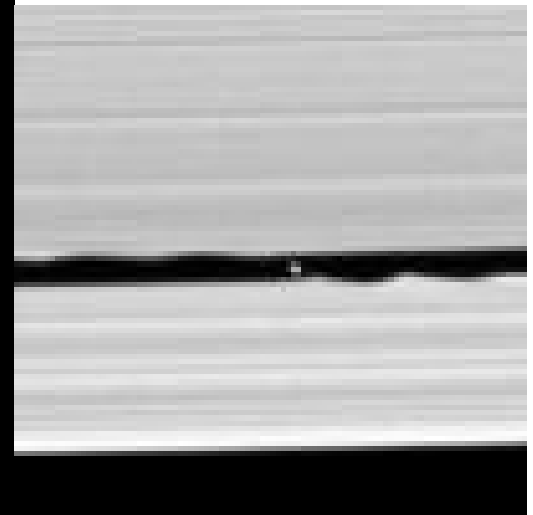
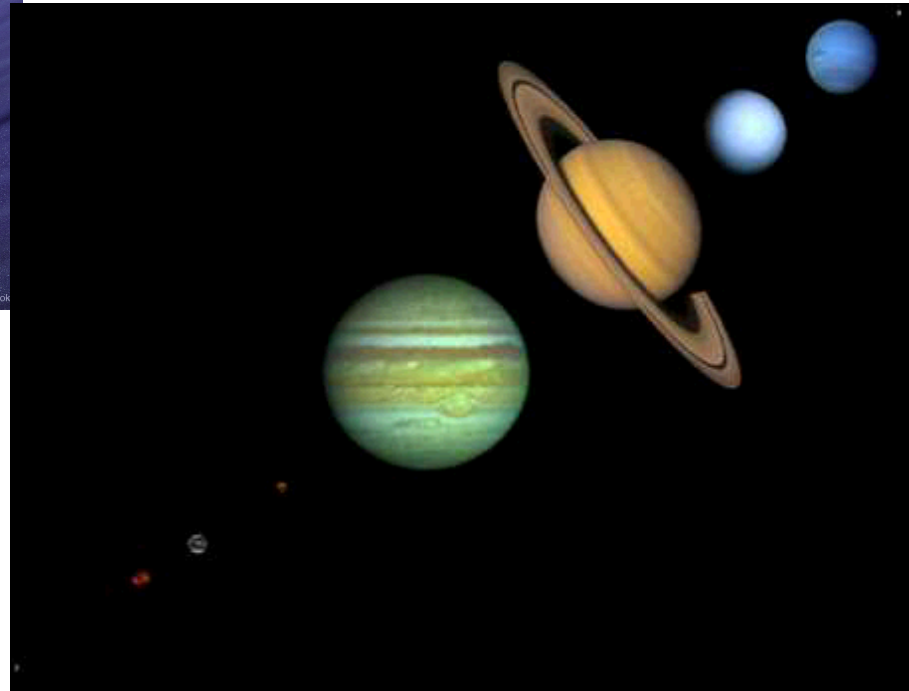
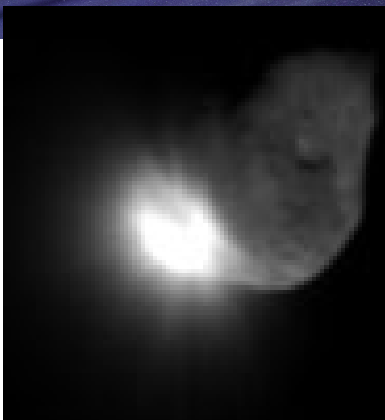
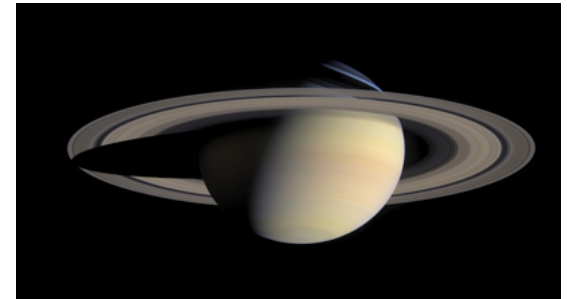
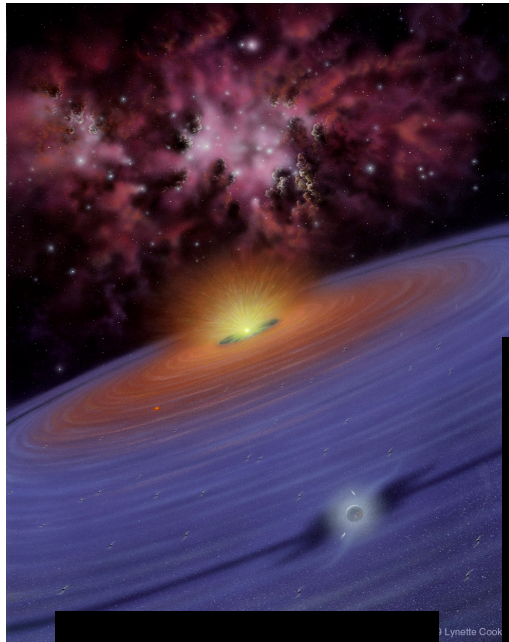


Summer School on Planet Formation

Niels Bohr Institute; Copenhagen, Denmark; 2015 August

Planets & Planetary Systems

Jack Lissauer
NASA Ames



Planets & Planetary Systems

- **Our Solar System**
- **Extrasolar Planets**
- NASA's *Kepler Mission*
- Giant Planet Formation

Observable Planetary Properties

Orbit

Mass, distribution of mass

Size

Rotation rate & direction

Shape

Temperature

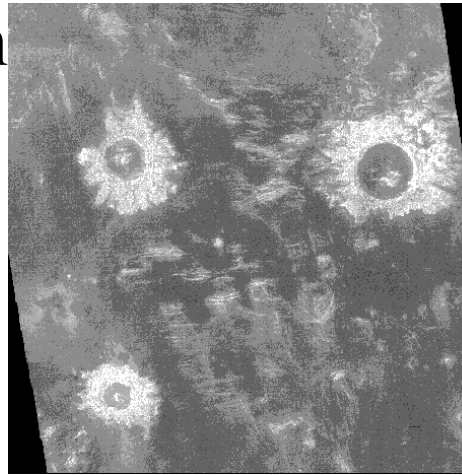
Magnetic field

Surface structure & composition

Atmospheric structure & composition



Phobos & Mars *Phobos 2*



Radar image of craters on Venus

Magellan



Jupiter w/ Europa's shadow *Cassini*

Solar System Planets

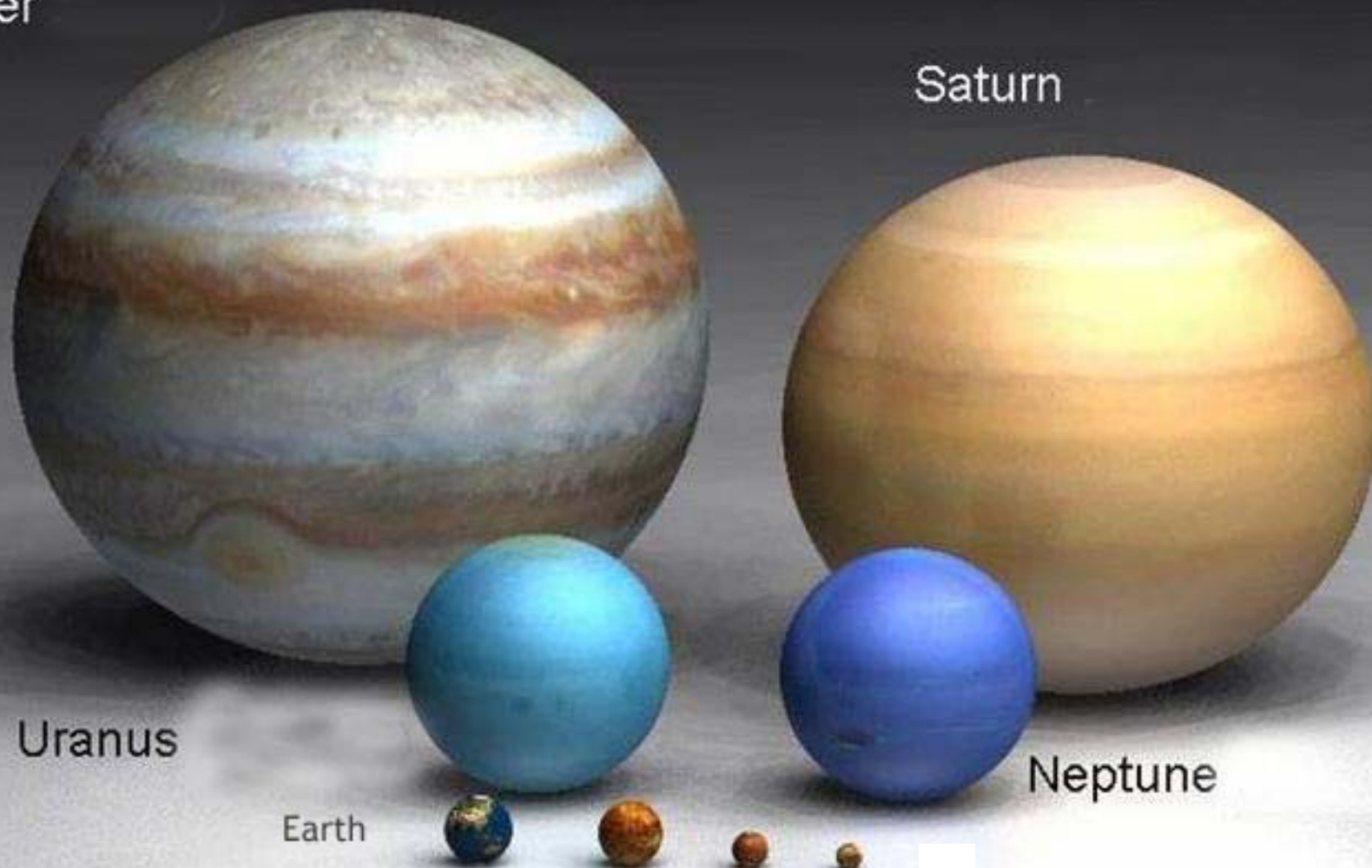
Jupiter

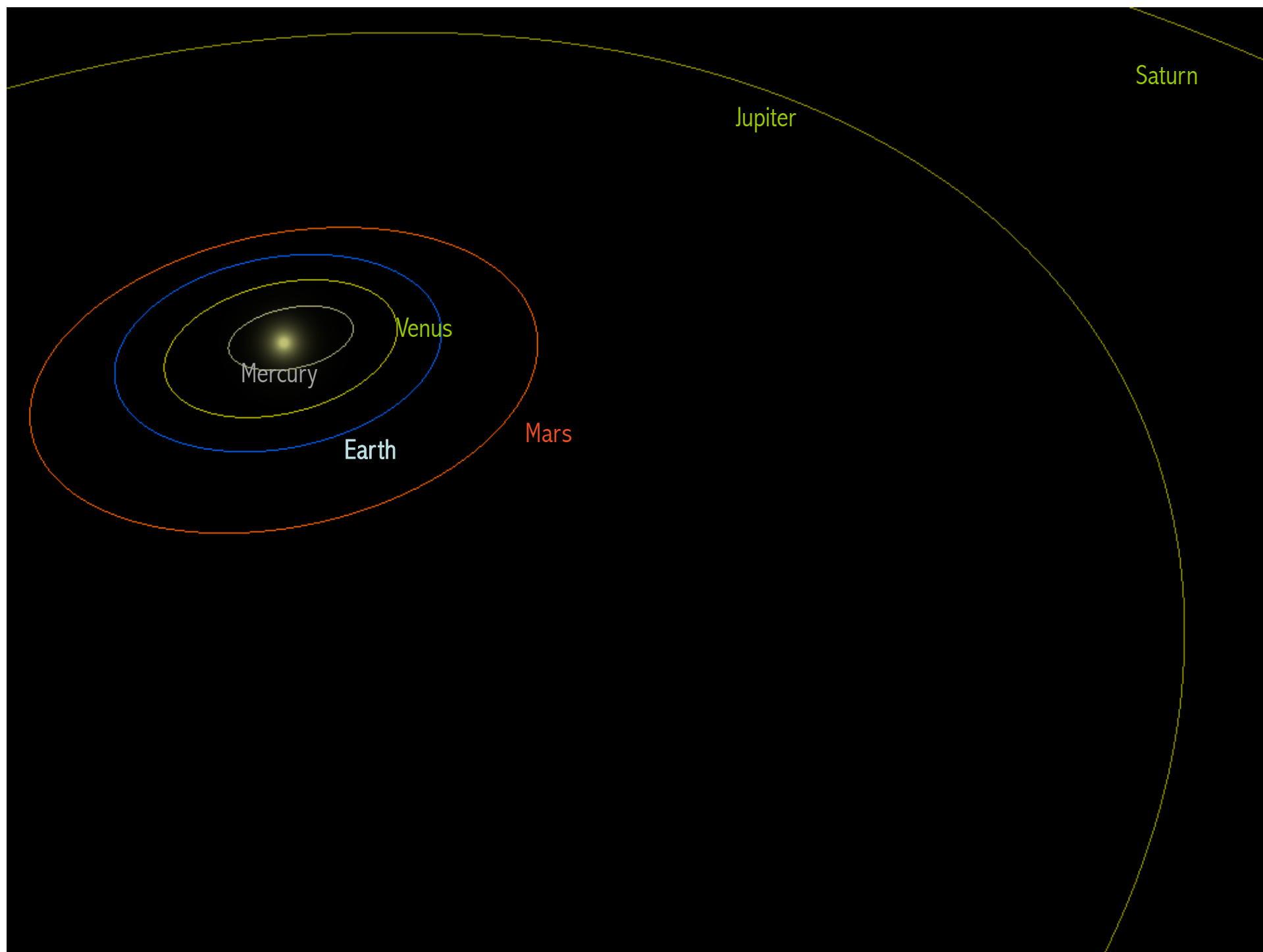
Saturn

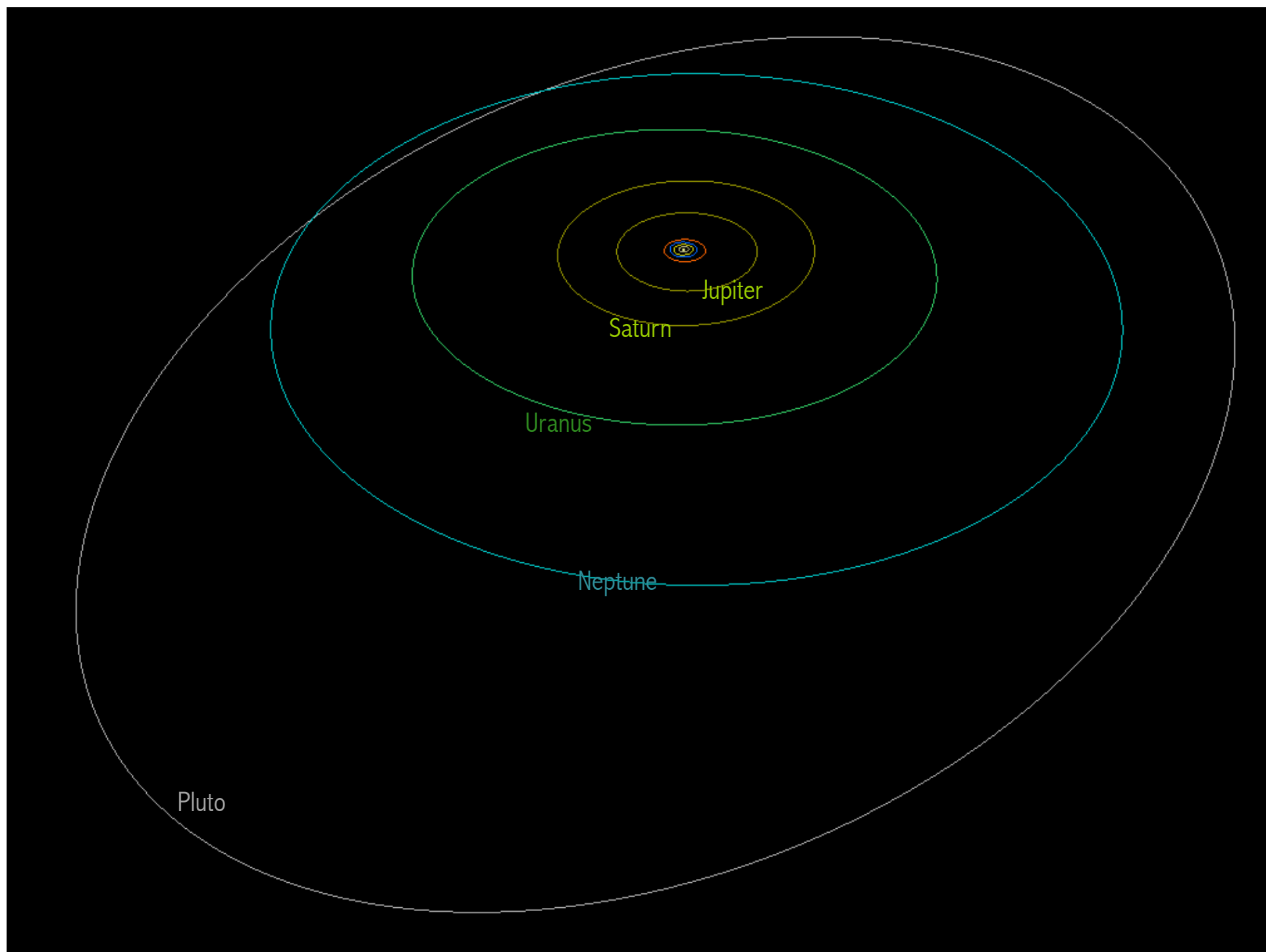
Uranus

Neptune

Earth







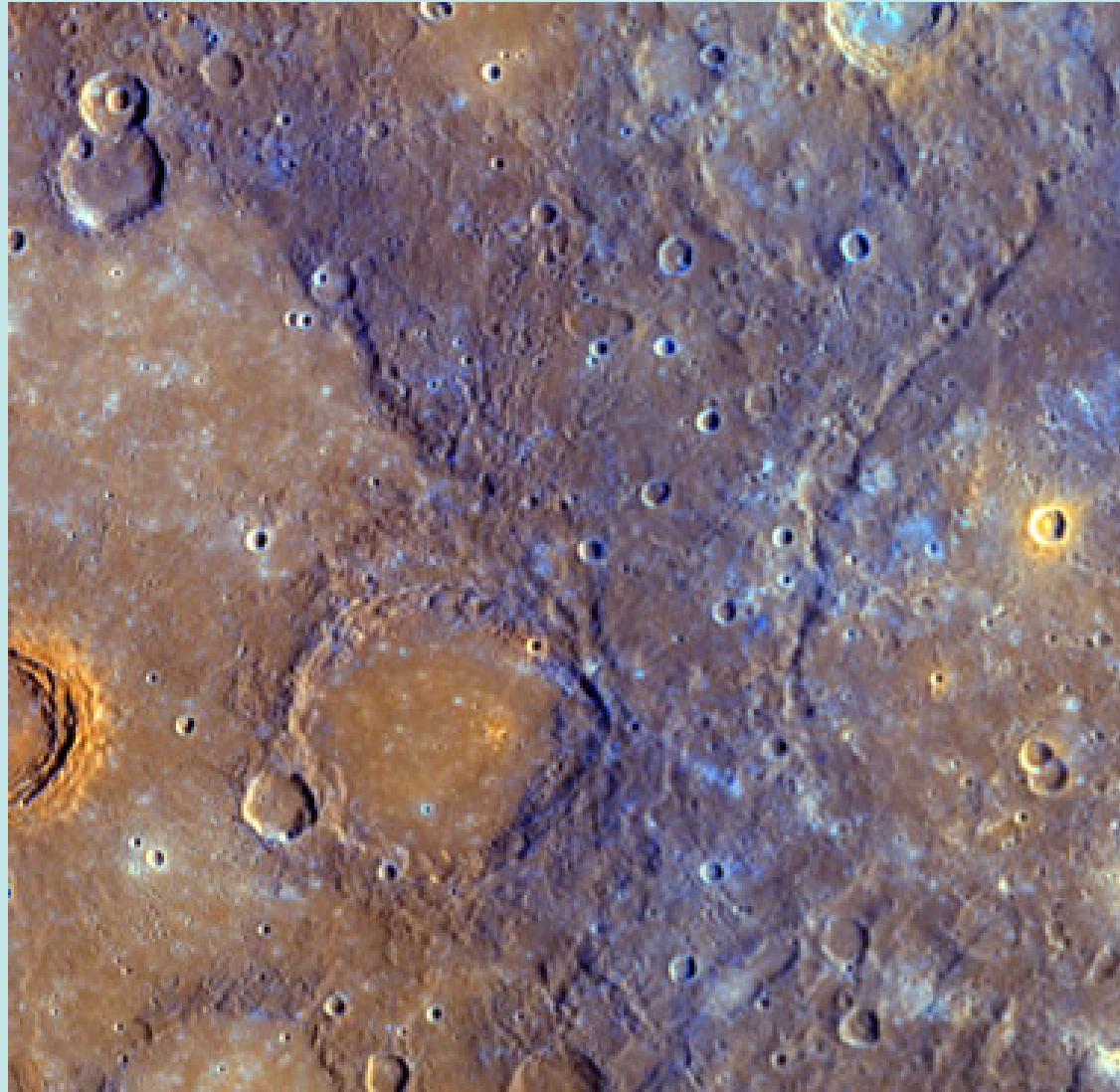
Mercury

Mariner 10 mosaic



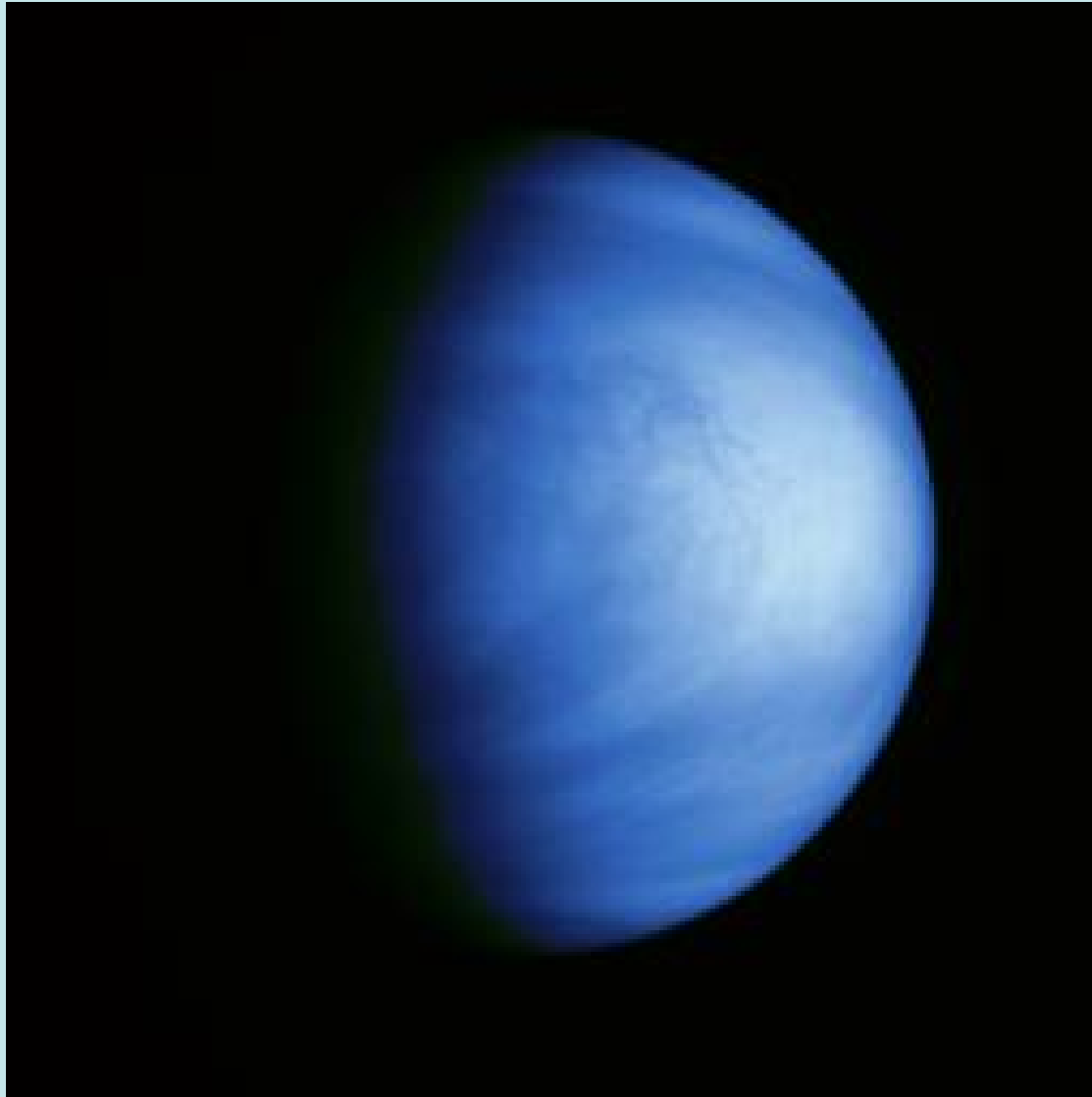
Mercury

MESSENGER close-up (enhanced color)



