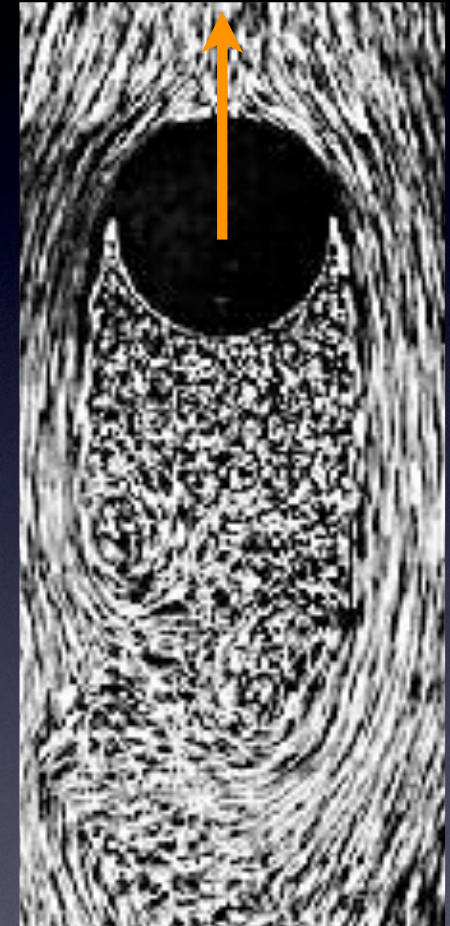
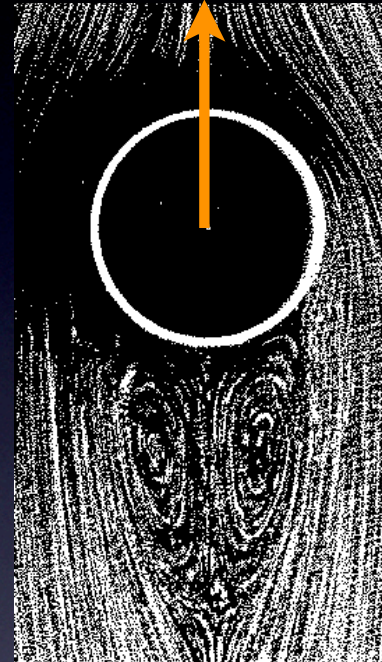
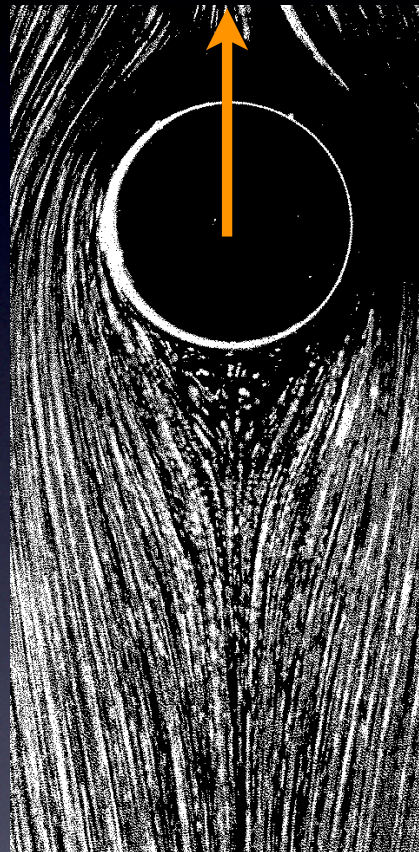
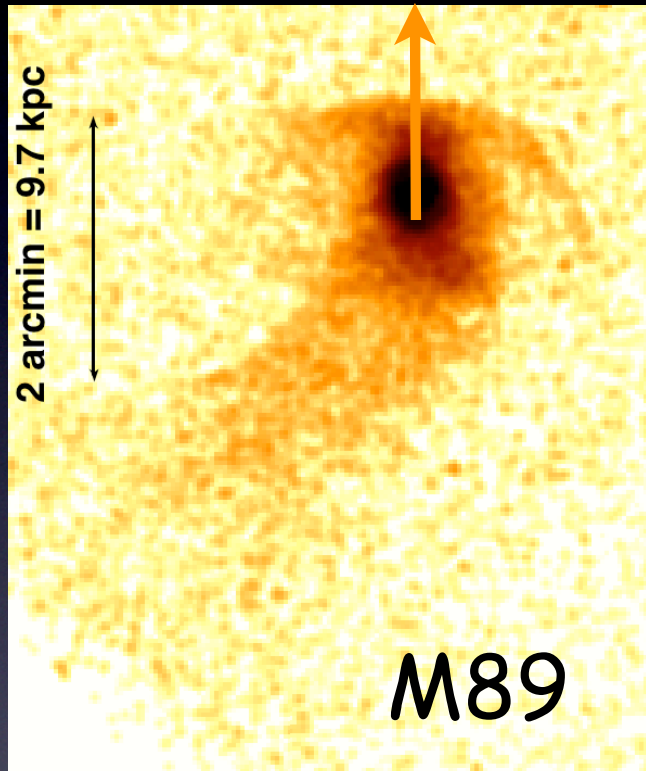


Shaken and stirred - but mixed? Viscous flow experiments in the ICM

Kraft et al., Roediger et al., in prep.

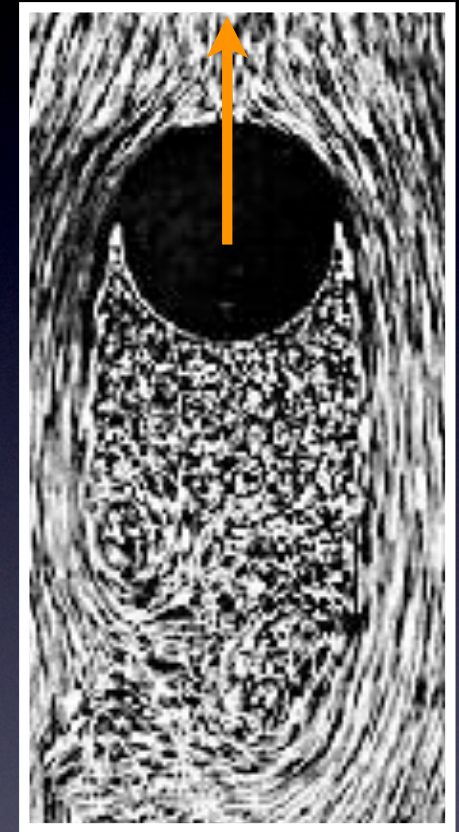
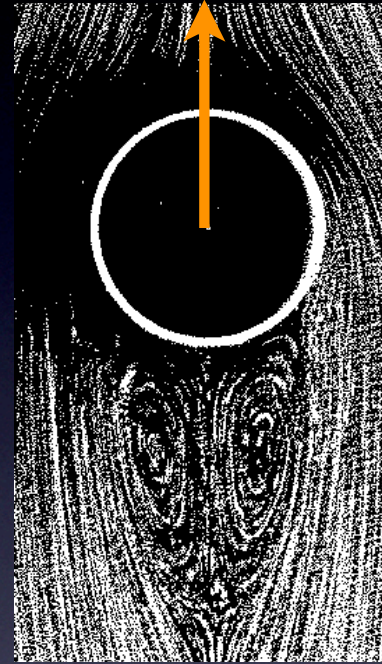
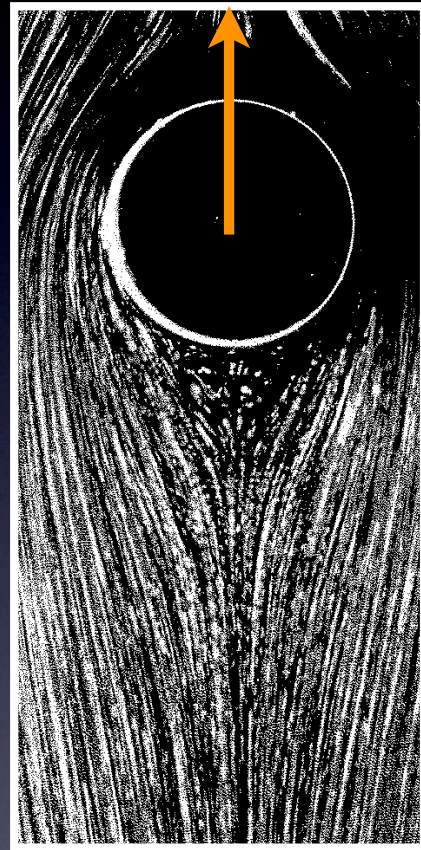
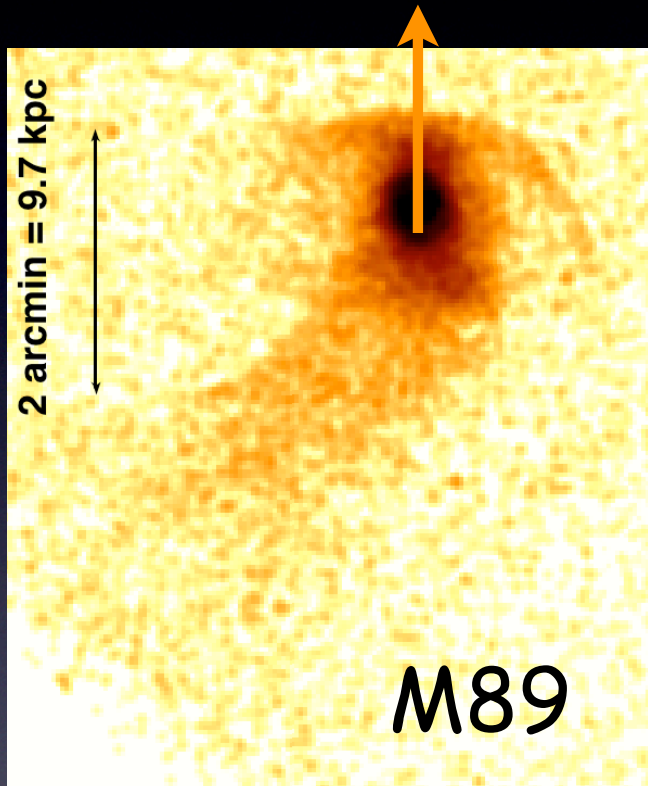


van Dyke - Album of Fluid Motion

*Elke Roediger (Hamburger Sternwarte), Ralph Kraft,
Bill Forman, Paul Nulsen, Marie Machacek, Scott Randall, Christine
Jones (CfA), Eugene Churazov (MPA)*

