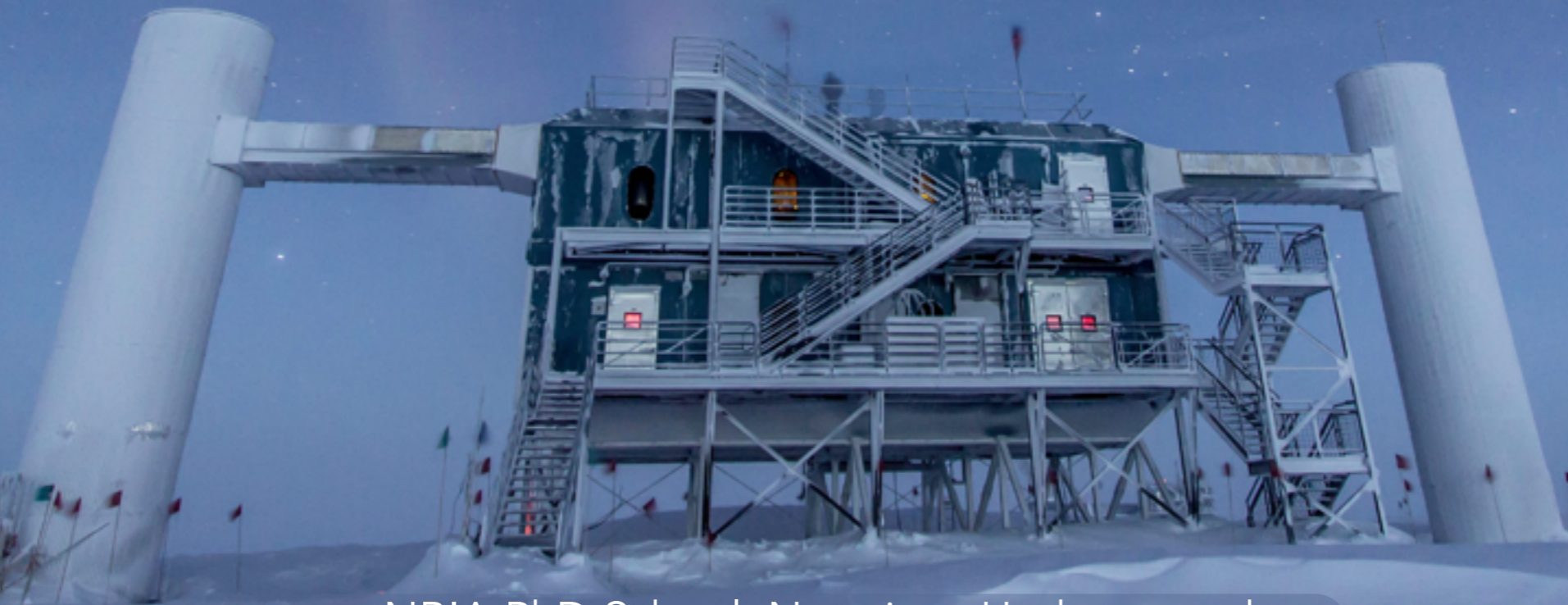


Neutrino Oscillations w/ PINGU

D. Jason Koskinen
University of Copenhagen
Niels Bohr Institute



NBIA PhD School: Neutrinos Underground
and in the Heavens II

August, 2016

VILLUM FONDEN



*Felipe Pedreros, IceCube/NSF

Neutrino Oscillations w/ PINGU, IceCube/DeepCore and other stuff too

D. Jason Koskinen
University of Copenhagen
Niels Bohr Institute



NBIA PhD School: Neutrinos Underground
and in the Heavens II

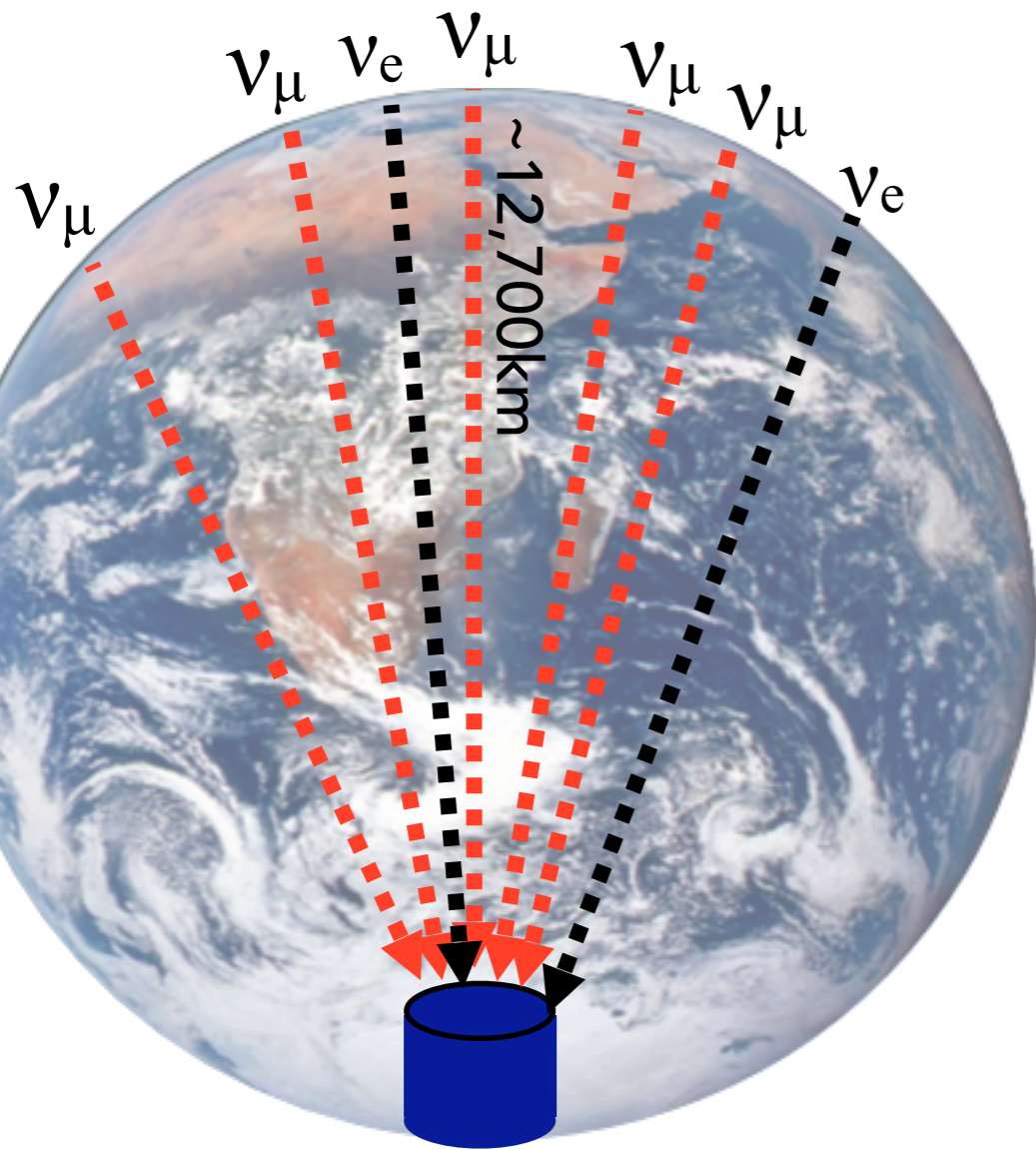
August, 2016

VILLUM FONDEN

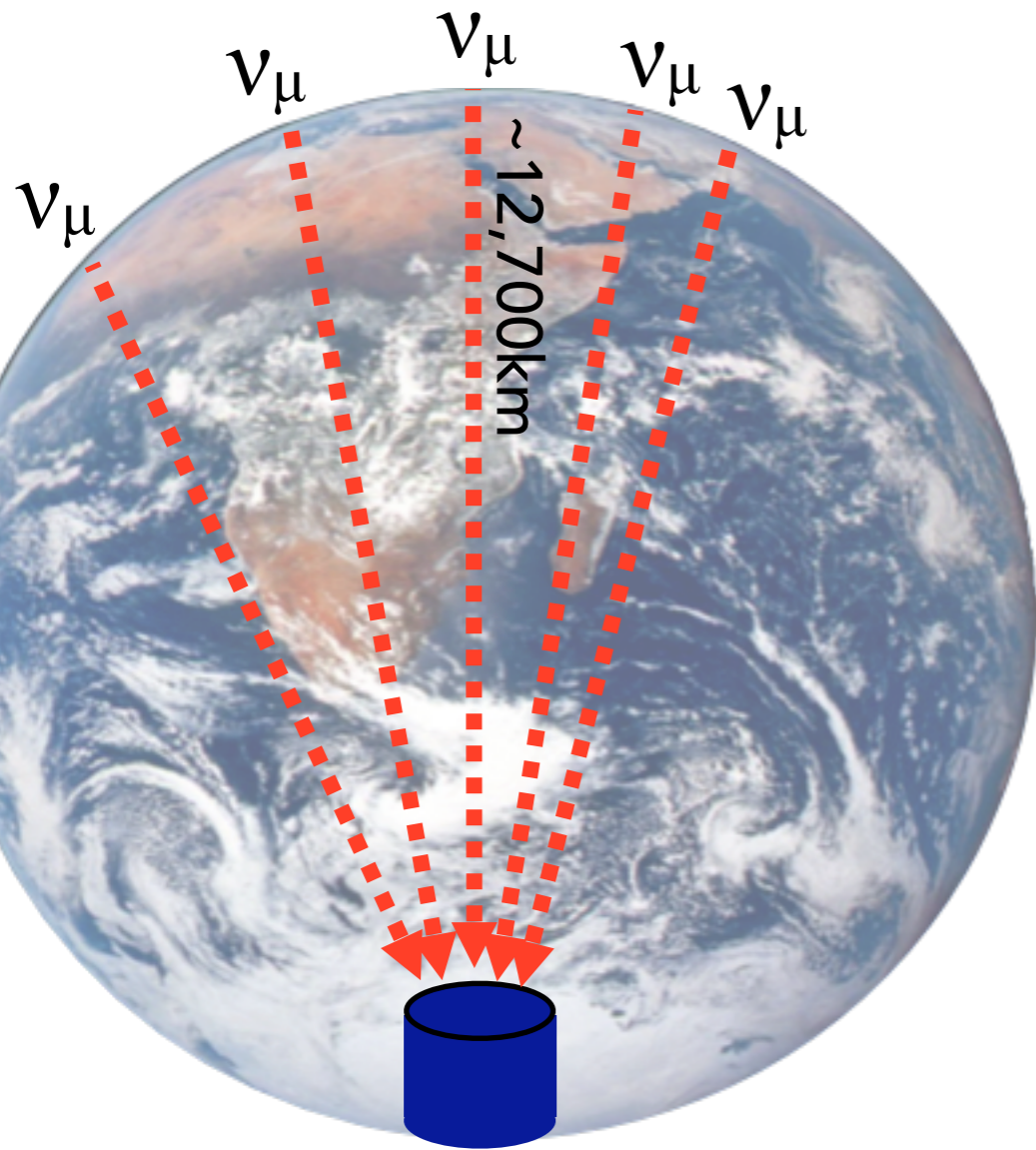


*Felipe Pedreros, IceCube/NSF

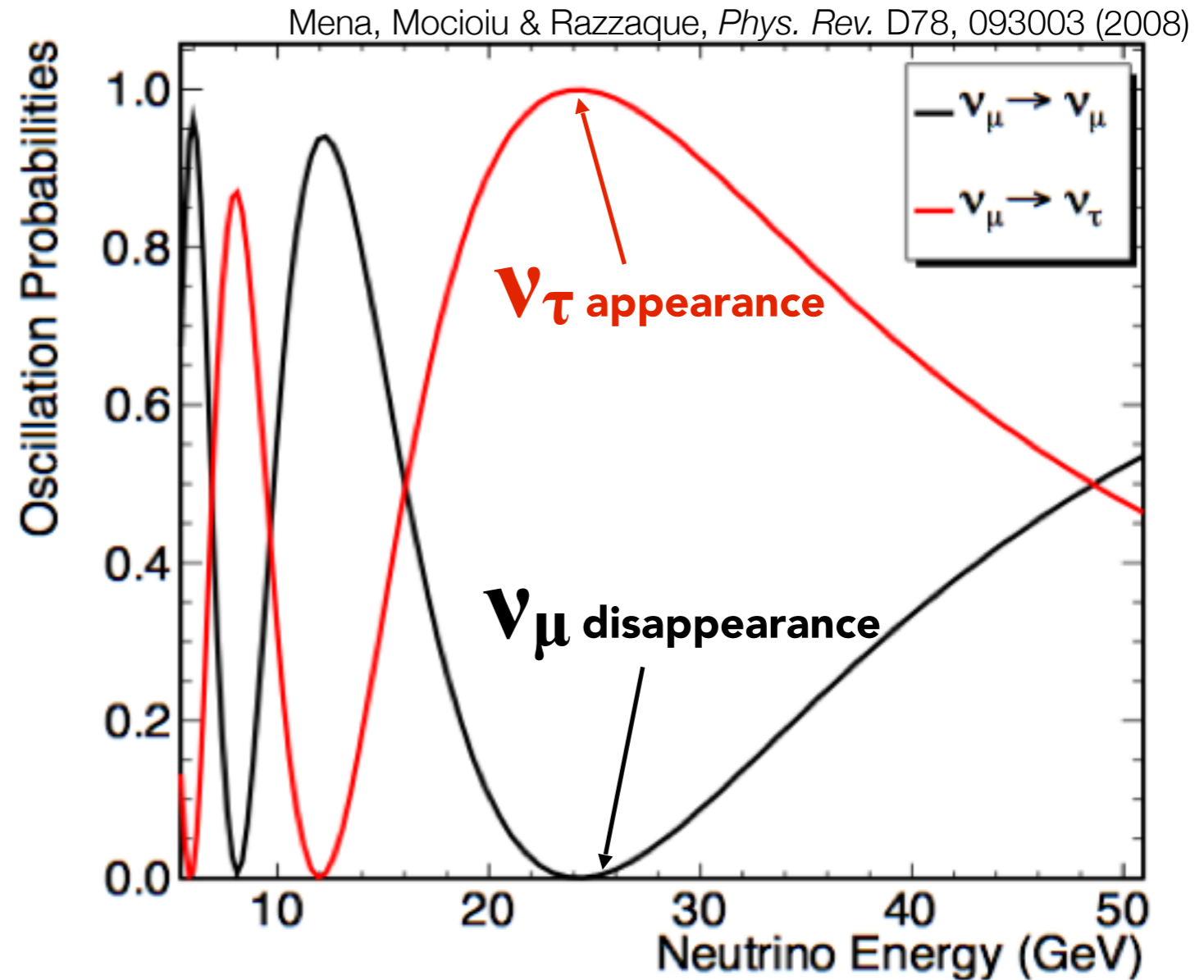
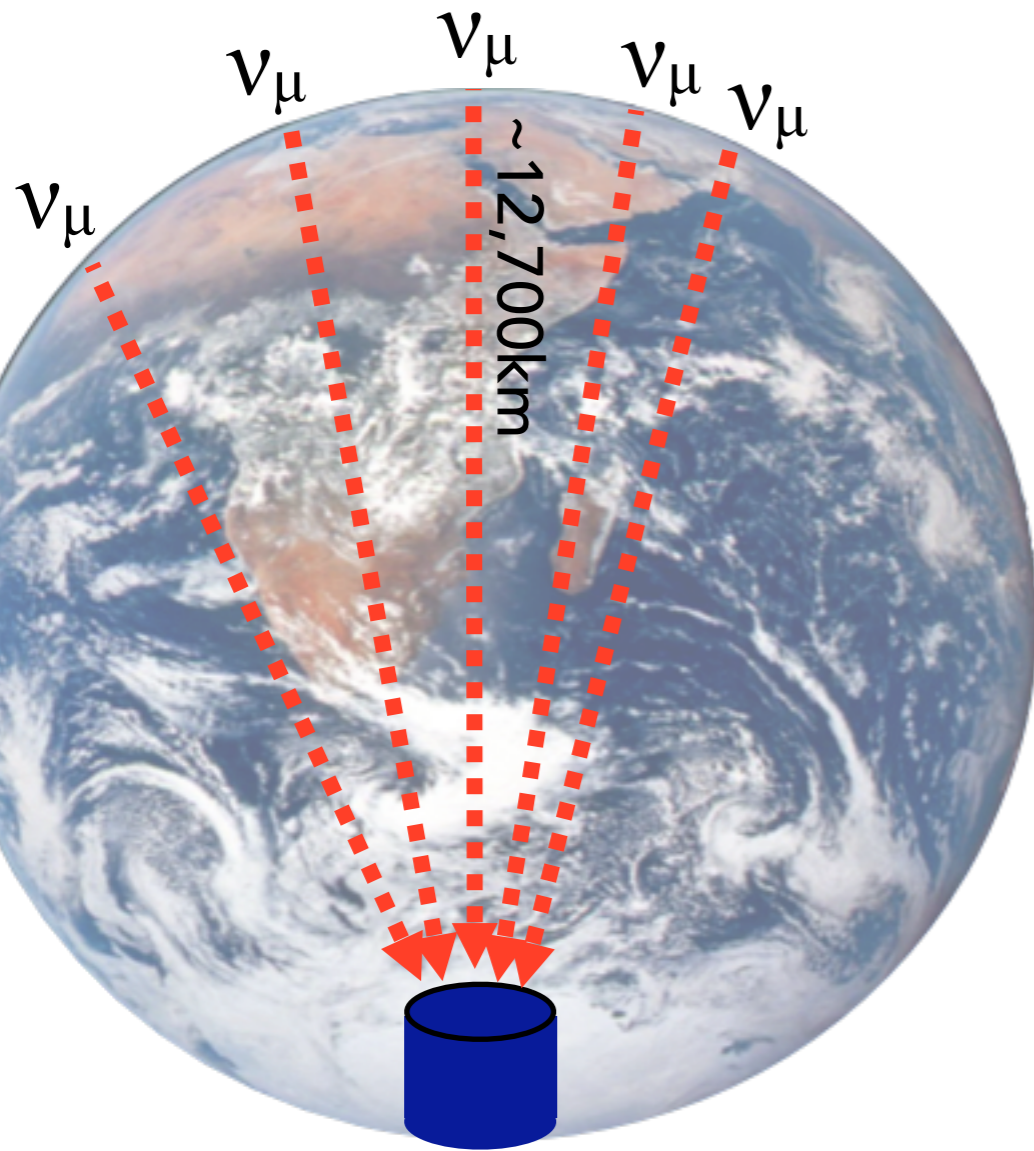
Atmospheric Neutrino Oscillation



Atmospheric Neutrino Oscillation



Atmospheric Neutrino Oscillation



- Osc. probabilities $P(\nu_\mu \rightarrow \nu_\mu)$ at earth diameter baselines produce 1st oscillation maximum/minimum at ~ 25 GeV

Charged particles traveling through water/ice produce Cherenkov radiation

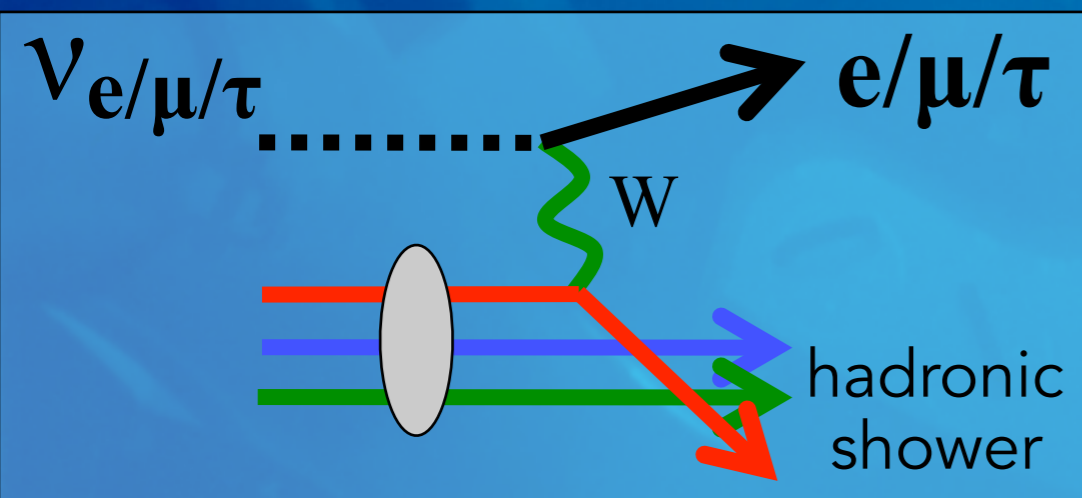
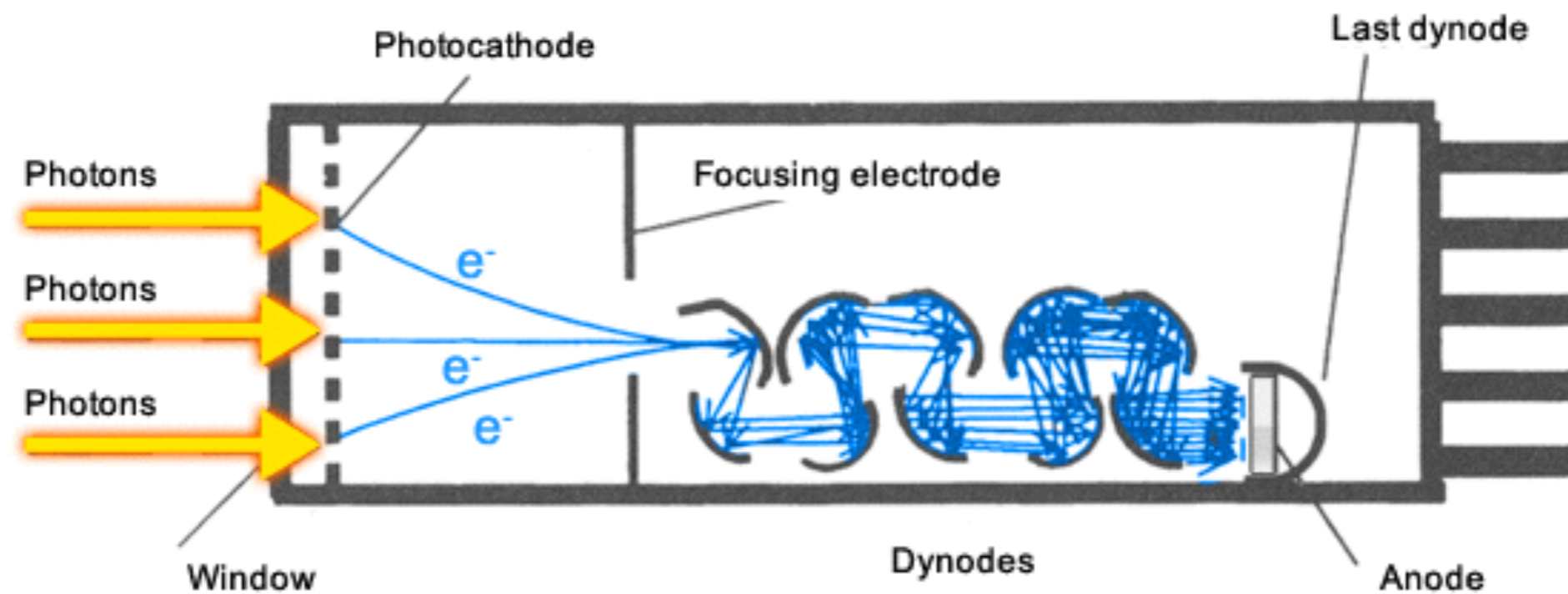


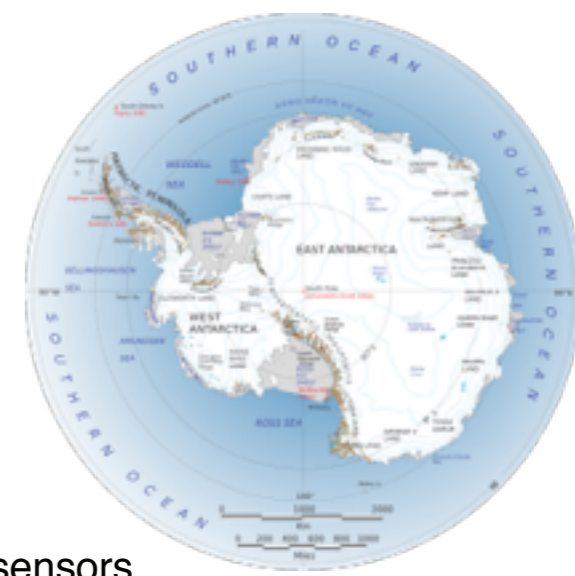
Photo-Multiplier Tubes



Photomultiplier tubes (PMTs)

*Australian Microscopy & Microanalysis Research Facility

IceCube/DeepCore



IceCube Lab

Ice Top

81 Stations
324 optical sensors

50 m

IceCube Array

86 strings including
8 DeepCore strings
5160 optical sensors

1450 m

DeepCore

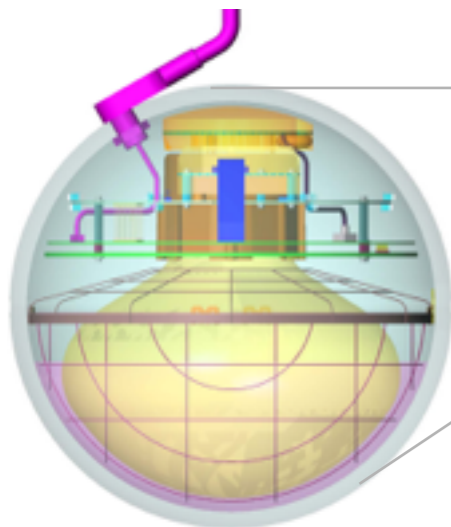
8 strings-spacing optimized
for lower energies
480 optical sensors

2450 m

Eiffel Tower
324 m

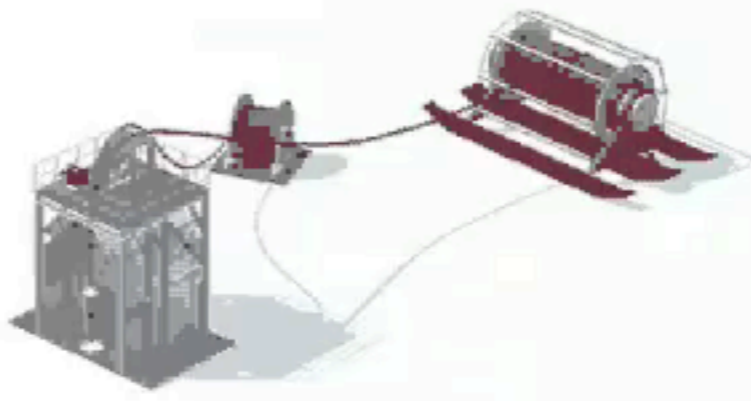
2820 m

Bedrock

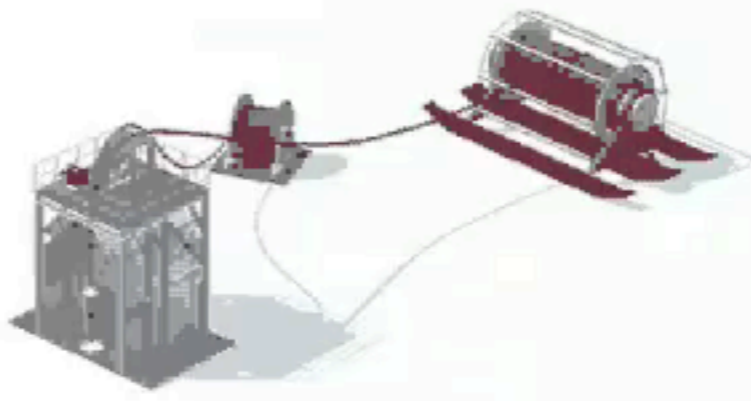


IceCube Digital Optical
Module (DOM)

IceCube Hot Water Drill Animation



IceCube Hot Water Drill Animation



NuMu

6.08e+04

44.43 deg

357.53 deg

100/446 shown, max E (GeV) == 56675.77

100/444 shown, max E (GeV) == 1.58

NuMu

6.08e+04

44.43 deg

357.53 deg

100/446 shown, max E (GeV) == 56675.77

100/444 shown, max E (GeV) == 1.58

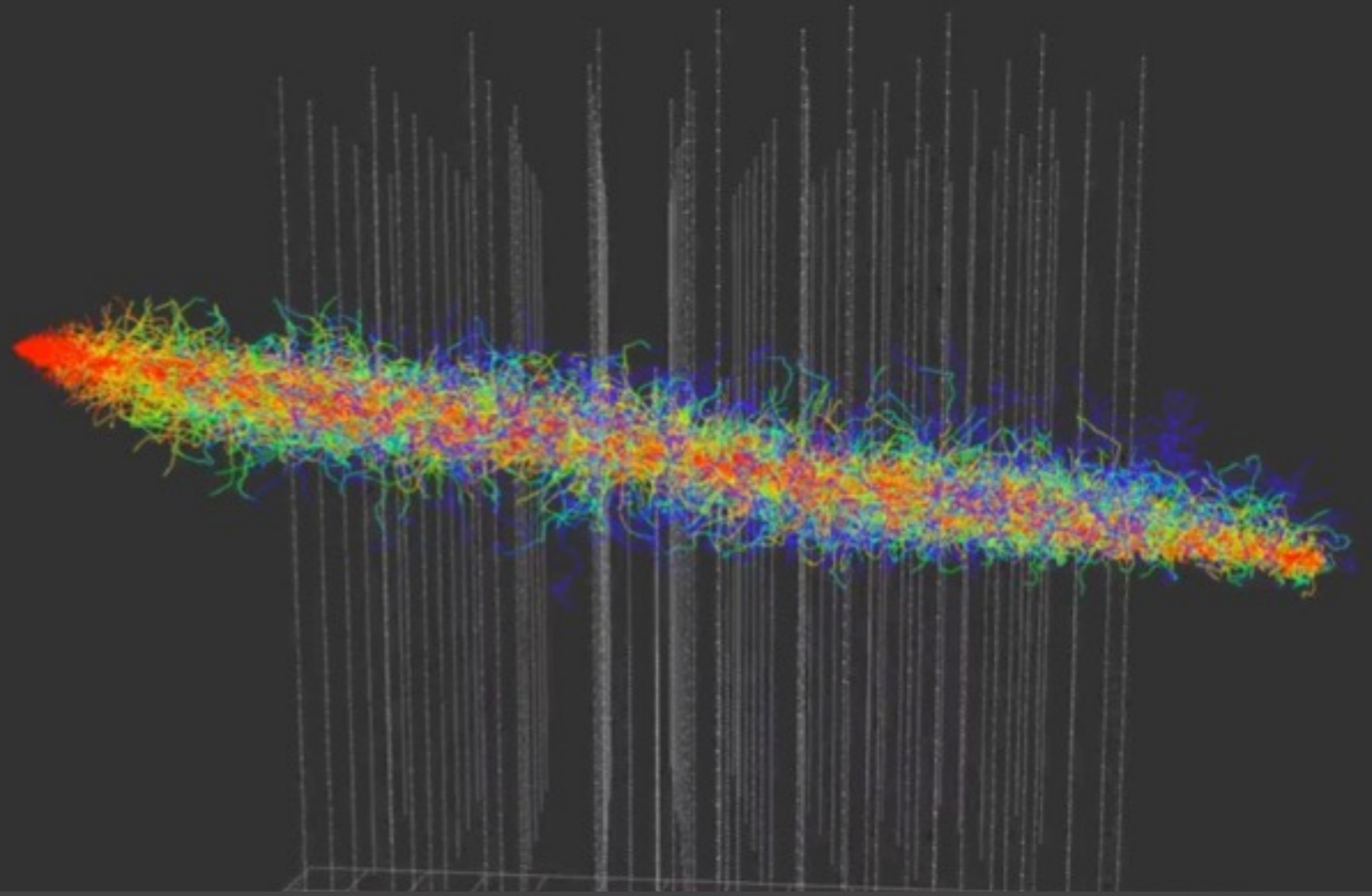
Track topology

(e.g. induced by
muon neutrino)

Good pointing

IceCube: lower bound on
energy for through-going
events

DeepCore: well
contained and provide
good energy via muon
track length



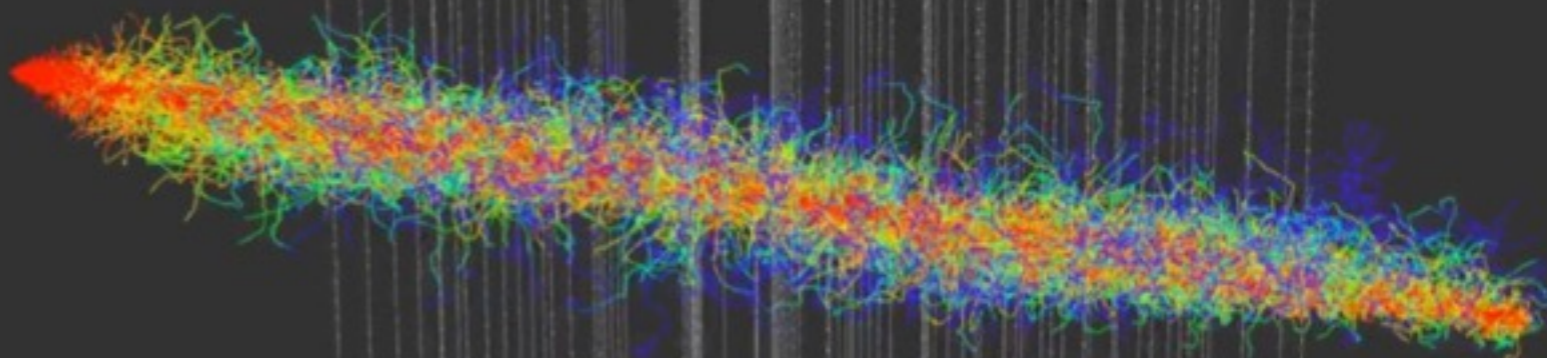
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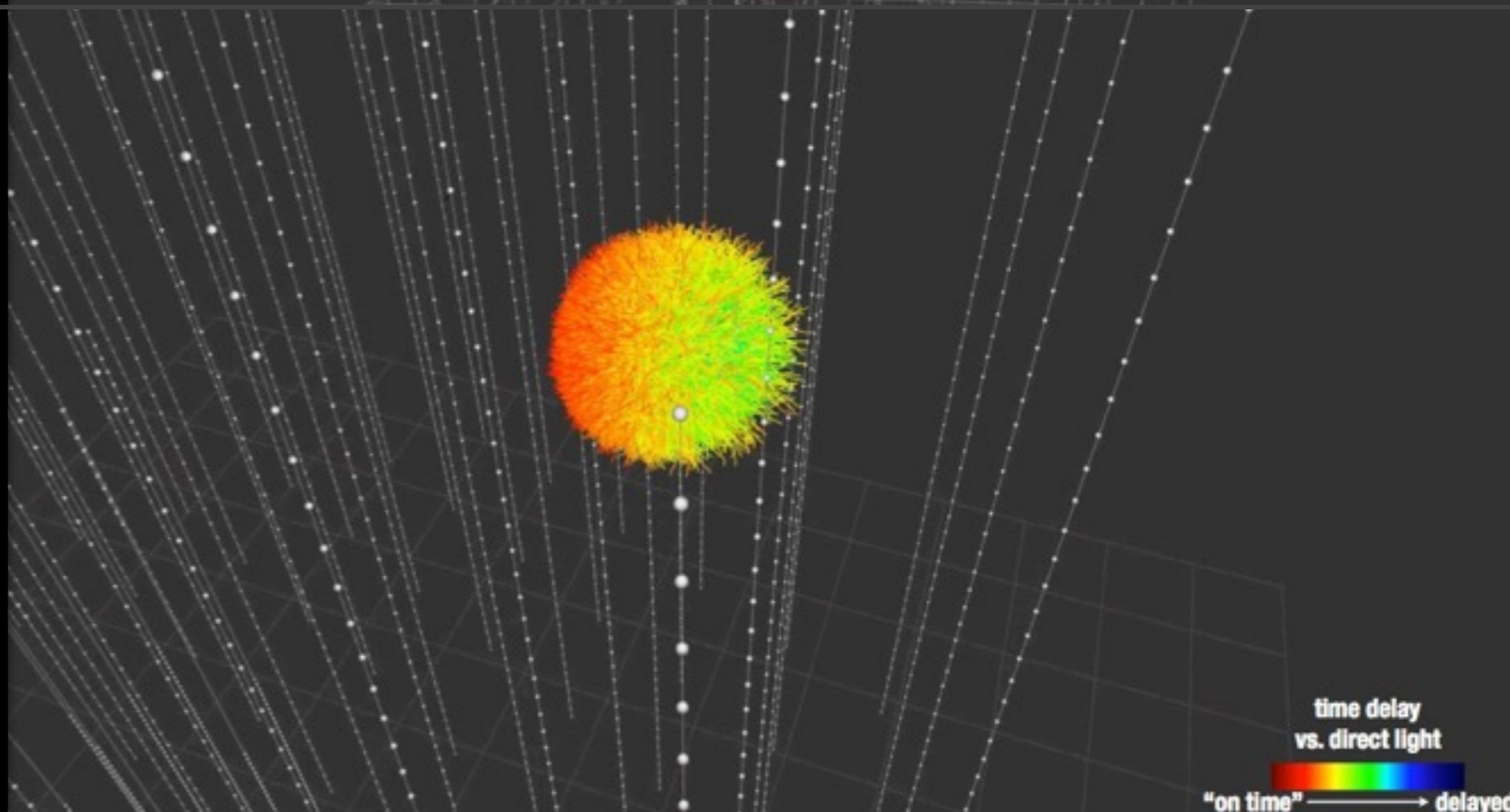
Cascade topology