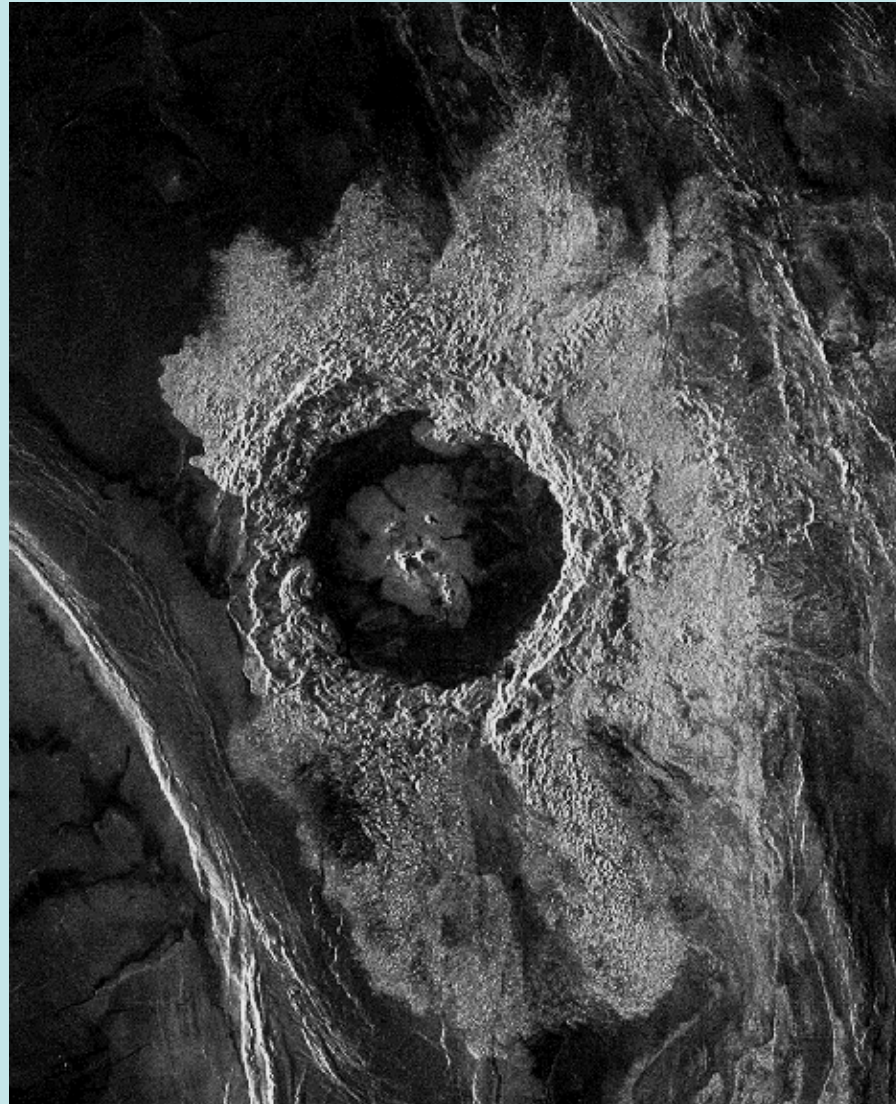
Venus

Violet light - Galileo spacecraft image



Venus

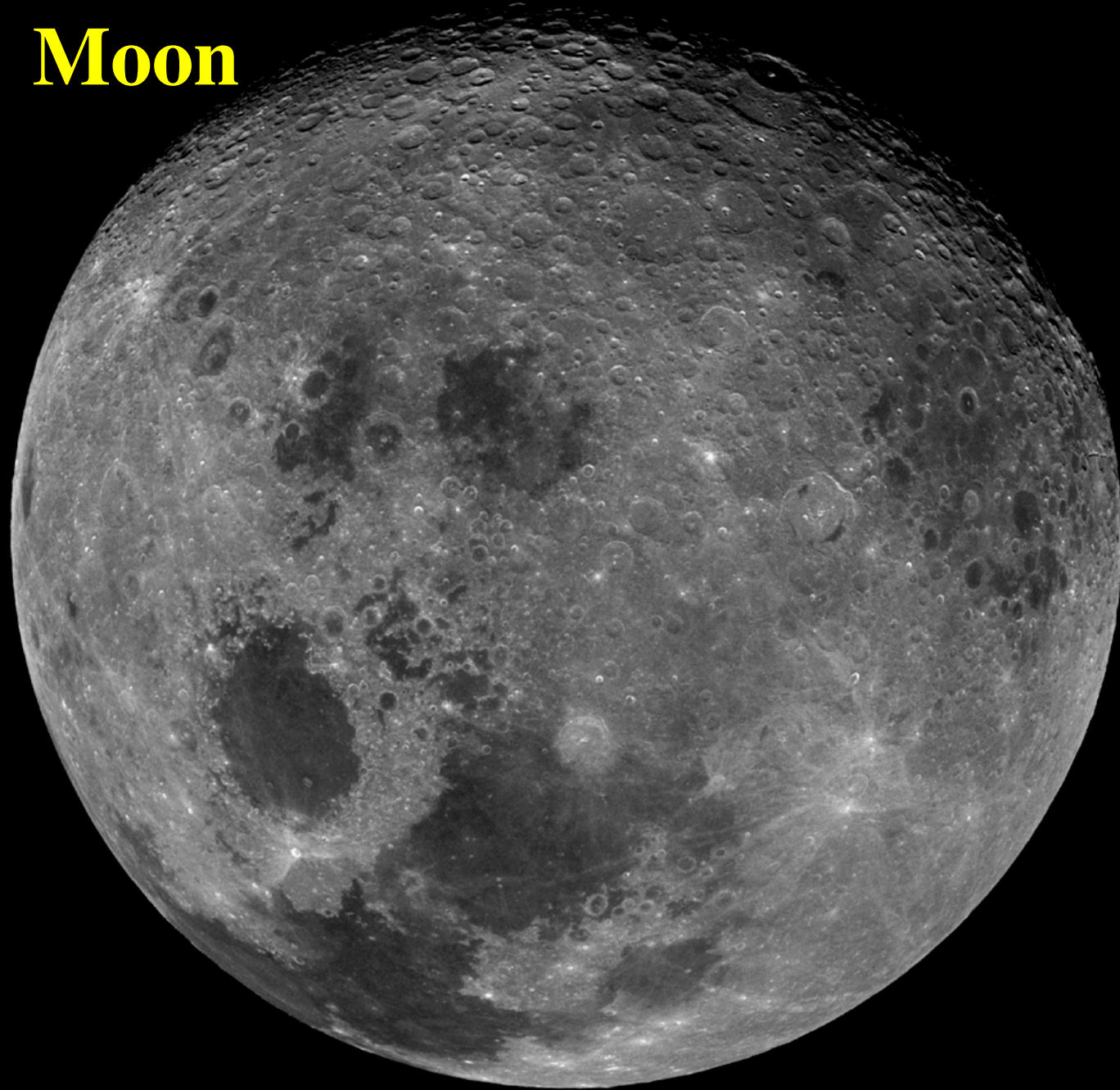
RADAR - Magellan spacecraft image





Earth

Moon



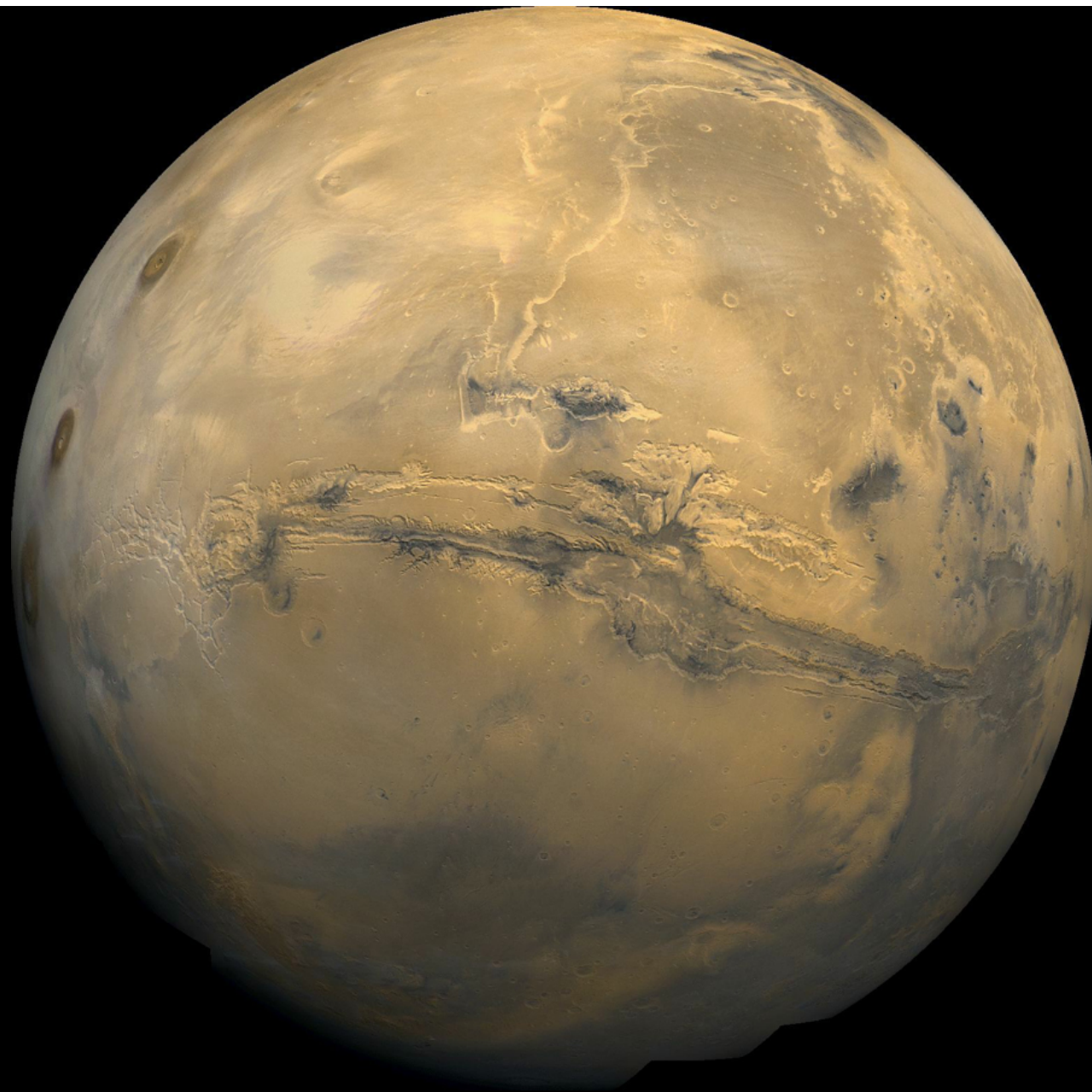
Meteor Crater

Arizona, USA Smithsonian 1938

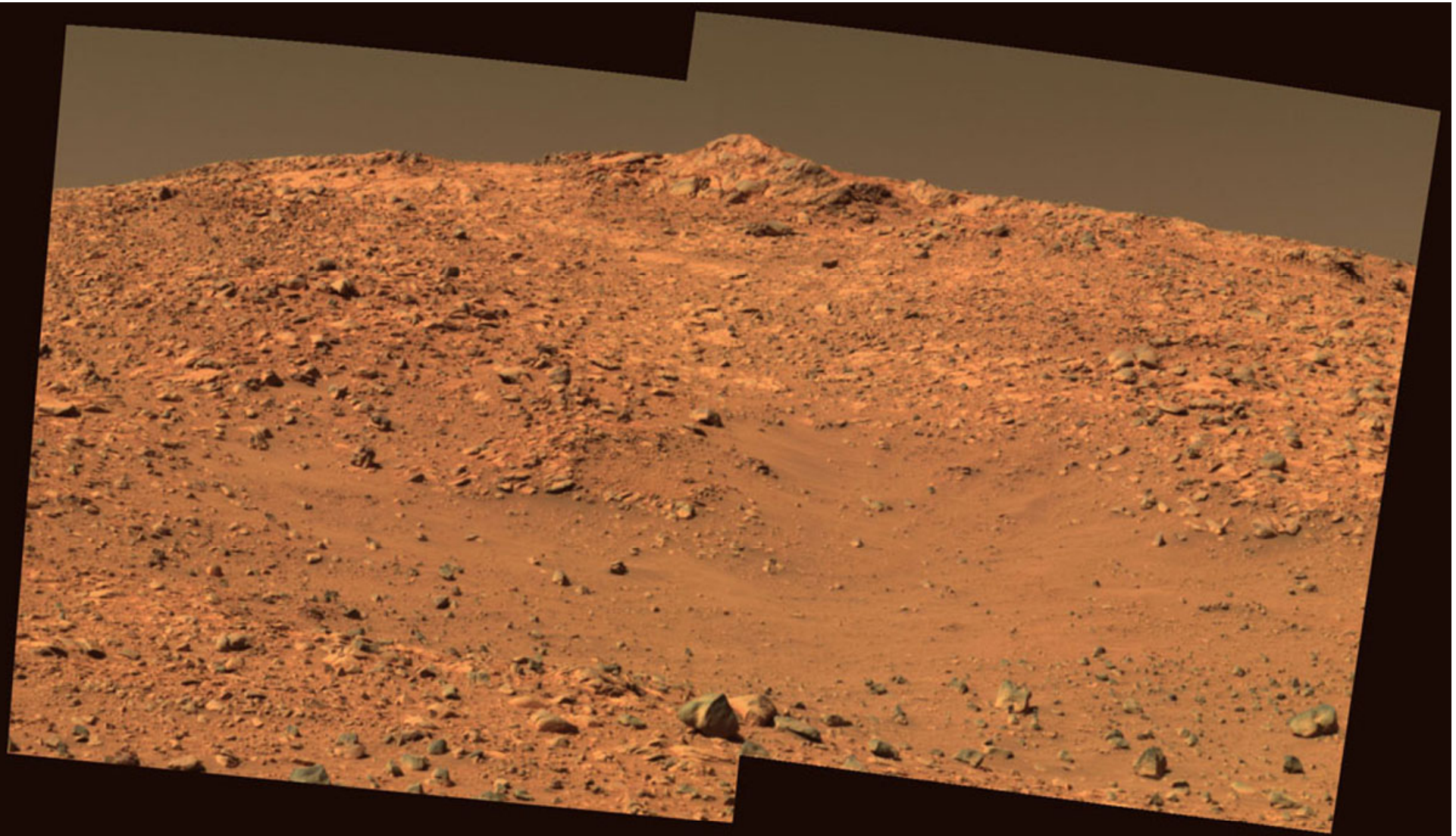




Asteroid Eros

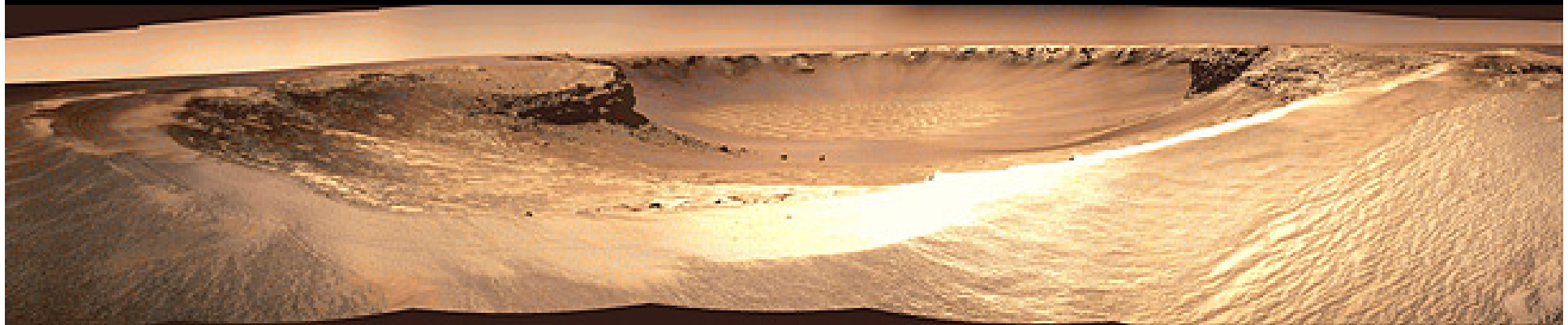


Mars



True-color composite image, taken by Mars Exploration Rover Spirit, shows Martian landscape.

Mars Surface - Opportunity Rover at Victoria Crater





Asteroid Ida and its moon, Dactyl

Sikhote-Alin (eastern USSR) **Iron IIB** Fall 1947 Feb 12



Photo: Jackie Beckett

© AMNH 2003

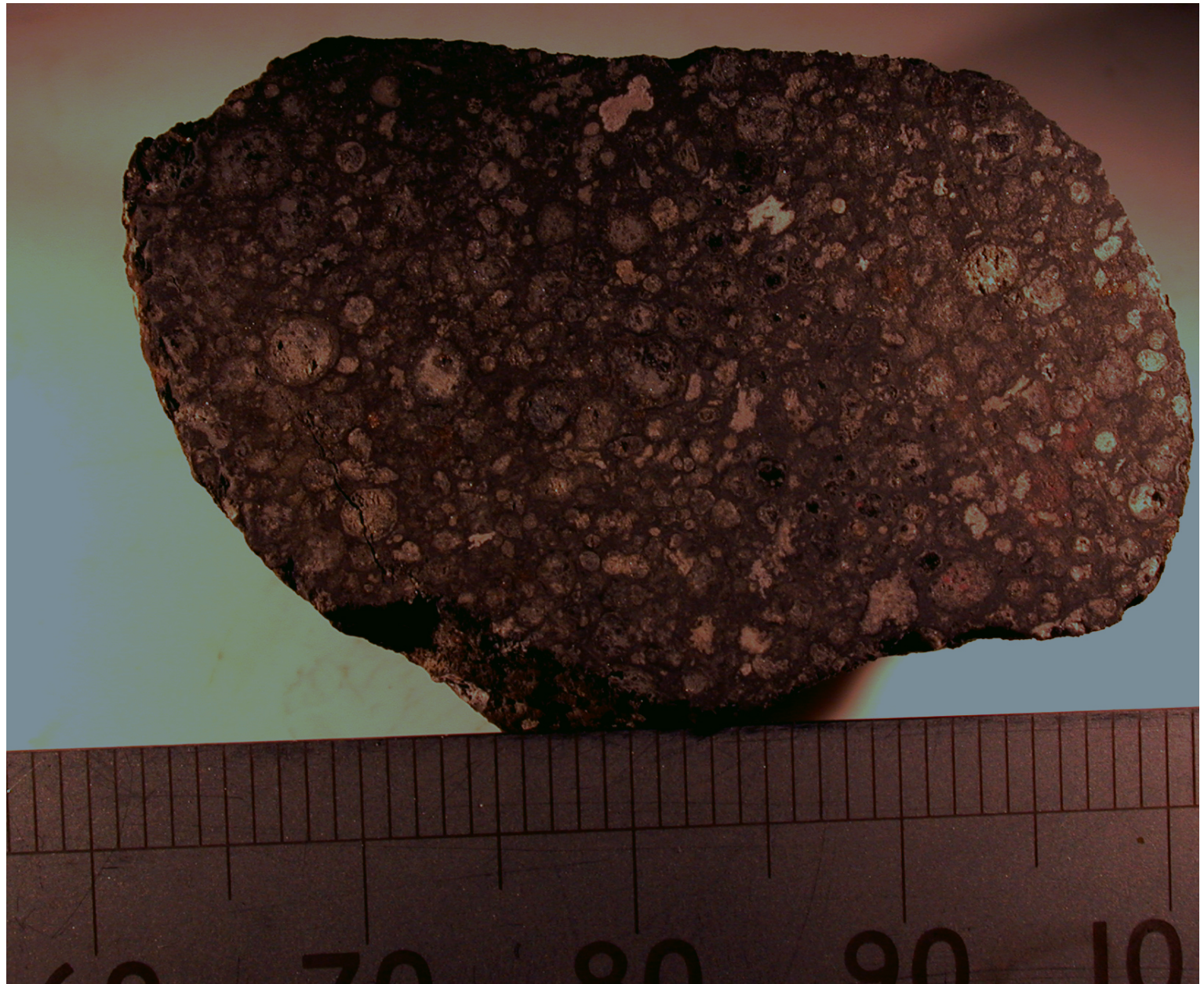
Cape York (Greenland) *"Ahnighito"* **Iron** 31 metric tons

