchel

- Maybe artifact of smearing
- Localized anti-branes may not have this problem
- ♦ Localized BH with anti-D3 charge in KS exists

Can be anywhere on S^3 Could be smeared

Smeared BH with negative charge does not exist Bena, Buchel, Dias



- Nobody could have predicted it a-priori !
- No a-posteriori physical reason for accepting it
- Several highly nontrivial calculations that could have worked either for or against - all worked against

What would help

- Localized anti-D3 in KS
- Localized BH in KS
 - Non-BPS solution, 2 variables
 - Separation of scales
- No smeared BH solution

 → no localized BH solution
 - Is this always true ? If not why ?
- Solution for smeared anti-M2, anti-D2 black holes in CGLP, A8
 - Would confirm whether anti-D3 story is generic or not
 - One variable shooting or relaxation straightforward.

What would help

- Metastable supertube solution
 - cannot smear \rightarrow 2 variables !
 - supertube charges: (-,-) or (+,-)
- Numerics ? ... BlackFold ? ...
 Separation of scales ? ... Inverse scattering ? ... Perturbative ?
 - first fully-backreacted microstate geometry of a nonextremal BH with macroscopic horizon
 - existence of gazillions of microstates resolve info paradox
 mechanism that keeps them from collapsing into BH (which nobody else has ⁽³⁾)



Conclusions

- Probe antibranes uplift AdS to dS
- Probe antibranes give stationary near-extremal fuzzballs
- Backreacted antibranes have singularity
- No reason to accept it. So far all evidence against.

If unphysical:

- A lot of string cosmology and phenomenology to be revisited.
- SCAPEZILLA: AdS landscape ≠ dS landscape
- Find other ways to uplift AdS to dS (Kahler uplifting? nonperturbative effects ? nothing ?)
- Find other ways to build non-extremal fuzzballs (JMaRT-type centers ? motion on moduli space ? inverse scattering ? numerics ?)