

Scattering Amplitudes from Ambitwistor Strings

Ricardo Monteiro

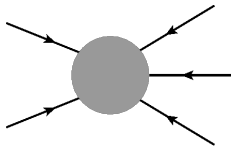
Queen Mary University of London

Current Themes in High Energy Physics and Cosmology,
Copenhagen, 16 August 2018

with Y. Geyer (IAS), L. Mason (Oxford), P. Tourkine (CERN):
1507.00321, 1511.06315, 1607.08887

with Y. Geyer (IAS): 1711.09923, 1805.05344

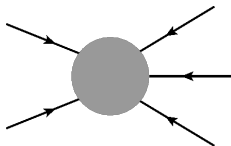
Scattering Amplitudes



Calculable with Feynman diagrams:

- **good**: general, clear physical picture.
- **bad**: inefficient, symmetries obscured.

Scattering Amplitudes



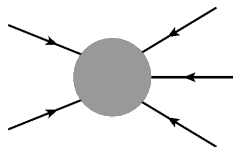
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Modern approaches:

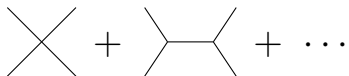
- analyticity, kinematic variables (e.g. on-shell), symmetries.
- relations between theories (e.g. gravity vs gauge theory).
- new formulations of QFTs.

Why Strings?

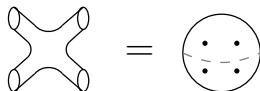
QFT from low-energy string theory: alternative to Feynman expansion.

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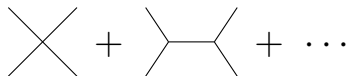
particle scattering
(many Feynman diagrams)



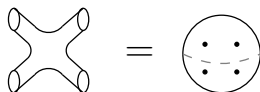
string scattering
(one “world-sheet”, 2D CFT)

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Old idea: calculate QFT amplitudes from string theory.

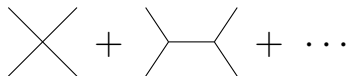
[Green, Schwarz, Brink 82]

Insights: UV divergences, algebraic structure,
gravity (closed strings) v.s. gauge theory (open strings), ...

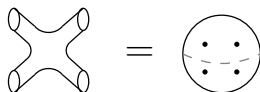


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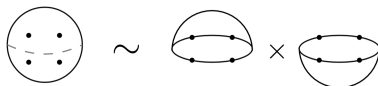


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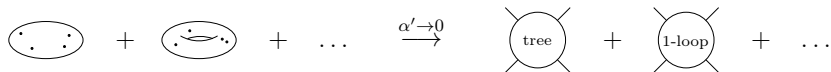
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Hard: higher loop corrections (simpler at low energy), dropping susy.

World-sheet Models of (Massless) QFTs

String theory: field theory is $\alpha' \rightarrow 0$, massive modes decouple, $m_n^2 = c_n/\alpha'$.



Target space is spacetime.

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The diagram shows a sequence of world-sheet surfaces on the left, separated by plus signs and followed by an ellipsis. The first is a sphere with four external legs (two on the left, two on the right). The second is a torus with four external legs. An arrow labeled $\alpha' \rightarrow 0$ points to the right. On the right, a sequence of Feynman diagrams is shown, also separated by plus signs and followed by an ellipsis. The first is a tree-level diagram (a circle with four external legs). The second is a 1-loop diagram (a circle with a loop inside and four external legs).

Target space is spacetime.

Is there **truncated** version just for QFT?

Ambitwistor strings: no α' , only massless states.

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Target space is space of (complex) null geodesics = *ambitwistor space*.

Formulas for amplitudes based on *scattering equations*.

Worksheet Models \longleftrightarrow Scattering Equations

Old story

Twistor string theory [Witten 03] \longrightarrow RSV formula [Roiban, Spradlin, Volovich 04]

$D = 4$. SYM, SUGRA [Hodges, Cachazo, Geyer, Skinner, Mason 12].

Only tree level (unwanted states). [Berkovitz, Witten 04]

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New story

Ambitwistor string theory [Mason, Skinner 13] \longleftarrow CHY formulas [Cachazo, He, Yuan 13-14]

Any D . Many theories of massless particles.

Loop-level progress! Unwanted states absent or projected out.