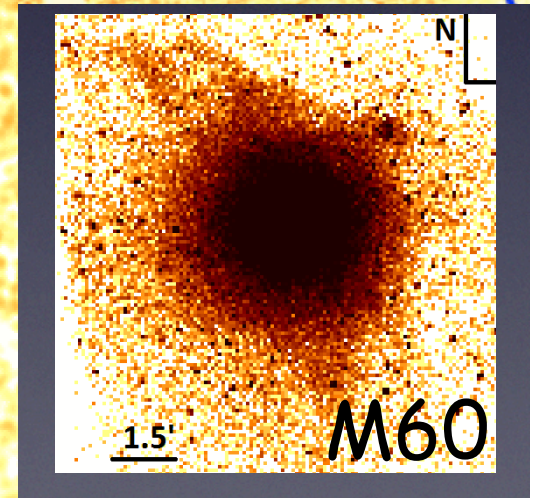
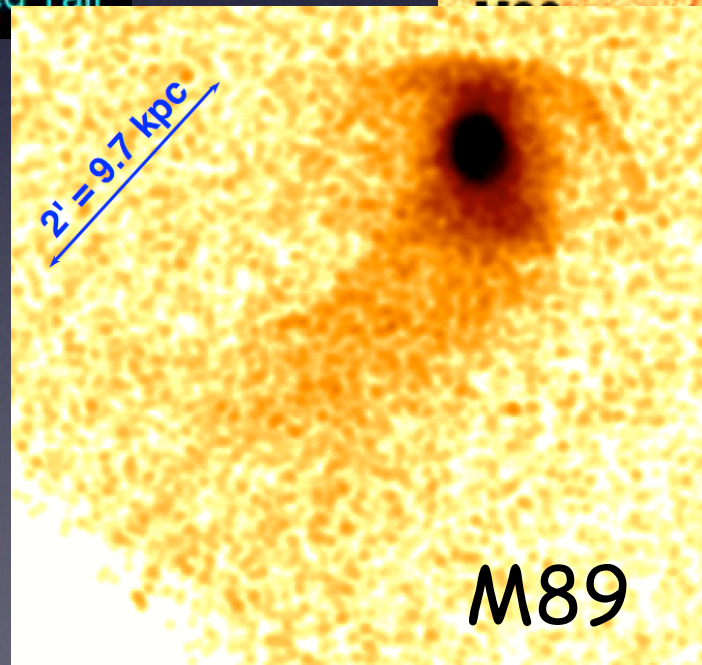
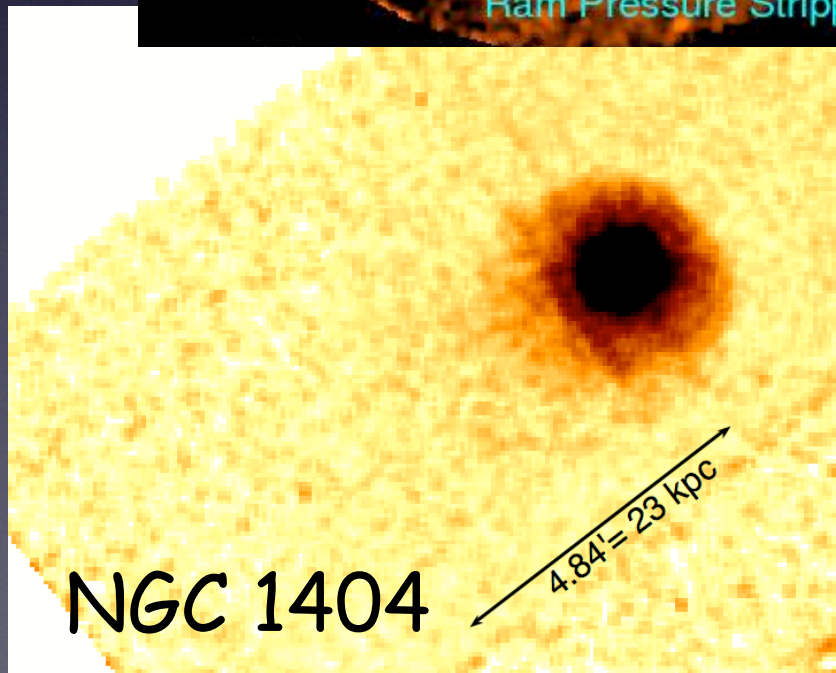
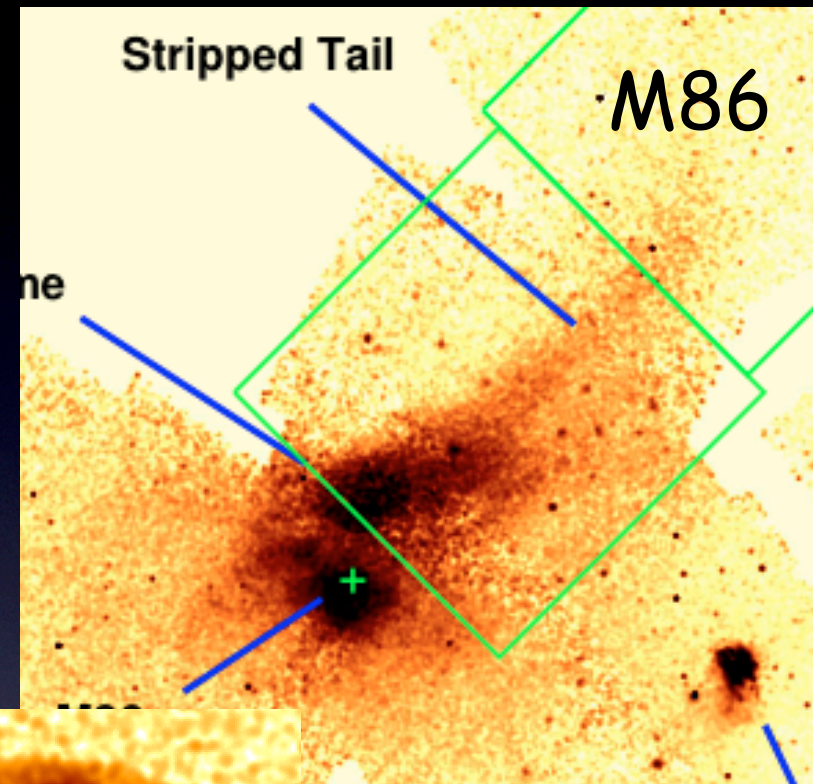
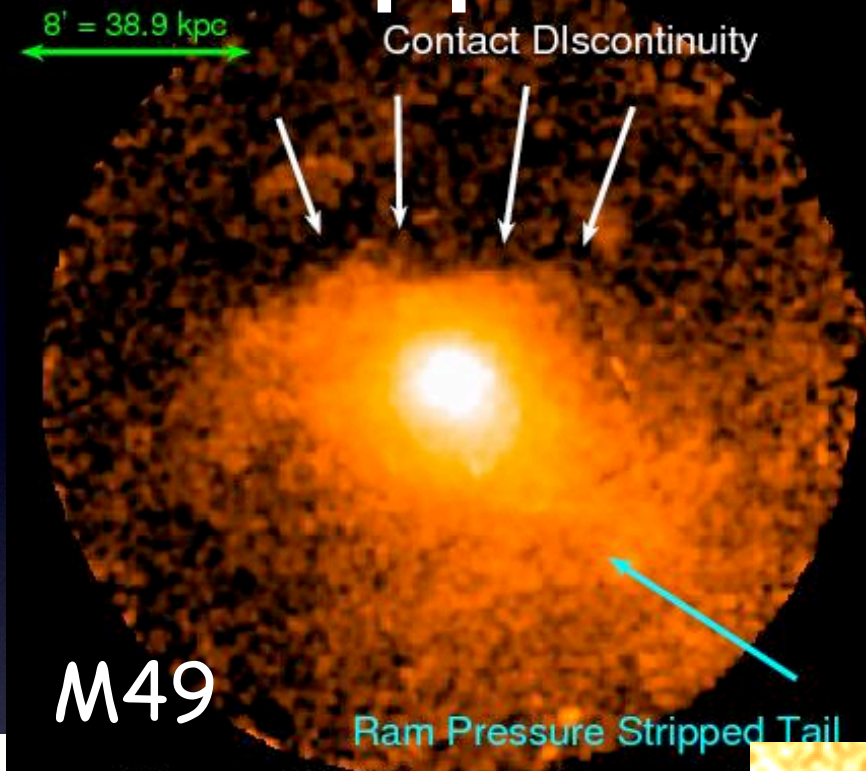
Gas-stripped elliptical galaxies

Kraft et al., Roediger et al., in prep.

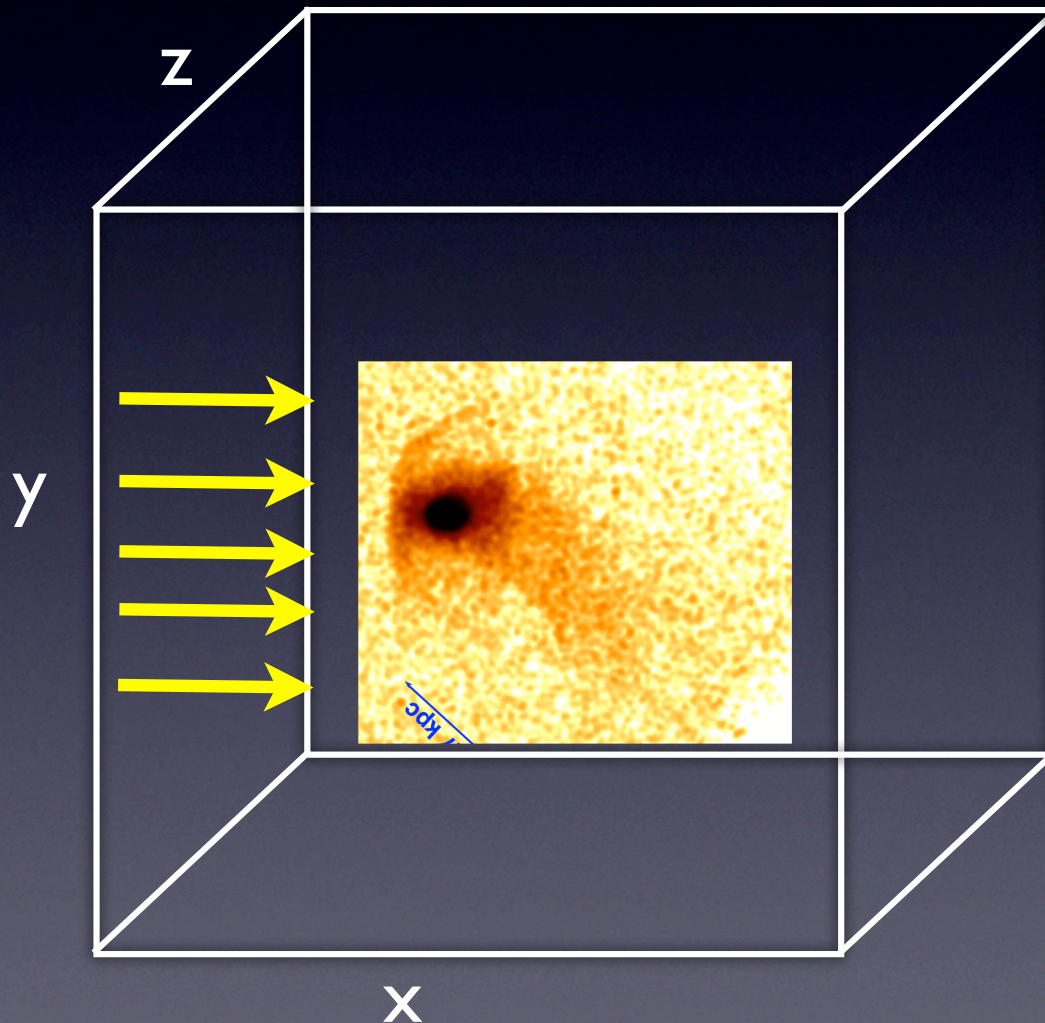


van Dyke - Album of Fluid Motion

Gas-stripped elliptical galaxies



Simulating galaxy stripping

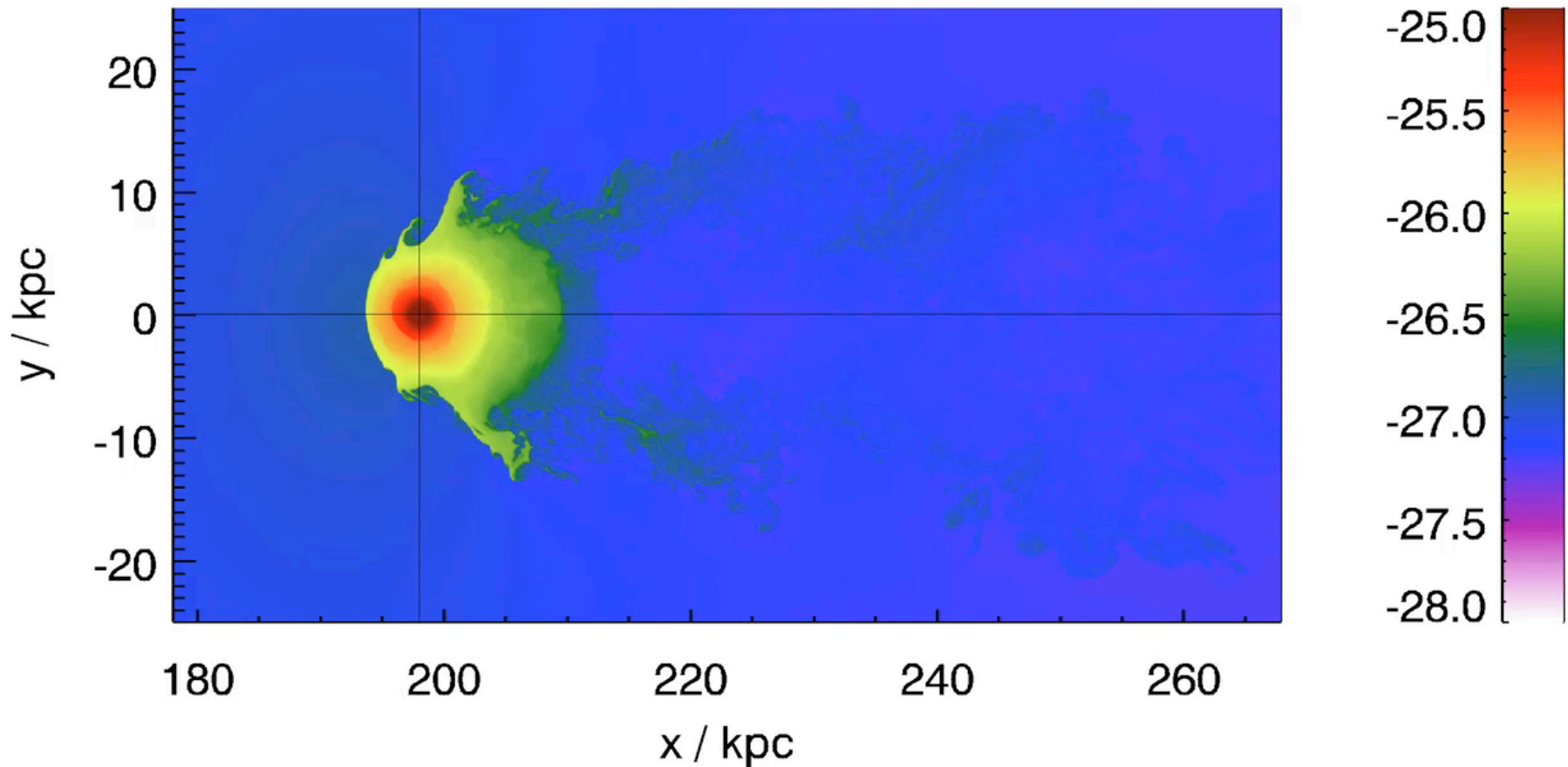


- FLASH code
- rest frame of galaxy
- galaxy potential tailored to M89
- ambient ICM head wind mimicks M89's infall into Virgo

