

NBIA Summer School on Astrophysical Plasmas – From Planets to Galaxies



The Niels Bohr
International Academy

August 28 – September 01,
2017, Copenhagen

- Name: MYWISH ANAND
- Currently: M.Sc. Computational Physics (1st year)
- Institution: Central University of Punjab, India
- Contact Info: Mywish@physics.org



- Research Motivations: Computational Astrophysics, Galaxy Evolution, Magnetic Reconnection
- Most relevant courses: Classical Electrodynamics, Relativity
- Programming Languages: C and C++
- Software & Plotting Packages: MATLAB, Scilab, GNU PLOT
- Programming Experience: around 1 year
- Research Experience: around 10 months

My Curiosity and Learning Goals from the School ...



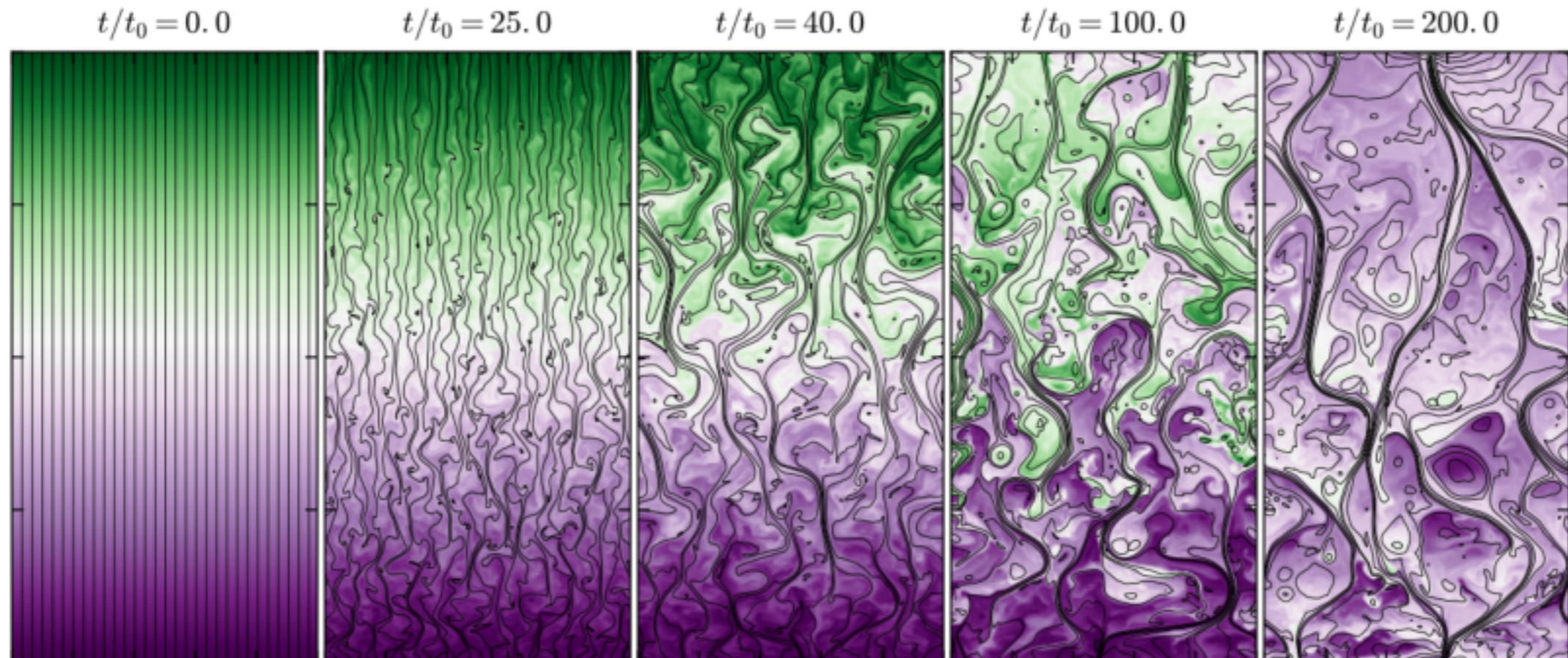
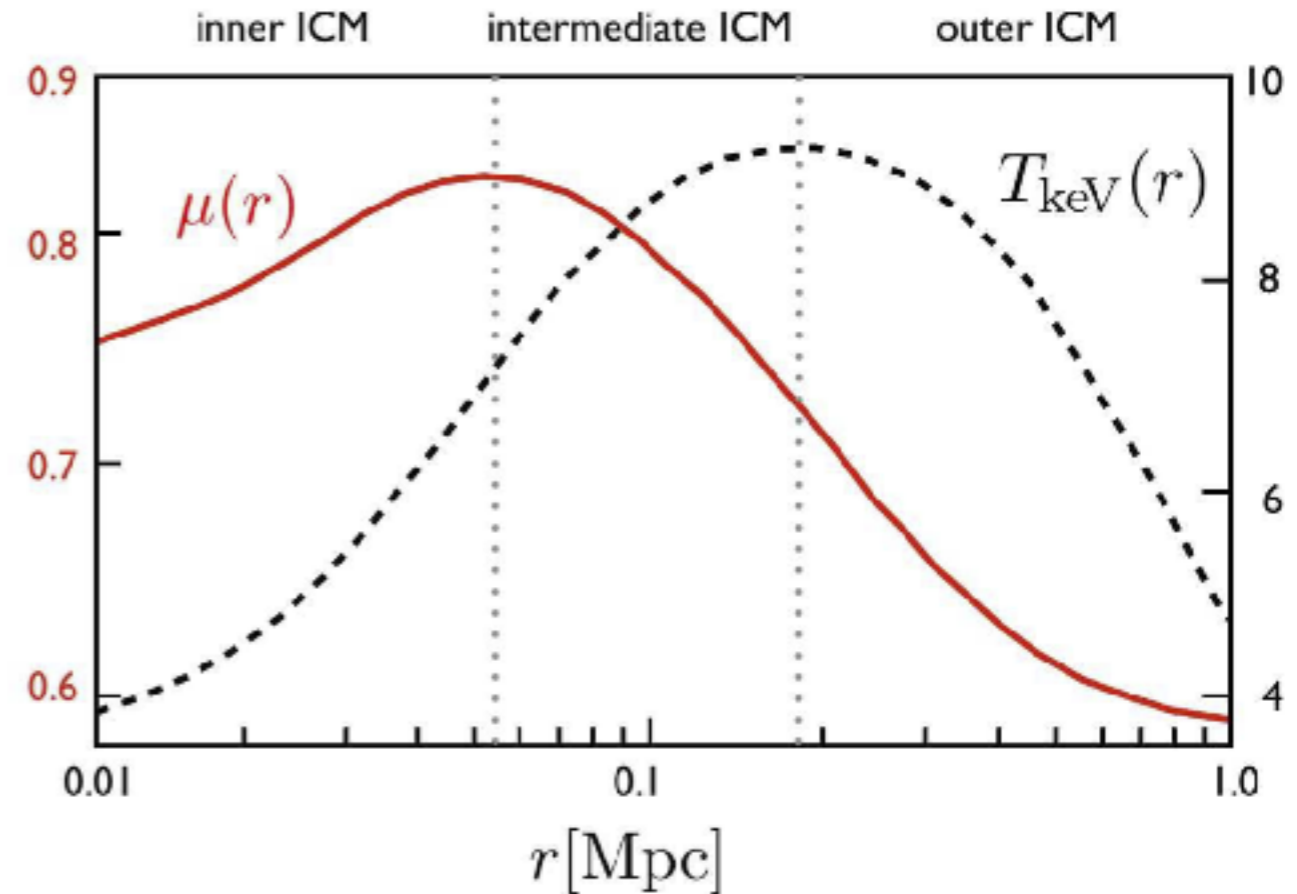
Plasma instabilities in the intracluster medium

Name: Thomas Berlok

Advisor: Prof. Martin Pessah

Interested in Astrophysical fluid dynamics

Currently working on understanding the physics of the low density gas between galaxies in galaxy clusters.



Asmita Bhandare

(Max Planck Institute for Astronomy)

Research Interests: Star formation, formation and evolution of circumstellar disks.



bhandare@mpia.de

- ◆ B.Sc. in Physics @ University of Pune, India in 2012.
-

- ◆ M.Sc. in Astrophysics @ University of Bonn, Germany in 2015.

Thesis: Effects of inclined star-disk encounter on protoplanetary disk size.

Supervisor: Susanne Pfalzner, MPIfR, Bonn.

- ◆ PhD in Astronomy @ MPIA, Heidelberg, Germany since July 2016.

Thesis: Numerical simulations of low mass star formation.

Supervisors: Thomas Henning (MPIA), Rolf Kuiper (University of Tuebingen) and Christian Fendt (MPIA).