Outline

- Tree level
- One loop
- Two loops

Tree Level

Scattering Equations

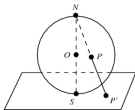
[Cachazo, He, Yuan '13]

[Previous history: Fairlie, Roberts '72, Gross, Mende '88, Witten '04]

Consider n massless particles, $k_i^2 = 0$, $i = 1, \dots, n$, $\sum_{i=1}^n k_i = 0$.

$$E_i = \sum_{j \neq i} \frac{k_i \cdot k_j}{\sigma_i - \sigma_j} = 0, \quad \forall i$$

- kinematic invariants $s_{ij} = (k_i + k_j)^2 = 2 k_i \cdot k_j \rightarrow$ points $\sigma_i \in \mathbb{CP}^1$



$SL(2, \mathbb{C})$ invariance, $\sigma \rightarrow \frac{A\sigma + B}{C\sigma + D}$

- $(n-3)!$ solutions $\sigma_i^{(A)}$.

Scattering Equations

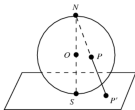
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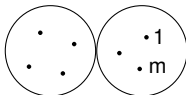
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- factorisation: $(k_1 + \dots + k_m)^2 \rightarrow 0 \iff \sigma_1, \dots, \sigma_m \rightarrow \sigma_*$



[Dolan, Goddard 13]

CHY Formulas

[Cachazo, He, Yuan '13]

Tree-level scattering amplitude:

$$\mathcal{A} = \int_{\mathfrak{M}_{0,n}} d\mu \mathcal{I}$$

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Direct evaluation:

$$A = \sum_{A=1}^{(n-3)!} \frac{\mathcal{I}}{J} \Big|_{\sigma_i = \sigma_i^{(A)}}$$

Scattering equations hard to solve, but no need for that!

[Dolan, Goddard; Cachazo, Gomez; Baadsgaard et al; Huang et al; Sogaard, Zhang; Cardona, Kalousios; Fu et al; . . .]

Examples: Yang-Mills theory and Gravity

[Cachazo, He, Yuan '13]

Ingredients in Yang-Mills (gluon) $e^{ik \cdot x} \epsilon_{\mu} T^a$, Gravity (graviton, ...) $e^{ik \cdot x} \epsilon_{\mu\nu}$:

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gauge invariant $(\epsilon_j^{\mu} \rightarrow \epsilon_j^{\mu} + \alpha k_j^{\mu})$ on $E_j = 0$.

- **colour** (T^{a_i}) :
$$C(a_i, \sigma_i) = \frac{\text{tr}(T^{a_1} T^{a_2} \dots T^{a_n})}{(\sigma_1 - \sigma_2)(\sigma_2 - \sigma_3) \dots (\sigma_n - \sigma_1)} + \dots$$
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Amplitudes are $\mathcal{A} = \int_{\mathfrak{M}_{0,n}} d\mu \mathcal{I}$. Clear n -particle structure in D dimensions!

Yang-Mills theory: $\mathcal{I}_{\text{YM}} = \text{Pf}' M(\epsilon_i) \times \mathcal{C}(a_i)$

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\Rightarrow Gravity \sim YM²

cf. Kawai-Lewellen-Tye relations '86

Bern-Carrasco-Johansson relations '08

Other Examples

Many more theories
of massless particles!

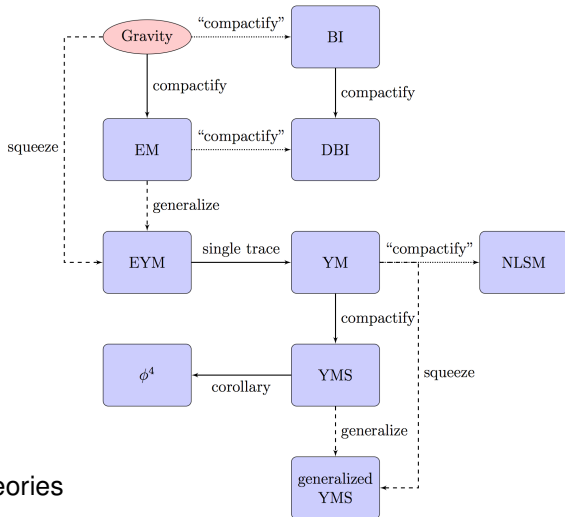
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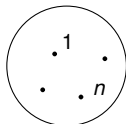
Lessons:

- relations between theories
- messy Feynman rules \neq messy amplitudes

Geometry of Scattering Equations

$SL(2, \mathbb{C})$ invariant differential on \mathbb{CP}^1 :

$$P_\mu(\sigma) = d\sigma \sum_{i=1}^n \frac{k_{i\mu}}{\sigma - \sigma_i}$$



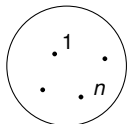
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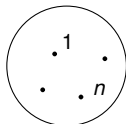
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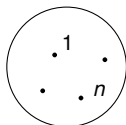
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Any D ? Twistor \rightsquigarrow Ambitwistor space [Mason, Skinner 13]

= space of null geodesics of (complexified) spacetime.