(e.g. induced by electron neutrino)

Good energy resolution

IceCube: some pointing

DeepCore: poor pointing, more difficult to ID and reconstruct



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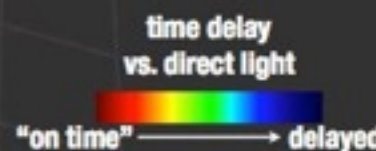
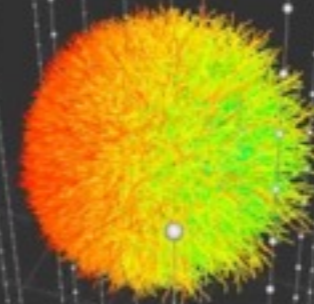
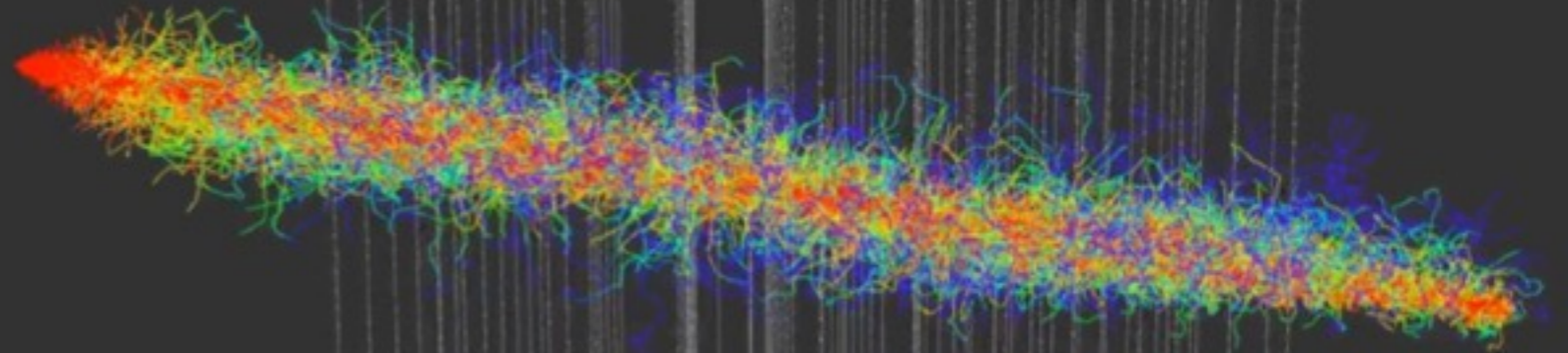
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Good energy resolution

IceCube: some pointing

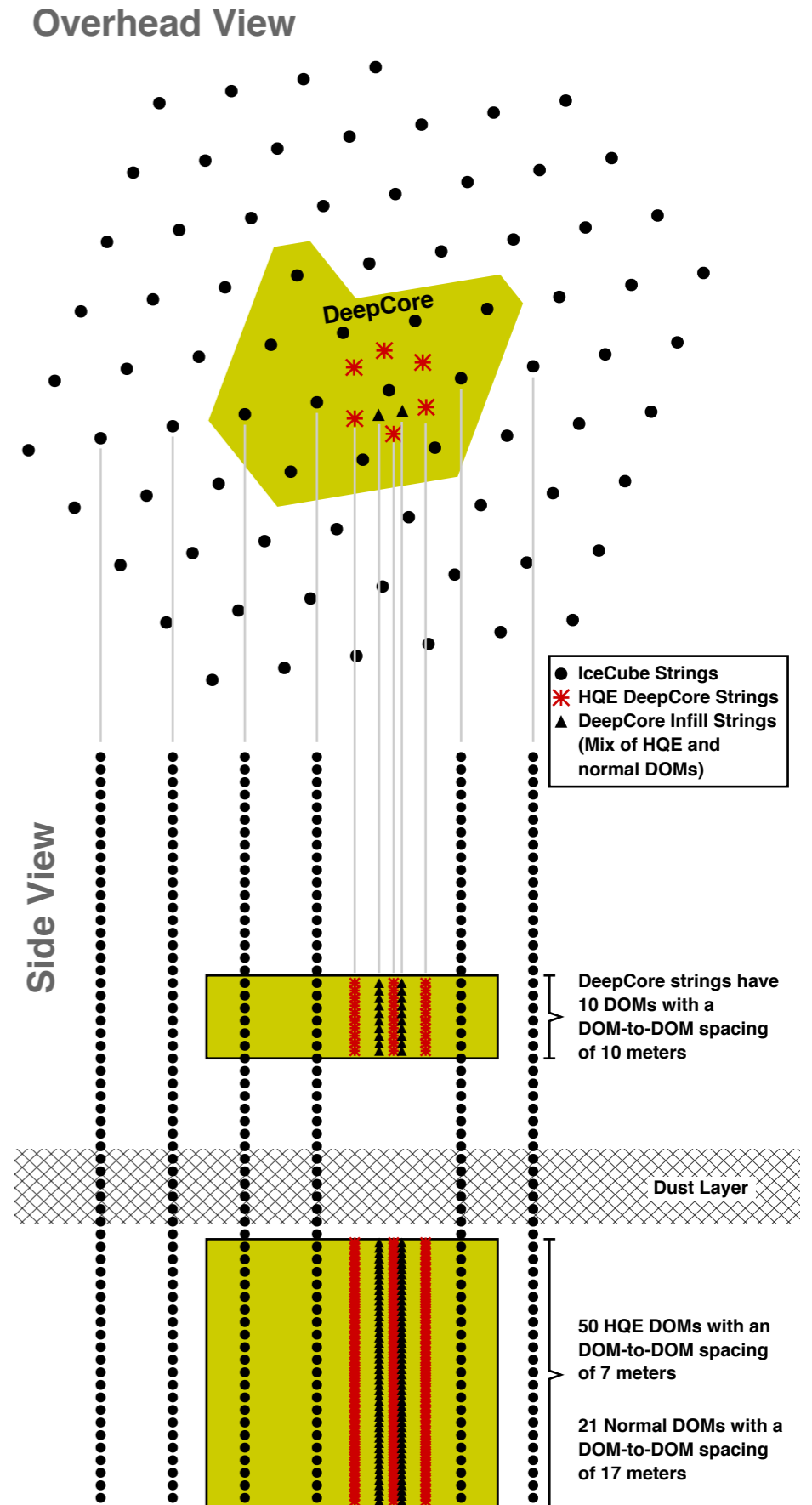
DeepCore: poor pointing, more difficult to ID and reconstruct

ν and $\bar{\nu}$ are, essentially, indistinguishable



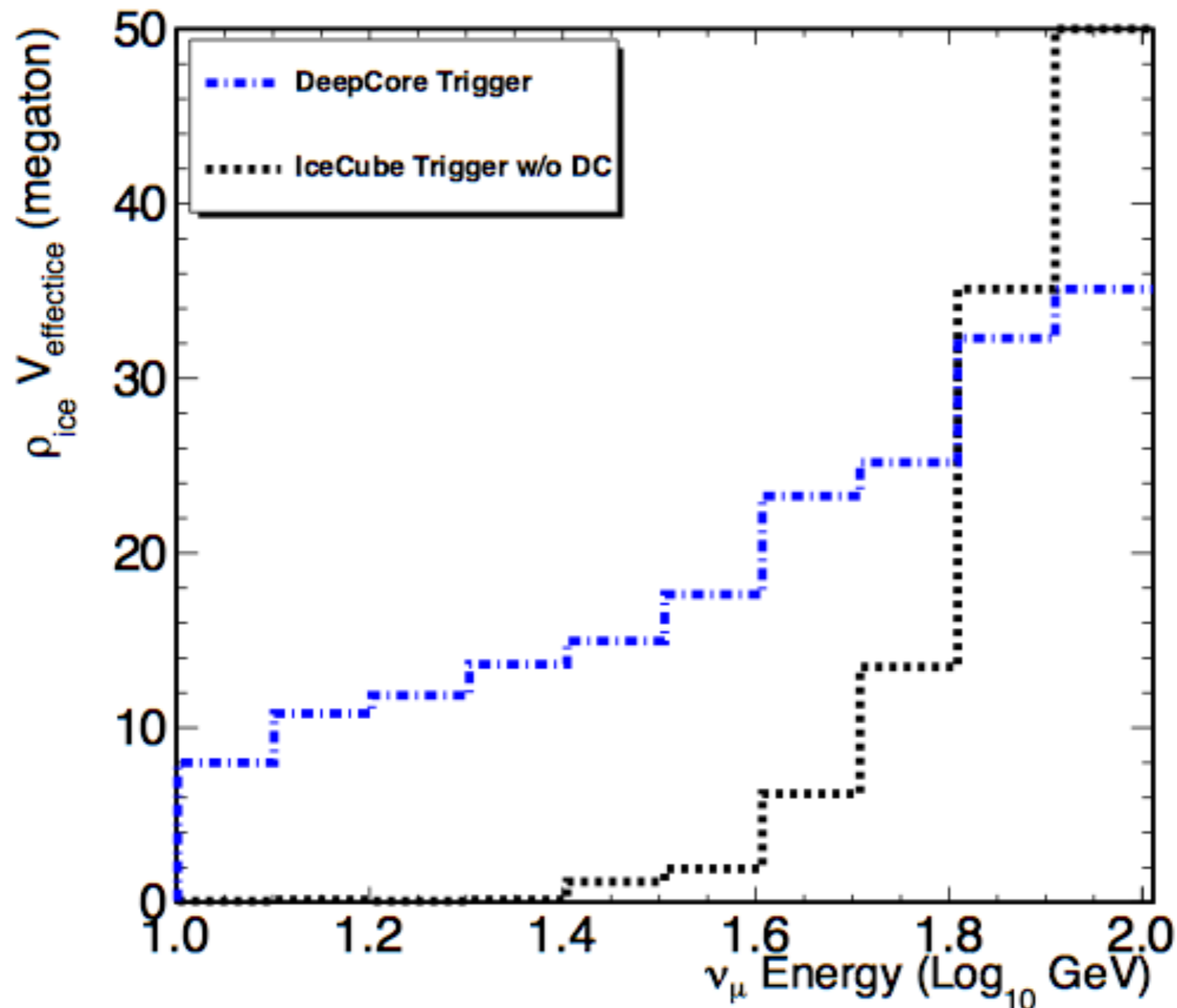
IceCube/DeepCore

- DeepCore
 - Increased sensitivity at energies less than 100-200 GeV
 - 8 special strings plus 12 closest IceCube-standard strings
 - Denser DOM and string spacing
 - Deepest and clearest Ice
 - Higher efficiency photon sensors
 - Lower trigger threshold
- IceCube is not only a high energy neutrino detector, but also a cosmic ray muon veto for any inner detectors



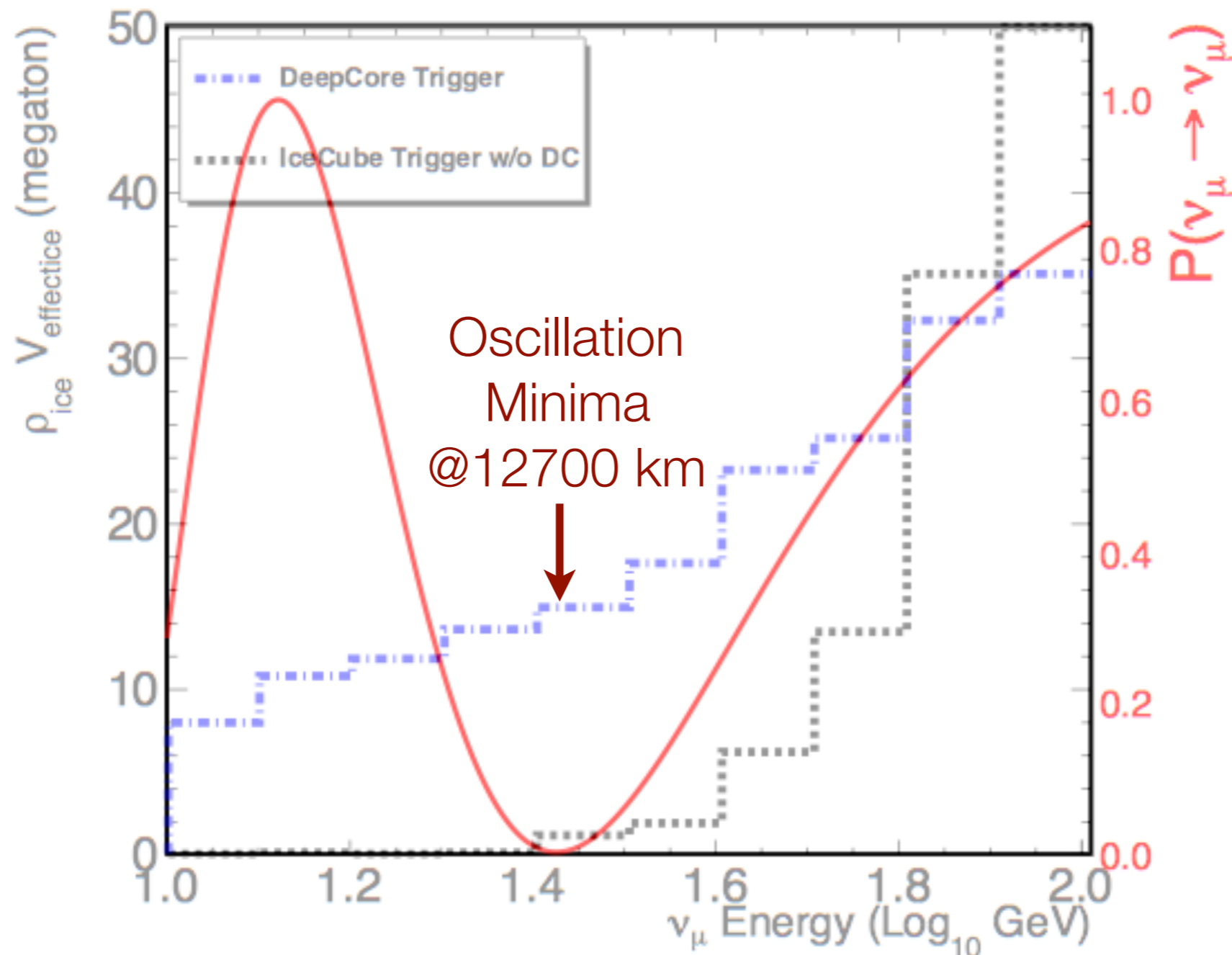
Oscillation w/ DeepCore

- IceCube + DeepCore collects $> 100k$ isotropic neutrinos *at trigger level*, tens of thousands have undergone oscillation. Even single percent final analysis efficiency contains 1,000s of atm. ν events/year

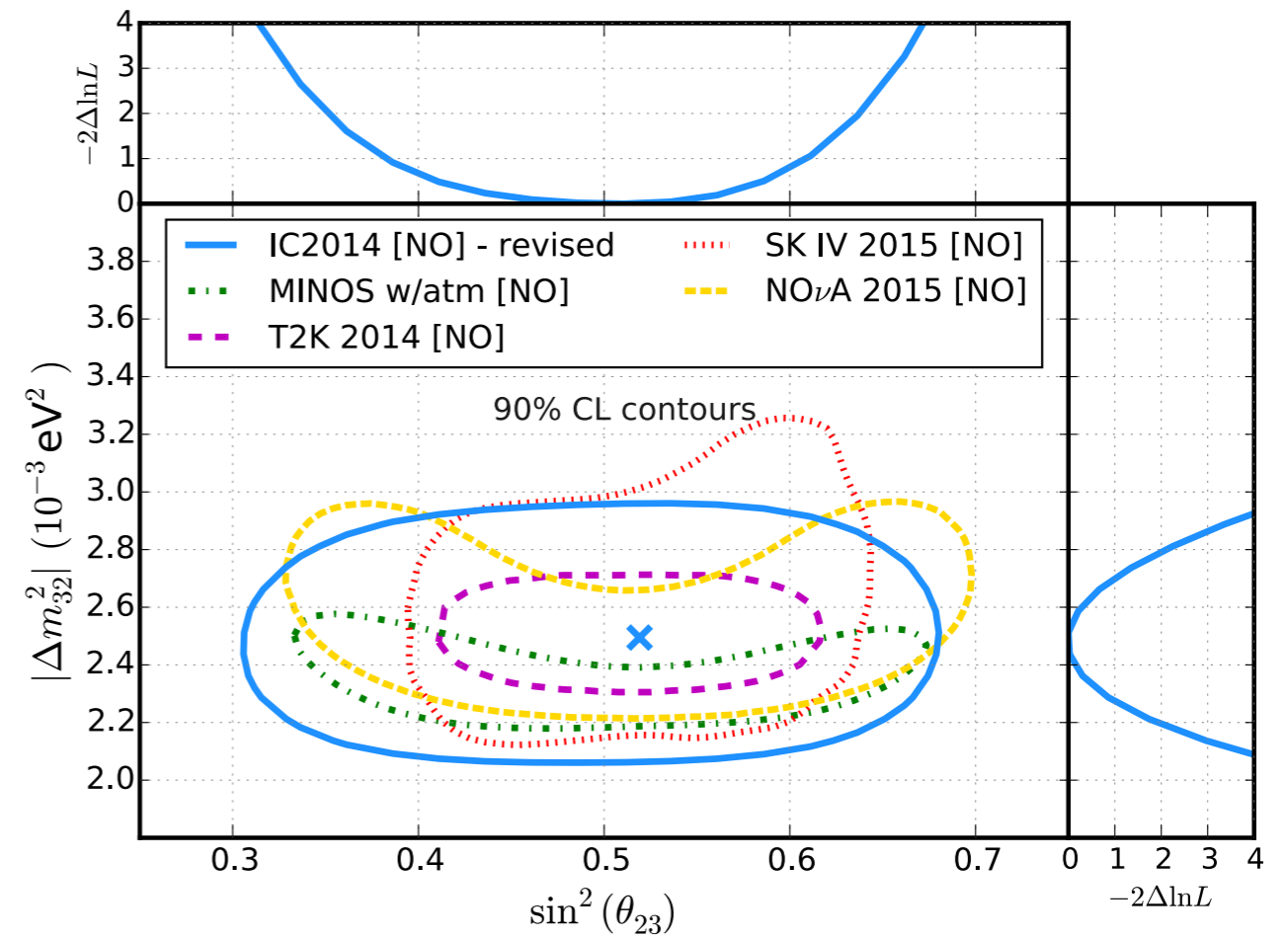
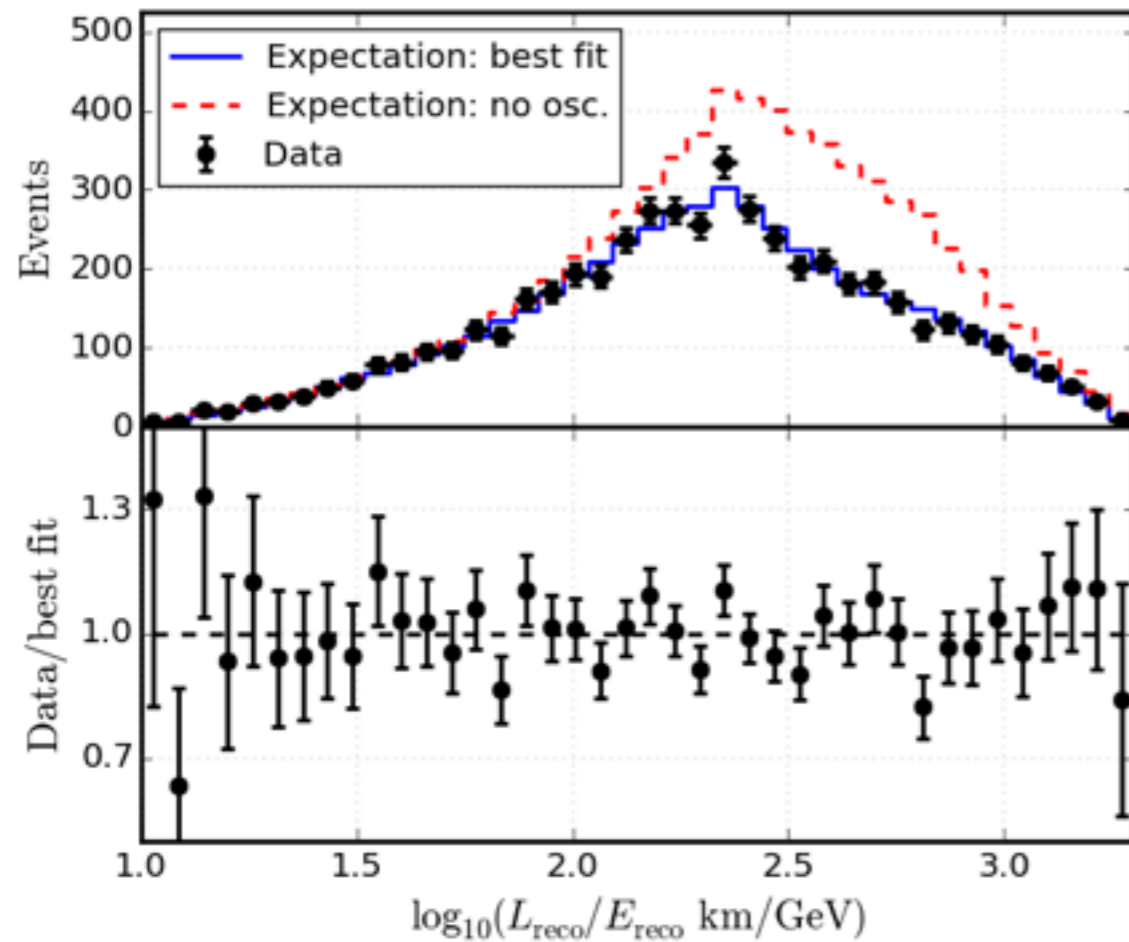


Oscillation w/ DeepCore

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Current Results



*DJK, Neutrino 2016

Measuring Parameters

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix}$$

underlying nature of
weak mixing

$$C_{12} = \cos\theta_{12} \quad S_{12} = \sin\theta_{12}$$

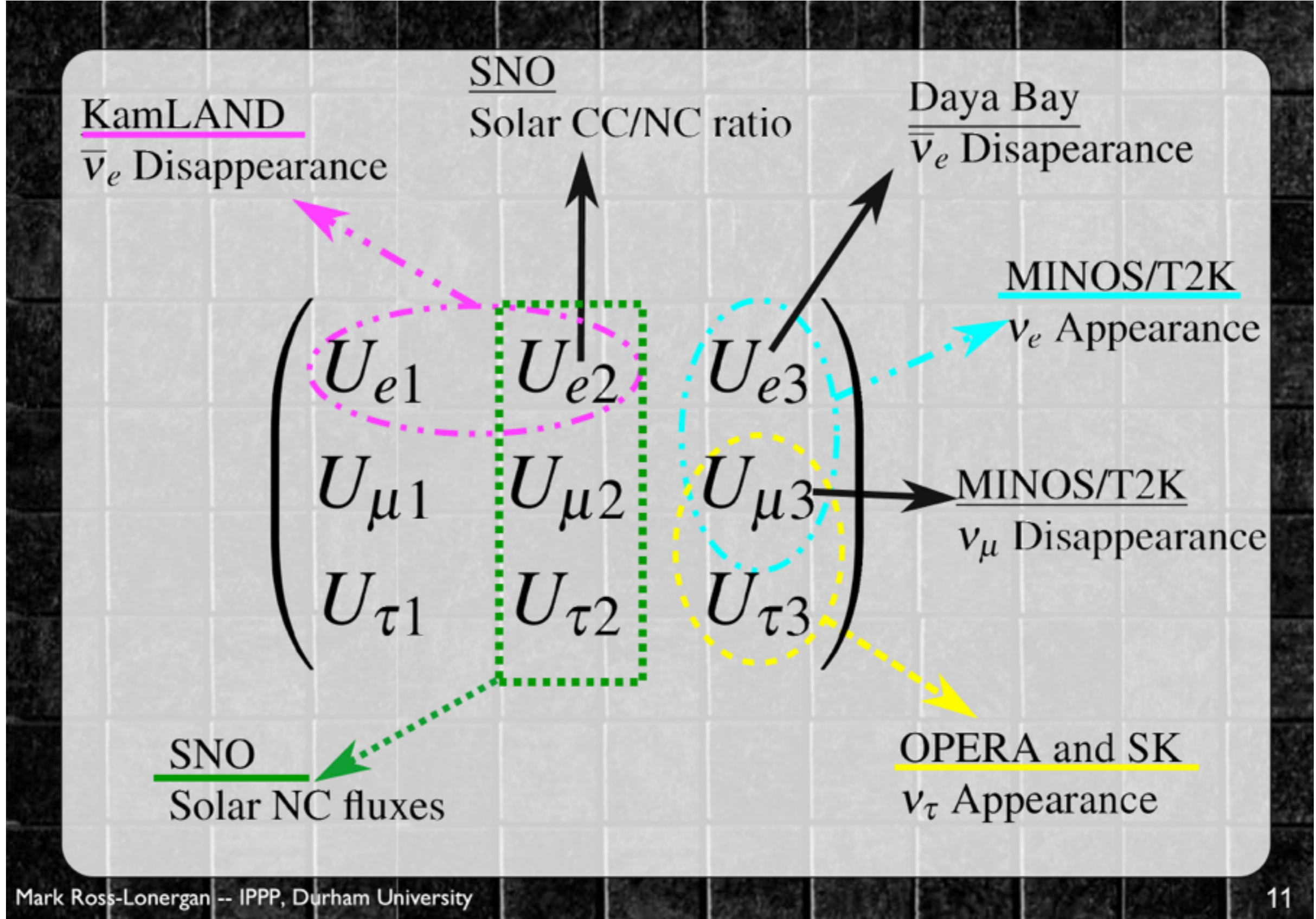
$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix}$$

Experimentally
measured
values

Three angles and one Charge-Parity phase

What Is Being Measured?

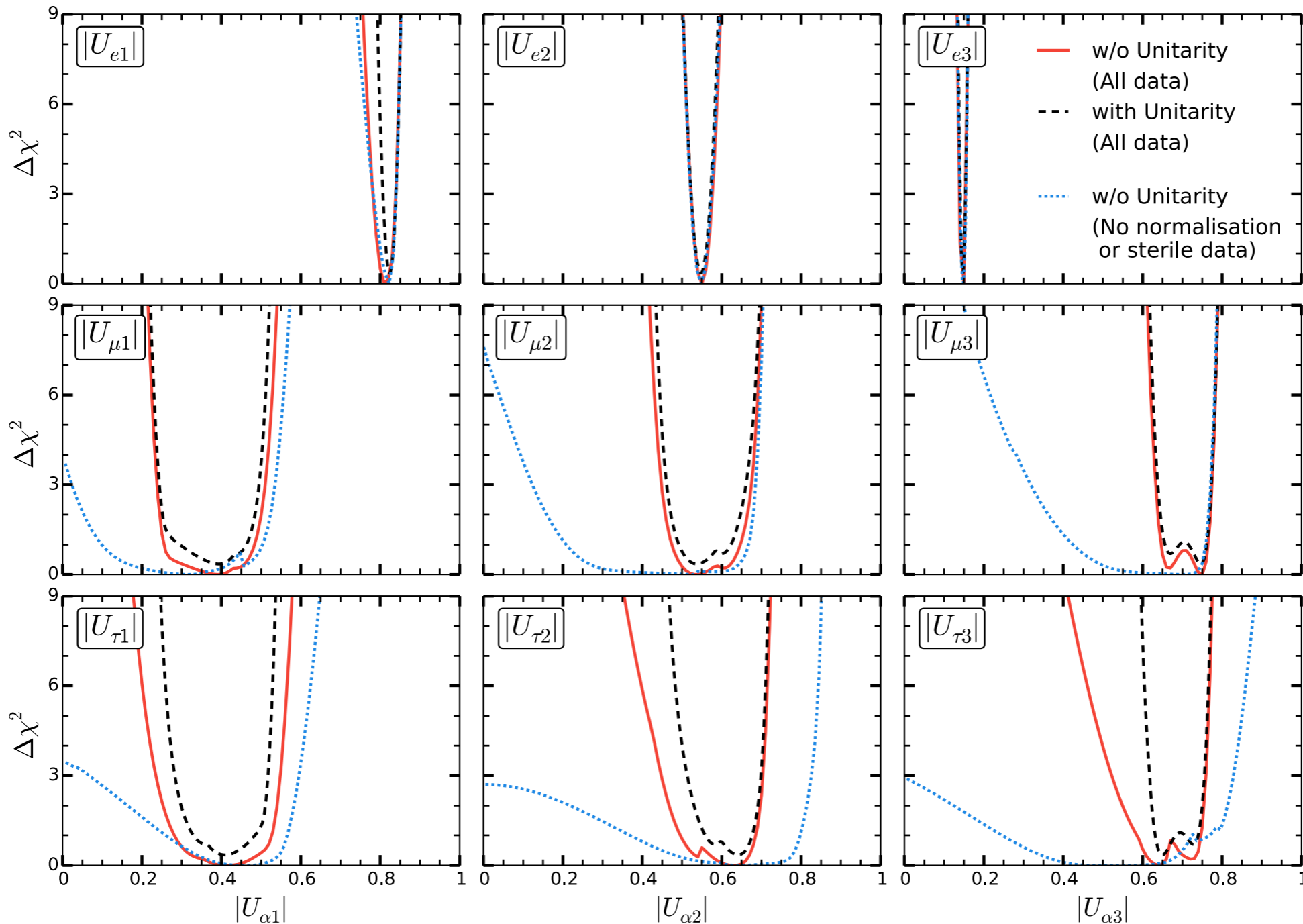
*NOW2014



Unitarity

- Minimal assumption direct experimental constraints for PMNS unitarity can be improved upon

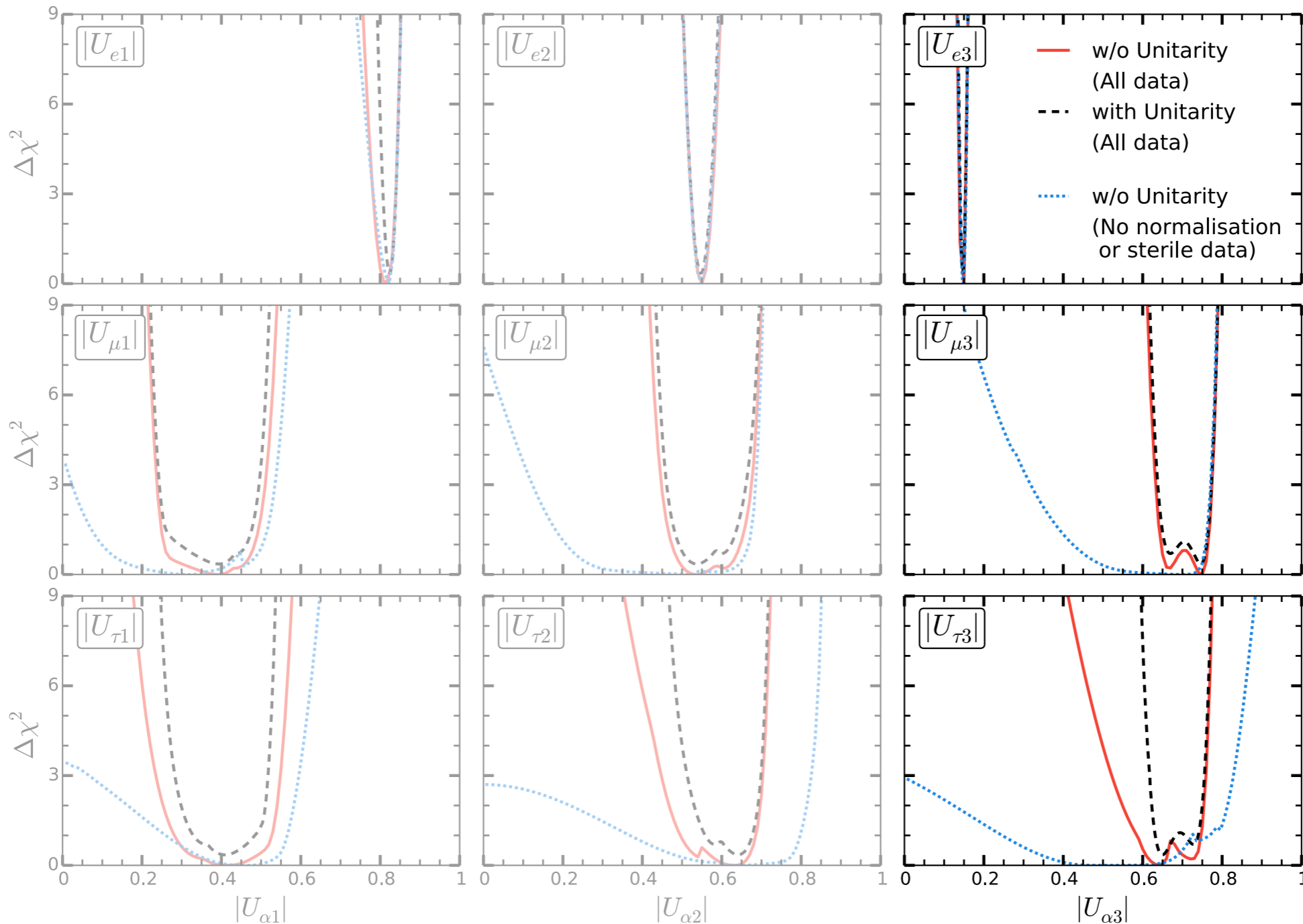
*arXiv:1508.05095



Unitarity

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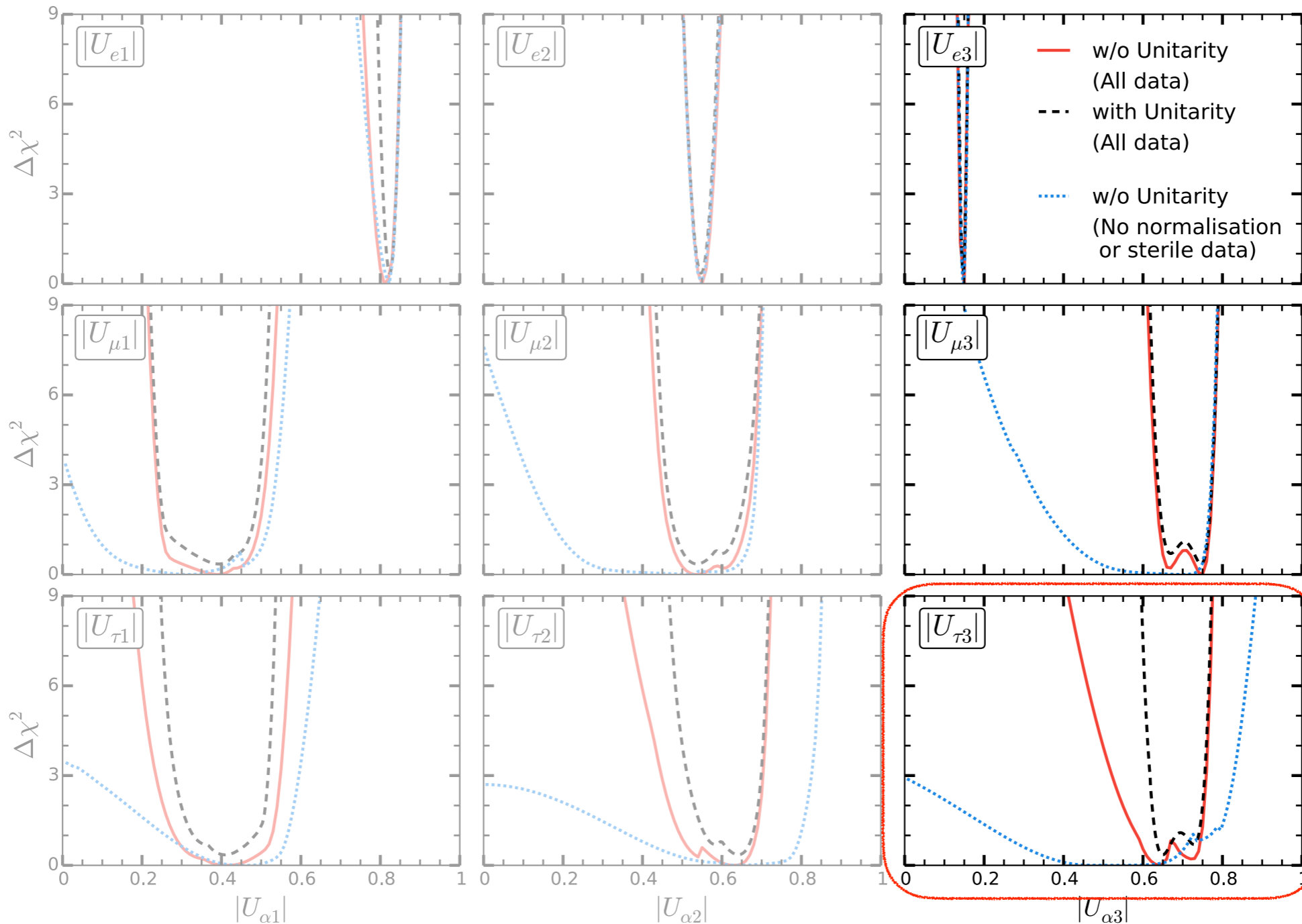
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Unitarity