inviscid galaxy stripping

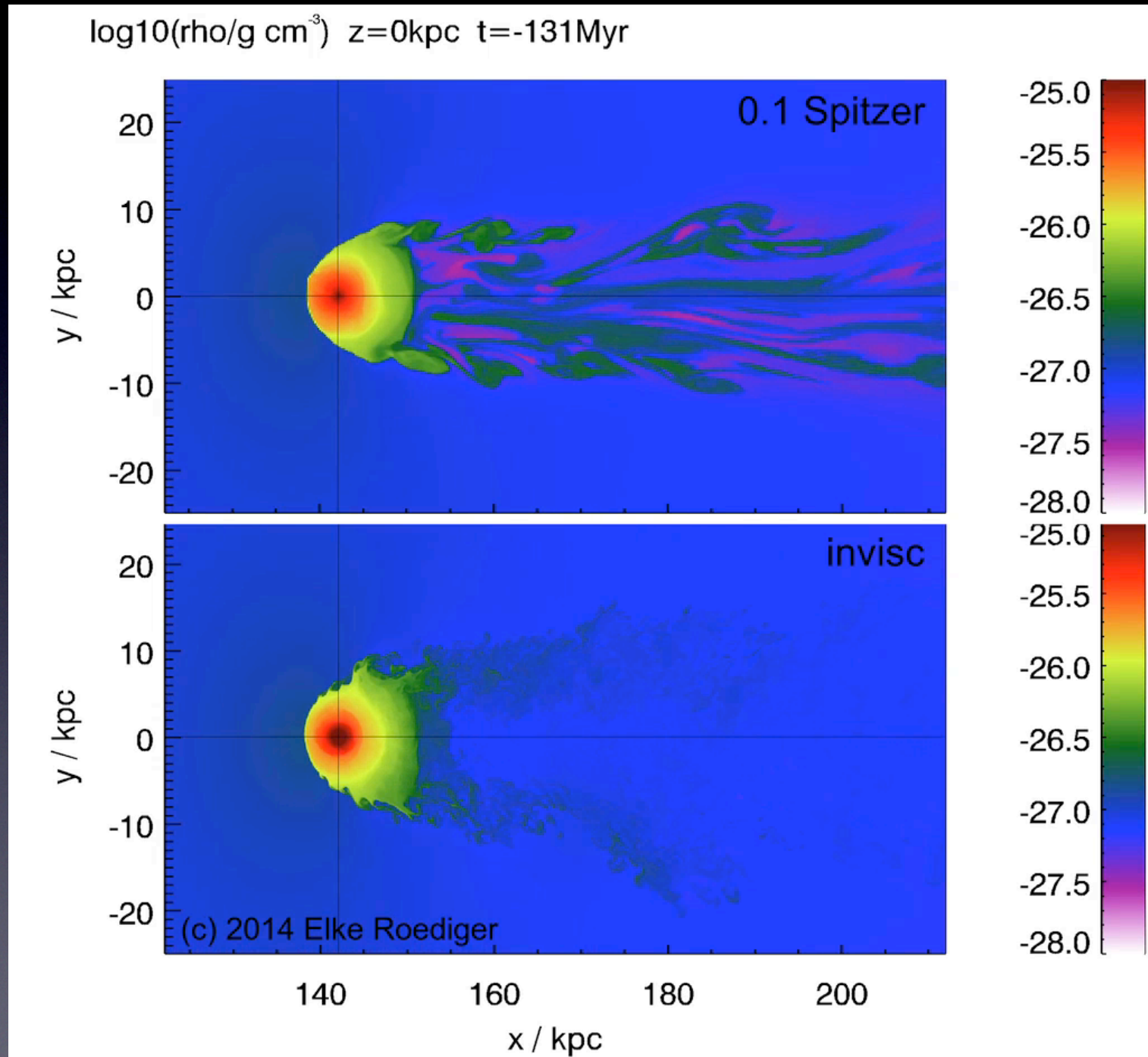
density slices

$\log_{10}(\rho/\text{g cm}^{-3})$ $z=0\text{kpc}$ $t=-185\text{Myr}$



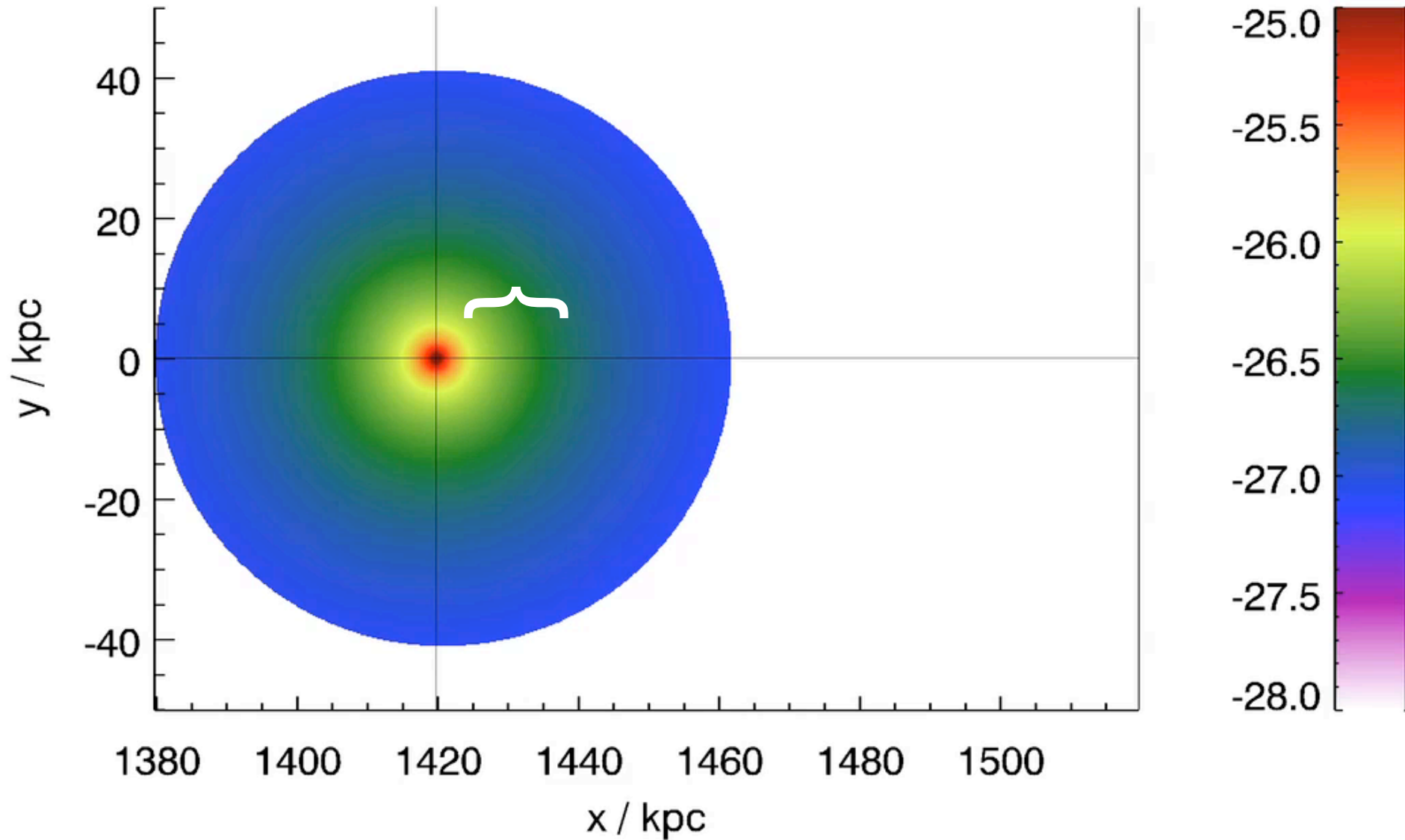
inviscid and viscous galaxy stripping

density slices



stripping of an extended atmosphere

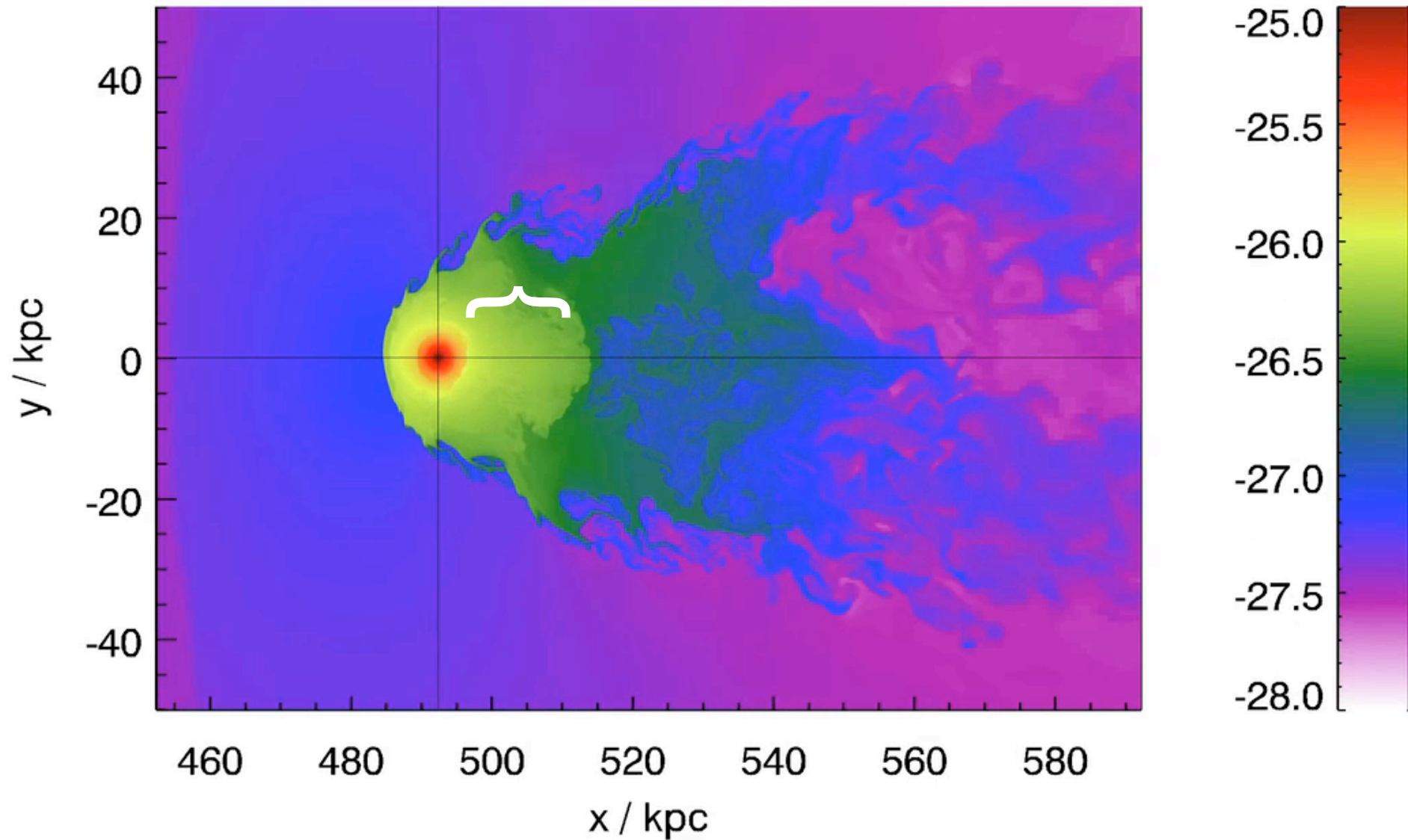
$\log_{10}(\rho/\text{g cm}^{-3})$ $z=0\text{kpc}$ $t=-1500\text{Myr}$