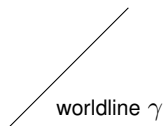
Strings in Ambitwistor Space

[Mason, Skinner 13]

Worldline action for massless particle:

$$S_p = \int_{\gamma} p_{\mu} dx^{\mu} - \frac{1}{2} e p^2 \quad d = d\lambda \partial_{\lambda}$$

- e enforces $p^2 = 0$ ($m^2 = 0$).
- gauge freedom: $\delta x^{\mu} = \alpha p^{\mu}$, $\delta p_{\nu} = 0$, $\delta e = d\alpha$.



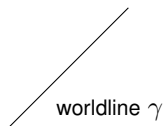
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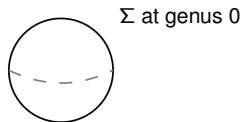
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Chiral complexification: “string”

$$S = \frac{1}{2\pi} \int_{\Sigma} P_{\mu} \bar{\partial} X^{\mu} - \frac{1}{2} e P^2 \quad \bar{\partial} = d\bar{\sigma} \partial_{\bar{\sigma}}$$

- $P_{\mu} = d\sigma p_{\mu}(\sigma)$.
- e enforces $P^2 = 0$, same gauge freedom ($d\alpha \rightsquigarrow \bar{\partial}\alpha$).



Ambitwistor space: (X^{μ}, P_{ν}) with $P^2 = 0$, $(X^{\mu}, P_{\nu}) \sim (X^{\mu} + \alpha P^{\mu}, P_{\nu})$.

Quantisation of Ambitwistor String

[Mason, Skinner 13]

$$\text{Action: } S = \frac{1}{2\pi} \int_{\Sigma} P_{\mu} \bar{\partial} X^{\mu} - \frac{1}{2} e P^2$$

$$\text{Amplitude: } \mathcal{A} = \left\langle \prod_{i=1}^n V_i \right\rangle$$

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Integrate $X^{\mu} \Rightarrow \bar{\partial} P_{\mu} = 2\pi i \sum_j k_{j\mu} \delta^2(\sigma - \sigma_j)$

$\Rightarrow P_{\mu}$ is meromorphic with simple poles at σ_j with residues $k_{j\mu}$.

Quantisation of Ambitwistor String


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Riemann sphere $\mathbb{CP}^1 \Rightarrow P_{\mu} = d\sigma \sum_i \frac{k_{i\mu}}{\sigma - \sigma_i}$ 

$\Rightarrow \text{Res}_{\sigma_i} P^2 = 2E_i = 0$ are scattering equations $\Rightarrow \mathcal{A} = \int_{\mathfrak{M}_{0,n}} d\mu \mathcal{I}$

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
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Combine with world-sheet matter to reproduce various CHY formulas.

[also Ohmori 15; Casali, Geyer, Mason, RM, Roehrig 15]

E.g., add system of fermions \rightsquigarrow Pfaffian.

Ambitwistor Strings vs. Ordinary Strings

Ambitwistor strings:

- chiral, massless states
- $\mathcal{A} = \int (\prod_i d\sigma_i \delta(E_i)) \mathcal{I}_L(\sigma_i) \mathcal{I}_R(\sigma_i)$

Ordinary (closed) strings:

- not chiral, infinite tower of massive states
- $\mathcal{A}_{\text{st}}(\alpha') = \int (\prod_i d\sigma_i d\bar{\sigma}_i) F(|\sigma_i - \sigma_j|) \mathcal{I}_L(\sigma_i) \mathcal{I}_R(\bar{\sigma}_i)$

Formulas: field theory limit is $\mathcal{A}_{\text{st}}(\alpha') \xrightarrow{\alpha' \rightarrow 0} \mathcal{A}$. Non-trivial!