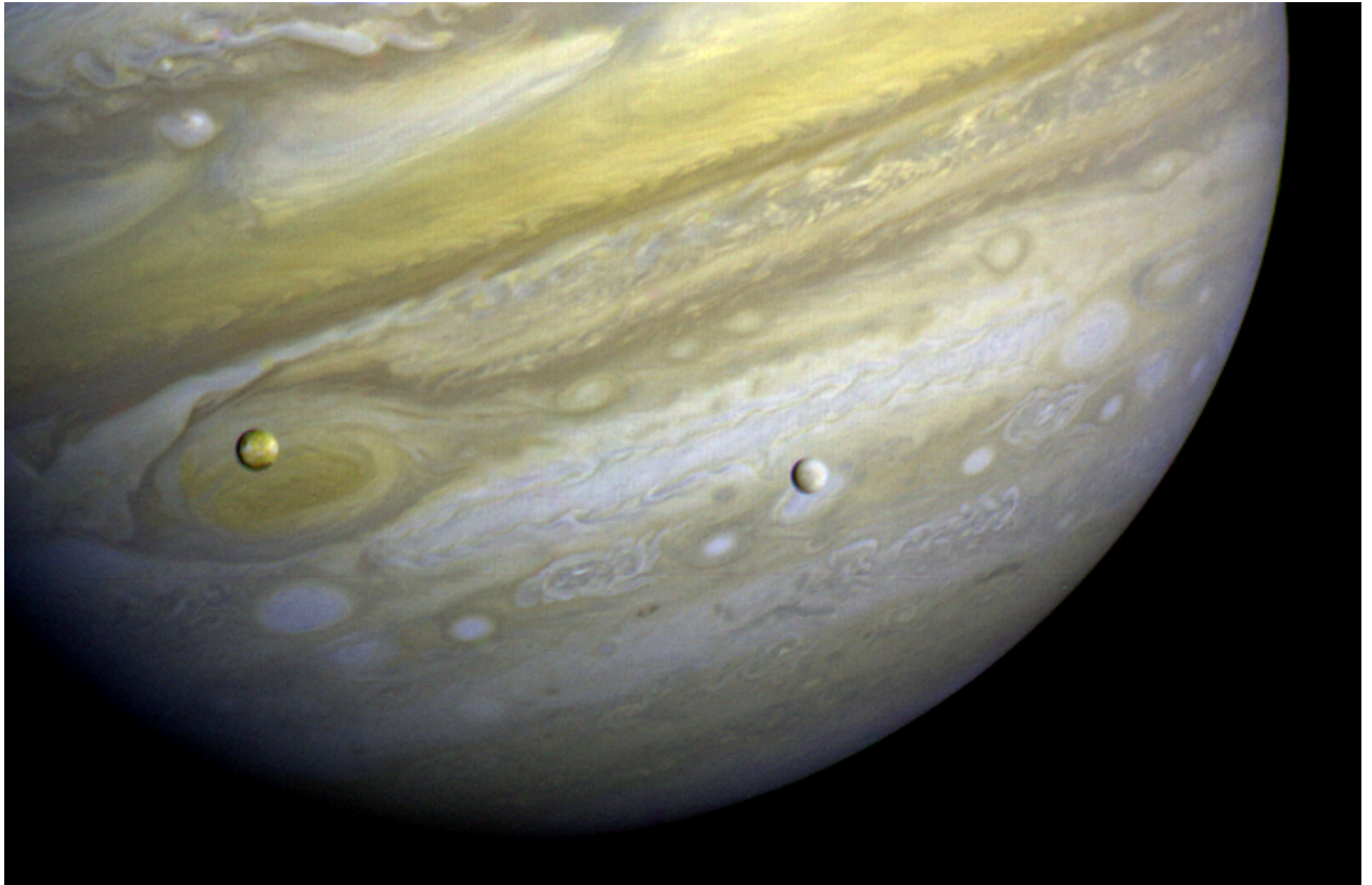


Photo: © AMNH

Allende CV3 Carbonaceous Chondrite Meteorite

Close-up view.
**This piece is
39 mm long.**
**Note CAIs &
chondrules.**





Jupiter, Io & Europa



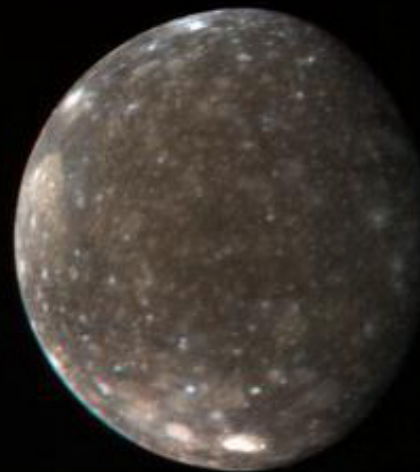
Io



Europa



Callisto



Ganymede

Jupiter's four largest satellites

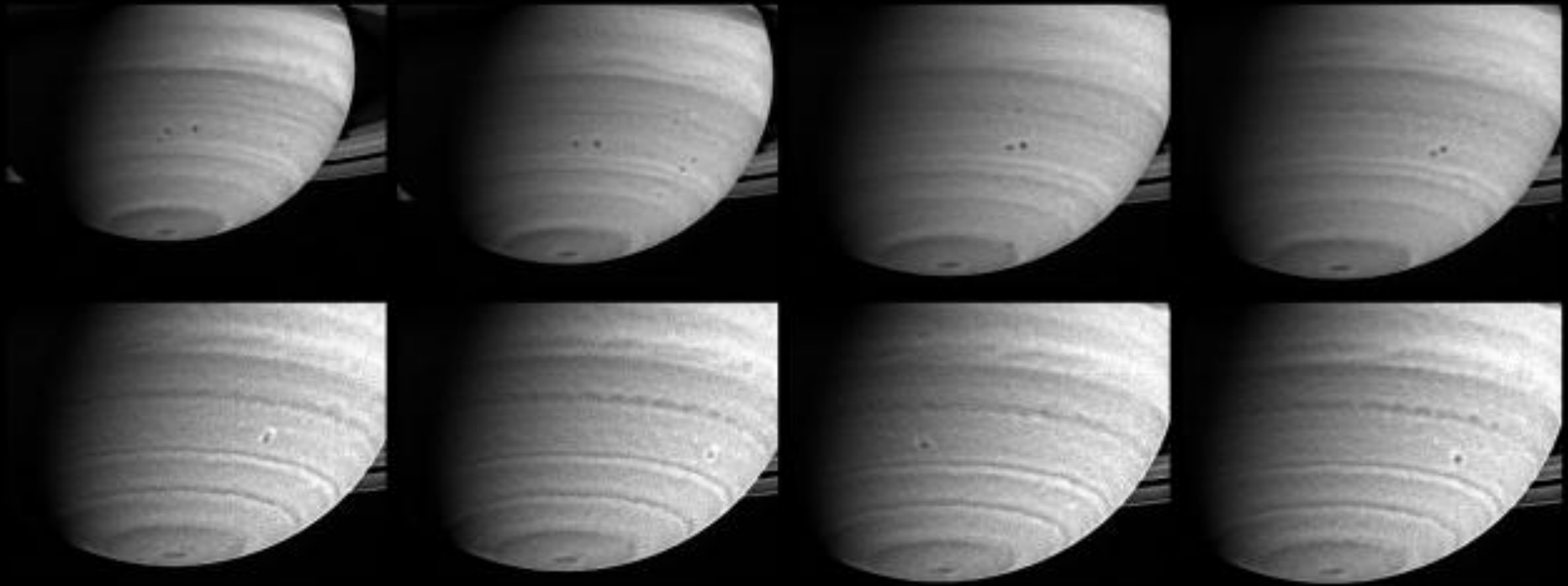
Saturn



Hubble
Heritage

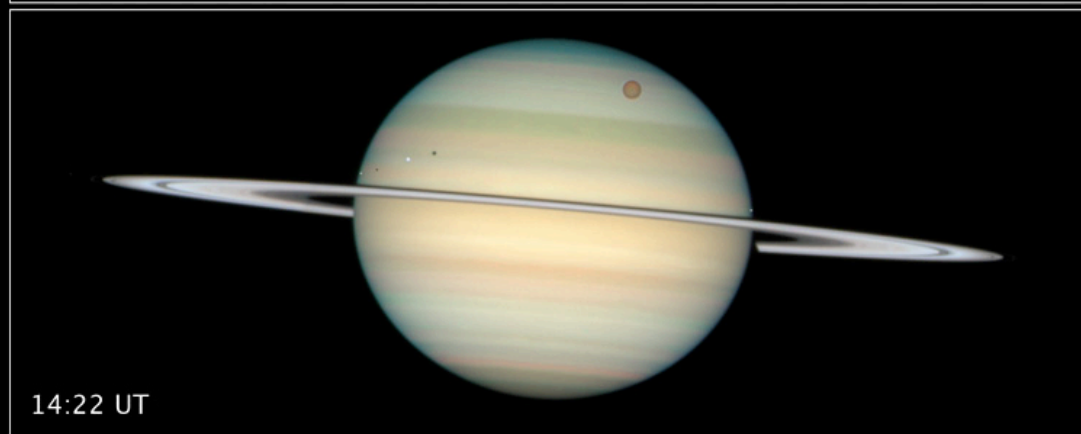
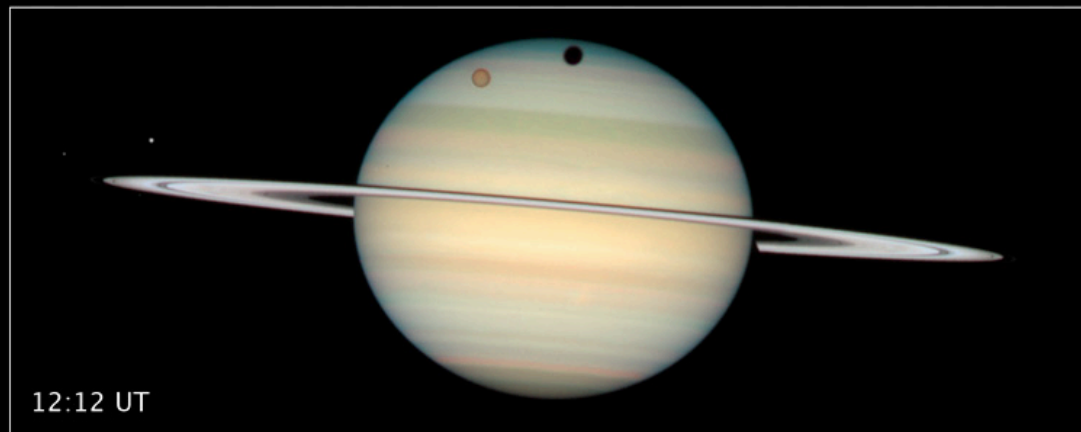
Merging Storms on Saturn

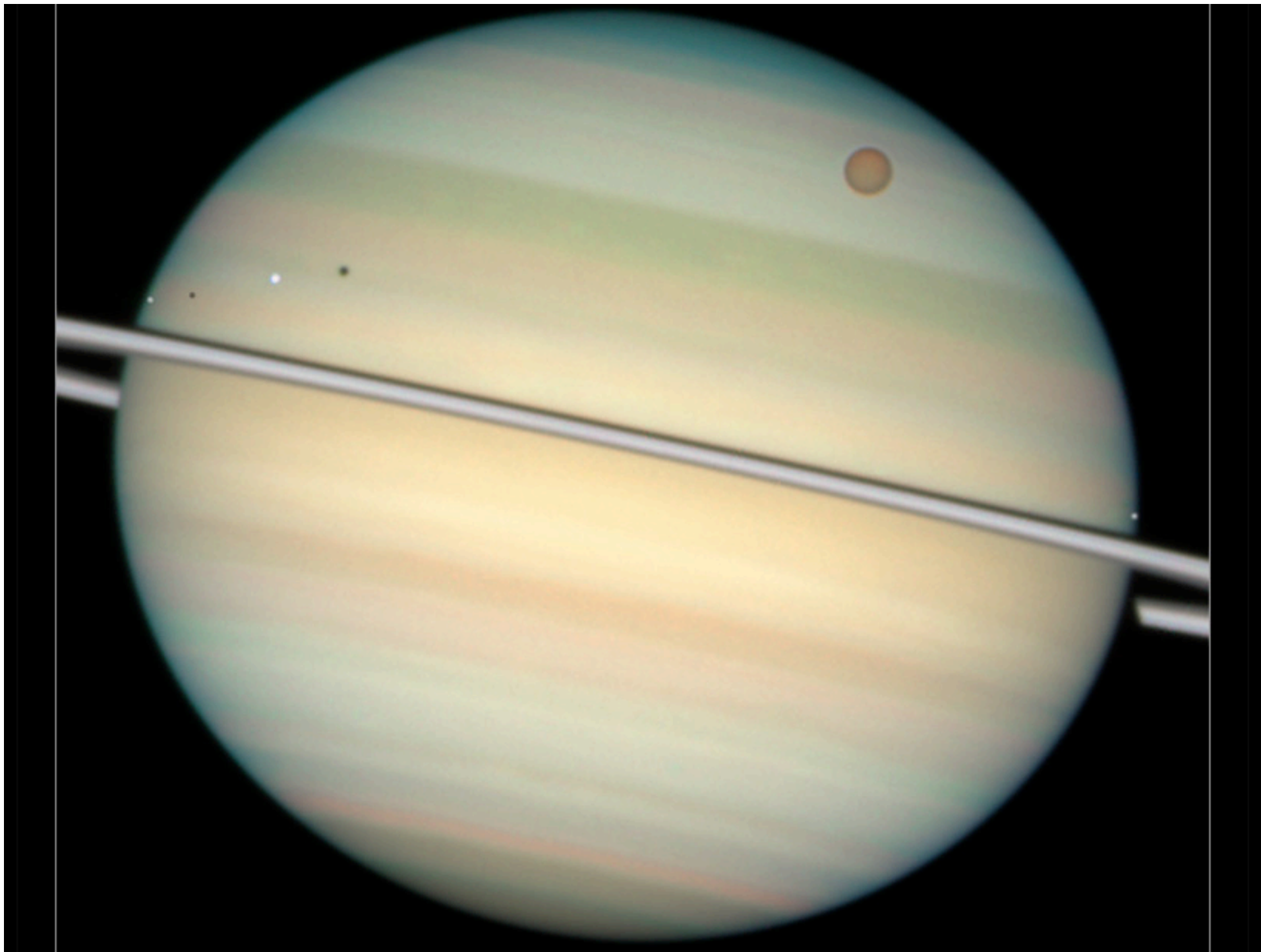
(2004 - Cassini Images)



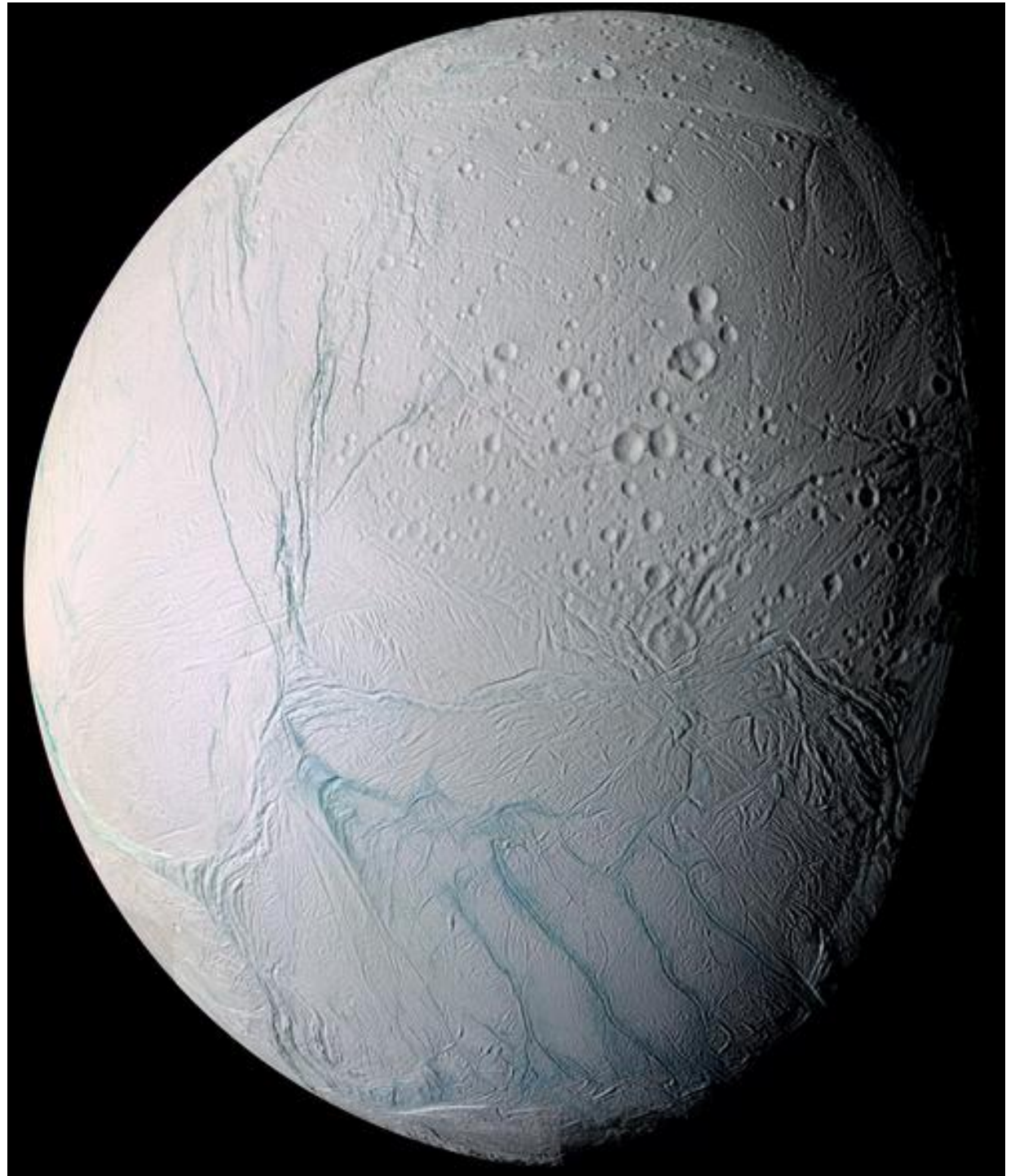
Saturn February 24, 2009

Hubble Space Telescope WFPC2



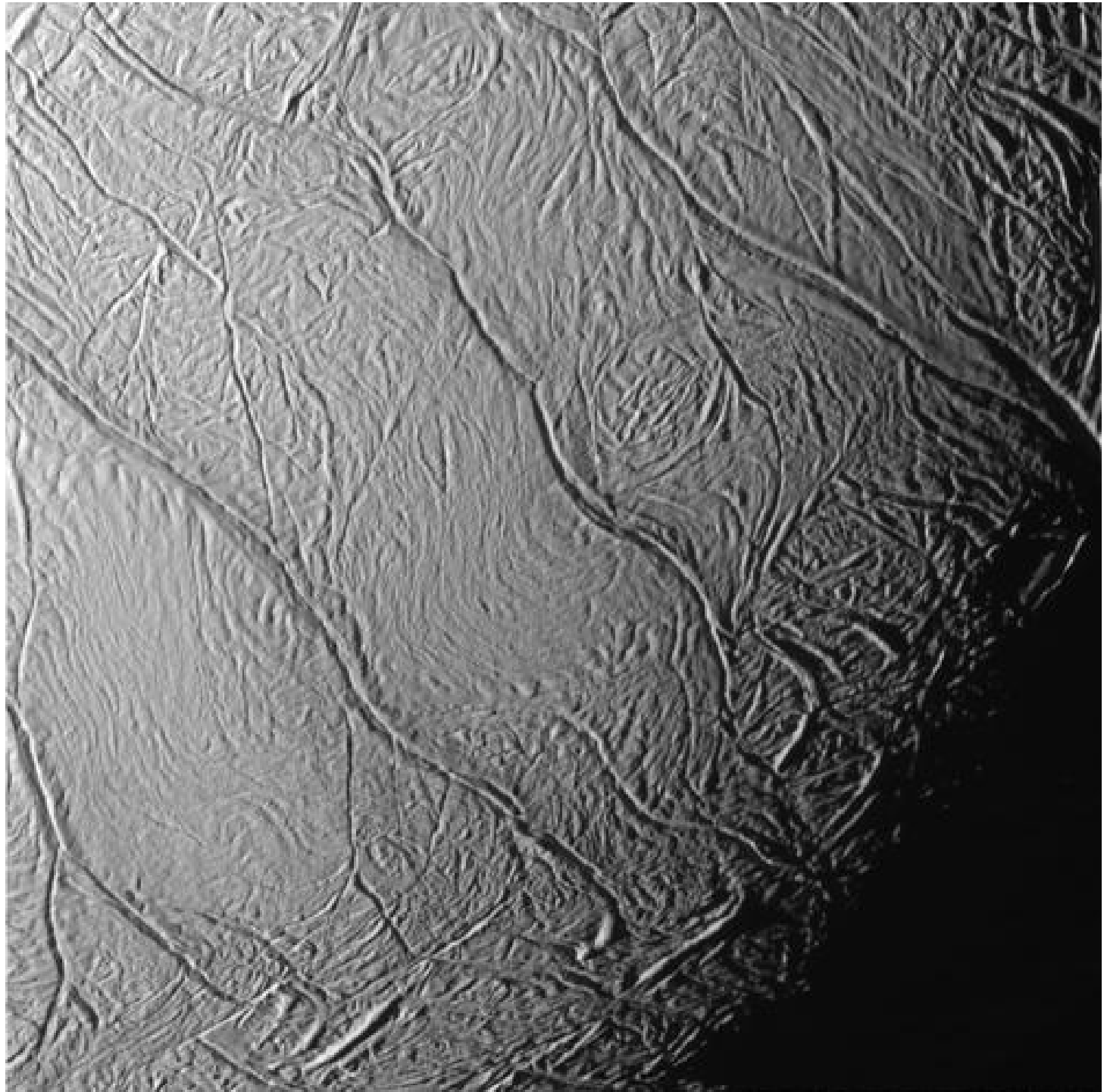


Enceladus

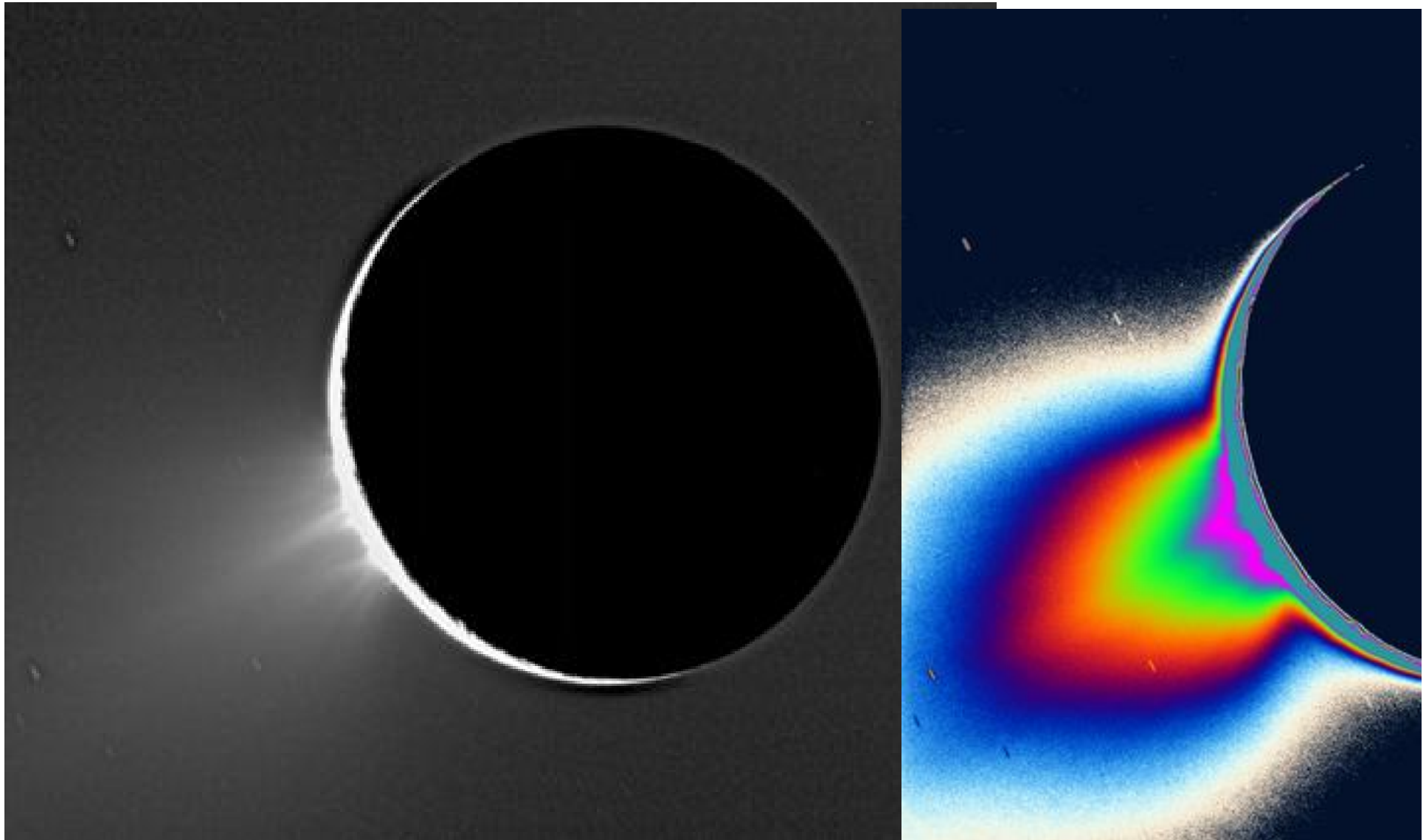


Enceladus:

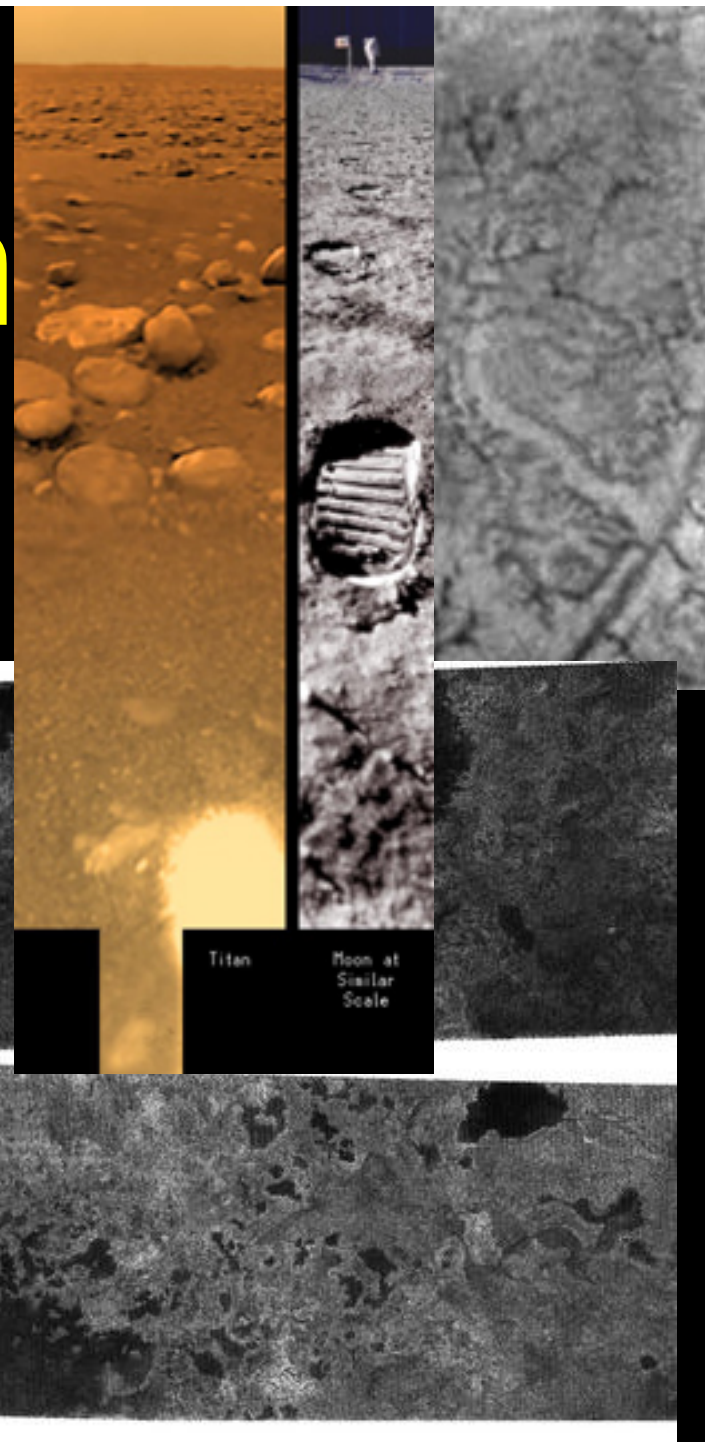
‘Tiger Stripes’