density slices

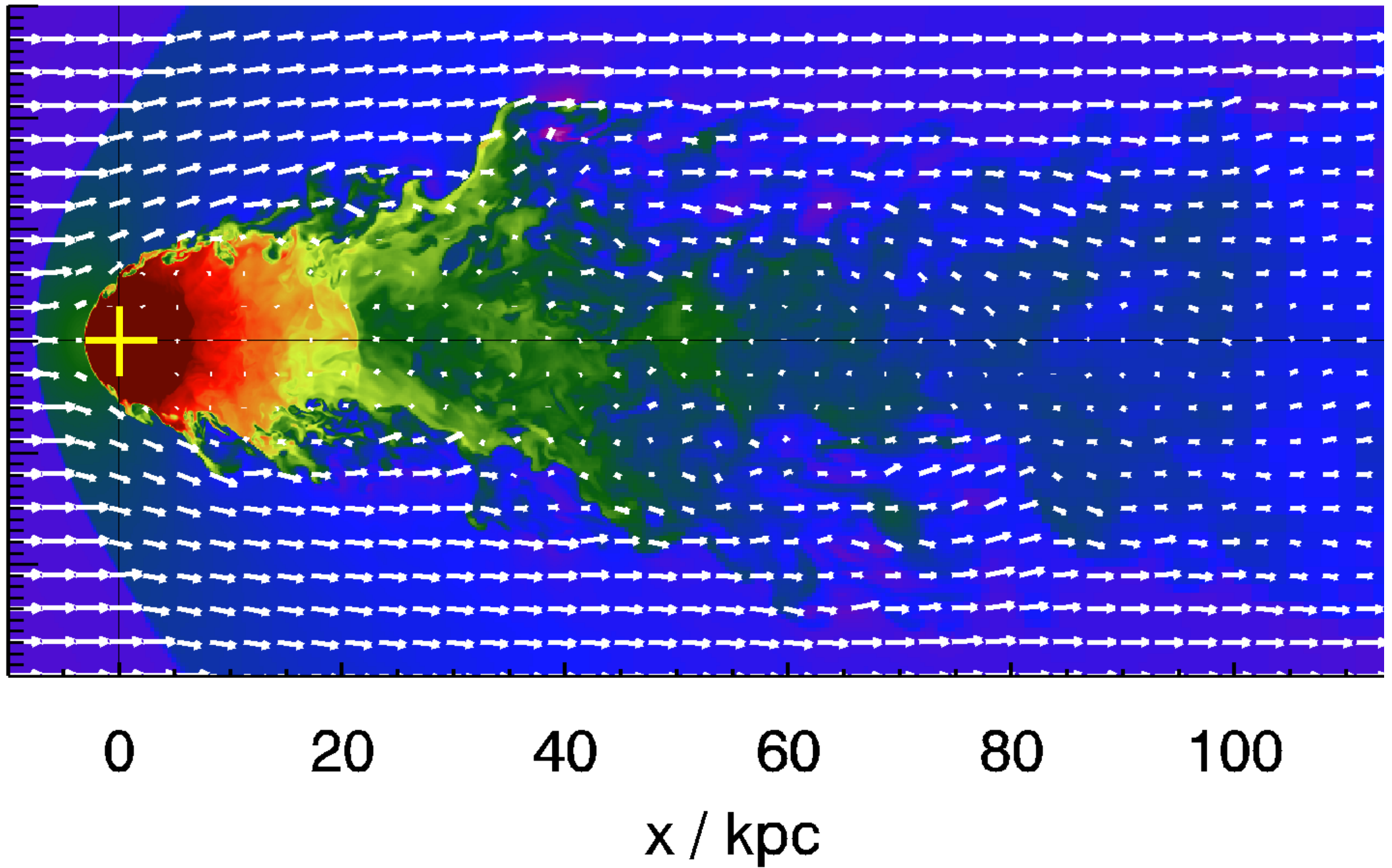
stripping of an extended atmosphere

$\log_{10}(\rho/\text{g cm}^{-3})$ $z=0\text{kpc}$ $t=-385\text{Myr}$

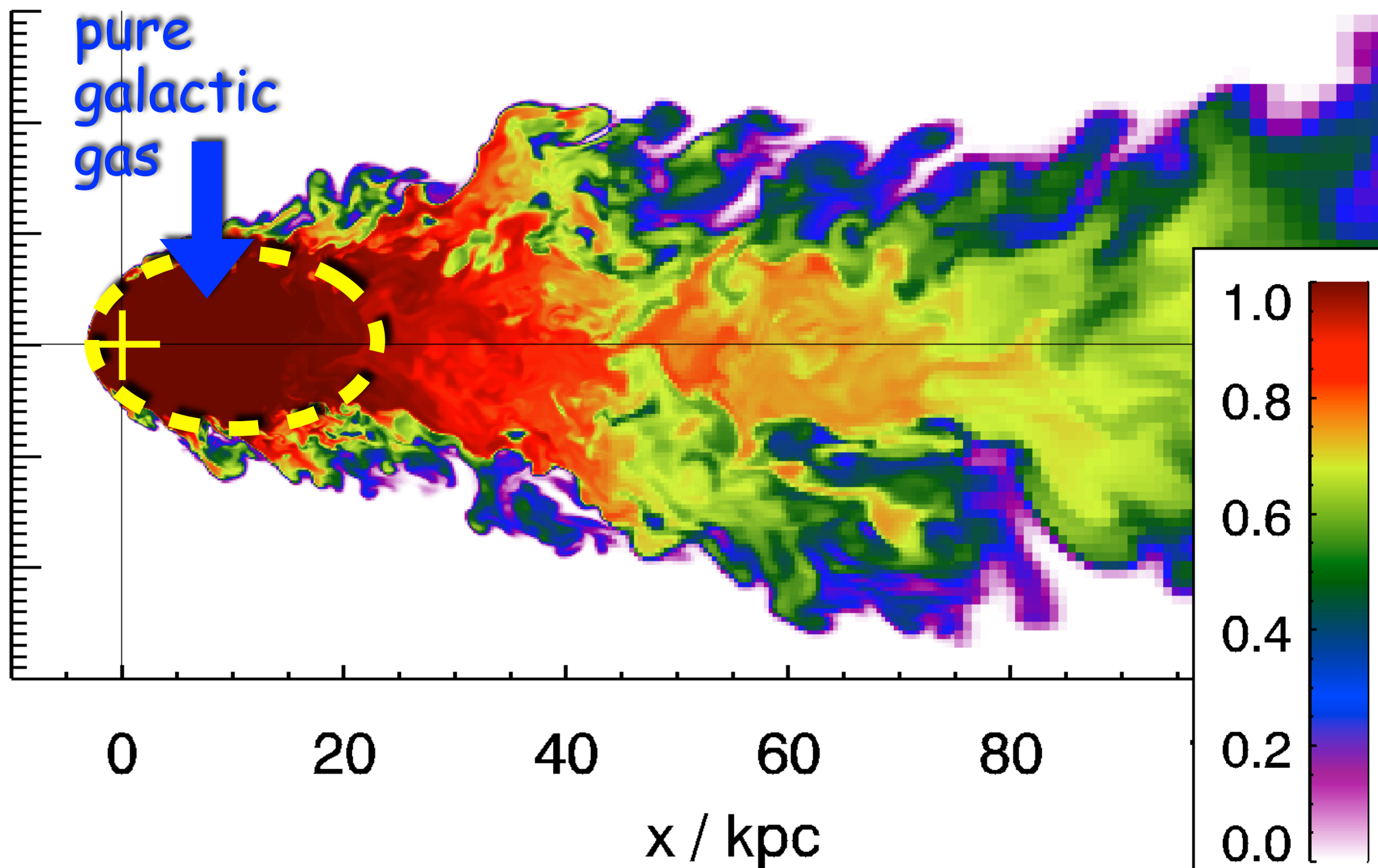


density slices

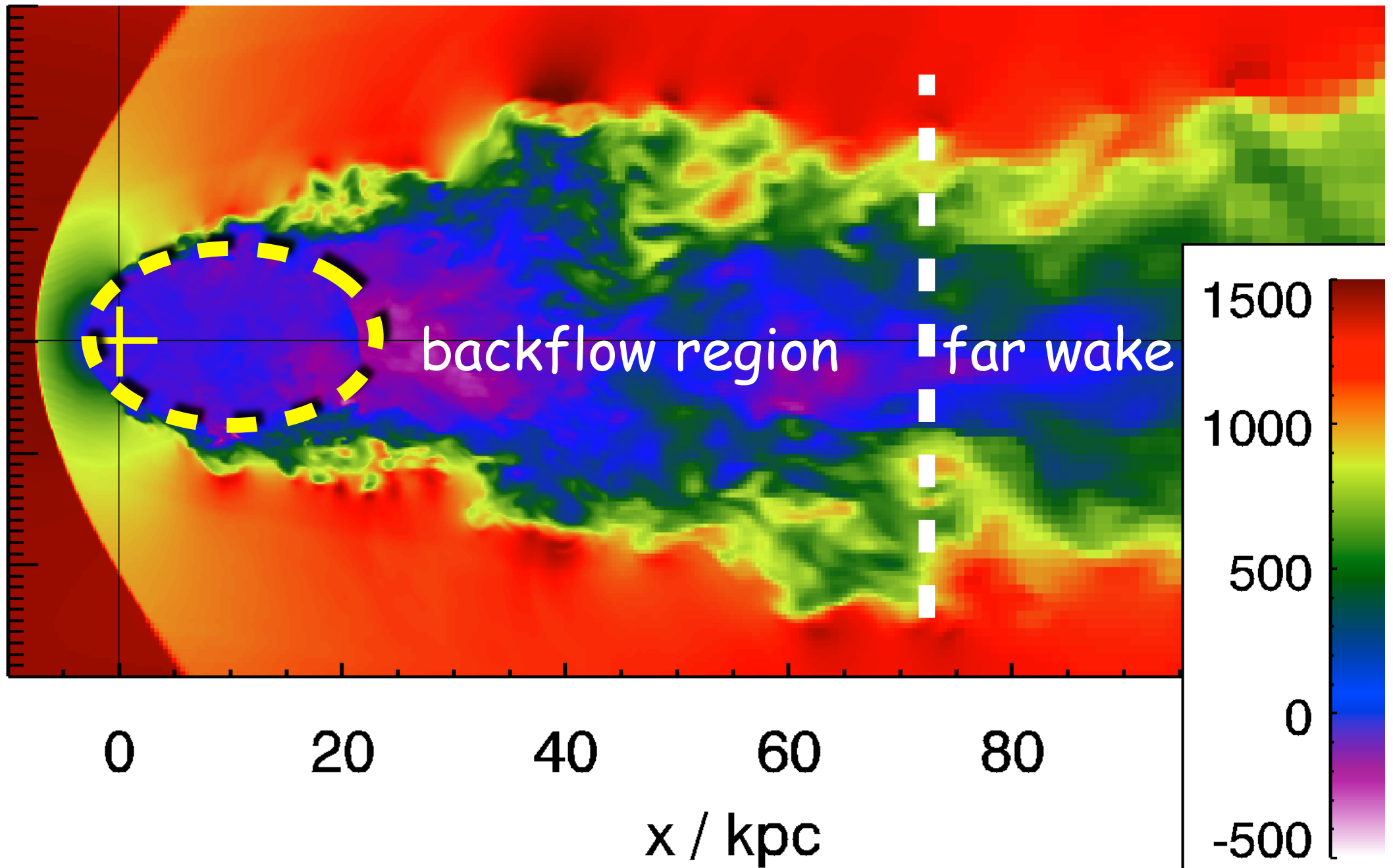
density



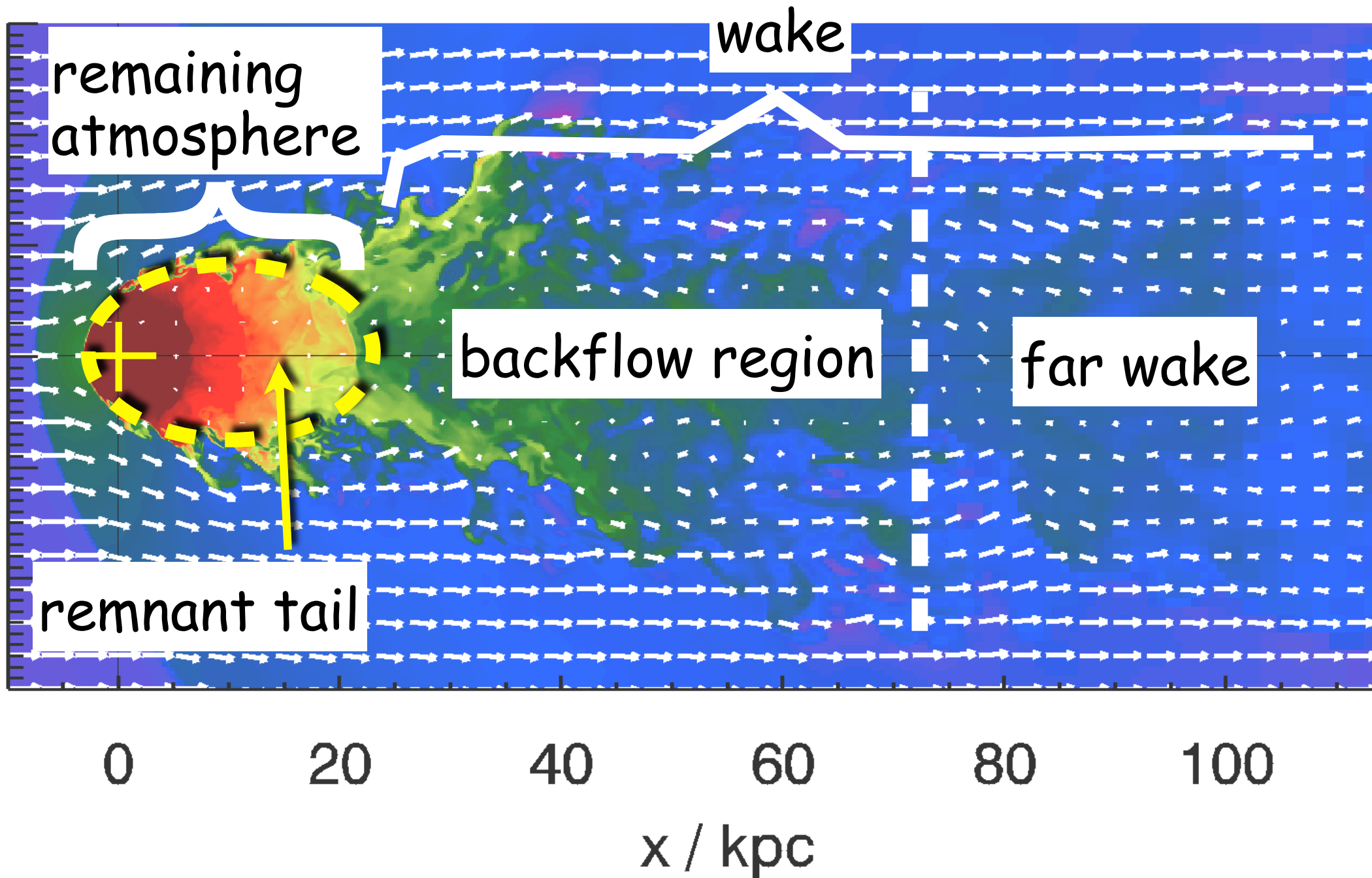
fraction of galactic gas