[e.g. Bjerrum-Bohr, Damgaard, Tourkine, Vanhove 14, Mizera 17]

Ambitwistor Strings vs. Ordinary Strings

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Theories:

Surprise understanding from $\alpha' \rightarrow \infty$ with alternative quantisation.

[Siegel 15, +Huang, Yuan 16, Casali, Tourkine 16, +Herfray 17, Azevedo, Jusinkas 17, ...] [Gross, Mende 88]

Open problem type II superstring $\xrightarrow{\alpha' \rightarrow 0}$ type II ambitwistor string.

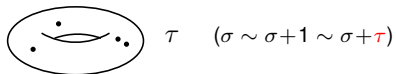
Higher Genus

Geometry at Higher Genus

Good theory. Loop corrections?

[Adamo, Casali, Skinner 13]

Example: one loop needs torus

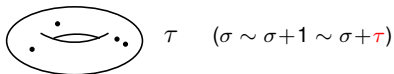


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Recall P_μ is meromorphic with simple poles at σ_i with residues $k_{i\mu}$.

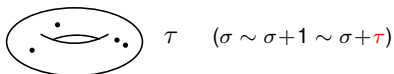
- sphere:
$$P_\mu = d\sigma \sum_i \frac{k_{i\mu}}{\sigma - \sigma_i}$$
- torus:
$$P_\mu = \ell_\mu d\sigma + \sum_i k_{i\mu} \varpi_{i,*}$$

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loop momentum!

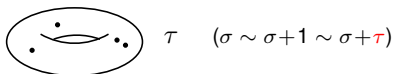
- genus g :
$$P_\mu = \sum_{l=1}^g \ell_\mu^{(l)} \omega_l + \sum_i k_{i\mu} \varpi_{i,*}$$
 (g holomorphic diff's)

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Scattering equations $P^2 = 0$ should localise both σ_i and parameters τ_{IJ} .

Higher-genus surfaces hard (Jacobi θ functions)... Practical?

One Loop

From the Torus to the Nodal Sphere

[Geyer, Mason, RM, Tourkine 15]

Torus scattering equations should localise $\{\sigma_i, \tau\}$, but too hard to solve.

Loop integrand should be easier, like tree level.

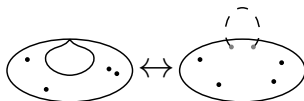
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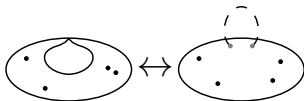


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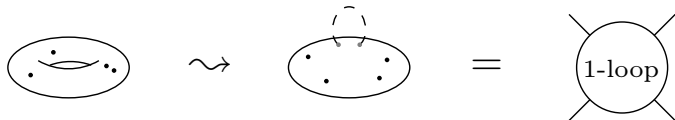
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How to get there? **Residue theorem on τ integration:** localises on $\tau = i\infty$.



Different approach: use elliptic parametrisation.

[Cardona, Gomez 16]

Residue Argument

Torus amplitude: $\mathcal{A}^{(1)} = \int d^D \ell \int_{\mathfrak{M}_{1,n}} d\mu \mathcal{I}^{\text{torus}}$



modulus τ

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Measure: $d\mu \sim d\tau \delta(u) \left(\prod_i d\sigma_i \delta(E_i) \right)$

View as residue integral, e.g., $\int d\sigma \delta(F(\sigma)) \cdots = \oint \frac{d\sigma}{2\pi i \underline{F(\sigma)}} \cdots$

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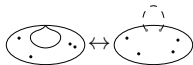
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Use $q = e^{2\pi i \tau}$ and localise at $q = 0$ ($\tau = i\infty$).

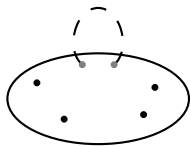


$$\int_{\mathfrak{M}_{1,n}} d\mu \mathcal{I}^{\text{torus}} \sim \oint \frac{dq \prod_i d\sigma_i}{q \underline{u} \prod_i \underline{E_i}} \mathcal{I}^{\text{torus}} = - \oint \frac{dq \prod_i d\sigma_i}{\underline{q} \underline{u} \prod_i \underline{E_i}} \mathcal{I}^{\text{torus}} = - \left[\oint \frac{\prod_i d\sigma_i}{\underline{u} \prod_i \underline{E_i}} \mathcal{I}^{\text{torus}} \right]_{q=0}$$

New One-Loop Formula

[Geyer, Mason, RM, Tourkine 15]

Final result:
$$\mathcal{A}^{(1)} = \int d^D \ell \frac{1}{\ell^2} \int_{\mathcal{M}_{0,n+2}} d\mu^{(1)} \mathcal{I}^{(1)}$$



Like tree level, but now for one-loop integrand!