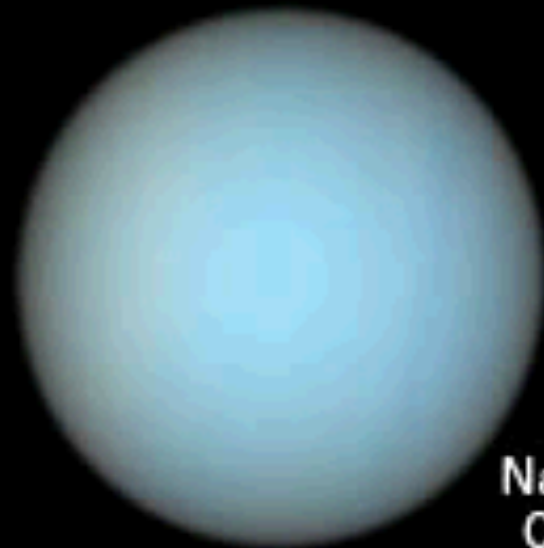
Plumes from Enceladus



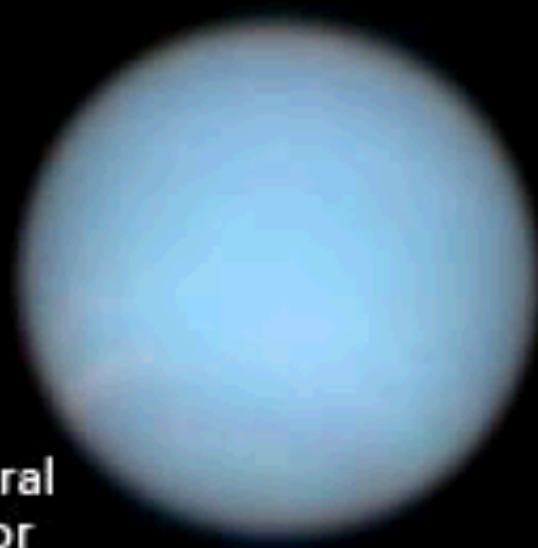
Titan



HST/ACS images 2003



Uranus



Neptune

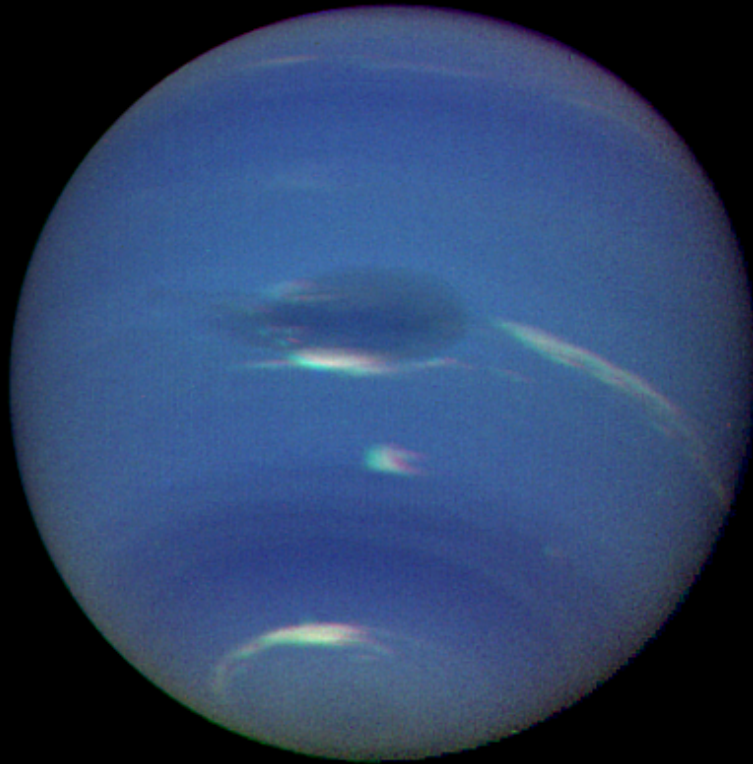
Natural
Color



Enhanced
Color



Neptune - Voyager

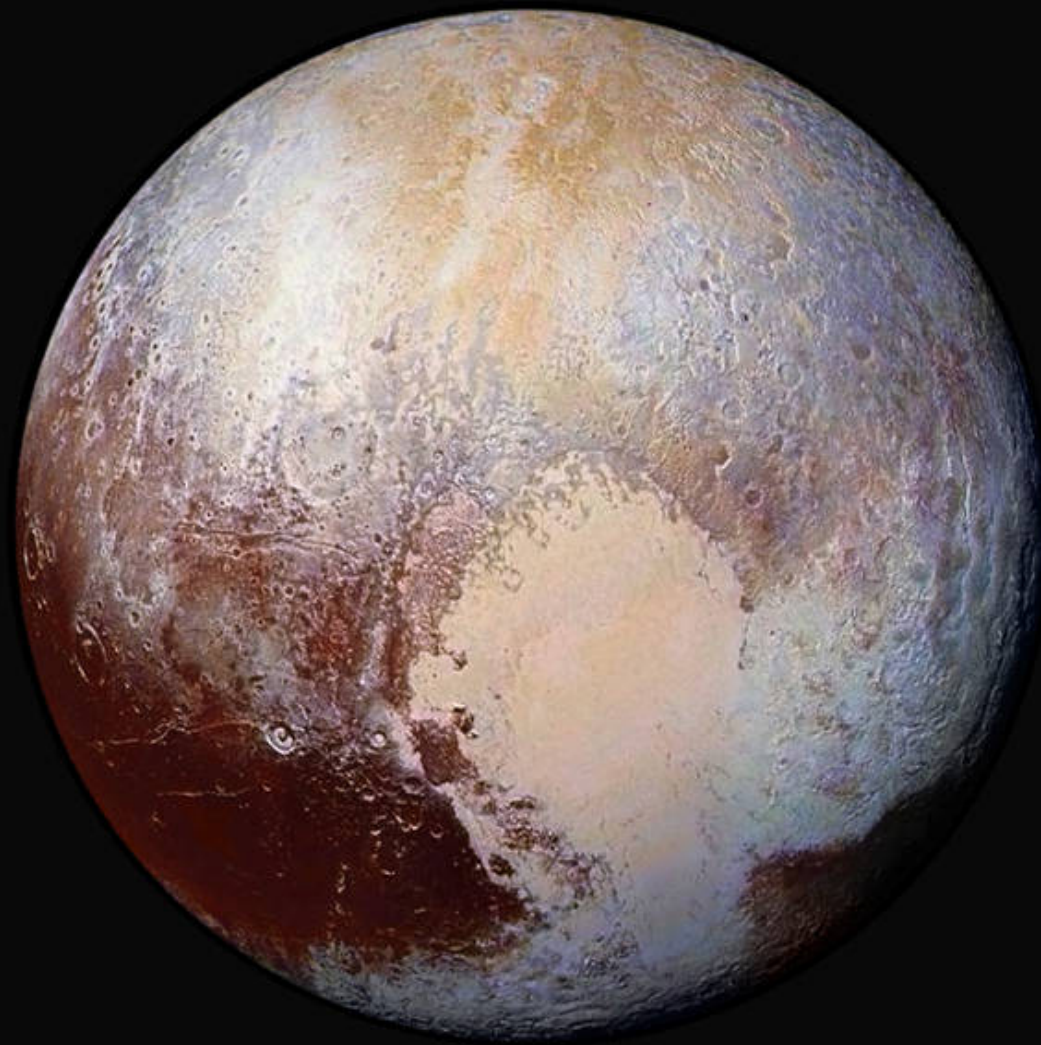


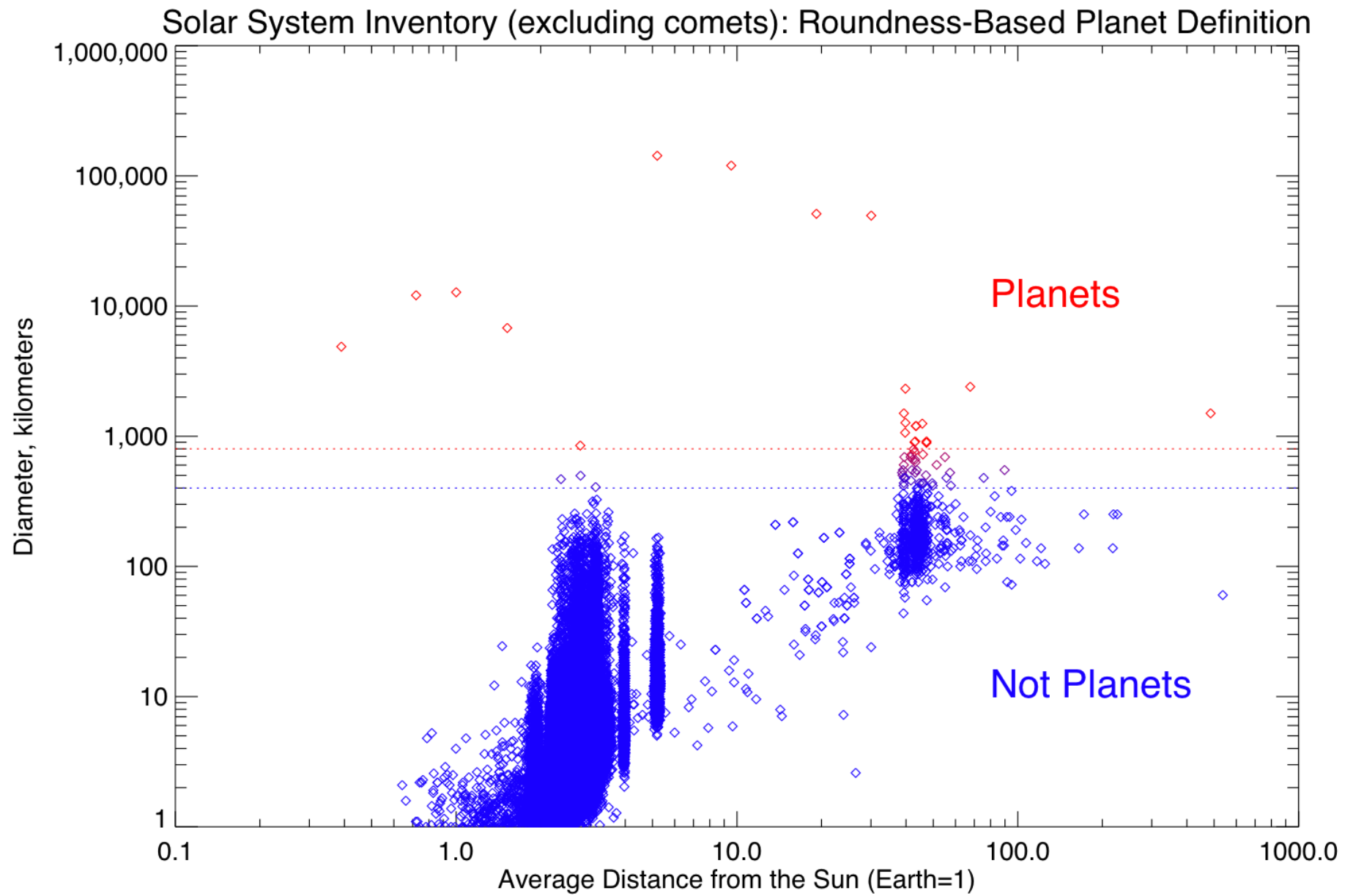
Triton – Neptune's large moon



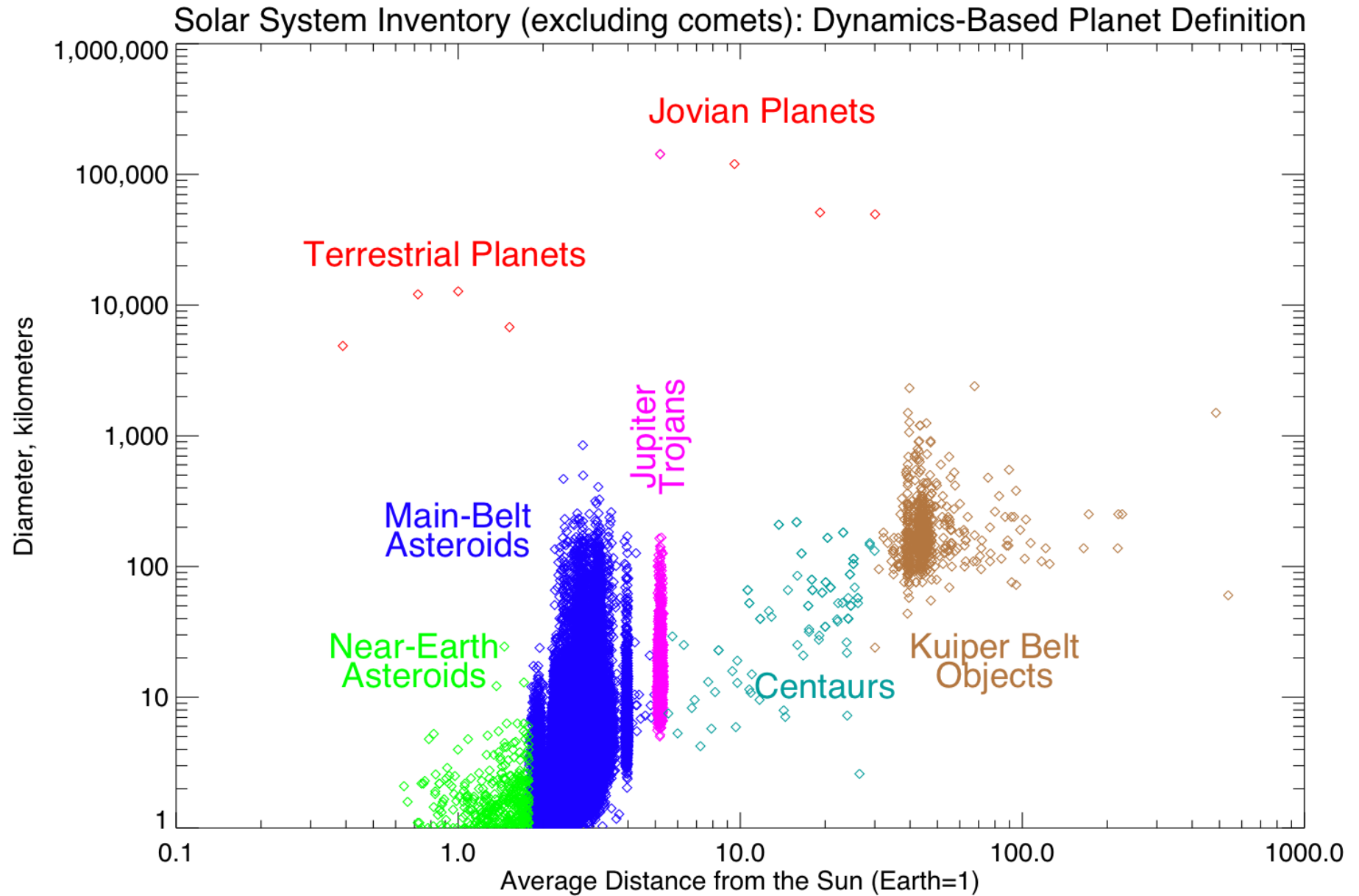


Pluto – New Horizons (false color)





John Spencer, SwRI



John Spencer,

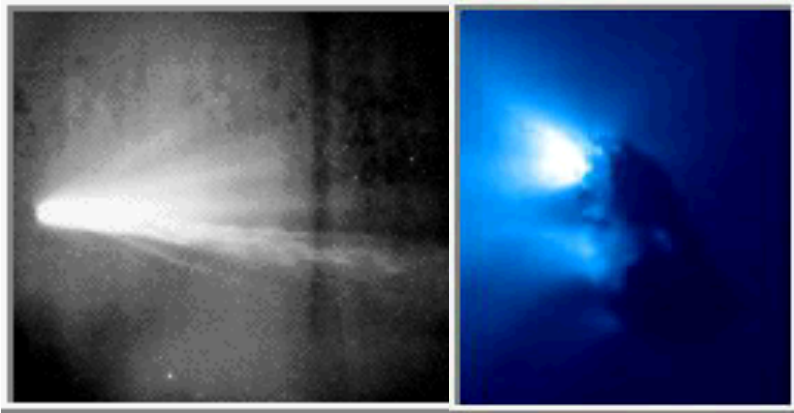
Small Bodies: Asteroids, Comets, KBOs

- Asteroid = Minor planet with $a < 6$ AU



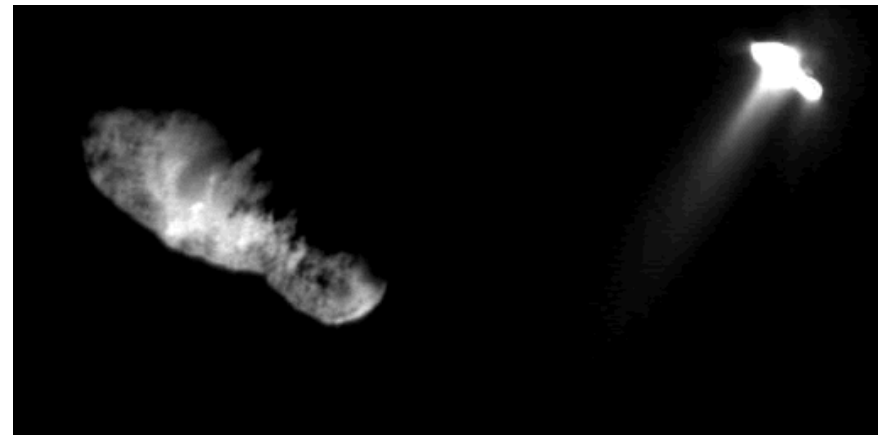
Asteroid
433 Eros

- Comet = Diffuse coma, tails



Comet 1P/Halley

Comet 19P/Borrelly



- Kuiper belt/KBOs

Comets Hyakutake & Hale-Bopp

1996





Comet Halley

Comet Halley

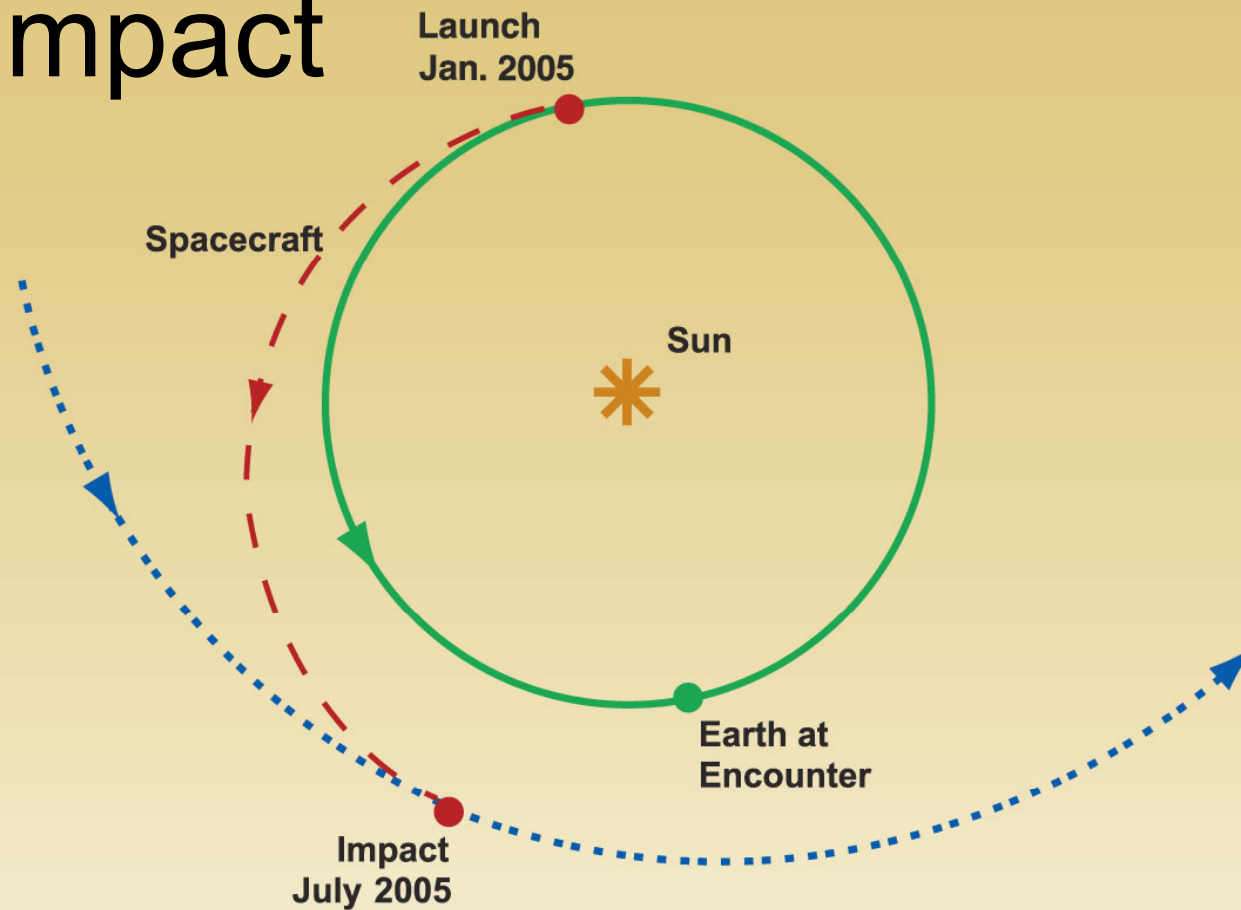
Close-up -- *Giotto* Spacecraft



P/Wild 2 Stardust 20 m res.

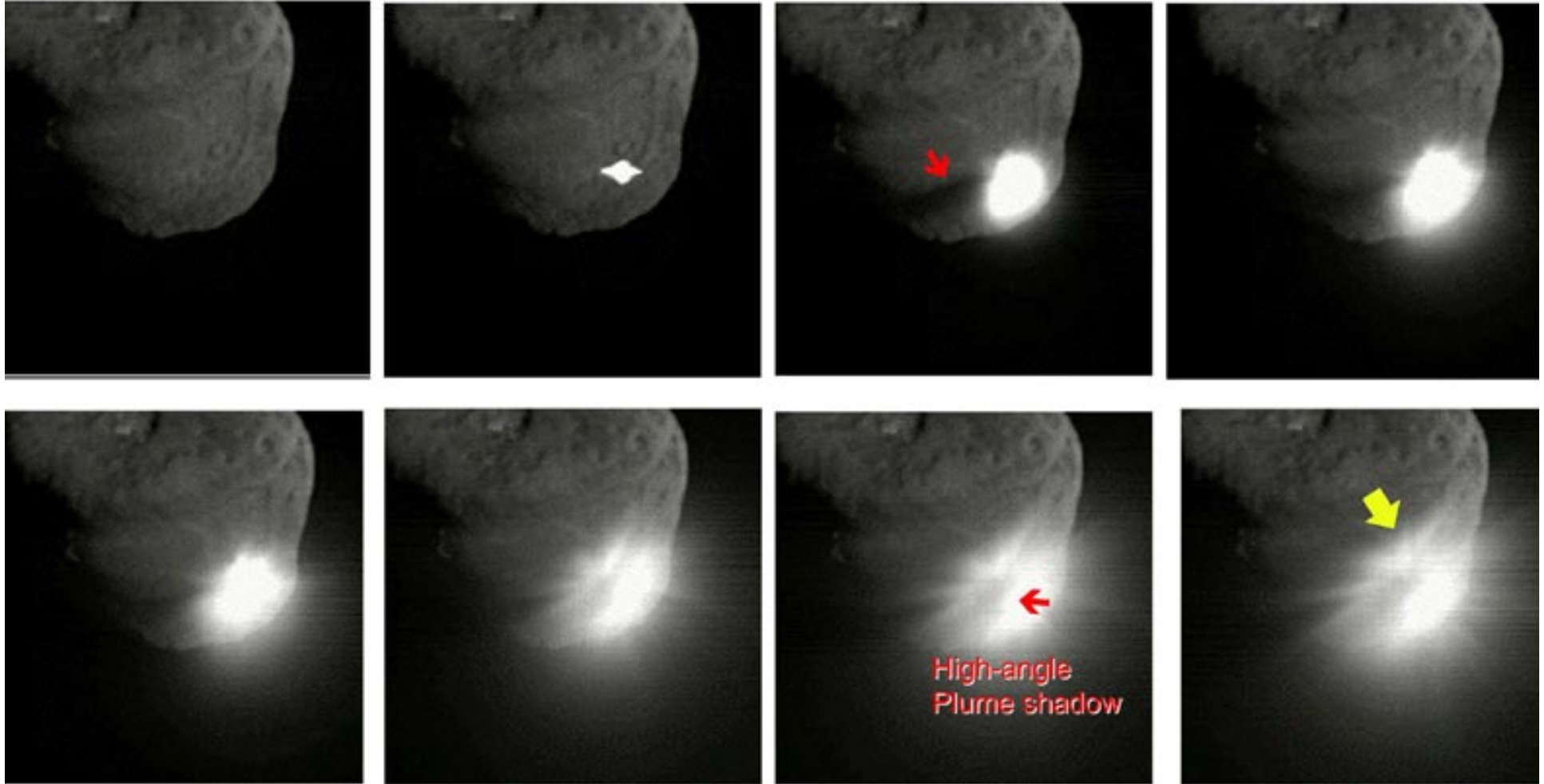


Deep Impact

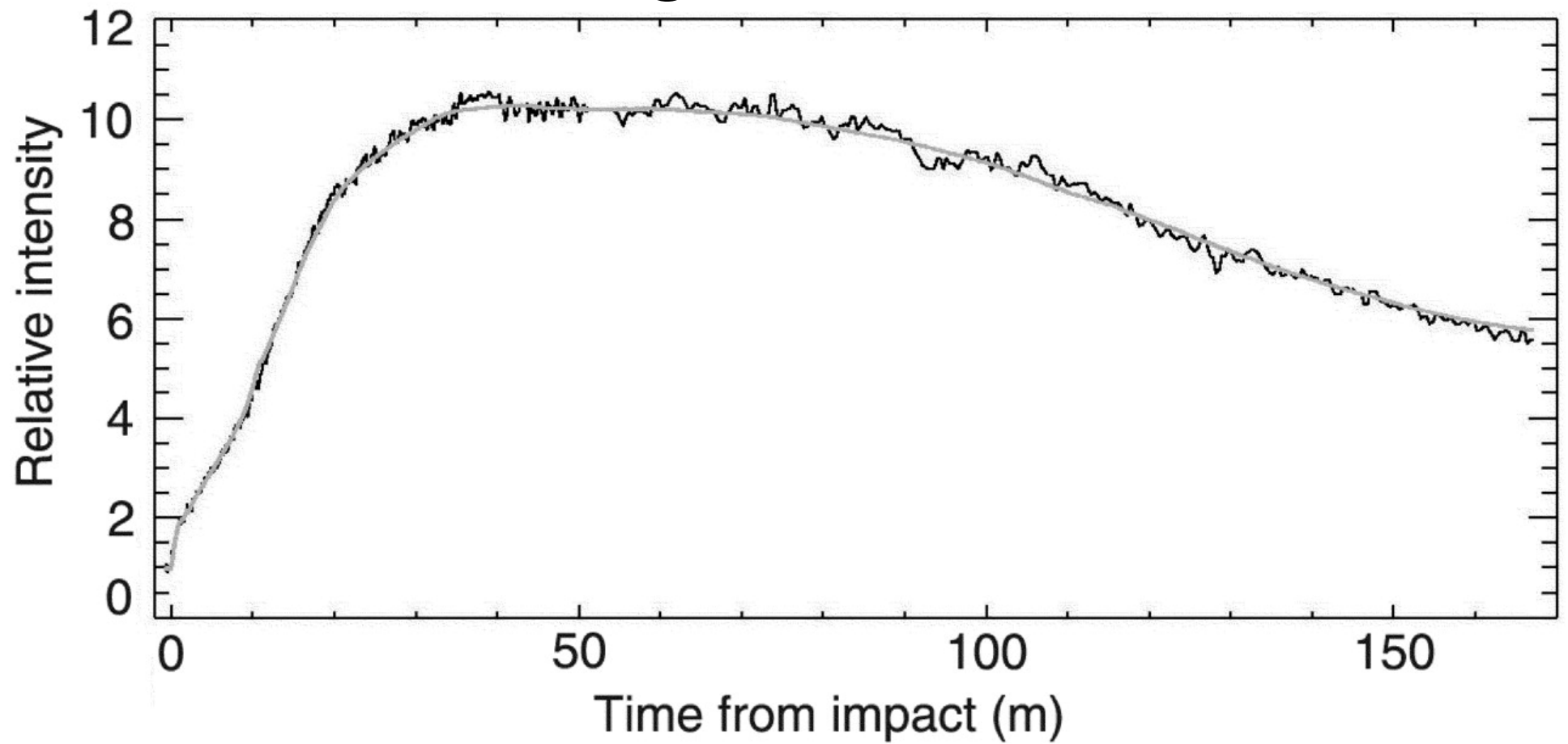


- Tempel 1 Orbit (5.5 year Period)
- - - - - Spacecraft Transfer Orbit (from Earth to Tempel 1)
- Earth Orbit