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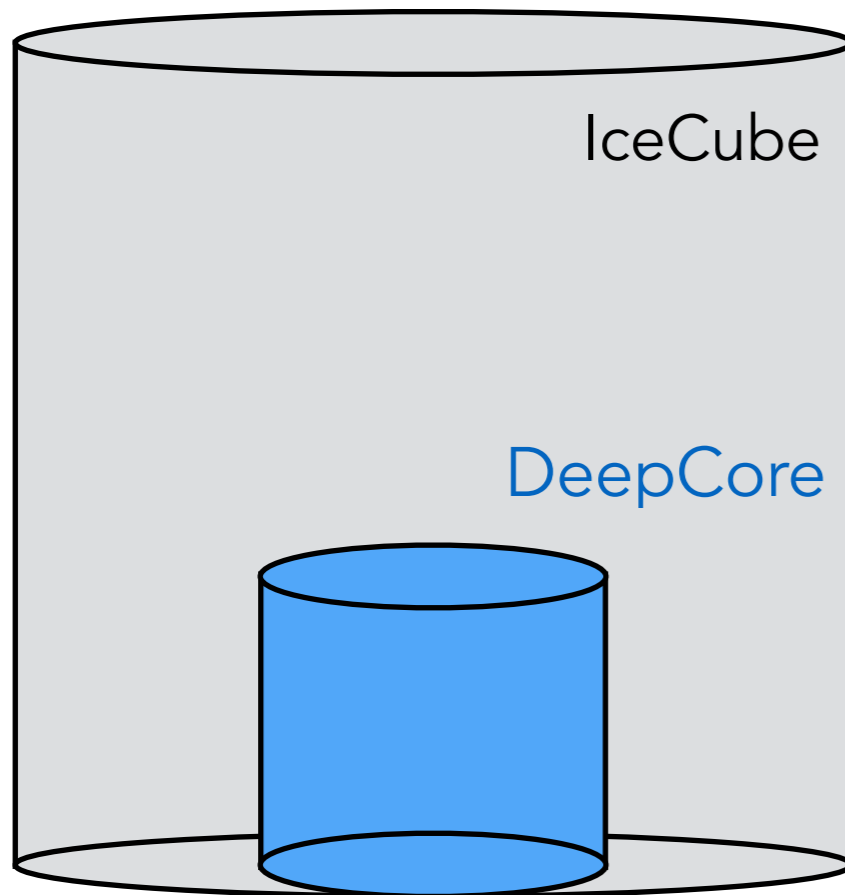


From OPERA
 $|U_{\tau3}|^2 > 0$
 *arXiv:1507:01417

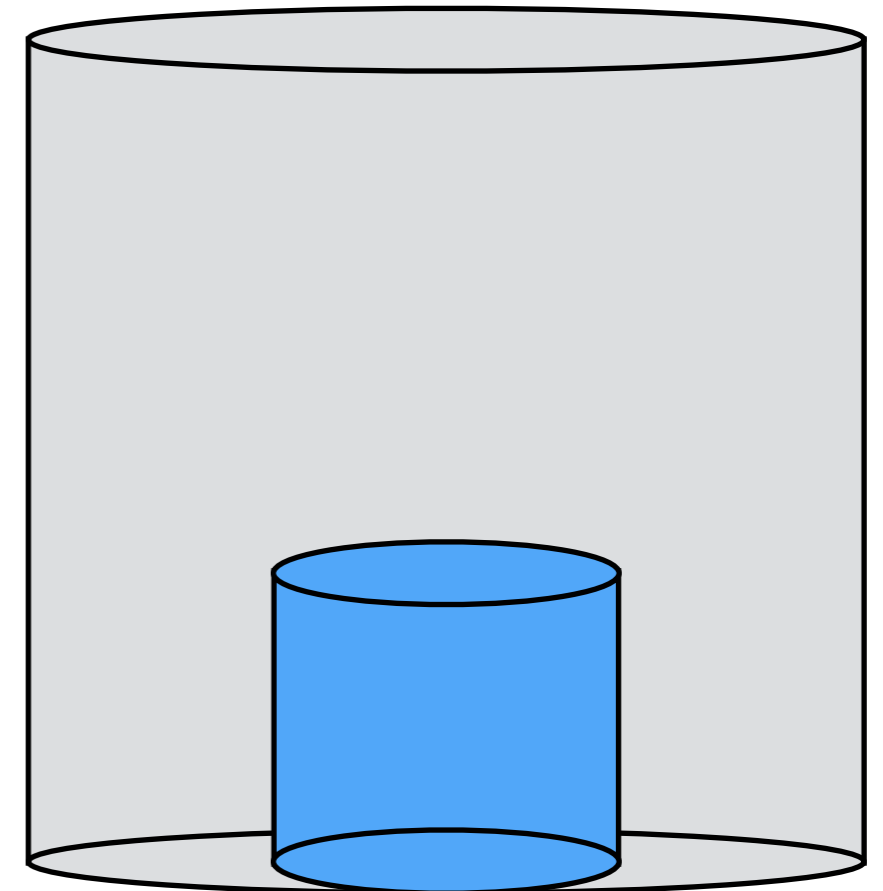
Experimental Approach

- IceCube-DeepCore will collect the largest sample of oscillated ν_τ **ever**

Signal



Background

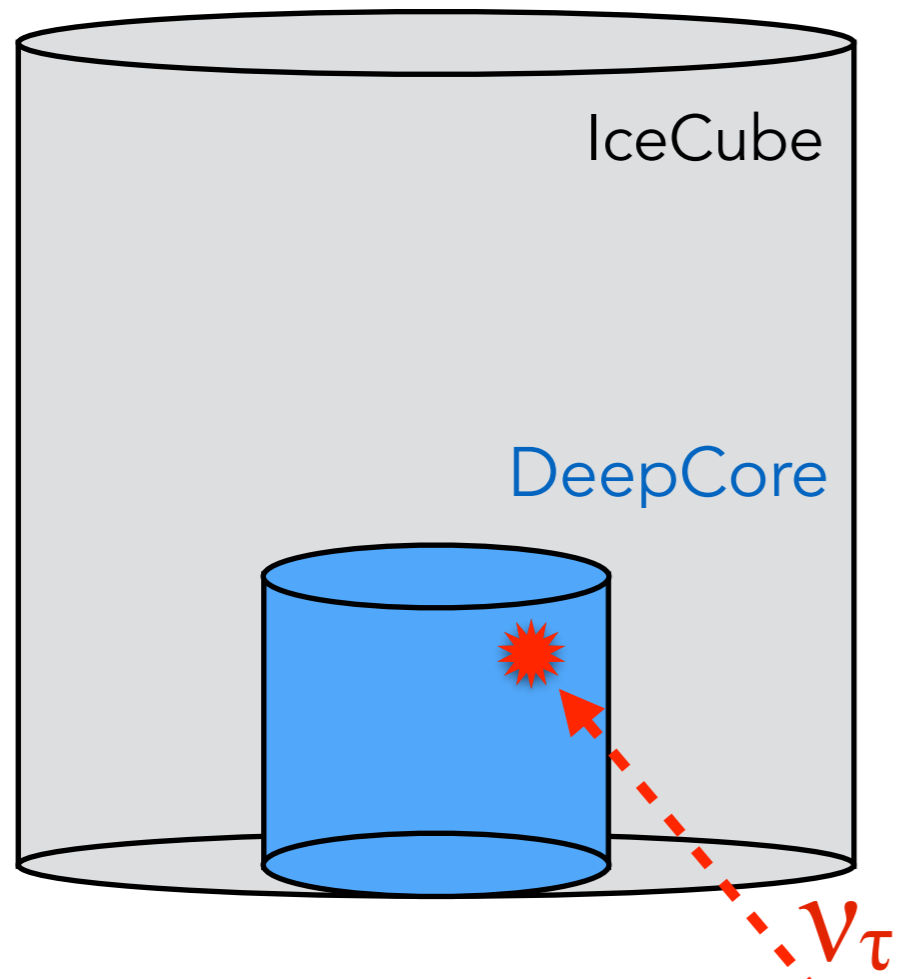


Experimental Approach

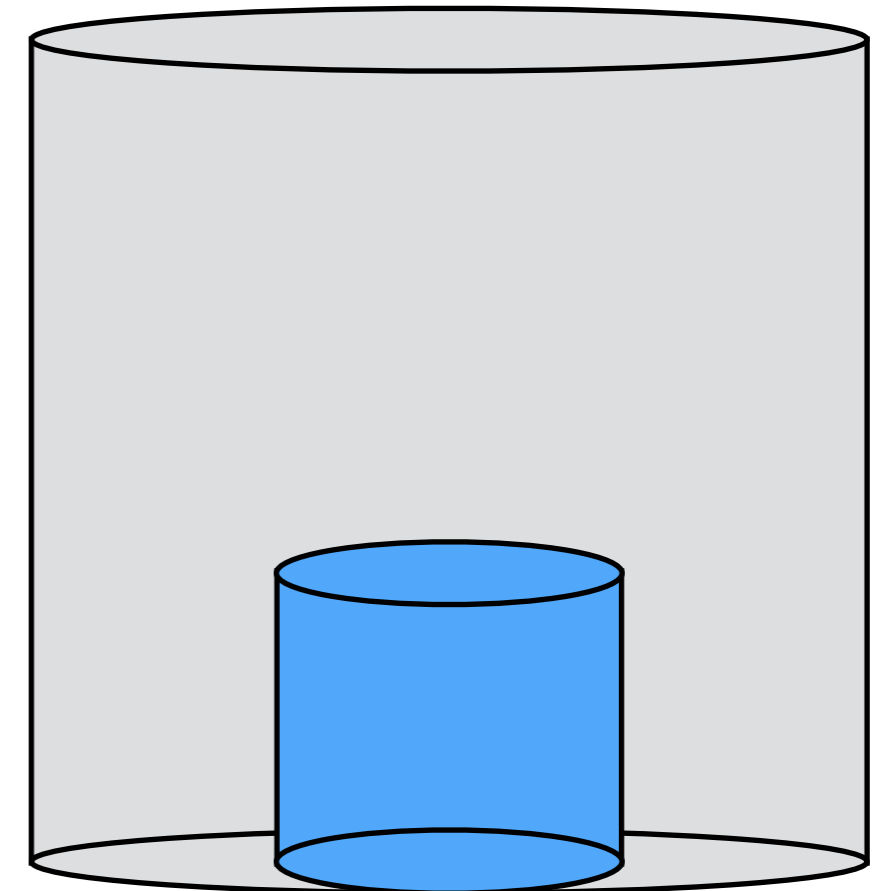
- IceCube-DeepCore will collect the largest sample of oscillated ν_τ **ever**

Signal

- ν -cascade at $\mathcal{O}(25)$ GeV



Background

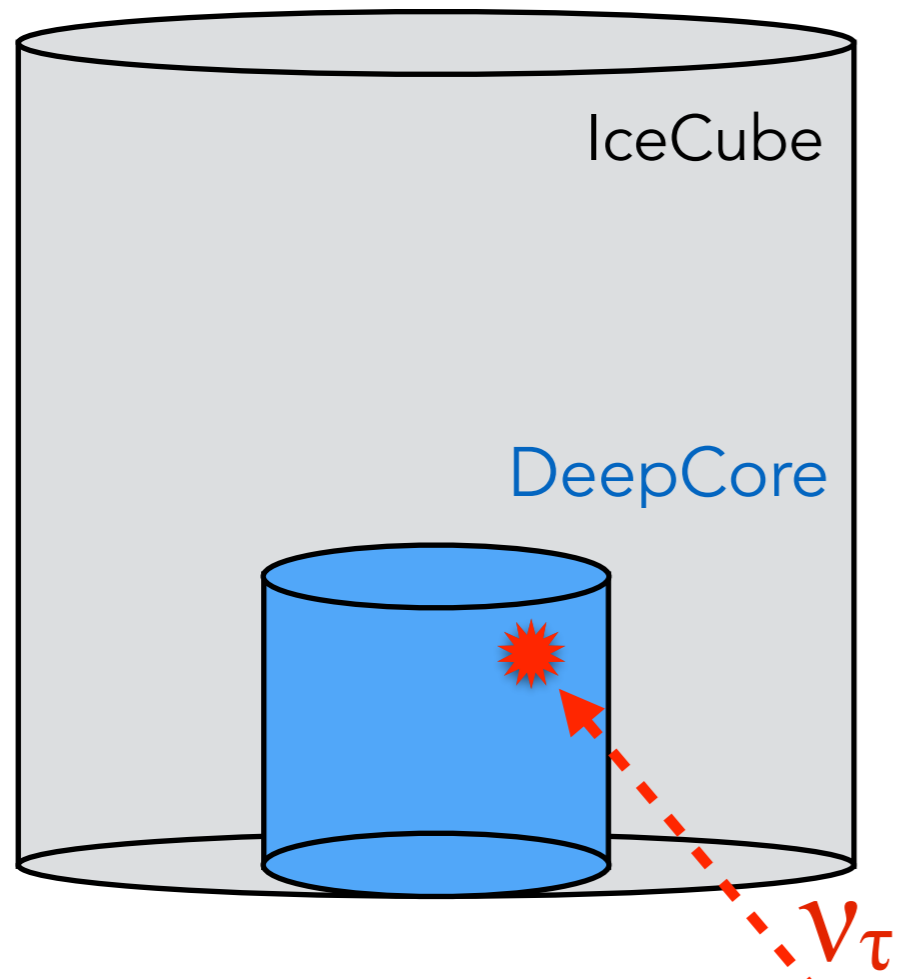


Experimental Approach

- IceCube-DeepCore will collect the largest sample of oscillated ν_τ **ever**

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- ν -cascade at $\mathcal{O}(25)$ GeV

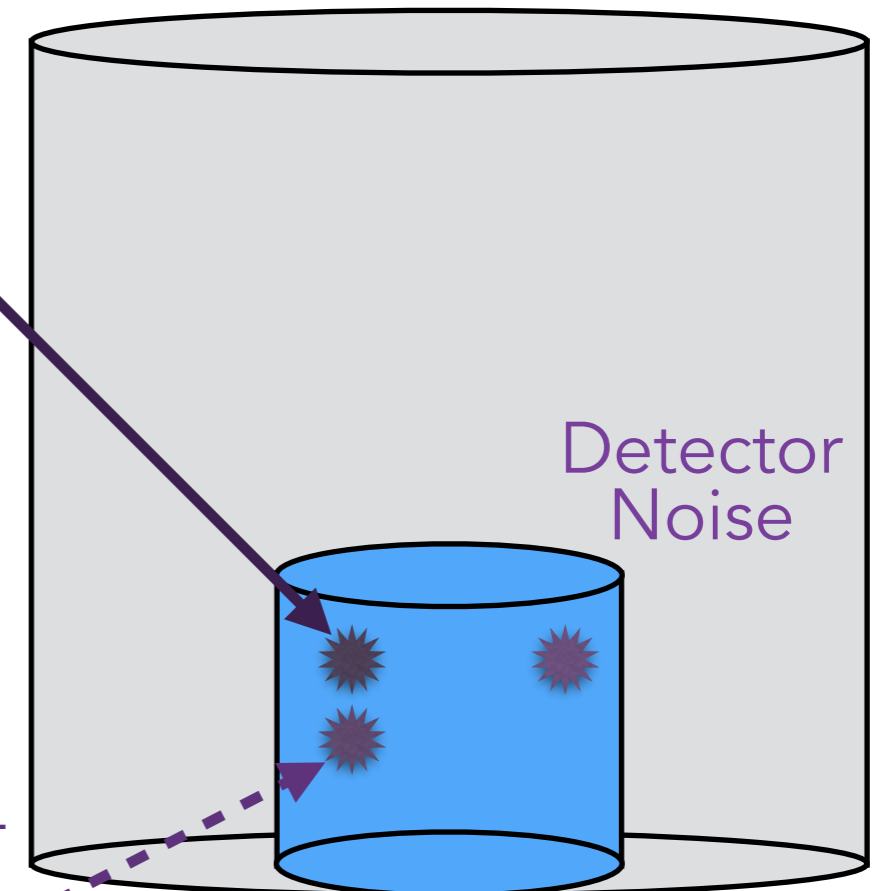


Background

- Pure-noise events
- 'Ghost' cosmic ray muons
- Irreducible background from cascade-type neutrino interactions

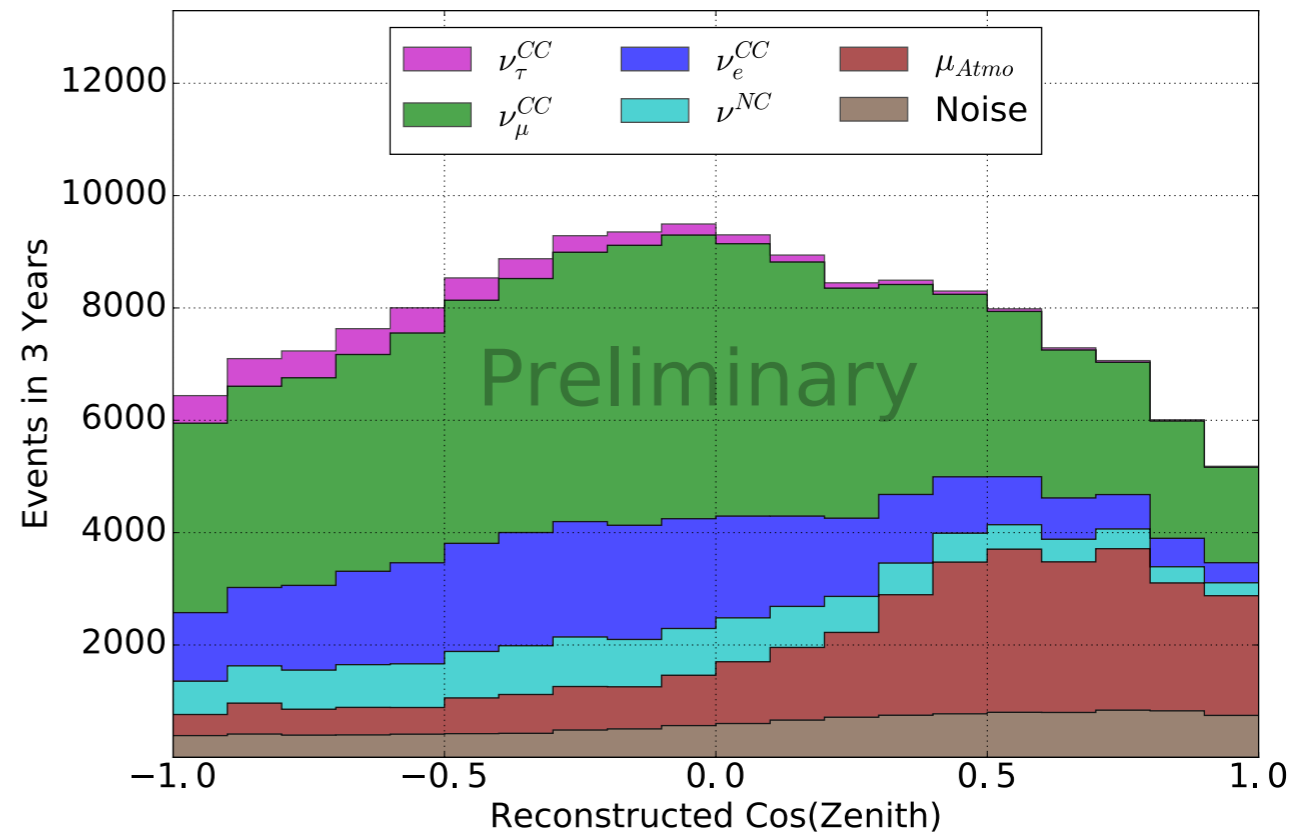
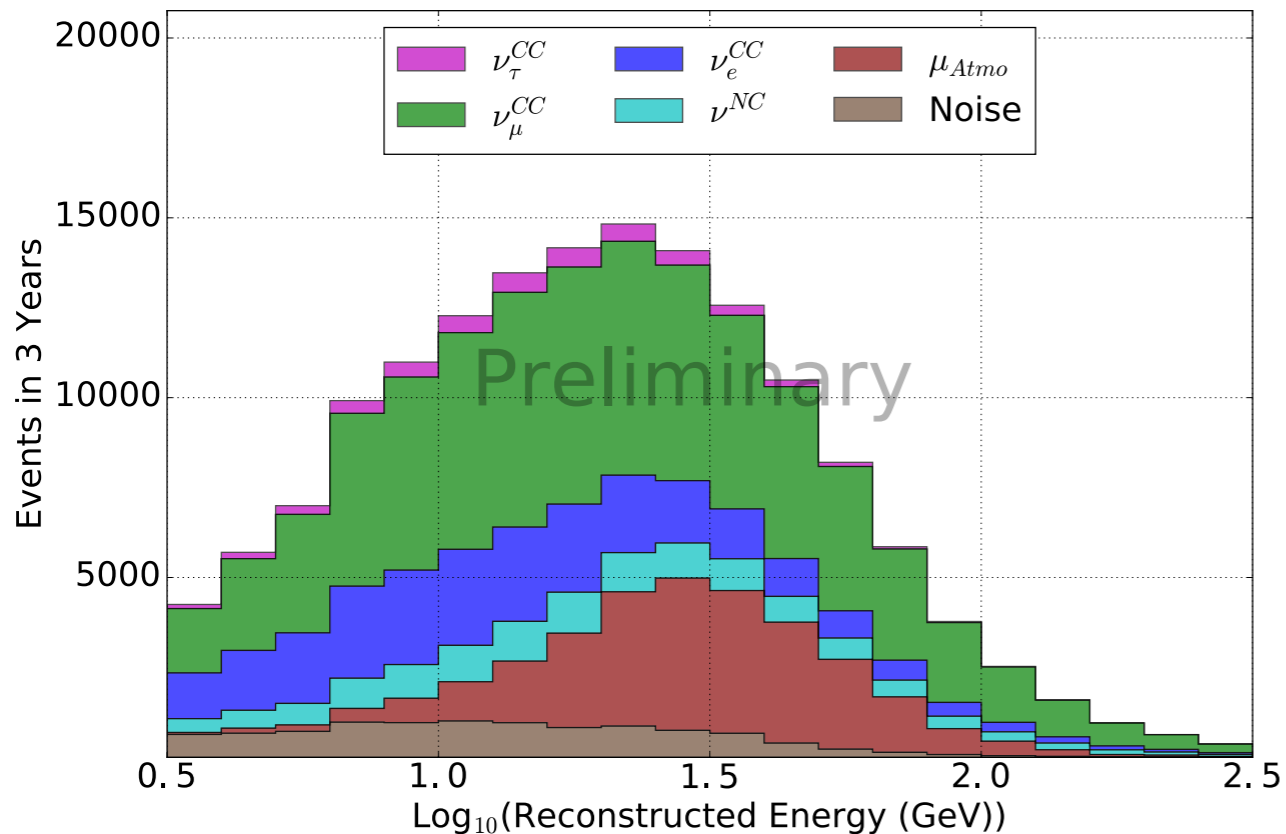
Cosmic Ray Muon

$\nu_{e,\mu,\tau}$ -NC +
 ν_e -CC



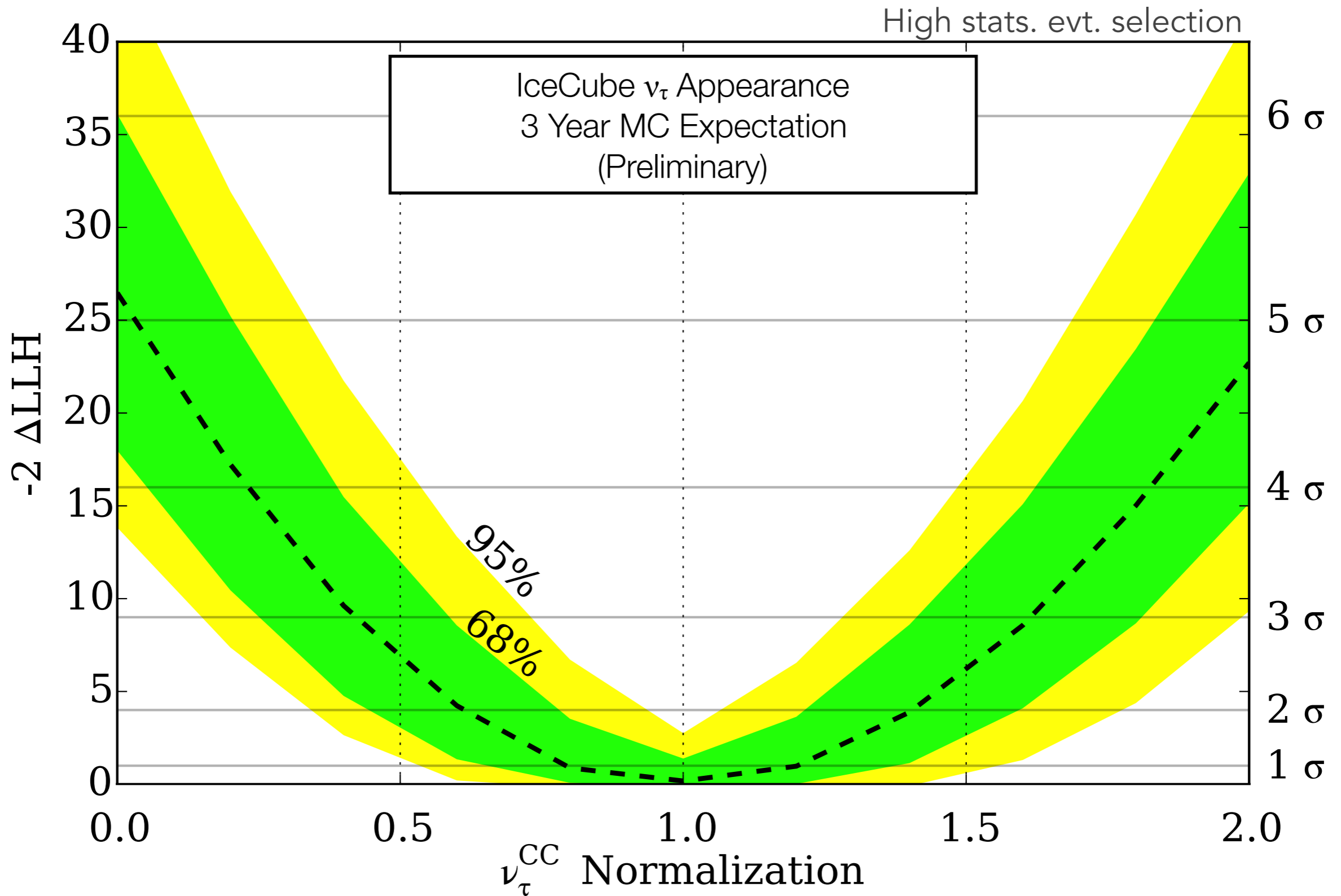
Event Rate for ν_τ Appearance

- Expected rate is low compared to background
 - Kinematic suppression to the ν_τ cross-section versus $\nu_{e,\mu}$
 - τ -lepton decays quickly w/ final state neutrino resulting in missing energy



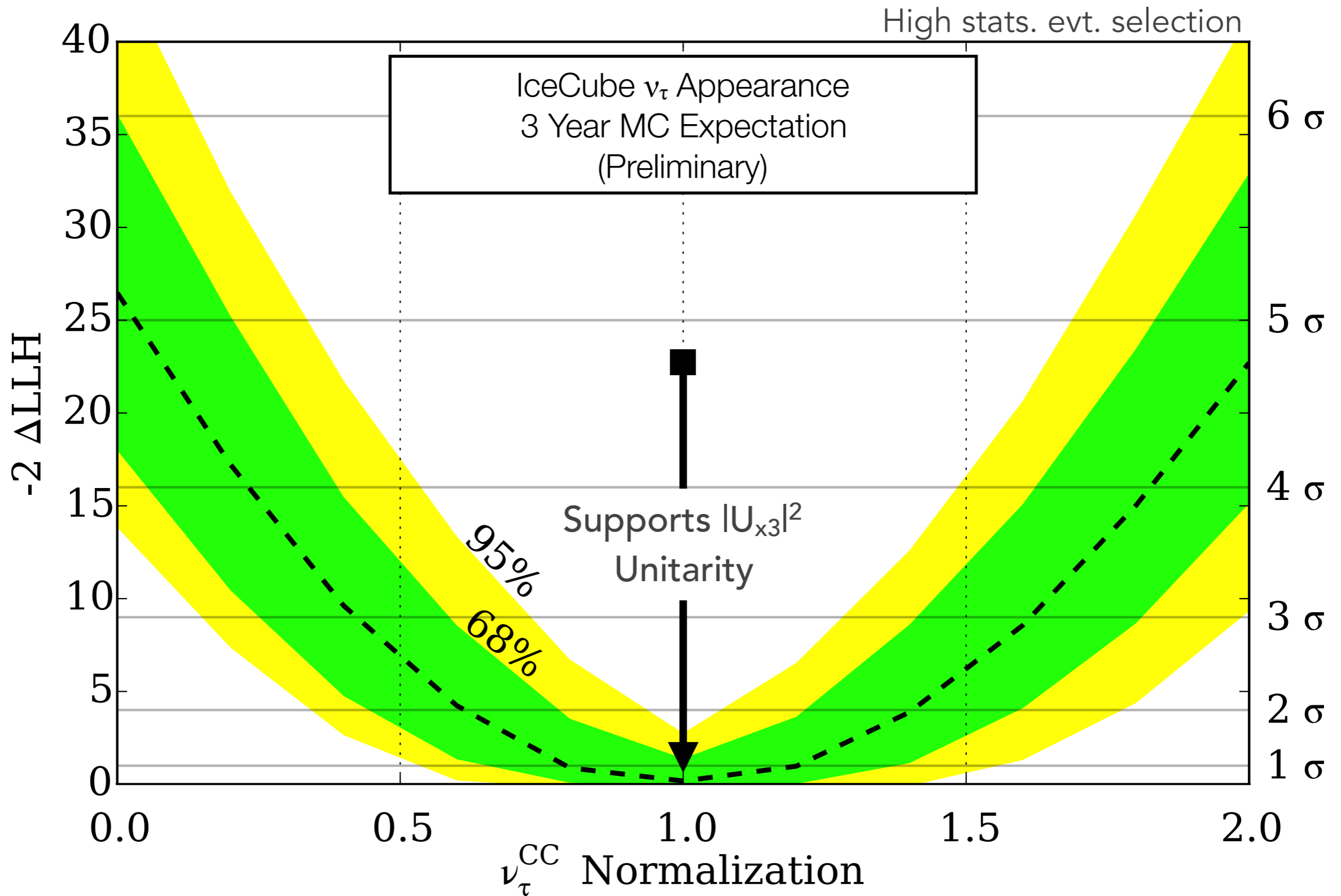
*M. Larson, Neutrino 2016 (poster)

