

Gravity and Thermodynamics

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A new point of view in King function

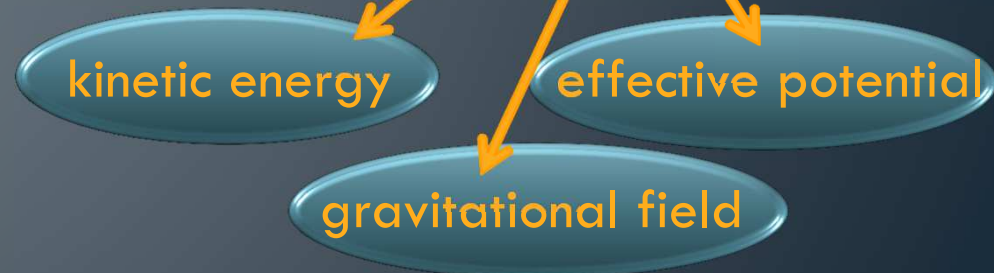
King distribution function for Globular Clusters

$$f(\varepsilon) = \begin{cases} A \left[e^{-\frac{\varepsilon}{kT}} - e^{-\frac{\psi}{kT}} \right] & \varepsilon \leq \psi \\ 0 & \varepsilon > \psi \end{cases}$$



$$f(\varepsilon) = B e^{-H/k\theta}$$

$$H = \varepsilon + m\varphi + \phi$$



It is possible to include the competing effects of stellar encounters and evaporation of stars in a Boltzmann-like distribution function thanks to an effective potential ϕ , which reduces the phase space accessible to particles

Kinetic and thermodynamic variables

Kinetic variables defined using the kinetic part of the Hamiltonian

Thermodynamic variables defined using the full Hamiltonian

Intensive

Variable

$$T = R\theta$$

$$P = R\Pi$$

$$\langle \mu_0 - \varepsilon \rangle = R\langle \alpha_0 - H_0 \rangle$$

Extensive

Conjugate

$$S$$

$$V$$

$$N$$

Equation of state

$$\Pi V = Nk\theta$$

$$PV = NkT$$

First principle

$$dU = \theta dS - \Pi dV + \alpha dN + N\langle d\phi \rangle$$

$$dU_k = TdS - PdV + \langle \mu_0 \rangle dN + N(d\langle \mu_0 \rangle - \langle d\mu_0 \rangle)$$

Gibbs-Duhem relation

$$N\langle d\phi \rangle = Sd\theta - Vd\Pi + Nd\alpha$$

$$N(d\langle \mu_0 \rangle - \langle d\mu_0 \rangle) = SdT - VdP + Nd\langle \mu_0 \rangle$$

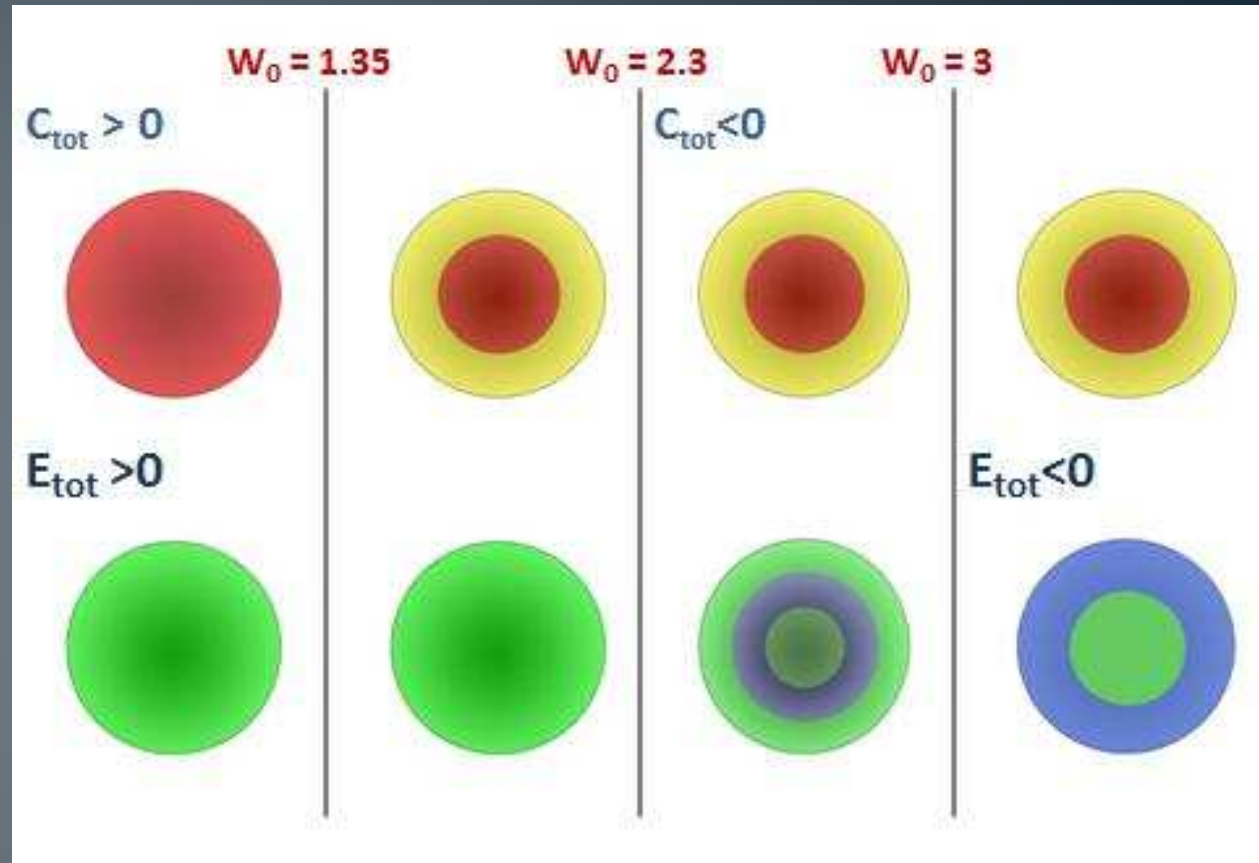
Euler equation

$$U = \theta S - \Pi V + \alpha N$$

$$U_k = TS - PV + \langle \mu_0 \rangle N$$

Configurations and evolution

- $W_0 > 1.35$
regions with
negative specific
heat
- $W_0 > 2.3$
intermediate
regions with
negative energy
- $W_0 > 3$
negative total
energy



- $W_0 < 1.35$ GCs don't evolve towards **gravothermal catastrophe**
- $W_0 < 3$ GCs evolve towards **disruption**