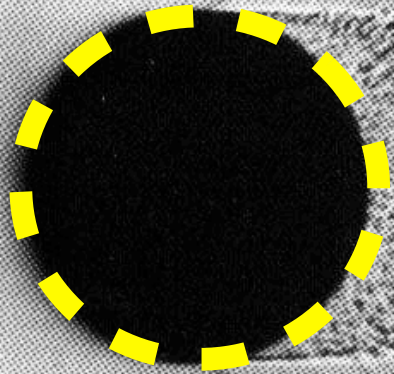


v_x in km/s



density



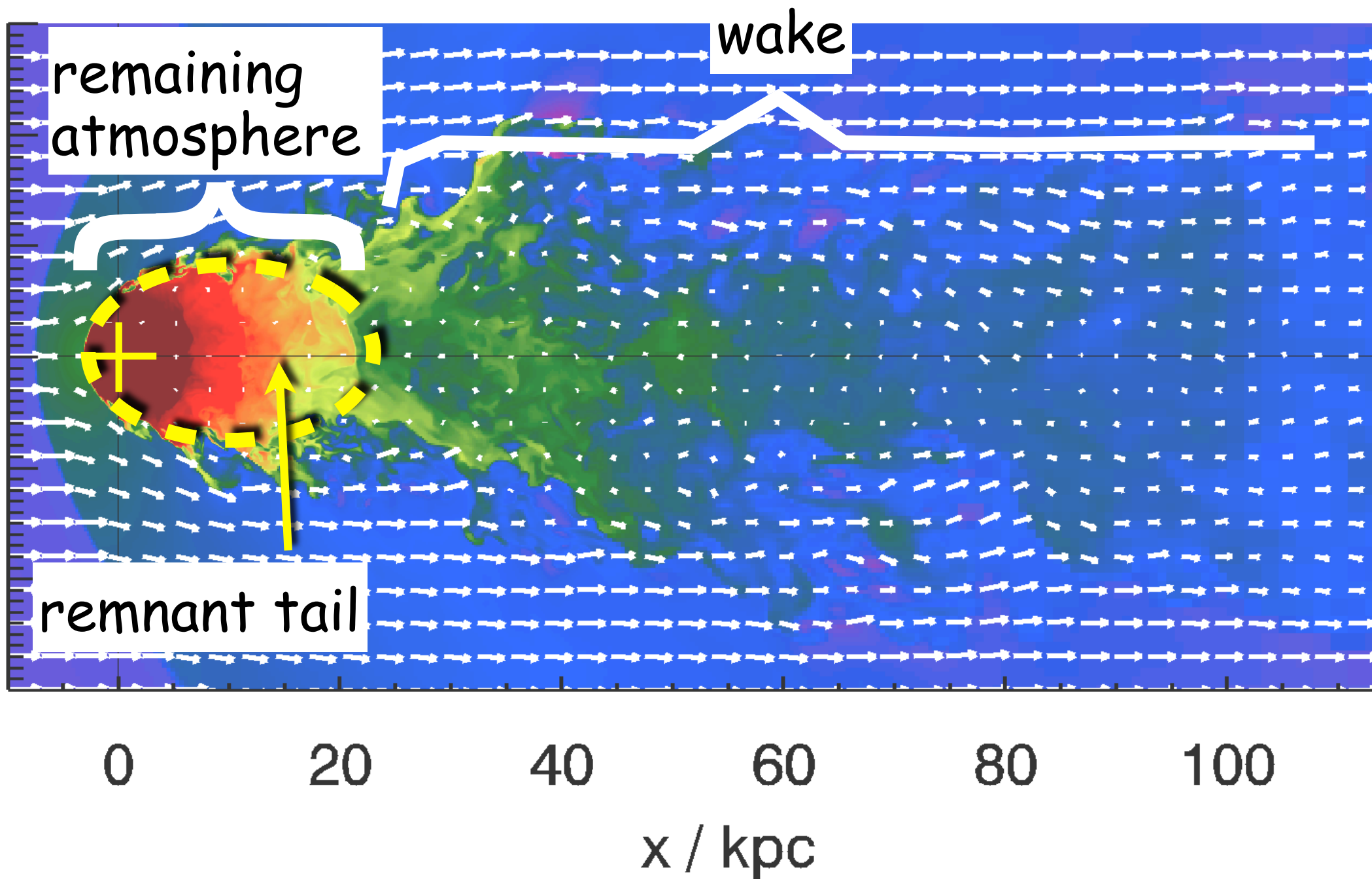


backflow
region

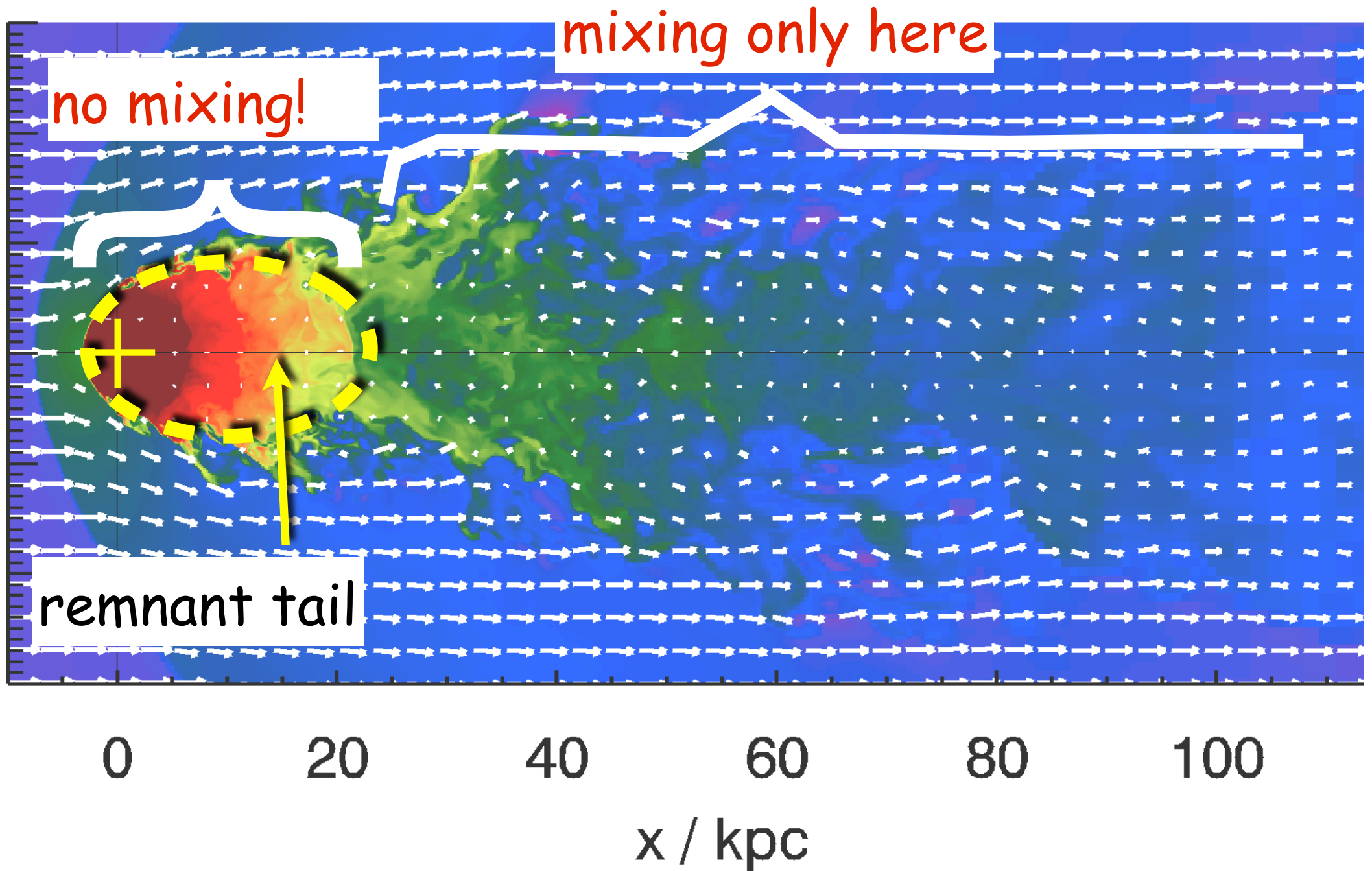


far wake

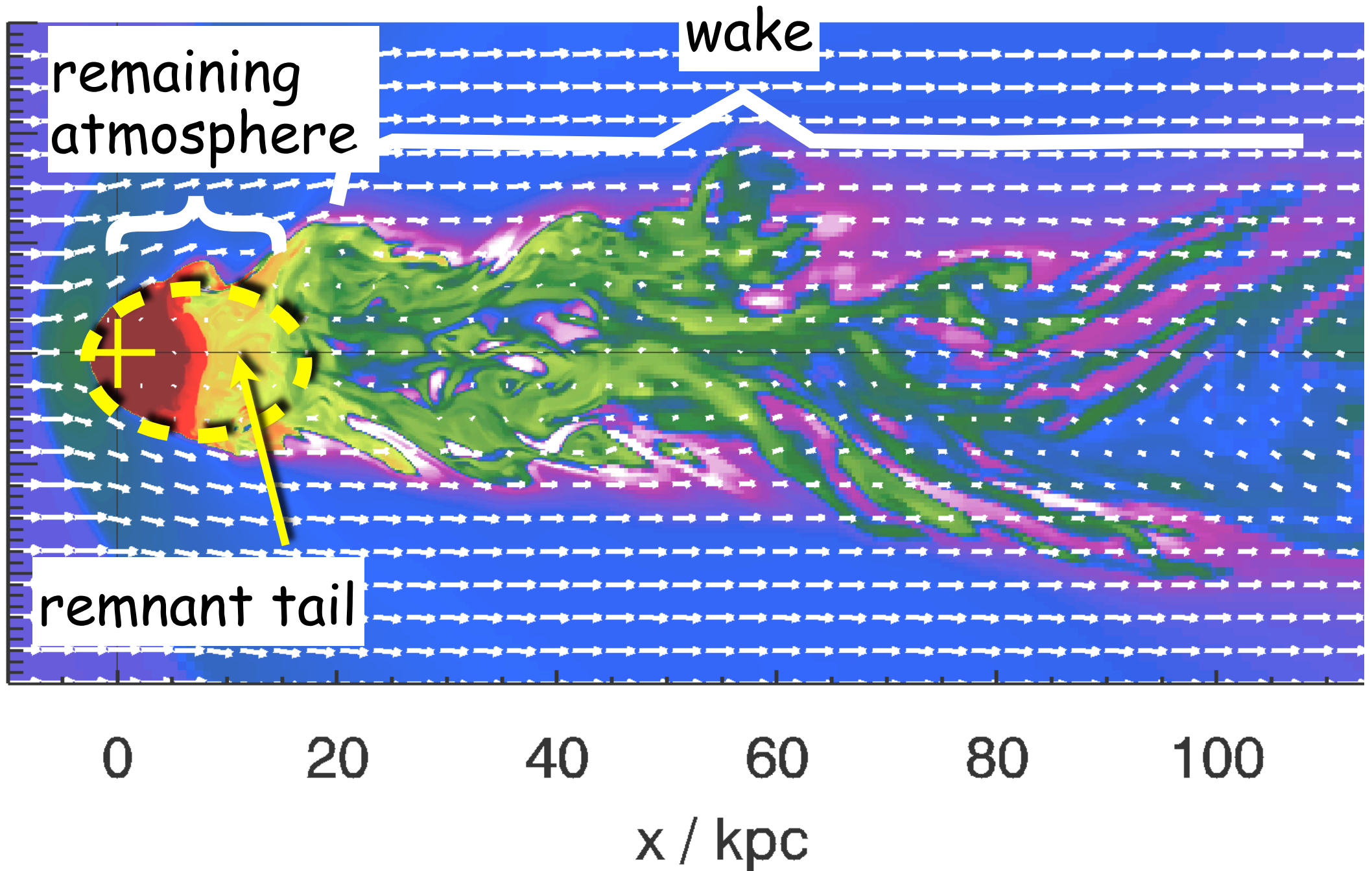
density



density

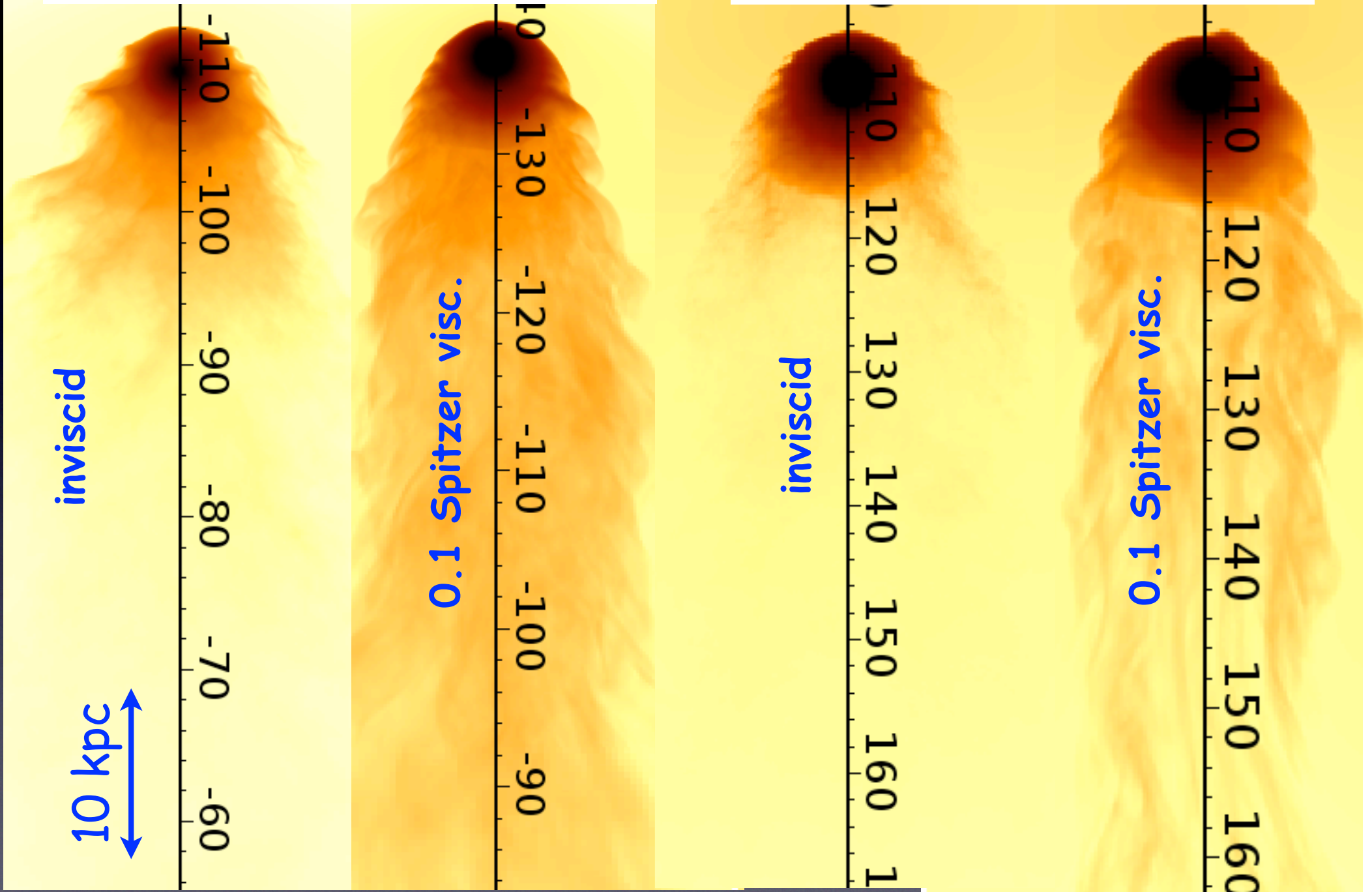


density - 0.1 Spitzer viscosity



extended atmosphere