IceCube/DeepCore ν_τ Appearance



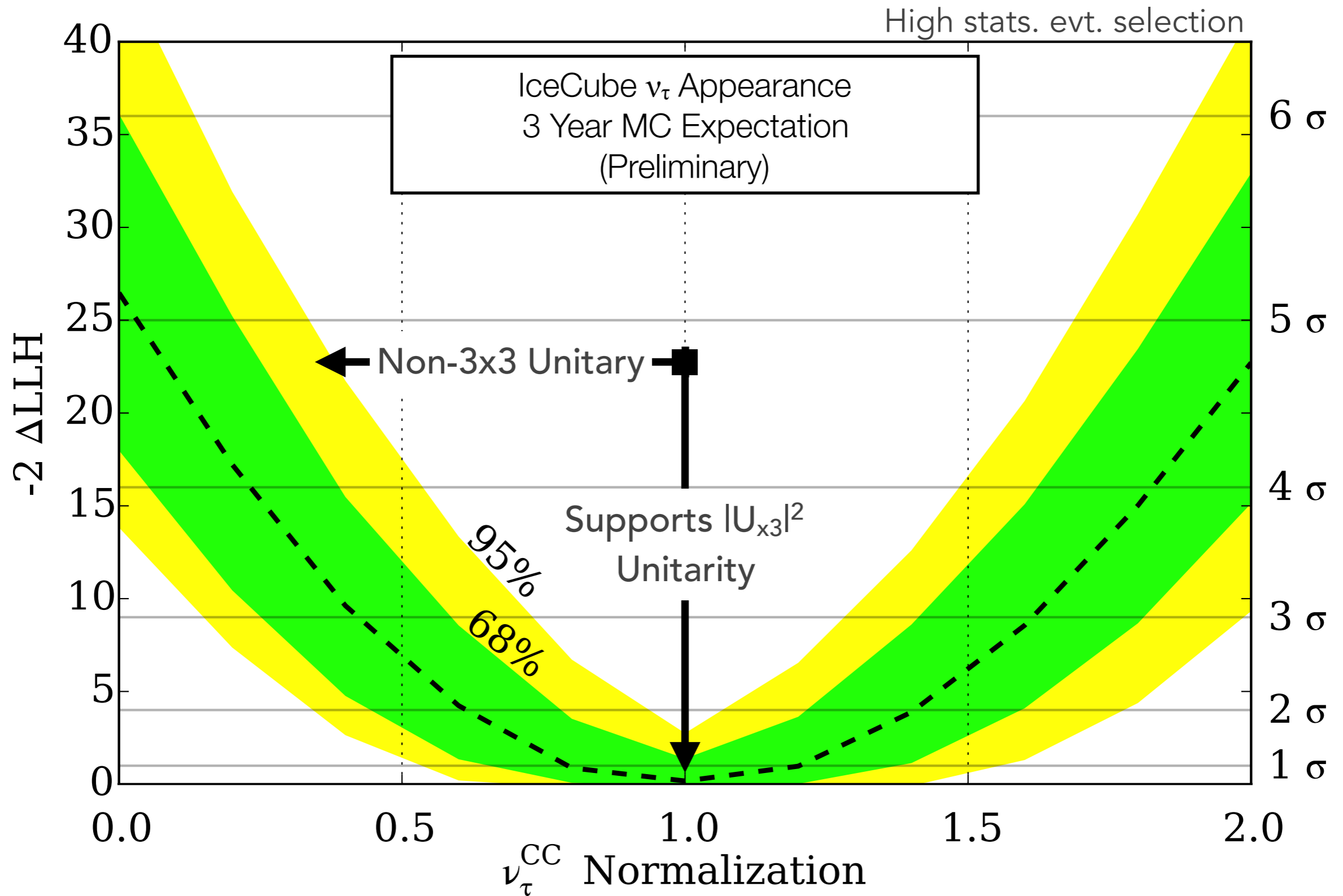
*DJK, Neutrino 2016

IceCube/DeepCore ν_τ Appearance



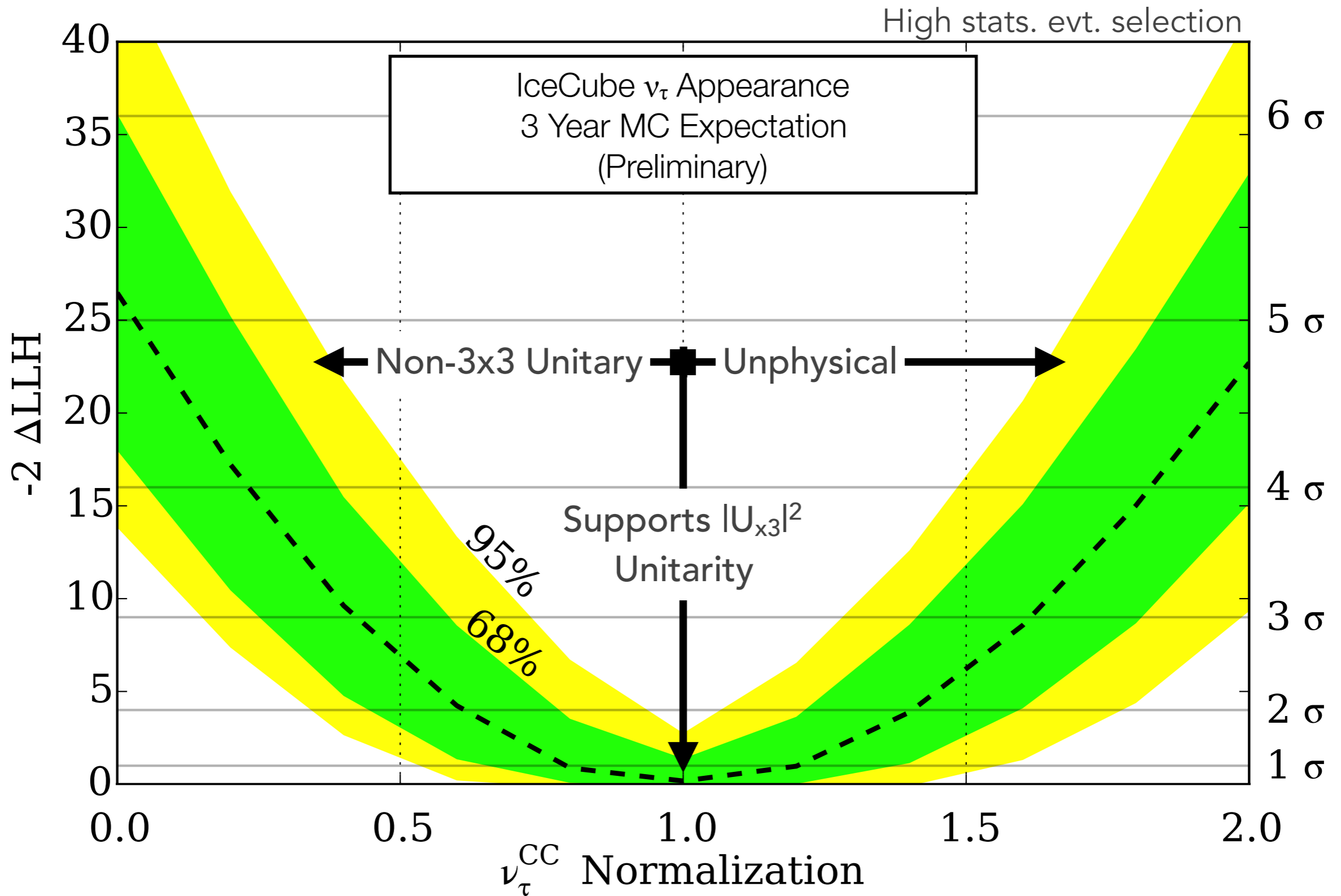
*DJK, Neutrino 2016

IceCube/DeepCore ν_τ Appearance



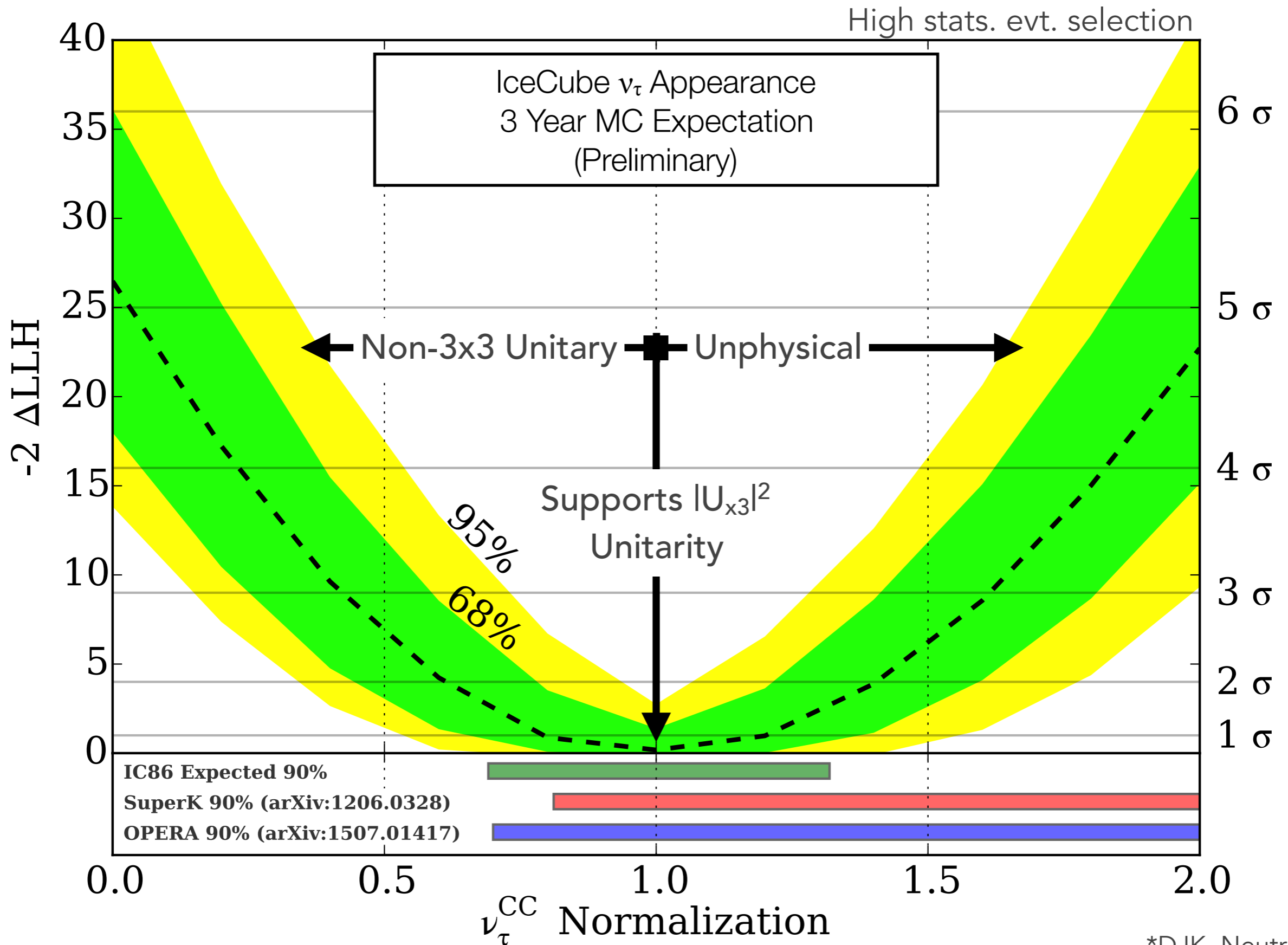
*DJK, Neutrino 2016

IceCube/DeepCore ν_τ Appearance



*DJK, Neutrino 2016

IceCube/DeepCore ν_τ Appearance



*DJK, Neutrino 2016

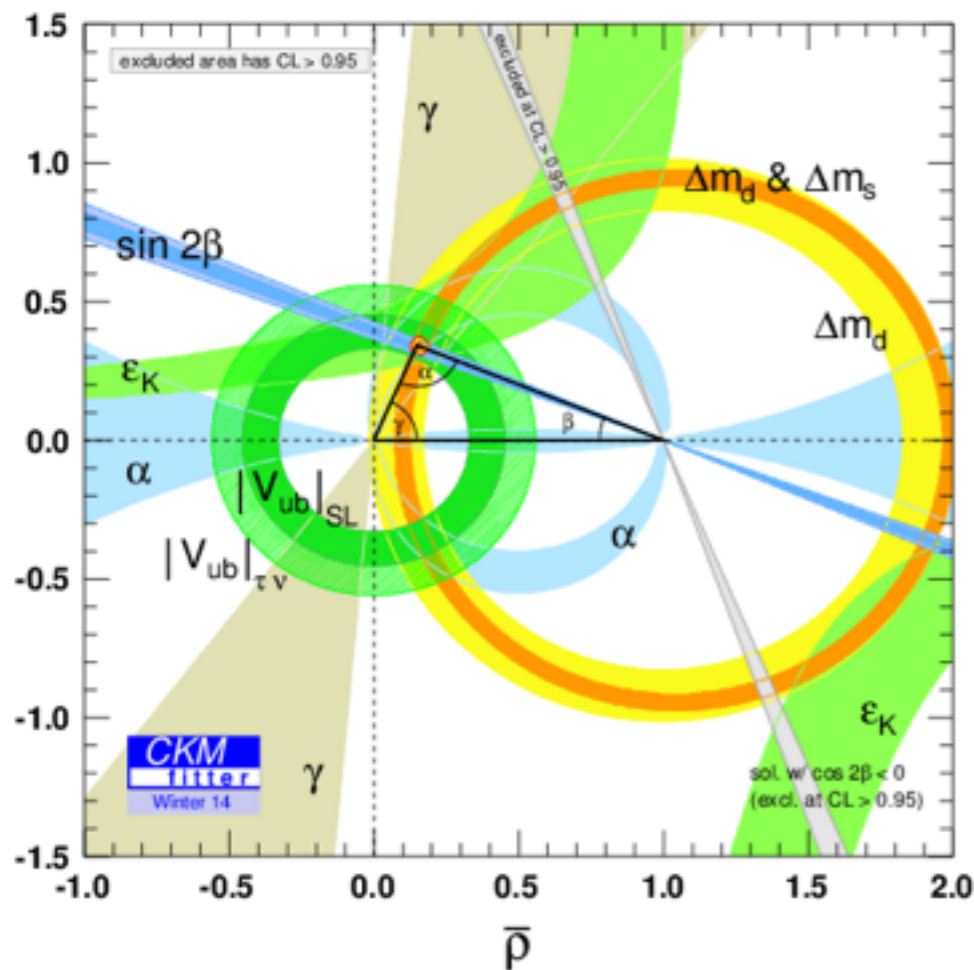
Fundamental Mixing Today

Quarks (CKM)

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

Neutrino (PMNS)

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix}$$



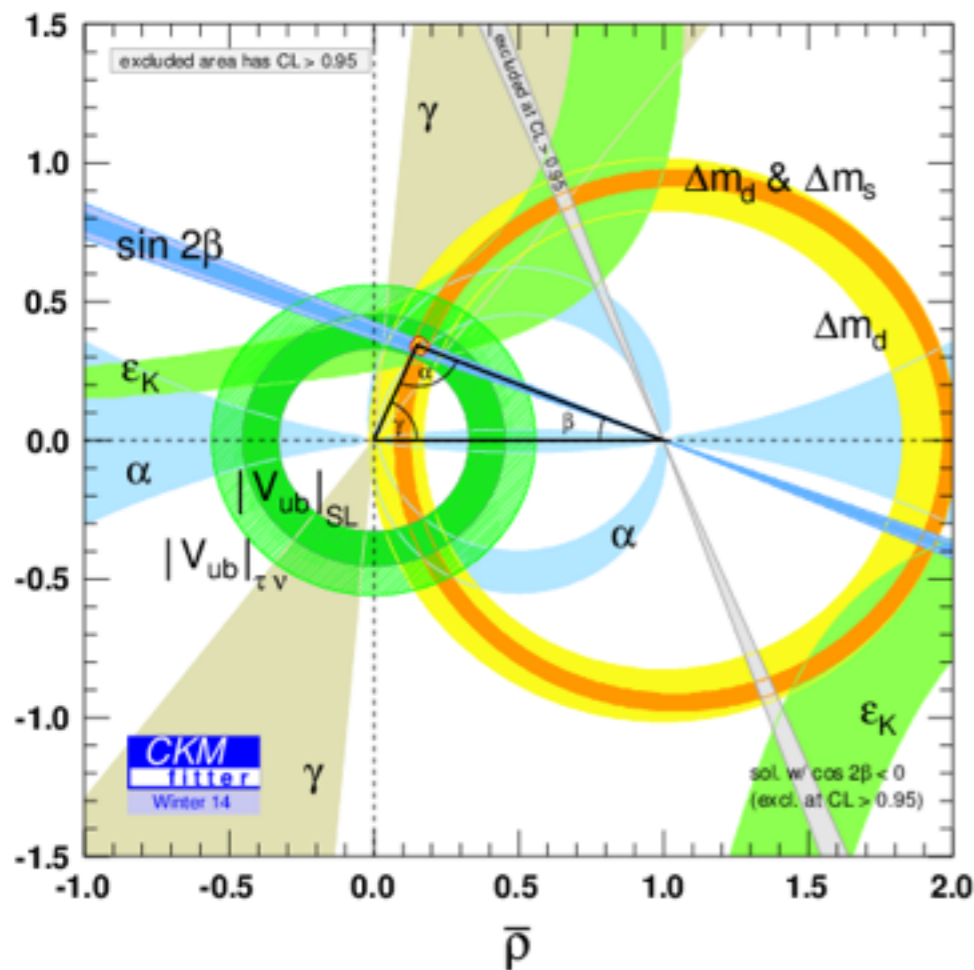
Confirms Unitarity

Currently Assumes Unitarity

Fundamental Mixing Today

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$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

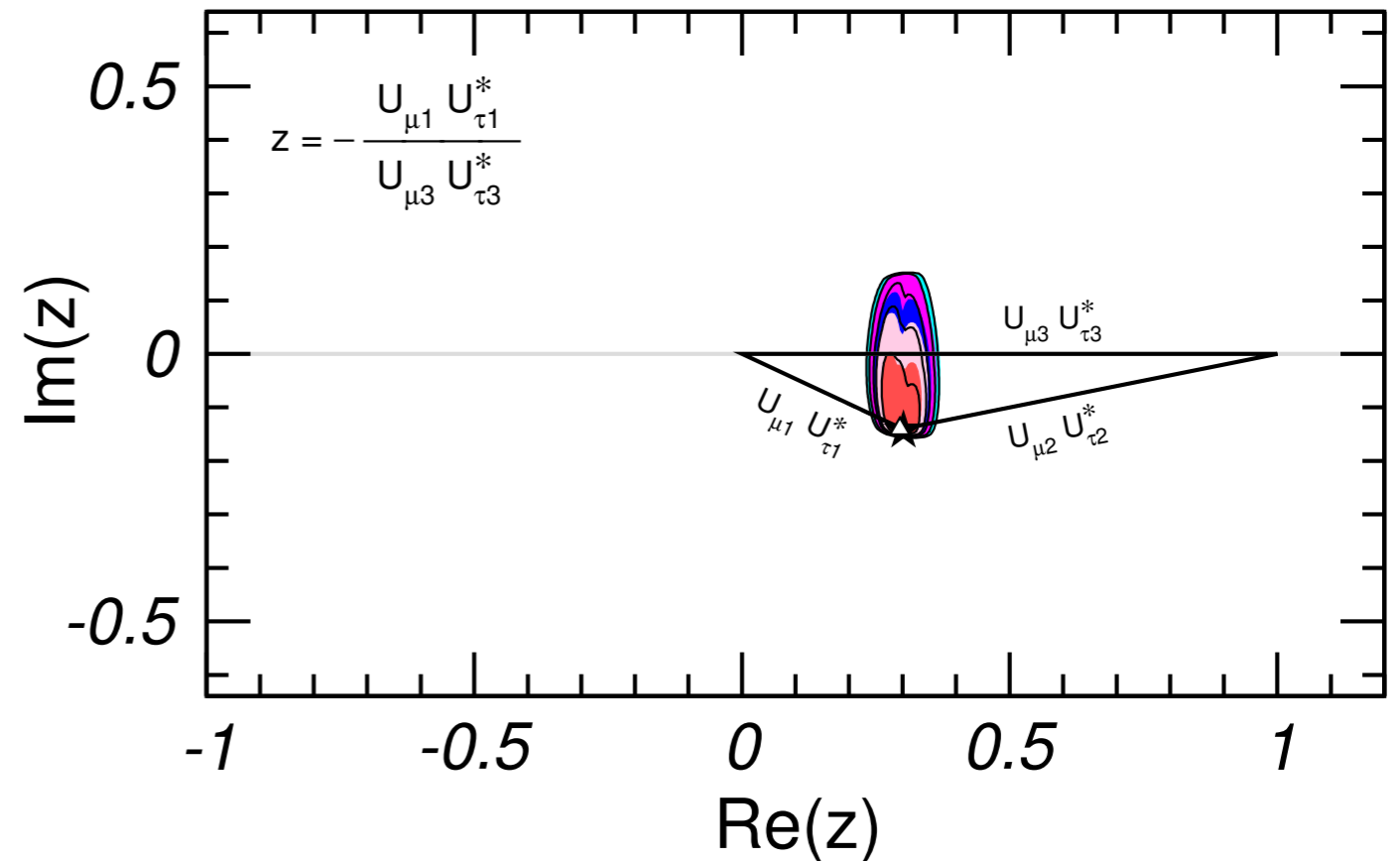


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NuFIT 2.1 (2016)



Currently Assumes Unitarity

'Current' Experimental Landscape

Accelerator Based

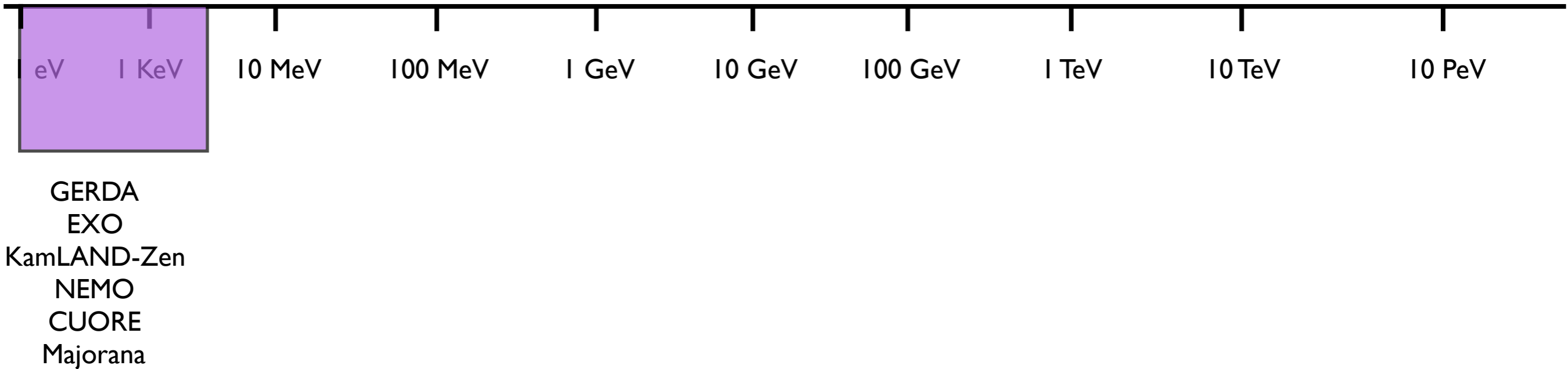


Non-Accelerator Based

*Boxes provide sense of scale for physics sensitive regions

'Current' Experimental Landscape

Accelerator Based

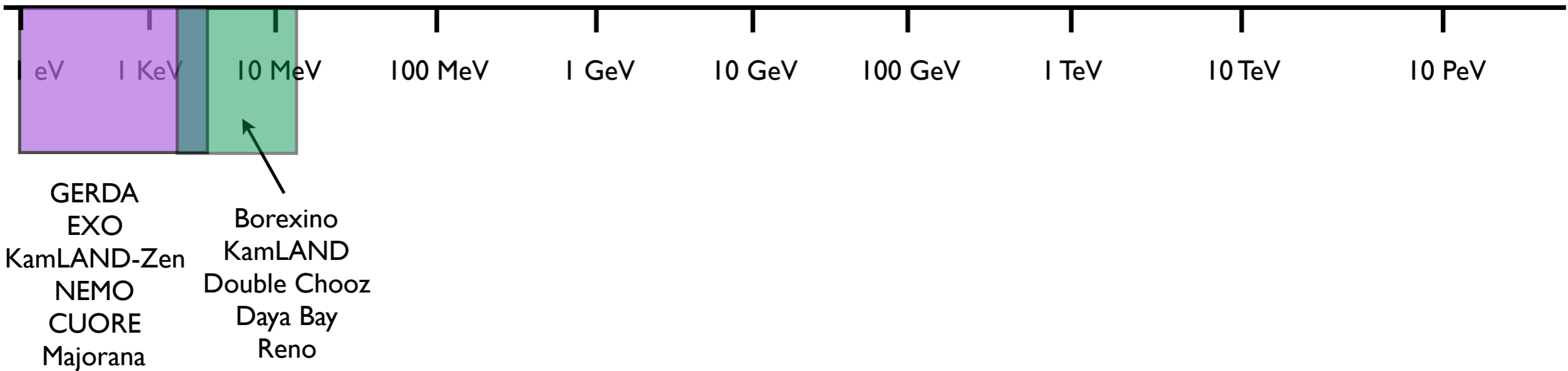


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Accelerator Based

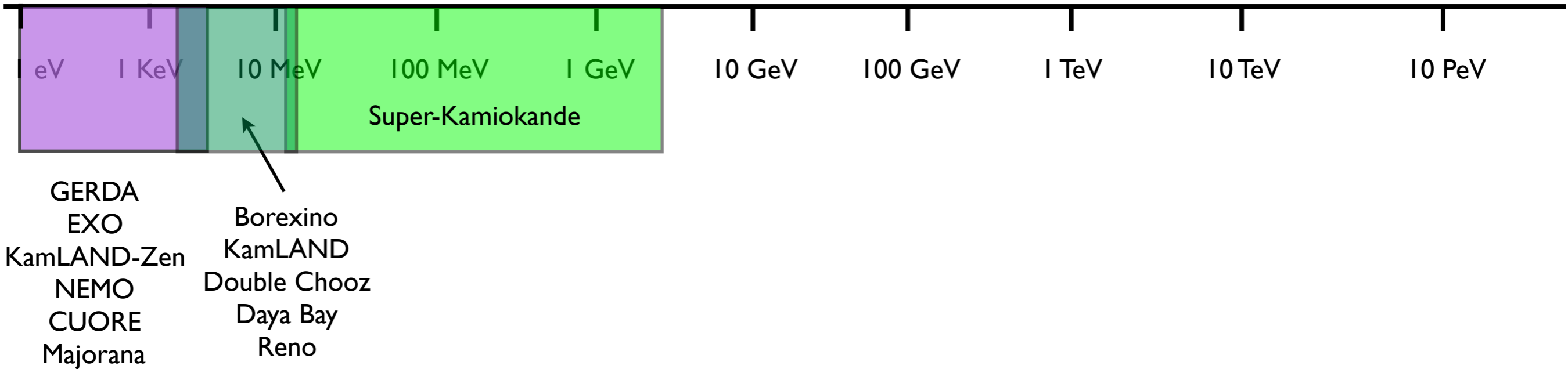


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Accelerator Based

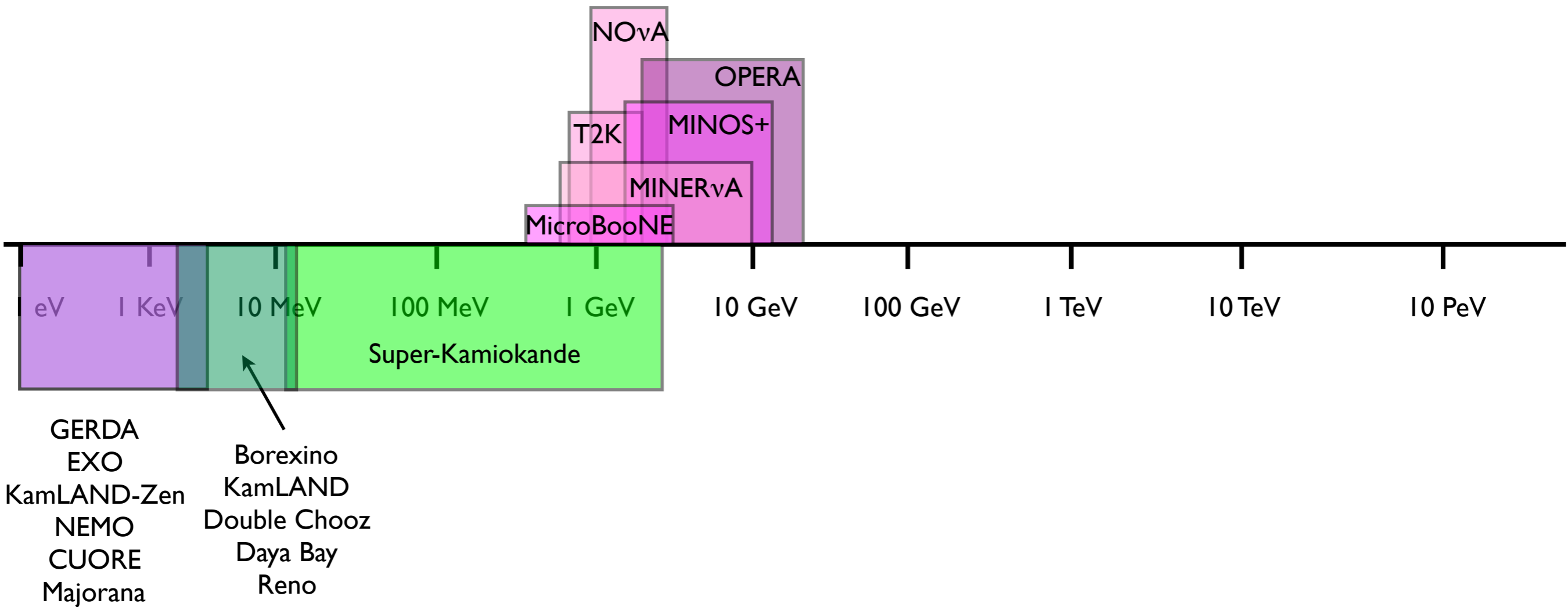


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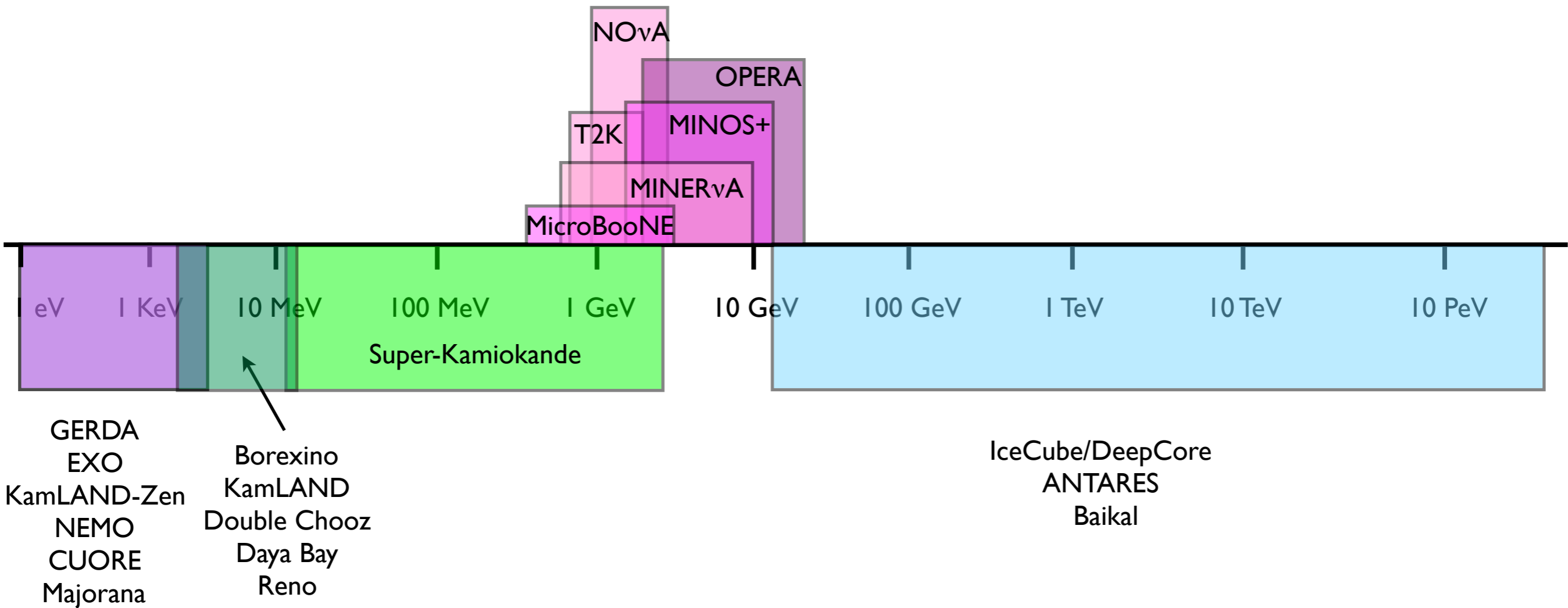


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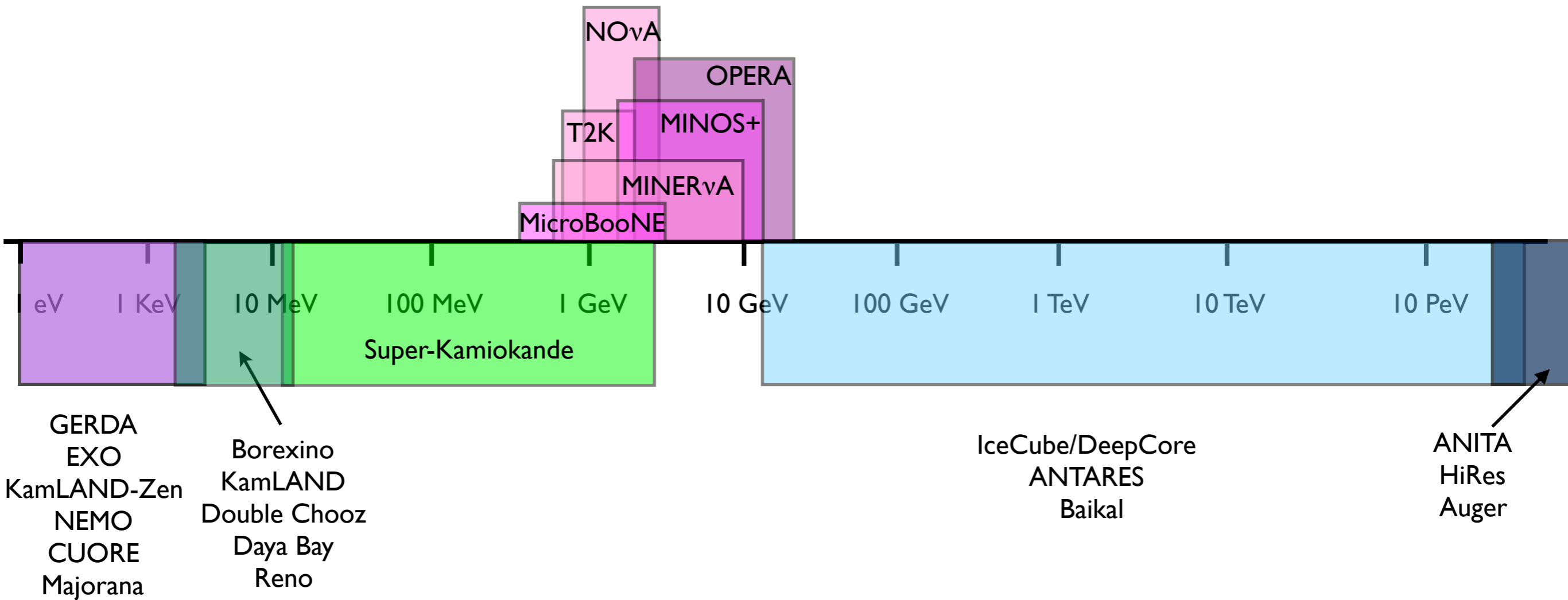


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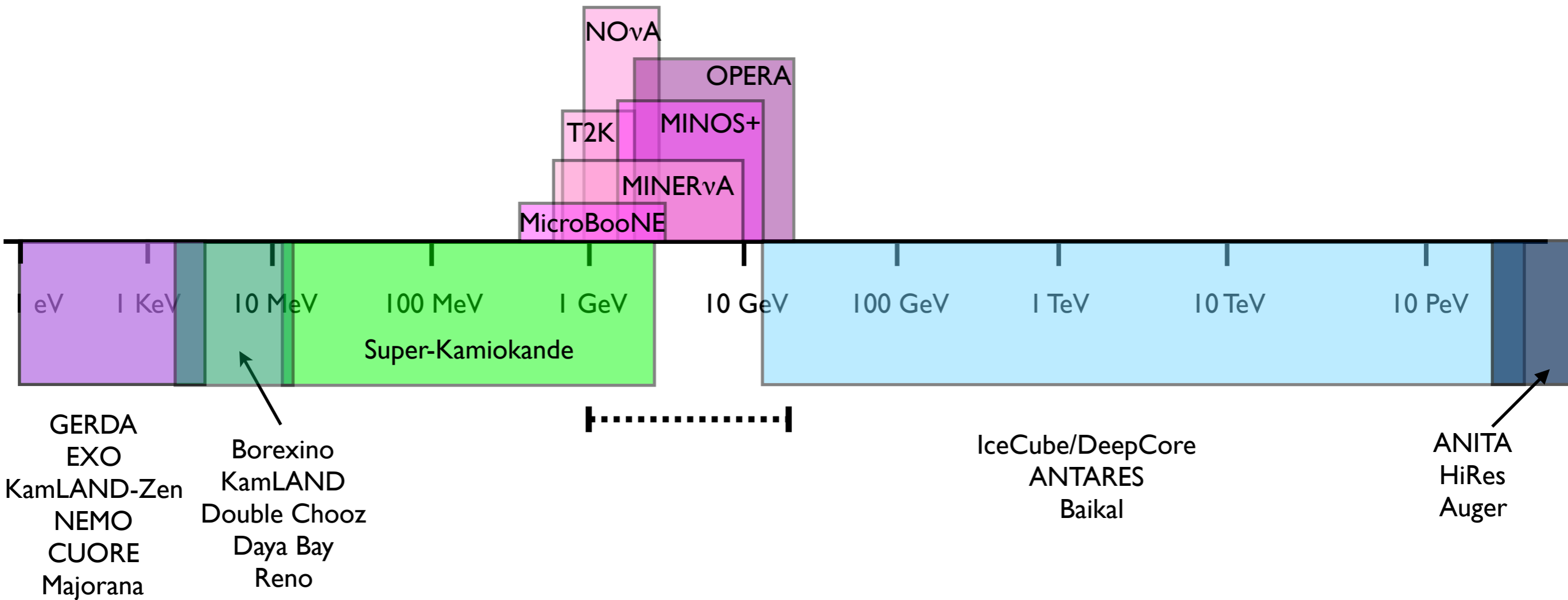


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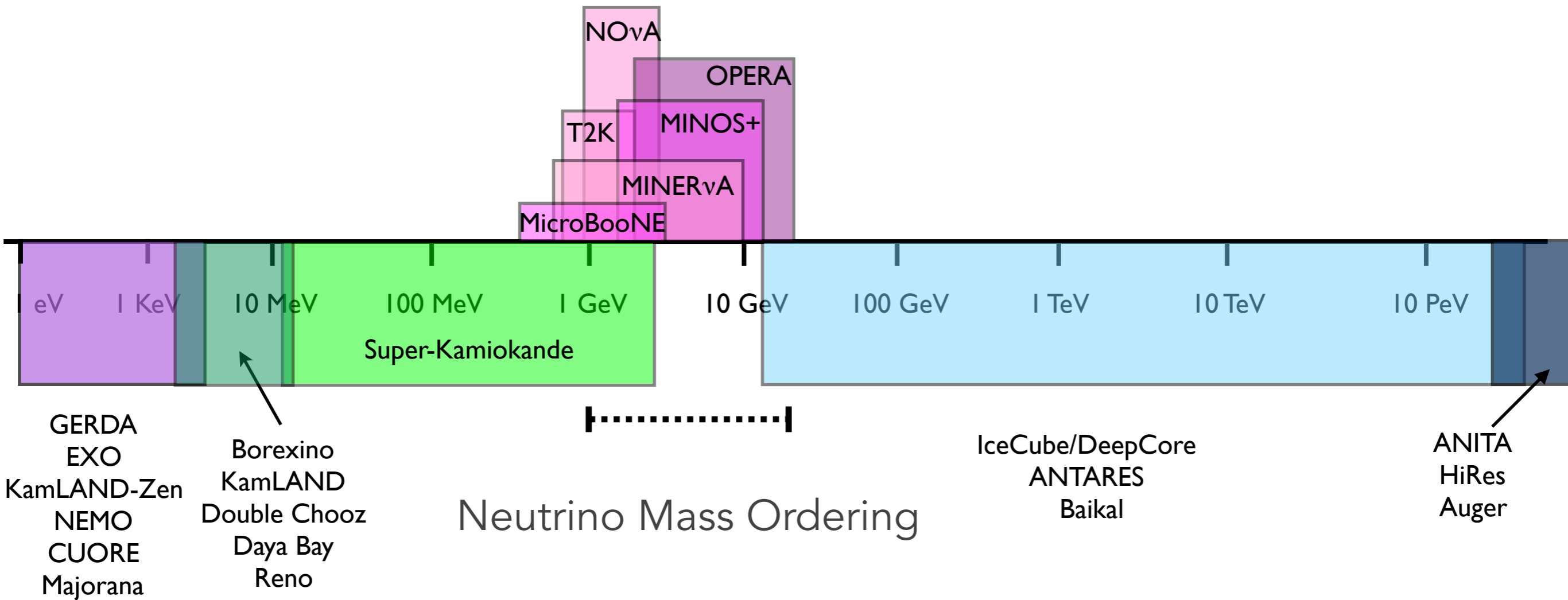


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'Current' Experimental Landscape