- two new “particles”: loop momentum insertions $\pm \ell$

- $$P_\mu = d\sigma \left(\frac{\ell_\mu}{\sigma - \sigma_{+\ell}} + \frac{-\ell_\mu}{\sigma - \sigma_{-\ell}} + \sum_i \frac{k_{i\mu}}{\sigma - \sigma_i} \right) = \ell_\mu \omega + d\sigma \sum_i \frac{k_{i\mu}}{\sigma - \sigma_i}$$

- $$\mathcal{I}^{(1)} = \lim_{\tau \rightarrow i\infty} \mathcal{I}^{\text{torus}}$$

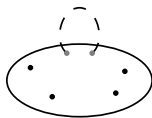
- one-loop scattering equations depend on ℓ_μ

$$P^2 - \ell^2 \omega^2 = 0 \quad \Leftrightarrow \quad P^2 - \ell^2 \omega^2 \quad \text{has no poles at } \sigma_i, \sigma_\pm$$

New Formalism

CHY-type expression for loop integrand:

$$\mathcal{A}^{(1)} = \int d^D \ell \frac{1}{\ell^2} \int_{\mathfrak{M}_{0,n+2}} d\mu^{(1)} \mathcal{I}^{(1)}$$



Previous derivation applied to “type II supergravity in 10D”.

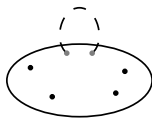
More general?

- Higher genus is very restrictive (modular invariance).
- Perturbative QFT less restricted than a string theory!

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Proposal: generic loop expansion is nodal expansion on sphere.

Look directly for $\mathcal{I}^{(1)}$ in nodal sphere, in any D .

New CHY-type Formulas

[Geyer, Mason, RM, Tourkine 15]

Yang-Mills: $\mathcal{I}_{\text{YM}}^{(1)} = \mathcal{I}_{\text{kin}}^{(1)}(\epsilon) \mathcal{I}_{\text{colour}}^{(1)}$

Gravity: $\mathcal{I}_{\text{Grav}}^{(1)} = \mathcal{I}_{\text{kin}}^{(1)}(\epsilon) \mathcal{I}_{\text{kin}}^{(1)}(\check{\epsilon})$

Double copy!

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(similar to tree level)

Can write “vertex operator” for node. [Roehrig, Skinner 17]

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New formula: $\mathcal{A}^{(1)} = \int d^D \ell \frac{1}{\ell^2} \mathfrak{I}(\ell), \quad \mathfrak{I}(\ell) = \int_{\mathfrak{M}_{0,n+2}} d\mu^{(1)} \mathcal{I}^{(1)}$

Explicit integrand $\mathfrak{I}(\ell)$? **No need to solve the scattering equations.**

[Baadsgaard, Bourjaily, Bjerrum-Bohr, Damgaard, Feng 15; He, Yuan 15, ...]

New Propagator Structure

New formula:
$$\mathcal{A}^{(1)} = \int d^D \ell \frac{1}{\ell^2} \int_{\mathfrak{M}_{0,n+2}} d\mu^{(1)} \mathcal{I}^{(1)}$$

Puzzle! Only one $\frac{1}{\ell^2}$, rest depends only on $\ell \cdot K$, $\ell \cdot \epsilon$...

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Shifted Integrand

- use $\frac{1}{\prod_i D_i} = \sum_i \frac{1}{D_i \prod_{j \neq i} (D_j - D_i)}$, $D_i = (\ell + K_i)^2$
- shift each term $\frac{1}{D_i} \rightarrow \frac{1}{\ell^2}$

Bubble:

$$\frac{1}{\ell^2(\ell + K)^2} = \frac{1}{\ell^2(2\ell \cdot K + K^2)} + \frac{1}{(\ell + K)^2(-2\ell \cdot K - K^2)}$$

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Democratic placement of ℓ in “canonical” integrand.

Good for obtaining loop integrands from trees! “Q-cuts”

[Baadsgaard et al 15, CHY 15]

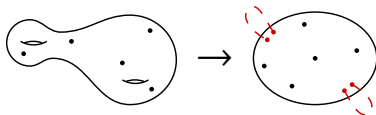
Consequences for colour-kinematics duality.