compact atmosphere



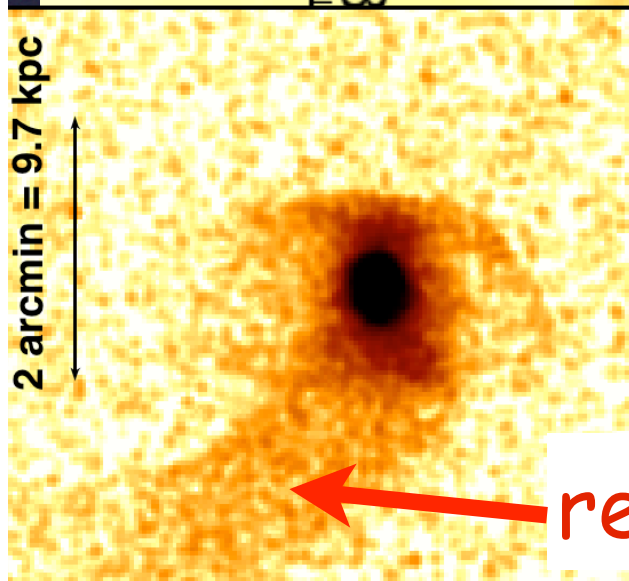
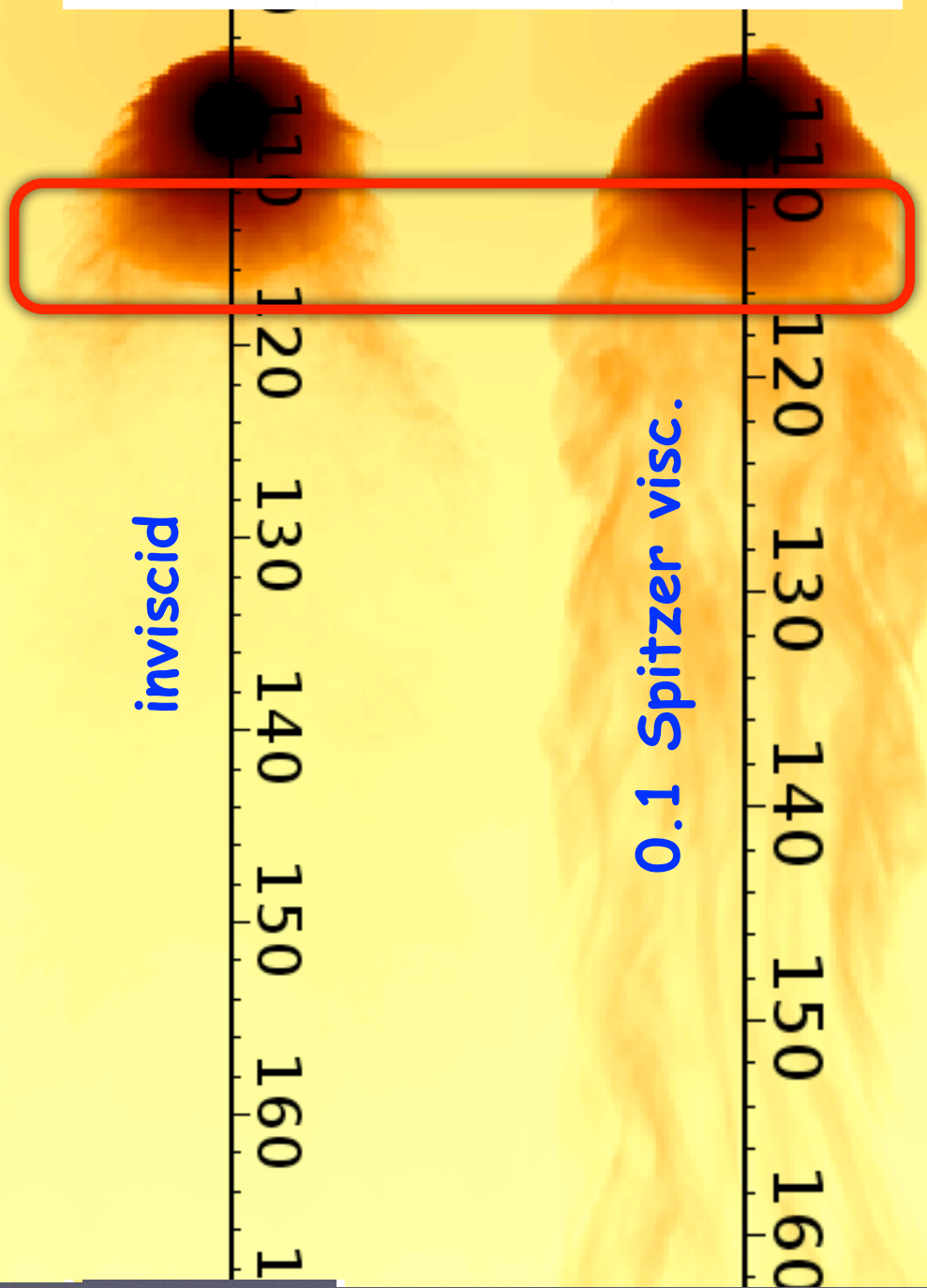
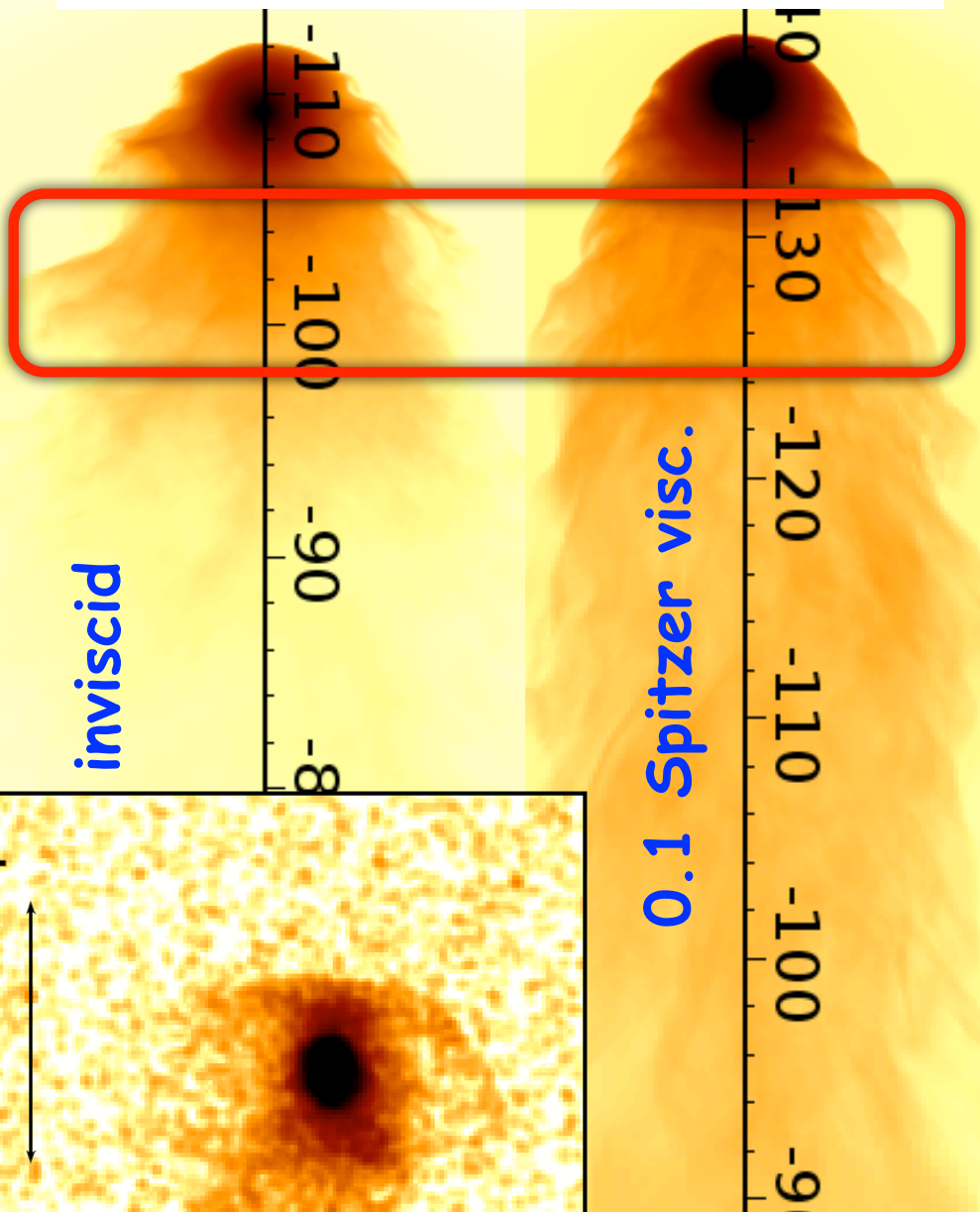
Mock X-ray (0.7-1.1keV)

Elke Roediger



extended atmosphere

compact atmosphere

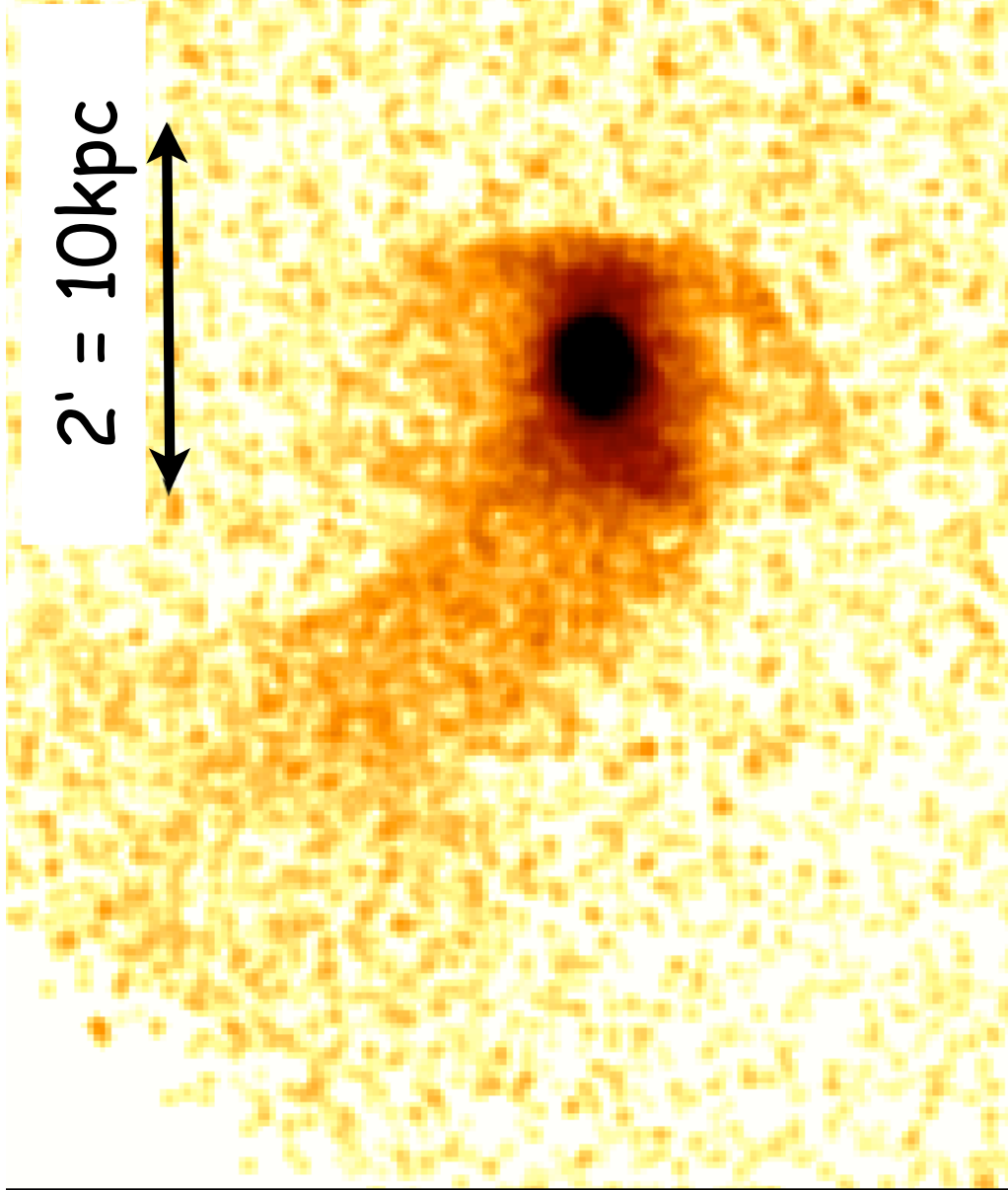


remnant tail

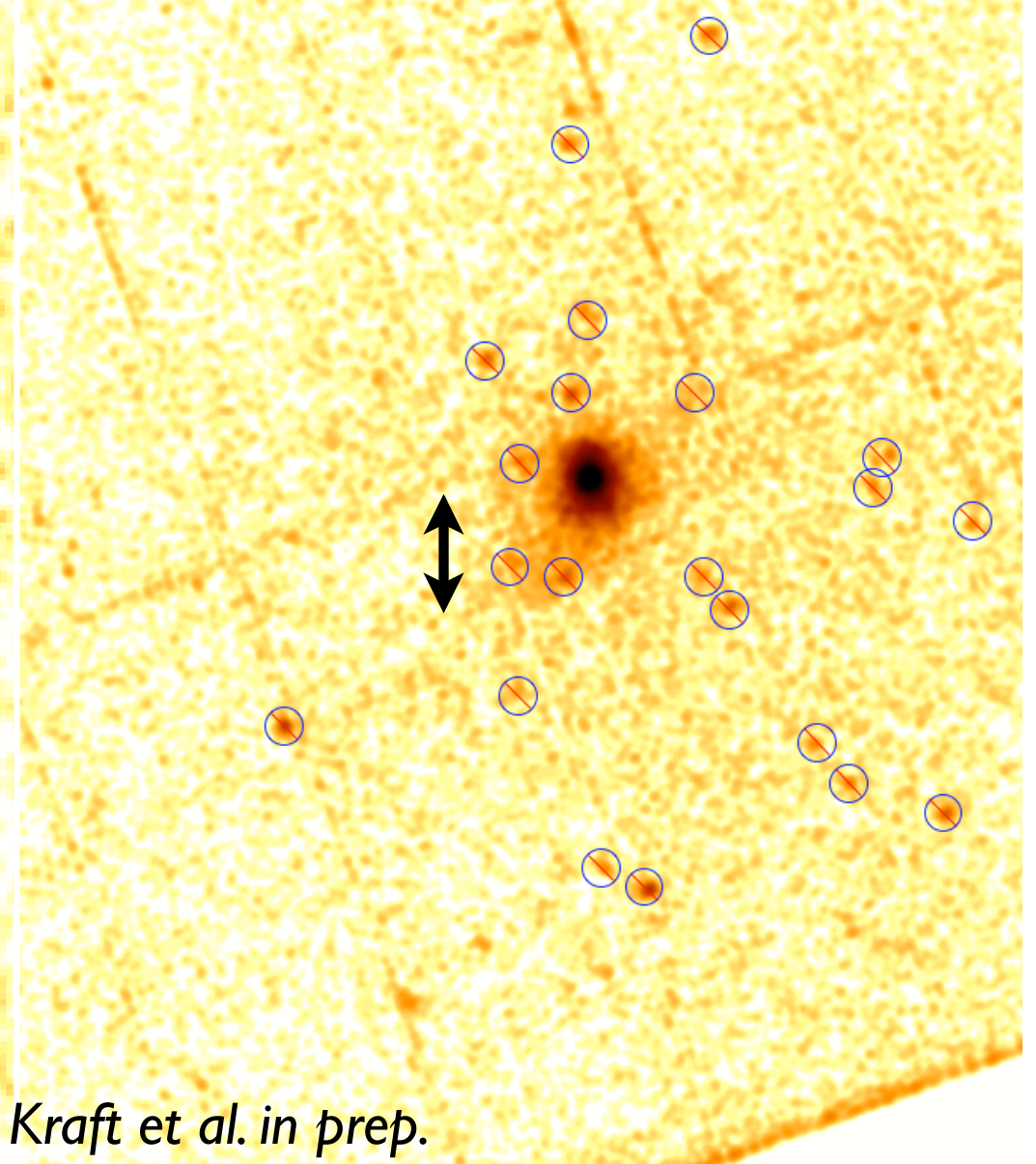
Mock X-ray (0.7-1.1keV)