Accelerator Based

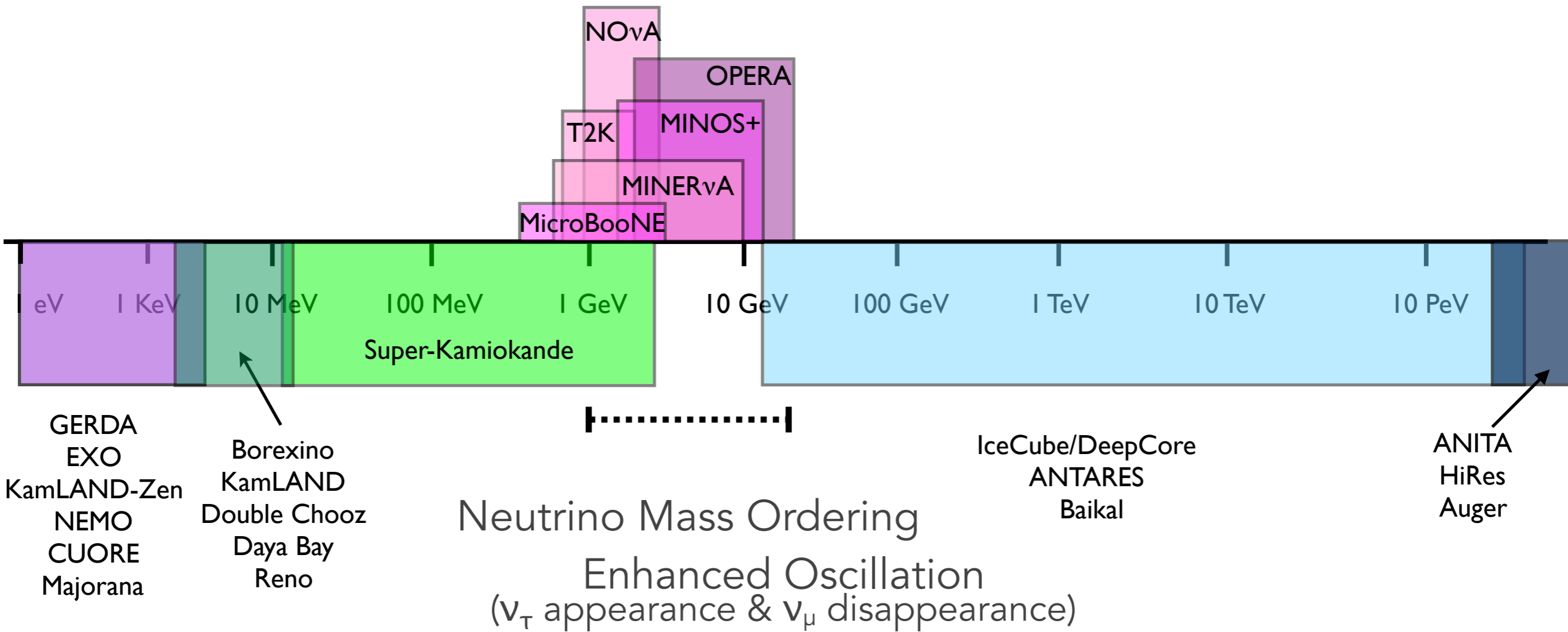


Non-Accelerator Based

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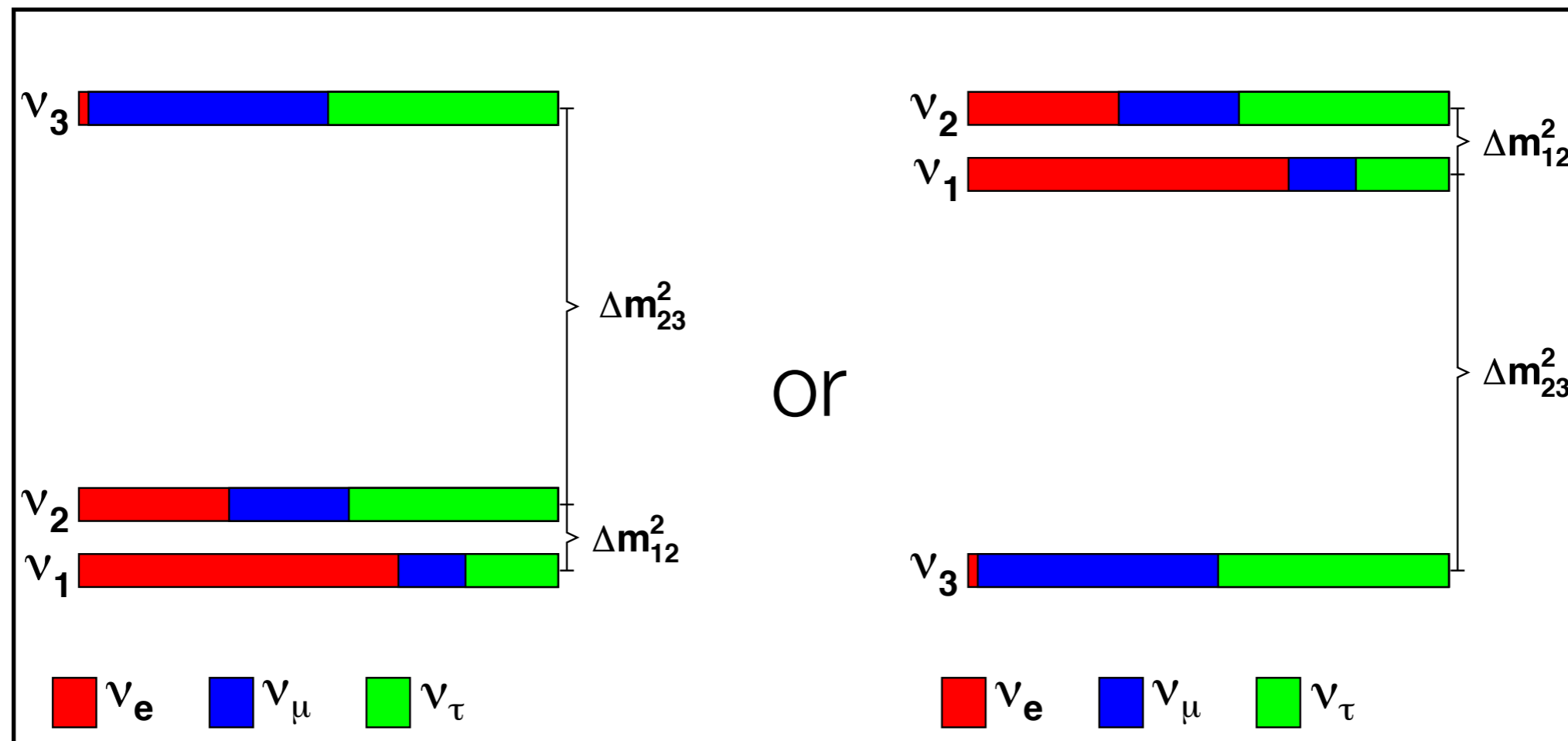
Accelerator Based



Non-Accelerator Based

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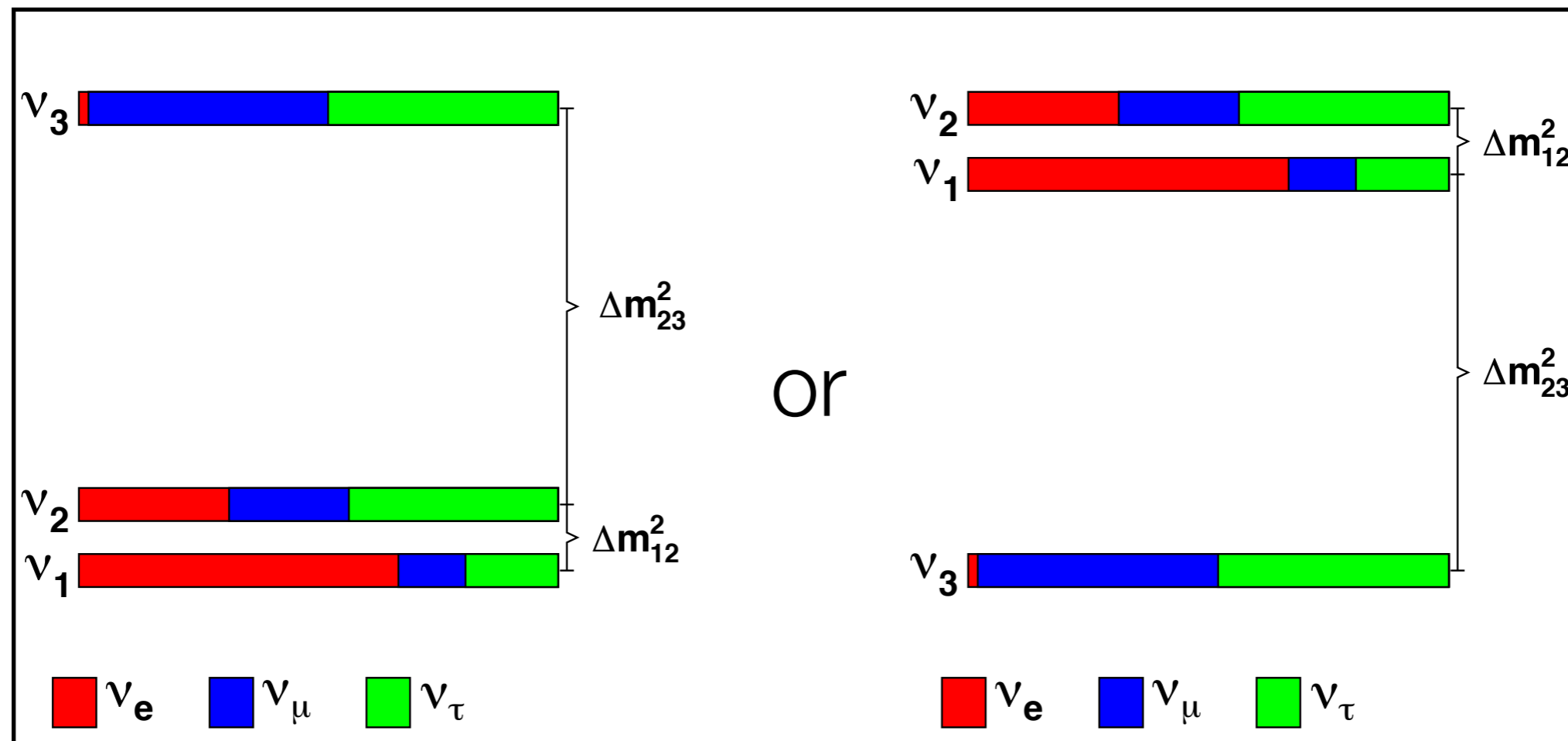
Neutrino Hierarchy/Ordering



n.b. "Neutrino Mass Ordering" and "Neutrino Mass Hierarchy" are used interchangeably

Neutrino Hierarchy/Ordering

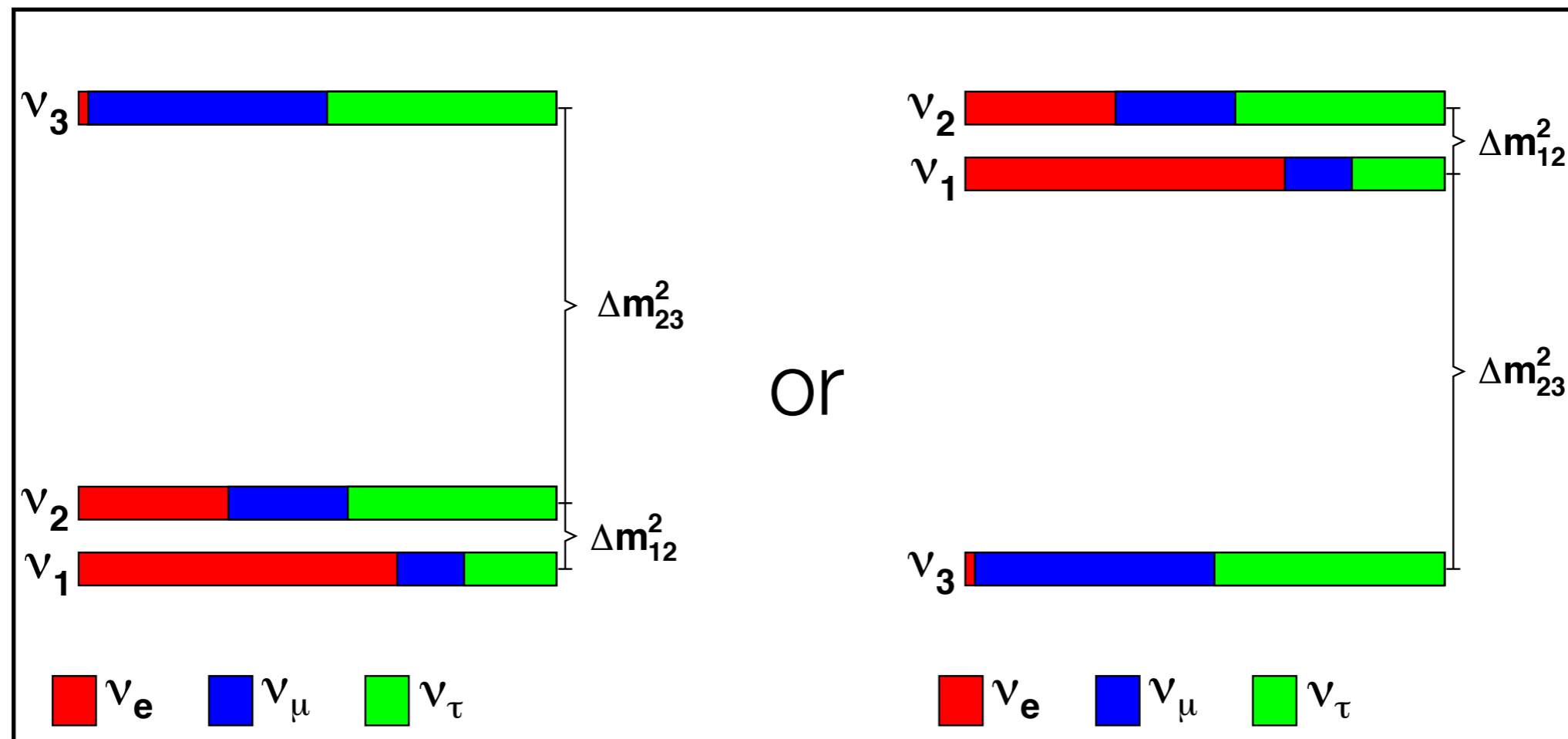
- A big detector with sensitivity at $\mathcal{O}(1)$ GeV energy is in the range of measuring the neutrino hierarchy/ordering



n.b. "Neutrino Mass Ordering" and "Neutrino Mass Hierarchy" are used interchangeably

Neutrino Hierarchy/Ordering

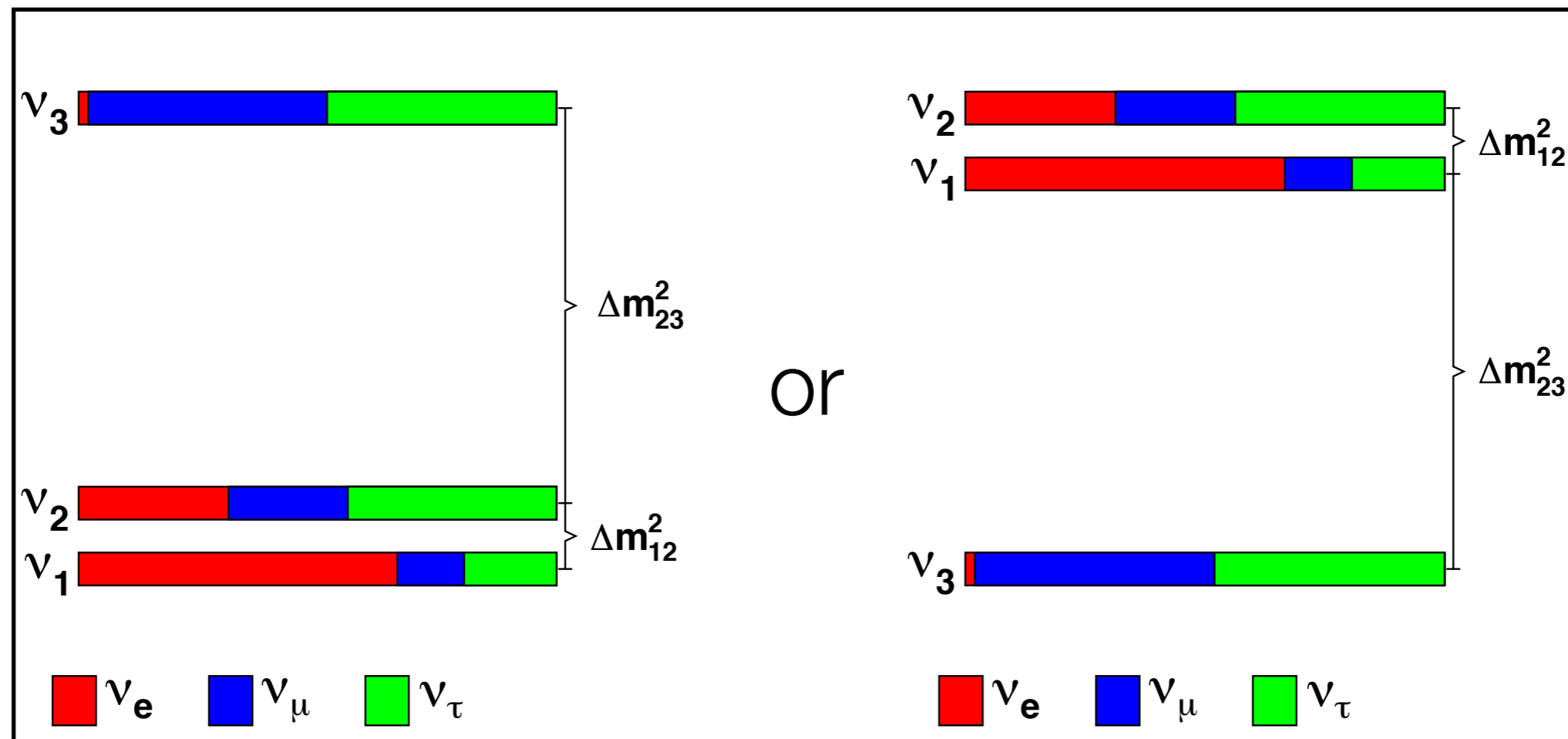
- A big detector with sensitivity at $\mathcal{O}(1)$ GeV energy is in the range of measuring the neutrino hierarchy/ordering
- Ok, fine. But, why would anyone?



n.b. "Neutrino Mass Ordering" and "Neutrino Mass Hierarchy" are used interchangeably

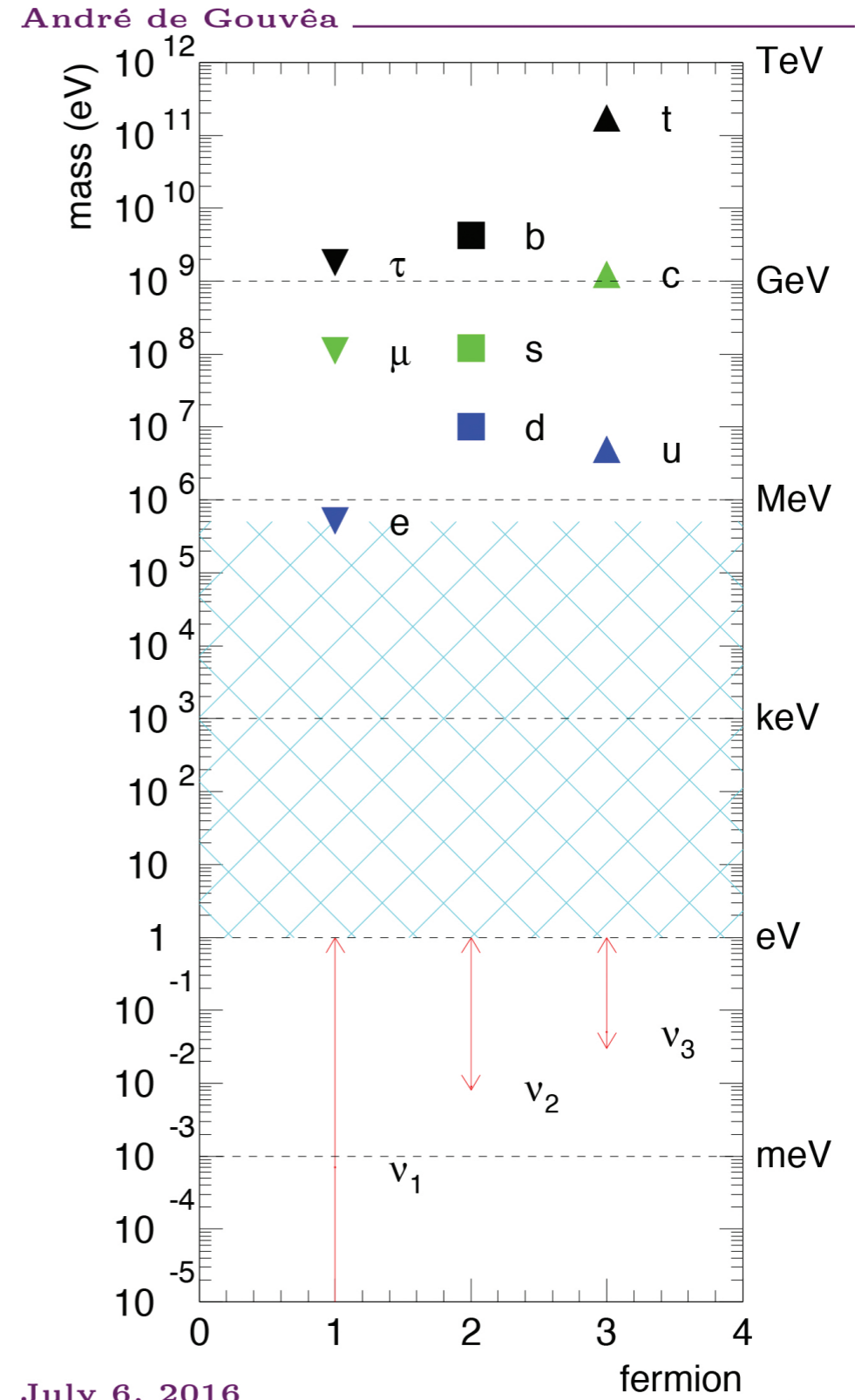
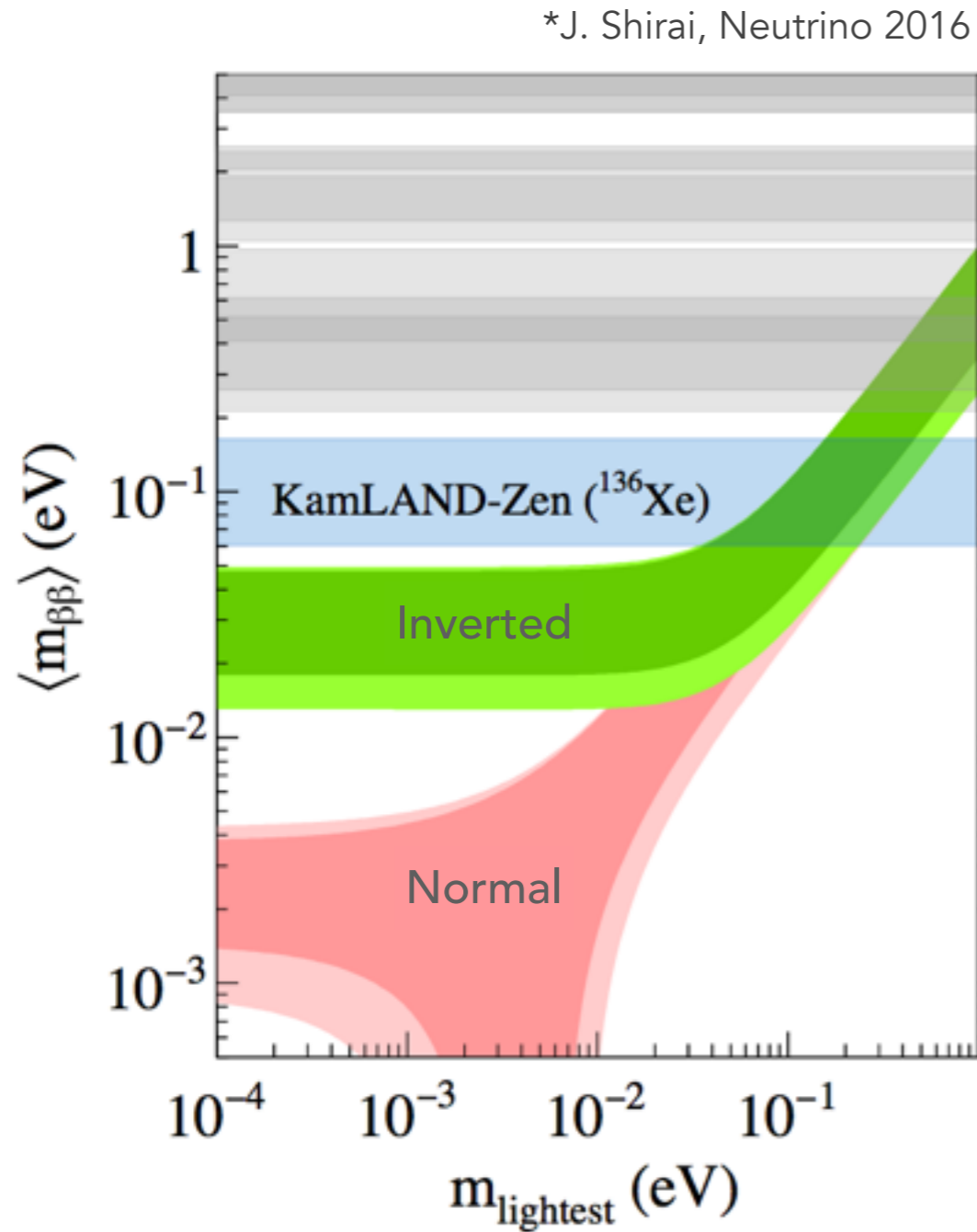
Neutrino Hierarchy/Ordering

- A big detector with sensitivity at $\mathcal{O}(1)$ GeV energy is in the range of measuring the neutrino hierarchy/ordering
- Ok, fine. But, why would anyone?
 - Besides, of course, fundamental physics is fundamental



n.b. "Neutrino Mass Ordering" and "Neutrino Mass Hierarchy" are used interchangeably

Neutrino Mass



- Why is the neutrino mass ordering (hierarchy) relevant?

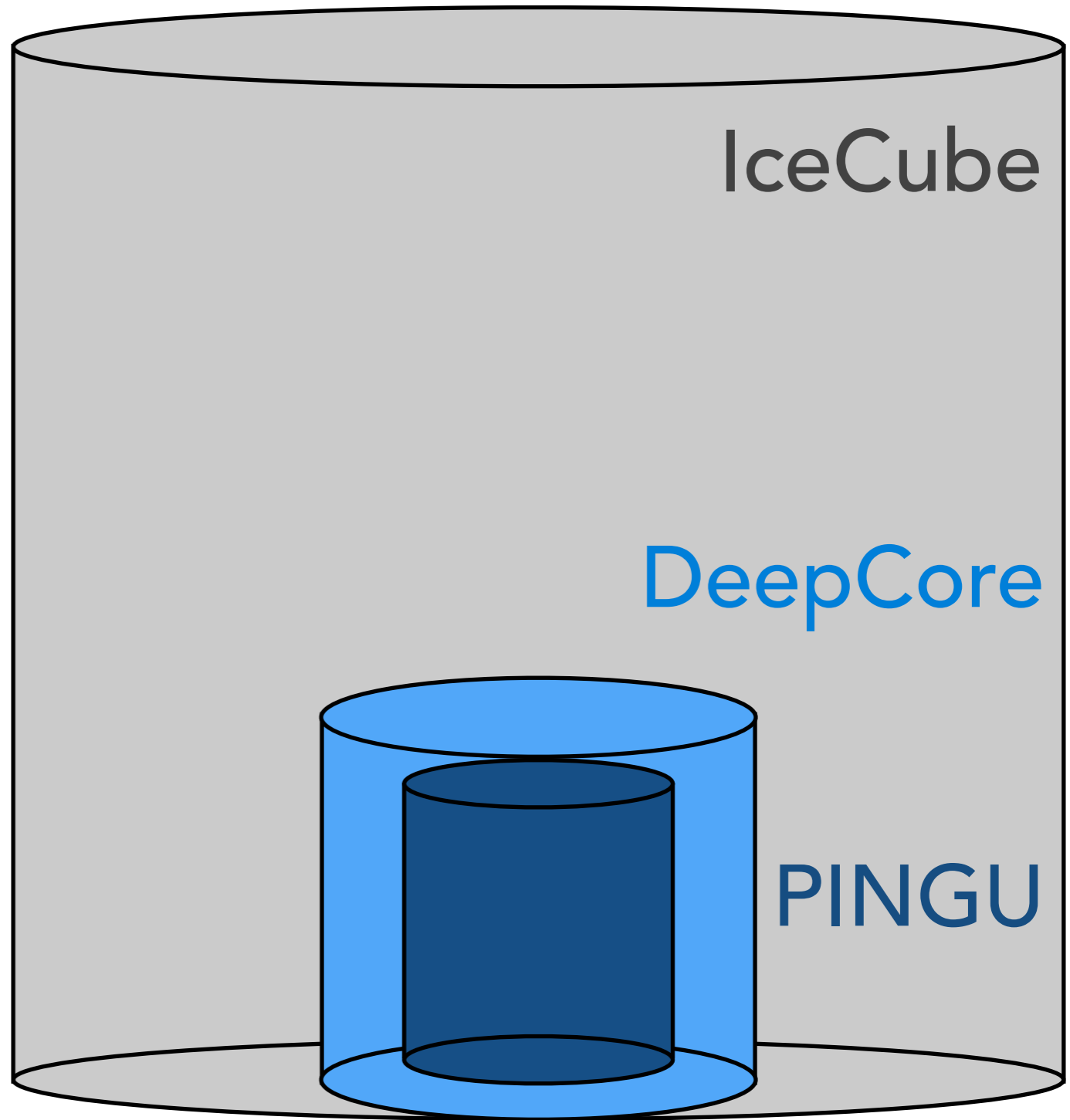
P_{recision}

I_{ceCube}

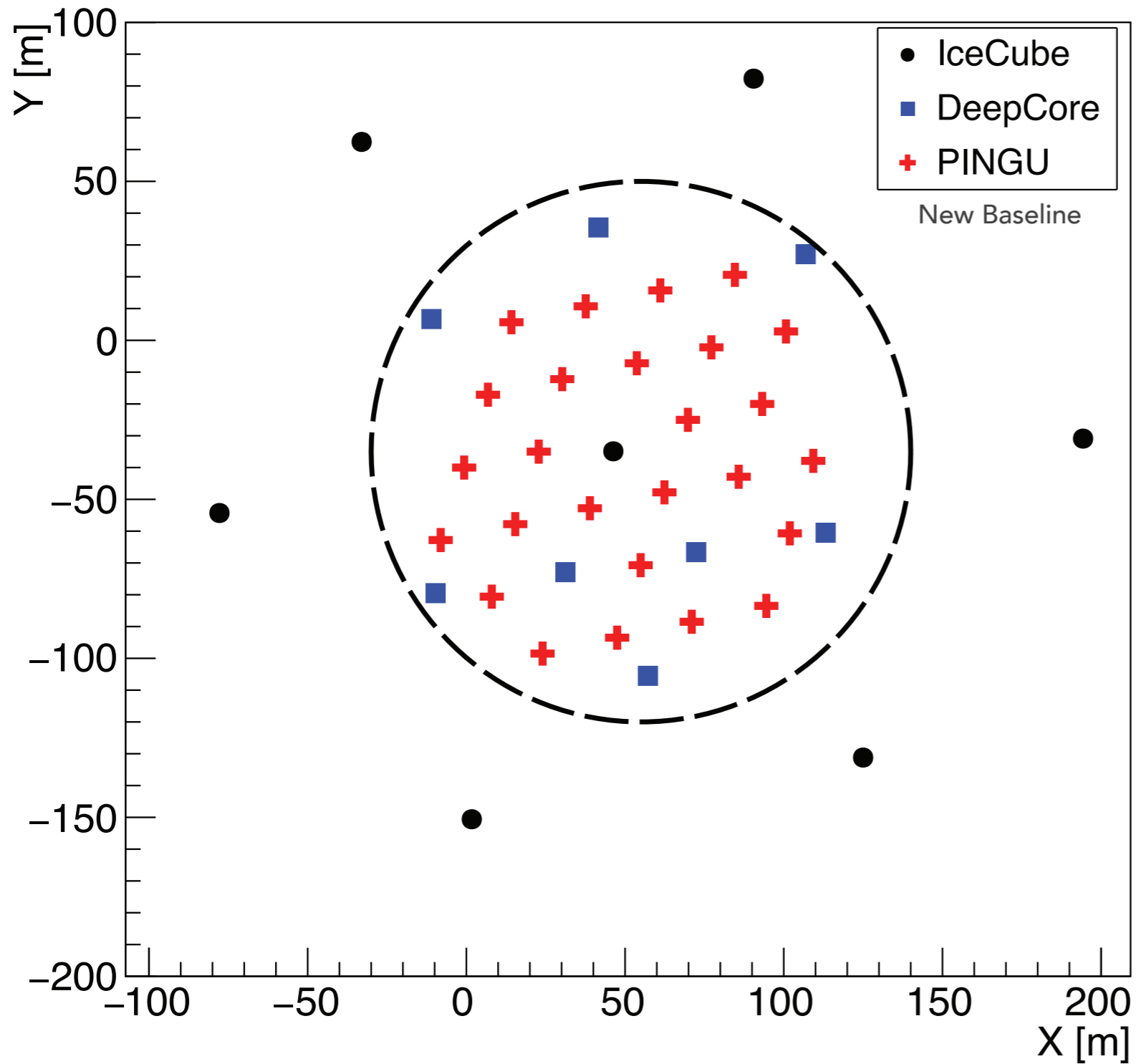
N_{ext}

G_{eneration}

U_{pgrade}



Geometry Optimization



Current

26 strings
192 DOMs/string
1.5 m DOM-DOM spacing

PINGU Simulated Event (Old)



- 9.28 GeV Neutrino, 4.9 GeV muon, 4.5 GeV cascade
- Older PINGU geometry w/ $\sim 1/3$ the number of DOMs/string, but illustrative of the potential

PINGU Simulated Event (Old)

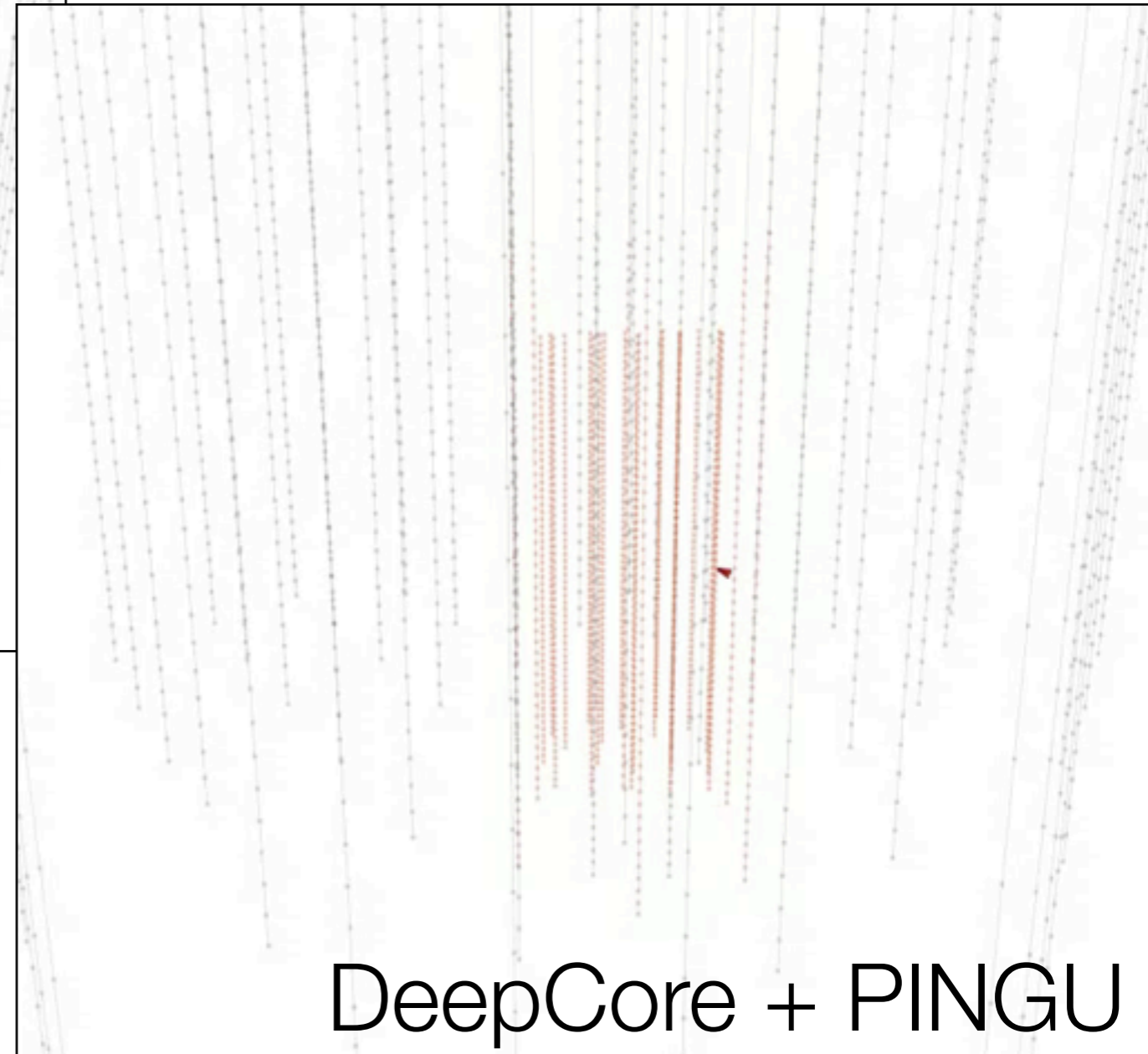


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PINGU Simulated Event (Old)