[He, Schlotterer 16-17, Geyer, RM 17]

Two Loops

Status at Two Loops

Idea:



Done

- heuristic: 4-pts in SUGRA, SYM [Geyer, Mason, RM, Tourkine 16] [genus 2: Adamo, Casali 15]
- detailed: n -pts in SUGRA, SYM [Geyer, RM 18]
genus 2 follows closely superstring [D'Hoker, Phong 01, 05]

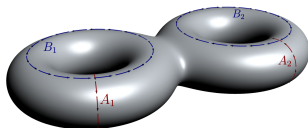
In progress

- n -pts in non-SUSY gravity, YM

Genus 2

There are 2 holomorphic differentials ω_l , $l = 1, 2$.

Pick homology basis:



Normalise ω_l with A-cycles: $\oint_{A_i} \omega_j = \delta_{ij}$

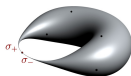
Period matrix: $\tau_{IJ} = \oint_{B_I} \omega_J = \tau_{JI} \rightarrow 3$ modular parameters

$$q_{11} = e^{i\pi\tau_{11}}, \quad q_{22} = e^{i\pi\tau_{22}}, \quad q_{12} = e^{2i\pi\tau_{12}}$$

Bi-nodal sphere: $q_{11} = q_{22} = 0$



Recall genus 1: $q = 0$



From Genus 2 to the Bi-Nodal Sphere

Solution for P_μ :
$$P_\mu = \ell_{1\mu} \omega_1 + \ell_{2\mu} \omega_2 + \sum_i k_{i\mu} \varpi_{i,*}$$

- $P^2(\sigma) = 0$ at genus 2:
- $E_i = 0$ (no poles at σ_i),
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Genus-2 scattering equations would fix $\{\sigma_i, q_{IJ}\}$ but too hard to solve.

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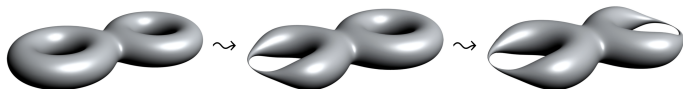
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Residue argument

Genus 1: from $u = 0$ into $q = 0$.

Genus 2: from $u_{11} = u_{22} = 0$ into $q_{11} = q_{22} = 0$. (more subtle)



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Amplitude:
$$\mathcal{A}^{(2)} = \int d^D l_1 d^D l_2 \int_{\mathfrak{M}_{2,n}} d\mu \mathcal{I}^{\text{genus-2}} = \int \frac{d^D l_1 d^D l_2}{(l_1)^2 (l_2)^2} \int_{\mathfrak{M}_{0,n+4}} d\mu^{(2)} \mathcal{I}^{(2)}$$

From Genus 2 to the Bi-Nodal Sphere

New formula: $\mathcal{A}^{(2)} = \int \frac{d^D l_1 d^D l_2}{(l_1)^2 (l_2)^2} \int_{\mathfrak{M}_{0,n+4}} d\mu^{(2)} \mathcal{I}^{(2)}$



Subtle points of residue argument:

- leftover equation of type $u_{IJ} = 0$: $u_{11} + u_{22} + u_{12} = 0$
 $\Rightarrow P^2 - \ell_1^2 \omega_1^2 - \ell_2^2 \omega_2^2 + (\ell_1^2 + \ell_2^2) \omega_1 \omega_2 = 0$ (no poles at $\sigma_i, \sigma_{1\pm}, \sigma_{2\pm}$)

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$$\int_{|q_{12}| \leq 1} \cdots = \int_{|q_{12}| \leq 1} \left(\frac{1}{1-q_{12}} - \frac{q_{12}}{1-q_{12}} \right) \cdots \stackrel{\text{modular}}{=} \int_{q_{12}} \frac{1}{1-q_{12}} \cdots$$

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$$\mathcal{I}_{\text{SUGRA}}^{(2)} = \frac{1}{1-q_{12}} \mathcal{I}_{\text{kin}}^{(2)}(\epsilon) \mathcal{I}_{\text{kin}}^{(2)}(\tilde{\epsilon}) \rightsquigarrow \mathcal{I}_{\text{SYM}}^{(2)} = \mathcal{I}_{\text{kin}}^{(2)}(\epsilon) \mathcal{I}'_{\text{colour}}^{(2)}$$

Get $\mathcal{I}'_{\text{colour}}^{(2)}$ from current algebra on bi-nodal sphere (cf. heterotic string).