Chandra ACIS-S
0.7-1.1 keV (200 ks)

$2' = 10\text{kpc}$



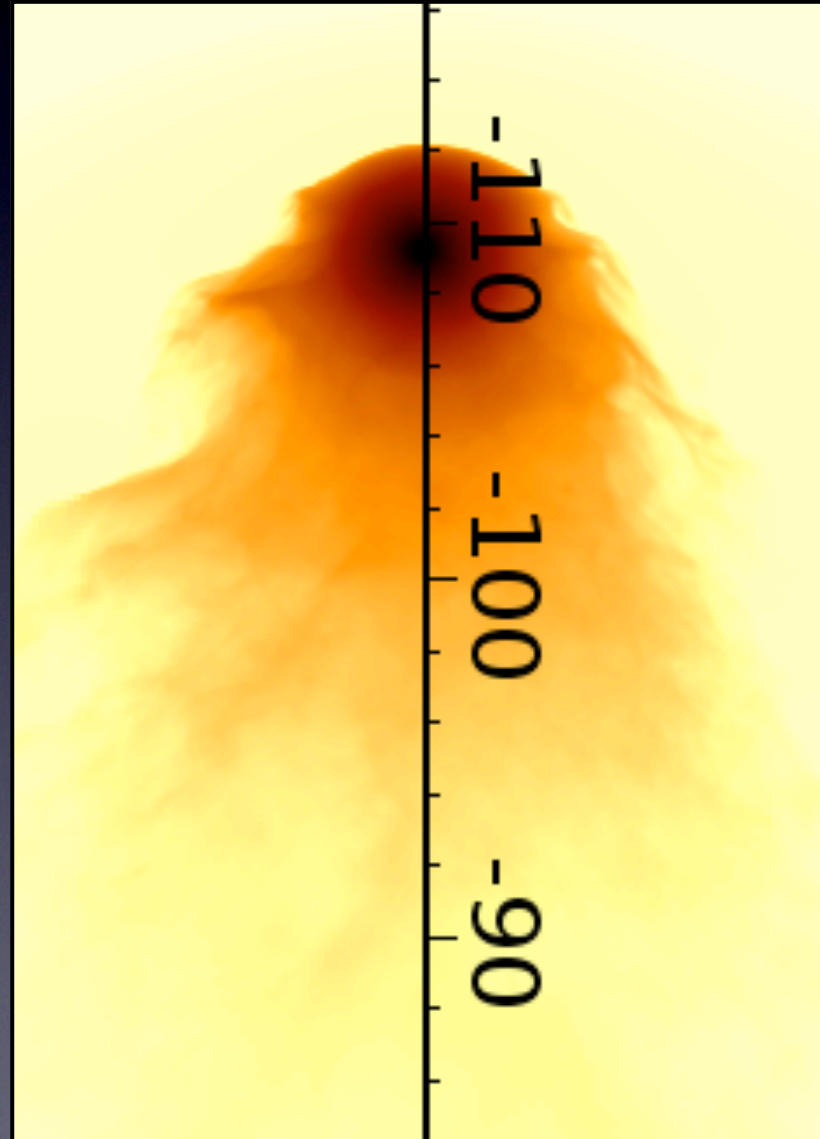
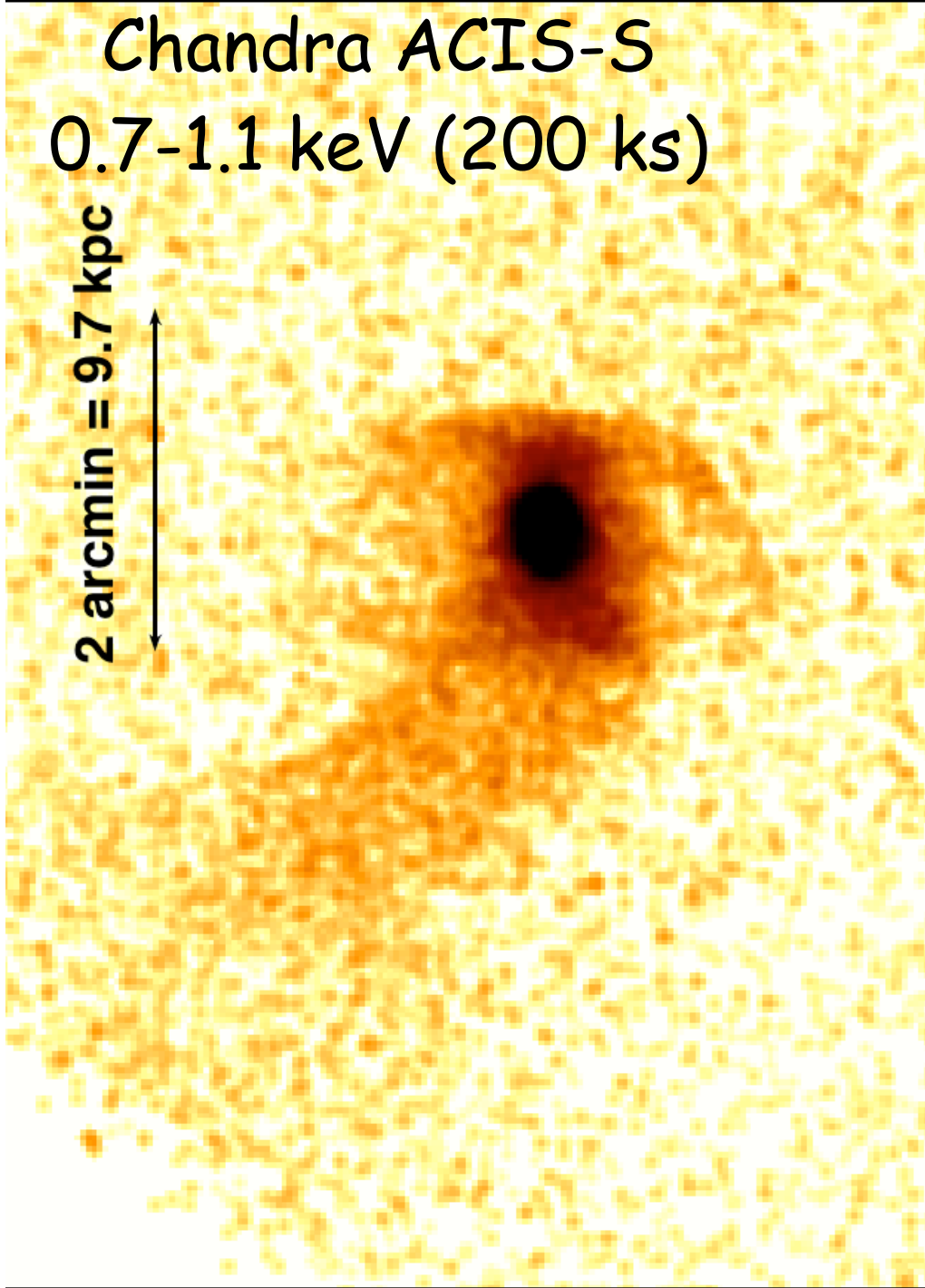
XMM pn
0.5-1.2 keV (18 ks clean)



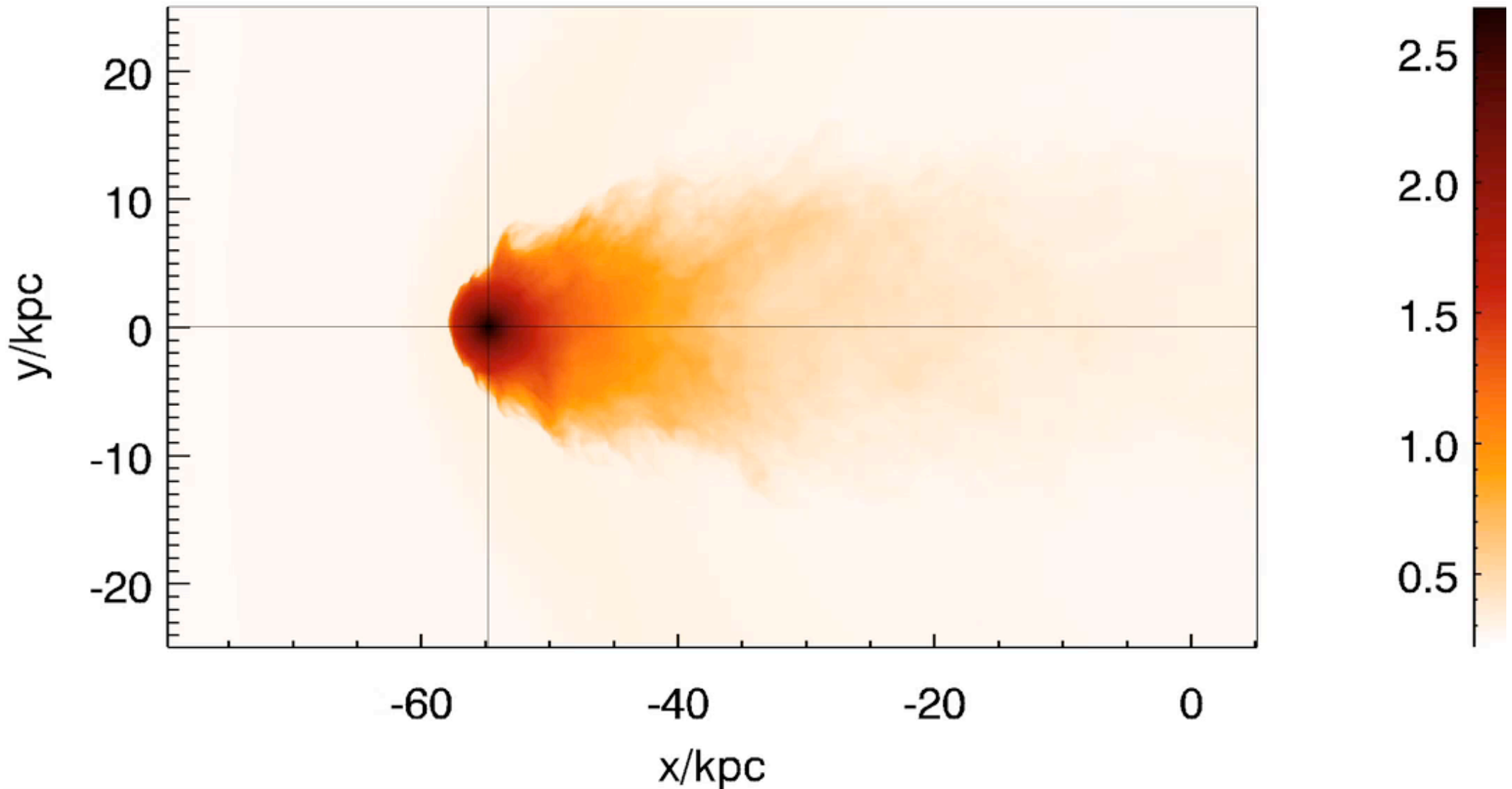
Kraft et al. in prep.

Chandra ACIS-S
0.7-1.1 keV (200 ks)