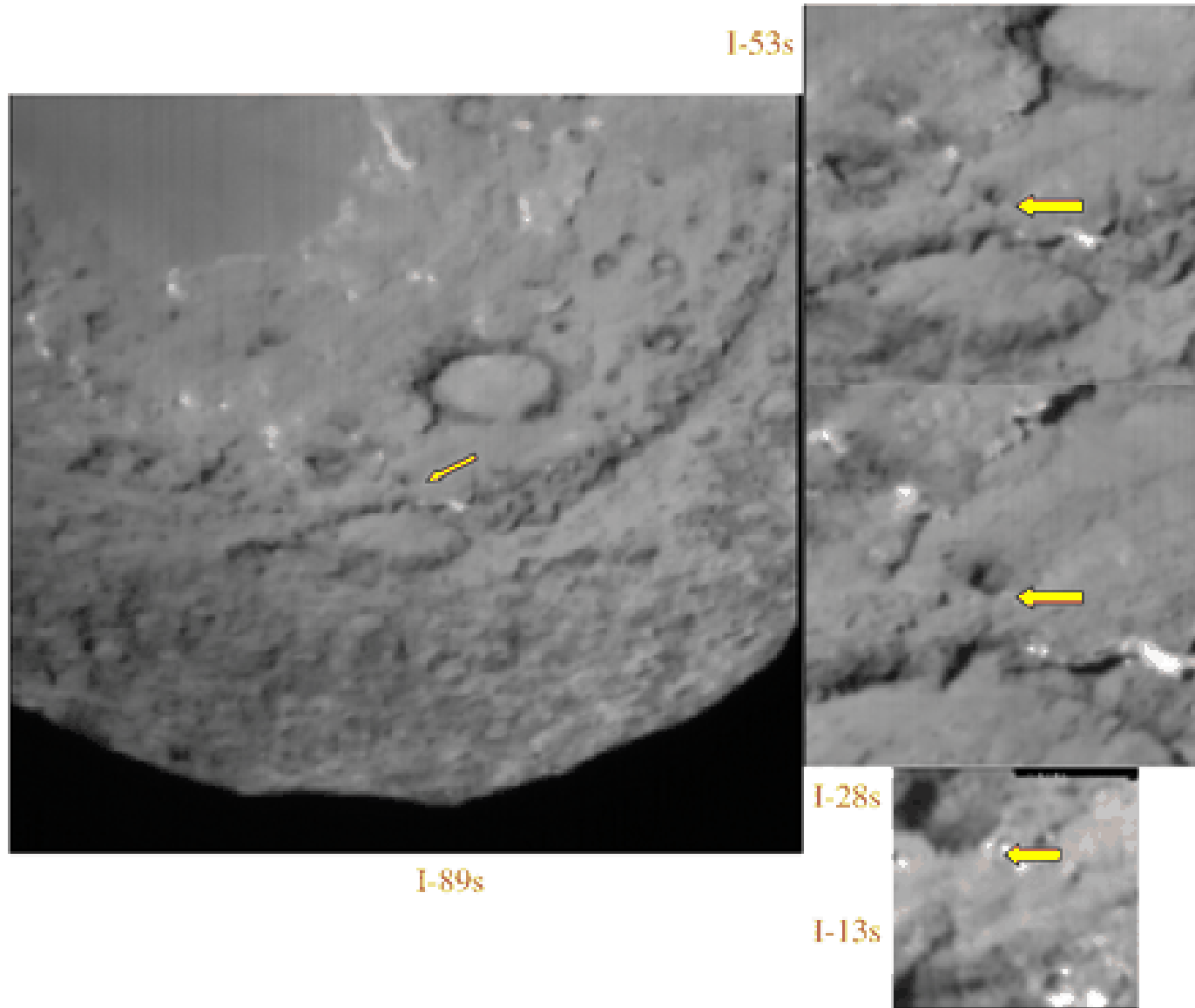
The Impact (images span 6 seconds)



Lightcurve



Impact Site prior to the impact

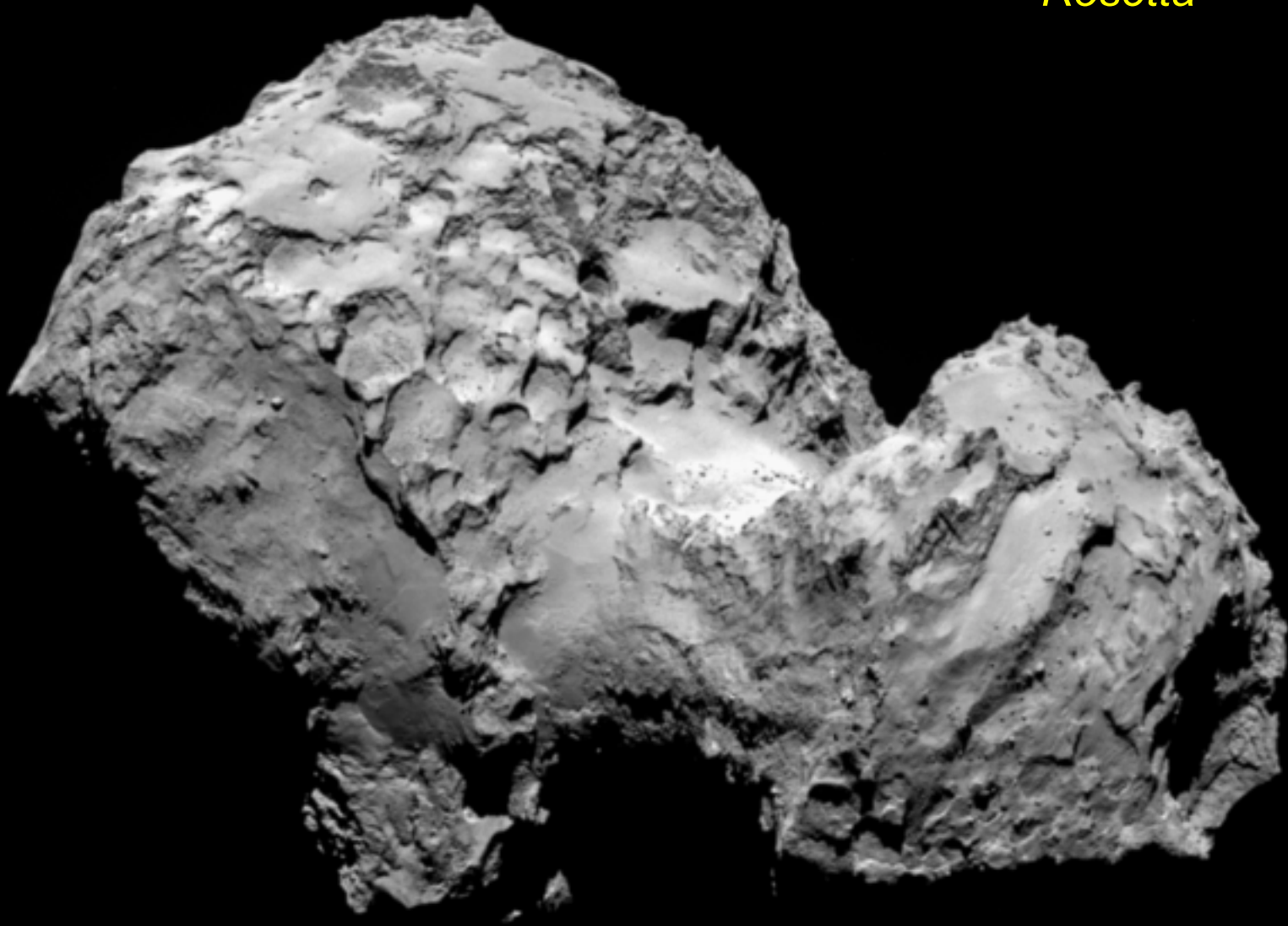


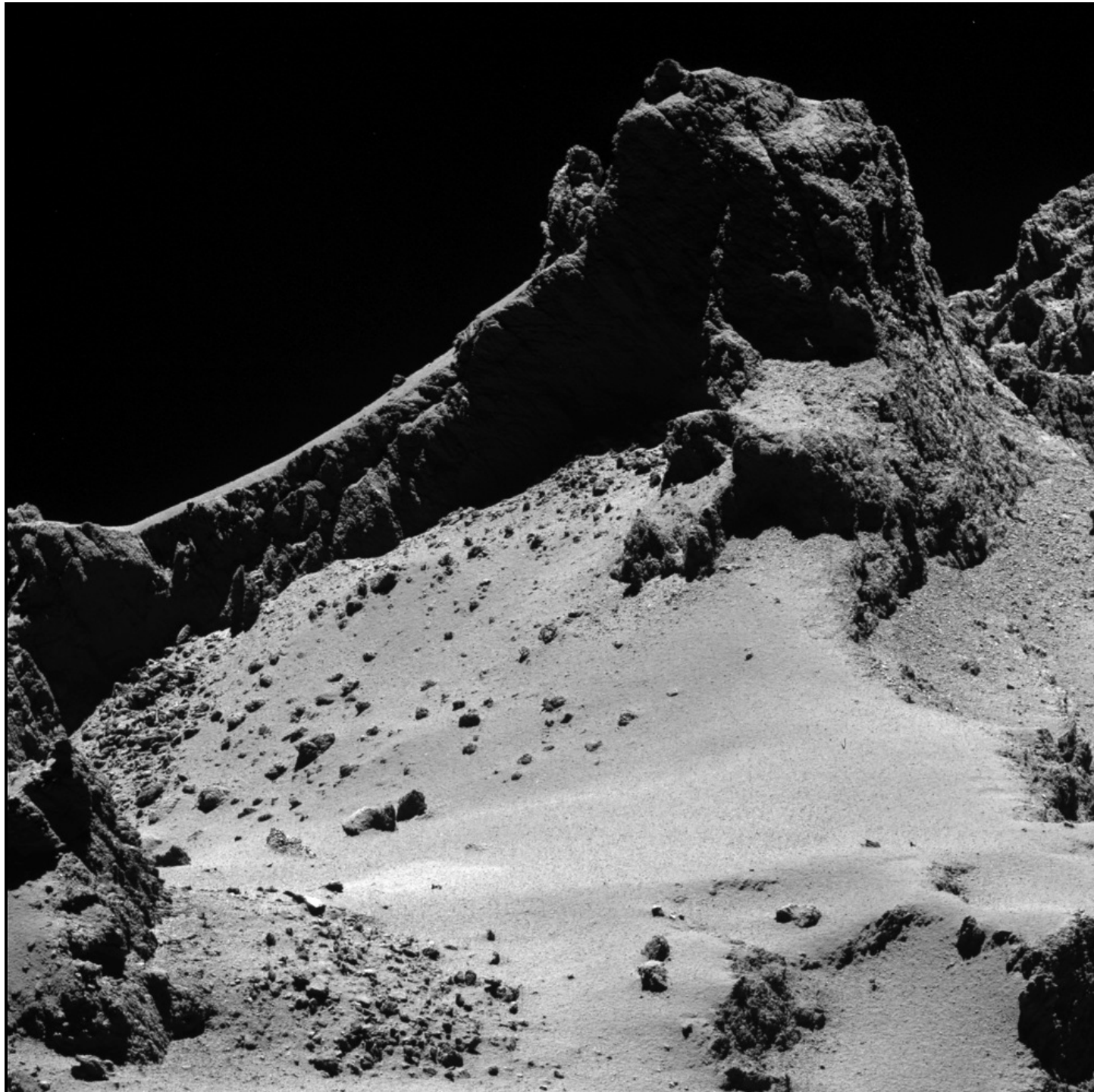
45 minutes after Impact



Comet 67P/Churyumov–Gerasimenko

Rosetta



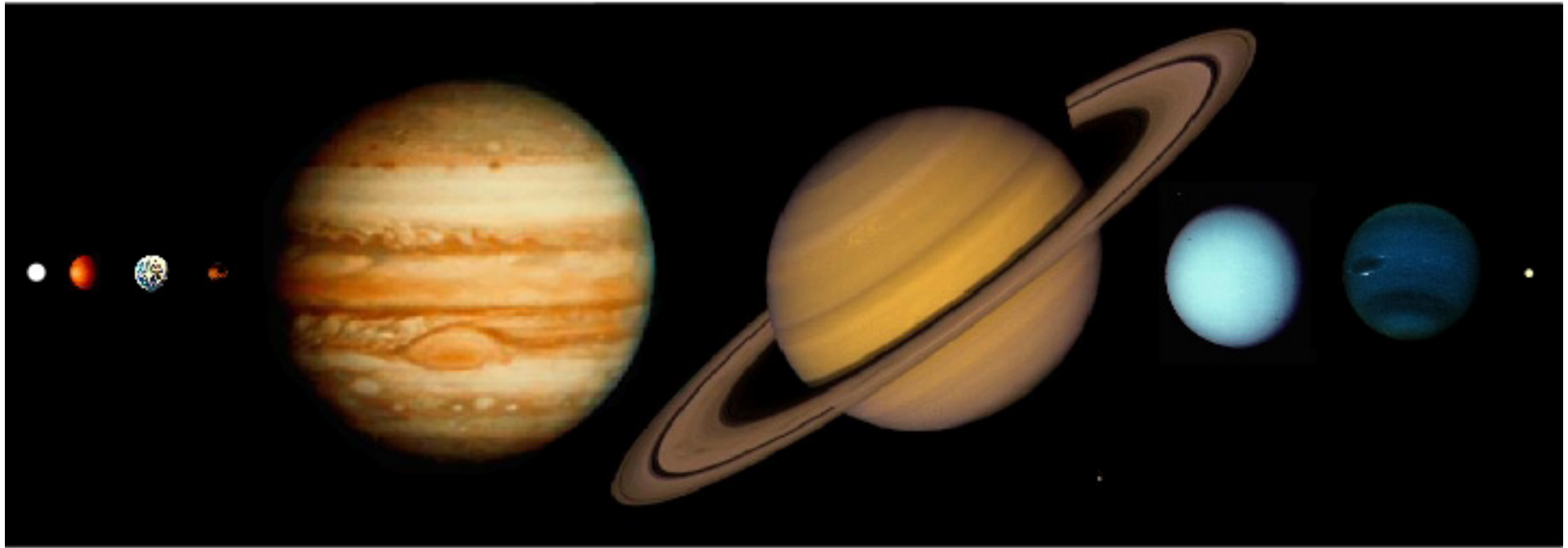


Comet 67/P

Close-up

Rosetta from 8 km

PLANETS IN OUR SOLAR SYSTEM



| | Terrestrials | | | | | Gas giants | | | | |
|---------|--------------|-------|-------|------|---------|------------|--------|---------|-------|--|
| | Mercury | Venus | Earth | Mars | Jupiter | Saturn | Uranus | Neptune | Pluto | |
| mass | 0.055 | 0.82 | 1.00 | 0.11 | 318 | 95 | 14 | 17 | .0002 | |
| radius | 0.38 | 0.95 | 1.00 | 0.53 | 11.2 | 9.4 | 4.0 | 3.9 | 0.18 | |
| area | 0.15 | 0.90 | 1.00 | 0.28 | 126 | 89 | 16 | 15 | 0.03 | |
| volume | 0.06 | 0.85 | 1.00 | 0.15 | 1408 | 844 | 64 | 59 | 0.006 | |
| density | 0.98 | 0.95 | 1.00 | 0.71 | 0.24 | 0.12 | 0.24 | 0.32 | 0.20 | |

(all values are relative to Earth)

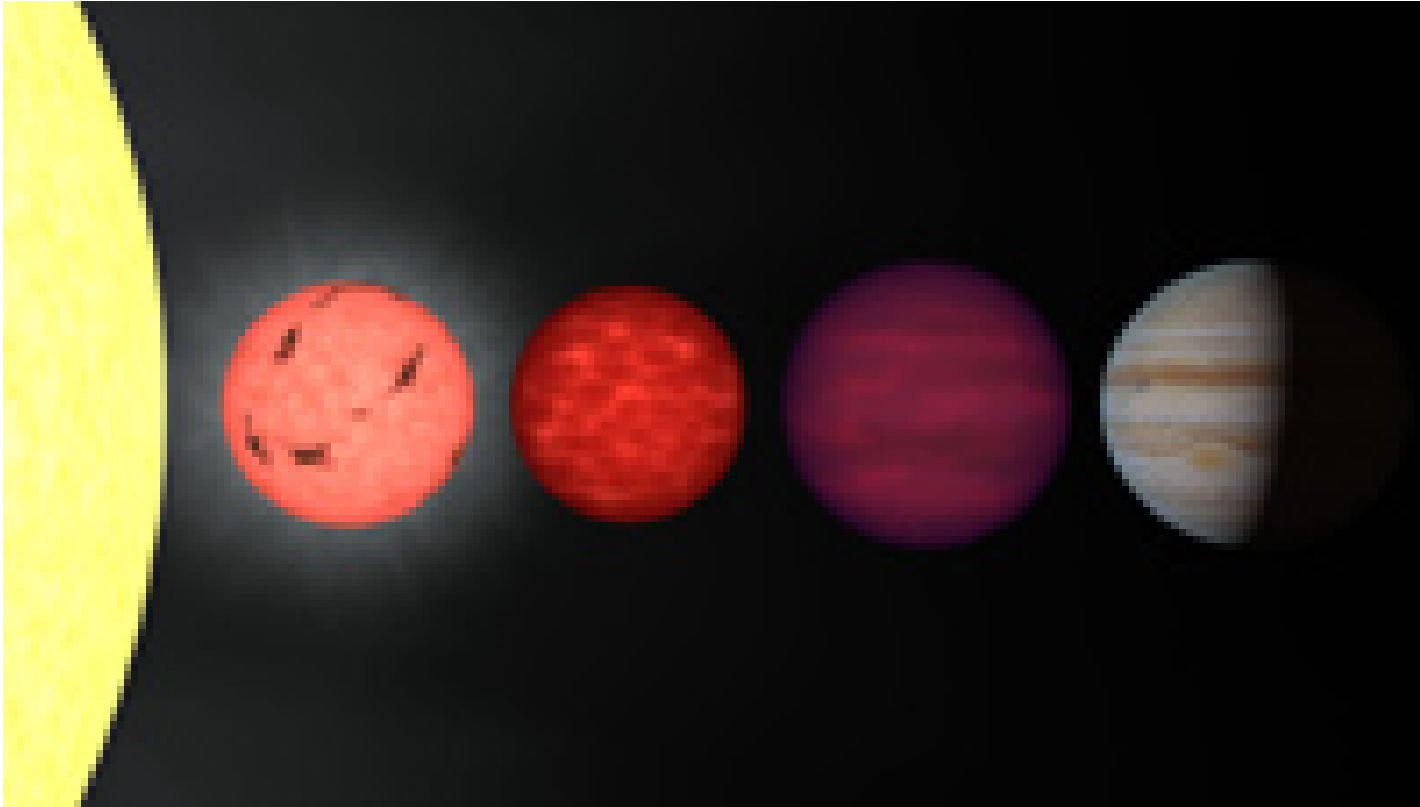
Sun



Earth



Stars, Brown Dwarfs & Planets



Sun, 75, 60, 30 & 1 Jupiter mass objects

Extrasolar Planets

- Detection Techniques
 - Pulsar timing
 - **Radial velocities (Doppler technique)**
 - Astrometry (motion on the sky plane)
 - **Transit photometry (eclipses)**
 - Microlensing (bending of light)
 - Direct imaging
- *Kepler results* (next lecture)

FEBRUARY 5, 1996 \$2.95

**Dole Drops,
Clinton Rises**

TIME

IS ANYBODY OUT THERE?