All Genus Plausibility



genus g

$$\dim_{\mathbb{C}}(\mathfrak{M}_{g,n}) = n + 3g - 3$$

degenerations

$-g$

g -nodal sphere

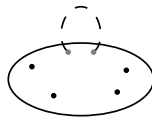
$$\dim_{\mathbb{C}}(\mathfrak{M}_{0,n+2g}) = (n + 2g) - 3$$

Conclusion

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- Ambitwistor strings describe perturbative QFTs.
- Formulas for YM and gravity, SUSY/no-SUSY.

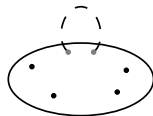
loop expansion = nodal expansion



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Many open questions

- All-loop story?
- Double copy from gauge theory to gravity?
- 4D formalism? Efficient manipulation, inc. loop integration?
- Beyond scattering amplitudes? Beyond perturbation theory?
- Insights into string theory?

Extra slides

KLT Relations: String Theory Origin

[Kawai, Lewellen, Tye '86]

Vertex operators: $V_{\text{closed}}(\epsilon^{\mu\nu} = \epsilon^\mu \tilde{\epsilon}^\nu) \sim V_{\text{open}}(\epsilon^\mu) \bar{V}_{\text{open}}(\tilde{\epsilon}^\nu)$ $s_{ij} = (k_i + k_j)^2$

$$\mathcal{A}_3^{\text{grav}} = A^{\text{YM}}(123) \tilde{A}^{\text{YM}}(123) \quad \mathcal{A}_4^{\text{grav}} = \frac{\sin \pi \alpha' s_{12}}{\pi \alpha'} A^{\text{YM}}(1234) \tilde{A}^{\text{YM}}(1243)$$

Field theory limit is $\alpha' \rightarrow 0$.

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In general (tree level)

[Bern, Dixon, Perelstein, Rozowsky '98]

$$\mathcal{A}_n^{\text{grav}} = \sum_{P_n, P'_n} A^{\text{YM}}(P_n) S_{\text{KLT}}[P_n, P'_n] \tilde{A}^{\text{YM}}(P'_n) \quad S_{\text{KLT}} \sim s_{ij}^{n-3}$$

Useful at loop level via unitarity cuts.

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Recall YM colour decomposition: colour traces or colour factors.

$$\mathcal{A}_n^{\text{YM}} = \sum_{\text{non cyclic}} A^{\text{YM}}(1, 2, \dots, n) \text{tr}(T^{a_1} T^{a_2} \dots T^{a_n}) = \sum_{\alpha \in \text{cubic}} N_\alpha c_\alpha$$

with $c_\alpha = f^{abc} f^{\dots} \dots f^{\dots}$, $f^{abc} = \text{tr}([T^a, T^b] T^c)$,

but Jacobi identities: $c_\alpha \pm c_\beta \pm c_\gamma = 0$

Summary of Double Copy

$$\mathcal{A}_{\text{grav}}(\epsilon_i^{\mu\nu}) \sim (\text{prop})^{-1} \mathcal{A}_{\text{YM}}(\epsilon_i^\mu) \times \mathcal{A}_{\text{YM}}(\tilde{\epsilon}_i^\nu) \quad \Big| \text{colour stripped}$$

KLT relations

$$\mathcal{A}_{\text{grav}} = \sum_{P_n, P'_n} \mathcal{A}_{\text{YM}}(\epsilon, P_n) S_{\text{KLT}}[P_n, P'_n] \mathcal{A}_{\text{YM}}(\tilde{\epsilon}, P'_n)$$

BCJ double copy

$$\mathcal{A}_{\text{YM}} = \sum_{\alpha \in \text{cubic}} \frac{n_\alpha(\epsilon) c_\alpha}{D_\alpha} \quad \mathcal{A}_{\text{grav}} = \sum_{\alpha \in \text{cubic}} \frac{n_\alpha(\epsilon) n_\alpha(\tilde{\epsilon})}{D_\alpha}$$

CHY formulas

$$\mathcal{A} = \int d\mu \mathcal{I} \quad \mathcal{I}_{\text{YM}} = \text{Pf}' M(\epsilon) \times \mathcal{C} \quad \mathcal{I}_{\text{grav}} = \text{Pf}' M(\epsilon) \times \text{Pf}' M(\tilde{\epsilon})$$