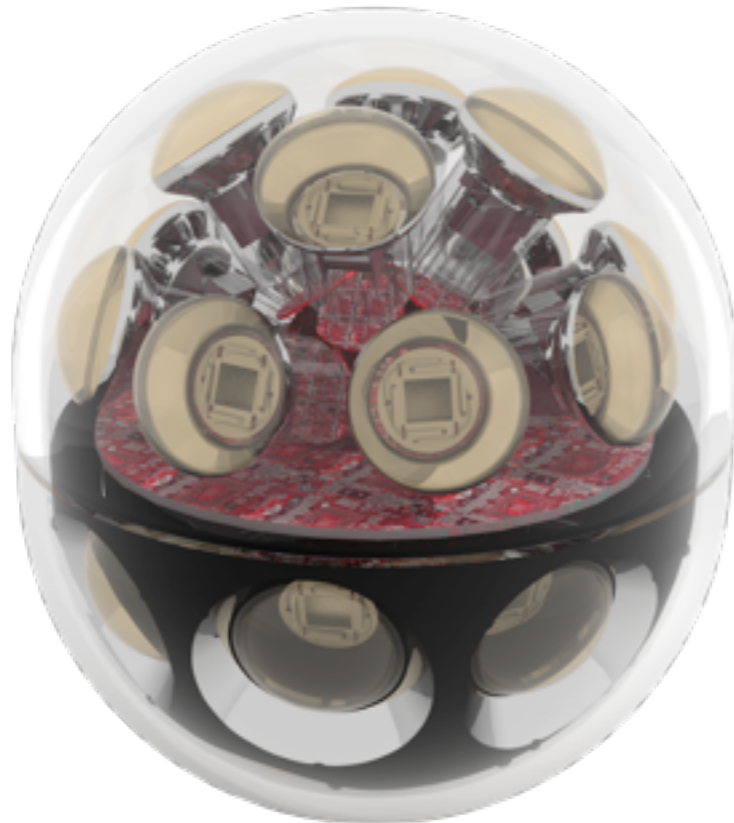


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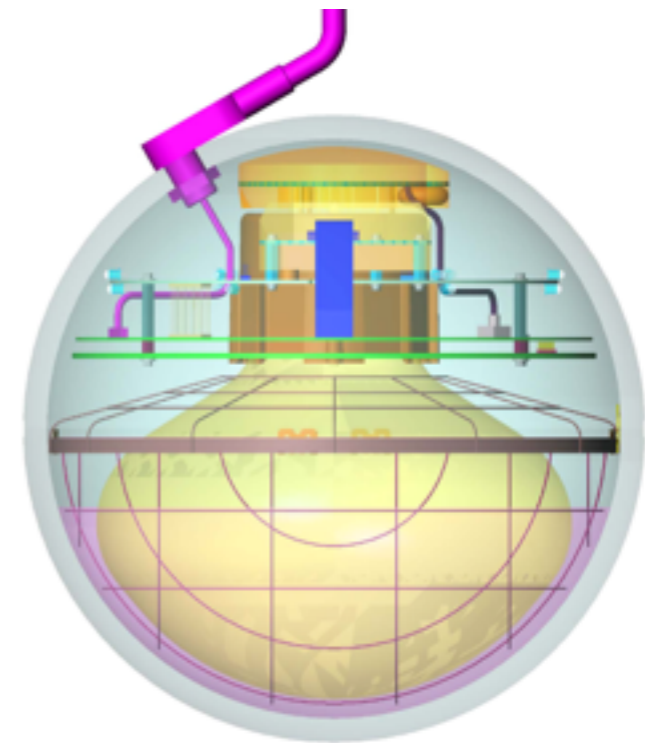


PINGU Technology Options

- Conventional 10" single PMT DOM
- Possibility for using multi-PMT DOM (mDOM)
 - 3" PMTs providing almost 4π angular coverage
 - Up to factor 2 increase in photon collection vs. 10" PMT

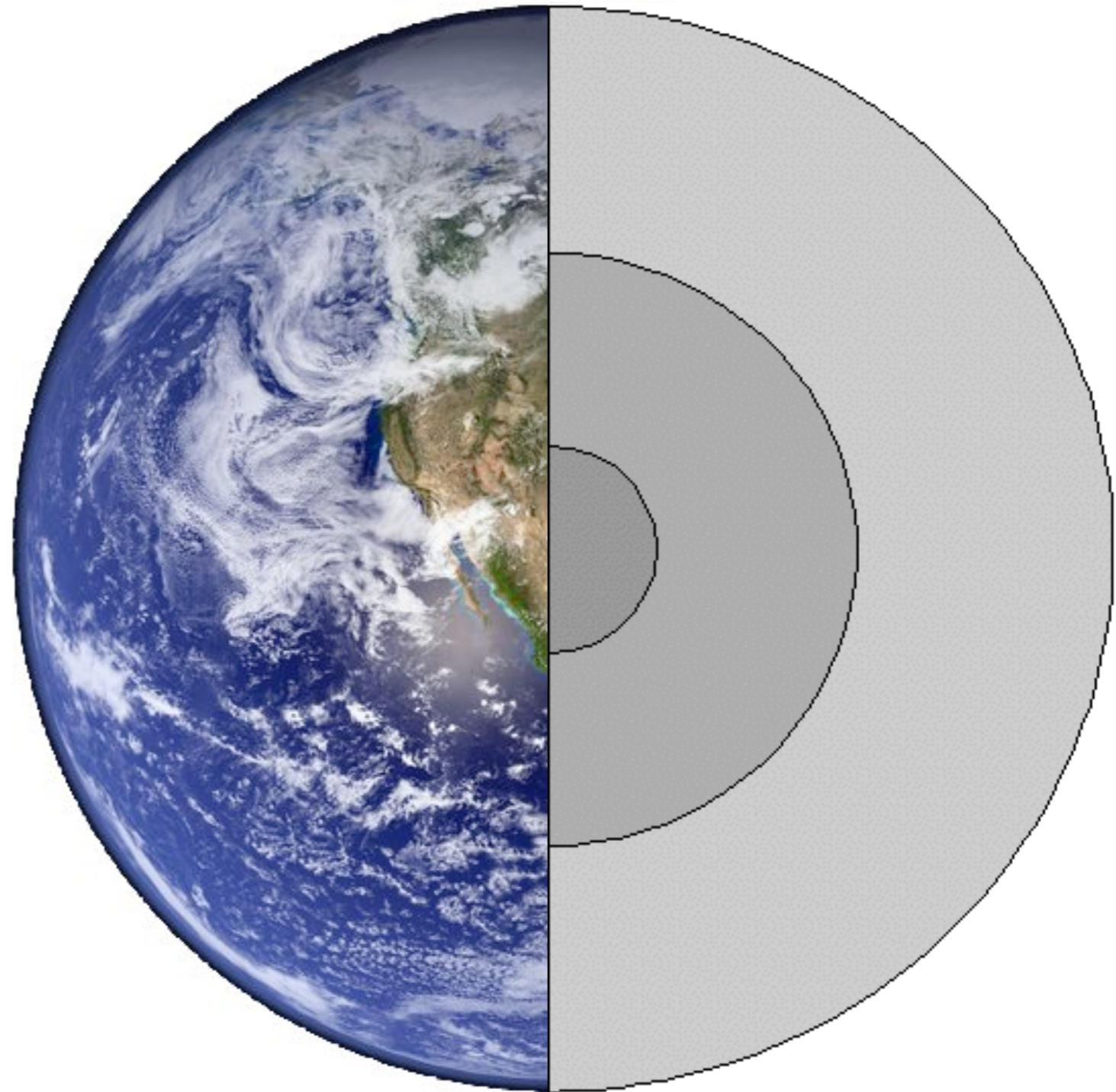


*L. Classen, ICRC 2015 (proc.)

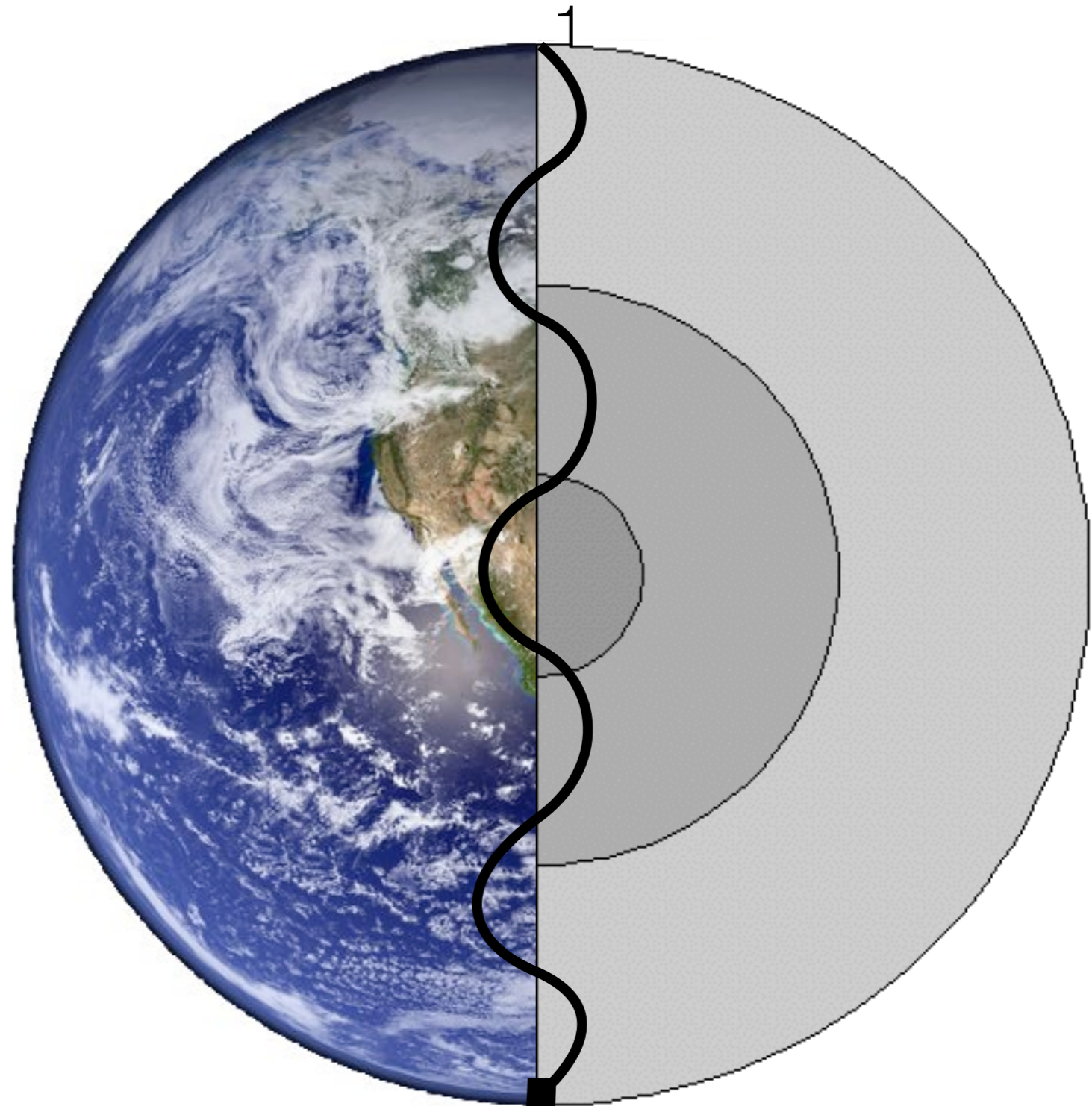
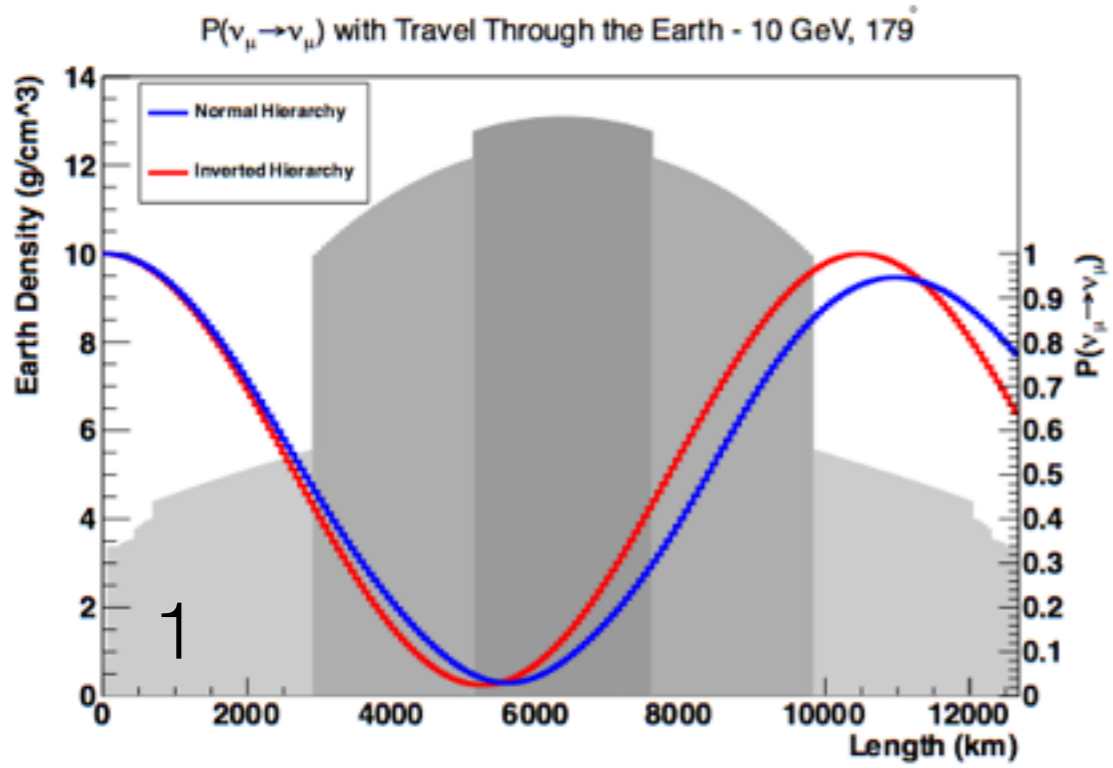


IceCube
Digital Optical Module (DOM)

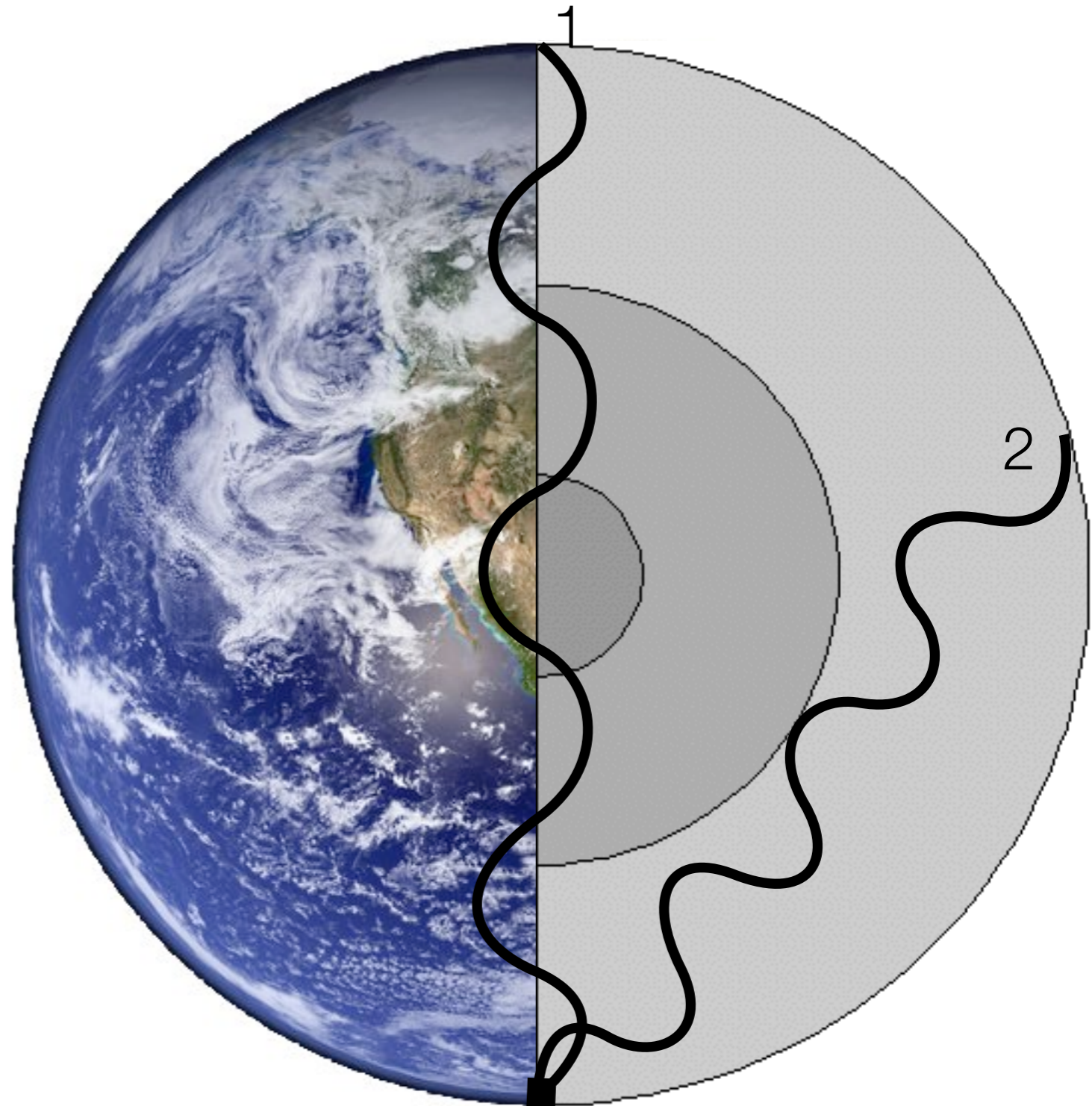
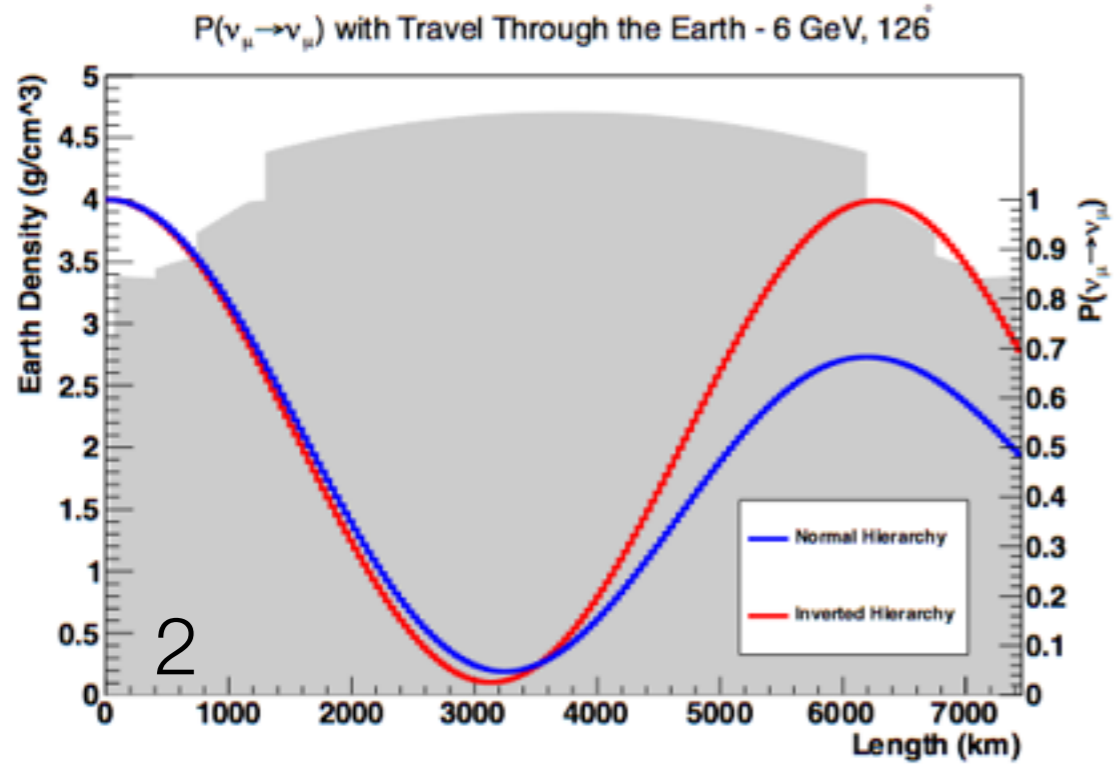
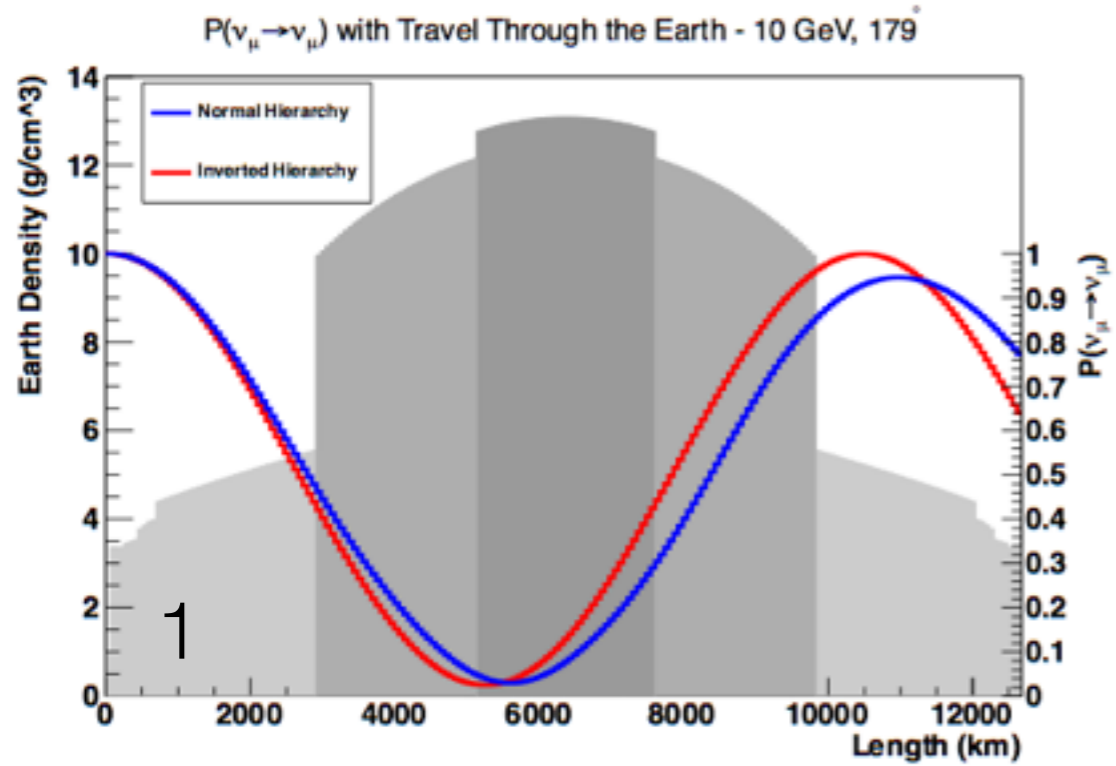
PINGU Neutrino Mass Ordering



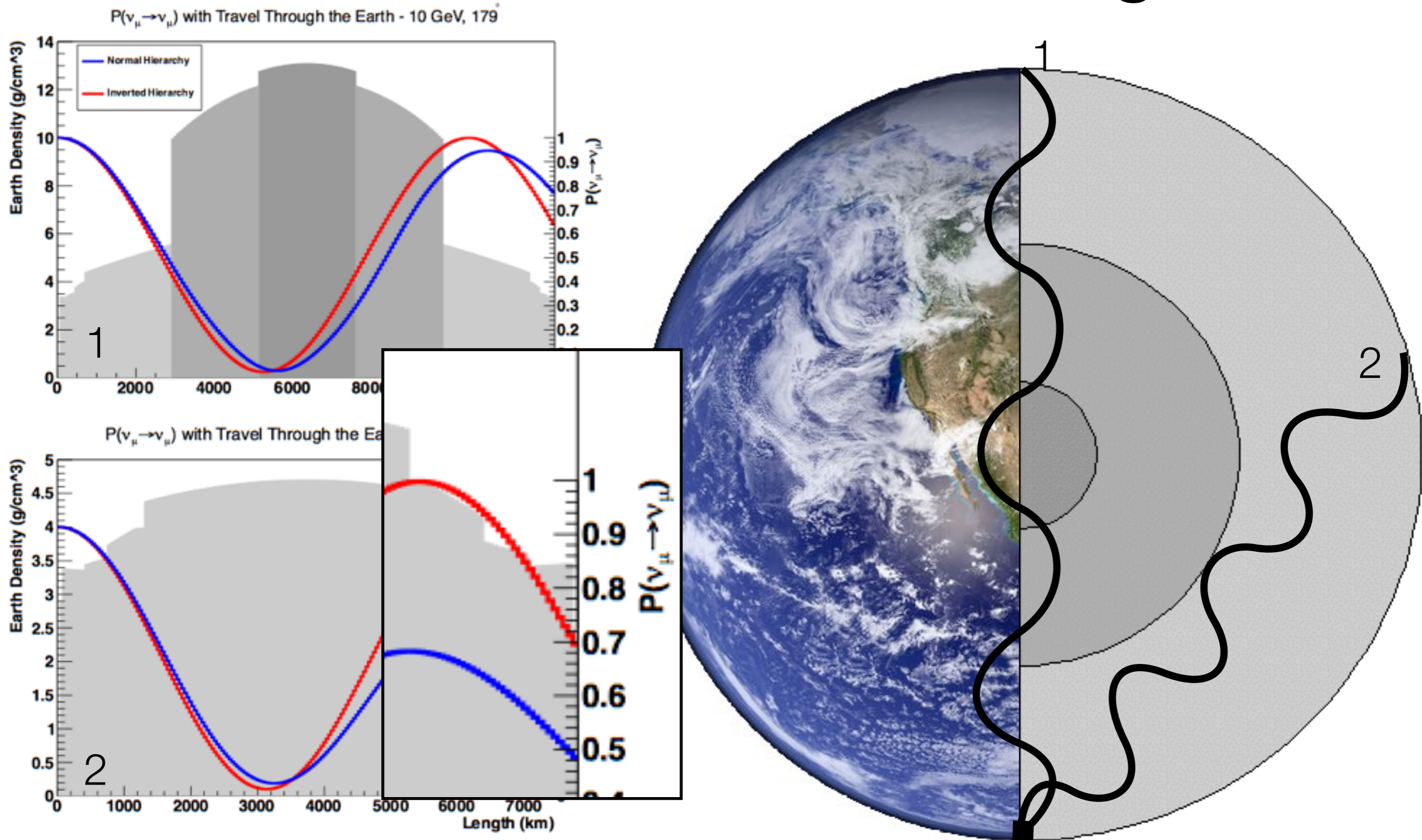
PINGU Neutrino Mass Ordering



PINGU Neutrino Mass Ordering

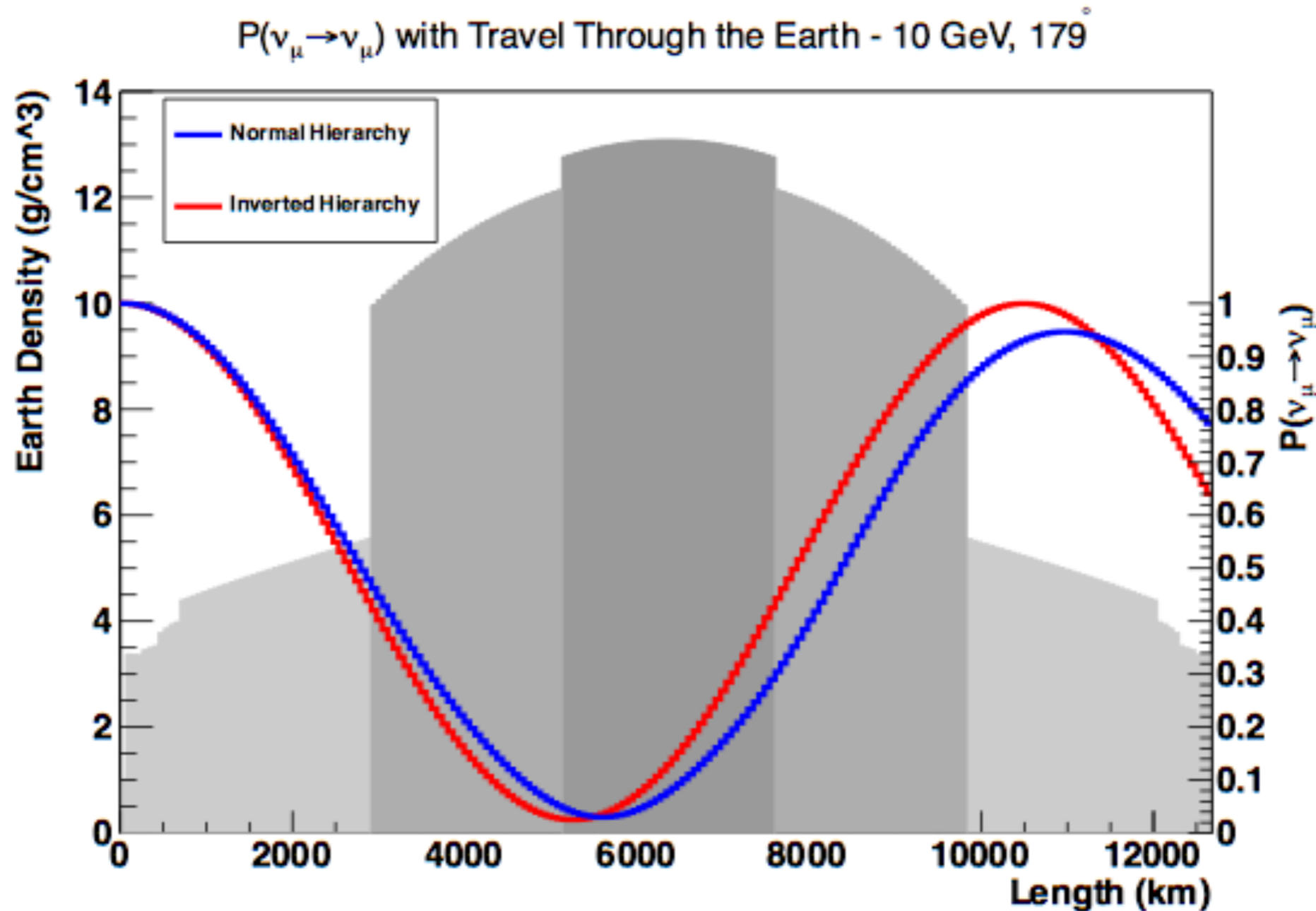


PINGU Neutrino Mass Ordering



- Inverted/Normal ordering has up to 20% different in oscillation probability for specific energies and zenith angles (baselines)

Neutrino Ordering w/ No Magnet

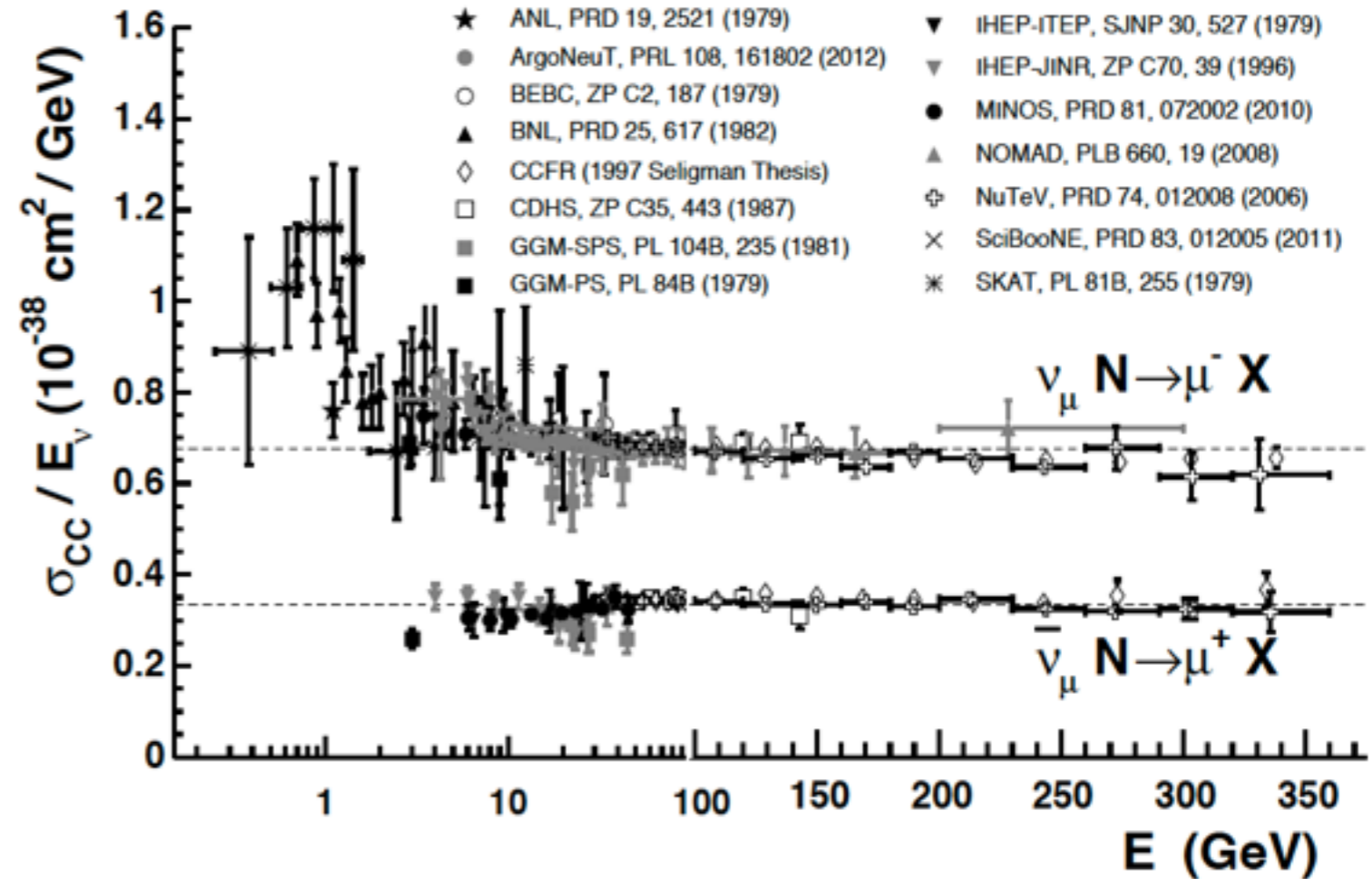
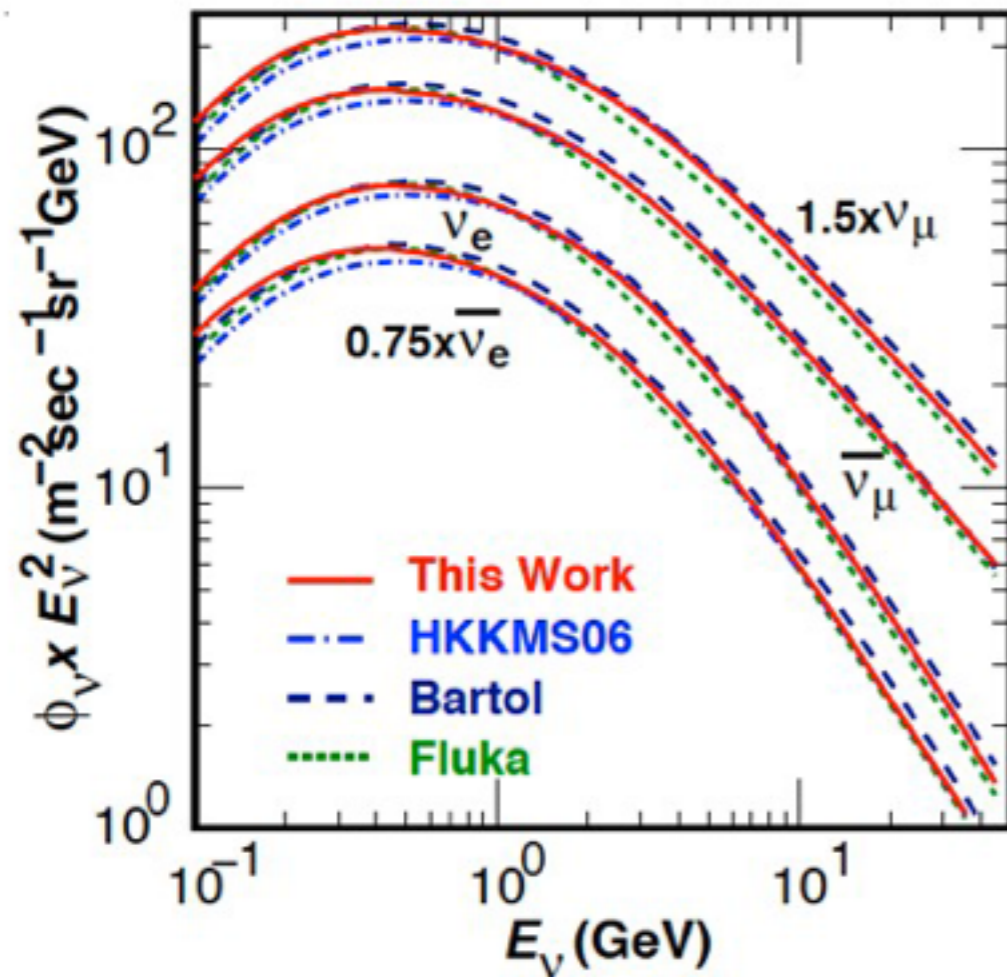