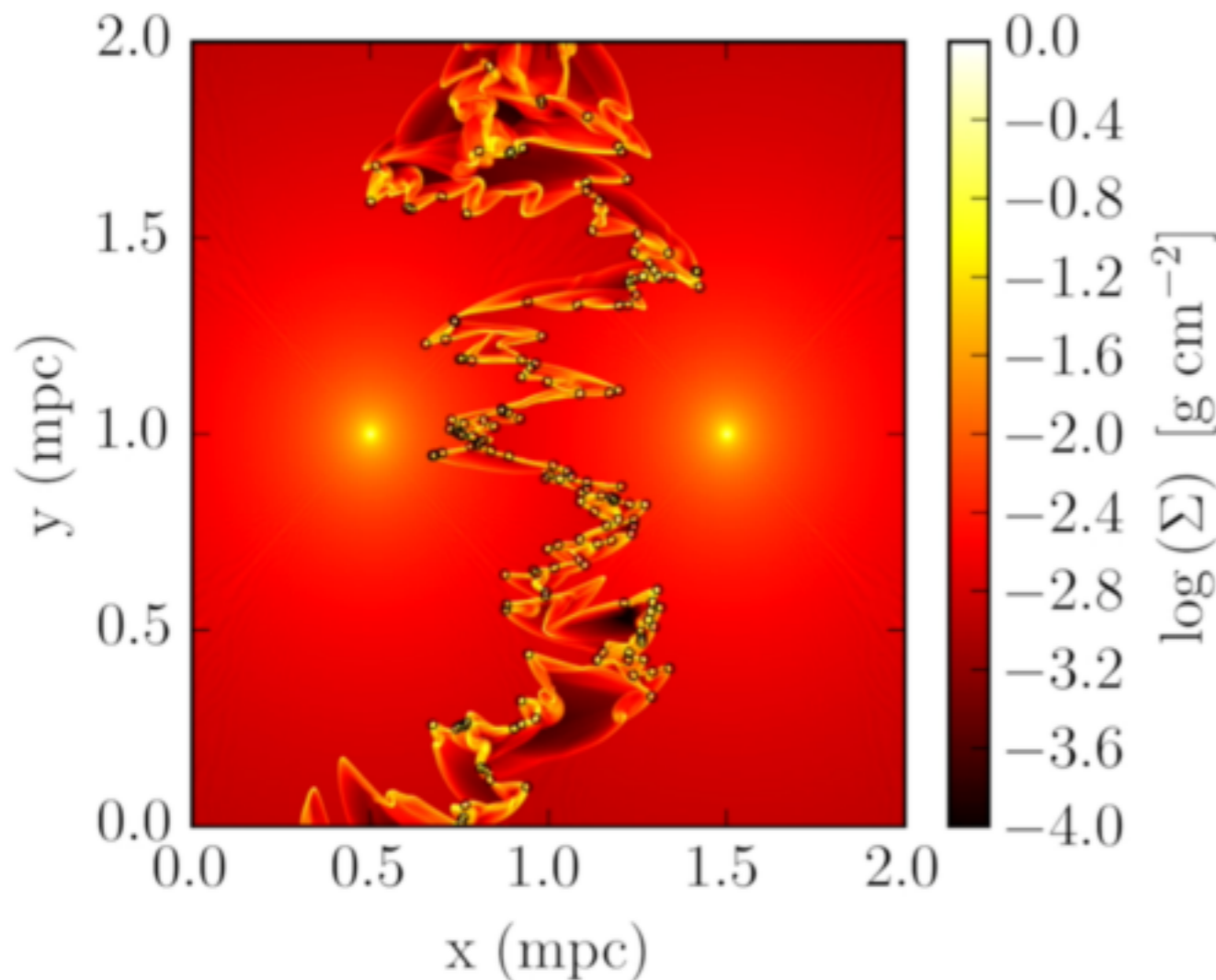
Diego Calderón



PhD student

Pontificia Universidad Católica de Chile

Adviser: Professor Jorge Cuadra



Research interests

Colliding wind binaries

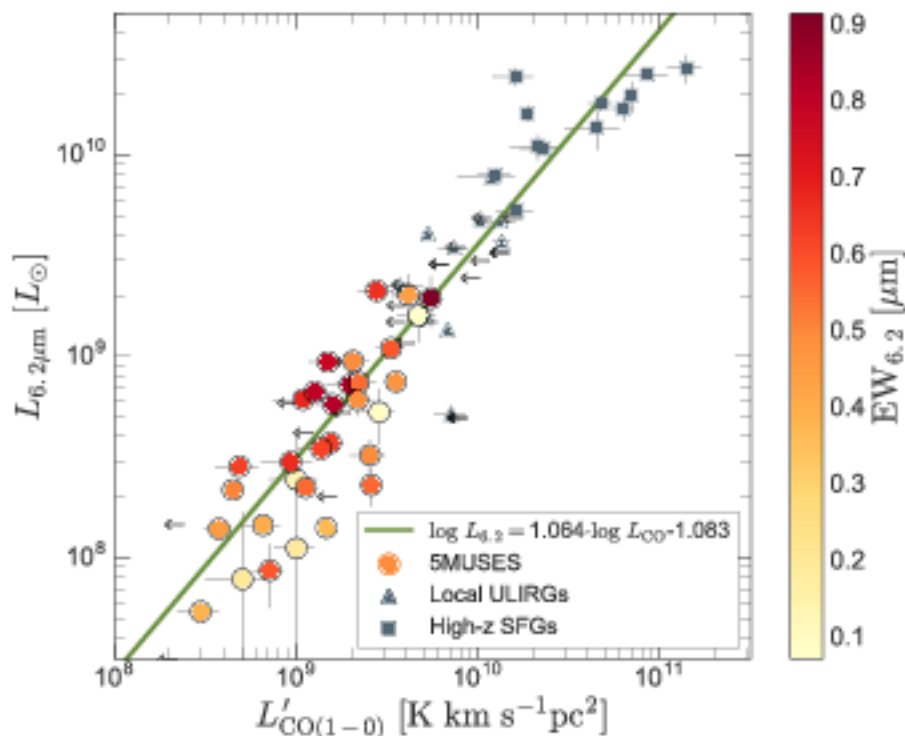
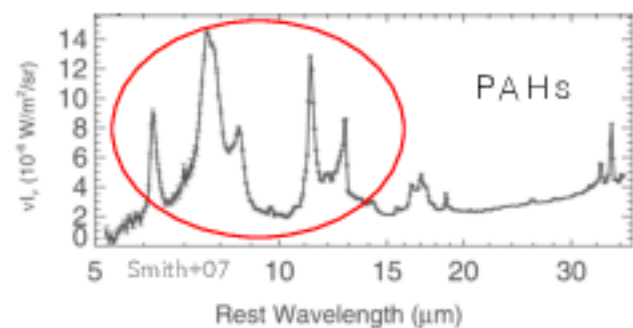
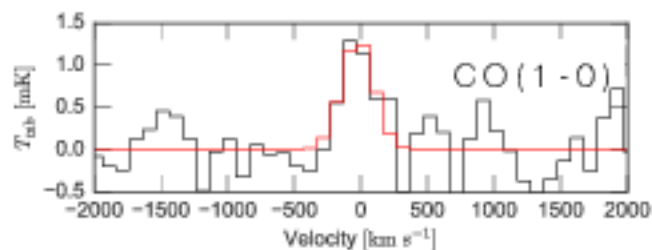
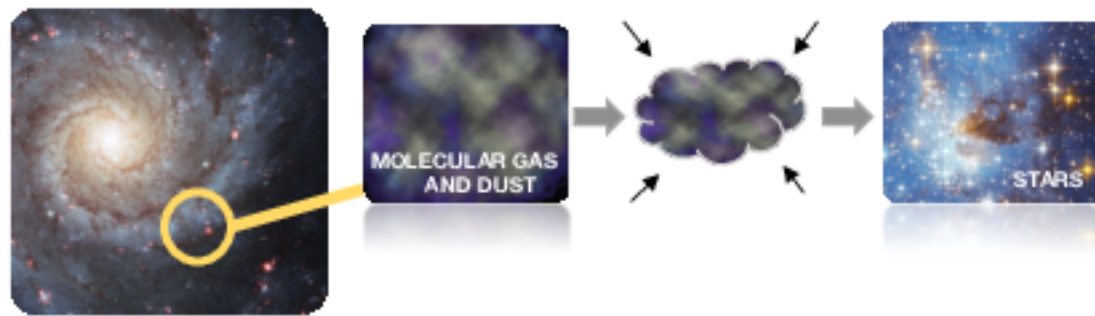
Galactic Centre

Numerical hydrodynamics

Isabella Chi Gieseler Cortzen

Ph.D. student at Dark Cosmology Centre
University of Copenhagen

- Examine star formation in galaxies across cosmic time
- Linking molecular gas mass (CO lines) and warm dust emission (PAHs)
- 5MUSES galaxies with CO(1-0) observations:
Bridge the gap between local and high-redshift galaxies



Personal Info Slides

Davide Decataldo

Scuola Normale Superiore of Pisa
(Italy)

davide.decataldo@sns.it





Scuola Normale Superiore of Pisa



Master Degree in **Astrophysics**

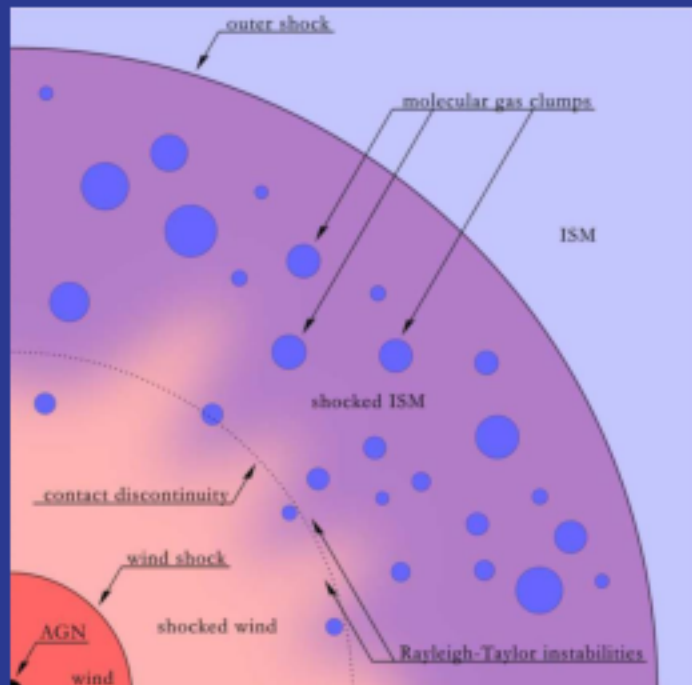
- ▶ Fluid Dynamics
- ▶ Stellar Physics
- ▶ Cosmology