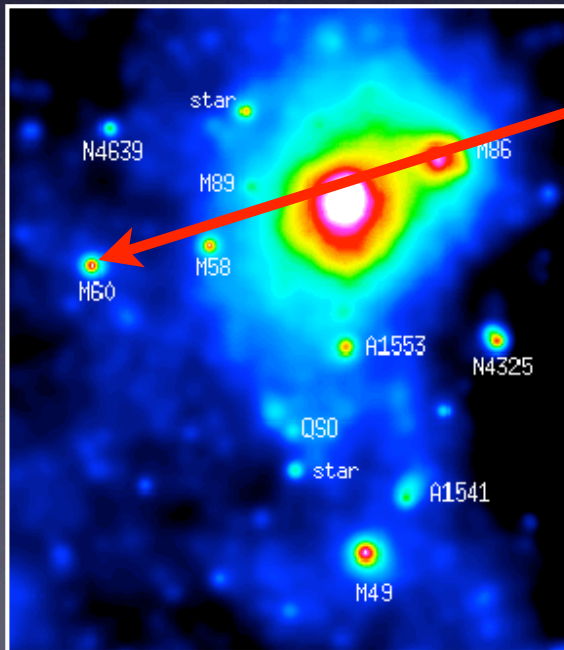
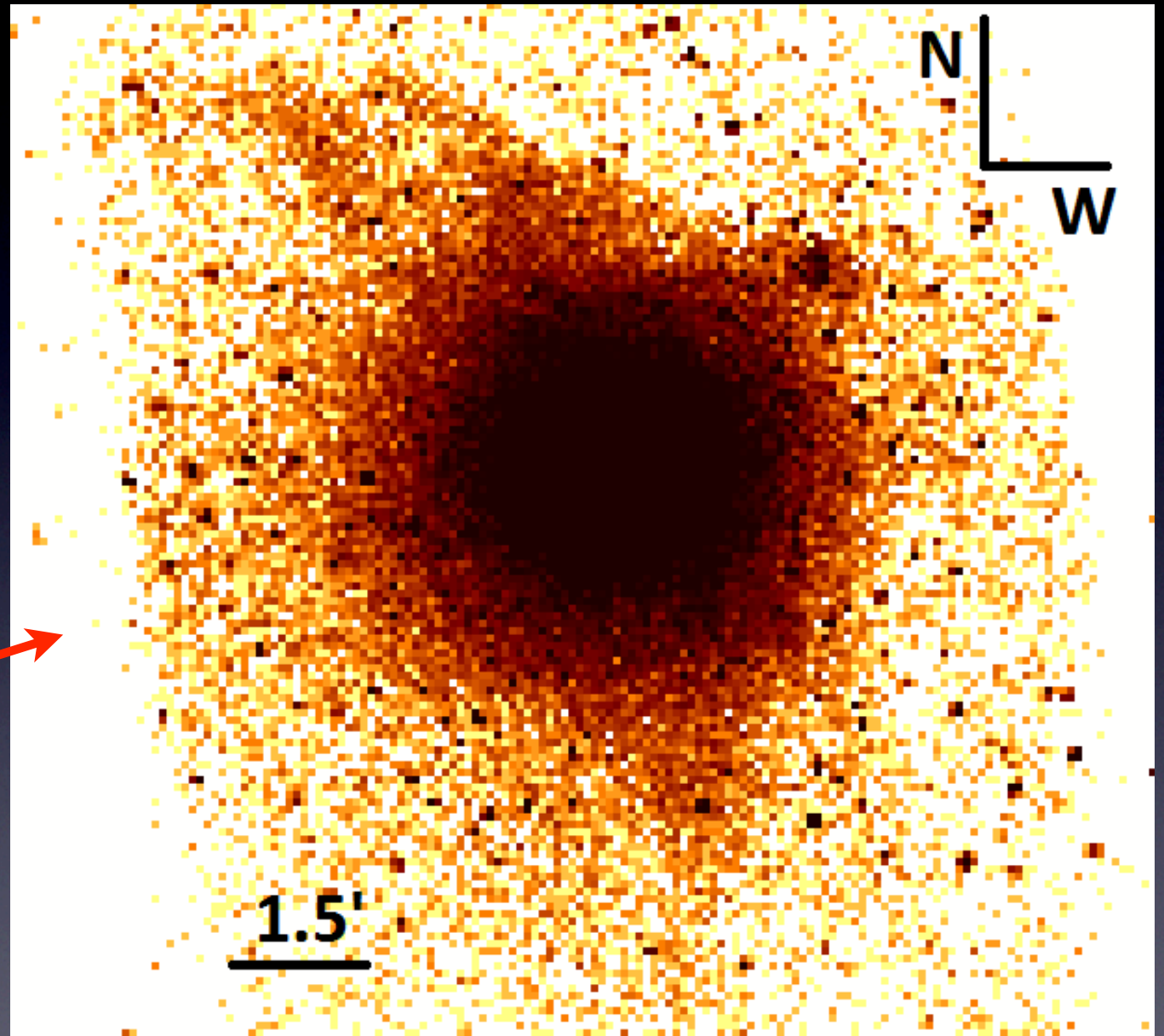
2 arcmin = 9.7 kpc



0.7-1.1keV band; $\log \text{LOS/kpc} = (-24, 24) + \text{obs. cluster BG}$ $i=0$ $t=40\text{Myr}$

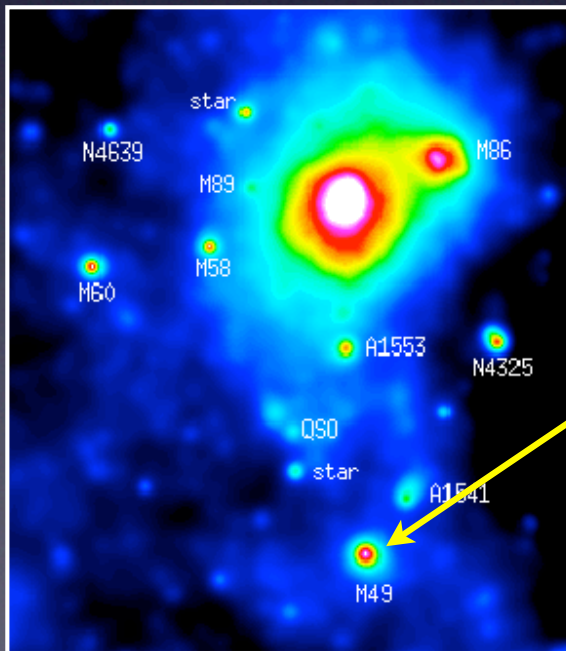
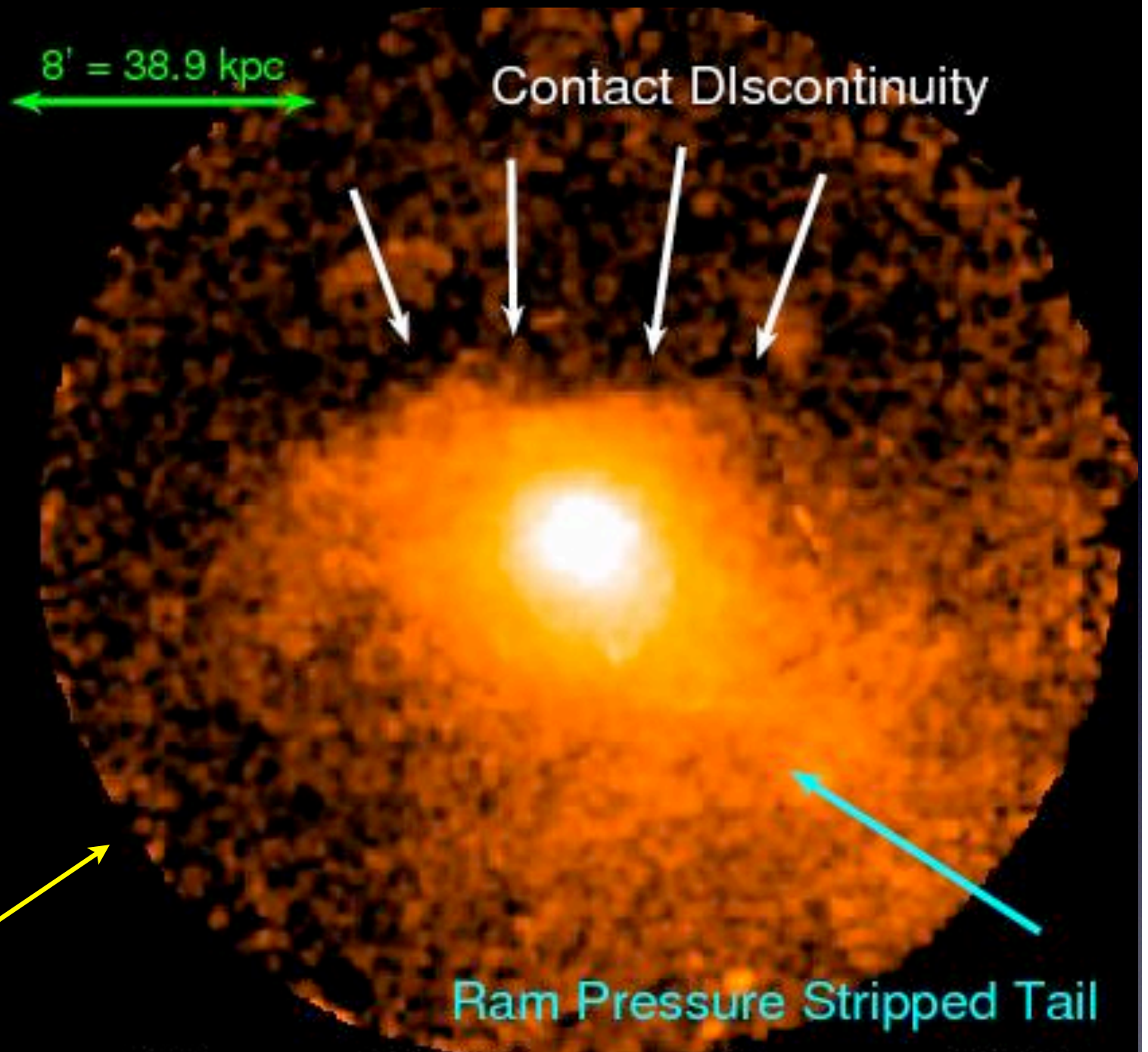


NGC 4649 (M60)



Chandra mosaic (Wood+ 2014, submitted)

NGC 4472 (M49)



XMM-Newton image (0.5-2.0 keV) (Kraft+ 2011)

Chandra
mosaic,
0.5–1.5 keV

NGC 1399
(Fornax center)

upstream edge

NGC 1404

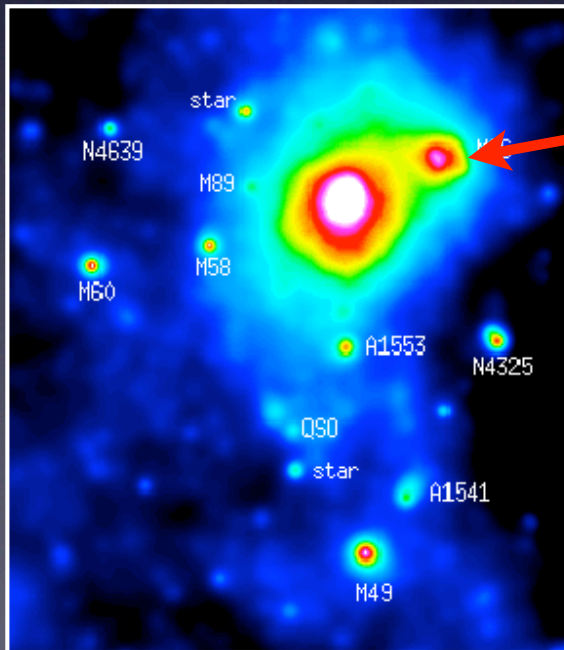
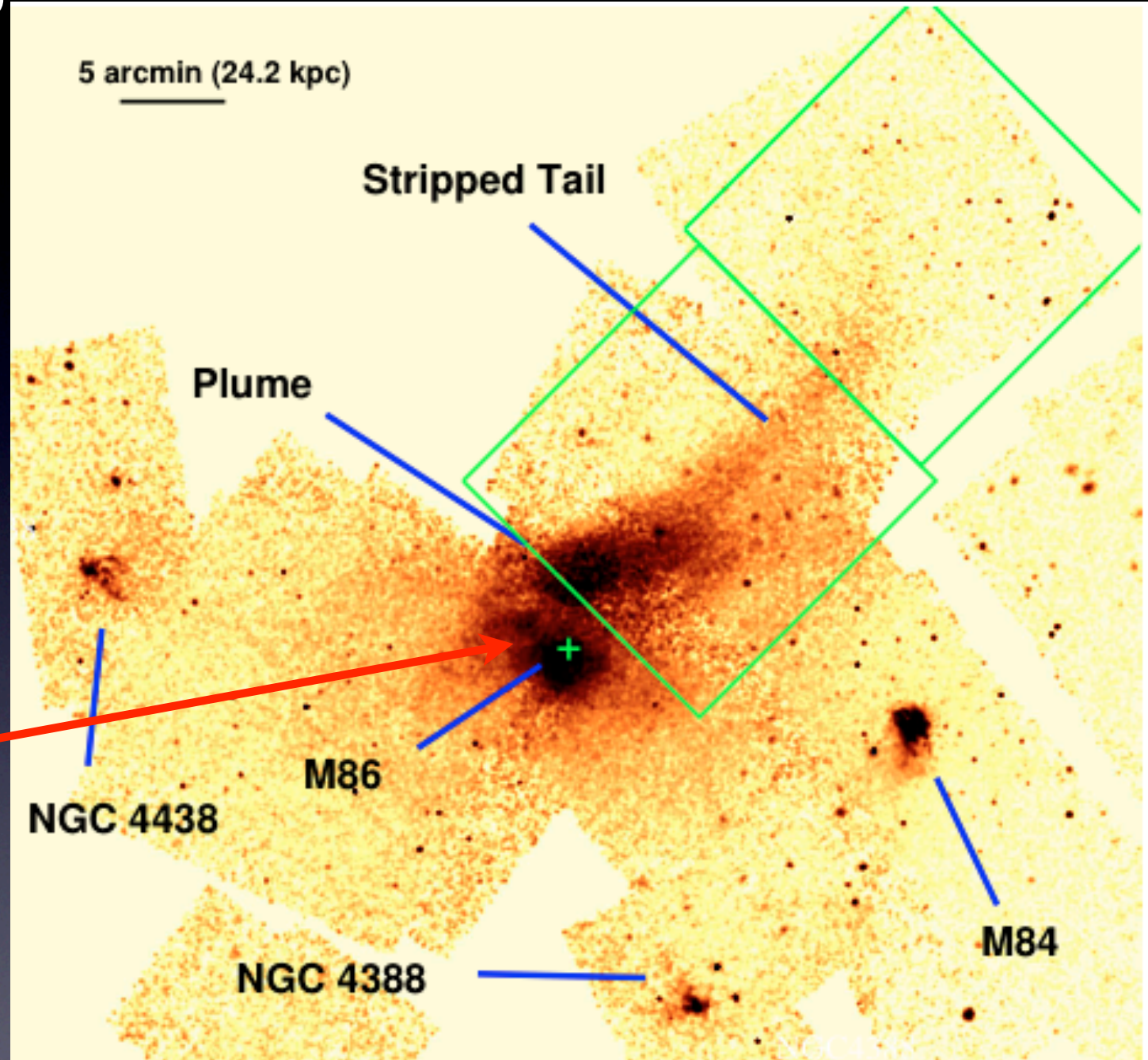
tail

1.32' = 7.7 kpc

9.8' = 57 kpc

NGC 1404
in Fornax
Machacek+ 2005

NGC 4406 (M86)



Chandra mosaic (Randall+ 2008)

N4438

SDSS, optical

CHANDRA,

Randall et al. 2008

H α ,

Kenney et al. 2008

4'

20 kpc

HI, Oosterloo et al. 2005

M86

N4388

Summary

cluster dynamics
establish opportunity
for ICM mixing

ICM properties allow/
suppress/prevent it



observables for *correct dynamics*

first evidence against substantial isotropic viscosity
in Virgo cluster

pilot studies done, more to come!