How the discovery of two planets
brings us closer to solving the
most profound mystery in the cosmos


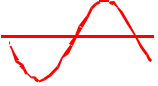
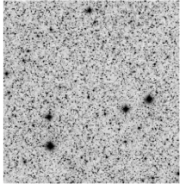

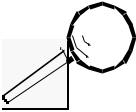
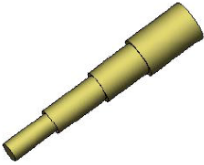


724404 1

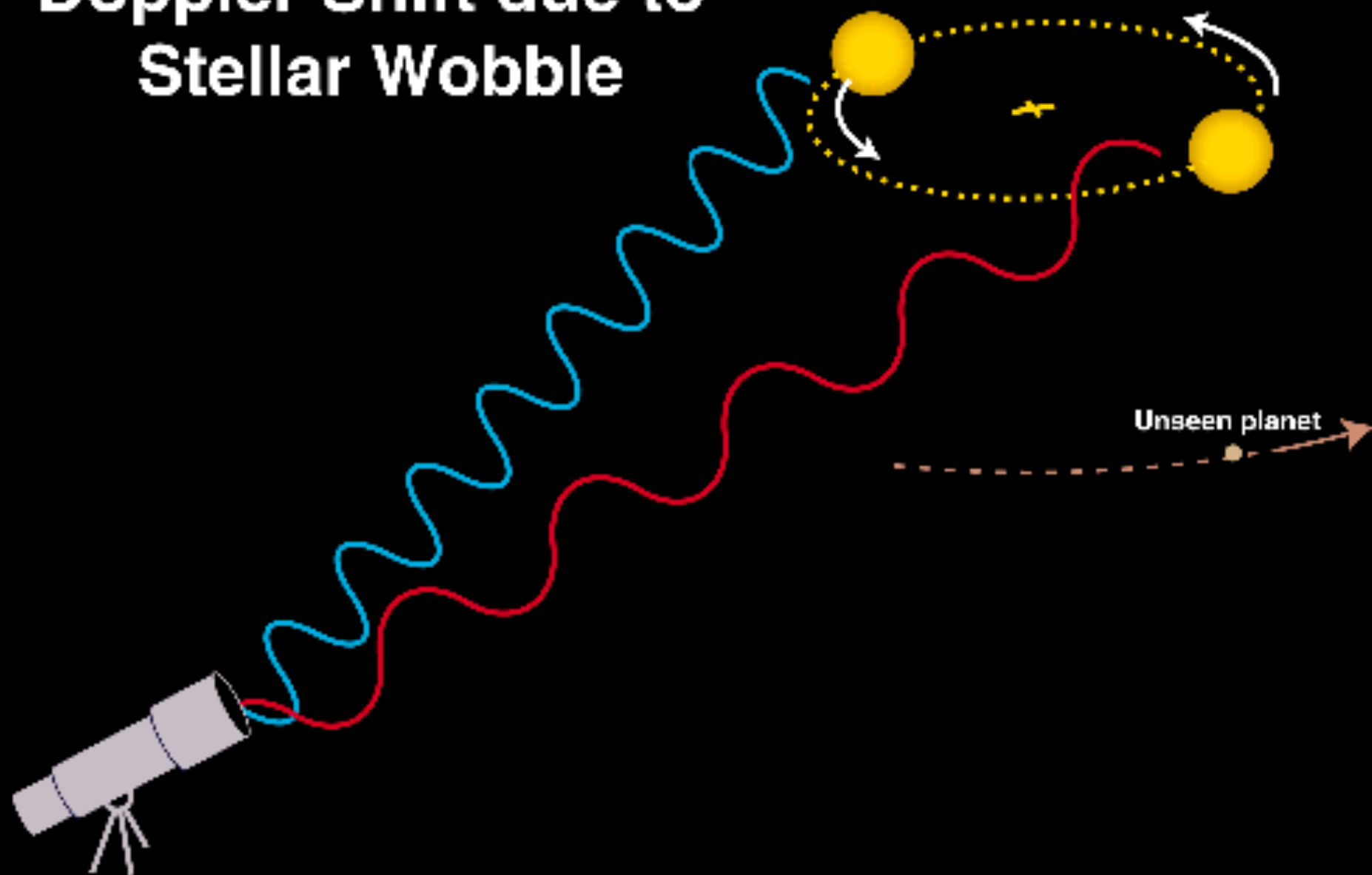
10099

0

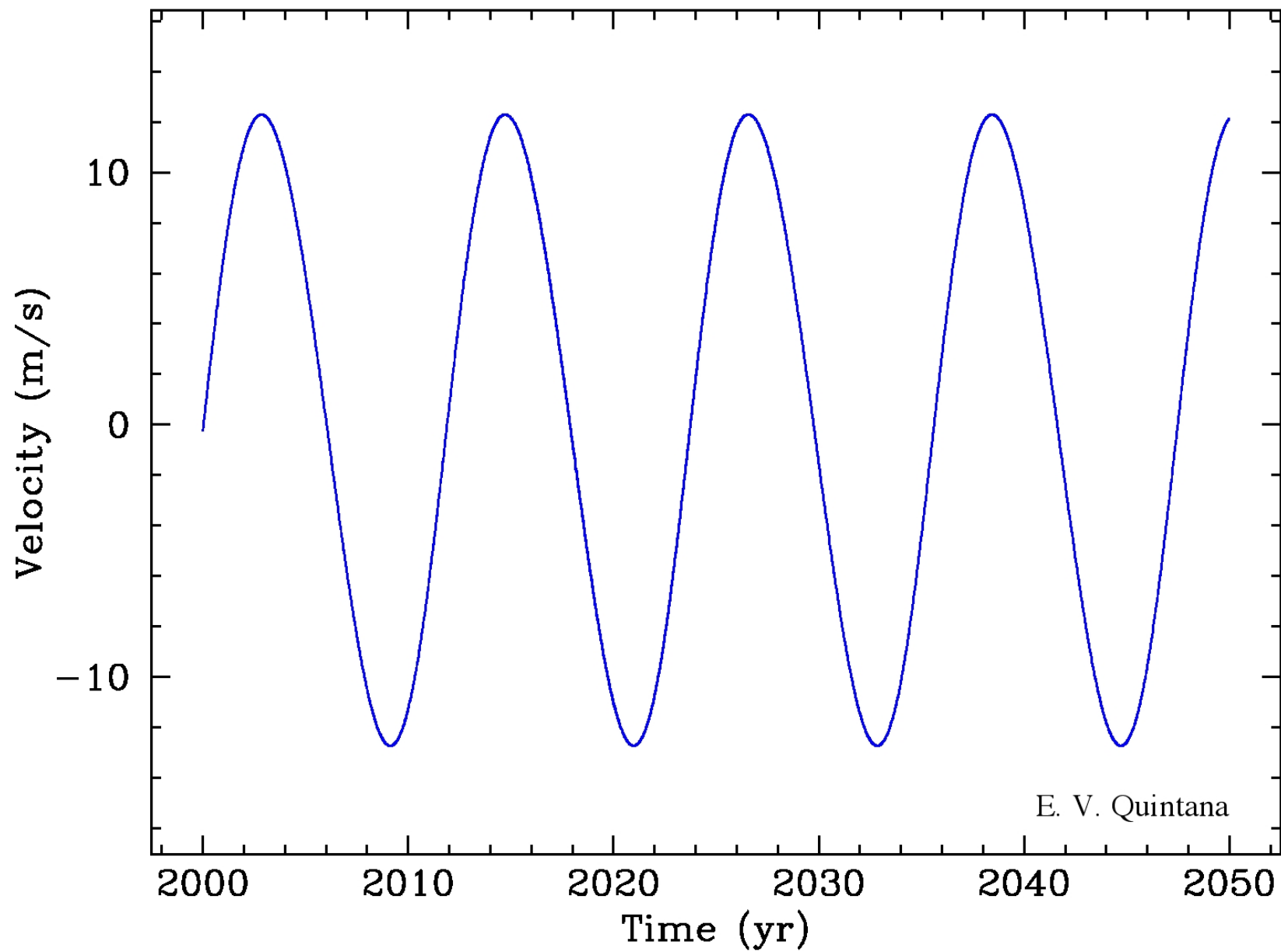
TECHNIQUES FOR FINDING EXTRASOLAR PLANETS

| | <u>Method</u> | <u>Yield</u> | <u>Mass Limit</u> | <u>Status</u> |
|---|--|---------------------------------------|---------------------------------------|--|
|  | Pulsar Timing | $m/M ; \tau$ | Lunar | Successful (3+~2) |
|  | Radial Velocity | $m \sin i ; \tau$ | super-Earth | Successful (~600) |
|  | Astrometry Ground: Space: | $m ; \tau ; D_s ; a$ | Jupiter sub-Jupiter | Ongoing Ongoing |
|  | Transit Photometry Ground Space, 27 cm Space, 1 m | $A ; \tau ; \sin i=1$ | sub-Jupiter sub-Neptune Mercury | Successful (~200) CoRoT (~30) Kepler (>1000 + >3000) |
|  | Microlensing: Ground | $f(m, M, r, D_s, D_L)$ super-Earth | super-Earth | Successful (~40) |
|  | Direct Imaging Ground Space | $albedo * A ; \tau ; D_s ; a ; M$ | Saturn Earth | Successful (>20) Being studied |

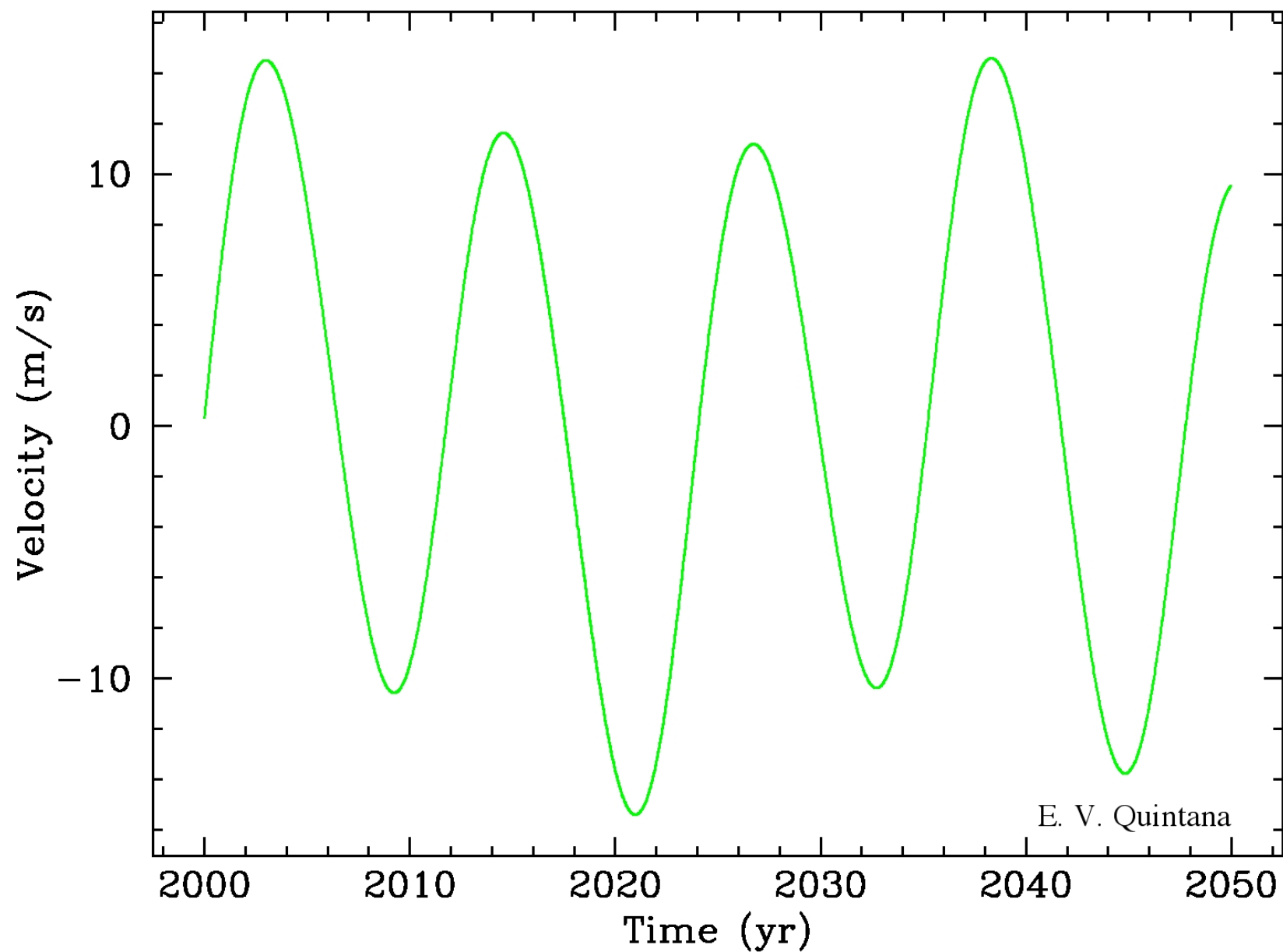
Doppler Shift due to Stellar Wobble



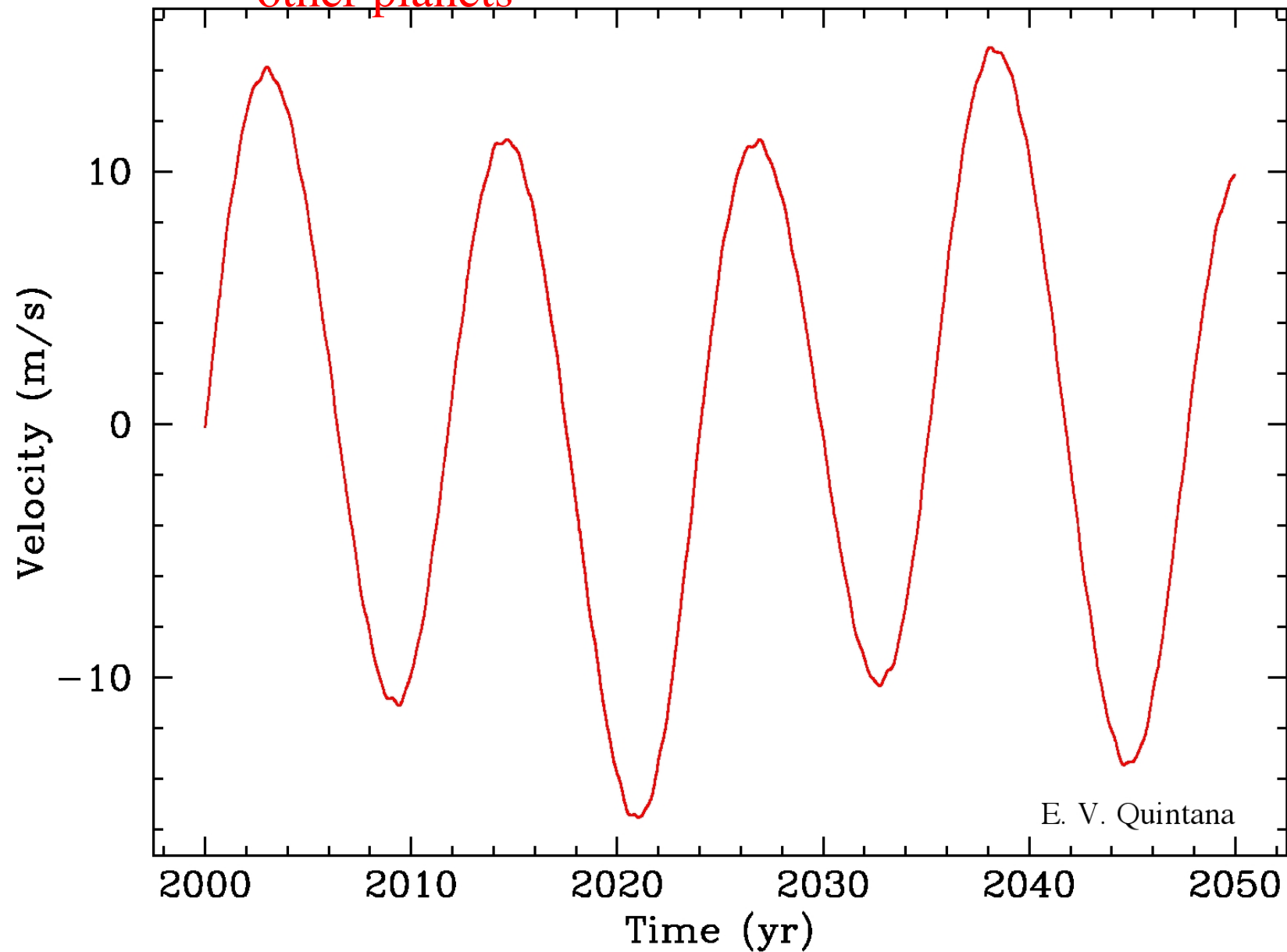
Radial velocity of Sun in response to Jupiter



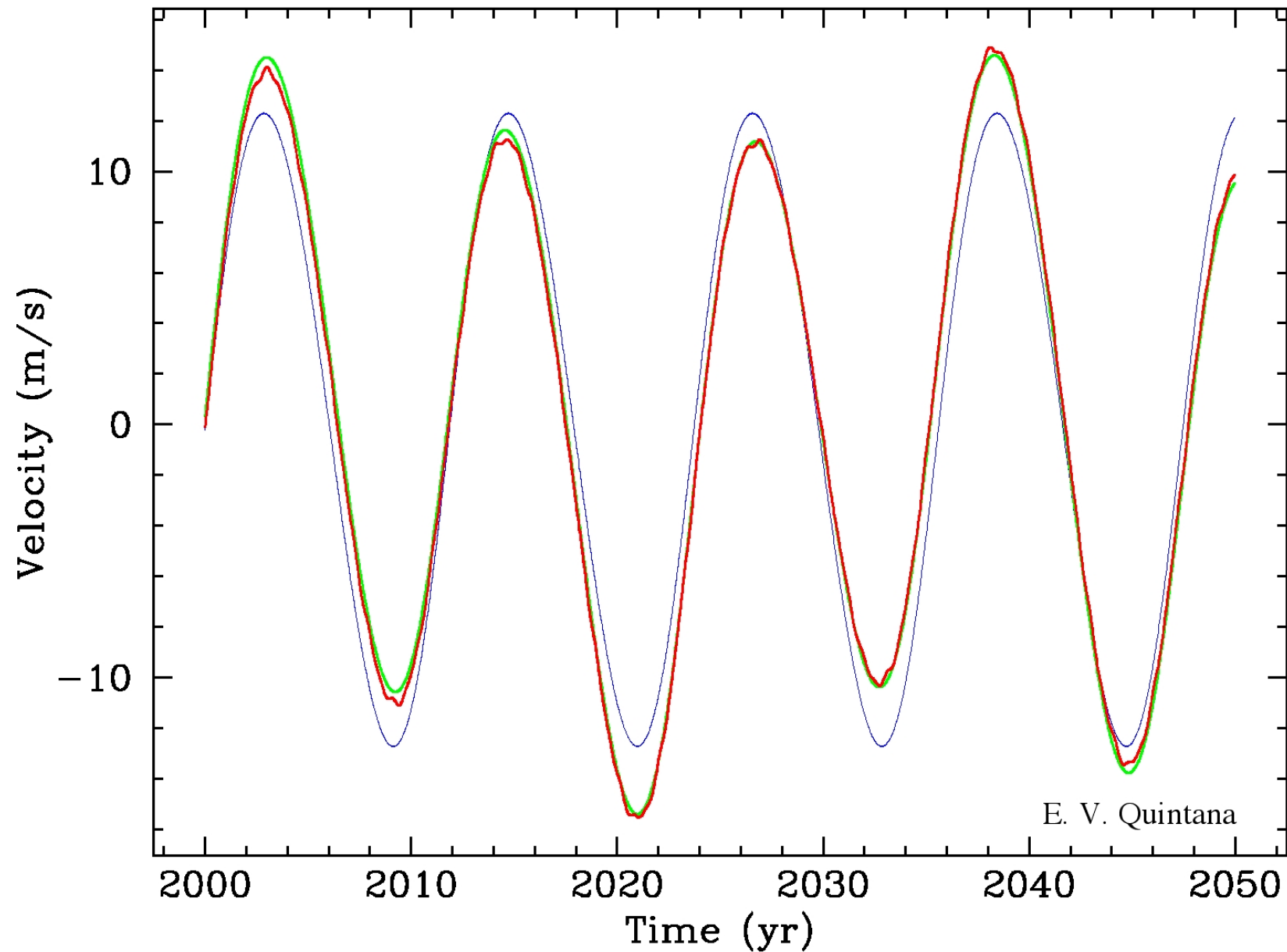
Radial velocity of Sun in response to Jupiter + Saturn



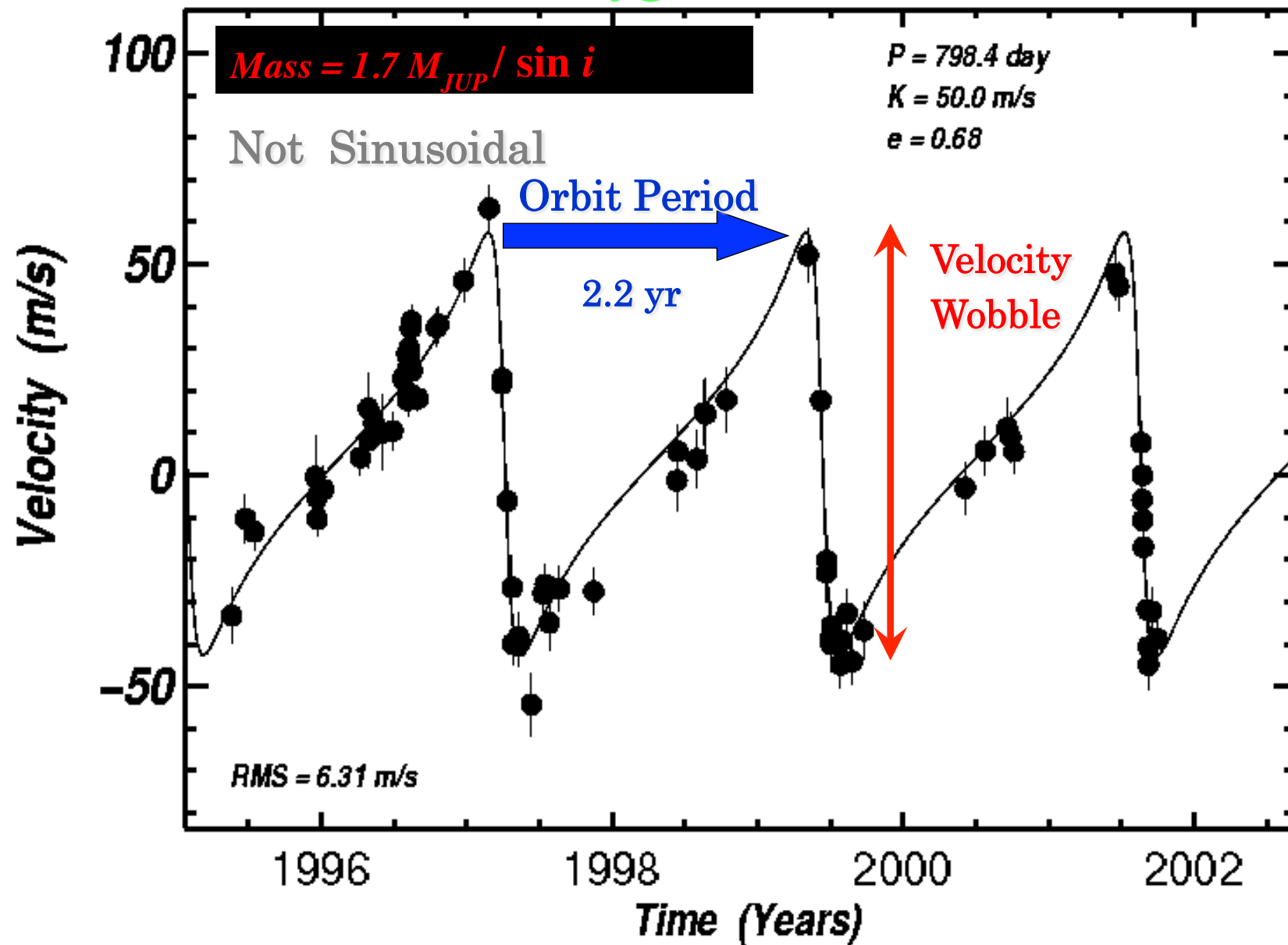
Radial velocity of Sun in response to **Jupiter + Saturn +
other planets**



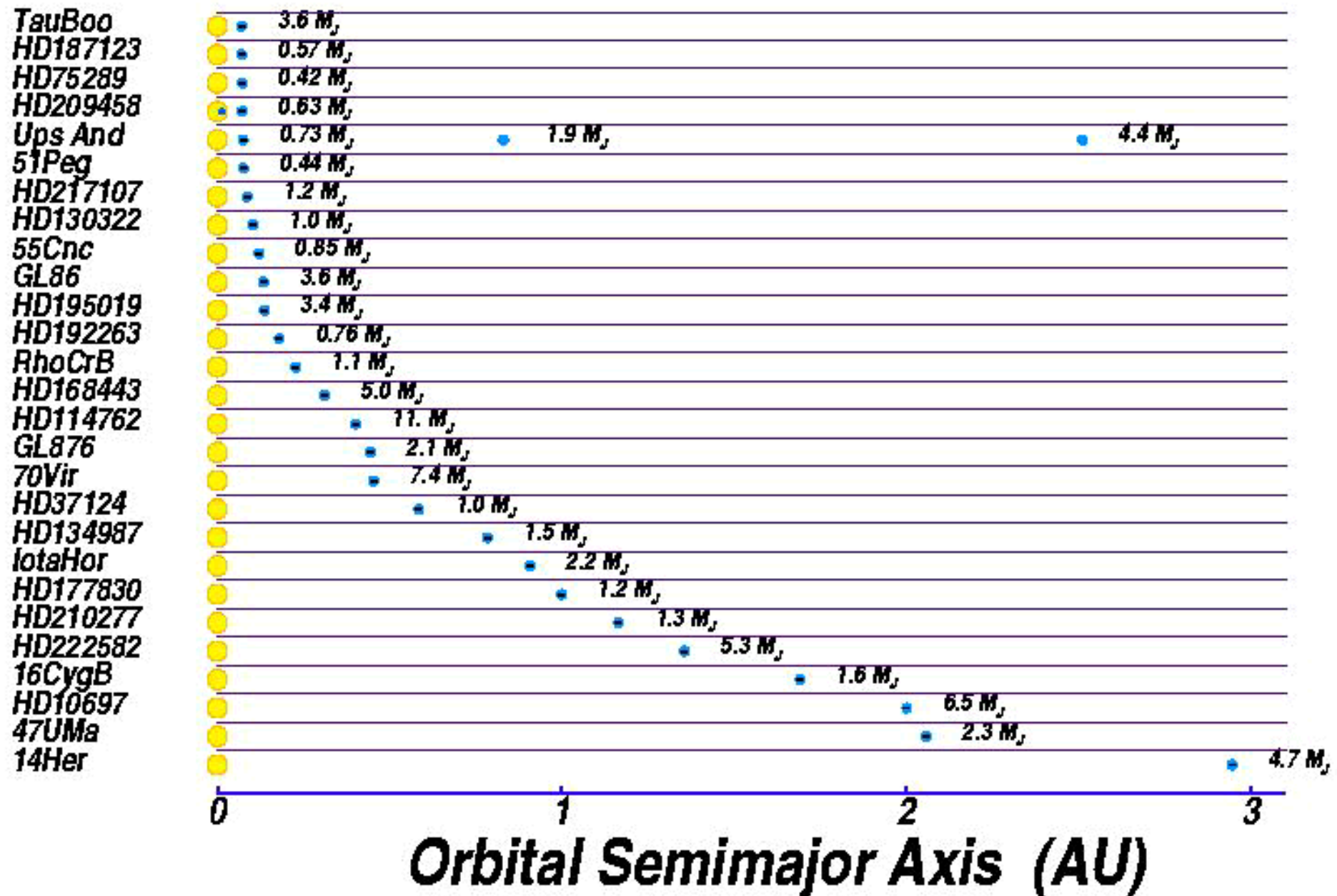
Radial velocity of Sun in response to Jupiter + Saturn + other planets