- Neutrinos will see either enhancement or suppression of oscillation probability, but anti-neutrinos will experience an opposite sign modulation of the exact same magnitude

Neutrino Ordering w/ No Magnet

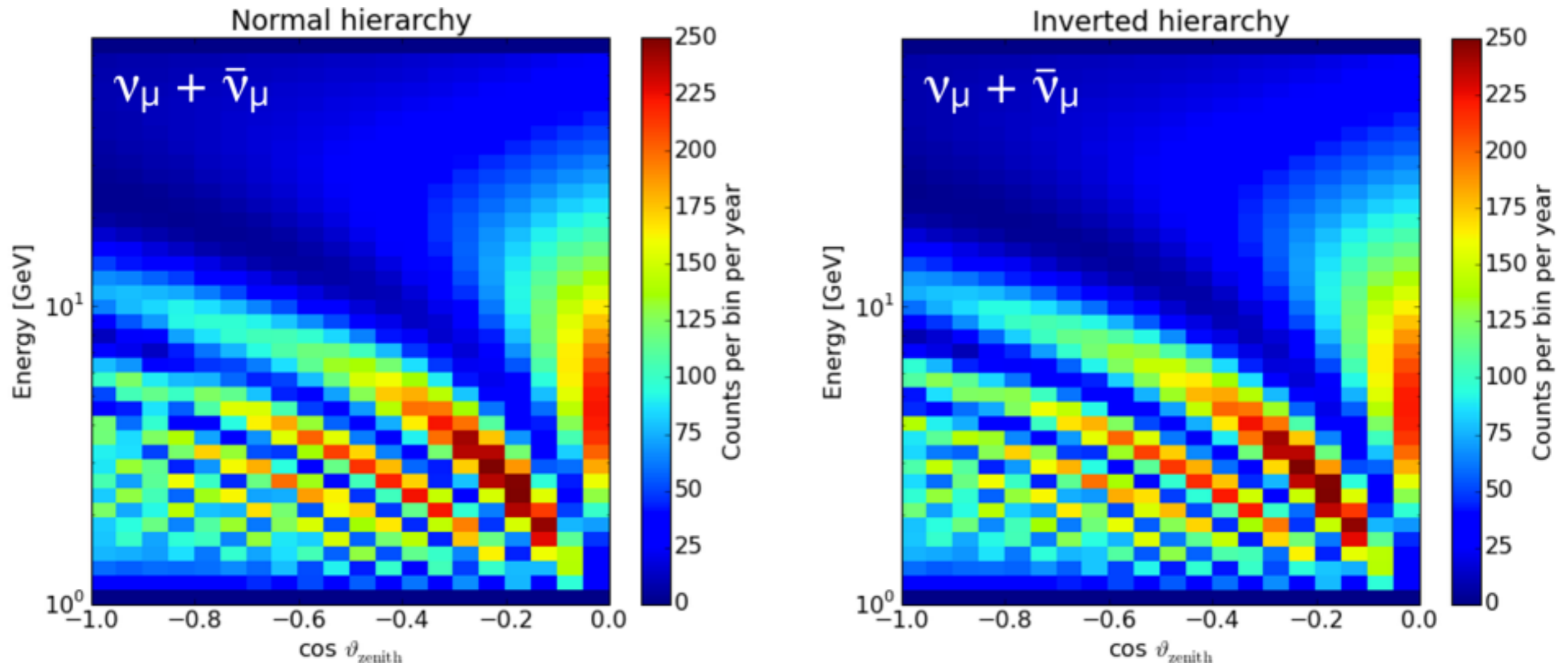
arXiv:0203272

*G. Zeller, PDG 2012



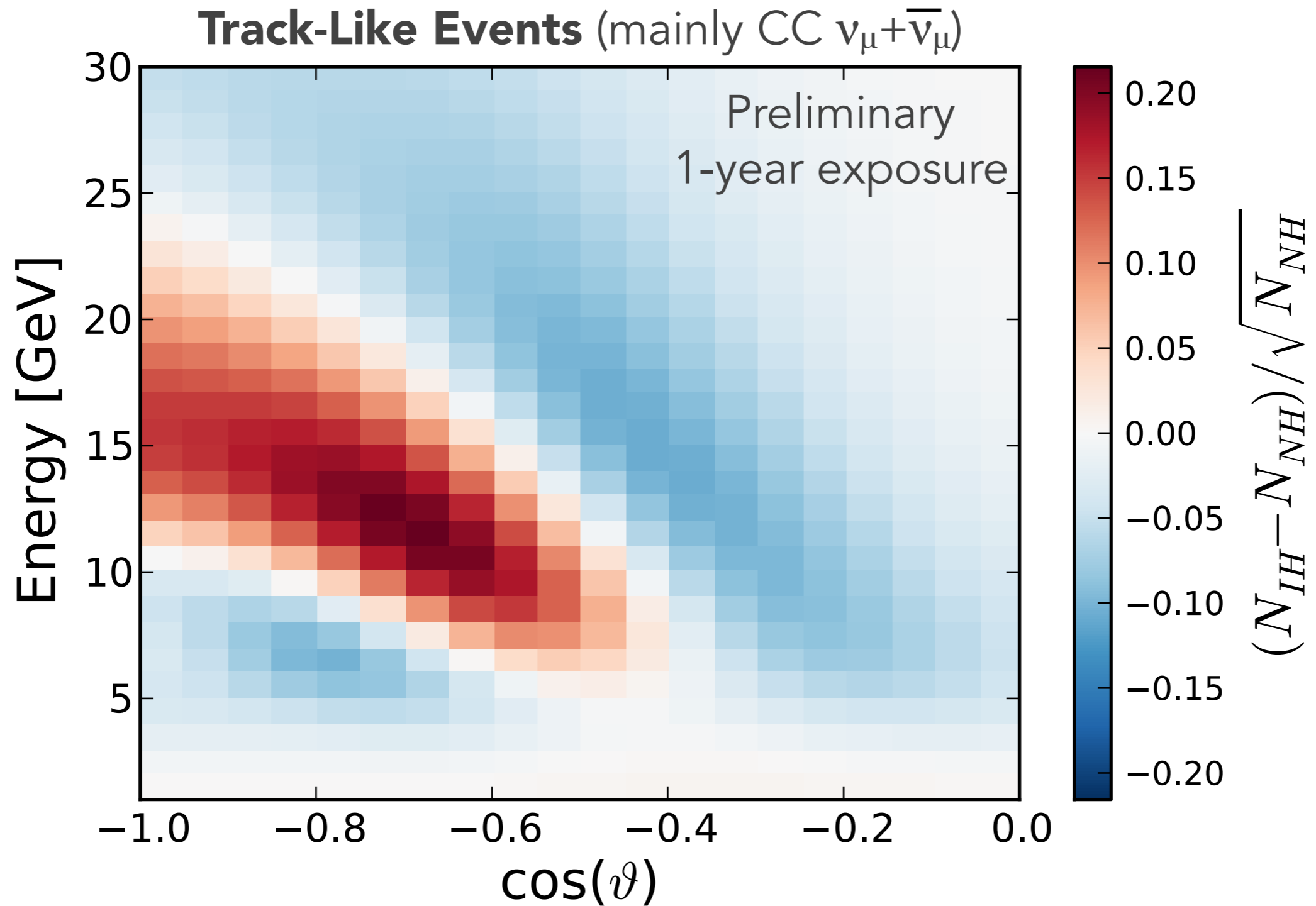
- PINGU has no magnet to separate neutrino from anti-neutrinos, but there is a cross-section and flux difference between neutrinos and anti-neutrinos

Oscillation Pattern



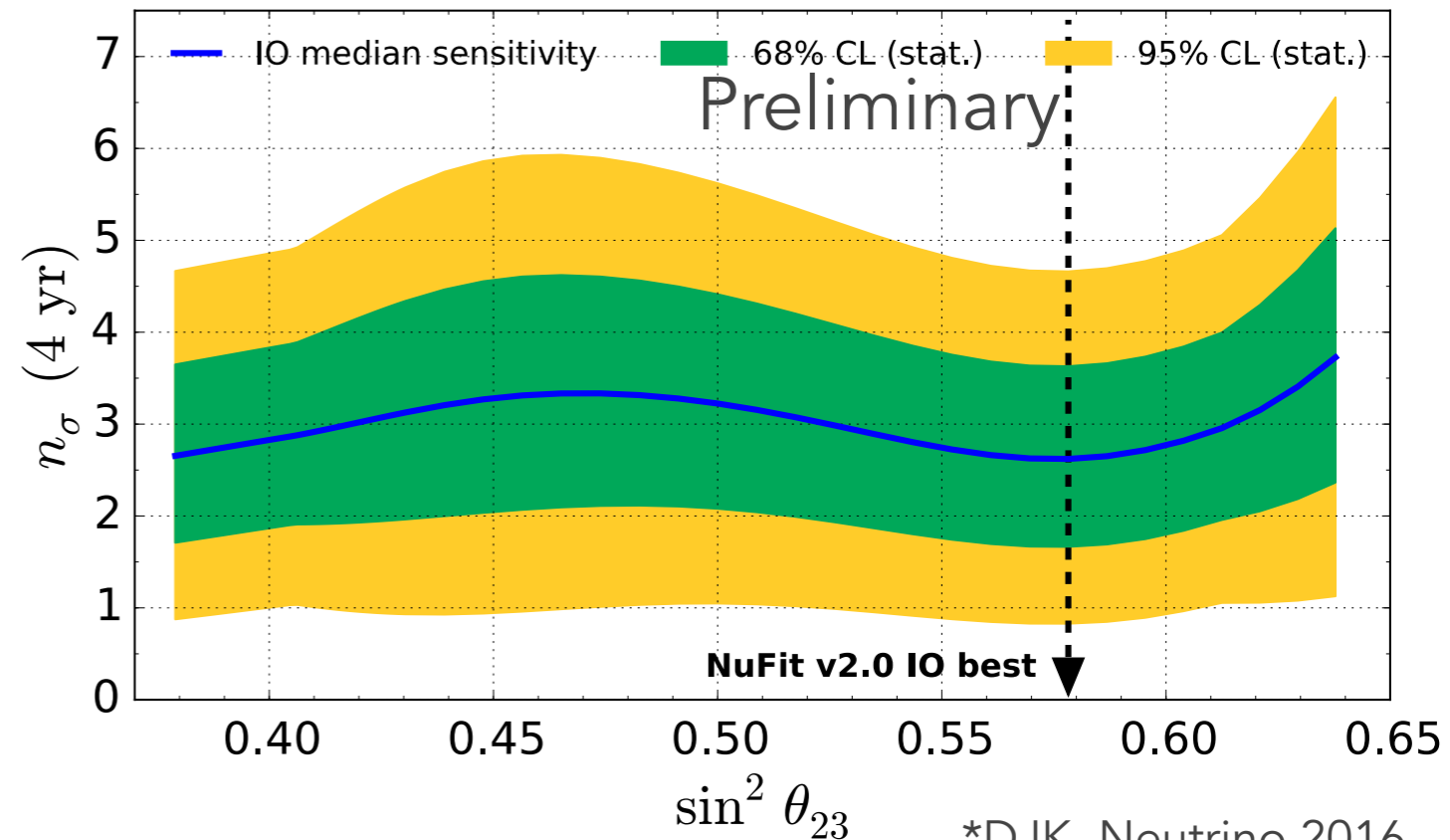
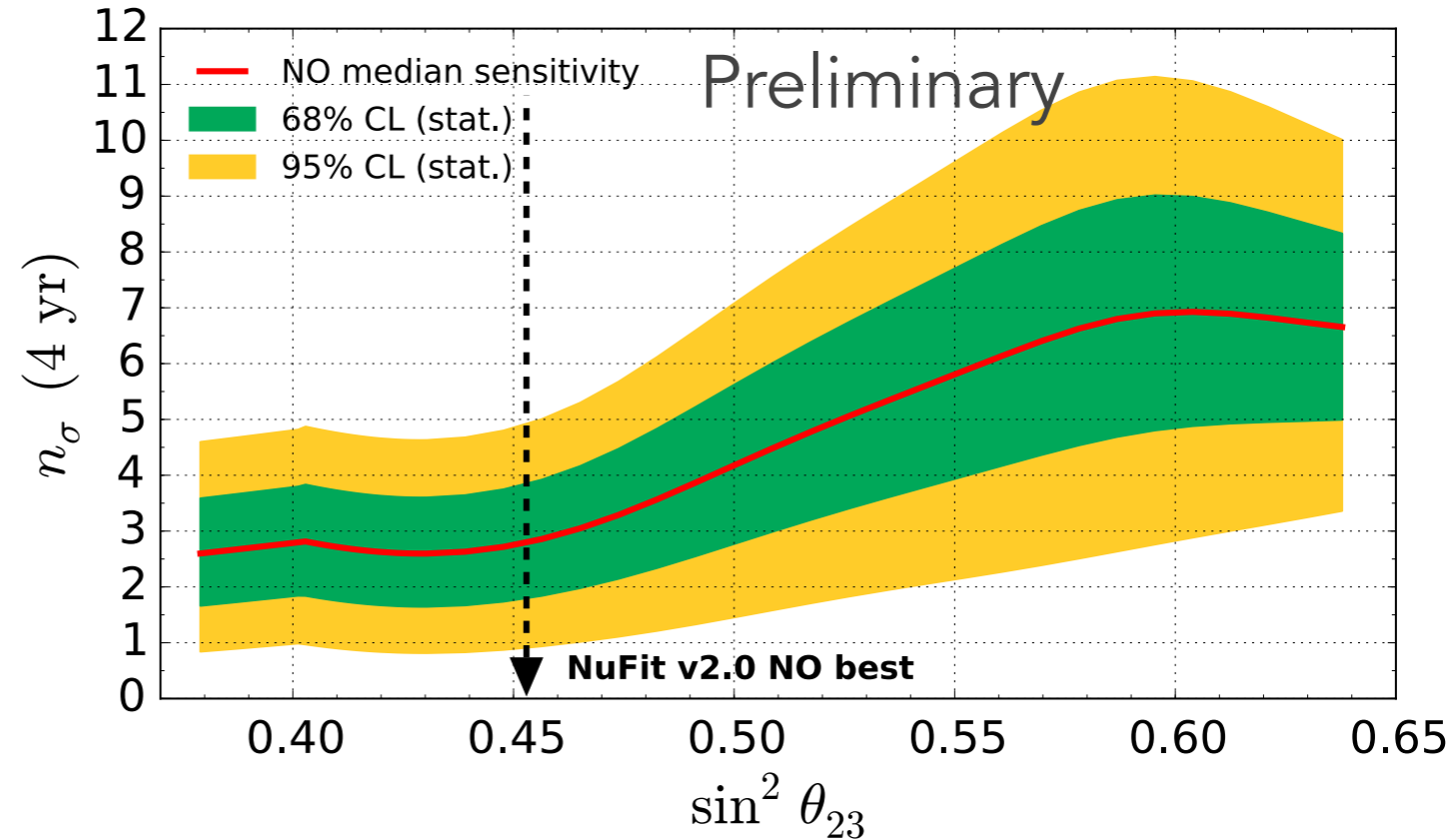
- Even before including detector effects and reconstruction smearing, the event rate histograms are quite similar

Neutrino Mass Hierarchy by Eye



PINGU NMO Sensitivities

- Brazilian flag sensitivities for the NMO analysis
- Sensitivity is mostly insensitive to the value of δ_{cp} , which is in contrast to other beam-based experiments



*DJK, Neutrino 2016

Statistics - A slight aside

- Use method outlined in Akhmedov, Razzaque, Smirnov - [arXiv:1205.7071](https://arxiv.org/abs/1205.7071)

$$S_{tot} = \sqrt{\sum_{ij} \frac{(N_{ij}^{IH} - N_{ij}^{NH})^2}{N_{ij}^{NH}}}$$

$i = \cos(\text{zenith})$
 $j = \text{energy}$
 $V^{eff} = \text{effective volume}$

$$N_{i,j}^{NH} = P(\nu_{\mu})_{i,j}^{NH} * \Phi(\nu_{\mu})_{i,j} * \sigma(\nu_{\mu})_j * V_{i,j}^{eff} + P^{NH}(\bar{\nu}_{\mu})_{i,j} * \Phi(\bar{\nu}_{\mu})_{i,j} * \sigma(\bar{\nu}_{\mu})_j * V_{i,j}^{eff}$$

- Essentially bin, sum, and subtract one hierarchy from the other. It works because:

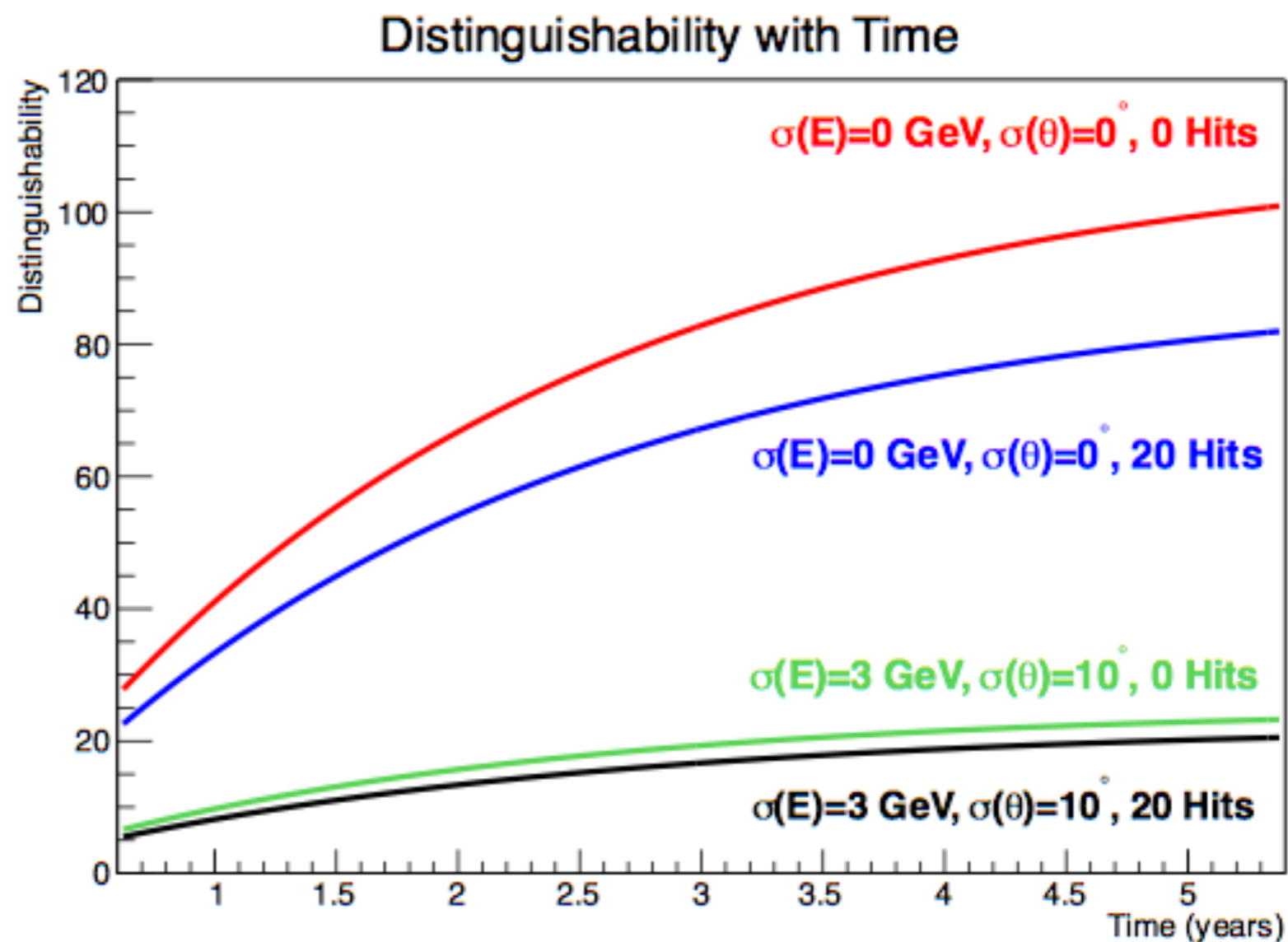
Probability : $P(\nu_{\mu})^{IH} + P(\bar{\nu}_{\mu})^{IH} \neq P(\nu_{\mu})^{NH} + P(\bar{\nu}_{\mu})^{NH}$

Flux : $\Phi(\nu_{\mu}) > \Phi(\bar{\nu}_{\mu})$

Cross – Section : $\sigma(\nu_{\mu}) > \sigma(\bar{\nu}_{\mu})$

How Long?

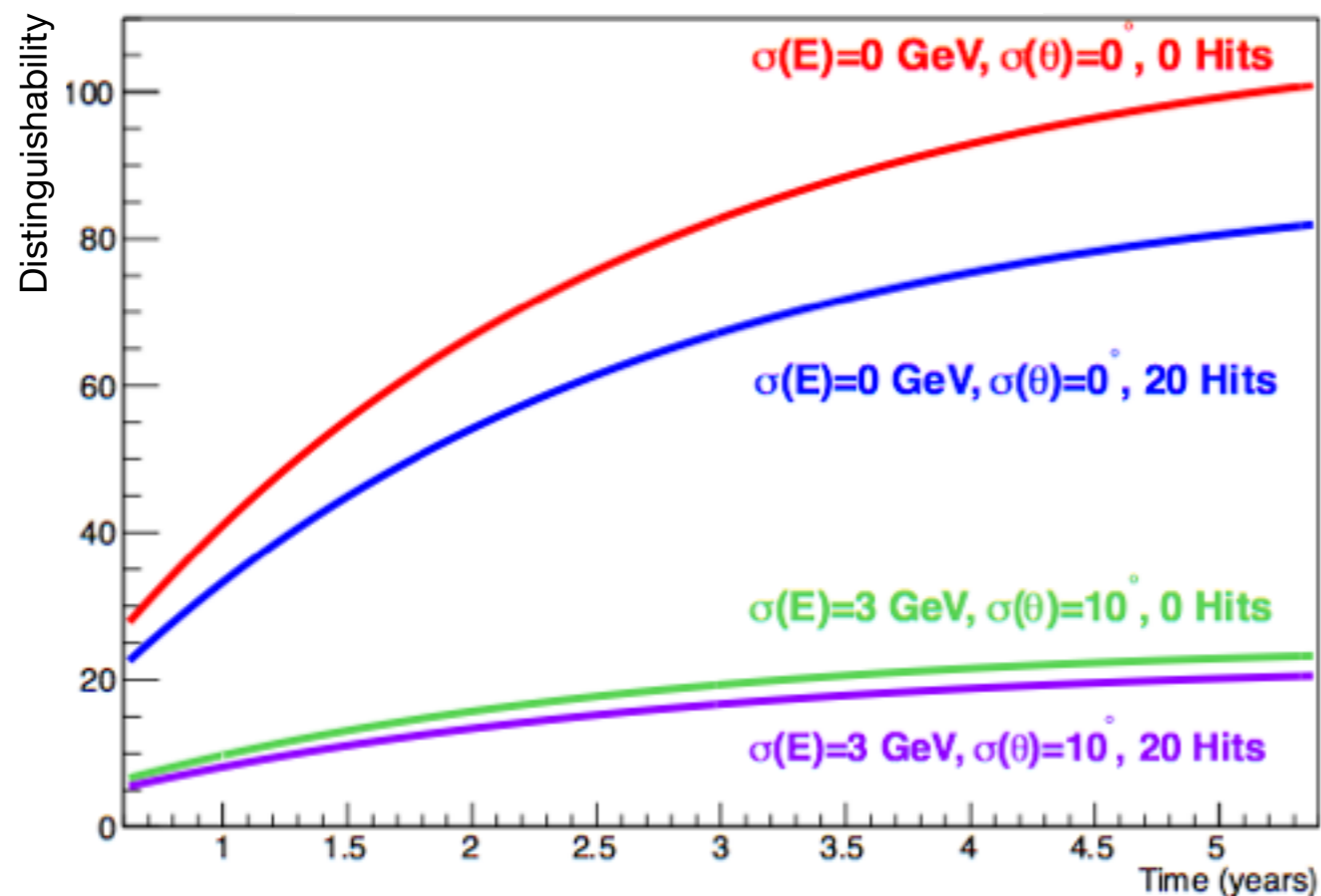
- The effect of all of the caveats needs to be determined
- Apply no conditions
- Apply 20 hit “reconstructability” cut
- Apply detector resolution
- Apply resolutions and 20 hit cut



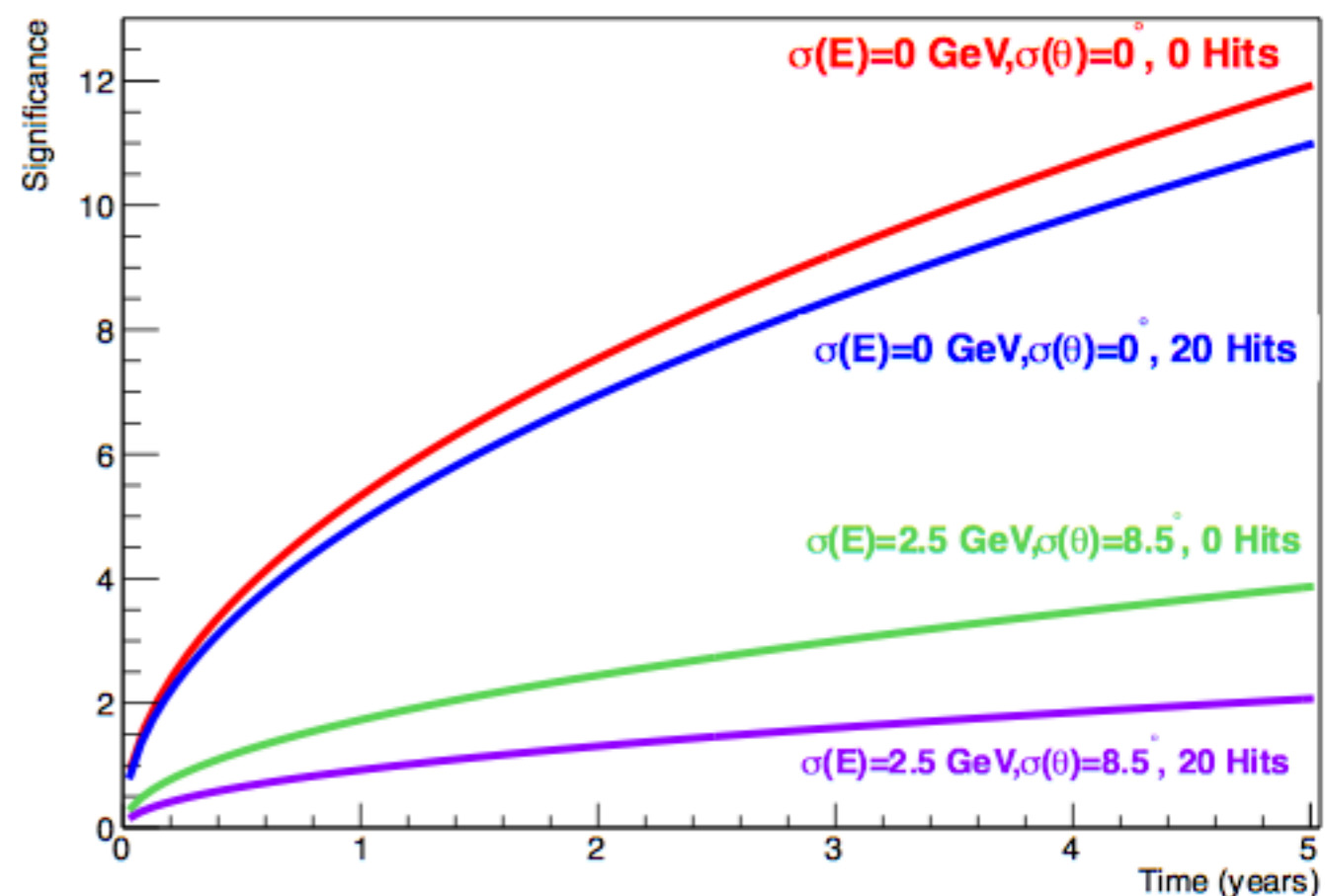
- Since this is a “discussion” ...
- The Akhmedov et. al. method may be optimistic
 - Uses a chi-squared like statistic, but the discrete aspect of the neutrino hierarchy fails regularity condition of Wilks’s theorem
- Alternative possibility is a MC method*
 - Create many sets smearing the reconstructed angle and/or energy for a specific hierarchy
 - Compare likelihood of smeared set(s) to a normal hierarchy template and an inverted hierarchy template

- The statistical power of PINGU makes systematics a critical factor sooner rather than later for hierarchy
 - PINGU specific - angular reco, energy reco, ice modeling...
 - Neutrino field at large - MC neutrino generators, cross-sections, atmospheric neutrino flux...

Akhmedov Method

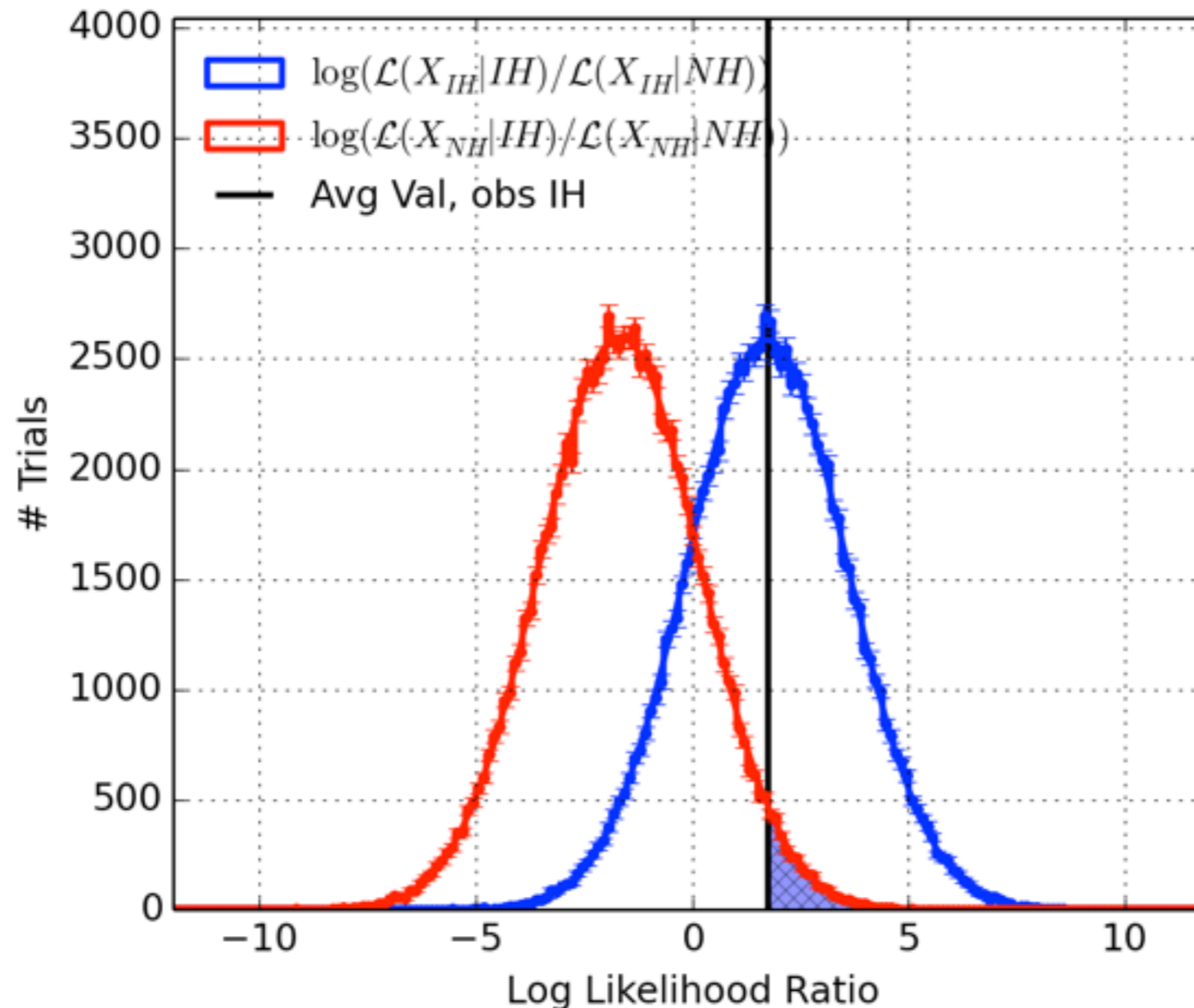


MC Method



Differentiation Between Inverted/Normal

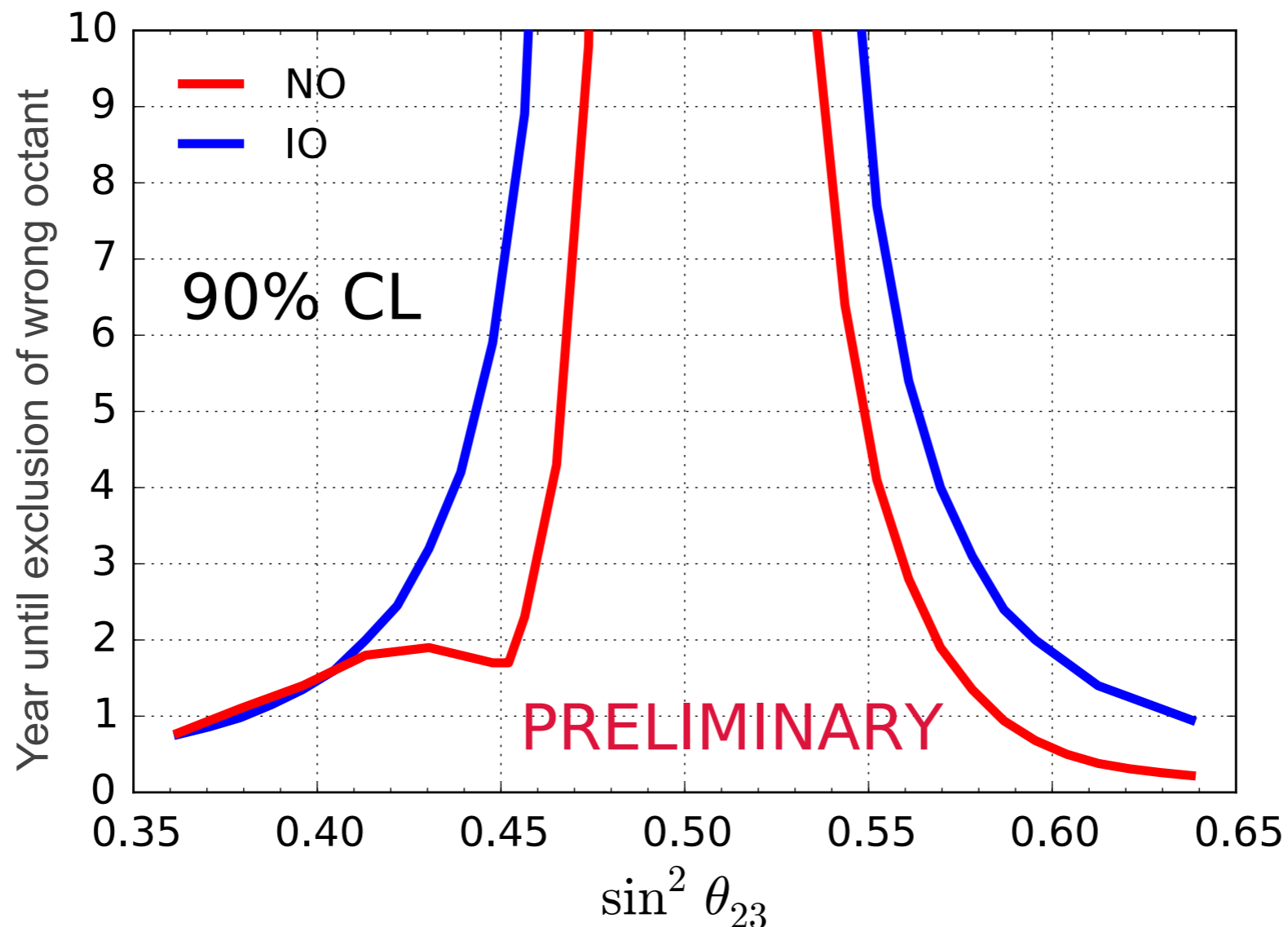
- Use a likelihood ratio with many simulated trials