Master Thesis

Evolution of Molecular Clumps in Quasar Outflows

Supervisor: Prof. Andrea Ferrara

Evolution of Molecular Clumps in Quasar Outflows



Experiences:

- 2015 **Cosmology Summer School**
Jan Kochanowski University of Kielce (Poland)
- 2015 **Recomputing the early universe**
DAVID X International Workshop
Scuola Normale Superiore of Pisa (Italy)
- 2017 **Galaxy Formation Summer School**
Anargyrios and Korgialenios School of Spetses (Greece)

PhD at Scuola Normale Superiore

Current work:

Evaporation of molecular clouds

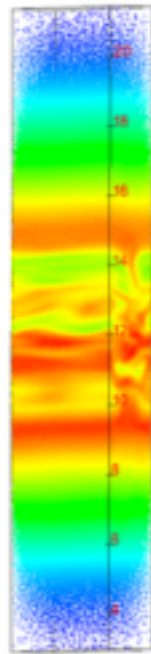
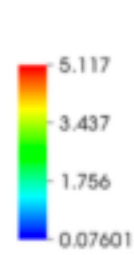
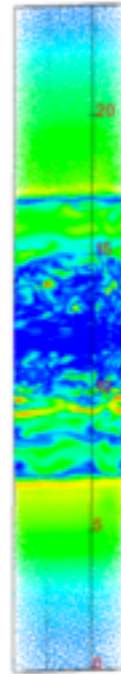
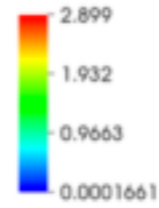
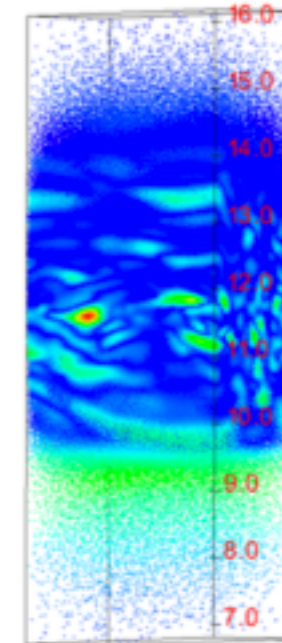
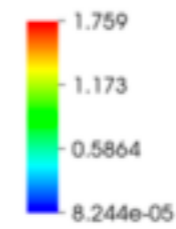
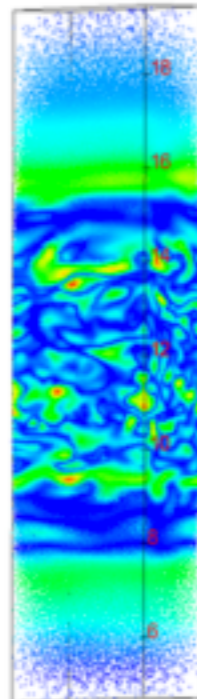
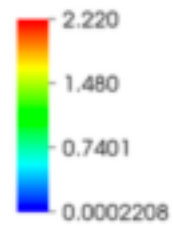
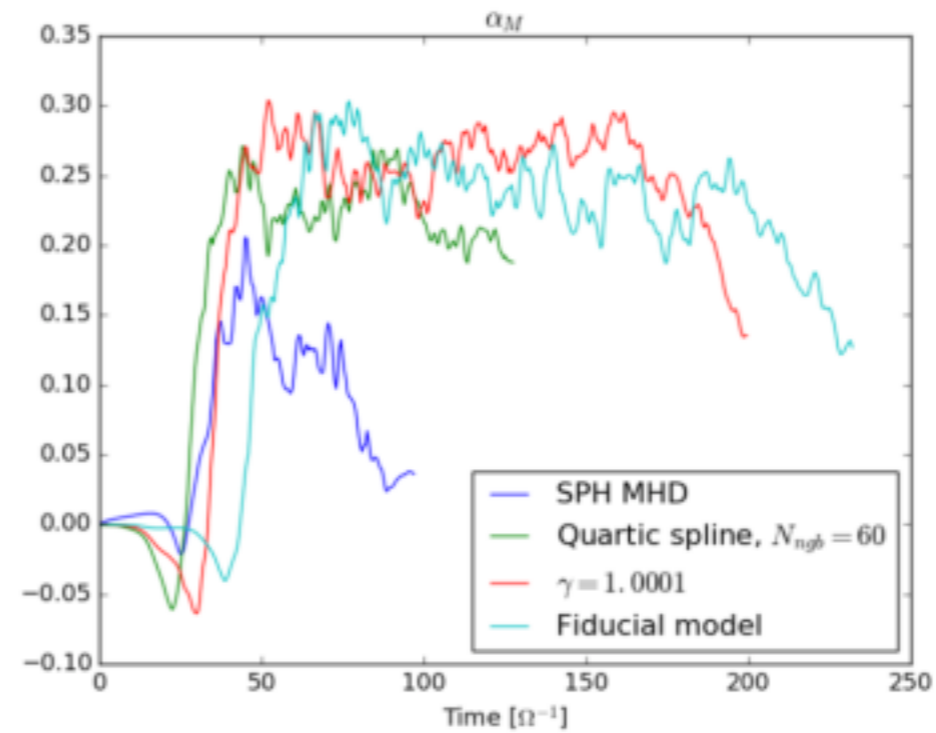
Evaporation of molecular clouds



Semi-analytical model



Simulations with RAMSES

(a) SPH MHD, $100\Omega^{-1}$ (b) Quartic spline, $128\Omega^{-1}$ (c) $\gamma = 1.0001$, $200\Omega^{-1}$ (d) Fiducial model, $200\Omega^{-1}$ (e) Time evolution of α_M

NBIA Summer School on Computational Astrophysics



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International Academy

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2017, Copenhagen

- Name: Kristian Ehlert
- Currently: M.Sc. (2nd yr)
- Advisor: Prof. Christoph Pfrommer
- Institution: University of Heidelberg
- Contact Info: kristian.ehlert@gmail.com

- Research Interests: jets, cosmic rays, clusters, cosmology
- Most relevant courses: Astro. Fluid Dyn., Gal. Extragal. Astro., Fundamentals of Simulational Methods
- Programming Languages: C, python
- Visualization Packages: Paraview
- Programming Experience: 2+ years

I would like to learn more about the theory and numerics of astrophysical plasmas.

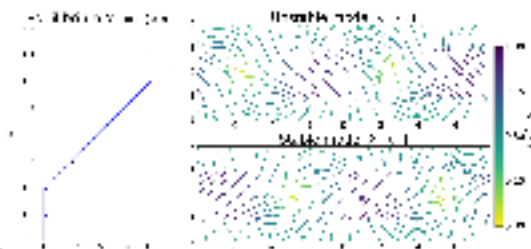
Adrian Fraser, University of Wisconsin-Madison

Current research: Effect of stable modes on shear-flow turbulence

Hydrodynamic shear flow:

- Stable modes excited nonlinearly, reach significant amplitude in saturation
- Strongly modifies flow and transport

⇒ Important in turbulence models

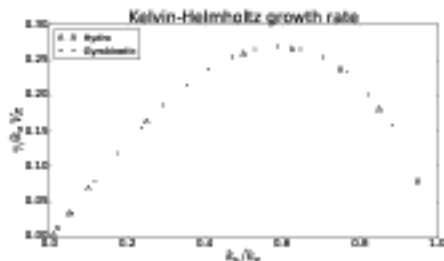


Fraser et al. arXiv:1610.06142

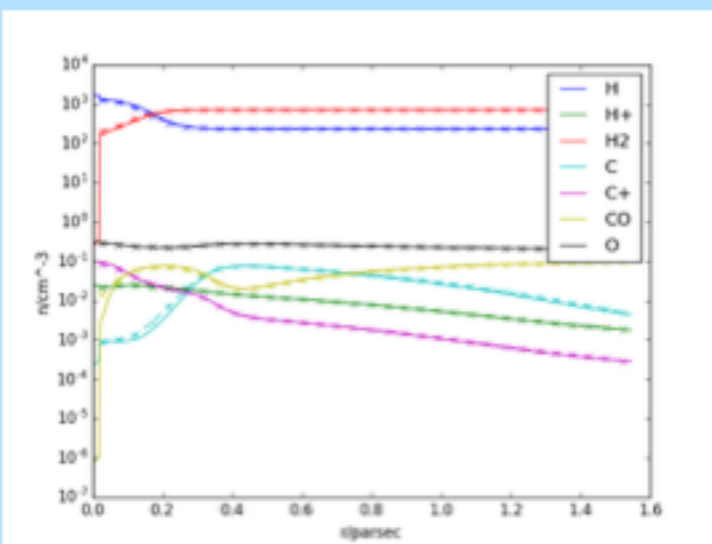
Gyrokinetic shear flow:

- Benchmark to hydro cases
- Numerical tools yield eigenmode characterization and saturation amplitude

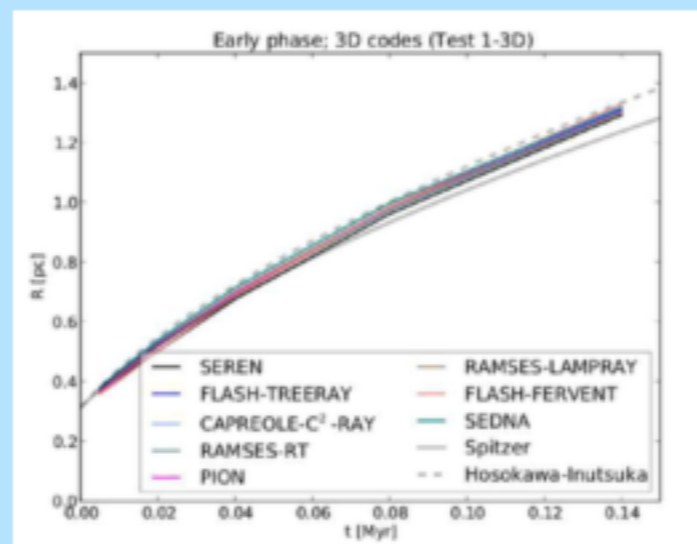
→ Stable modes excited nonlinearly, reach significant amplitude



Radiative transfer in RAMSES simulations

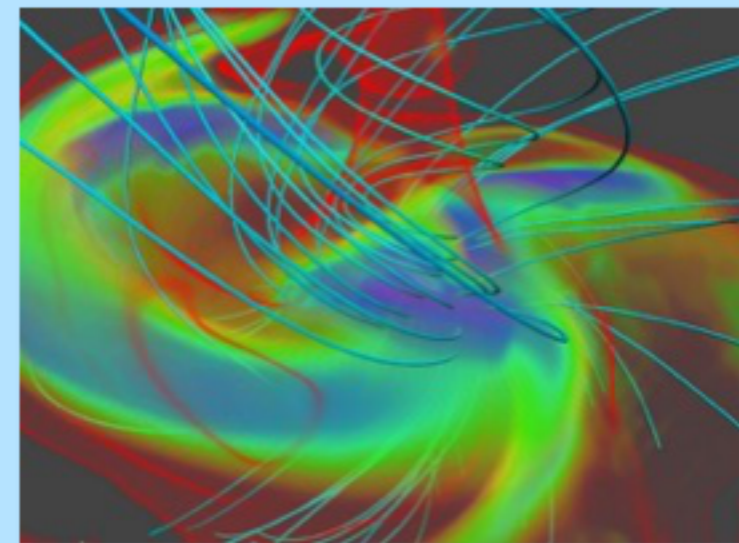


photodissociation region benchmark

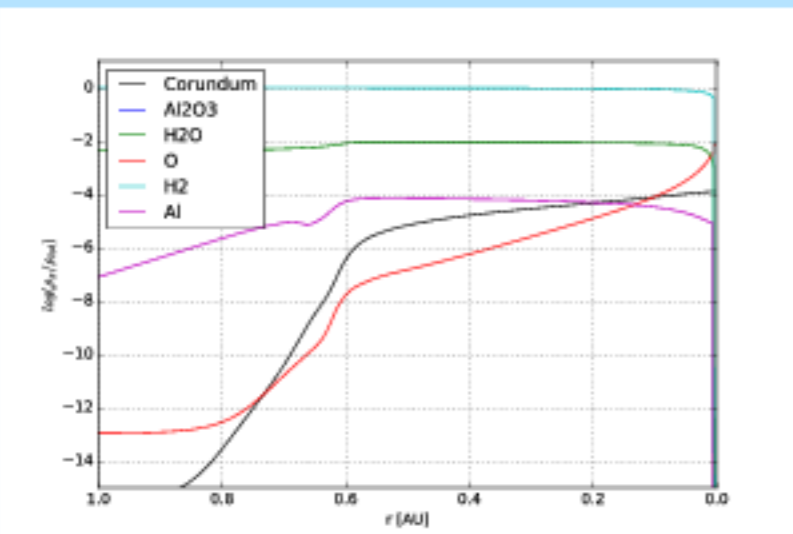


D-type expansion of ionization front

To be used in
Zoom-in simulations of star formation

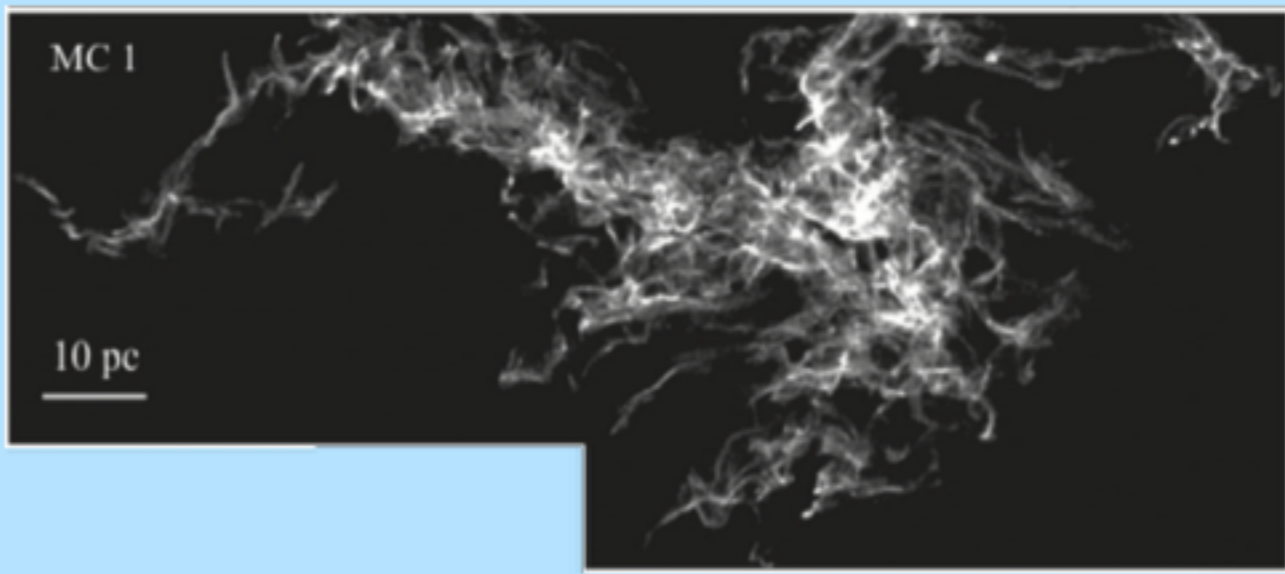


Formation of CAIs



Chemical evolution of corundum in in-falling
gas during protostellar collapse phase

Simulations of GMC evolution





Carlos Gómez-Guijarro

PhD Student (2nd year) - Dark Cosmology Centre, Niels Bohr Institute

cguijarro@dark-cosmology.dk

High-z starburst galaxies

Evolutionary connection with massive elliptical galaxies

Development of Massive Elliptical Galaxies





Institute for Research in
Fundamental Sciences



IC1613 dwarf

- ❖ First name: Seyed Azim (Azim)
- ❖ Last name: Hashemi
- ❖ Birthday: 1/4/1991
- ❖ Master Student at Sharif University of Technology, Tehran, IRAN.
- ❖ Research Assistant at IPM, Tehran, IRAN
- ❖ Major: Physics and Astrophysics.
- ❖ Research Interest : Astrophysics and Cosmology (Stellar evolution, Star Formation History(SFH), Dust Production , Exoplanet, Dwarf Galaxies).
- ❖ Email: hashemi.seyedazim@gmail.com

Francisco (Paco) Holguin



Graduate student in the Department of Astronomy
University of Michigan-Ann Arbor

Current research: Simulations of Cosmic Ray Driven Galactic Winds

Marija Jankovic

- PhD student at Imperial College London
- Modelling of protoplanetary disks:
 - self-consistent pseudo-viscous accretion of disks with calculations of pseudo-viscosity from MHD conditions in the disk
 - Can close-in Super Earths form in situ?
 - simulated observations of massive disks
 - What can we currently learn from marginally resolved disks?
- BSc, MSc in physics at University of Belgrade, Serbia