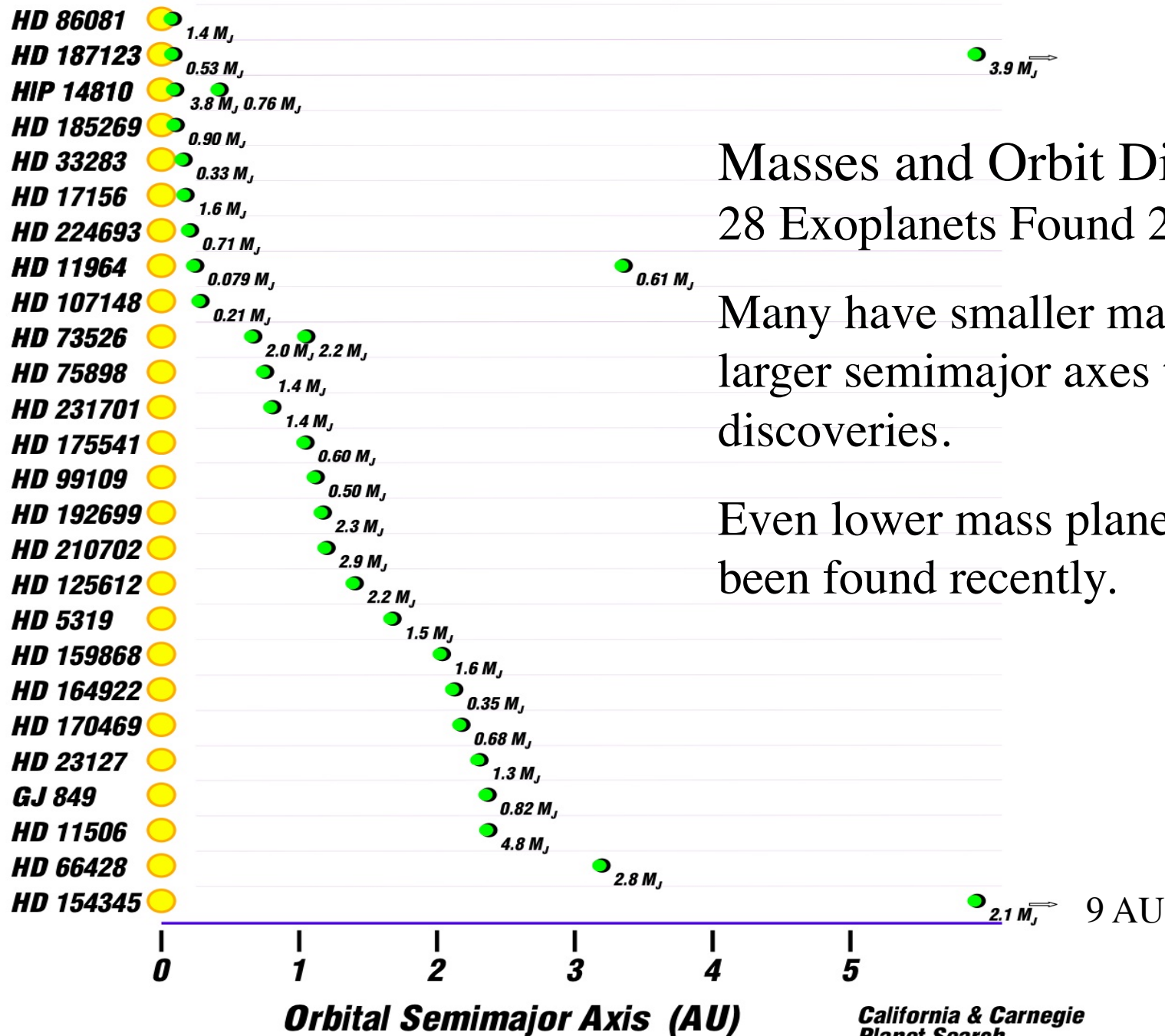


16 Cygni B



First 29 Exoplanets Discovered by Radial Velocity Surveys

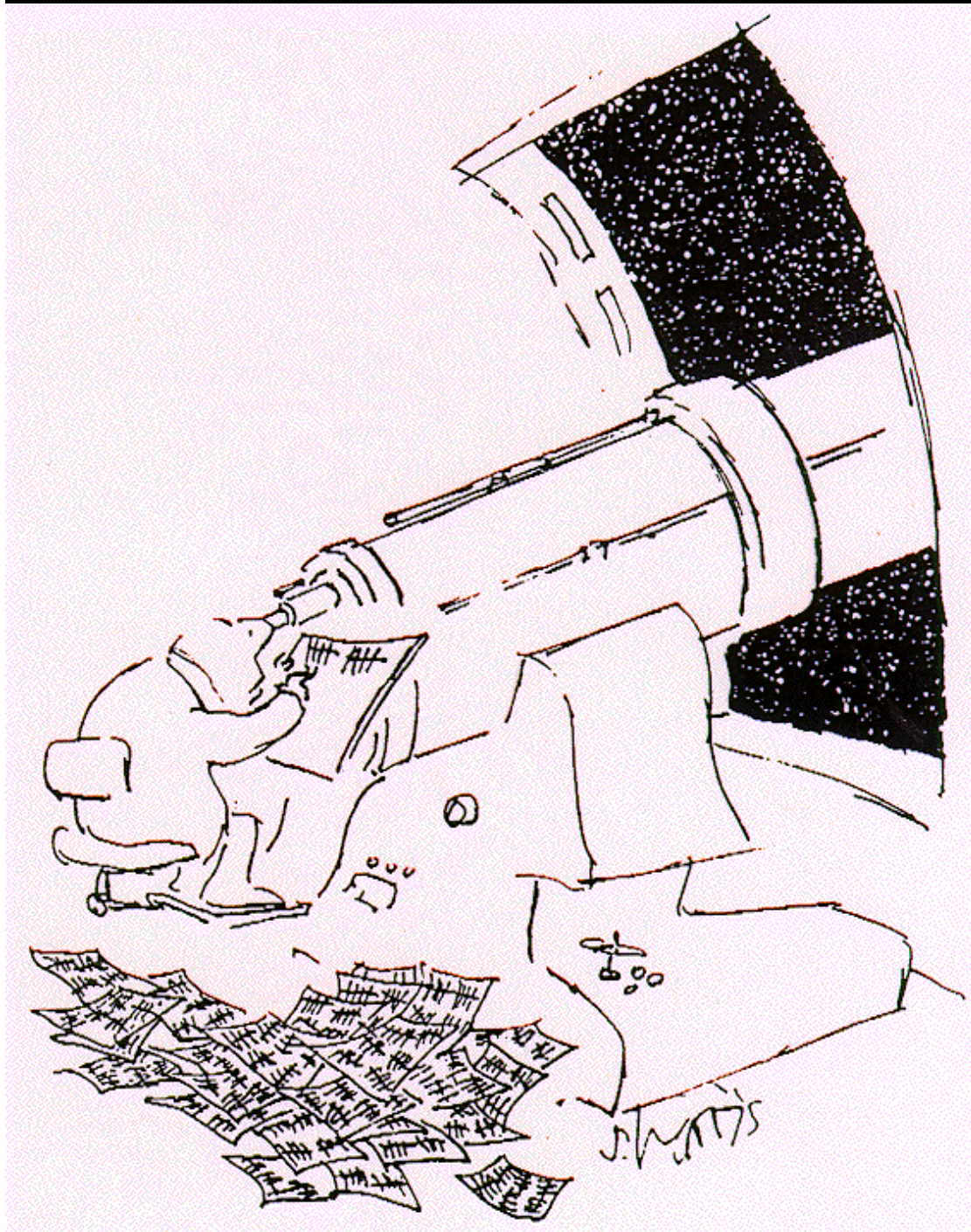




Masses and Orbit Distances of
28 Exoplanets Found 2006-2007

Many have smaller masses and/or
larger semimajor axes than early
discoveries.

Even lower mass planets have
been found recently.



How Many Known Extrasolar Planets?

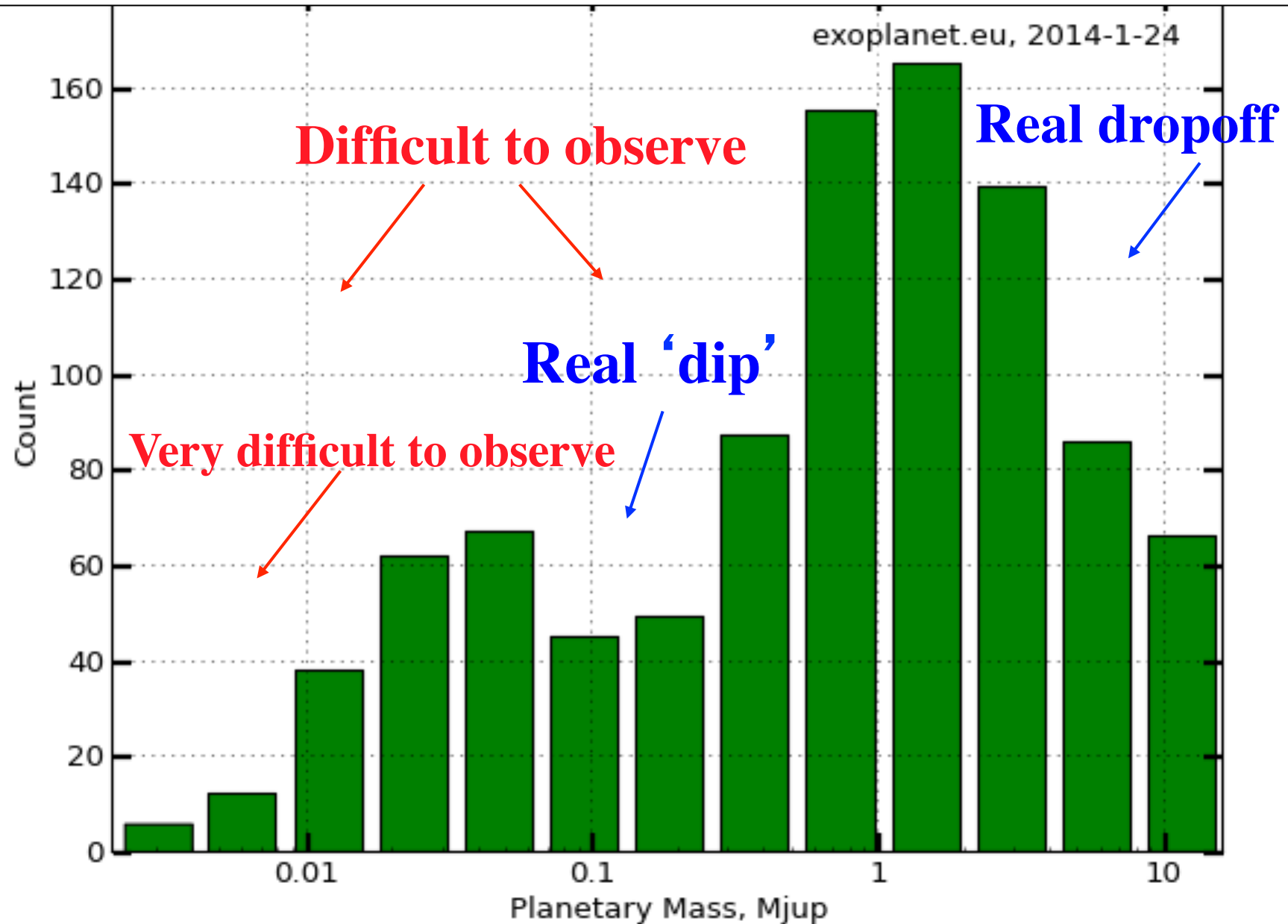
"What's one and one and
one and one and one and
one and one and one and
one and one?"

"I don't know," said Alice.
"I lost count."

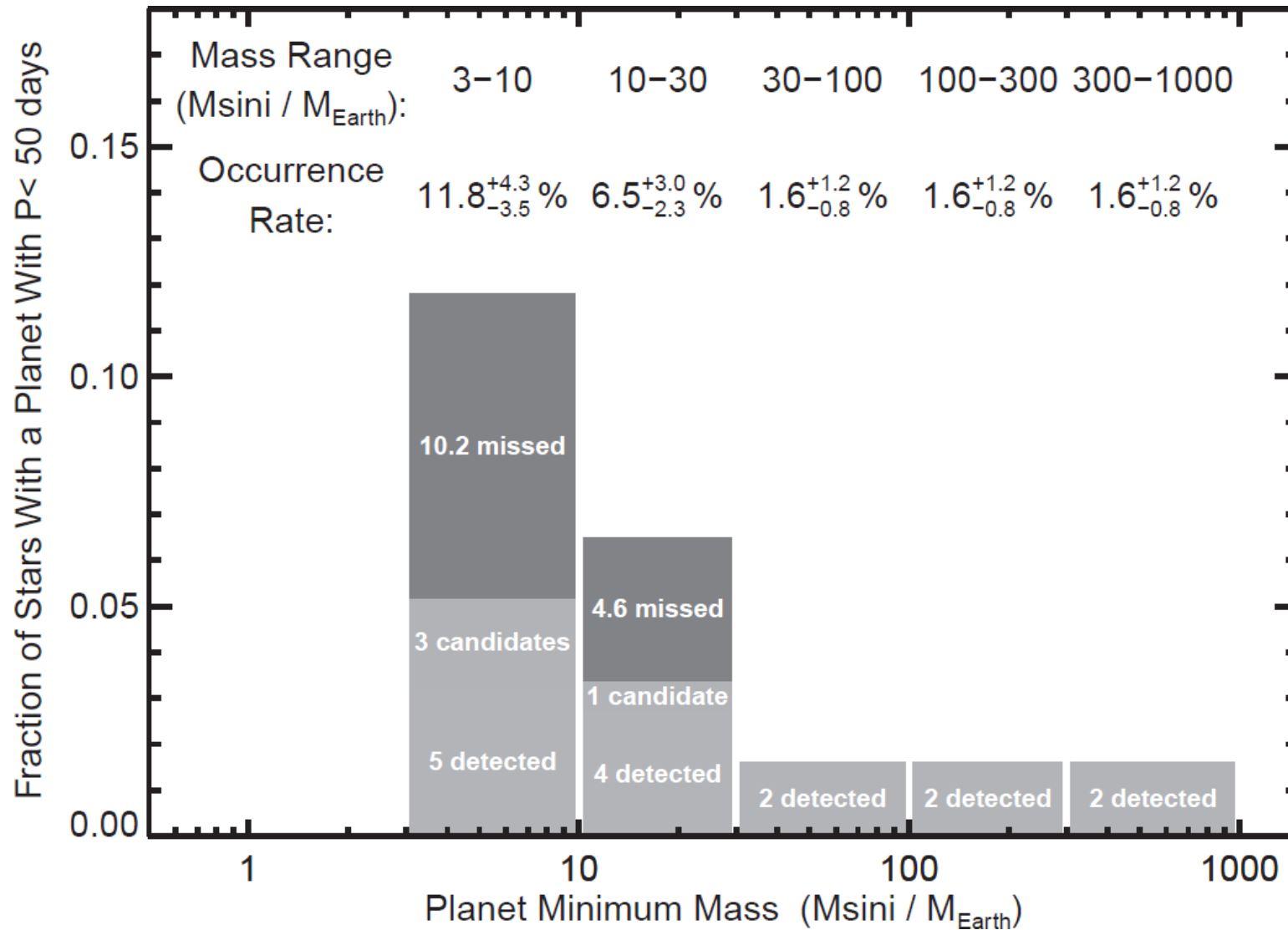
"She can't do addition,"
said the Red Queen.

Lewis Carrol,
Alice in Wonderland

Mass Distribution of RV Planets

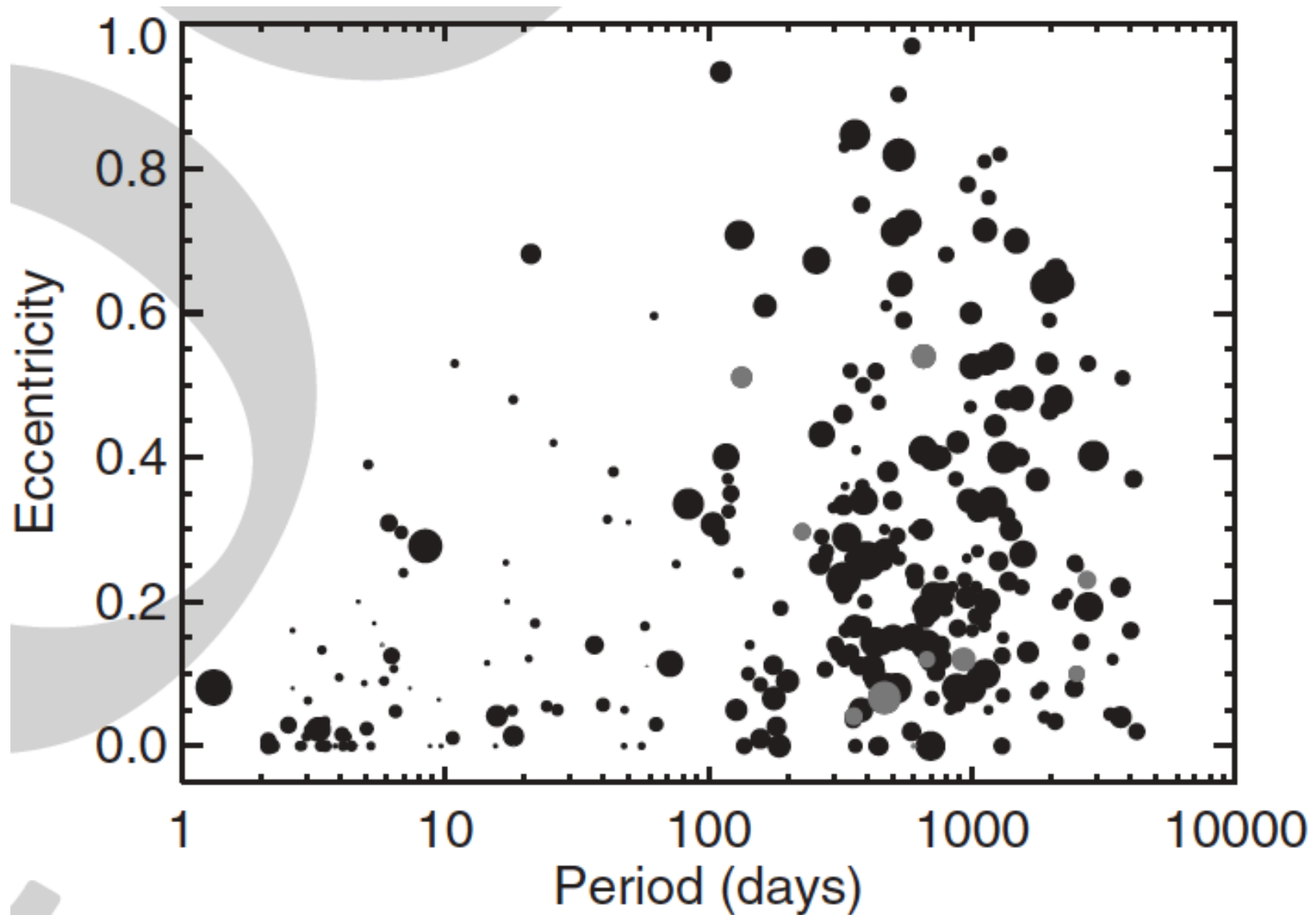


Small Planets are Numerous



Doppler/Keck (Howard et al. 2010)