Back to our regularly
scheduled physics

PINGU Octant

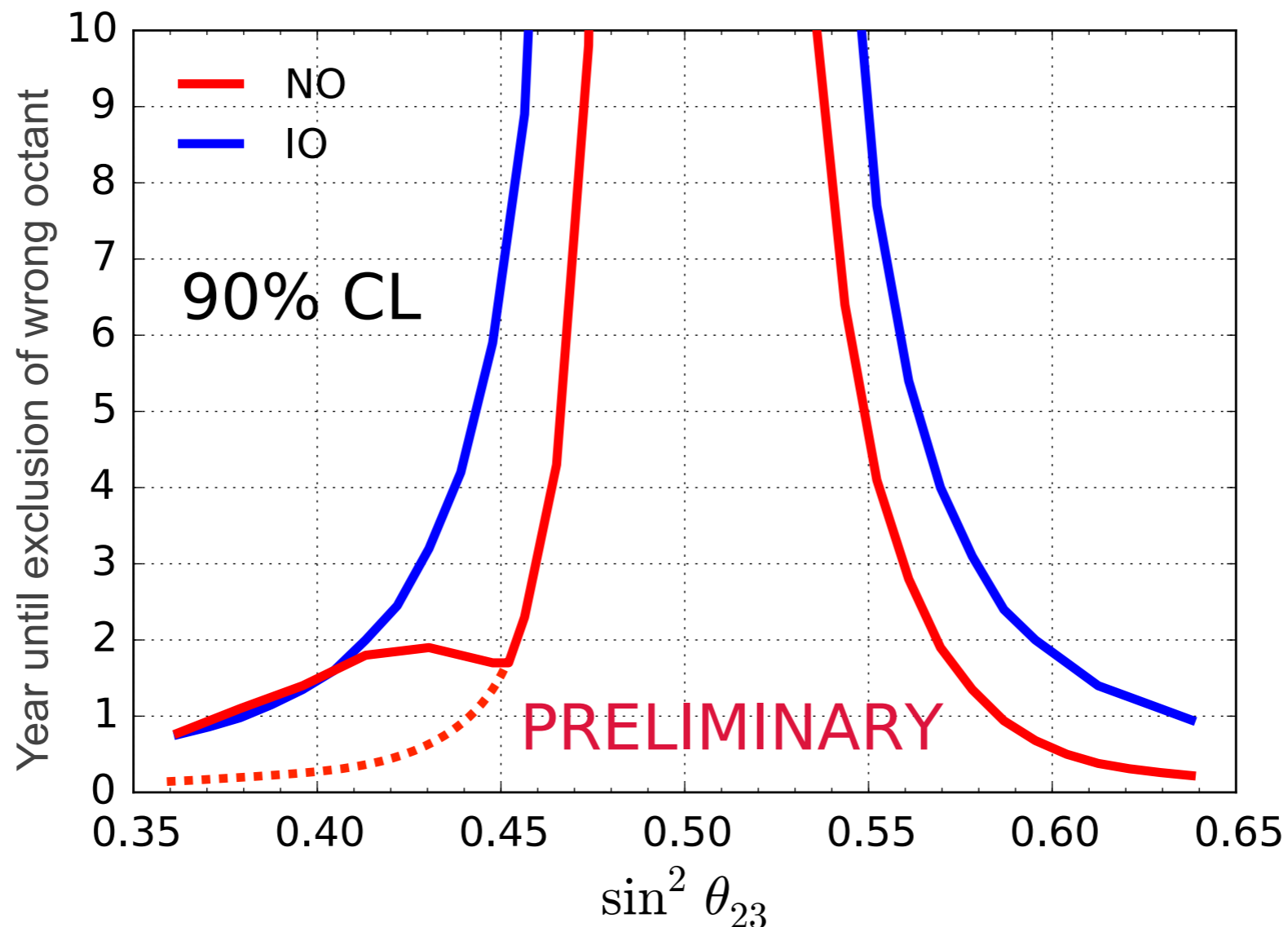
- Number of years to exclude the wrong θ_{23} octant
 - Compare goodness-of-fit between best-fit over the entire range and best-fit where the search is restricted to opposite octant
 - IO less sensitive because MSW effect is in the anti- ν channel



*DJK, Neutrino 2016

PINGU Octant

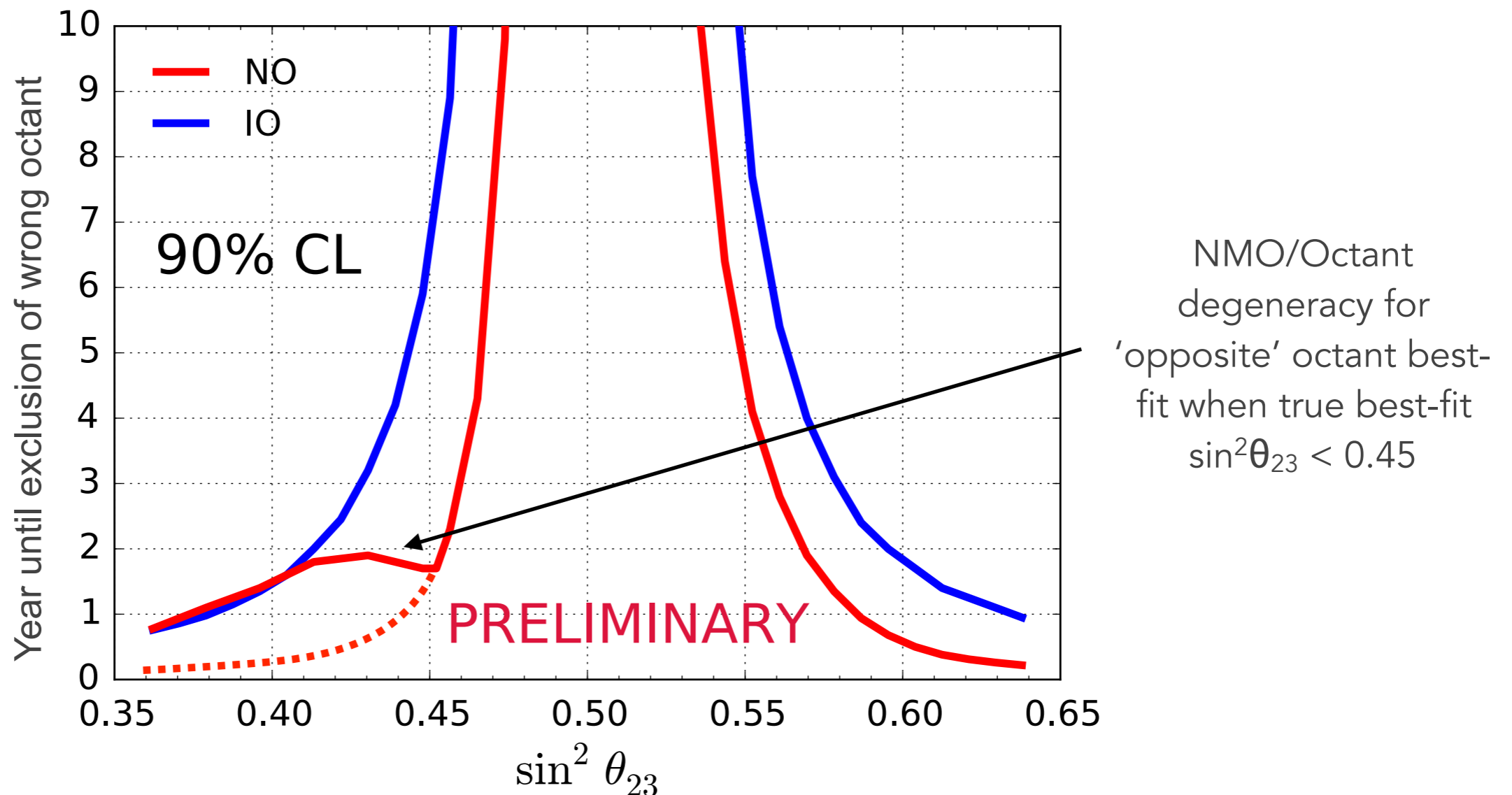
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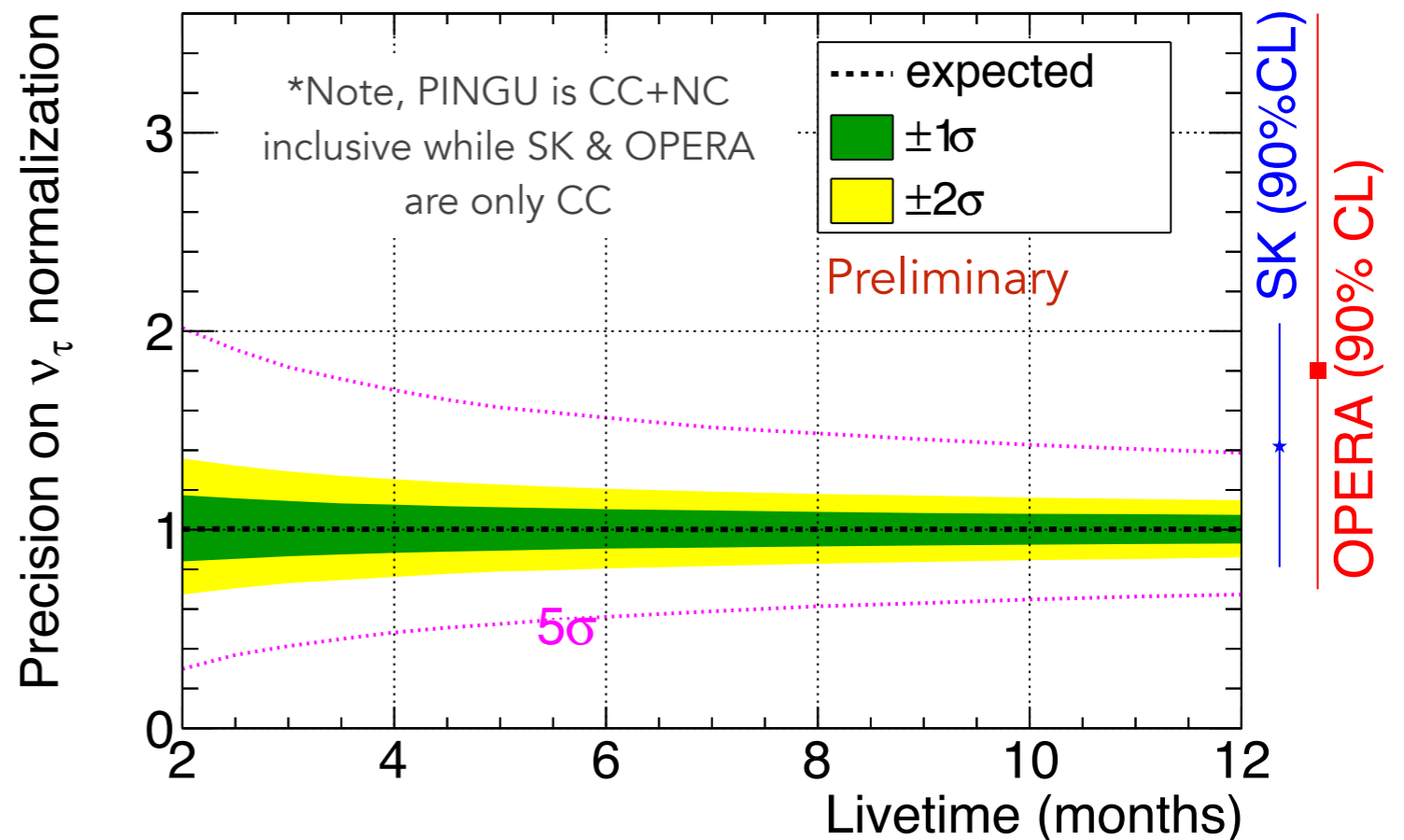
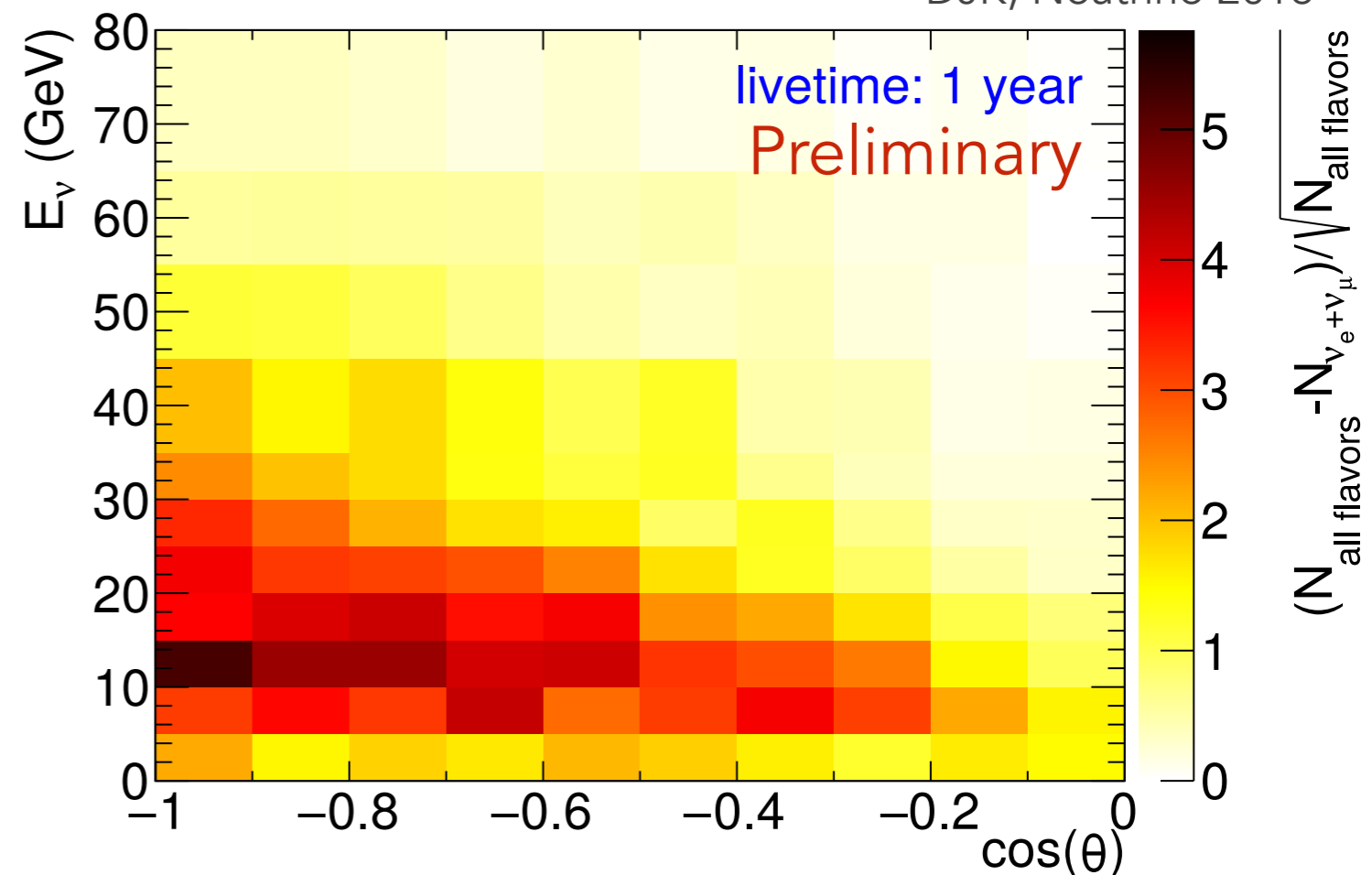
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*DJK, Neutrino 2016

ν_τ Appearance in PINGU

- Similarity to a DeepCore measurement
 - Direct measure of $|U_{\tau 3}|^2$
 - Energy and zenith angle excess in cascade channel
- PINGU analysis currently uses same initial Boosted Decision Tree as NMO, but secondary selection for 'cascades'



Conclusion

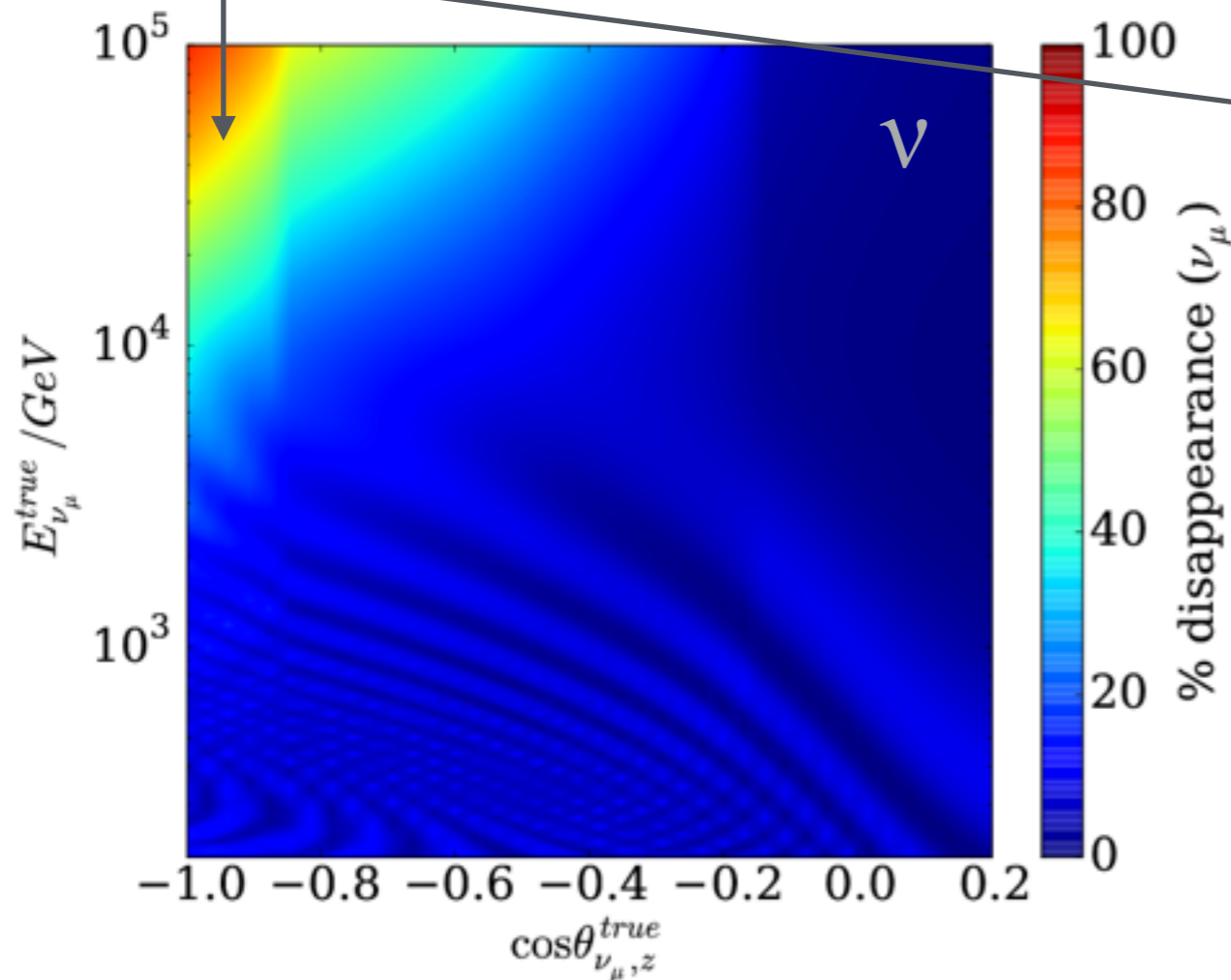
- IceCube/DeepCore is producing results using high-purity track-like events and working towards improved and new results with multiyear datasets which will allow for the most sensitive probes of ν_τ -appearance, which is a direct test of neutrino mixing unitarity
- Moving further into precision neutrino physics, e.g. δ_{cp} , requires extended precision measurements of PMNS unitarity
- PINGU can resolve the Neutrino Mass Ordering at 3σ in 3-4 years and greatly enhances the reach of IceCube/DeepCore physics portfolio (ν_τ appearance, octant, precision θ_{23} , ...)

Backup

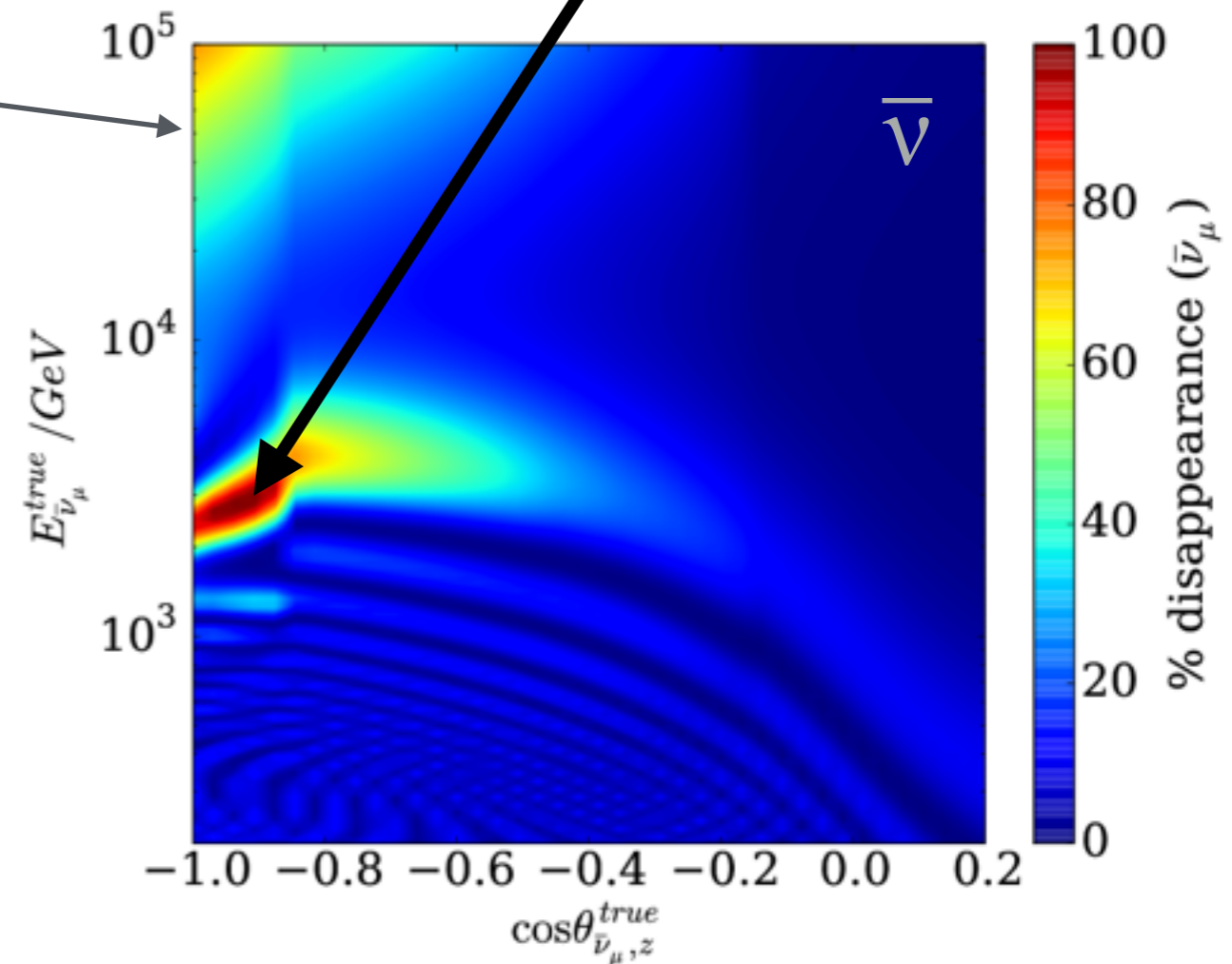
Sterile Neutrinos

- Sterile neutrino signatures extend in energy beyond the conventional reactor and accelerator searches
- At $\Delta m^2 \simeq 1 \text{ eV}^2$ there is a matter induced resonance at $\bar{\nu}$ energies of $\mathcal{O}(1) \text{ TeV}$ for 3+N models, or ν for N+3

Earth induced opacity

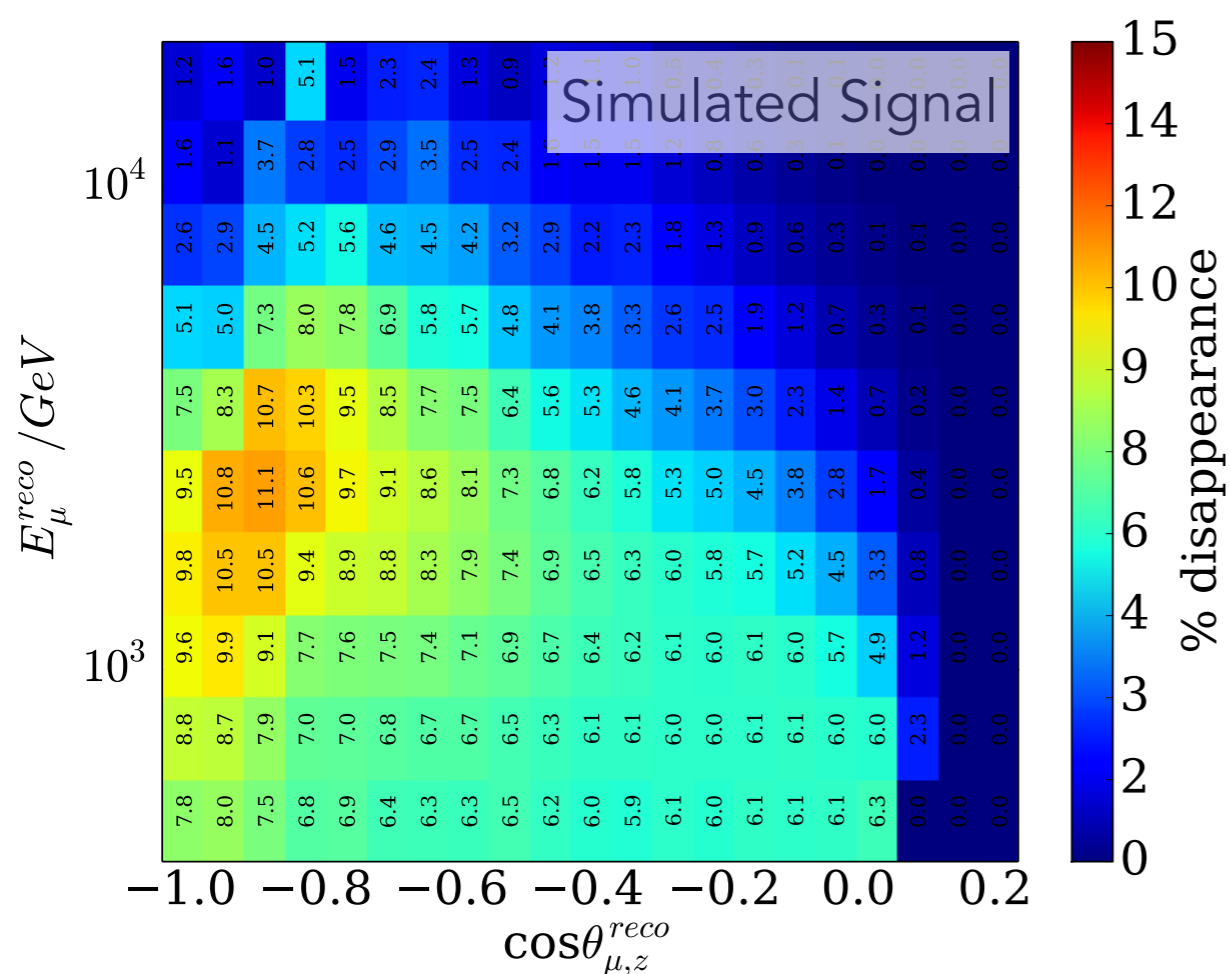


3+1 sterile including MSW



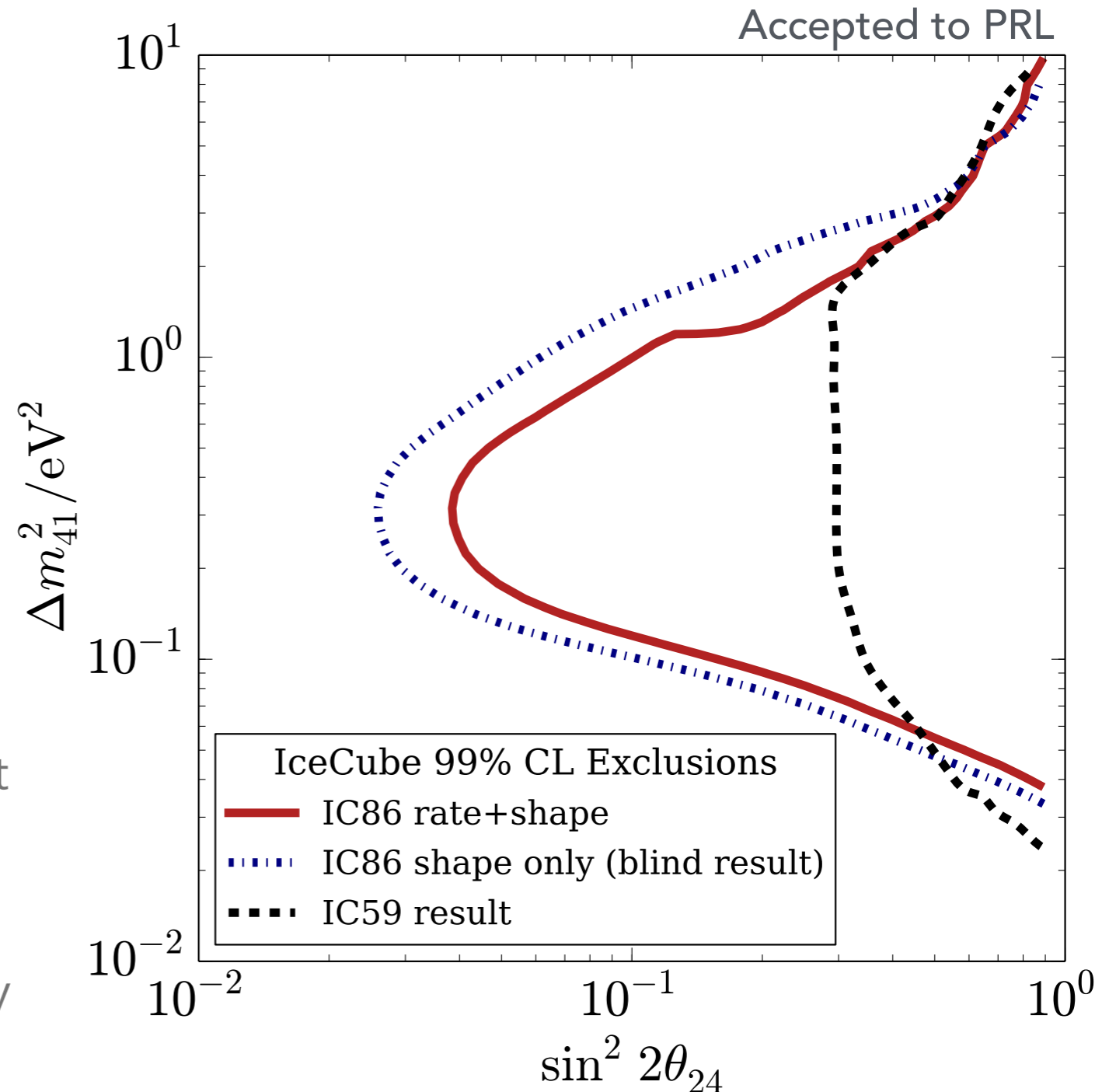
Sterile Search Approach

- Two separate diffuse ν_μ event selections of 1-year livetime (IC59 and IC86-1) were used to search for a sterile signal [deployment map in backup]
- The pronounced sterile feature for $\bar{\nu}$ is smeared out by:
 - Reconstruction uncertainty
 - The ν -induced muons are uncontained
 - Signal is combination of $\nu + \bar{\nu}$

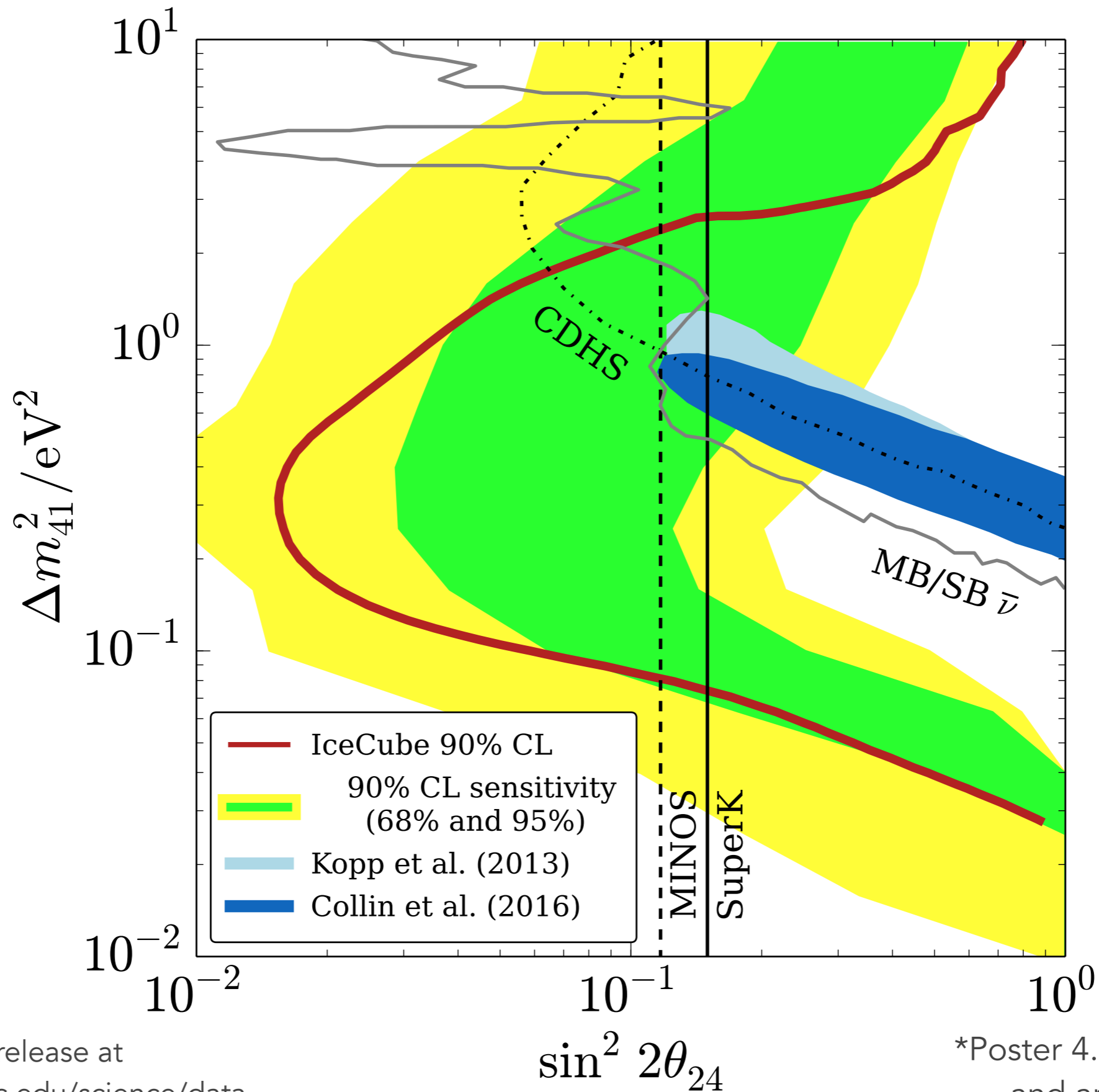


No Sterile Signature

- Primary result is IC86 “rate+shape”, complemented by IC59 and IC86 “shape only”
- “rate+shape” is a *posteriori* inclusion of a 40% prior on the atm. ν flux normalization
 - From rapid oscillations, results at $\Delta m^2 \gtrsim 5 \text{ eV}^2$ with an unphysical flux normalization are highly degenerate with a no-sterile result
 - The loose prior constrains the flux normalization to be physical and breaks degeneracy between many high Δm^2 results



IceCube Sterile ν Result



Public IC86-1 data release at
<https://icecube.wisc.edu/science/data>

*Poster 4.088 by C. Argüelles
and arXiv:1605.01990