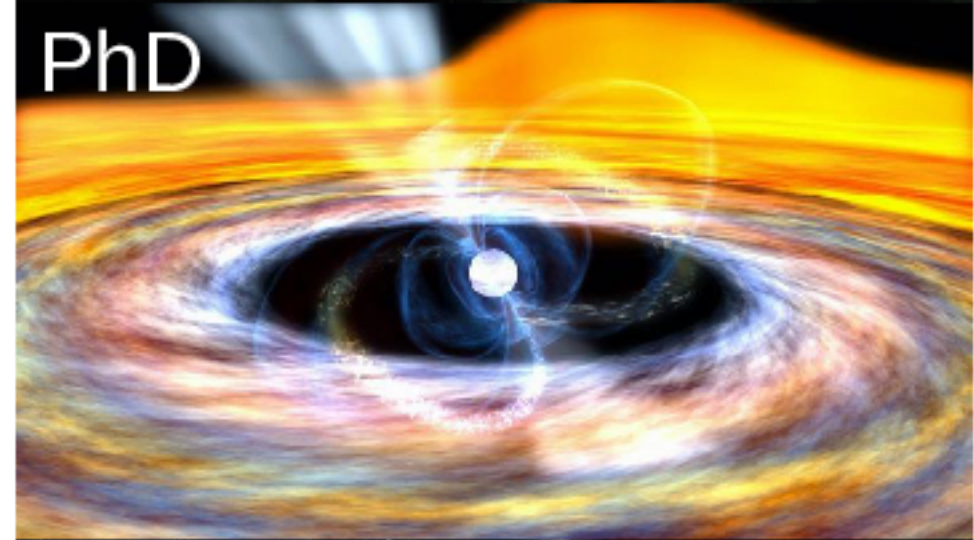
Jeff Jennings

Ludwig Maximilians
University Munich

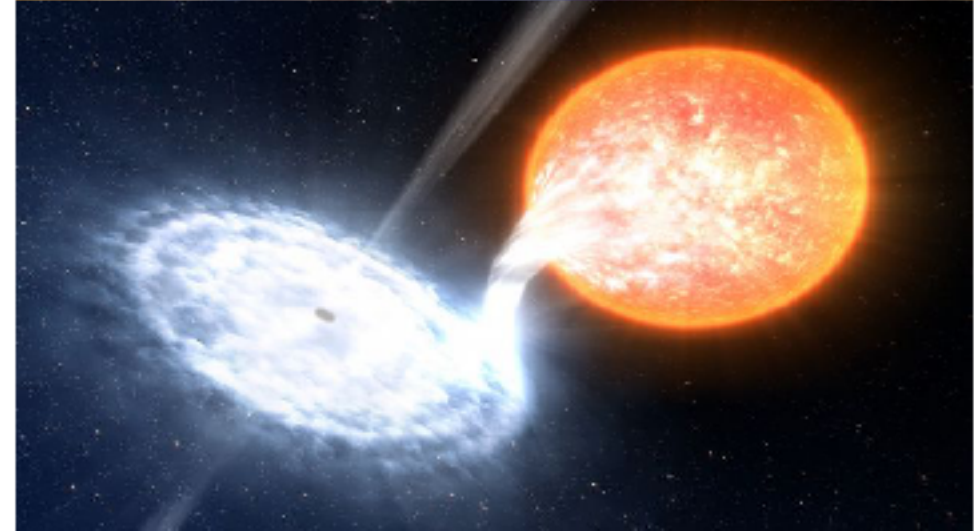
Masters



PhD



Images courtesy:
University of Copenhagen/Lars Buchhave
NASA/Goddard Space Flight Center/Dana Berry
ESO/L. Calçada
NASA/JPL-Caltech



Star formation through dynamical accretion

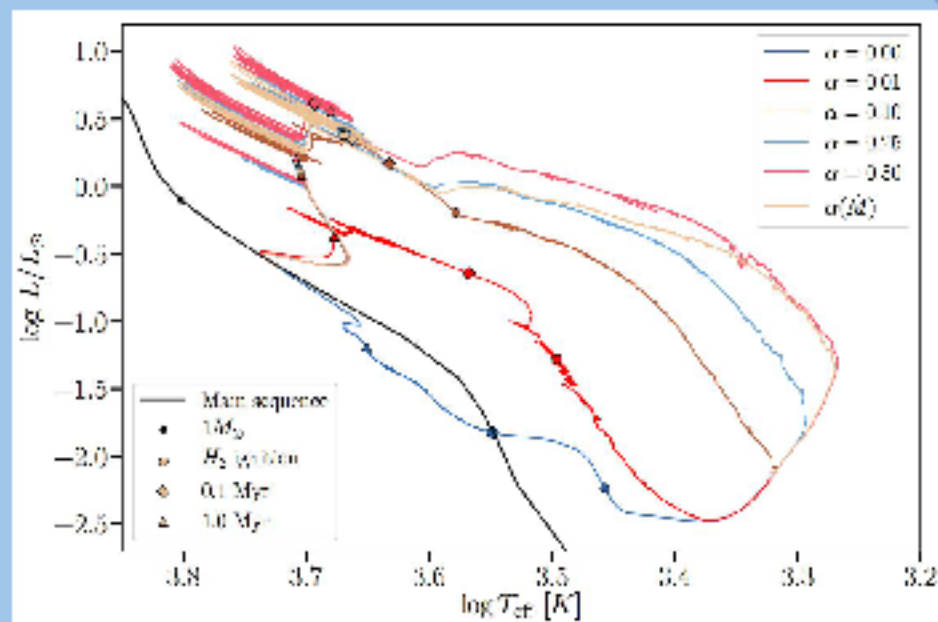
Master project
Sigurd Sigersen Jensen
University of Copenhagen

Methods:

- RAMSES accretion profiles
- MESA 1D stellar evolution code

Goals:

- Improve understanding of accreting models of star formation
- Improve stellar isochrones - solve luminosity spread problem?
- Implement MESA into RAMSES -> Irradiation feedback to protoplanetary disk etc



Star formation through dynamical accretion

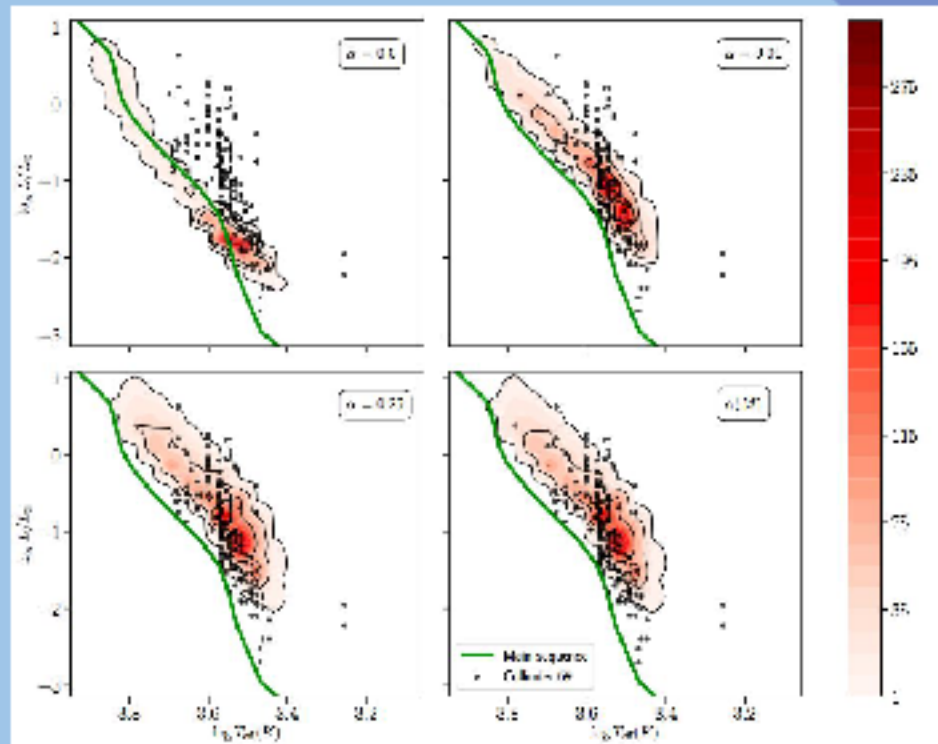
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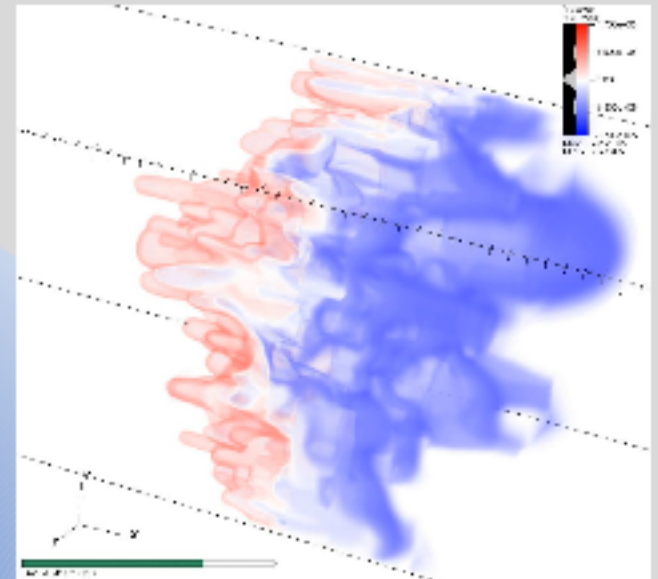
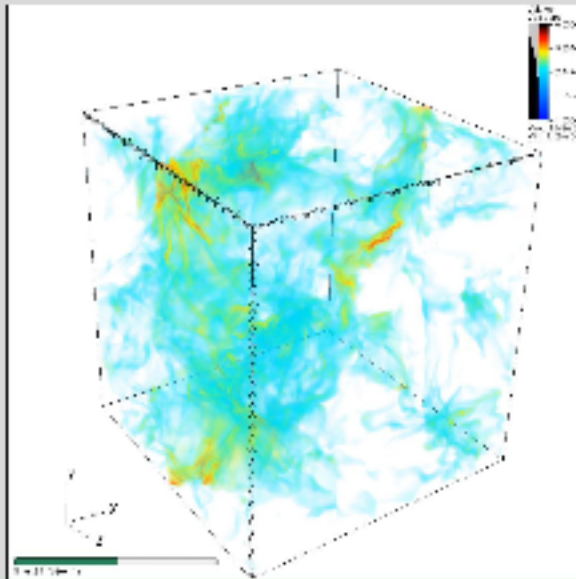
Goals:

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- Implement MESA into RAMSES -> Irradiation feedback to protoplanetary disk etc



Resolution requirements for modelling molecular gas formation

gas formation



Prabesh Raj Joshi

I. Physikalisches Institut,

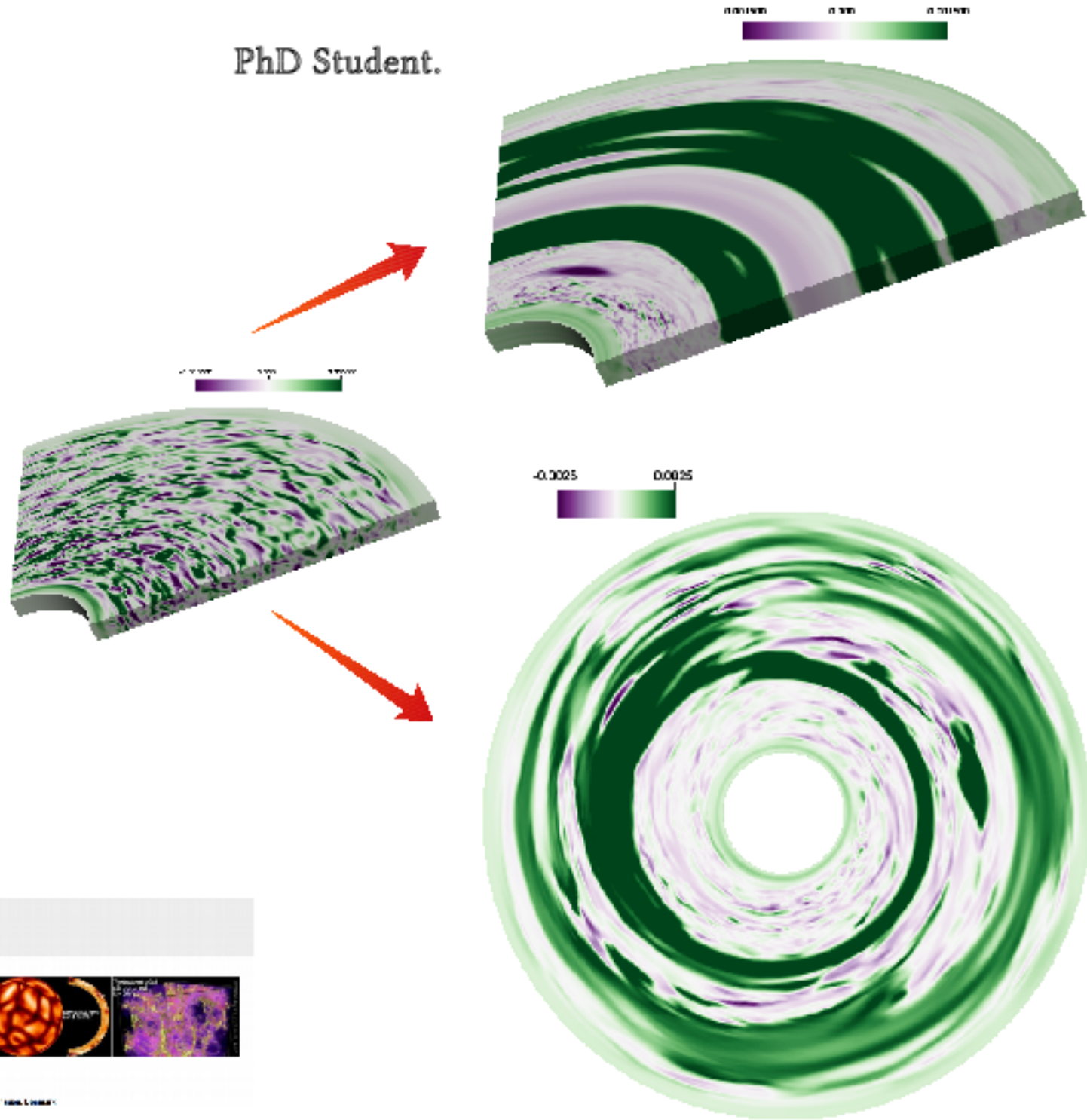

Universität zu Köln

joshi@ph1.uni-koeln.de

Supervisor : Prof. Dr. Stefanie Walch




HALL-MHD
in protoplanetary disks.

FARGO3D

A 3D code for MHD in disks. See the website for more information.
URL: www.fargo3d.org



NIRVANA code



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NBIA Summer School on Astrophysical Plasmas



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28/08 to 01/09/2017

Michael Küffmeier

PhD student

Sup: Åke Nordlund & Troels Haugbølle

University of Copenhagen -

NBI/StarPlan

kueffmeier@gmail.com



Research Interest: Star and Planet Formation

Phd: Sample of stars and protoplanetary disks forming in different environments in a Giant Molecular Cloud

I would like to learn more about:

Processes beyond Giant Molecular Cloud scales

Regimes non-suitable for magnetohydrodynamics

AHMAD LALTI

Second year masters student at the American university of Beirut.

My research work involves the study of secondary instabilities in rotating flows.

Specifically, we are exploring the effect of the third dimension (usually ignored) on the development of a non-axisymmetric secondary instability from an axisymmetric primary instability.

Introduction slide

NBIA summer school

Academic details

- Master student, Astrophysics and Astronomy, Observatory of Paris, France
- PhD at Ecole Normale Supérieure de Lyon, France, starting in september

Master thesis & PhD topic

"Star formation : study of the interstellar dust dynamics"

Personal information

Ugo Lebrouilly

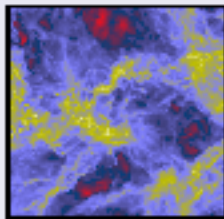
E-mail :

ugo.lebrouilly@ens-lyon.fr

Phone : +33 (0)9 440 023

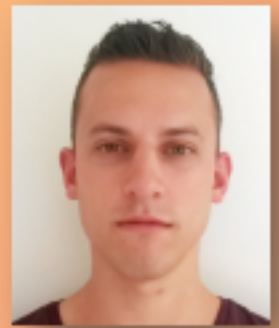


Molecular cloud, RAMSES simulation (OBSPM database)



NBIA Intro slide

Raphaël Mignon-Risse



- M.Sc. Student, Astrophysics & Astronomy
University Paris-Diderot (France)
→ Ph.D. Student, CEA Saclay (France), starting in 10/2017

- Former internships :

Accretion disks in LMXBs & stability

→ Hydrodynamics, microphysics

- Last internship topic:

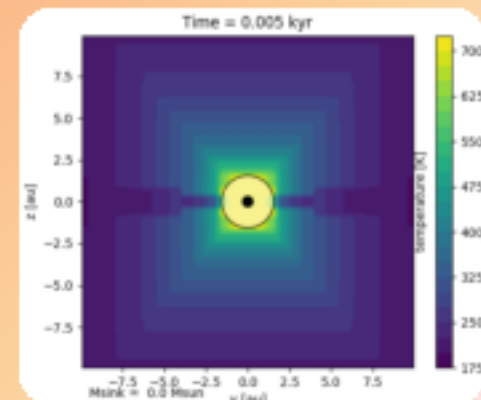
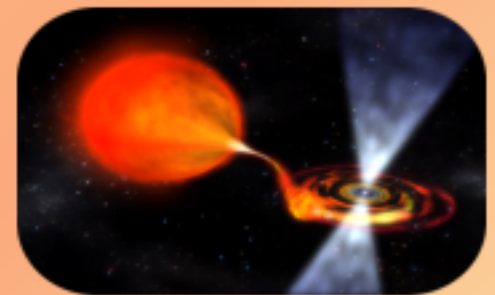
Effect of radiative transfer in massive stars formation

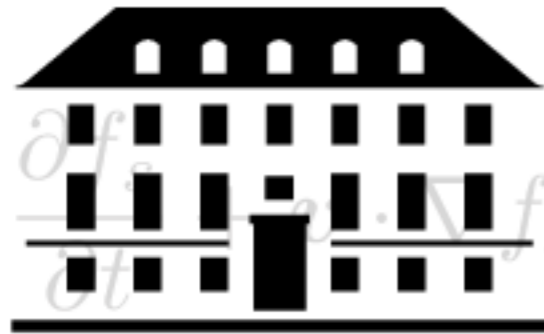
→ Important feedback

- Ph.D. topic:

Study of massive stars formation
via numerical simulations

→ Computational astrophysics, dust environment





The Niels Bohr
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UNIVERSITY OF
COPENHAGEN

$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho v) = 0$
Gopakumar Mohandas

$\rho \left(\frac{\partial v}{\partial t} + v \cdot \nabla v \right) = -\nabla P - \rho \nabla \Phi_g + \mathbf{J} \times \mathbf{B} + \mathbf{F}_{\text{visc}}$
PhD Student

Supervisor: Dr. Martin Pessah

$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$
Research Interests:

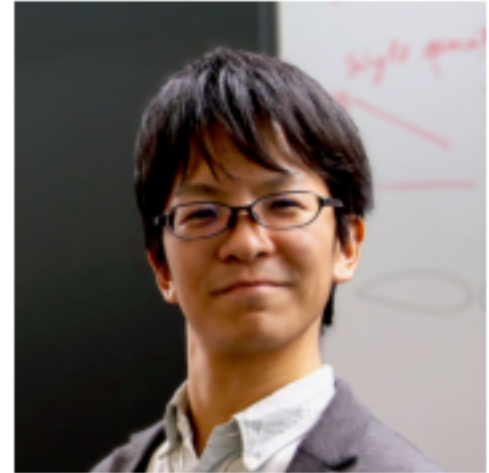
$$\nabla \cdot \mathbf{B} = 0$$

- Astrophysical Fluid Dynamics & Plasma Physics
- Accretion Disk Physics

$\left(\frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla \right) \frac{P}{\rho^\gamma} = \frac{\mathcal{R}}{\mu} \rho T$
• Stability of differentially rotating plasmas

Self-Introduction

- Shoji Mori
- Tokyo Institute of Technology (Japan)
- Research Interest: Planet formation, Protoplanetary disks, MHD, MRI
- Method: MHD simulation
- Like: Beer, Soccer, Snowboarding
- Paper: “Electron Heating in the MRI : Implications for Turbulence Strength in Outer Regions of PPDs ”



me



beer

Website: <http://www.geo.titech.ac.jp/lab/ida/mori/>

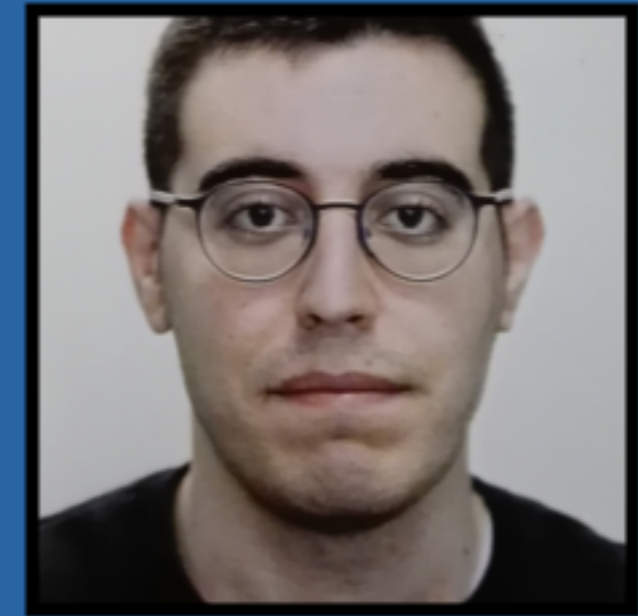
NBIA Summer School on Computational Astrophysics



The Niels Bohr
International Academy

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Copenhagen

- Name: Matteo Pais
- Currently: PhD (2nd yr)
- Advisor: Prof. Christoph Pfrommer
- Institution: University of, Institute, Country
- Role in School: Student
- Contact Info: matteo.Pais@h-its.org



- Research Interests: Galaxy Formation, Galactic dynamos
- Most relevant courses: Fluid Dynamics, Computational Physics, Theoretical astrophysics, General Relativity.
- Programming Languages: Python, C, Mathematica
- Visualization Packages: Paraview
- Programming Experience: 1 year

I would like to learn more about Plasma physics applied to astrophysics and consolidate my knowledge in this field.

Henrik Rhodin

Personal

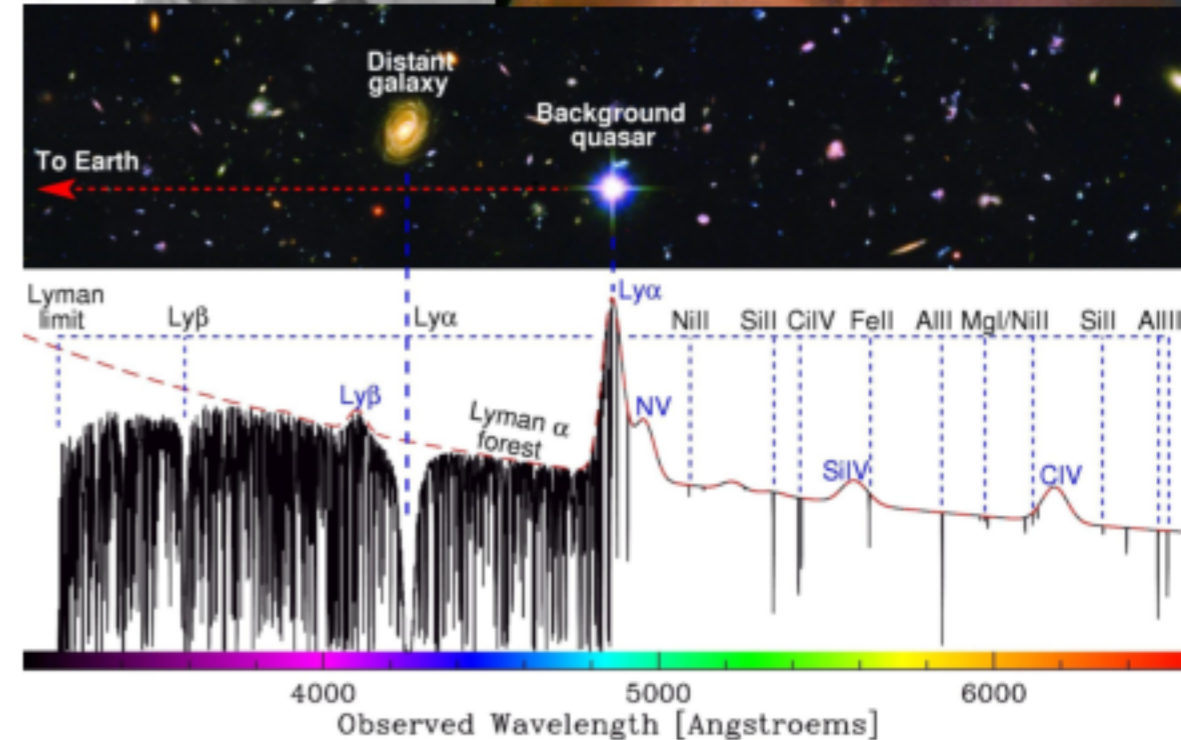
- Born in Sweden, lived in France, studied in Switzerland
- Astrophysics in Edinburgh, Lund, Copenhagen
- Ph.D at Dark Cosmology Centre under supervision of Lise Christensen

Work

- Observational aspects of galaxy evolution
- Understanding the connection between absorbers and their galaxy counterparts
- How are metals dispersed to large distances? - there must be a direct connection to large scale flows of gas

Miscellaneous

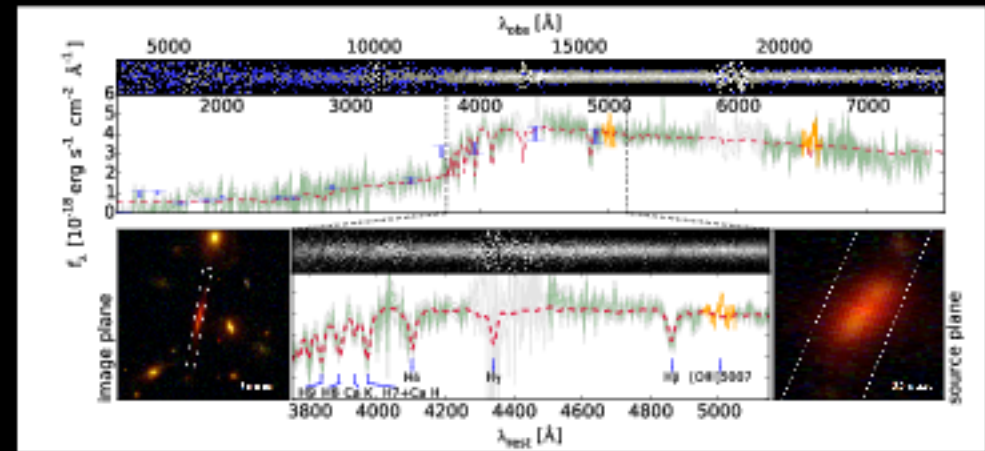
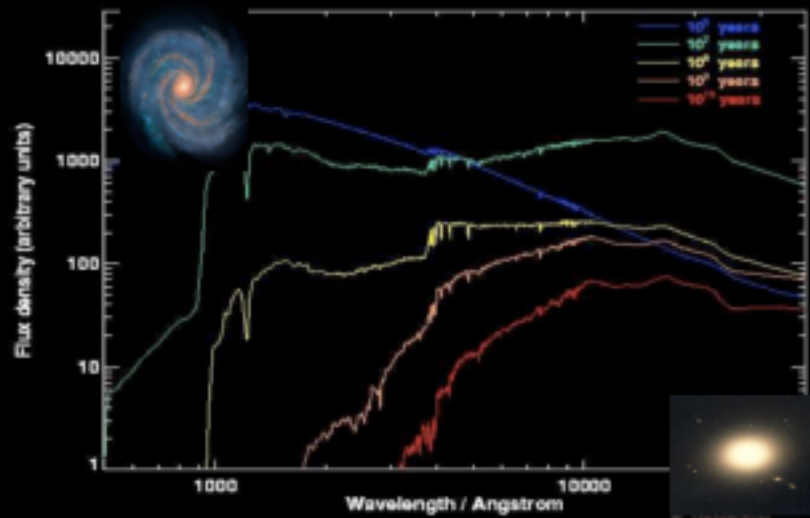
- Enjoy long discussions, good food and company (see *Kräftskiva*)
- Philosophy on the side
- Hate bicycles (and cycling)