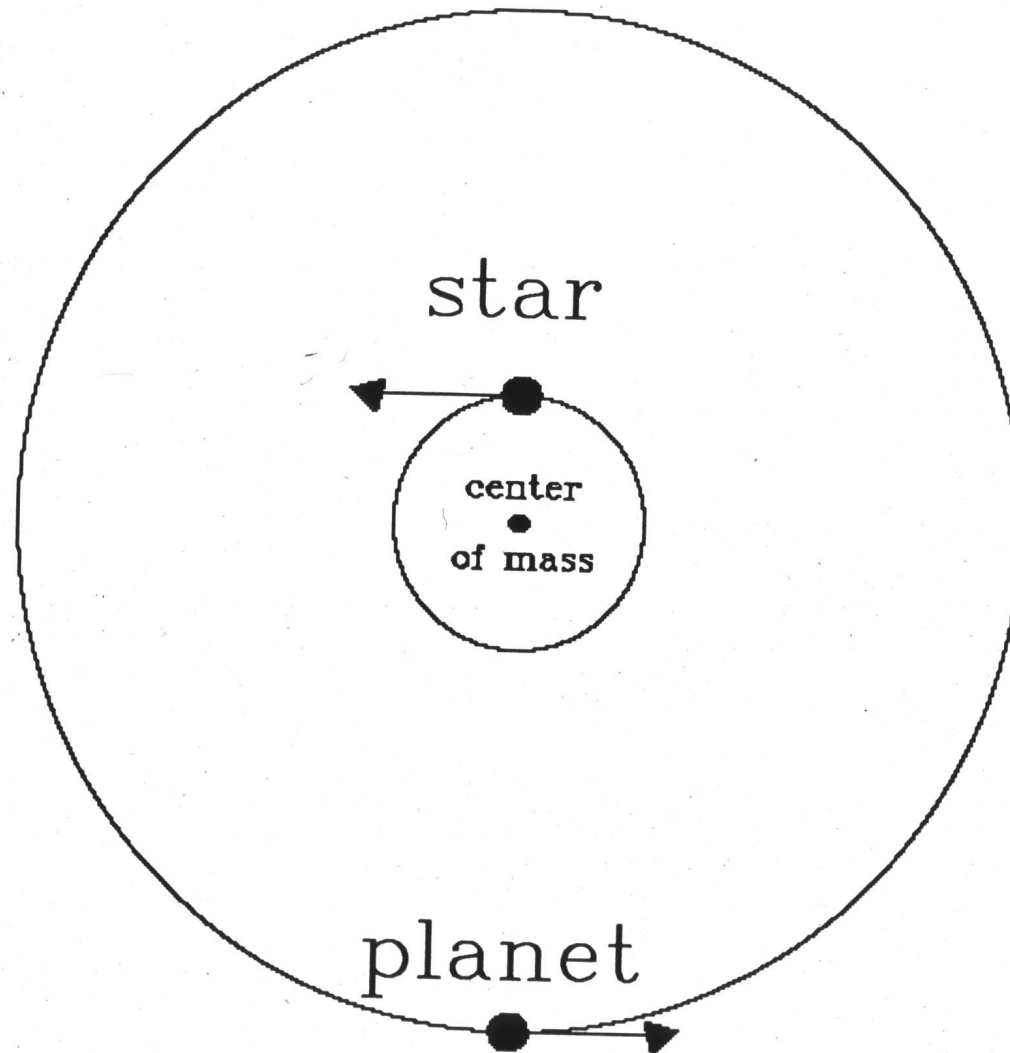
Eccentricities of RV Planets



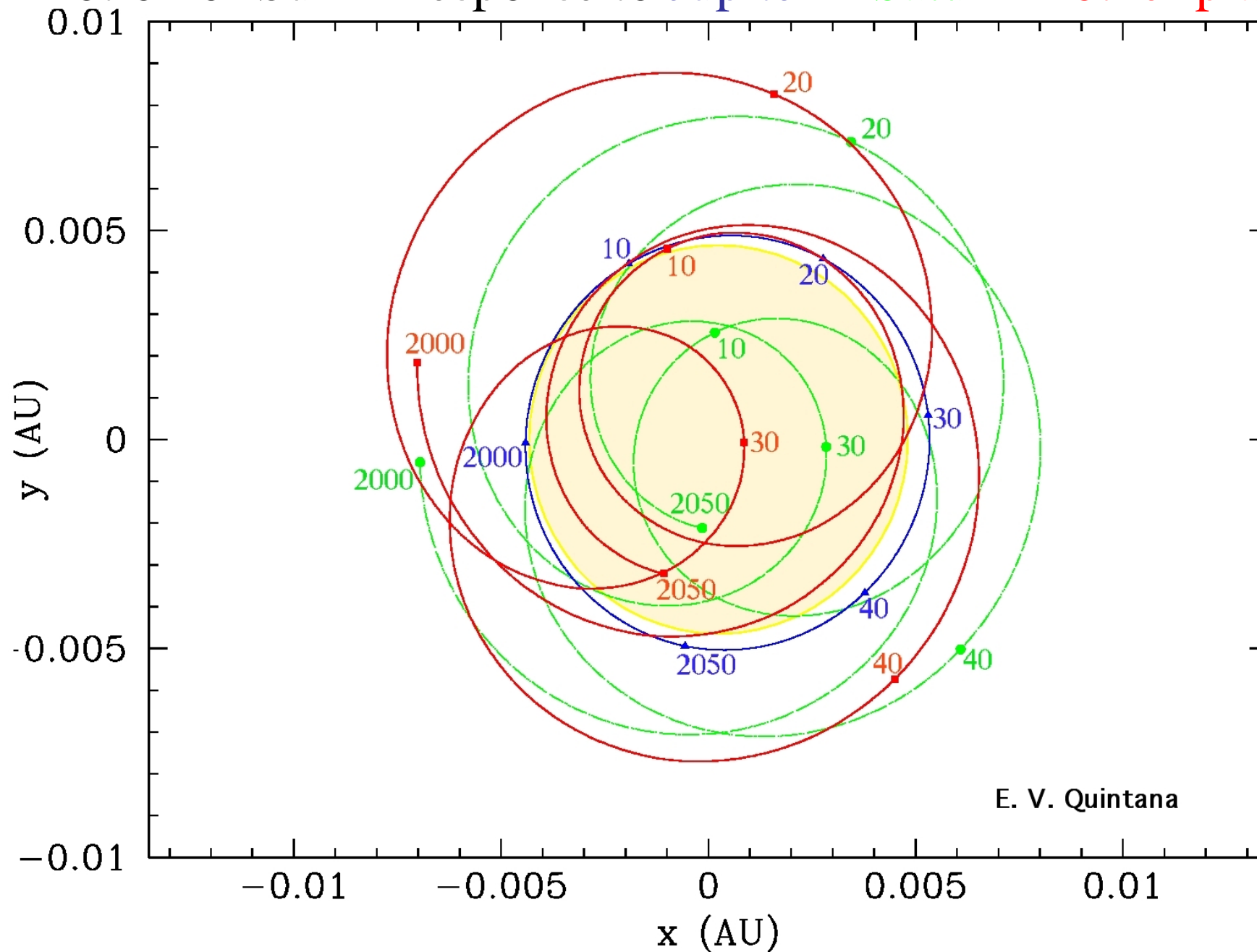
Black = single planets

Gray = planets in multis

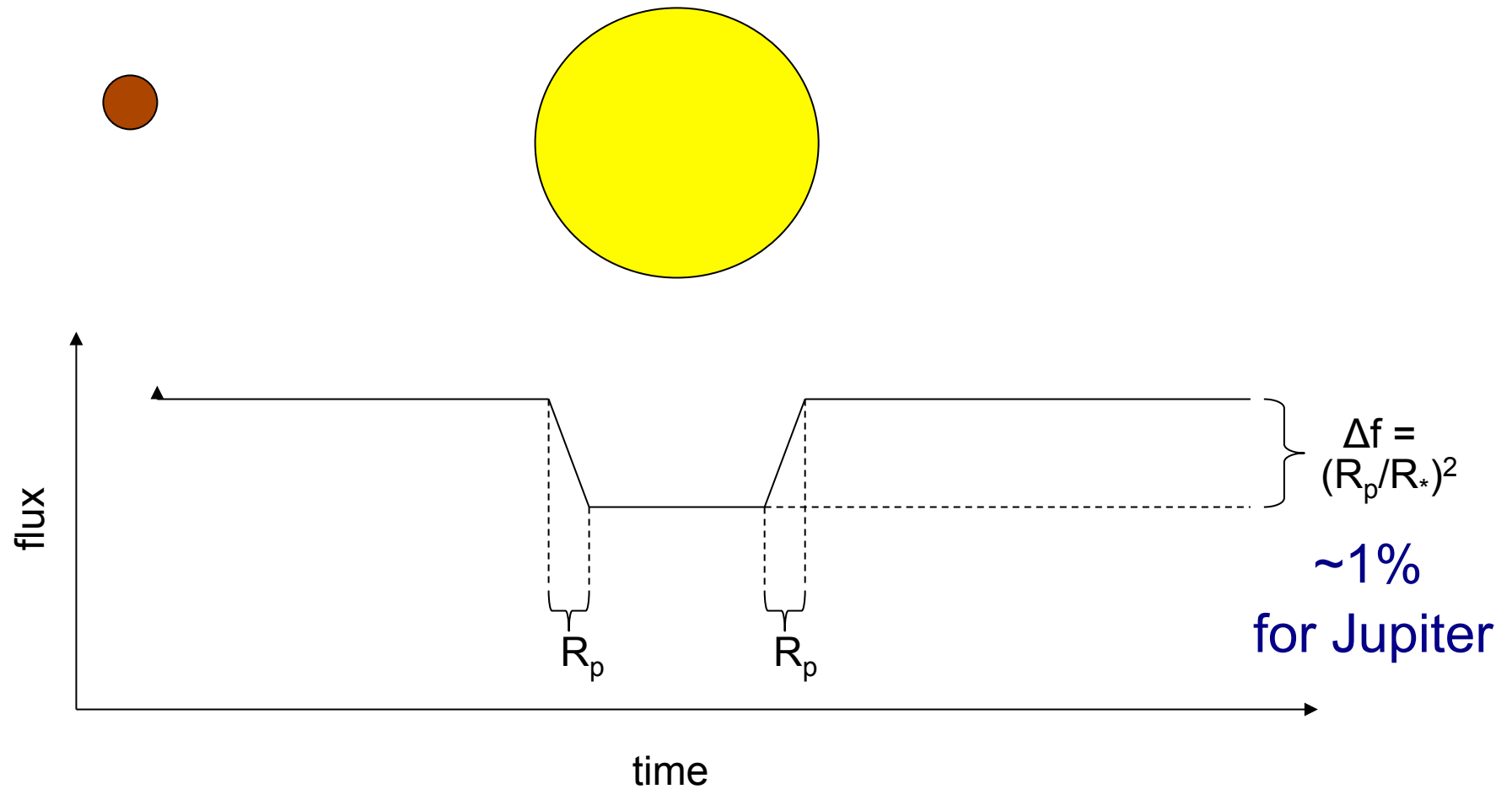
Astrometry



Motion of Sun in response to Jupiter + Saturn + other planets



Transit Lightcurve



2004 Venus Transit at Sunrise



2004 Venus Transit at Sunrise



Next Venus transit: December 2117

Mercury Transit of 2016 May 09

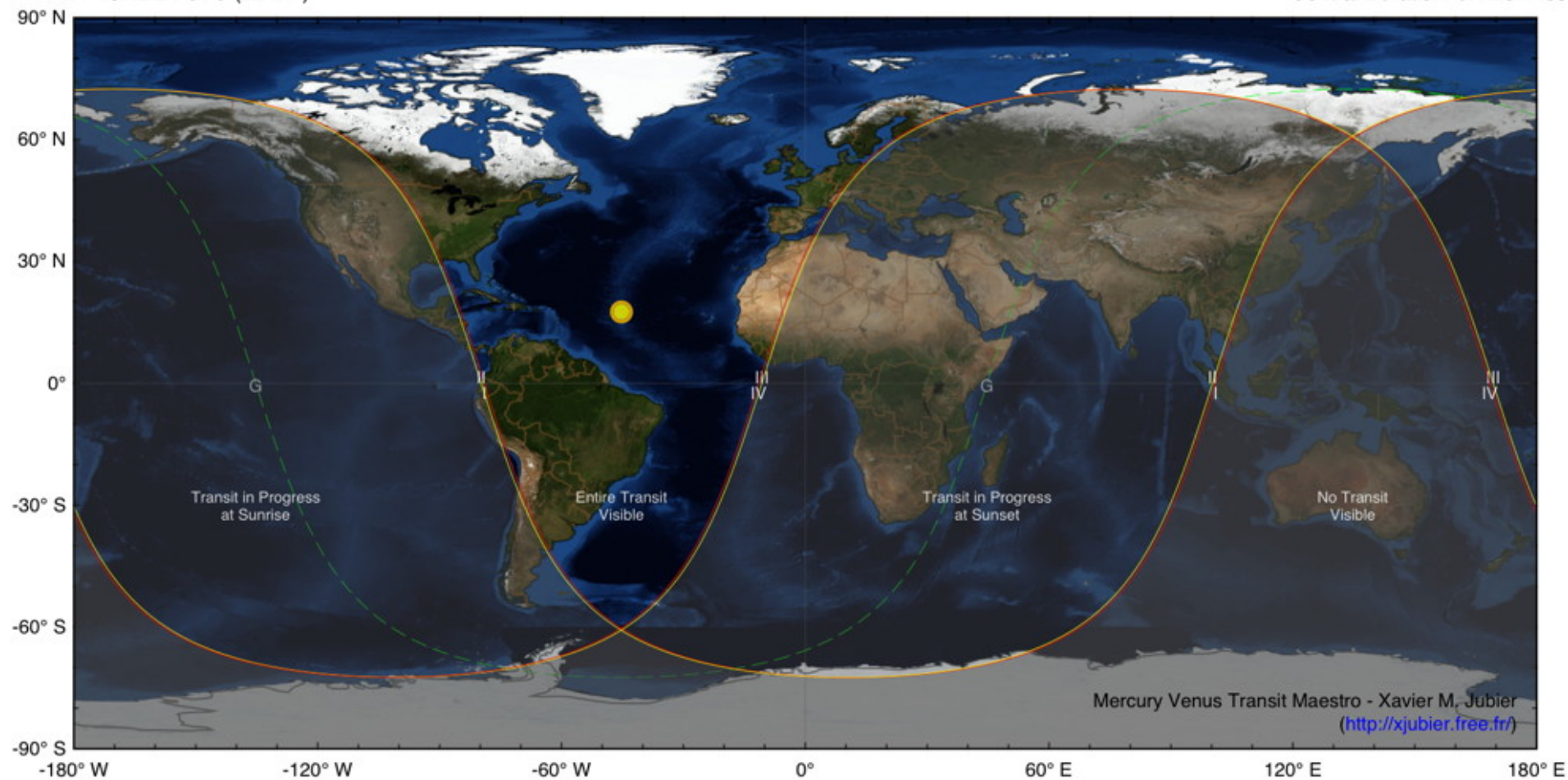
Greatest Transit: 14:57:24.8 UT J.D.: 2457518.123203
 ΔT : 68.30s

Transit Geocentric Contacts

I: 11:12:18 UTC (83.2°)
II: 11:15:30 UTC (83.5°)
G: 14:57:25 UTC (153.8°)
III: 18:39:12 UTC (224.1°)
IV: 18:42:24 UTC (224.4°)

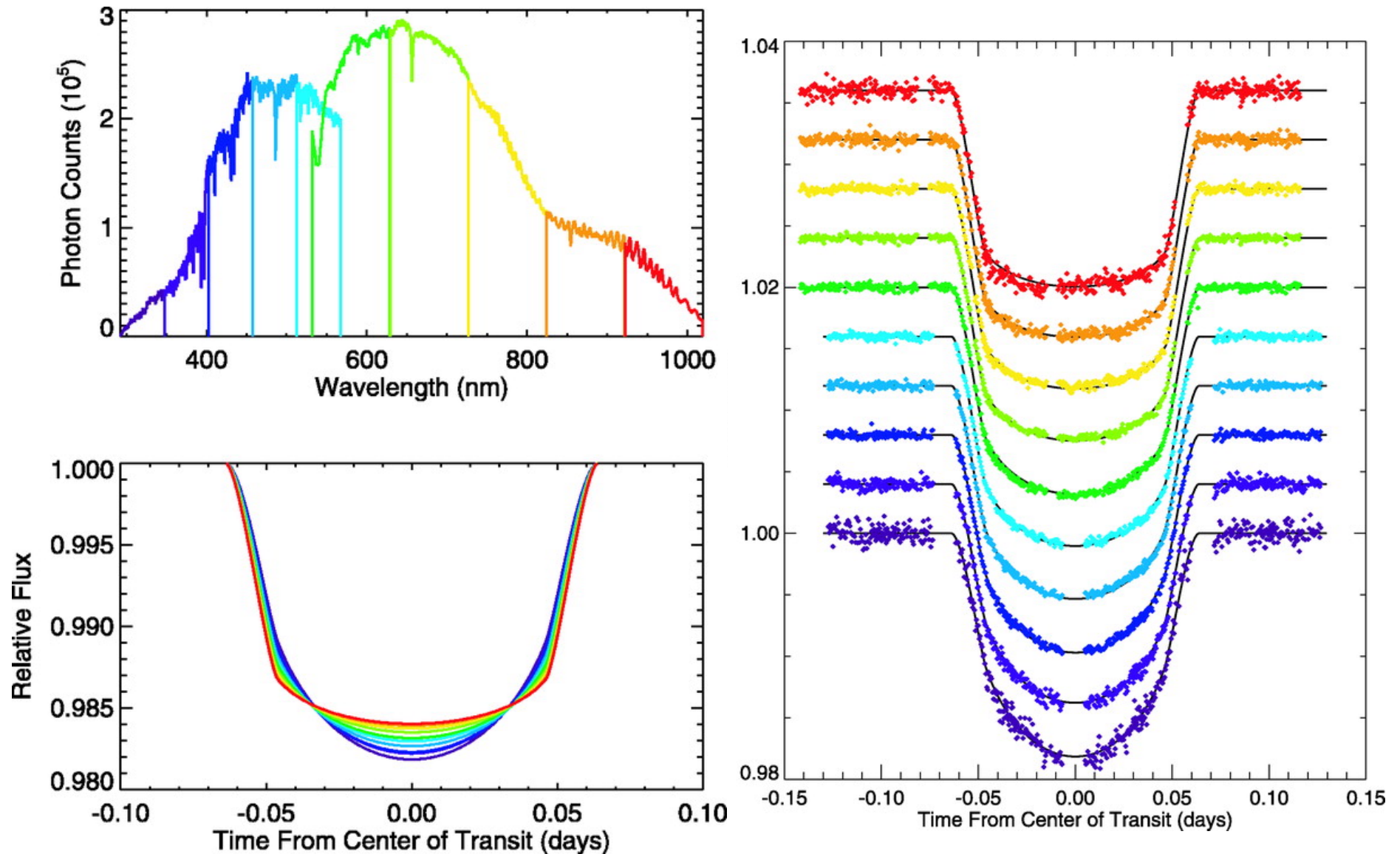
Geocentric Data

Minimum separation: 318.5"
General Duration: 07h30m05s
Central Duration: 07h23m43s



HD 209458 b Transits at Differing Wavelengths - HST

(Knutson et al. 2007)



Direct imaging of exo-planets is Difficult:

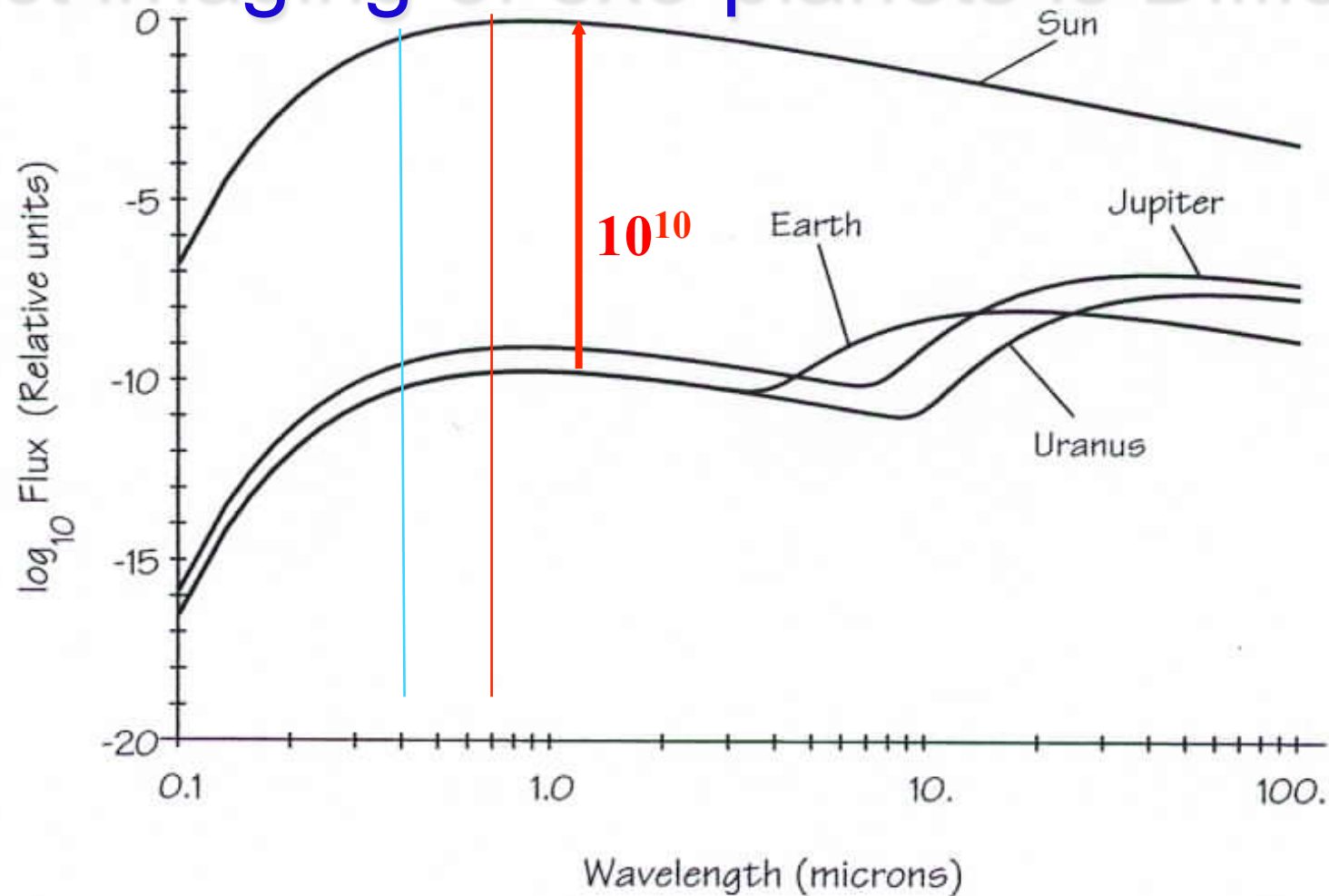
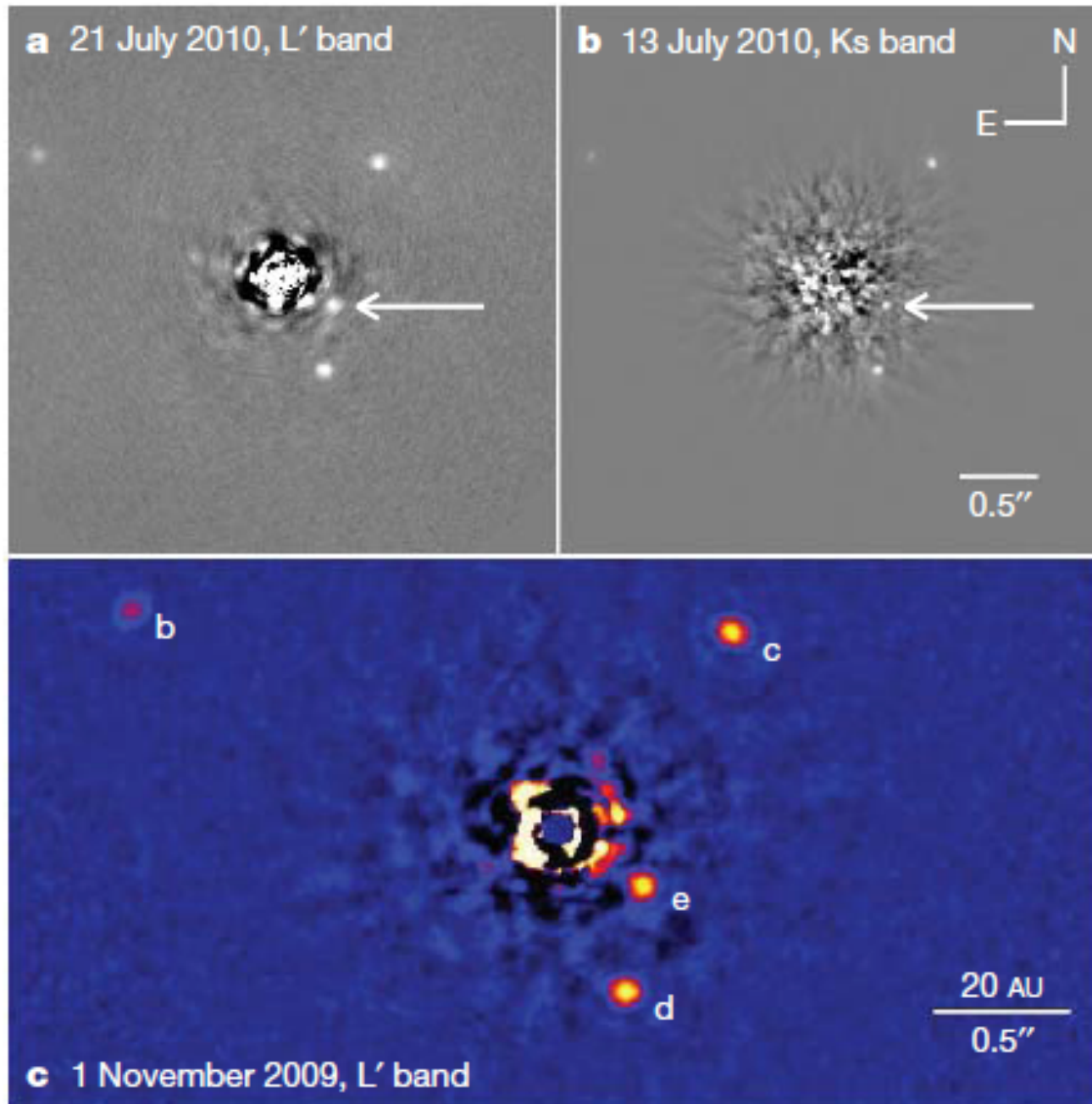
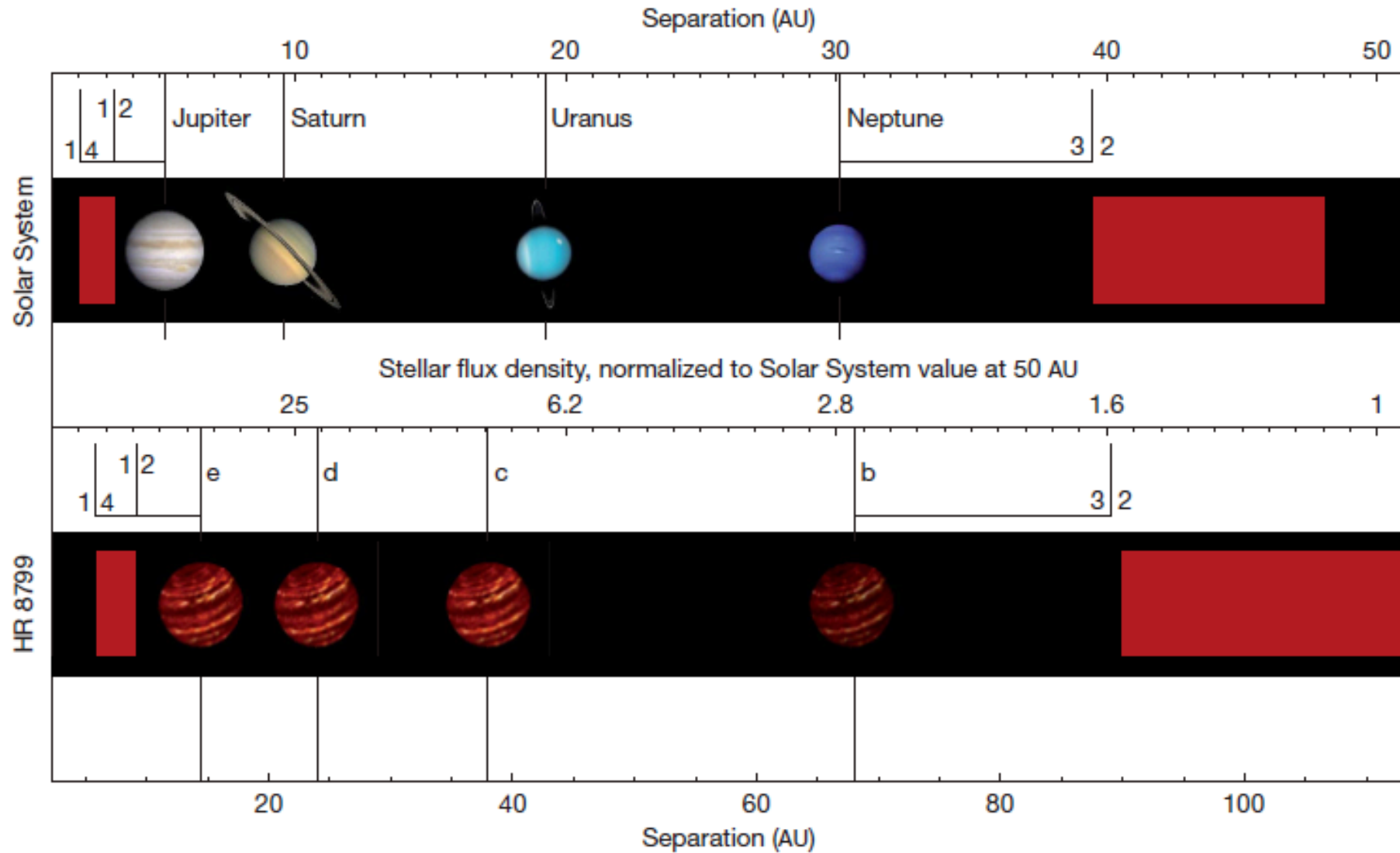


Figure 4-1. The spectral energy distributions of the Sun, Jupiter, Earth, and Uranus as they would appear at 5 pc, averaged over a 10% spectral bandpass. Note the decreased ratio of solar to planetary flux in the thermal infrared, compared to visible wavelengths.

HR 8499: 4 distant giant planets



HR 8499 vs Solar System



RV Extrasolar Planets

- **~ 0.7% of Sun-like (late F, G & early K dwarf) stars have planets more massive than Saturn within 0.1 AU**
 - Transiting planets known to be gas giants; HD 149026b is metal-rich
 - Models suggest these planets migrated inwards
- **~ 7% of Sun-like stars have planets more massive than Jupiter within 2 AU**
 - Many of these planets have very eccentric orbits
- **< 2% of M dwarf stars have planets more massive than Jupiter within 1 AU**
- **Stars with higher metallicity are more likely to host giant planets**
- **Stars with one detectable planet are more likely to host more detectable planets**
- **> a few % of stars have Jupiter-like companions ($0.5 - 2 M_{\text{Jup}}$, $4 \text{ AU} < a < 10 \text{ AU}$), but > 25% do not**
- **Brown dwarf desert; low-mass planets most common**



**Artist Conceptions
(Science fiction)**