The extreme life of massive galaxies:

Live Fast, Die young

Dead old Galaxies



Extreme compactness

Size comparison of Milky Way with ultracompact galaxy



Mikkel Stockmann
PhD Student
Dark Cosmology Centre



TOMAS TAMFAL

PHD STUDENT

UNIVERSITY ZÜRICH

CENTER FOR THEORETICAL ASTROPHYSICS AND COSMOLOGY



**University of
Zurich**^{UZH}

PERSONAL

- BORN IN OPAVA (CZ)
- LIVED IN OPAVA, BAD TÖLZ (GER), MANNHEIM (GER) AND NOW IN ZÜRICH (CH)

EDUCATION

- STUDIED PHYSICS AT THE ETH IN ZÜRICH
- STARTED MY PHD IN DECEMBER 2016

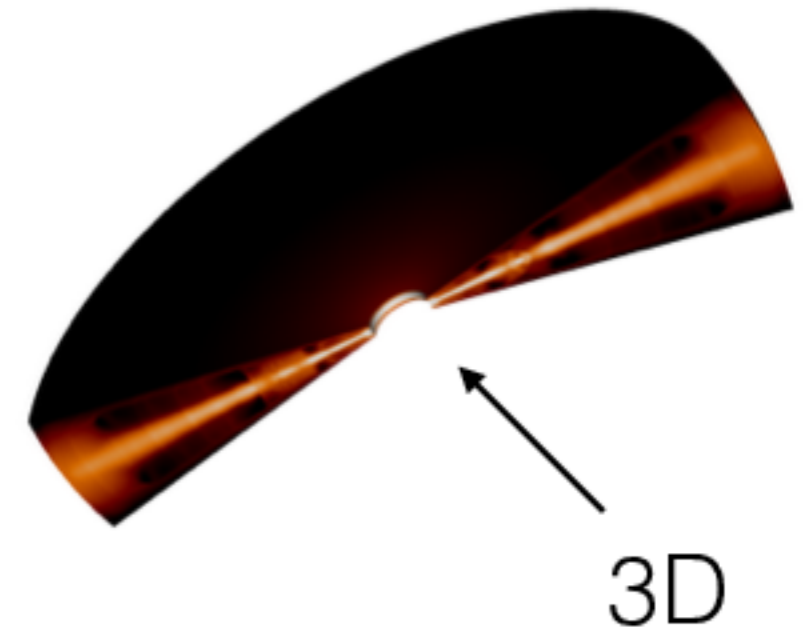
RESEARCH INTERESTS

- DUST COAGULATION IN PROTOPLANETARY DISKS
- BLACK HOLES IN DWARF GALAXIES

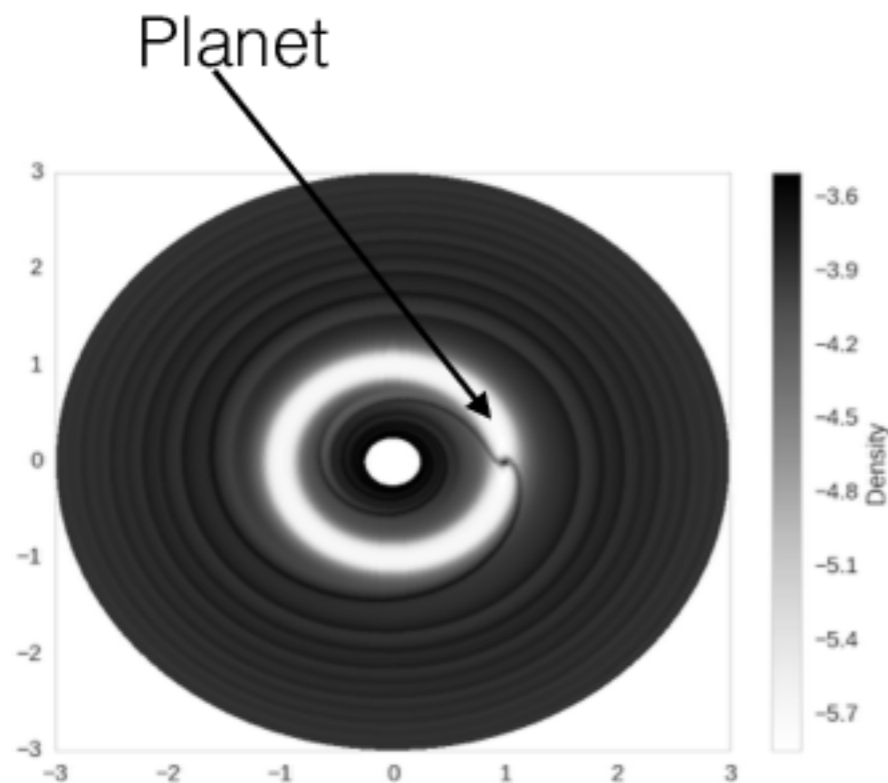
HOBBIES

- SPORT, BOOKS, TRAVEL

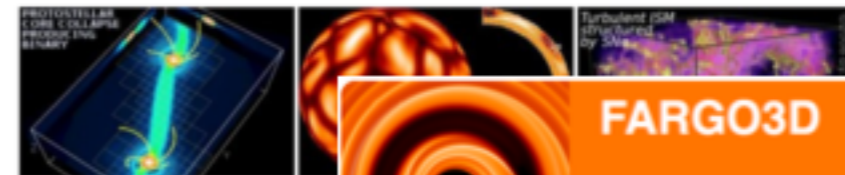
Philipp Weber
PhD Student
NBIA Copenhagen



Dust in Protoplanetary Disks



NIRVANA code



FARGO3D

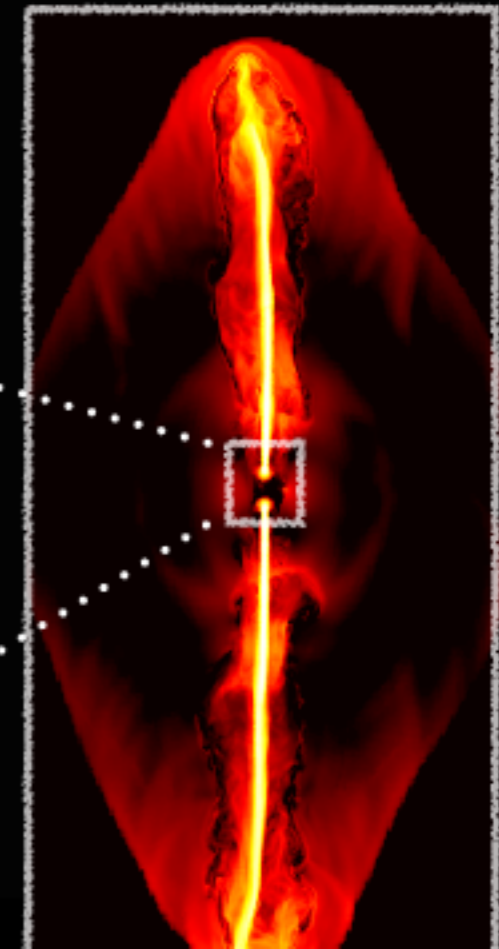
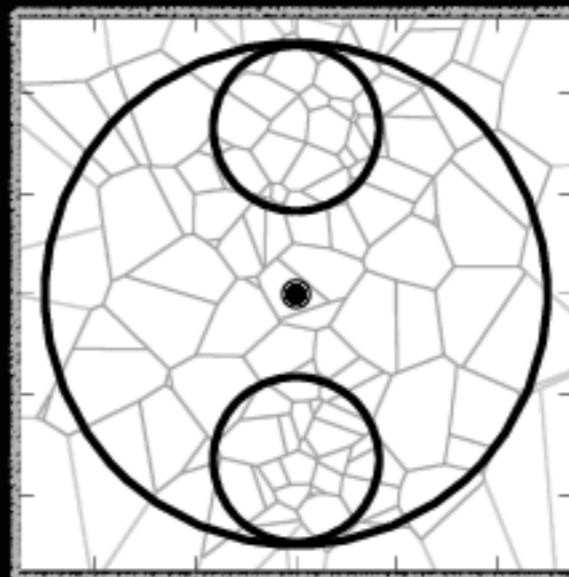
A versatile HD/MHD code that runs on clusters of CPUs or GPUs, with special emphasis on protoplanetary disks.



Rainer Weinberger



- Cosmological simulations of galaxy formation
- Simulating jets from supermassive black holes
- Finite volume MHD simulations with the Arepo code



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- Name: Georg Winner
- Currently: PhD (1st yr)
- Advisor: Prof. Christoph Pfrommer
- Institution: Heidelberg Institute for Theoretical Studies
- Role in School: Student
- Contact Info: georg.winner@h-its.org



- Research Interests: Cosmic Rays, Galaxies, Cosmology
- Most relevant courses: Plasma and Cosmic Ray Physics, Physics of Galaxy Clusters, Fundamentals of Simulation Methods
- Programming Languages: C/C++, Python
- Programming Experience: 2 years

I would like to learn more about
plasma physics and numerical codes